

Report
of the Working Subgroup
for the Strategic Long-range Plan of JINR

A. Ioffe, V. Aksenov
(co-chairmen of the WSG)

Working subgroup (since June 2018)

Co-chairmen:

- V.L. Aksenov (Russia)** NRC “Kurchatov Institute”, Moscow;
Frank Laboratory of Neutron Physics JINR, Dubna
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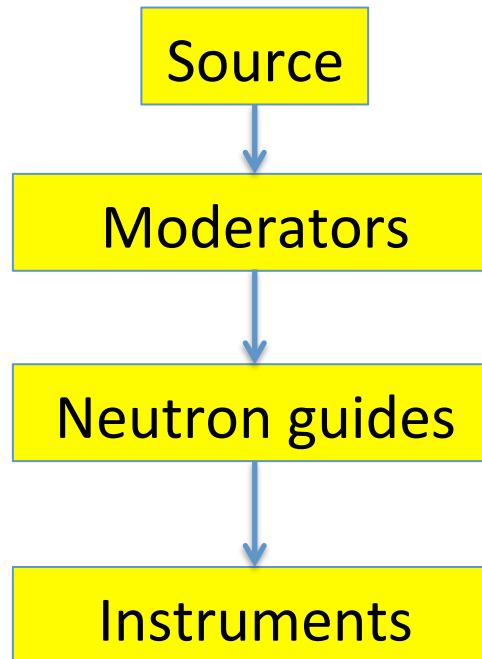
Members:

- J. Carpenter (USA)** – Argonne National Laboratory, Chicago
- N. Kucerka (Slovakia)** – Frank Laboratory of Neutron Physics JINR, Dubna
- T.V. Kulevoy (Russia)** – National Research Center “Kurchatov Institute” – ITEP, Moscow
- A.V. Lopatkin (Russia)** – N.A. Dollezhal Research and Development Institute of Power Engineering (NIKIET), Moscow
- F. Mezei (Sweden)** – European Spallation Source, Lund
- P. Mikula (Czech. Rep.)** – Nuclear Physics Institute – Řež, Member of the JINR PAC on Condensed Matter Physics
- D. Nagy (Hungary)** – Wigner Research Centre for Physics, Budapest, Chairman of JINR PAC on Cond. Matter Physics
- V. Nesvizhevsky (France)** – Institute Laue-Langevin, Grenoble, Member of JINR PAC on Nuclear Physics
- L. Rosta (Hungary)** – Budapest Neutron Center, Budapest
- S.F. Sidorkin (Russia)** – Institute for Nuclear Research of Russian Academy of Sciences, Troitsk Moscow Reg.
- I.T. Tretyakov (Russia)** – N.A. Dollezhal Research and Development Institute of Power Engineering (NIKIET), Moscow

Activities:

- Meeting on June 15, 2018
- Int. Workshop “Advanced ideas and experiments for DNS-IV”
(December 6 – 7, 2018, Dubna)
- Meeting on January 23, 2019

Directions of work:



