

# Holography: a theory lab for strong coupling

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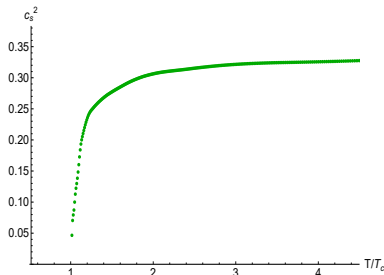
*"On the other hand, it might be helpful to explore a theory that deviates from the unknown truth in the opposite direction from that of the conventional theory."*

*Enrico Fermi*

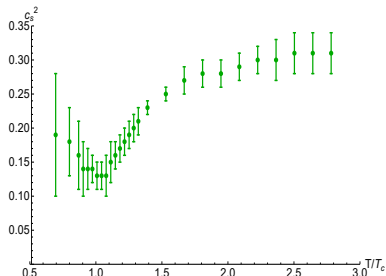


# Phase structure at strong coupling

- Systems at strong coupling exhibit various phase structures
- Pure gluon system  $\rightarrow$  1<sup>st</sup> order phase transition (left)
- Gluons + quarks  $\rightarrow$  smooth crossover (right)



G. Boyd *et al.* Nucl. Phys. B **469**, 419 (1996)



S. Borsanyi *et al.* JHEP **1009**, 073 (2010)

# Phase structure at strong coupling

- Lattice methods do not reach real time dynamics easily
- Compute the spectrum of linearized perturbations
- Compute transport coefficients and non-hydrodynamic modes
- Compute the non-linear time evolution
- Investigate the dynamical appearance of diverse phases
- Check linear and non-linear stability

- Does spinodal instability appear for a holographic system with a 1<sup>st</sup> order phase transition?
- Does the phase separation effect appear dynamically?
- Are there black hole solution with inhomogeneous horizons?
- How do non-hydrodynamic degrees of freedom behave in the critical region?
- Do diffusive modes appear?

- **Top-down construction**

$\mathcal{N} = 4$  broken to  $\mathcal{N} = 2^*$  SUSY theory. Known, but complicated dual gravity description

A. Buchel, S. Deakin, P. Kerner, J. T. Liu, Nucl. Phys. B **784**, 72 (2007)

- **Bottom-up construction**

Assuming AdS/CFT dictionary, try to model gravity+matter background to approach as closely as possible to your favourite physics

J. Erlich, *et al.* Phys. Rev. Lett. **95**, 261602 (2005)

S. S. Gubser, A. Nellore, Phys. Rev. D **78** (2008) 086007

U. Gürsoy, *et al.* JHEP **0905**, 033 (2009)

# Holographic set-up $\rightarrow$ bottom-up approach

- **Boundary:** add a source for an operator  $O_\phi$  in a  $\text{CFT}_d$

$$\mathcal{L} = \mathcal{L}_{\text{CFT}} + \Lambda^{d-\Delta} O_\phi$$

- **Bulk:** a gravity-scalar system in  $D = d + 1$

$$S = \frac{1}{2\kappa_D^2} \int_{\mathcal{M}} d^D x \sqrt{-g} \left[ R - \frac{1}{2} (\partial\phi)^2 - V(\phi) \right] + S_{\text{GH}} + S_{\text{ct}}$$

with the potential

$$V(\phi) = 2\Lambda_C (1 + a\phi^2)^{1/4} \cosh(\gamma\phi) + b_2 \phi^2 + b_4 \phi^4 + b_6 \phi^6$$

