



KM3NeT

Opens a new window on our universe

Data Acquisition and Trigger in KM3NeT

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Dubna, Russia

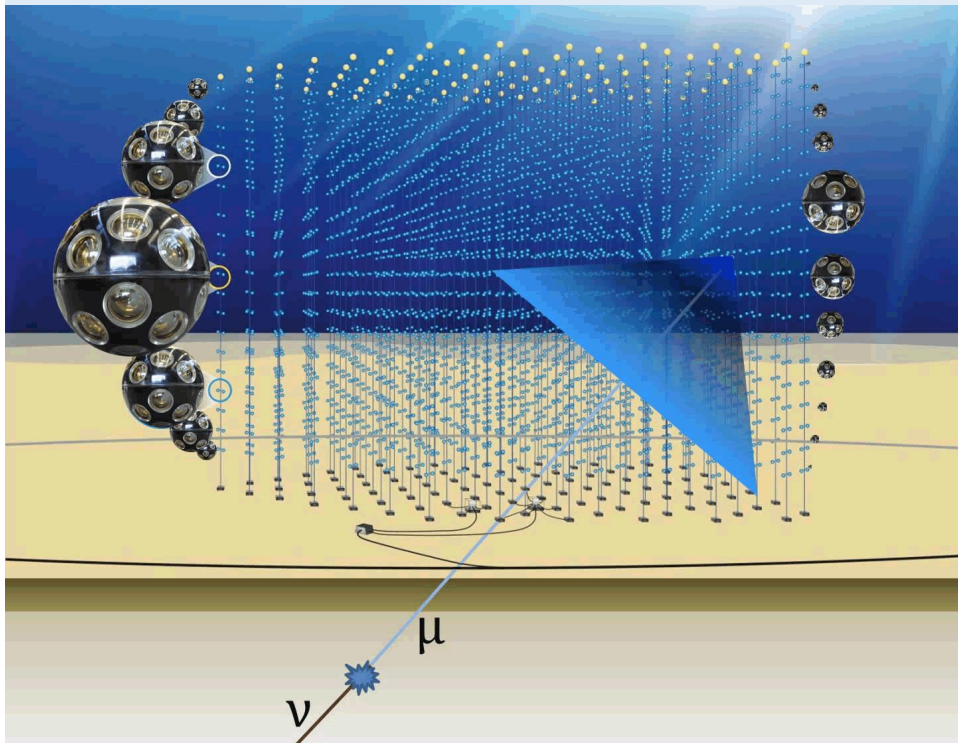


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KM3NeT in short

Cherenkov light from the charged products of neutrino interactions in sea-water are detected by a sparse array of photo-multiplier tubes



31 3" PMTs in glass sphere
Multi-PMT DOM

ARCA
 $\sim 1 \text{ km}^3$
Astrophysical neutrinos (TeV-PeV)
Italian Site

ORCA
 $\sim 5.7 \text{ Mton}$
Atmospheric neutrinos ($\sim 1\text{-}100 \text{ GeV}$)
Neutrino mass hierarchy
French Site

DAQ Constraints & Requirements

Size :
115 Detection Units per block
18 DOMs/DU
31 PMTs/DOM

~64k PMTs

Background photons:

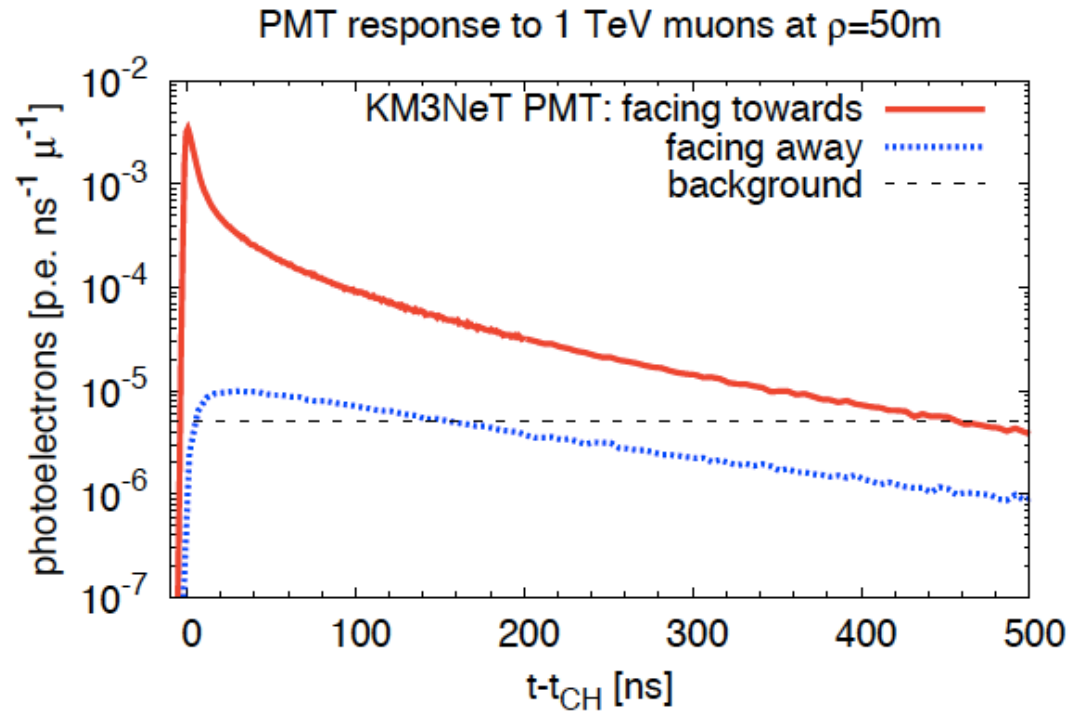
⁴⁰ K decay	~7 kHz/PMT constant
Bioluminescence	~100 kHz/PMT bursts
Atmospheric muons	~100 Hz/block

Expected signal (reconstruction level) :
ARCA : 3 cosmological neutrinos/day
ORCA : 50.000 atmospheric neutrinos/year

Signal to noise
< 10⁻⁶

Requirements:
Infrastructure to transfer all data to shore for processing
Rejecting the overwhelming background
Don't loose the sparse signals
...and do this in a hostile and dynamic environment – the Sea

But why transfer all data?? – All data to shore



Photons from neutrino interaction products have long-range nanosecond correlations, due to favorable optical properties of sea-water.

