

Large Volume Neutrino Telescopes

Cherenkov light from the charged products of neutrino interactions in sea-water are detected by a sparse array of photo-multiplier tubes

Two *general* event types:

Tracks - Charged current (CC) v_{μ} and $v_{ au}$

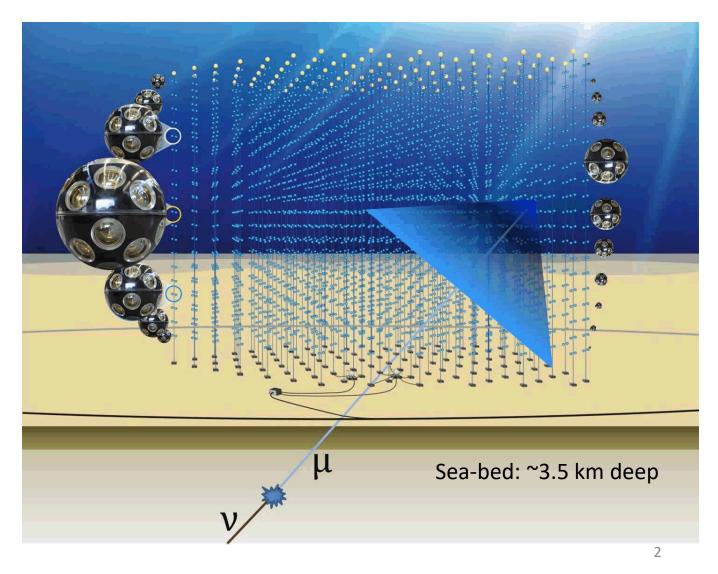
interactions

Showers - Neutral current v interaction

- v_e CC electromagnetic shower

- Vertex of CC interaction

- τ decay shower



ARCA & ORCA

High Energy Neutrino Astronomy:

ARCA: Astroparticle Research with Cosmics in the Abyss

Large Detector: ~1 km³ total

Sparsely instrumented: 36 m vertical spacing, 95 m horizontal

TeV-PeV Energies

Astrophysical Neutrinos

Same technology & layout, dimensions scaled

Neutrino Physics:

ORCA: Oscillations Research with Cosmics in the Abyss

'Smaller' detector: 5.7 Mton

More densely instrumented: 9m vertical spacing, 20m horizontal

GeV energies

Atmospheric neutrinos





KM3NeT Design

Detection Units:

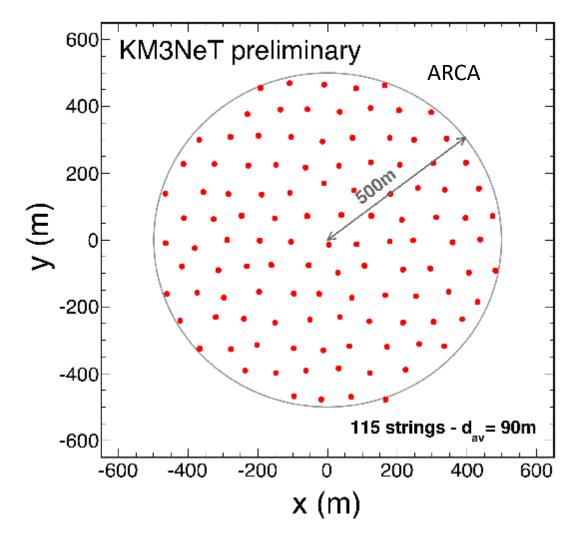
18 optical modules per vertical string
~36m or 9m between optical modules
Lowest optical module ~100m or 40m above seabe
Two Dyneema® ropes
Backbone: 2 copper conductors; 18 fibres (+spares
Break out of cable at each optical module
Base module with DWDM at anchor
Cable for connection to seafloor network
Cost saving design

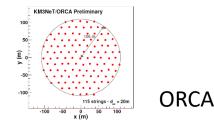
Infrastructure:

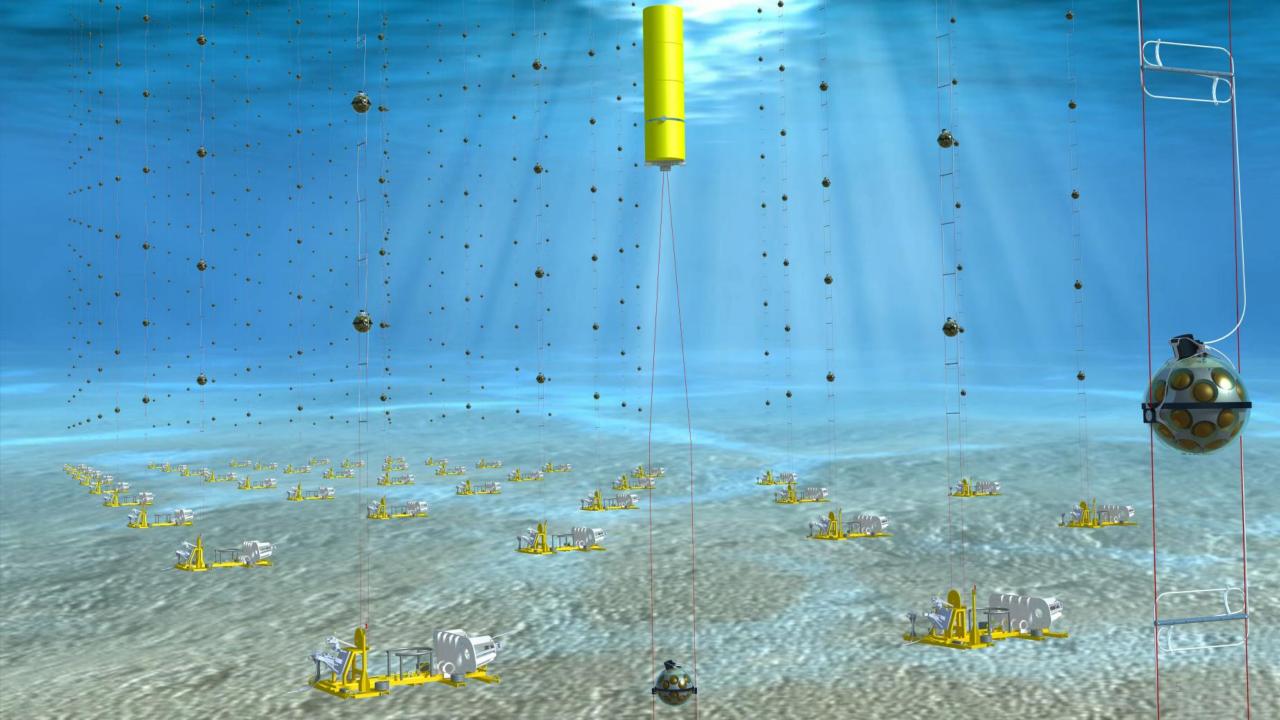
Detector building blocks of 115 detection units
Sea-bed infrastructure
(facility for long term high-bandwidth connection for sea-science, biology etc.)
Optical data transmission

All-data-to-shore

Filtering/Trigger on-shore in computer farm







Multi-PMT Concept

Segmented photocathode: 31 3" PMTs in a 17" sphere (equivalent to 3 10" PMTs)

+ All front-end and digitization electronics, slow control sensors and supporting mechanics



KM3NeT

Digital Optical Module

(DOM)



Advantages

- Large photocathode area
 - Directional Sensitivity
 - Photon Counting (1 vs 2 vs ... photons, background suppression)
 - Less overhead
 - Cost effective
- Minimal glass penetrations

PMTs

Main PMT Specifications:

➤ Timing ≤2ns (RMS)

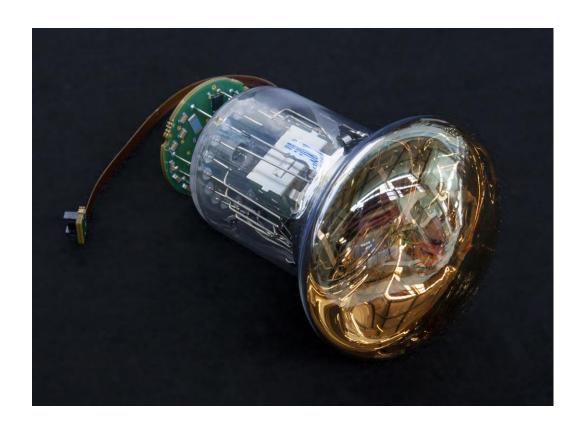
➤ QE@ 404 nm ≥23%

➤ QE@470 nm ≥18%

➤ Collection efficiency ≥90%

➤ Dark Count (0.3 p.e.) < 2 kHz

➤ Price/cm2 ≤10" PMT



Suppliers:

- Hamamatsu (R12199)
 - (currently used in KM3NeT phase-1 detector)
- ETEL (D792)
- HZC (XP53B20, development ongoing)
- Melts

PMT Bases

PMT base - KM3NeT design

- HV generation on the base
 - Cockroft-Walton circuit
 - Input 3.3 V
 - Output to -1400 V
 - Controlled by custom ASIC : Coco
- <u>Time-over-threshold</u> readout (ToT)
 - Custom ASIC: PROMiS
 - Pre-amplifier
 - Digitization on the base
 - LVDS signal output
- LOW power (140 mW for 31 PMTs)
- HV and threshold adjustable over I2C
- Each base has a unique electronic identifier
- 3.3 V, I2C, LVDS over thin kapton cable
- Adjustable for different PMT manufacturers



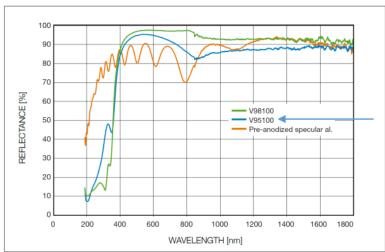
Negative HV on photo-cathode

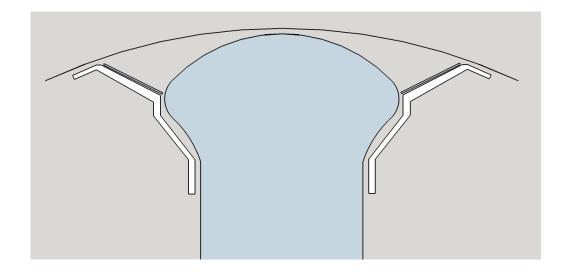
Gain: 3 *10⁶

HV tuned to set ToT to a specific value at fixed threshold

Reflector rings



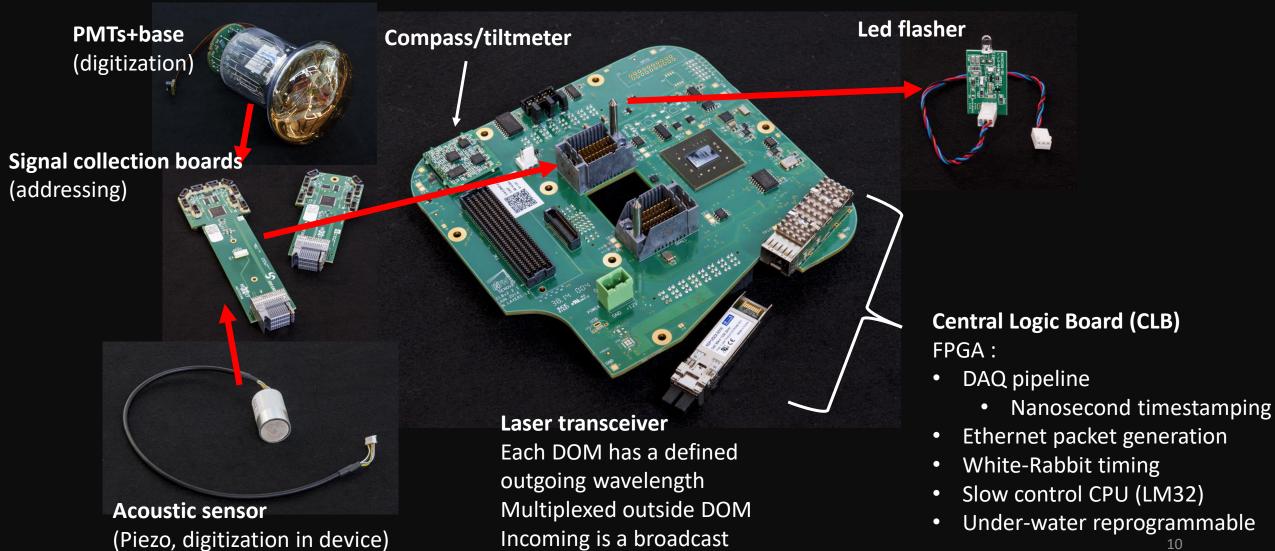




Reflector rings around PMTs increase light yield with 20-40 % and improve directionality