- $\Lambda_C = -d(d-1)/2$  is the cosmological constant

U. Gürsoy, *et al.* JHEP **0905**, 033 (2009)

S. S. Gubser, A. Nellore, Phys. Rev. D **78** (2008) 086007

# Equilibrium configurations

- Metric ansatz for a homogeneous configuration

$$ds^2 = e^{2A(r)} (-h(r)dt^2 + d\vec{x}^2) - 2e^{A(r)+B(r)} drdt$$

with  $\phi(r) = r$  the holographic coordinate

- Solve Einstein+matter equations
- The event horizon:  $h(r_H) = 0$
- Entropy and Hawking temperature

$$s = \frac{2\pi}{\kappa_D^2} e^{(d-1)A(r_H)}$$

$$T = \frac{e^{A(r_H)+B(r_H)} |V'(r_H)|}{4\pi}$$

- The free energy is defined by the action  $F = TS_{\text{on-shell}}$

U. Gürsoy, *et al.* JHEP **0905**, 033 (2009)

S. S. Gubser, A. Nellore, Phys. Rev. D **78** (2008) 086007

# Phase transitions in holography

- Finite  $T$  states correspond to various black hole solutions in the dual spacetime
- Phase structure is determined by the choice of  $a, \gamma$  and  $b_2, b_4, b_6$ , coefficients of  $V(\phi)$
- With  $a \neq 0$  confining models (IHQCD)
- It is possible to tune parameters to mimic
  - crossover e.g. QCD
  - 1<sup>st</sup> order phase transition e.g. pure gluon systems
  - 2<sup>nd</sup> order phase transition

U. Gürsoy, *et.al.* JHEP **0905**, 033 (2009)

S. S. Gubser, A. Nellore, Phys. Rev. D **78** (2008) 086007



# Equation of State (EoS)

- The free energy is defined by the action

$$F = TS_{\text{on-shell}}$$

- Configurations characterized by the horizon radius
- Condition for the 1<sup>st</sup> order phase transition  $F_{\text{BH}_1} = F_{\text{BH}_2}$
- Similar to Hawking-Page transition of pure *AdS*

S. W. Hawking, D. N. Page, Commun. Math. Phys. **87**, 577 (1983)

E. Witten, Adv. Theor. Math. Phys. **2**, 505 (1998)

# Example I: First order phase transition

- In  $d = 3 + 1$  we choose

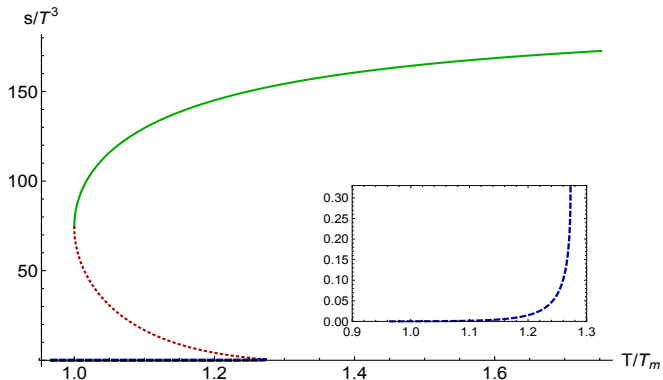
$$V_{1\text{st}}(\phi) = -12(1 + a\phi^2)^{1/4} \cosh(\gamma\phi) + b_2\phi^2 + b_4\phi^4 + b_6\phi^6$$

with  $a = 0$ ,  $\gamma = \sqrt{7/12}$ ,  $b_2 = 2.5$ ,  $b_4 = b_6 = 0$

- Conformal dimension of the scalar operator is  $\Delta = 3.41$
- Transition between two different black hole solutions
- An example of holographic 1<sup>st</sup> order phase transition
- No known physical counterpart

# Example I: First order phase transition

- There exists a critical temperature  $T_c \simeq 1.05 T_m$
- For the unstable region (red dashed line) we have  $c_s^2 < 0$



