

Moscow Institute of Physics and Technology NRC "Kurchatov Institute" - ITEP



Kirill Gubarev Generalizing 11D supergravity

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Motivation

- What are consistent backgrounds for the superstring?
 - 10D supergravity solutions (RNS string)
 - 10D generalized supergravity solutions (GS string)
- What are consistent backgrounds for the supermembrane?
 - 11D supergravity solutions (GS string)
 - some 11D generalized supergravity solutions? (GS string)



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Image: A matrix

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RNS string and Weyl invariance

RNS action

$$S = \frac{1}{4\pi\alpha'} \int d^2\sigma \sqrt{g} (G_{mn}\partial_J x^m \partial^J x^n + iB_{mn}\epsilon^{IJ}\partial_I x^m \partial_J x^n + \alpha' \Phi R^{(2)}), \tag{1}$$

to be Weyl invariant

$$=-\frac{1}{2\alpha'}\beta_{mn}(G)\partial_{J}x^{m}\partial^{J}x^{n}-\frac{i}{2\alpha'}\beta_{mn}(B)\epsilon^{IJ}\partial_{I}x^{m}\partial_{J}x^{n}-\frac{1}{2}\beta(\Phi)R^{(2)}=0,\qquad(2)$$

one loop calculation reproduces NS-NS supergravity equations

$$\beta^{1loop}(\Phi) = \alpha' \left(R - \frac{1}{12}H^2 + 4\nabla^m \nabla_m \Phi - 4(\nabla\Phi)^2\right) = 0,$$

$$\beta^{1loop}_{mn}(G) = \alpha' \left(R_{mn} - \frac{1}{4}H_{mkl}H_n^{kl} + 2\nabla_m \nabla_n \Phi\right) = 0,$$

$$\beta^{1loop}_{mn}(B) = \alpha' \left(\frac{1}{2}\nabla_k H^{kmn} - H^{kmn}\nabla_k \Phi\right) = 0,$$

(3)

where $H_{mnk} = 3\nabla_{[m}B_{nk]}$ and ∇_m covariant with respect to G_{mn} .





Рис.: Ordinary and generalized supergravities in 10D



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GS string and κ -symmetry

GS action

$$S = \int d^2 \sigma \sqrt{-G} - \int_{\Sigma} B, \qquad G = \det G_{IJ}, \qquad (4)$$

$$G_{IJ} = E_I^{\ a} E_J^{\ b} \eta_{ab} , \qquad E_I^{\ A} = \partial_I z^M E_M^{\ A}(z) , \qquad z^M = (x^m, \theta^\mu) .$$
(5)

 κ -symmetry transformation (IIB example)

$$\delta_{\kappa} z^{M} E_{M}{}^{a} = 0, \qquad \delta_{\kappa} z^{M} E_{M}{}^{\alpha i} = \frac{1}{2} (1+\Gamma)^{\alpha i}{}_{\beta j} \kappa^{\beta j}, \qquad \Gamma = \frac{1}{2\sqrt{-G}} \varepsilon^{IJ} E_{I}{}^{a} E_{J}{}^{b} \gamma_{ab} \sigma^{3}. \tag{6}$$



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GS string and κ -symmetry

Constraints on the supertorsion $T^A = dE^A + E^B \wedge \Omega_B{}^A$ and H = dB

$$H_{\beta j\gamma k\alpha i}(1+\Gamma)^{\alpha i}{}_{\delta l}=0\,, \qquad (7)$$

$$E_{I}^{a} \left[\sqrt{-G} G^{IJ} T_{\alpha i \beta j a} - \varepsilon^{IJ} H_{a \alpha i \beta j} \right] (1 + \Gamma)^{\alpha i}{}_{\gamma k} = 0, \qquad (8)$$

+ Bianchi identities are solved as

[Tseytlin, Wulff (2016)]

$$R - \frac{1}{12} H^2 + 4 \nabla^m X_m - 4 X_m X^m = 0, \qquad (9)$$

$$R_{mn} - \frac{1}{4} H_{mkl} H_n^{kl} + \nabla_m \mathbf{X}_n + \nabla_n \mathbf{X}_m = 0, \qquad (10)$$

$$\frac{1}{2}\nabla_k H^{kmn} - H^{kmn} X_k - \nabla^m X^n + \nabla^n X^m = 0, \qquad (11)$$

where $X_m = I_m + \nabla_m \Phi - B_{mn} I^n$ is a Killing vector.