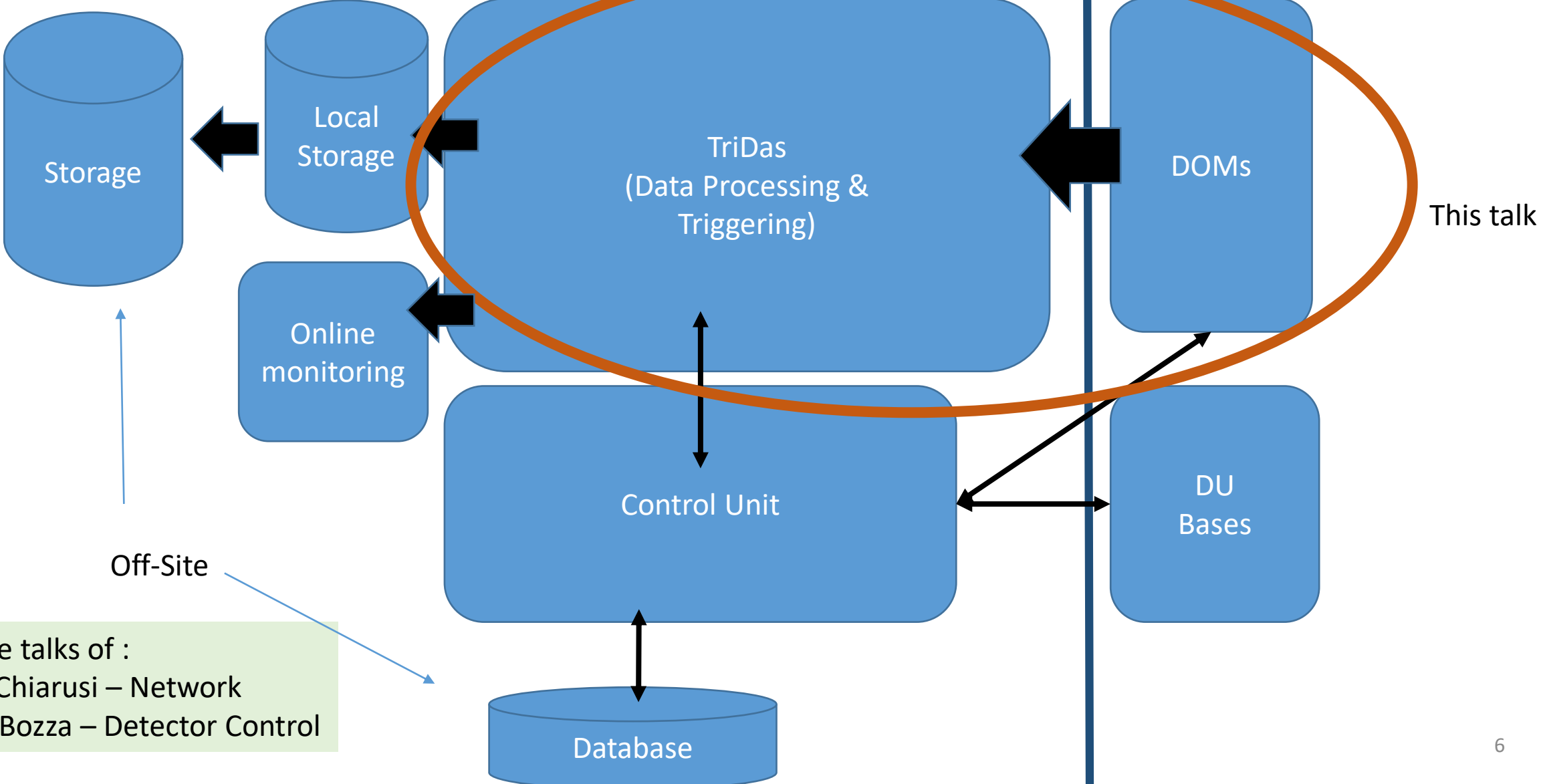
Require ns timing to exploit good water properties and suppress optical background

Suppress background (to “1st order”) :
use local coincidences

Find signal :
Use long range correlations

But you cannot do this off-shore
due to the detector scale!

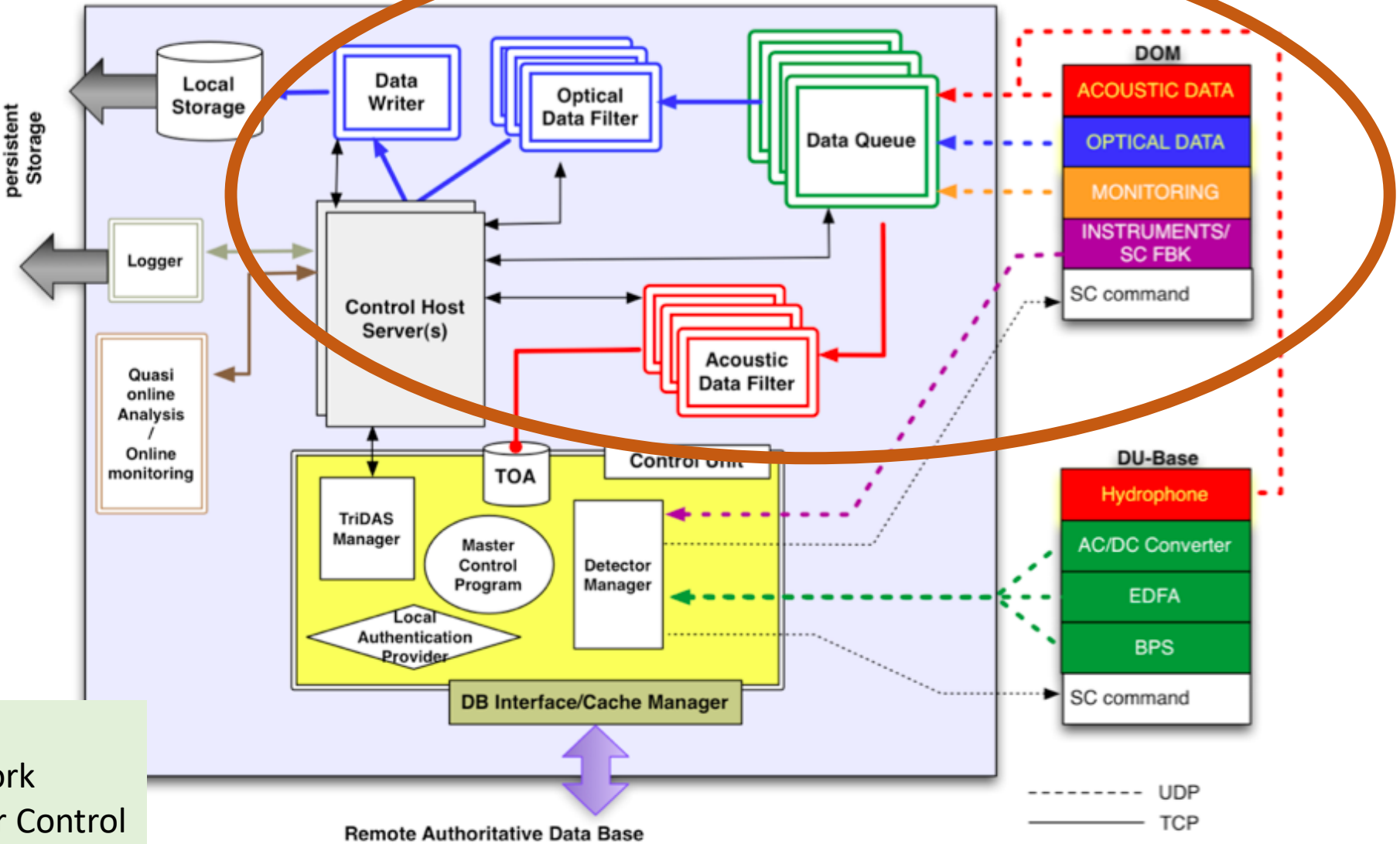
DAQ overview



Off-Site

See talks of :
T. Chiarusi – Network
C. Bozza – Detector Control

DAQ overview



See talks of :
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 C. Bozza – Detector Control