## Example II: Confining model IHQCD

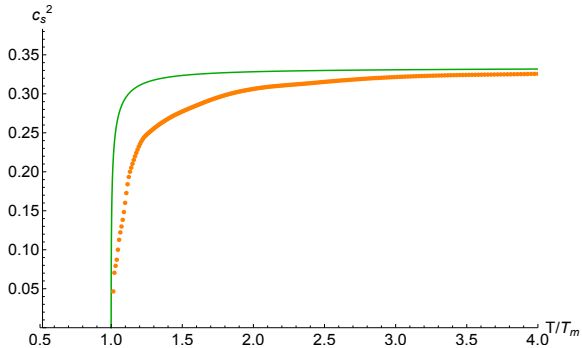
- In  $d = 3 + 1$  we choose

$$V_{\text{IHQCD}}(\phi) = -12(1+a\phi^2)^{1/4} \cosh(\gamma\phi) + b_2\phi^2 + b_4\phi^4 + b_6\phi^6$$

with  $a = 1$ ,  $\gamma = \sqrt{2/3}$ ,  $b_2 = 6.25$ ,  $b_4 = b_6 = 0$

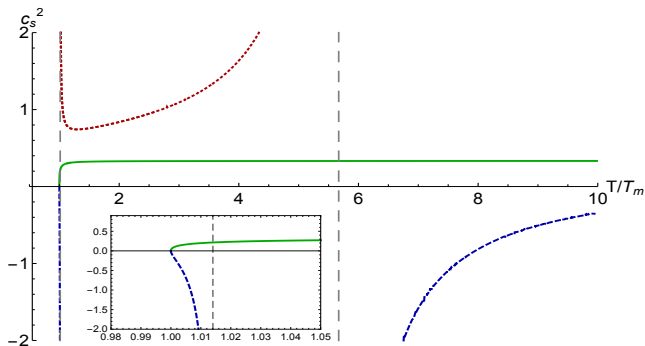
- Conformal dimension of the scalar operator is  $\Delta = 3.58$
- Transition between black hole and horizon-less geometry  
S. W. Hawking, D. N. Page, Commun. Math. Phys. **87**, 577 (1983)
- System motivated by the gluon dynamics
- Linear confinement in the meson spectrum, i.e.  $m_n^2 \sim n$

# Example II: Confining model IHQCD



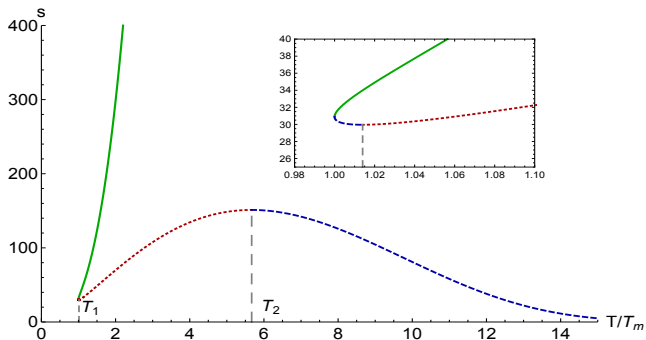
G. Boyd *et.al.* Nucl. Phys. B **469**, 419 (1996)

## Example II: Full holographic scan



- Below  $T_m$  no black hole solution exists
- Green line - stability region, blue dashed line - spinodal region
- Red dashed line - „dynamically unstable” region

## Example II: Entropy density



- The dynamically unstable region for  $T_1 < T < T_2$
- The limiting points  $T_1 = 1.014 T_m$  and  $T_2 = 5.67 T_m$

# Linear response and Quasinormal modes

- Perturb the system  $\mathcal{L} = \mathcal{L}_0 + h_{ij}\delta^3(x)\delta(t)T^{ij}(x)$  the response is the *retarded Green's* function

$$G_R(\omega, k) \propto i \int dt d^3x \theta(t) e^{ikx - i\omega t} \langle [T_{ij}(x, t), T_{kl}(0)] \rangle$$

- *Quasinormal modes*, i.e., solutions of linearized fluctuation equations correspond to poles of holographic retarded Green's functions. In general

$$\omega_n(k) = \Omega_n(k) - i\Gamma_n(k)$$

where  $n = 1, 2, 3, \dots$   $\Omega_n(k)$ —oscillation frequency,  
 $\Gamma_n(k)$ —damping rate. Stable modes have  $\Gamma_n(k) > 0$ .