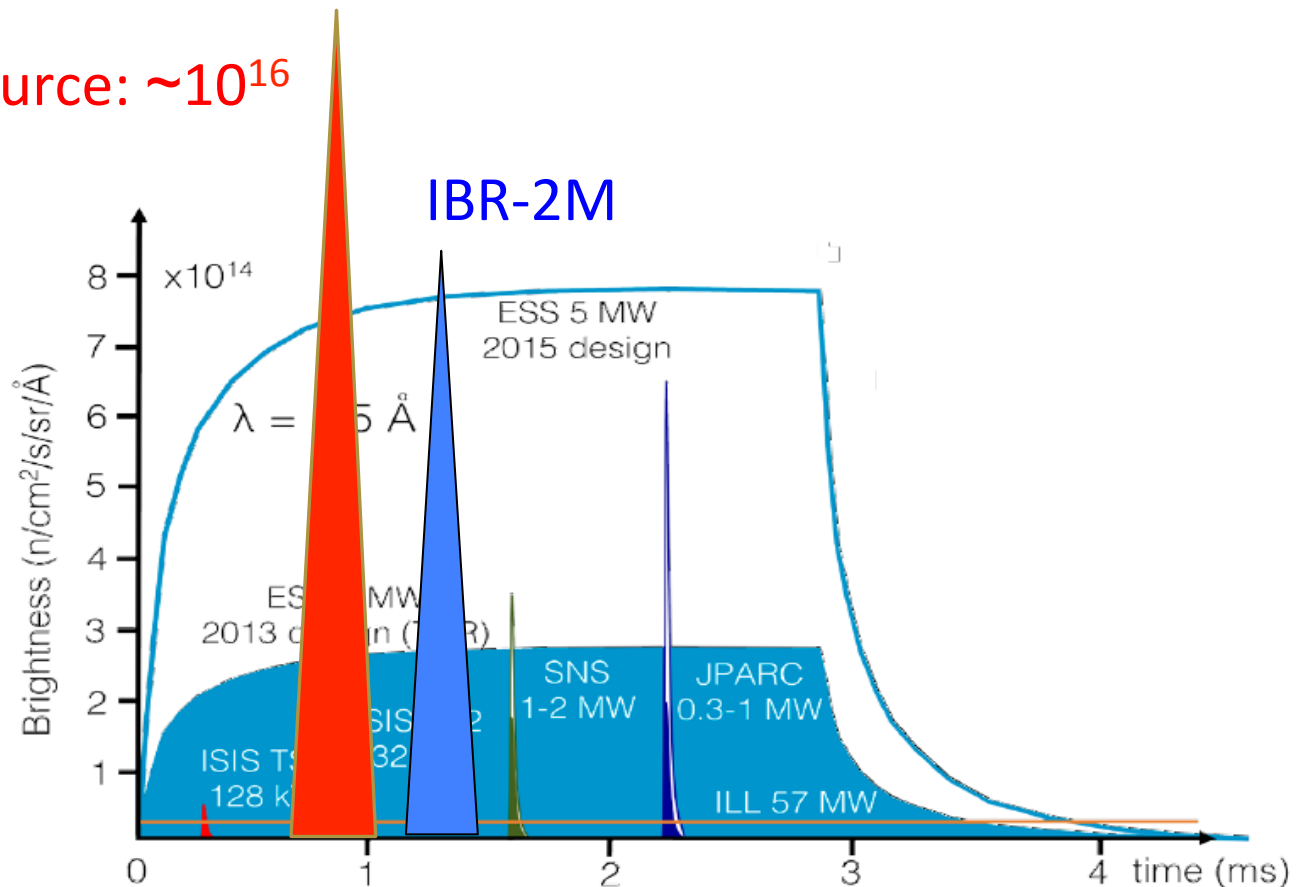
Requirements to the source:

Assumption: maximal allowable thermal power (10-15) MW.

⇒ Mean neutron flux 10^{14} n/s/cm²

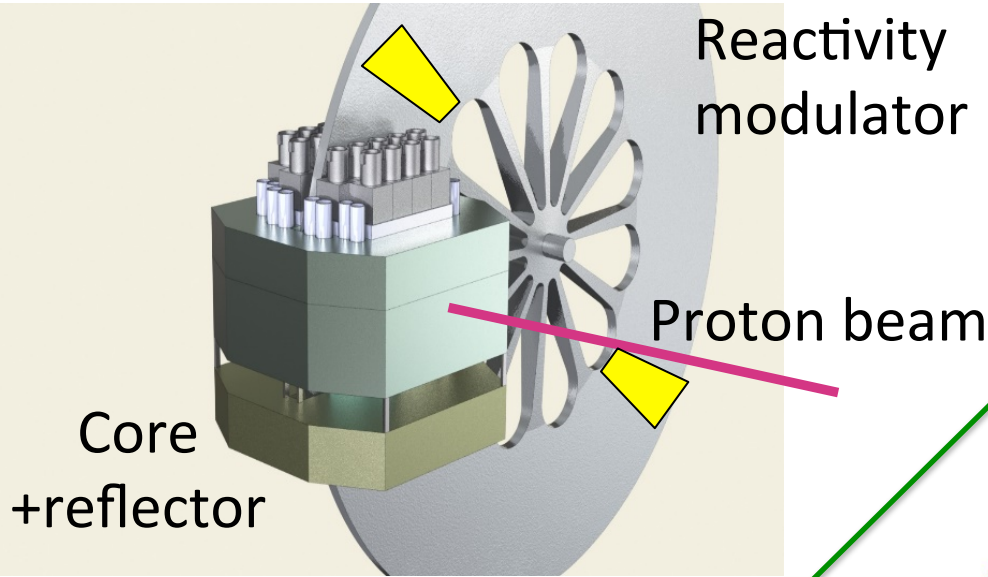
⇒ Corresponds to mean flux of the world-leading pulsed sources