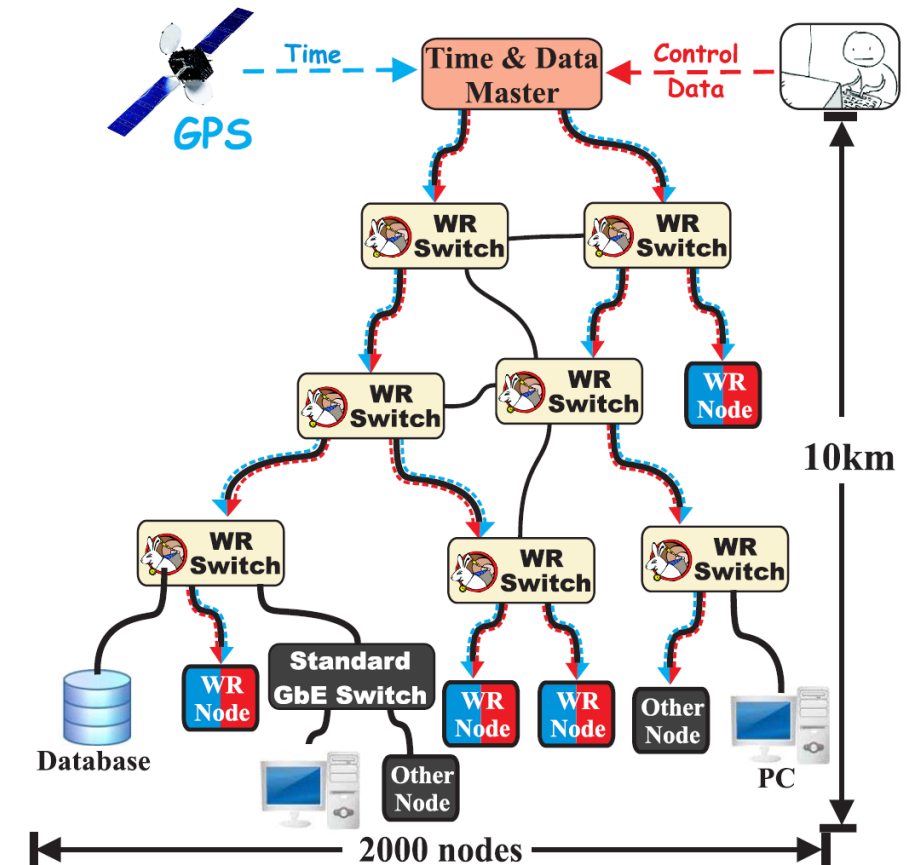
Data Streams

Type	Direction/Target	Size	Remark
Slow Control	Bidirectional, CLBs, Bases, Control Unit	minimal	To offshore : Broadcast, single channel To shore : individual channels
“Optical” Photon Hits	Unidirectional To Shore: DataQueue	11-23 Mbps/DOM	
“Acoustic” Hydrophone data	Unidirectional To Shore: DataQueue	4.6 Mbps/DOM	
“Monitoring” e.g. compass, countrates	Unidirectional To Shore: Database	minimal	
Timing	Bidirectional Bases, WR Master	minimal	White Rabbit, Part of Ethernet packets Broadcast to offshore

Time Synchronization: White-Rabbit

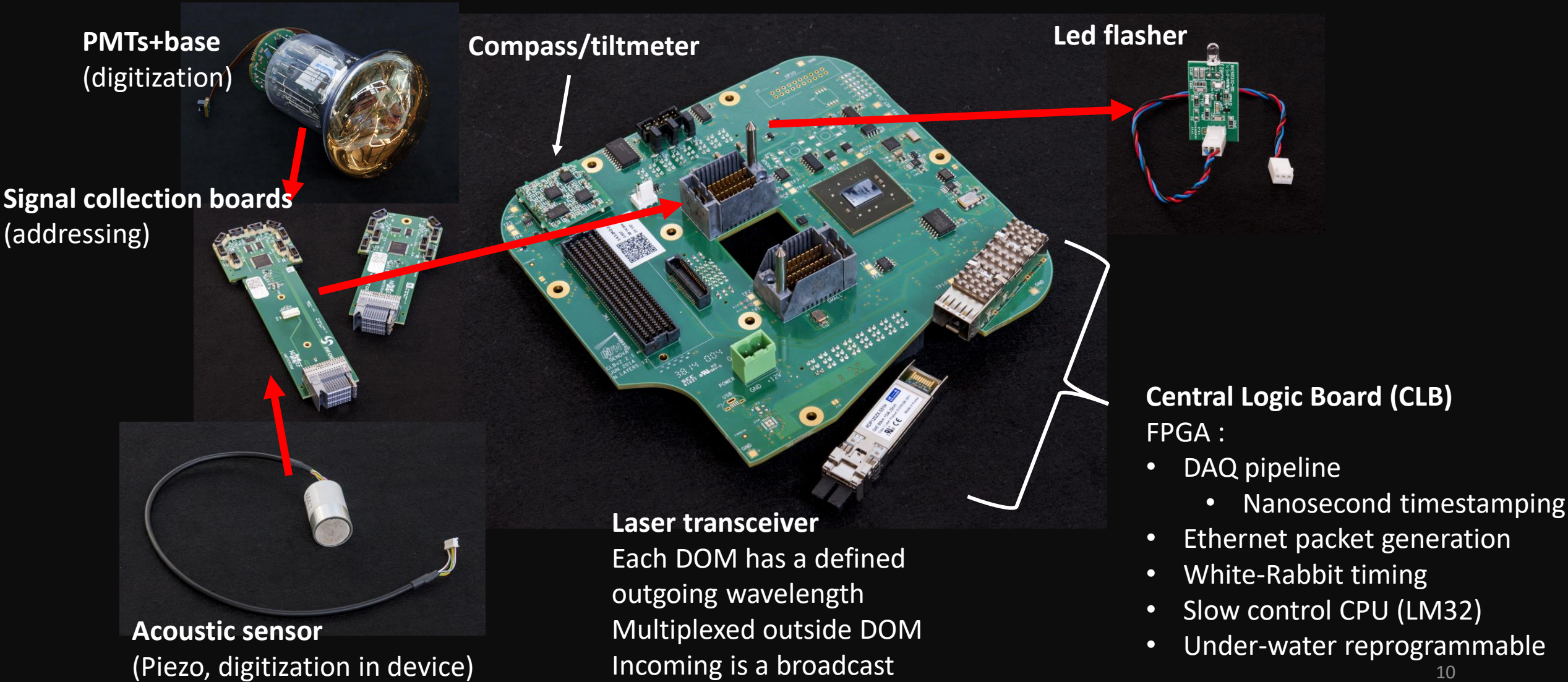
Enhanced Ethernet, providing:

- Synchronization: accuracy better than 1 ns precision (tens of ps stdev skew max)
- Deterministic, reliable and low-latency Control Data delivery



See talk of Pablo Marín JÍMENEZ

DAQ/Electronics in DOM



DOM Electronics: FPGA

IP/UDP (ethernet) packet creation

Multiple streams

(PMTs, acoustic, monitoring)