# Quasinormal modes

- Consider perturbations of the  $5D$  black hole background

$$g_{ab} = g_{ab}^{\text{BH}} + h_{ab}(r)e^{ikz-i\omega t}, \quad \Phi = \Phi^{\text{BH}} + \phi(r)e^{ikz-i\omega t}$$

- QNMs** are the solutions of linearized fluctuation equations that correspond to poles of holographic retarded Green's functions
- In general

$$\omega_n(k) = \Omega_n(k) - i\Gamma_n(k)$$

where  $n = 1, 2, 3, \dots$   $\Omega_n(k)$ —oscillation frequency,  
 $\Gamma_n(k)$ —damping rate.

- Stable** modes have  $\Gamma_n(k) > 0$
- A convenient normalization is:  $q = \frac{k}{2\pi T}$ ,  $\varpi = \frac{\omega}{2\pi T}$

P. K. Kovtun, A. O. Starinets, Phys. Rev. D **72**, 086009 (2005)

# Gauge invariant perturbations

- Gauge invariance at linearized level

$$h_{ab} \mapsto h_{ab} - \nabla_a \xi_b - \nabla_b \xi_a, \quad \phi \mapsto \phi - \xi^a \nabla_a \phi$$

- Five independent channels
  - sound and non-CFT modes, coupled

$$Z_1(h_{tt}, h_{tz}, h_{zz}, h_{xx} + h_{yy}), \quad Z_2(\phi, h_{xx} + h_{yy})$$

- twofold degenerated shear mode

$$Z_3(h_{xz}, h_{xt}), \quad Z_3(h_{yz}, h_{yt})$$

- scalar mode (EOM similar to external scalar field EOM)

$$Z_4(h_{xy})$$

- In our coordinate system  $\Phi(r) = r$  in the background

S. S. Gubser, A. Nellore, Phys. Rev. D **78** (2008) 086007

- The near-boundary ( $r \rightarrow 0$ ) behaviour

$$Z_1(r) \sim A_1 + B_1 r^{\frac{d}{d-\Delta}}, \quad Z_2(r) \sim A_2 r + B_2 r^{\frac{\Delta}{d-\Delta}}$$

imposes  $A_1 = A_2 = 0$  boundary condition

- Other modes have Dirichlet BCs at the conformal boundary
- Ingoing boundary conditions at the horizon
- Chebyshev discretization with high numerical precision
- Newton-Raphson method for the background
- Generalized eigenvalue problem for QNMs, i.e.  $\text{Det}M(\omega) = 0$

- Hydrodynamic mode is defined by

$$\lim_{k \rightarrow 0} \omega_H(k) = 0$$

- The sound mode

$$\omega(k) = \pm c_s k - \frac{i}{2T} \left( \frac{4}{3} \frac{\eta}{s} + \frac{\zeta}{s} \right) k^2 + O(k^3)$$

$\eta$ —shear viscosity,  $\zeta$ —bulk viscosity,  $s$ —entropy density,  
 $c_s$ —speed of sound,  $T$ —temperature

- In holographic models also *non-hydrodynamic* modes are present

P. K. Kovtun, A. O. Starinets, Phys. Rev. D **72**, 086009 (2005)

M. P. Heller *et al.* Phys. Rev. Lett. **110**, no. 21, 211602 (2013)

# Non-hydrodynamic degrees of freedom (DOF)

- Higher QNM's represent non-hydro DOF in QFT
- **Crossing** of the hydro and non-hydro modes happens when the hydrodynamic mode is more damped than the higher QNMs
- In the CFT this happens *only* in the shear mode for  
 $k \simeq 1.3(2\pi)T$  I. Amado *et.al.* JHEP **0807**, 133 (2008)
- Non-conformality affects the crossing phenomena in *qualitative* and *quantitative* way
- In  $\mathcal{N} = 4$  SYM case the non-hydro QNM's are linked with high order transport coefficients