Aluminium coated with silver and protective layers

DAQ/Electronics



IP/UDP (ethernet) packet creation

Multiple streams (PMTs, acoustic, monitoring)

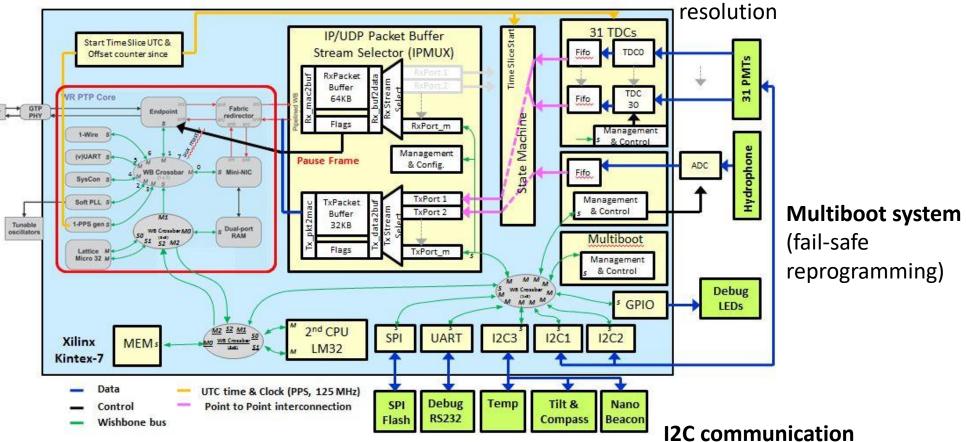
High-Speed TDCs

LVDS inputs from PMTs Hit time and ToT. $800 \text{ps}/\sqrt{12}$

DAQ/Electronics: FPGA

White Rabbit PTP core

for time synchronization and Ethernet over optical fibres.



LM32 CPU

Control from shore (Slow control, DAQ pipeline, White Rabbit, other sensors) Implements software state machine

UART

Serial terminal Tunneled over ethernet

I2C communication

HV and thresholds Compass/tiltmeter Led Flasher Temperature/Humidity

DAQ – Datastream from DOM

Digitized LVDS pulses are converted to t0 (leading edge) and ToT (width of pulse) by TDCs



Continuous datastream from PMTs is converted into 'hits': t0, ToT and PMT ID – 6 bytes



All hits for a specific duration (100 ms) are collected in 'frames'



Frames are formatted into IP/UDP packets and sent over 1Gb optical link

On- shore switching infrastructure and farm collects frames for all DOMs and assembles timeslices (100 ms snapshot)

Trigger farm looks for correlated hits.

Interesting timeslices are stored



Selected Mechanics

Cooling structure (mechanical support and passive cooling)

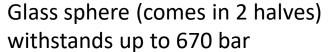




3D printed support structure (SLS)

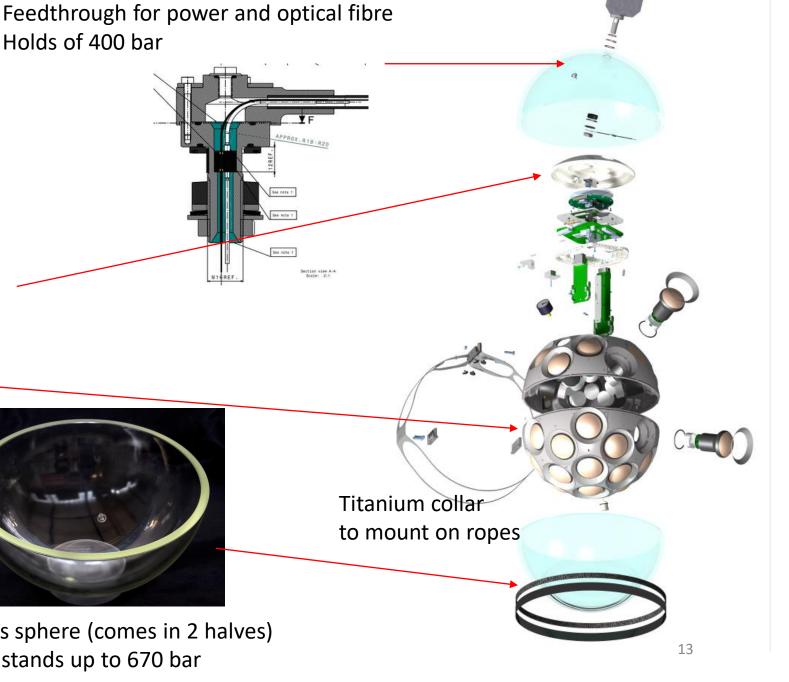
- Defines PMT, piezo, led, ... positions
- Barrier for optical gel