High-Speed TDCs

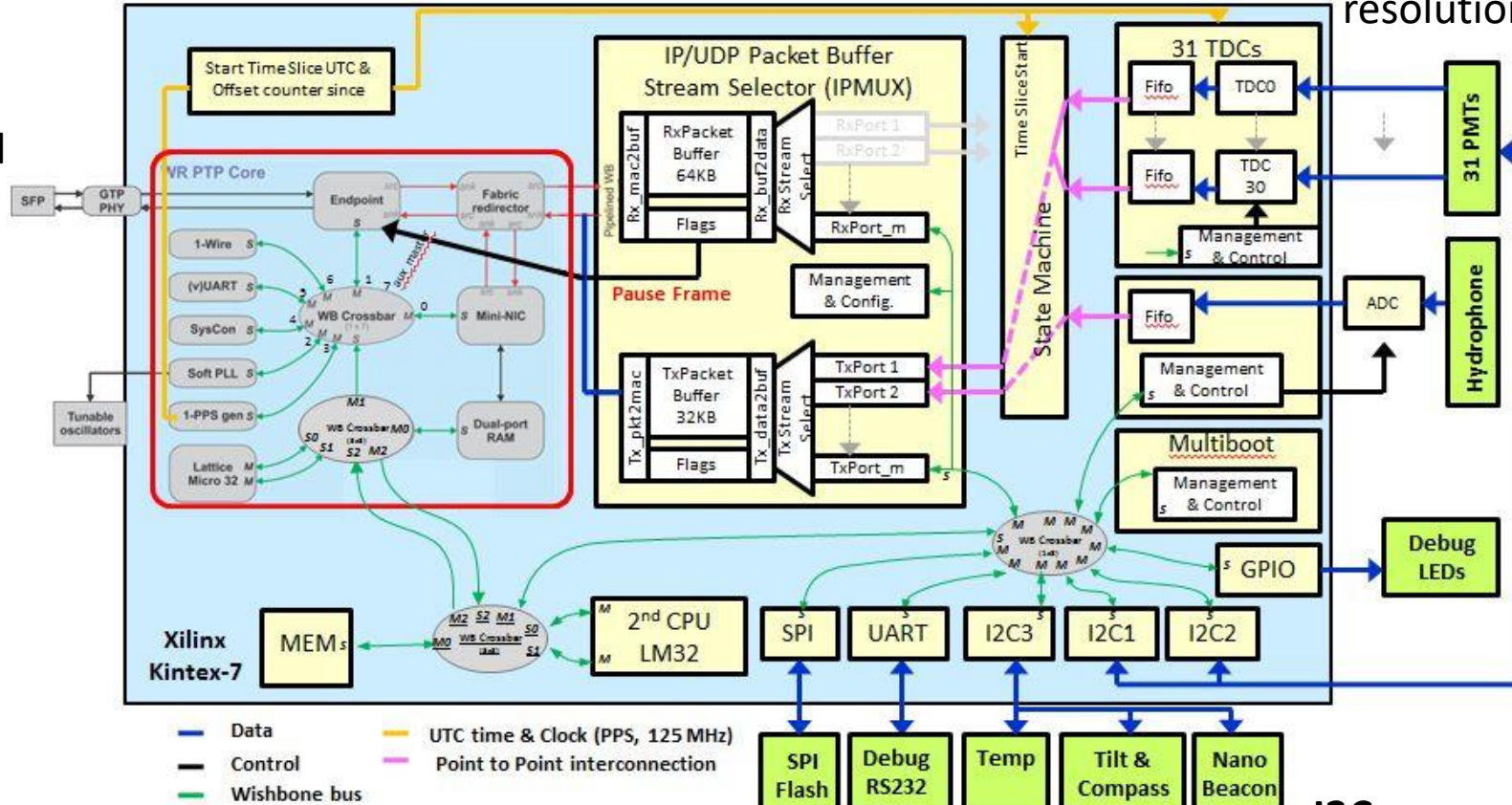
LVDS inputs from PMTs

Hit time and ToT. $800\text{ps}/\sqrt{12}$

resolution

White Rabbit PTP core

for time synchronization and Ethernet over optical fibres.



Multiboot system
(fail-safe reprogramming)

LM32 CPU

Control from shore

(Slow control, DAQ pipeline, White Rabbit, other sensors)

Implements software state machine

UART

Serial terminal

Tunneled over ethernet

I2C communication

HV and thresholds

Compass/tiltmeter

Led Flasher

Temperature/Humidity

Data format

Time-multiplexing:

Frame : all data from a DOM in 100 ms
 Timeslice: all frames for the same time
 Remember: DOMs are synchronized!!

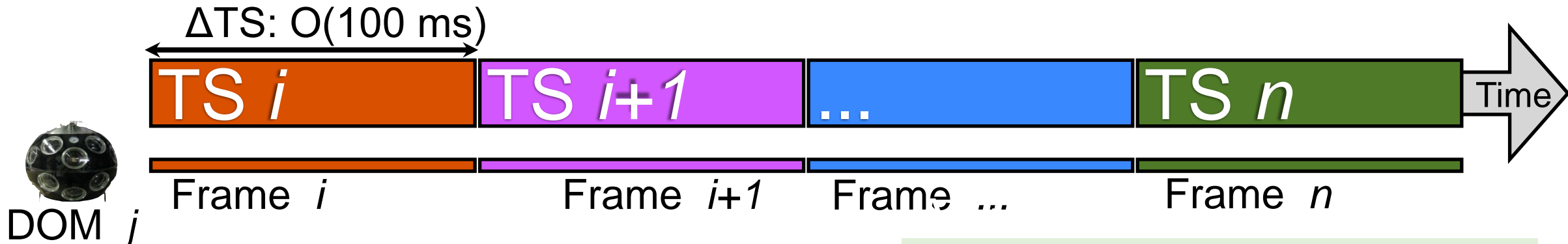
Time-multiplexing allows routing data with all benefits of Ethernet (addressing, load distribution, etc.)

Header of a frame :

		CLB Common Header																															
Offsets	Octet	0								1								2								3							
Octet	Bit	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31
0	0	Data Type																															
4	32	RunNumber																															
8	64	UDP Sequence Number																															
12	96	Timestamp																															
16	128																																
20	160	DOM Identifier																															
24	192	Status I																															
28	224	Status II																															
32	256	Status III																															
36	288	Status IV																															

UDP Packet # (for timeslice)
 Which DOM ?

Coarse time : seconds+16 ns ticks



Data is send over Ethernet via UDP packets

Timeslice spread out over multiple UDP packets and need reassembly (DataQueue)

Optical Data

Custom PMT bases

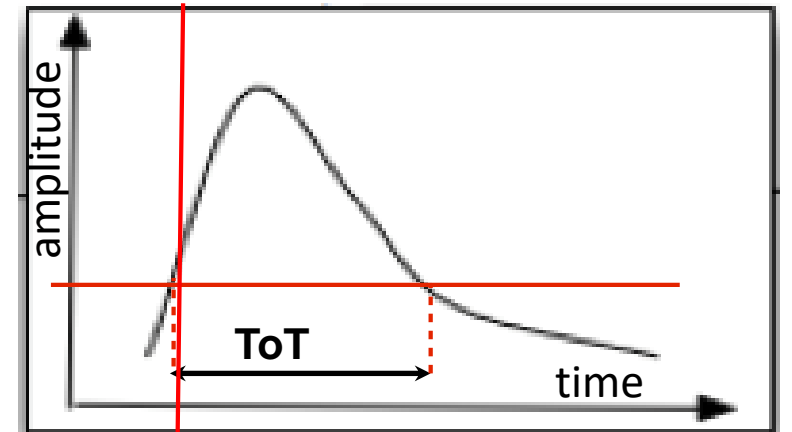
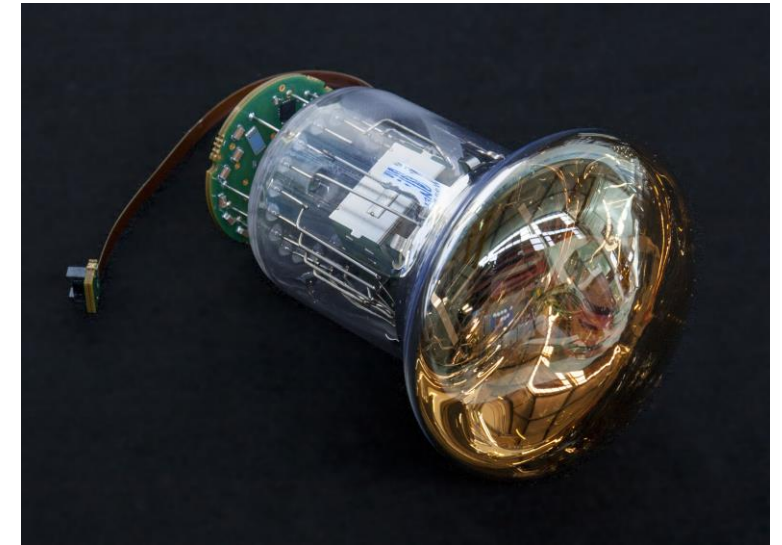
LVDS signal high as long as signal of threshold (adjustable)

FPGA registers

- For each PMT channel
- Time of threshold crossing - “fine time”
 - nanoseconds
 - Coarse time already in header
- Duration of time over threshold

CLB Optical Format Structure

Size (bit)	Description
448	DAQ Common Header
8	TDC channel
32	Time Stamp
8	Pulse Width
8	TDC channel
32	Time Stamp
8	Pulse Width
...	...
8	TDC channel
32	Time Stamp
8	Pulse Width



Note:

Low scattering (high information!) relaxes need for pulse shape!