New source: $\sim 10^{16}$



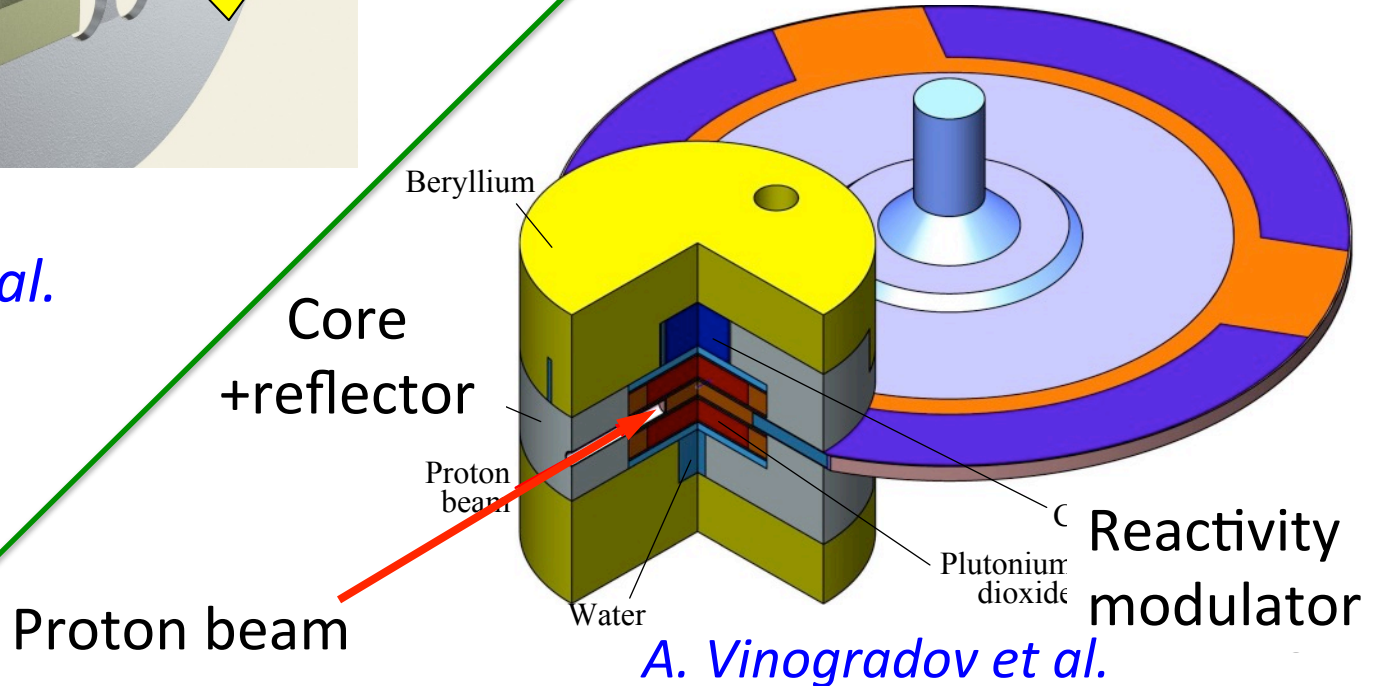
Pulsed neutron source with neutron multiplier: two alternative approaches