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Рис.: Ordinary and generalized supergravities in 10D



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RNS string and scale invariance

For the scale invariance of RNS string G_{mn} and B_{mn} must satisfy

[Arutynov, Frolov, Hoare, Roiban, Tseytlin (2015)]

$$\frac{1}{\alpha'}\beta_{mn}^{1loop}(G) = R_{mn} - \frac{1}{4}H_{mkl}H_n^{kl} = -\nabla_m X_n - \nabla_n X_m,$$

$$\frac{1}{\alpha'}\beta_{mn}^{1loop}(B) = \frac{1}{2}\nabla_k H^{kmn} = H^{kmn}X_k + \nabla^m Y^n - \nabla^n Y^m = 0,$$
(12)

which is true for the G_{mn} and B_{mn} solving generalized supergravity equations ($Y_m = X_m$). The stress energy tensor

$$< T^{J}{}_{J} >= 2\nabla^{J} (X_{m} \partial_{J} x^{m} + \epsilon_{IJ} Y_{m} \partial^{J} x^{m})$$
⁽¹³⁾

that can be canceled only with non-local generalization of Fradkin-Tseytlin counterterm.

[Fernandez-Melgarejo, Sakamoto, Sakatani, Yoshida (2018)]



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Рис.: Ordinary and generalized supergravities in 10D



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 $\eta\text{-}deformation$ of $\textit{AdS}_5\times\mathbb{S}^5$ - the real way the generalized supergravity was discovered

- GS string on $AdS_5 \times S^5$ is integrable [Bena, Polchinski, Roiban (2004)]
- GS string is integrable on Yang–Baxter (η -) deformed background AdS $_5 \times \mathbb{S}^5$

[Vicedo, Delduc, Magro (2013)]

$$S = -\frac{(1+\eta^2)^2}{2(1-\eta^2)} \int d\tau d\sigma P_{-}^{ab} \left[A_a \cdot d \circ \frac{1}{1-\eta R_g \circ d} (A_b) \right]$$
(14)

 η -deformed AdS₅ \times \mathbb{S}^5

does not solve ordinary supergravity equations [Arutyunov, Borsato, Frolov (2015)]
 solves equations of generalized supergravity [Arutynov, Frolov, Hoare, Roiban, Tseytlin (2015)]
 formally T-dual to HT background (IIA solution) [Hoare, Tseytlin (2015)]



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Рис.: Ordinary and generalized supergravities in 10D



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Double field theory

DFT action

$$S_{DFT} = \int d^{20} \mathbb{X} e^{-2d} \mathbb{Q}, \tag{15}$$

where the DFT curvature

$$\mathbb{Q} = \mathcal{H}^{AB} \mathcal{F}_{A} \mathcal{F}_{B} + \mathcal{F}_{ABC} \mathcal{F}_{DEF} \left(\frac{1}{4} \mathcal{H}^{AD} \eta^{BE} \eta^{CF} - \frac{1}{12} \mathcal{H}^{AD} \mathcal{H}^{BE} \mathcal{H}^{CF}\right) - \mathcal{F}_{A} \mathcal{F}^{A} - \frac{1}{6} \mathcal{F}_{ABC} \mathcal{F}^{ABC},$$
(16)