M. P. Heller, *et al.* Phys. Rev. Lett. **110**, no. 21, 211602 (2013)

# Spinodal instability $\rightarrow$ linear response theory

- When  $c_s^2 < 0$  we have purely damped hydro-modes

$$\omega \approx \pm i|c_s| k - \frac{i}{2T} \left( \frac{4}{3} \frac{\eta}{s} + \frac{\zeta}{s} \right) k^2 = \pm i|c_s| k - i\Gamma_s k^2$$

so for small enough  $k$  we have  $\text{Im } \omega > 0$

- This mode is present for a finite range of  $0 < k < k_{\text{max}}$
- The maximum momentum for the unstable mode is  $k_{\text{max}} = |c_s|/\Gamma_s$
- This appears for systems with a 1<sup>st</sup> order phase transition; *spinodal* instability

P. Chomaz, et al. M. Colonna, J. Randrup, Phys. Rept. **389**, 263 (2004)

# Examples of spinodal instabilities

- **Water**: superheated liquid and supercooled vapour

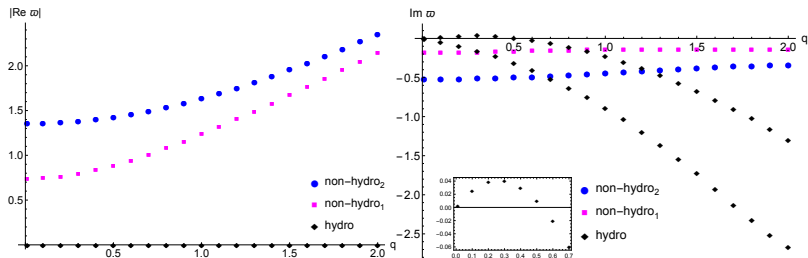


- Spinodal instability in nuclear matter liquid-gas transition  
**Nuclear multifragmentation: Xe+Sn @ 32 MeV/A**

B. Borderie *et al.* Phys. Rev. Lett. **86**, 3252 (2001)

P. Chomaz, M. Colonna, J. Randrup, Phys. Rept. **389**, 263 (2004)

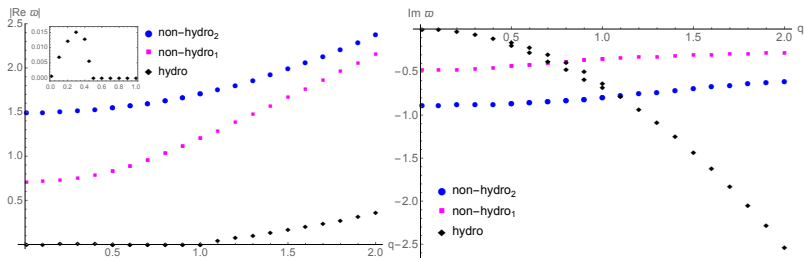
# Example I: Holographic spinodal instability



- Modes for  $T \simeq 1.06 T_m$  where  $c_s^2 \simeq -0.1$
- The hydrodynamic mode follows the thermodynamic instability
- Scale of the bubble =  $k$  for which  $\text{Im } \omega$  is maximal
- The maximal value of  $\text{Im } \omega$  is called the **growth rate**
- Non-hydrodynamic modes have weak  $k$ -dependence



# Example I: Diffusive modes

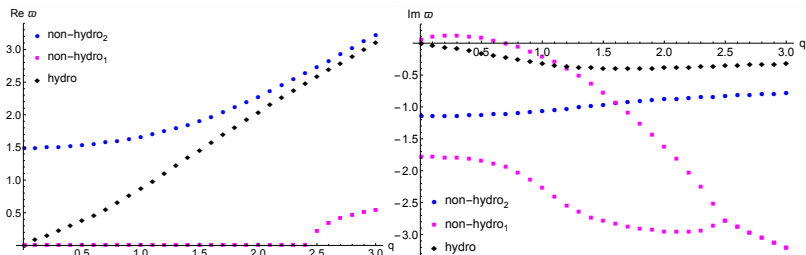


- Modes for  $T \simeq 1.00004 T_m$
- $\text{Re } \omega = 0$  for a range of momenta (here  $0.5 < q < 1$ )
- The sound mode becomes nonpropagating for this range of  $q$
- Crossing between hydro sound mode and non-hydro mode