Superbooster: pulsed reactor + spallation



E. Shabalin et al.

Accelerator-driven neutron source: spallation + multiplier

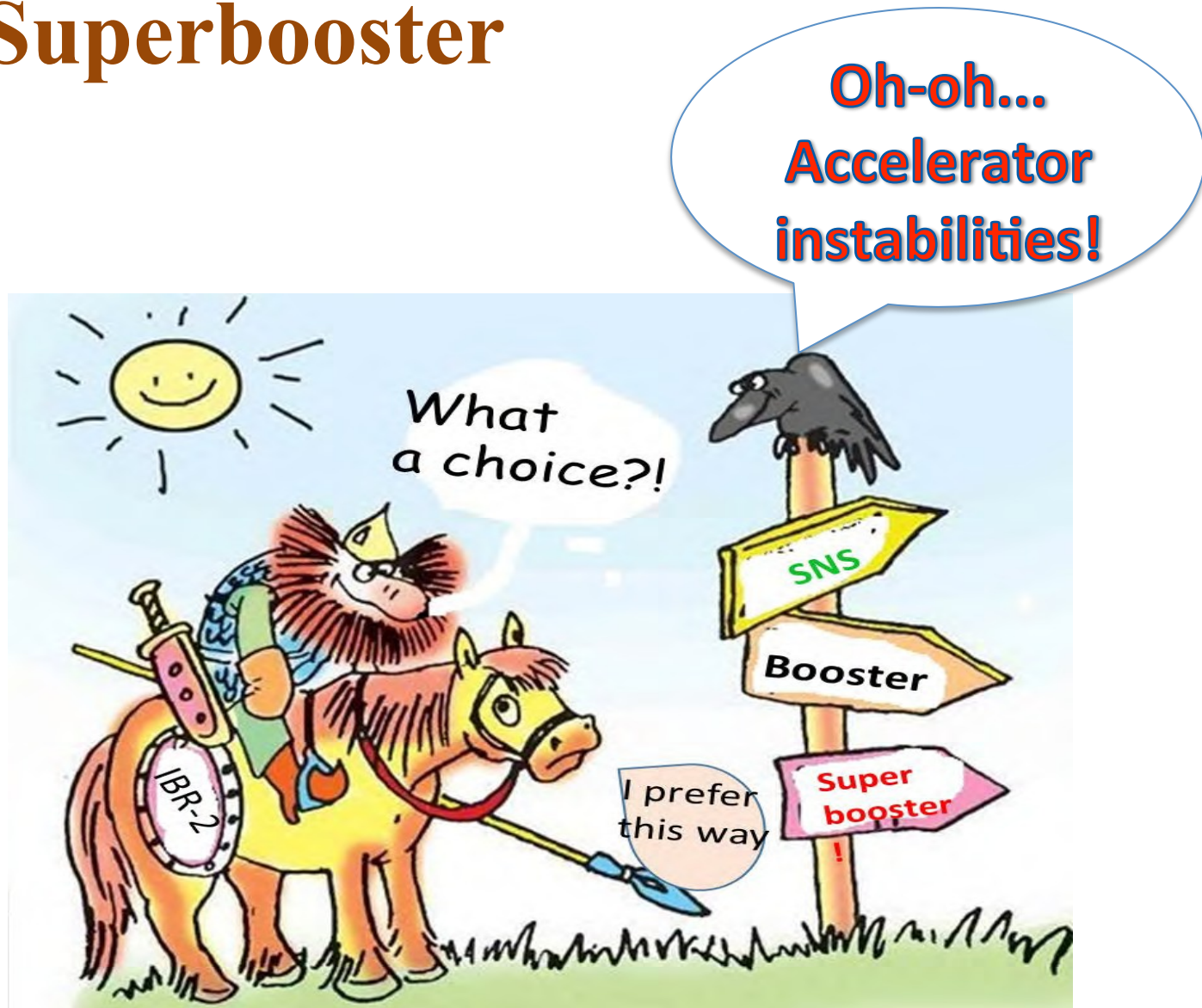


A. Vinogradov et al.

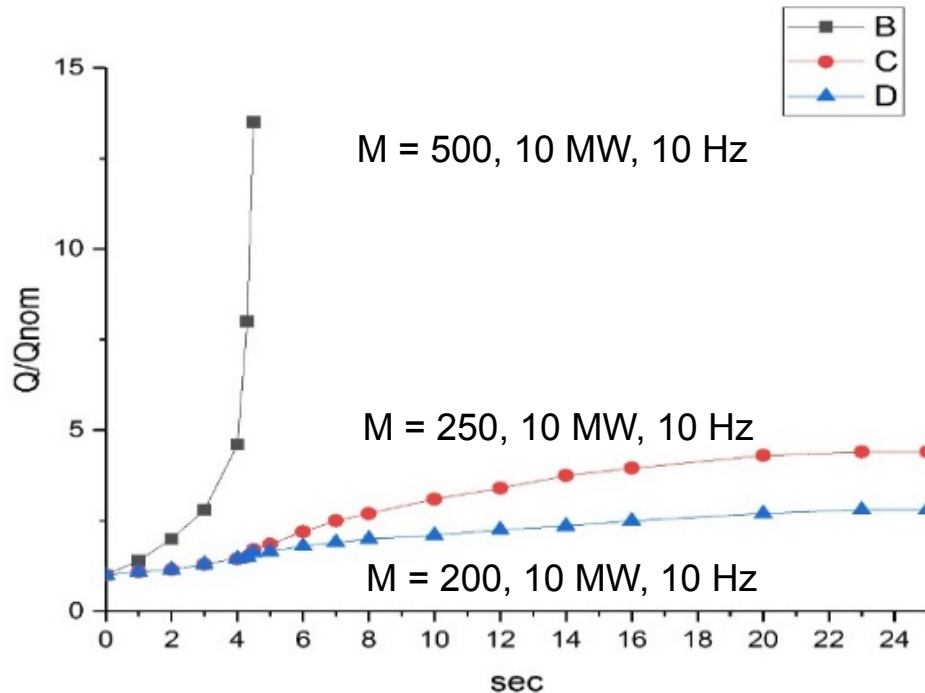
Superbooster



Superbooster



Superbooster NEPTUN: power deviation after proton beam breakdown



- It is permissible to triple the prompt power exceeding to the nominal value.
- A tenfold excess lead to damage of fuel elements.
- The reason is negative temperature reactivity feedback. When power is down, coolant decreases temperature of fuel that arises reactivity of the reactor core.