$$\mathcal{H}_{AB} = \begin{pmatrix} h_{ab} & 0\\ 0 & h^{ab} \end{pmatrix}, \quad \eta_{AB} = \begin{pmatrix} 0 & \delta^b_a\\ \delta^a_b & 0 \end{pmatrix}, \tag{17}$$

and DFT fluxes

$$\mathcal{L}_{E_A} E_B{}^M = \mathcal{F}_{AB}{}^C E_C{}^M, \quad \mathcal{L}_{E_A} d = \frac{1}{2} \mathcal{F}_A, \tag{18}$$

$$\mathcal{F}_{ABC} = 3E_{N[C}\partial_A E^N{}_{B]}, \qquad \mathcal{F}_A = E_{NA}\partial^B E^N{}_B + 2\partial_A d. \tag{19}$$

For the consistency \forall DFT fields f, g must satisfy section constraints

$$\eta^{MN}\partial_M f \partial_N g = 0, \quad \eta^{MN}\partial_M \partial_N f = 0.$$
⁽²⁰⁾



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10D supergravity from DFT

Supergravity parametrization

$$E_M{}^A = \begin{pmatrix} e_a^m & 0\\ -e_a^k B_{km} & e_a^m \end{pmatrix}, \qquad d = \phi - \frac{1}{4} \log G, \tag{21}$$

$$\partial_M = (\partial_m, \partial^m \equiv 0), \quad \partial_A = E^M_A \partial_M.$$
 (22)

The DFT EOMs \Leftrightarrow 10D supergravity EOMs

$$\delta d: \qquad R - \frac{1}{12} H^2 + 4 \nabla^m \nabla_m \phi - 4 (\nabla \phi)^2 = 0, \qquad (23)$$

$$\delta e^{m}{}_{a}: \qquad R_{mn} - \frac{1}{4} H_{mkl} H_{n}{}^{kl} + \nabla_{m} \nabla_{n} \phi + \nabla_{n} \nabla_{m} \phi = 0, \qquad (24)$$

$$\delta b_{mn}: \qquad \frac{1}{2} \nabla_k H^{kmn} - H^{kmn} \nabla_k \phi = 0.$$
⁽²⁵⁾



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Yang-Baxter deformation

Yang-Baxter deformation — local O(d,d) rotation

$$E'_{M}{}^{A} = O[\beta]_{M}{}^{N}E_{N}{}^{B}, \qquad O_{M}{}^{N} = \begin{pmatrix} \delta^{n}_{m} & -\beta^{nm} \\ 0 & \delta^{m}_{m} \end{pmatrix} = \exp(\beta^{mn}T_{mn}), \tag{26}$$

with bi-Killing ansatz

$$\beta^{mn} = r^{ab} k^m_a k^n_b. \tag{27}$$

Fluxes deform as

$$\delta \mathcal{F}_{ABC} \propto CYBE, \quad \delta \mathcal{F}_a = 2I^m b_{mn} e_a^n, \quad \delta \mathcal{F}^a = 2I^a,$$
(28)

where

$$I^{m} = \nabla_{k}\beta^{km} = \frac{1}{2}r^{ab}f_{ab}{}^{c}k_{c}{}^{m}, \quad f_{de}{}^{[a}r^{b]d}r^{c]e} = 0 \text{ (CYBE)}.$$
(29)



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DFT Bianchi identities and constraints on I_m

DFT Bianchi identities

$$0 = \partial_{[A} \mathcal{F}_{BCD]} - \frac{3}{4} \mathcal{F}_{[AB}{}^{E} \mathcal{F}_{CD]E},$$

$$0 = 2\partial_{[A} \mathcal{F}_{B]} + \partial^{C} \mathcal{F}_{CAB} - \mathcal{F}^{C} \mathcal{F}_{CAB},$$

$$0 = \partial^{A} \mathcal{F}_{A} - \frac{1}{2} \mathcal{F}^{A} \mathcal{F}_{A} + \frac{1}{12} \mathcal{F}^{ABC} \mathcal{F}_{ABC}.$$
(30)