Time of threshold crossing¹³

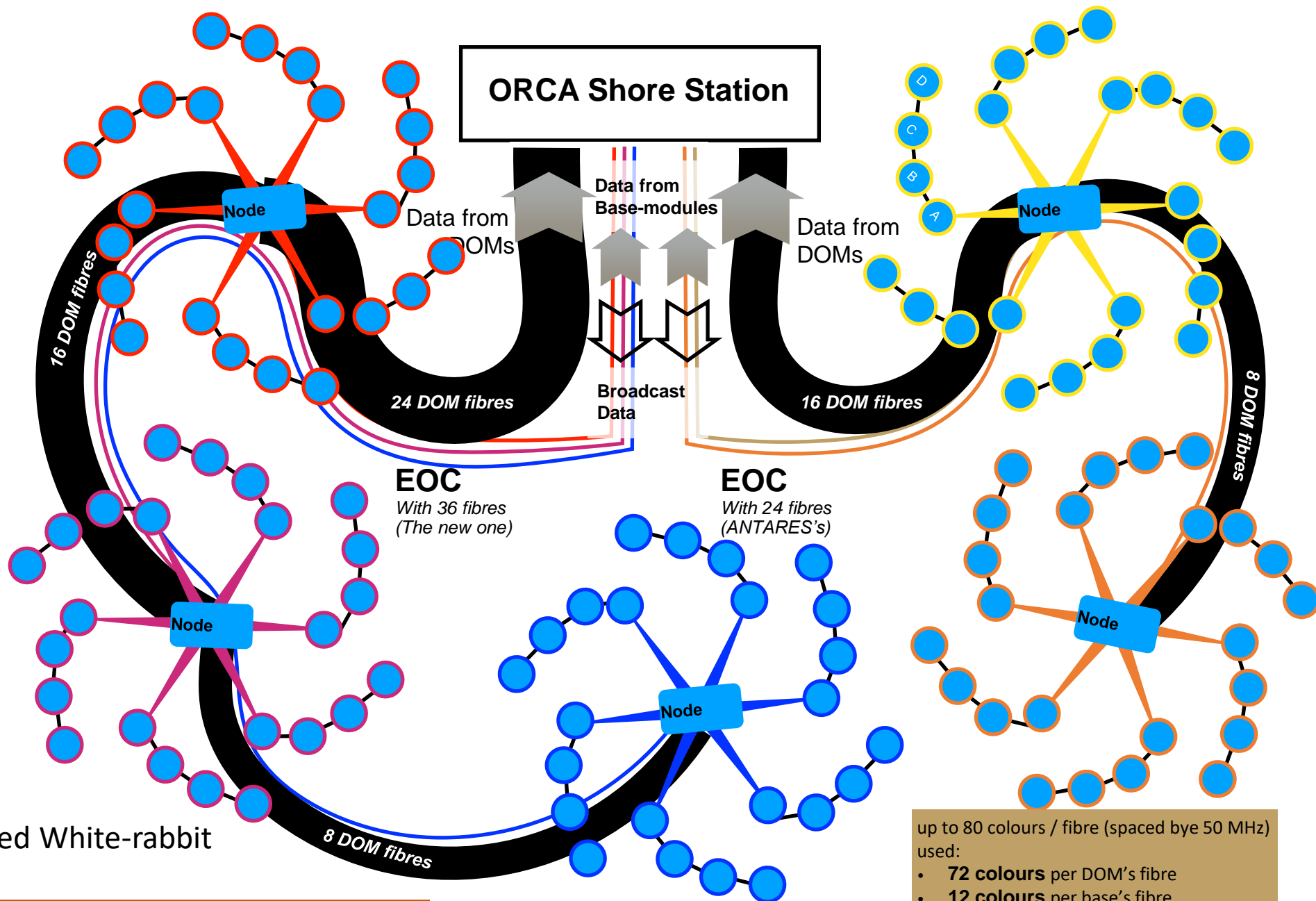
KM3Net :
Asymmetric network !

Broadcast :

Single channel to DOMs
(single wavelength)

Individual channels to shore
(wavelength multiplexing!)

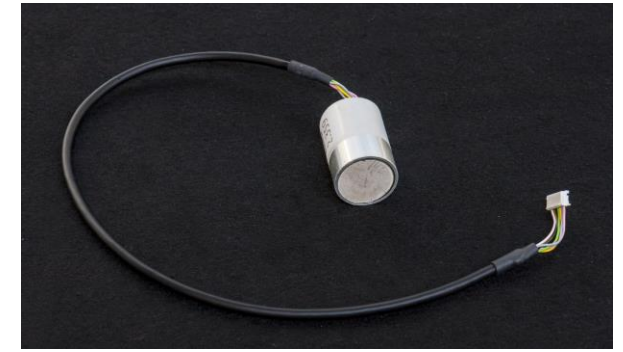
Datapath to shore does not need White-rabbit



up to 80 colours / fibre (spaced by 50 MHz) used:
• **72 colours** per DOM's fibre
• **12 colours** per base's fibre

See talk of T. Chiarusi on details of the onshore infrastructure

Some words on acoustic data



Continuous data stream, but sent in timeslices

Integration time > timeslice time

Processed per DOM

193.5 kHz sampling

12 bit

1 channel

Adjustable

20% of total data stream

(positioning)

Data Queue #1

Data Queue #2

⋮

Data Queue #n

Acoustic Data Filter #1

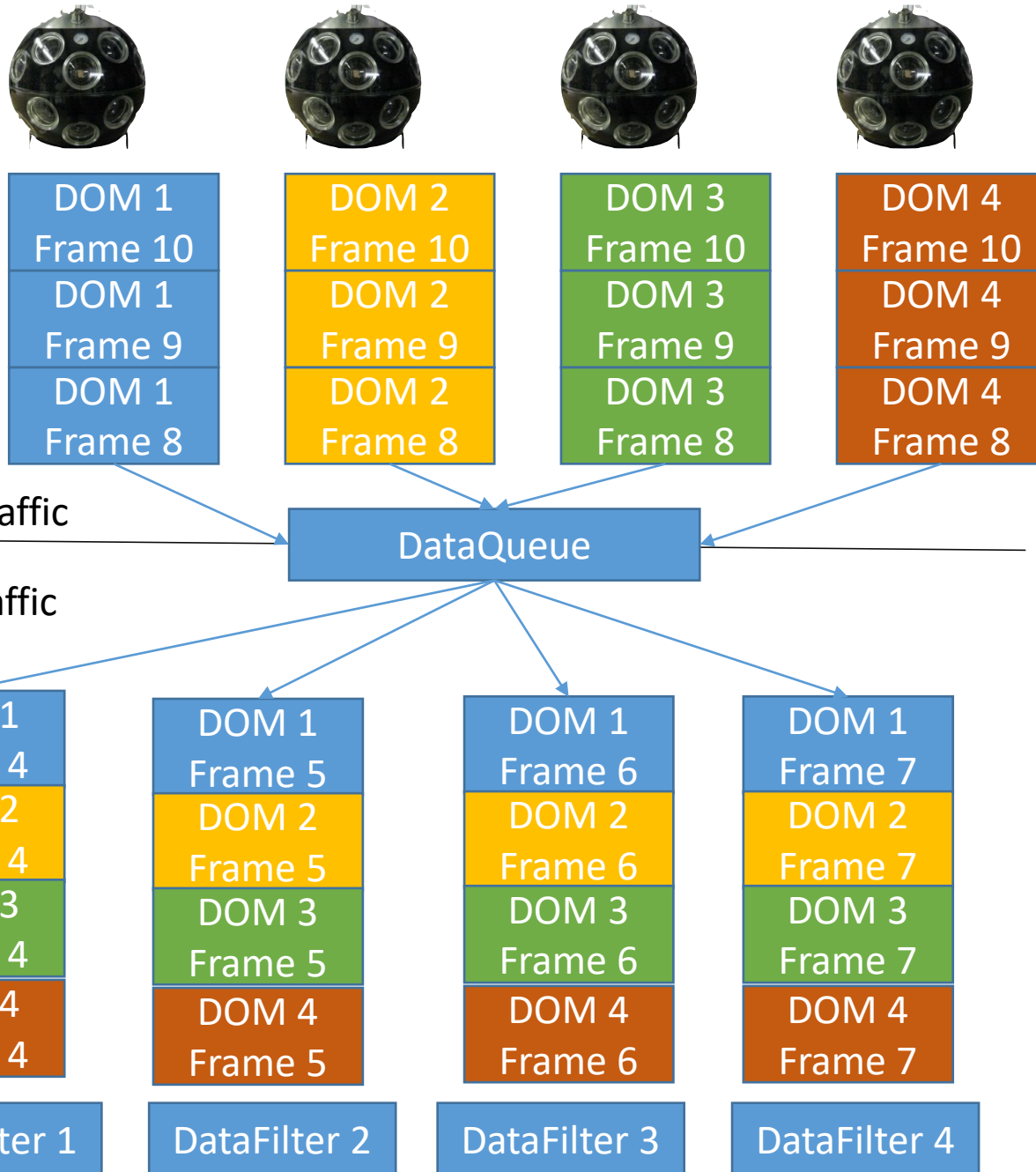
Acoustic Data Filter #2

⋮

Acoustic Data Filter #n

For acoustic neutrino detection another approach is required !!!