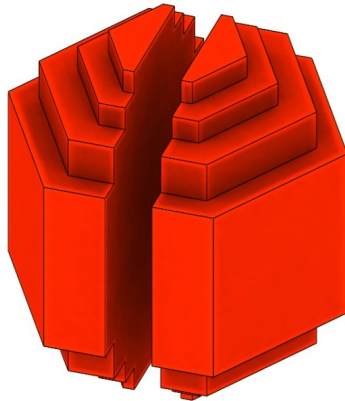
Each beam breakdown can lead to a long interruption of the subcritical system operation for the restoration of a normal regime.

=> Superbooster NEPTUN is highly depend on accelerator stability.

Decreasing of M up to 200 leads to increasing of E_p up to 2.5 GeV

Superbooster => IBR-3:

- no accelerator
- Power 10MW, $f=10\text{Hz}$, $\tau = (200-300)\mu\text{s}$
- Np core



Spherical core to reduce the amount of Np.

- Still the same neutron flux is expected

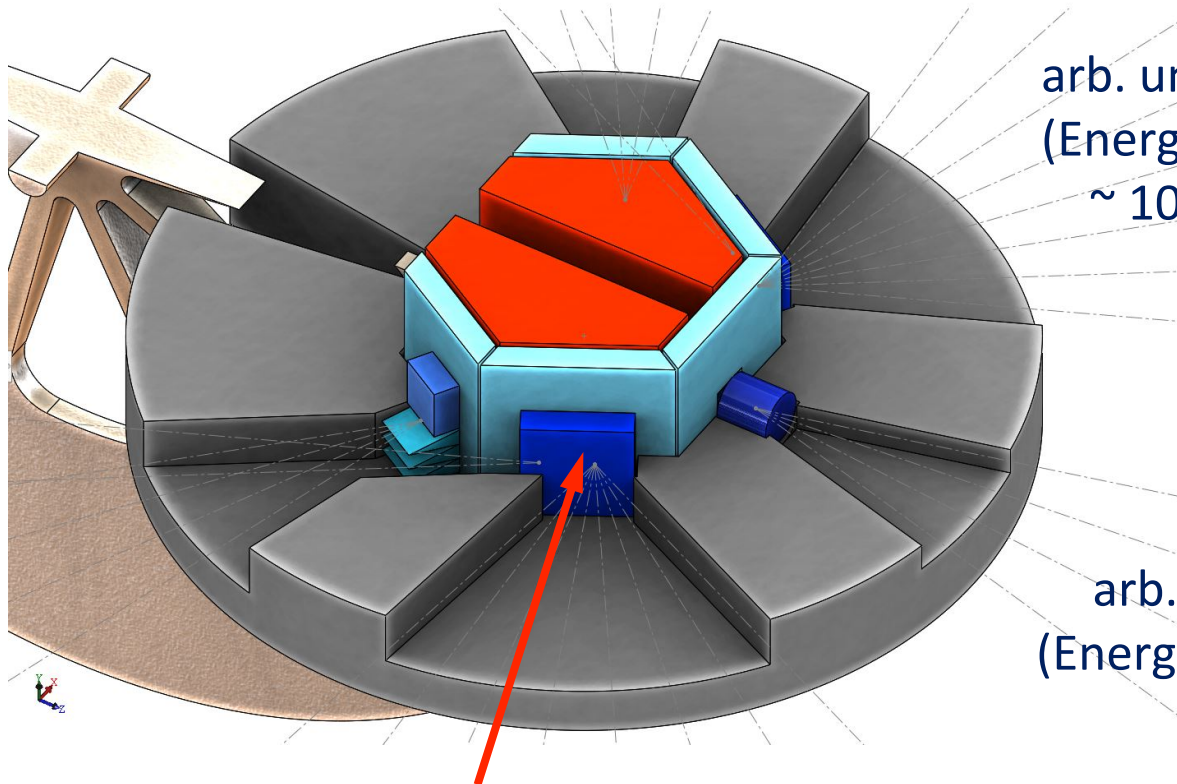
Moderators: neutron fluxes at "Neptun" neutron source

S. Kulikov et al.

arb. units: $2.86E-08$
(Energy range: 0 - 0.5 eV)