After the deformation

$$\begin{aligned} \mathcal{F}'_{ABC} &= \mathcal{F}_{ABC}, \\ \mathcal{F}'_{A} &= \mathcal{F}_{A} + X_{A}, \end{aligned}$$
 (31)

for the BIs to be satisfied X_M must obey

$$\mathcal{L}_X E_A'^M = 0, \quad \mathcal{L}_X d' = 0, \quad X_M X^M = 0.$$
 (32)



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10D generalized supergravity from flux shift in DFT



10D generalized supergravity EOMs

$$\delta d: \qquad R - \frac{1}{12} H^2 + 4 \nabla^m X_m - 4 X_m X^m = 0, \qquad (33)$$

$$\delta e^{m}{}_{a}: \qquad R_{mn}e^{na} - \frac{1}{4} H_{nkm}H^{nk}{}_{l}e^{la} + \nabla_{m}X_{n}e^{na} + \nabla_{n}X_{m}e^{na} = 0,$$
(34)

$$\delta b_{mn}: \qquad \frac{1}{2}\nabla_k H^{kmn} - H^{kmn} X_k - \nabla^m X^n + \nabla^n X^m = 0, \qquad (35)$$

where $X_m = I_m + \nabla_m \Phi - B_{mn} I^n$ is a Killing vector.

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Generalization to 11D case

In string theory, the appearance of generalized supergravity can be associated with a violation

Weyl invariance supergravity scale invariance generalized supergravity

In M-theory, there is no Weyl symmetry for membranes. This was one of the arguments why 11D generalized supergravity should not exist. However, when the supergravity is considered in a split form, there is a violation

GL(11) invariance supergravity GL(d)×GL(11-d) invariance generalized supergravity

that allows to build generalization.



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Exceptional field theory

Exceptional field theory (ExFT) - this is a generalization of 11-dimensional supergravity in split form d + (11 - d), explicitly covariant with respect to the *U*-duality group.



$$m_{MN} = \mathcal{E}_M{}^A \mathcal{E}_N{}^B m_{AB} \tag{36}$$

Parametrization that corresponds to supergavity

$$m_{MN} = \mathcal{E}_{M}{}^{A}\mathcal{E}_{N}{}^{B}m_{AB}, \quad \mathcal{E}_{M}{}^{A} = e^{\frac{1}{10}} \begin{bmatrix} e^{-\frac{1}{2}}e_{m}{}^{a} & e^{\frac{1}{2}}V^{a} \\ 0 & e^{\frac{1}{2}} \end{bmatrix}, \quad e = \det(e^{a}_{m}), \quad V^{m} = \frac{1}{3!}\varepsilon^{mnkl}C_{nkl}, \quad (37)$$

section condition

$$\epsilon^{MNKLP}\epsilon_{MQRST} \partial_{NK} \bullet \otimes \partial_{LP} \bullet = 0, \quad \partial_{5m} = \partial_m, \quad \partial_{mn} = 0 \quad (SUGRA), \tag{38}$$



Ansatz for fields

We consider theory on the background $M_{11} = M_7 \times M_4$

$$g_{\mu\nu} = e^{-2\phi(x^{m})} e^{\frac{2}{5}} \bar{g}_{\mu\nu}(y^{\mu}), \quad m_{MN} = e^{-\phi} e^{\frac{1}{5}} M_{MN},$$

$$A_{\mu}{}^{MN} = 0, \qquad B_{\mu\nu M} = 0.$$
(39)

Lagrangian

$$m \mathcal{L}' = Y_{AB} Y_{CD} m^{AC} m^{BD} - \frac{1}{2} Y_{AB} Y_{CD} m^{AB} m^{CD} + 32 Z^{ABC} Z^{DEF} m_{AD} m_{BE} m_{CF} + 32 Z^{ABC} Z^{DEF} m_{AC} m_{BD} m_{EF} - \frac{7}{3} \theta_{AB} \theta_{CD} m^{AC} m^{BD} + \bar{e} \mathcal{R}[\bar{g}_{(7)}],$$
(40)