'Penetrator' (KM3NeT design)

Holds of 400 bar



DOM integration

With 1.5 FTE:
1 DOM takes 3 days
... but 5 take a week





Integration, functional test, integration, acceptance test.

All components have their own identification (QR code) with associated database entry (e.g. PMT calibration)

QA/QC system tracks components Integrated in DOMs.



DU Overview

Dyneema ropes provide mechanical support

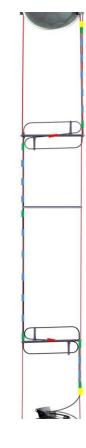
~37 m

DOMs are buoyant

VEOC

(vertical electrical optical cable)

- 7 mm PE tube
- 24 fibres
- 2 copper
- BOB at every DOM
- "pressure balanced"
- Closed system
- guided along ropes



Break-out-box (BOB):

- Oil-filled
- Two sections
 - VEOC side
 - DOM side
- One fibre split off
- Copper
- DC/DC converter (380 V to 12 V)
- At ambient pressure!!



~9 m





DU Base and Container

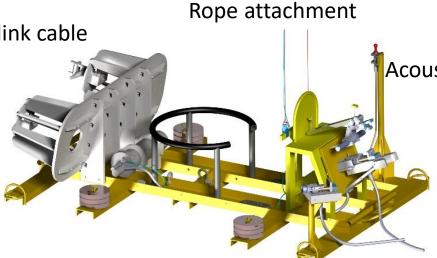
Base container

- Titanium cylinder
- Power distribution and protection (slow control)
- Multiplexing of DOM signals onto one fibre
- CLB (as in DOMs) with white-rabbit (ns timing)

Contents of base container



Storage of interlink cable



Acoustic receiver and transmitter

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Wet-mateable connector with guiding system

Optical schematics of base-container

Weight and sacrificial anodes

Deployment mechanics

DU Integration

Steps of DU integration

Attaching DOMs to VEOC

Splicing of optical fibres and connection

of copper

Closing of BOBs on DOM side

Leak tests (pressurized air)

Oil Filling

Attaching 'base penetrator' to bottom BOB

Leak tests (pressurized air and oil)