arb. units: $2.89E-08$
(Energy range: 0 - 0.5 eV)
 $\sim 10^{14}$ n/cm²/sec

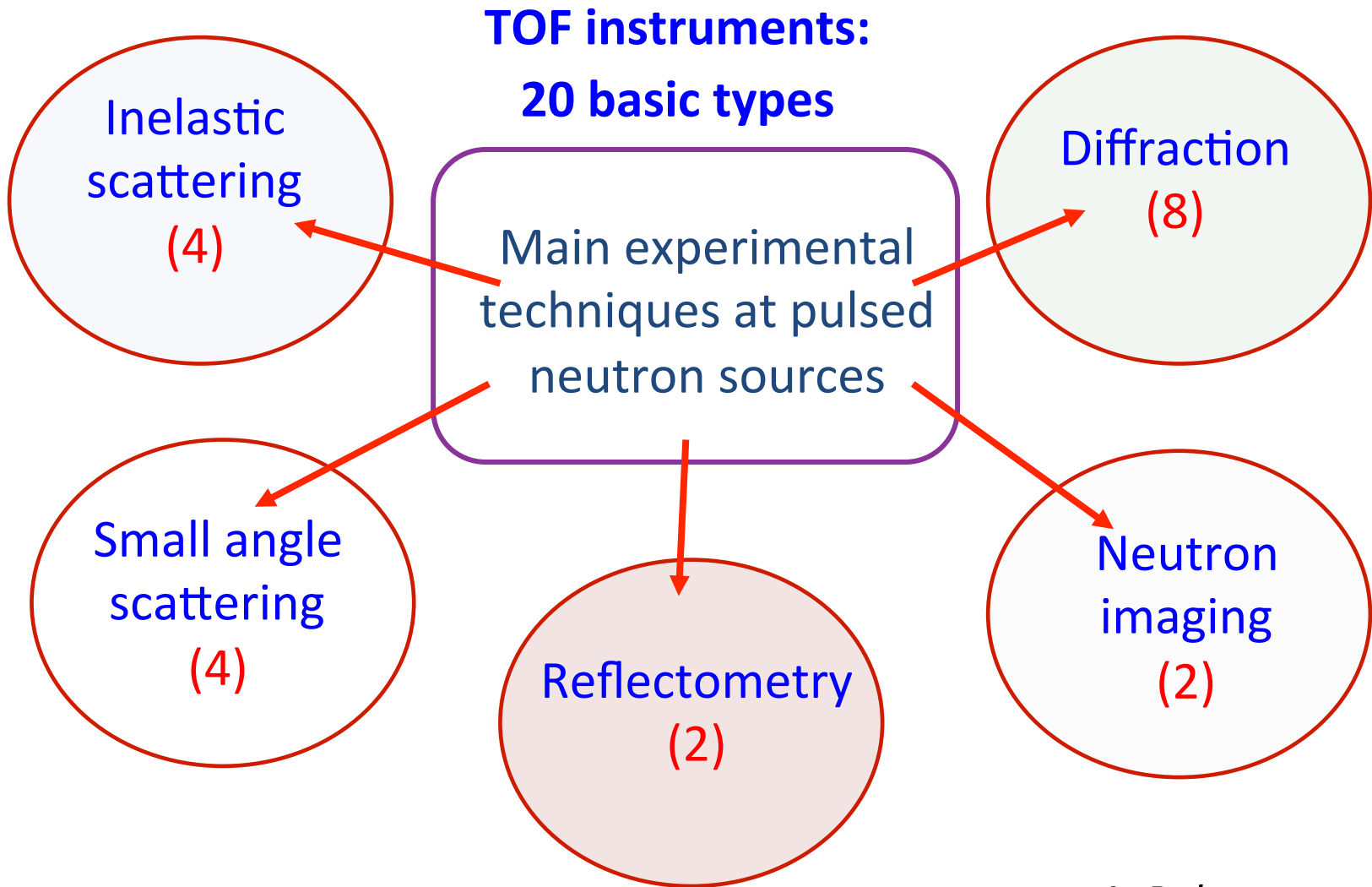
arb. units: $2.25E-08$
(Energy range: 0 - 0.5 eV)



Water moderators

(ambient temperature, size $d=120$ mm, $h=60$ mm)

Neutron instruments



Different requirements for different topics

Spectrum ↓	Diffraction	Spectroscopy	SANS, reflectometry, GISANS	Fund. physics	Imaging
Hot (0.3-1)Å	↕				
Thermal (1-3)Å	↕	↕			↕
Cold (3-20)Å	↕	↕	↕ SANS	↕	↕
Very cold (20-100)Å		↕	↕ GISANS	↕	
Ultra cold > 600Å				↕	

Beam divergences ↓	Diffraction	Spectroscopy	SANS	Reflect.	GISANS	Fund. physics	Imaging
Large 20mrad (1°)	↕	↕	↕	↕	↕	↕	
Small 1mrad (2-3')			↕	↕	↕		↕