Fluxes

$$\mathcal{L}_{E_{AB}} E^{M}{}_{C} = \mathcal{F}_{ABC}{}^{D} E^{M}{}_{D}, \qquad (41)$$

$$\mathcal{F}_{ABC}{}^{D} = \frac{3}{2} E_{N}{}^{D} \partial_{[AB} E^{N}{}_{C]} - E^{M}{}_{C} \partial_{MN} E^{N}{}_{[B} \delta^{D}{}_{A]} - \frac{1}{2} E^{M}{}_{[B|} \partial_{MN} E^{N}{}_{|A]} \delta^{D}{}_{C}, \tag{42}$$

$$\mathcal{F}_{ABC}{}^{D} = \frac{3}{2} Z_{ABC}{}^{D} - \frac{1}{2} \theta_{[AB} \delta_{C]}{}^{D} + \delta_{[A}{}^{D} Y_{B]C},$$
(43)



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Generalized Yang-Baxter deformation

$$E_{M}^{\prime A} = O[\Omega]_{M}{}^{N}E_{N}{}^{B}, \qquad O[\Omega] = \begin{bmatrix} \delta_{m}{}^{n} & 0\\ \frac{1}{3!}\epsilon_{mpqr}\Omega^{pqr} & 1 \end{bmatrix}.$$
 (44)

For the deformation parameters, we use the poly-Killing ansatz

$$\Omega^{mnk} = \frac{1}{6} \rho^{\alpha\beta\gamma} k_{\alpha}{}^{n} k_{\beta}{}^{n} k_{\gamma}{}^{k}, \qquad (45)$$

satisfying the condition (a generalization of the Yang-Baxter equation)

$$6\rho^{[i^2|i^7j_1}\rho^{|i^3i^4|j^2}f_{j_1j_2}^{|i^5]} + \rho^{j_1j_2[i^2}\rho^{i^3i^4i^5]}f_{j_1j_2}^{i^7} = 0.$$
(46)

Under such deformations, the fluxes transform as

$$\delta \mathcal{F}_{ABC}{}^{D} = \frac{1}{4} E^{m}{}_{C} E^{n}{}_{A} E^{k}{}_{B} E_{l}{}^{E} J^{lp} \epsilon_{kmnp}, \quad J^{mn} = k_{i_{1}}{}^{m} k_{i_{4}}{}^{n} \rho^{i_{1}i_{2}i_{3}} f_{i_{2}i_{3}}{}^{i_{4}}, \tag{47}$$

that under reduction gives

$$J^{mn} = \rho^{i_1 i_2 i_3} f_{i_2 i_3}^{i_4} k_{i_1}^{m} k_{i_4}^{n} \xrightarrow[i=(*,\alpha), m=(*,\bar{m})]{I^{\bar{m}} \equiv J^{*\bar{m}} = \rho^{*\alpha\beta} f_{\alpha\beta}^{\gamma} k_{\gamma}^{\bar{m}} (gCYBE).$$
(48)



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Constraints from SL(5) ExFT BI

SL(5) ExFT Bianchi identities

$$\frac{1}{2}\partial_{AB}\mathcal{F}_{DFC}^{E} + \frac{1}{2}\partial_{BC}\mathcal{F}_{DFA}^{E} - \frac{1}{2}\delta^{E}_{A}\partial_{CG}\mathcal{F}_{DFB}^{G} - \frac{1}{4}\delta^{E}_{C}\partial_{BG}\mathcal{F}_{DFA}^{G} + \frac{1}{4}\delta^{E}_{C}\partial_{AG}\mathcal{F}_{DFB}^{G} + \frac{1}{2}\delta^{E}_{B}\partial_{CG}\mathcal{F}_{DFA}^{G} - \frac{1}{2}\partial_{AC}\mathcal{F}_{DFB}^{E} - \mathcal{F}_{BGC}^{E}\mathcal{F}_{DFA}^{G} + \mathcal{F}_{AGC}^{E}\mathcal{F}_{DFB}^{G}^{G} + \mathcal{F}_{ABC}^{E}\mathcal{F}_{DFC}^{G}^{G} - \mathcal{F}_{ABC}^{G}\mathcal{F}_{DFG}^{E} - \frac{1}{2}\partial_{DF}\mathcal{F}_{ABC}^{E} = 0$$
(49)