Integration of base-container

Installation of electronics and optical components

Install feedthroughs/connectors in flange

Base container/VEOC connection

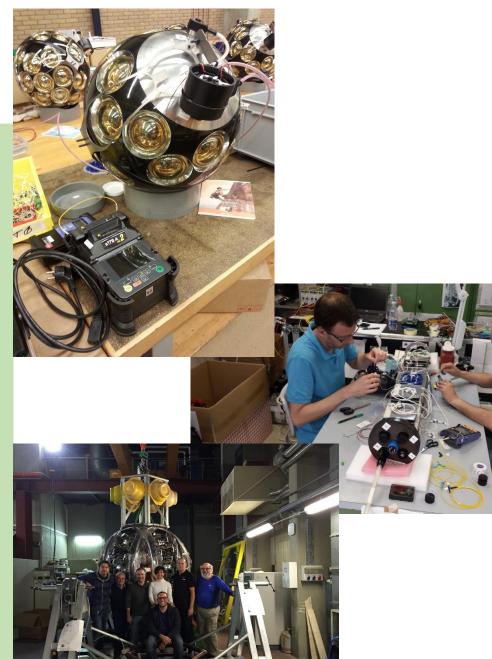
LOM (deployment mechanics) loading

Attaching ropes and guidance mechanics

Insert DOM into LOM

Attaching buoy

Preparation of base



DU Deployment





DU in LOM on sea-bottom

A ROV will pull the lever to release

DU installed on Launcher



Deployments

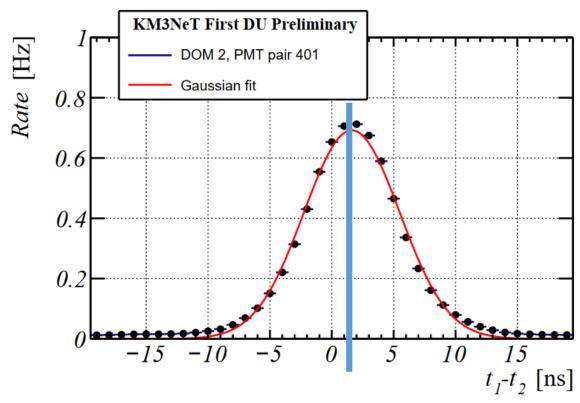
- 4 DUs deployed
 - 3 ARCA type
 - 2 in one operation
 - 2015, 2016
 - 1 ORCA type
 - 2017
 - Not counting test deployments
- 2 DUs recovered
 - 1 ARCA to investigate failure
 - 1 ORCA to facilitate undersea cable deployment

Up to 5 ORCA deployments foreseen after cable repair this fall (2018).

Operations foreseen at ARCA site to connect 2 DUs in Q4 2018

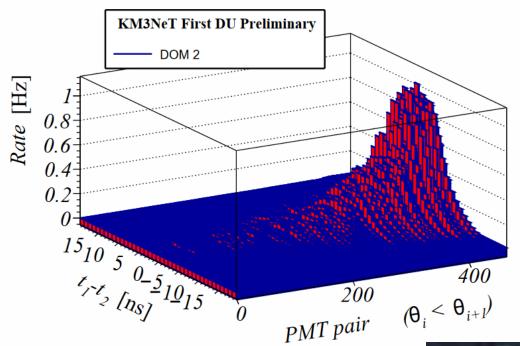
K40 time calibration

Time difference distribution between two PMTs in a DOM



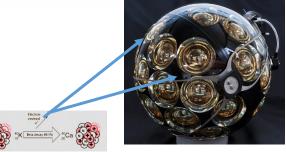
Time offset between the two PMTs

All PMT pairs

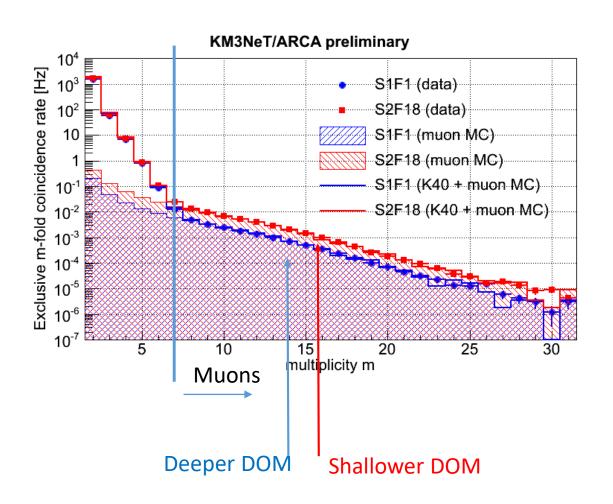


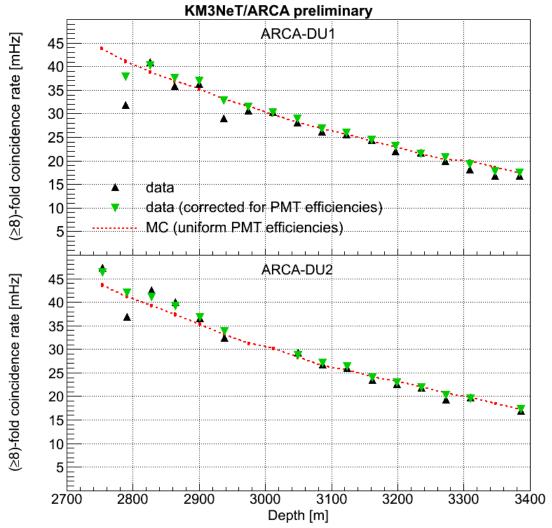
Information from k40 decay:

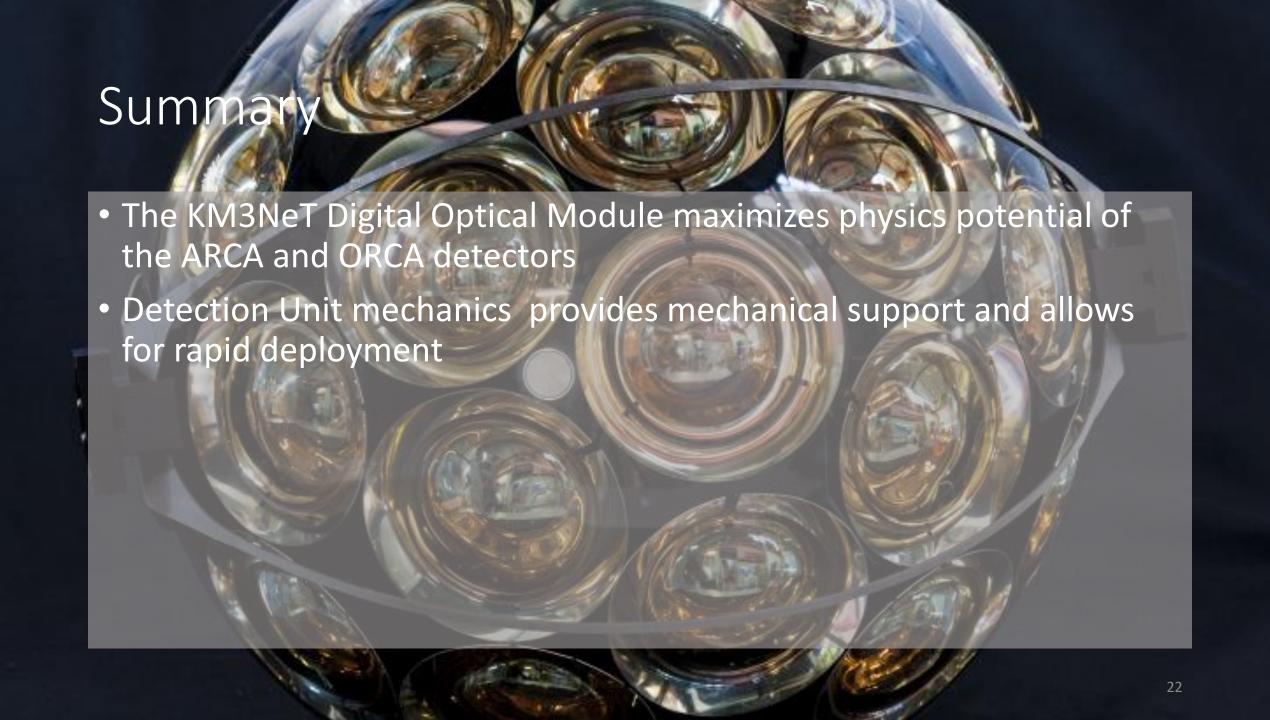
- time offset
- efficiency
- time spread



Muon depth dependence





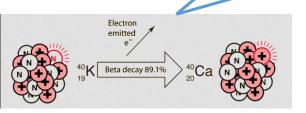


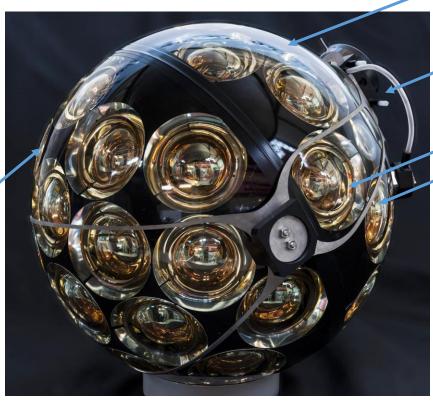
BACKUP

In-situ time calibration

Cherenkov photons

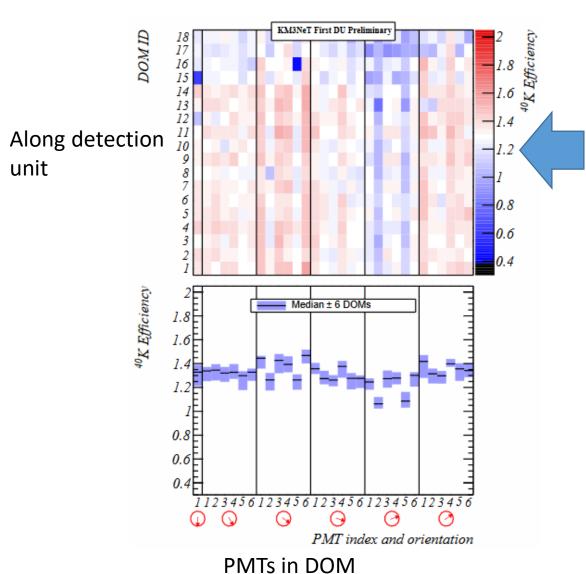
Photons from k40 decay cause coincident hits