(Optical) Datafilter



DataQueue

Assembles frames from UDP packets
Distributes all frames of a timeslice to a datafilter in a round-robin way.

More processing power required ??

Add more datafilters!!!

Datafilter sees all data(hits) in a timeslice

Flexibility of all-data-to-shore design

Design allows for multiple triggers to run in parallel and on-the-fly adjustment
(In operation for Antares, most in development for KM3NeT)

Event types:

Track
Shower
(Exotic ? (Monopole trigger))

Source Tracking

1D trigger with direction towards (lowers energy threshold)

- Sun, Galactic Center, Earth for Dark Matter
- Other potential sources

Foreseen (available in Antares)

Multi-Messenger Astronomy: Generate and receive alerts

- Supernovae

“Data-dump”

Dump all raw data memory buffers to disk
Allows to *look back* in time, depending on buffer (memory) size
e.g. GRB alerts

Implementation at DataFilter level

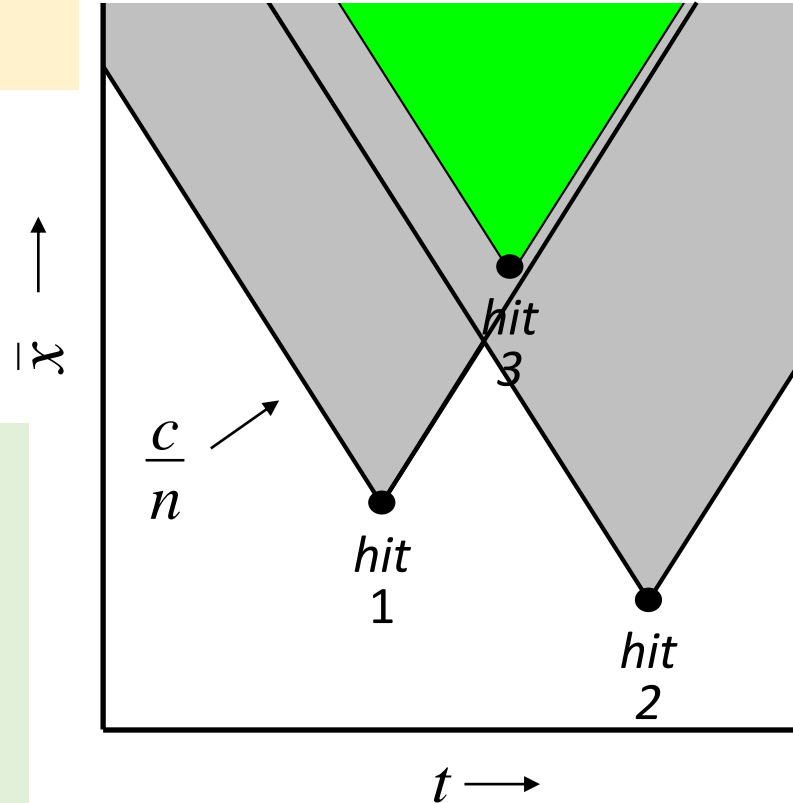
Trigger principles : causality & clustering

Basic condition : hits from a source must be causally connected

$$\frac{c}{n} |dt| \leq |d\bar{x}|$$

Clustering:

Find a cluster of at least N (e.g. 5) hits that are all pair-wise causally Connected ('Clique-problem')



hit
 \equiv
detected photon
at position \bar{x}
and time t

Non-transitive problem :

Hit (1,3) and (2,3) are causally related, but not hit (1,2)

Fine tuning example : Directional/1D causality

Assuming a direction, the causality criterion can be made more stringent

R can be constrained, as light attenuates!

R smaller than 3D distance

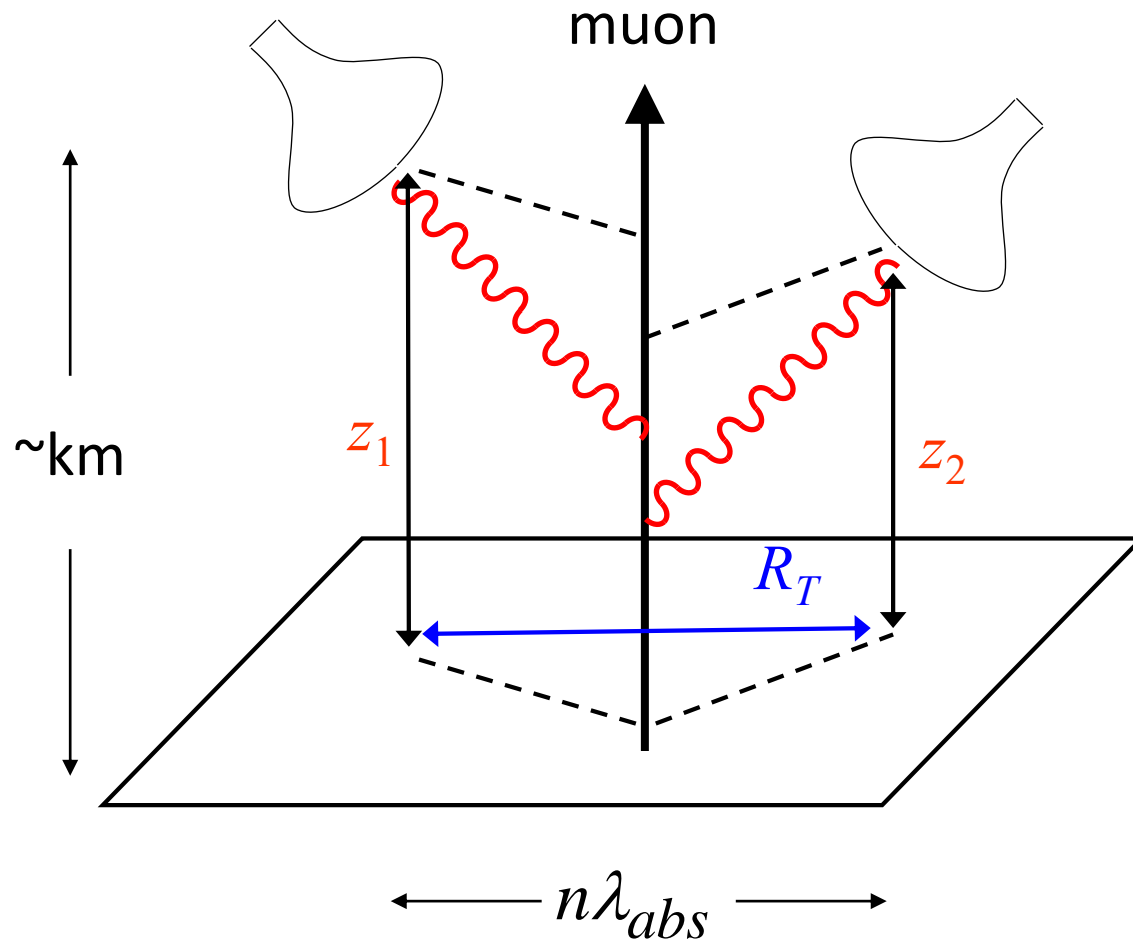
Stronger constraint on dT

Less random noise sources (PMTs)

$$\propto \frac{R}{D}$$

$$\propto \left(\frac{R}{D}\right)^2$$

Less CPU required, less chance of noise !



$$(z_1 - z_2) - R_T \tan \theta_c \leq c(t_1 - t_2) \leq (z_1 - z_2) + R_T \tan \theta_c$$

L0 vs L1

L0 hit : Any recorded hit

~7 kHz/**PMT**

L1 hit : 2 or more hits coincident (~20 ns) on DOM ~700 Hz/**DOM**

Sorting and cluster finding using L0 is computationally prohibitive

Using L1 hits increases energy threshold (e.g. 5 L1 hits mean ≥ 10 photons)

Solution (compromise?) :

Start with a L1 hit and use nearby L0 hits only

In ORCA, shower size < detector dimensions!