Intensity-hungry instruments →

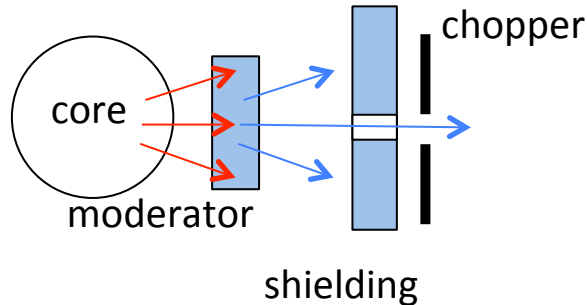
Brilliance-hungry instruments ←

Pulsed sources: pulse length (in μs) requirements

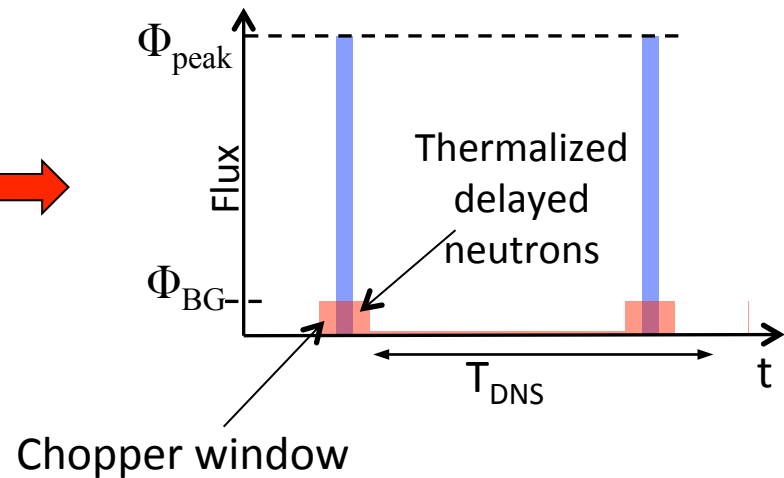
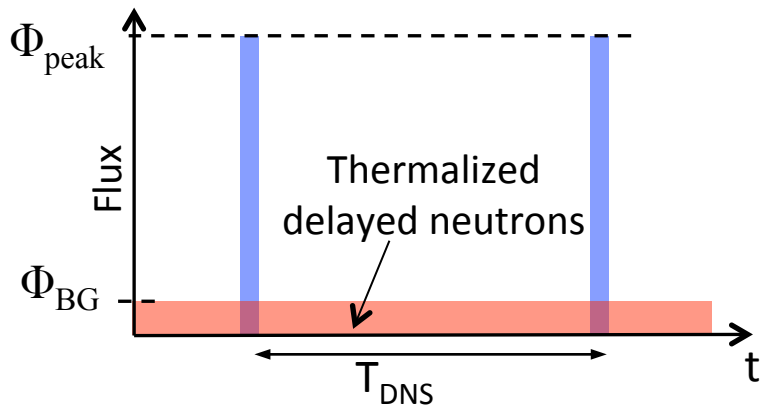
• Irradiation work:	∞
• SANS, NSE	2000 – 4000
• Reflectometry	500 – 2000
• Imaging	500
• Single crystal diffraction	100 – 500
• Powder diffraction	10 – 500
• Cold neutron spectroscopy	50 – 2000
• Thermal neutron spectroscopy	20 – 600
• Hot neutron spectroscopy	10 – 300
• Backscattering spectroscopy	10 – 100
• Electronvolt spectroscopy	1 – 10

Comment: There is a need for sufficiently long pulse to avoid intensity loss by excessive resolution!

Delayed neutrons: small amount, however a big problem...