Muon

Understanding efficiencies



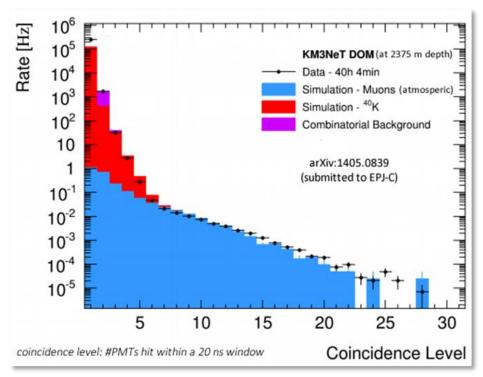
Coincidences give insight into relative efficiencies

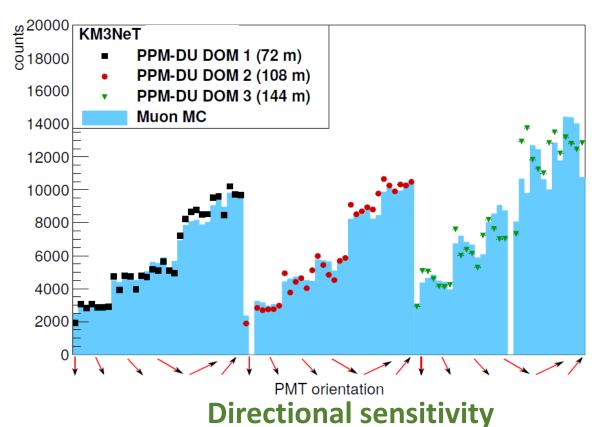
Vertical bands indicate influence of DOM mechanics



Photons are blocked

Photon counting and direction





Photon counting

(muons cause higher multiplicity coincidences)

(photons from muons come from above)

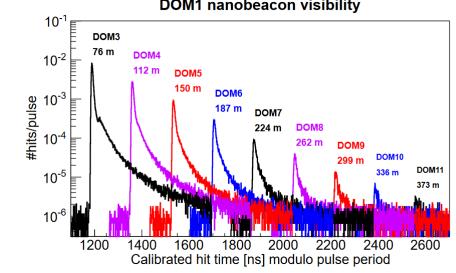
(data in these plots is from prototypes PPM-DOM (Eur. Phys. J. C (2014) 74:3056) and PPM-DU (Eur. Phys. J. C76 (2016) no. 2, 54)

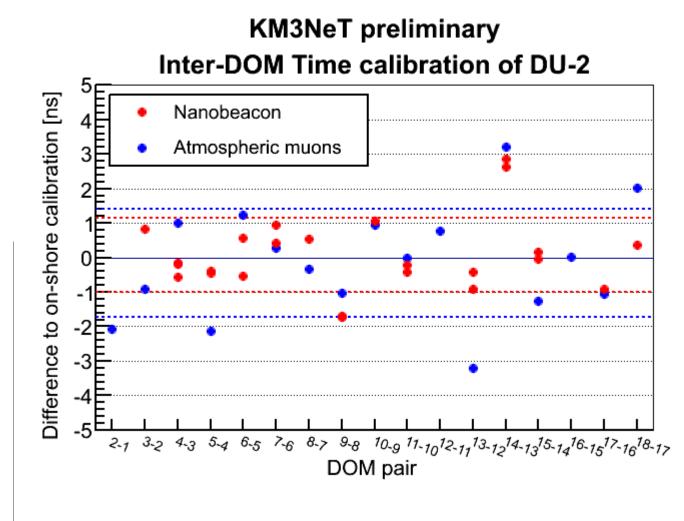
Inter-DOM calibration

Calibration between DOMs:

- Laser calibration in lab
- Led-flashers
- Atmospheric muons

KM3NeT First DU Preliminary DOM1 nanobeacon visibility





(Data from ARCA –DU1 & 2)

Reconstructed event