Optical Datafilter

- Assemble timeslices from frames received from DataQueues
 - Requires buffering and logic to deal with incoming stream
 - Calibrate
 - Translate DOM+PMT IDs to spatial coordinates
 - Level 1
 - Build L1
 - Find ≥ 2 coincidental hits in DOM
 - Level 2
 - Build L2
 - Restrict coincidences to nearby (angular) PMTs
 - Level 3
 - Muon trigger – 1D scan over N directions
 - Shower trigger – 3D causality (tuned)
 -
 -
 - Merge overlapping events
 - Send output to DataWriter
- Requires sorting, but only at DOM level
- Exploit Multi-PMT Design
- Multiple triggers in parallel on same data!!
-

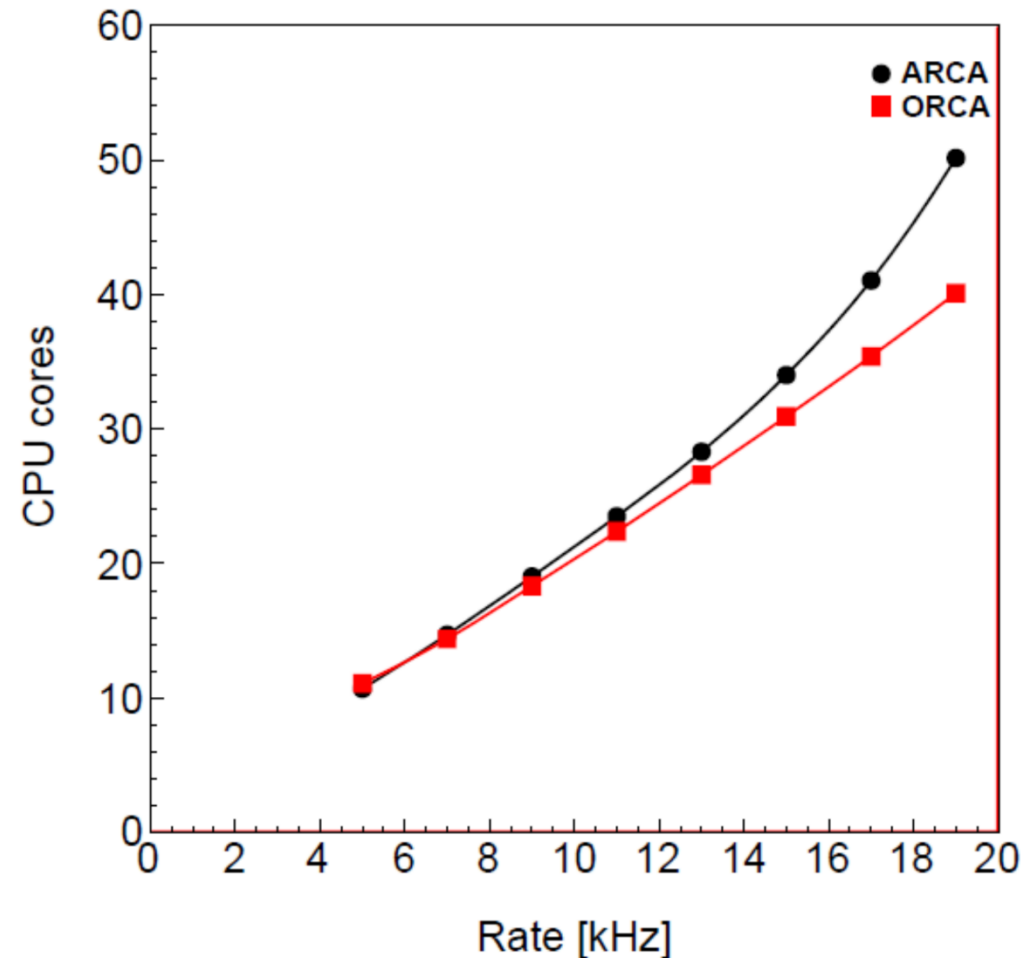
Required CPU power and performance

CPU time dominated
by sorting of hits
50 – 70 % depending on
algorithm, rates, etc.

Trigger settings optimized
by maximizing trigger efficiency
within bounds of storage/offline
processing

ORCA:
90% efficiency above 4 GeV

ARCA:
Efficiency depends on channel
(and definition)
Above 10 TeV all channels > 80 %
(≥ 5 hits)



ARCA
~100 Hz atm. muon events
~0 Hz pure noise

ORCA
~ 40 Hz atm. muon events
~ 20 Hz pure noise

Summary

- KM3NeT has designed* a Data Acquisition system that can deal with the large data volumes and unfavorable signal to noise ratio from the medium to maximize physics potential

*: and tested – ARCA and ORCA lines operated with this system

ARCA & ORCA

High Energy Neutrino Astronomy:

ARCA: Astroparticle Research with Cosmics in the Abbyss

Large Detector: $\sim 1 \text{ km}^3$ total

Sparsely instrumented: 36 m vertical spacing, 95 m horizontal

TeV-PeV Energies

Astrophysical Neutrinos

Same technology & layout, dimensions scaled \updownarrow

Neutrino Physics:

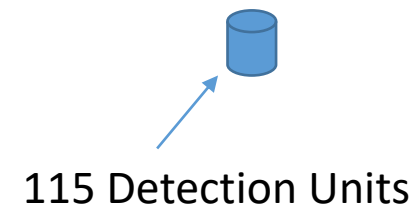
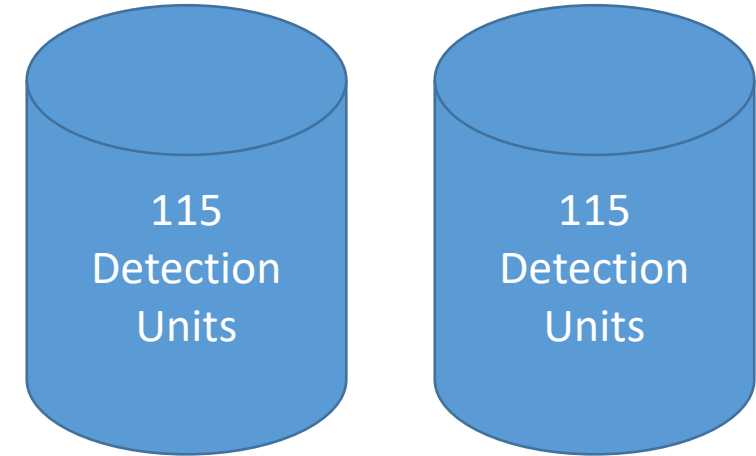
ORCA: Oscillations Research with Cosmics in the Abbyss

'Smaller' detector: 5.7 Mton

More densely instrumented: 9m vertical spacing, 20m horizontal

GeV energies

Atmospheric neutrinos



Multi-PMT Concept

Segmented photocathode : 31 3" PMTs in a 17" sphere
(equivalent to 3 10" PMTs)

+ All front-end and digitization electronics, slow control sensors
and supporting mechanics



KM3NeT
Digital Optical Module
(DOM)