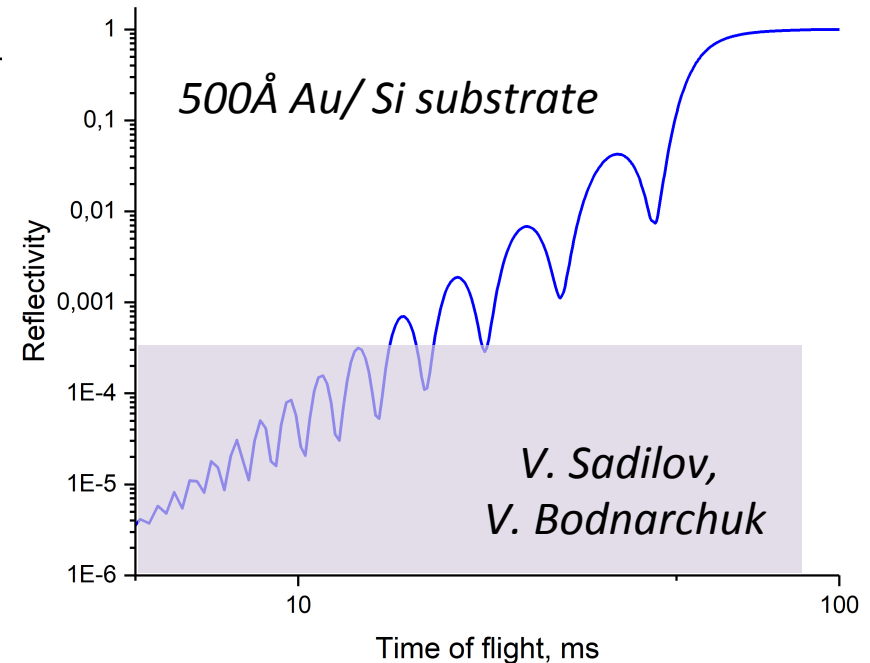
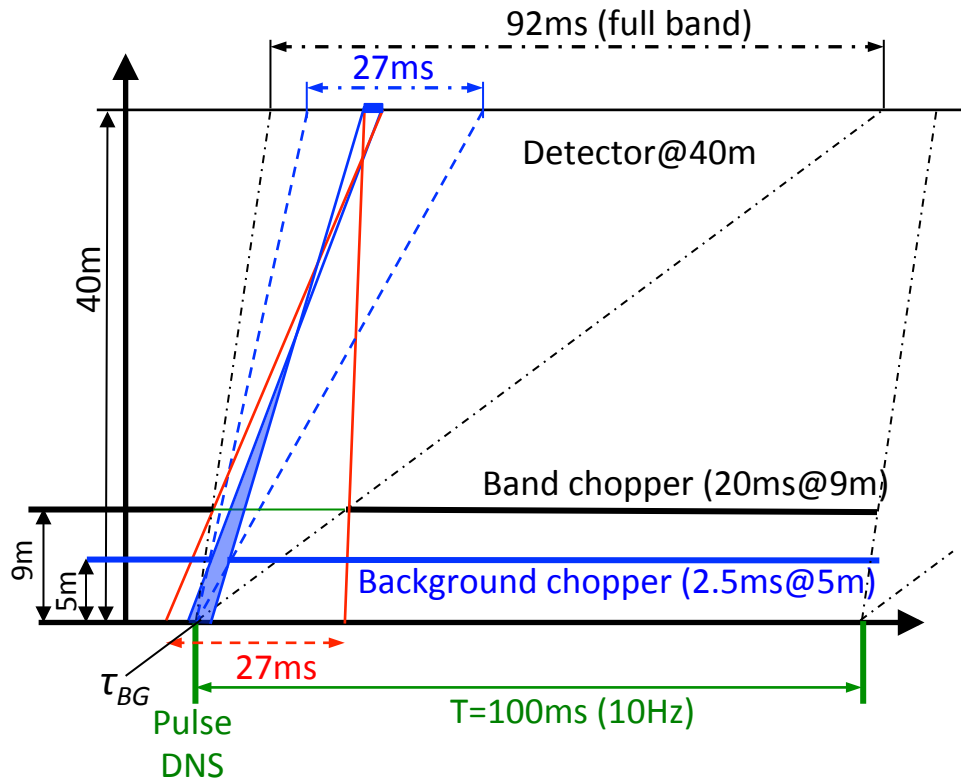


Solution: reduce the opening time of moderator to τ_{pulse} by a chopper:
=> no delayed neutrons outside the neutron pulse are emitted



Delayed neutrons: a possible solution

Chopper after the shielding: the time when detector is exposed to the moderated delayed neutrons is reduced to $\approx 1.15\tau_{BG}$ (broadening because of 5m to the source).



- Now only neutrons with a small λ -shift relative λ_0 are seen in the detector time channel.
- The spectrum inside τ_{BG} is practically the same Maxwellian as in the neutron pulse
- **These neutrons are not the BG anymore - they are a part of the signal.**

Int. Workshop “Advanced ideas & experiments for DNS-IV”

Recommendations for source:

1. The vector (in 2π) neutron flux density in the range of $(1\div 2)\cdot 10^{14}$ n/cm²/s at 10÷15 MW is taken as the key parameter of DNS-IV.
2. In cooperation with the Chief Designer (NIKIET), work on the conceptual project of DNS-IV should be carried out in two directions:
 - a) pulsed periodic reactor IBR-3;
 - b) pulsed neutron source driven by a proton accelerator with a multiplying target.
3. The fast neutron pulse duration in the range of 150÷200 μ s with 10 Hz repetition is considered to satisfy the majority of experimental research areas: diffraction, inelastic and small-angle scattering, reflectometry, tomography.
4. Provision should be made in the design of the source for a system of choppers to suppress background between pulses and to form an optimal pulse length.

Int. Workshop “Advanced ideas & experiments for DNS-IV”

Recommendations for source:

5. It is necessary to develop and provide a cost estimate of experimental stations including infrastructure and personnel costs.
6. The development of the concept of neutron moderators on DNS-IV should be considered as a top-priority task for 2019.
7. It is suggested to develop proposals for ultracold and very cold neutrons factory.

Conclusions

1. The WSG group started to work 6 months ago
2. Steady progress
3. Two alternative concepts; both are under development and will be evaluated by NIKIET during 2019.
4. Instrumentation
5. Moderators
6. Scientific program