R. A. Janik, J. J., H. Soltanpanahi, Phys. Rev. Lett. **117**, no. 9, 091603 (2016)

U. Gürsoy, S. Lin, E. Shuryak, Phys. Rev. D **88**, no. 10, 105021 (2013)

# Example II: Dynamical instability



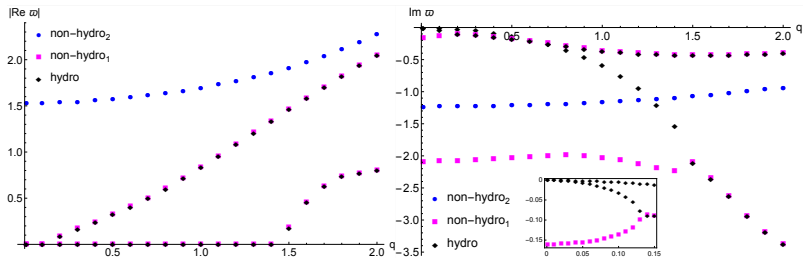
- Quasinormal modes at  $T = 1.027 T_m$
- System displays dynamical instability in spite of thermodynamical stability!
- The system is unstable against uniform ( $k = 0$ ) perturbations
- Possible implications for thermalization time

U. Gursoy, *et al.* Phys. Rev. D **94**, no. 6, 061901 (2016)

R. A. Janik, J. J., H. Soltanpanahi, Phys. Rev. Lett. **117**, no. 9, 091603 (2016)



# Example II: Ultralocality violation at $T = T_m$



- Small gap between hydro and non-hydro DOF at low  $q$
- Hydro and non hydro joining at  $q_J \simeq 0.14$  and  $q_J \simeq 1.5$
- At  $q_J$  the real part develops
- This structure is unique for  $T = T_m$

R. A. Janik, J. J., H. Soltanpanahi, JHEP 1606, 047 (2016)

# The crossing landscape

potential	sound $q_c$	shear $q_c$	$c_s^2$	$\zeta/s$
$V_{\text{QCD}}$	0.8	1.1	0.124	0.041
$V_{2\text{nd}}$	0.55	0.9	0.0	0.061
$V_{1\text{st}}$	0.8	1.15	0.0	0.060
$V_{\text{IHQCD}}$	0.14	1.25	0.0	0.512

- The crossing momentum  $q_c$  at  $T = T_c$  ( $V_{\text{QCD}}$ ,  $V_{2\text{nd}}$ ) and  $T = T_m$  ( $V_{1\text{st}}$ ,  $V_{\text{IHQCD}}$ )
- In contrast to the CFT case crossing happens in both channels
- Applicability of hydro is restricted near the transition (especially in the IHQCD model)

# Summary

- Thermodynamic instability  $\rightarrow$  dynamical instability
- Converse doesn't seem to be true!

U. Gursoy, *et al.* Phys. Rev. D **94**, no. 6, 061901 (2016)

- Non-trivial phase structure limits the applicability of hydrodynamics
- In most cases non-hydro degrees of freedom have very weak dependence on  $k \rightarrow$  „*ultralocality*”
- Non-linear time evolution  $\rightarrow$  the phase transition concept beyond the notion of thermodynamic equilibrium
- Spinodal instability in a evolving system

R. A. Janik, J. J. H. Soltanpanahi, Phys. Rev. Lett. **119**, no. 26, 261601 (2017)