Advantages

- Large photocathode area
- Directional Sensitivity
- Photon Counting
(1 vs 2 vs ... photons,
background suppression)
- Less overhead
- Cost effective
- Minimal glass penetrations 25

KM3NeT Design

Detection Units:

- 18 optical modules per vertical string
- ~36m or 9m between optical modules
- Lowest optical module ~100m or 40m above seabed
- Two Dyneema® ropes
- Backbone: 2 copper conductors; 18 fibres (+spares)
- Break out of cable at each optical module
- Base module with DWDM at anchor
- Cable for connection to seafloor network

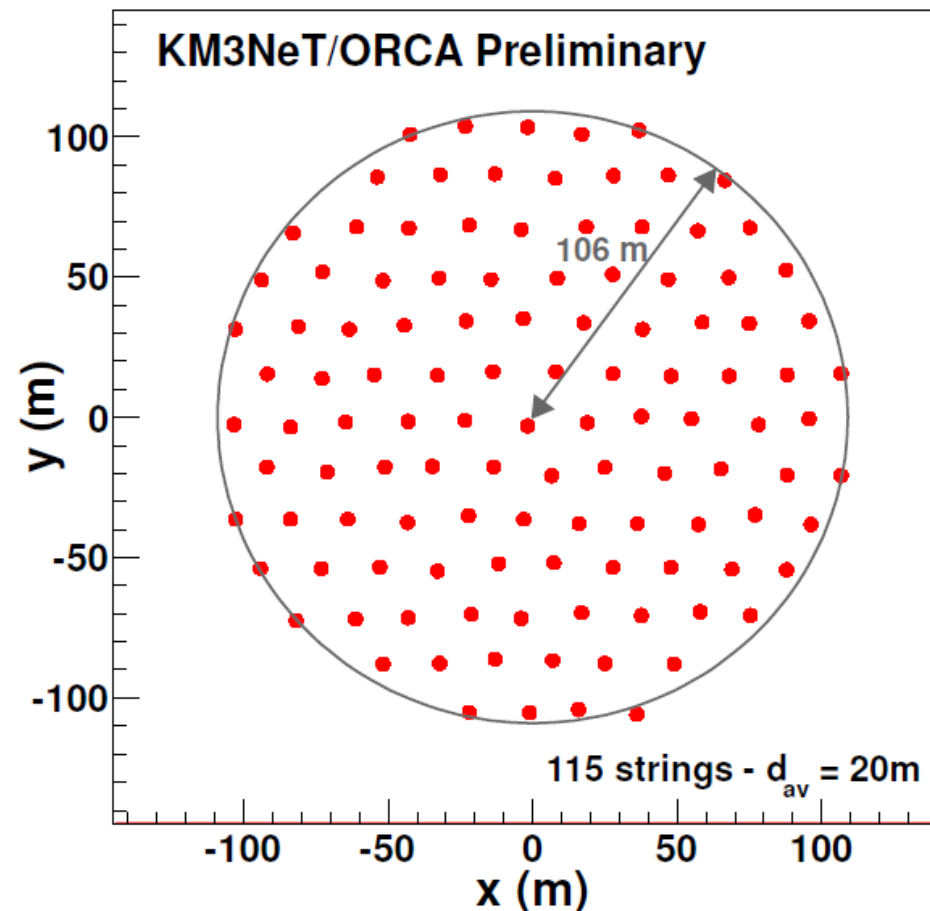
Cost saving design

Infrastructure:

- Building blocks of 115 strings
- Sea-bed infrastructure
(facility for long term high-bandwidth connection for sea-science, biology etc.)
- Optical data transmission

All-data-to-shore

Filtering/Trigger on-shore in computer farm



(‘ORCA’ layout)

153m or 612 m instrumented

DAQ State Machine

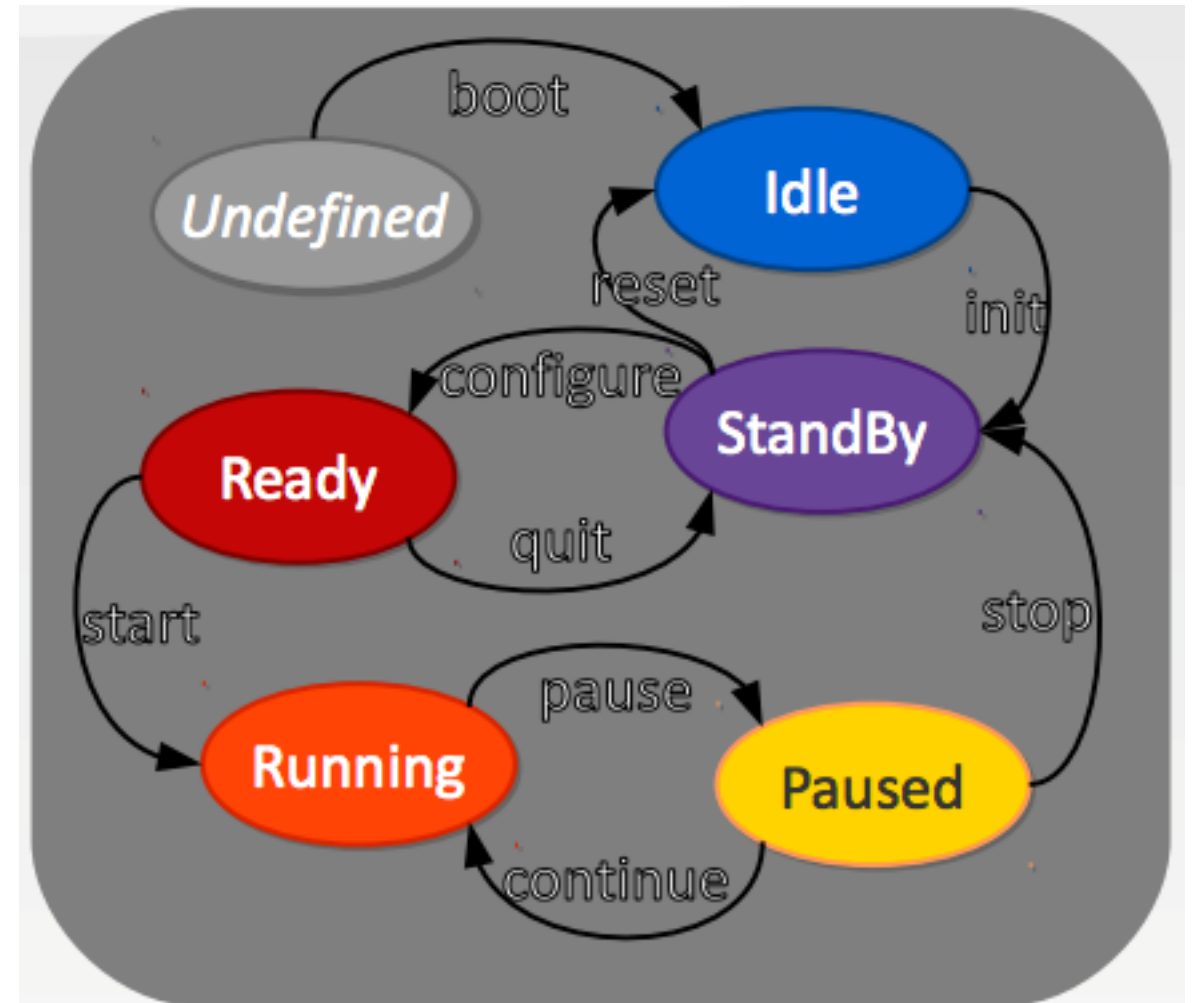
State machine implemented in all DAQ elements, off and onshore.
e.g. CLBs, DataFilters, ...

Transitions distributed and monitored from Control Unit

Actions and communications at transitions are defined

User interface simple :
ON,OFF, RUN

See talk of C. Bozza for details on detector control



Random Noise

$$R(m) \cong \binom{N}{m} \times f \times (f\Delta T)^{m-1}$$