also satisfied after the deformation

$$\mathcal{F}_{ABC}^{\prime D} = \mathcal{F}_{ABC}{}^{D} + \delta \mathcal{F}_{ABC}{}^{D}, \qquad (50)$$

if

$$L_{e_{a}}J^{kl} + J^{nl}\partial_{n}\phi e_{a}{}^{k} = 0, \qquad J^{mn}\partial_{n}\phi = 0,$$

$$\nabla_{m}(e^{-\phi}J^{mn}) = 0, \qquad J^{m[n}J^{kl]} = 0,$$

$$\nabla_{[m}Z_{n]} - \frac{1}{3}J^{kl}F_{mnkl} = 0,$$

$$\nabla_{k}\left(e^{-\phi}J^{k[l}V^{p]}\right) = 0,$$

$$\nabla_{k}(J^{(pl)}V^{k}) - \nabla_{k}(V^{(p}J^{l)k}) = 0.$$
(51)
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where

$$Z_m = \partial_m \phi - \frac{2}{3} \varepsilon_{mnkl} I^{nk} V^l.$$
(53)

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Construction of generalized supergravity equations



flux deformation

$$\delta \mathcal{F}_{mnk}{}^{l} = \frac{1}{4} \epsilon_{mnkp} J^{lp}.$$
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Generalization of 11D supergravity

Generalized equations

$$0 = \mathcal{R}_{mn}[h_{(4)}] - 7 \,\tilde{\nabla}_{(m}Z_{n)} - \frac{1}{3}h_{mn}(\nabla V) + 8(1 + V^{2}) \left(S_{mn}J^{k}{}_{k} - 2J^{k}{}_{(m}J_{n)k}\right) + 4V_{m}V_{n} \left(J^{kl}J_{kl} - 2J^{kl}J_{lk}\right) + 4V_{k}V_{l} \left(4J_{(m}{}^{k}J_{n)}{}^{l} - J^{k}{}_{(m}J^{l}{}_{n)} - 2S^{kl}S_{mn}\right) + 8V_{k}V_{(m} \left(2J^{l}{}_{n})J^{k}{}_{l} - 2S_{n}{}^{k}J^{l}{}_{l} + J^{kl}J_{n)l}\right),$$

$$0 = \frac{1}{7}e^{2\phi} \mathcal{R}[\bar{g}_{(7)}] + \frac{1}{6}(\nabla V)^{2} + \tilde{\nabla}^{m}Z_{m} - 6Z_{m}Z^{m} - 2J^{mn}J_{mn} + \frac{4}{3}J_{mn}J^{nm},$$

$$0 = \tilde{\nabla}^{m}F_{mnkl} - 6Z^{m}F_{mnkl} + 6\left(2J^{pm}C_{m[nk}J_{l]p} - J^{pm}J_{p[n}C_{kl]m}\right),$$

$$0 = \mathcal{R}_{\mu\nu}[\bar{g}_{(7)}] - \frac{1}{7}\bar{g}_{\mu\nu}\mathcal{R}[\bar{g}_{(7)}],$$
(55)

where $S^{mn} = J^{(mn)}$, $F_{mnkl} = 4\partial_{[m}C_{nkl]}$ and

$$\tilde{\nabla}_m = \nabla_m - \partial_m \phi \,. \tag{56}$$

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Example of solution - deformed $AdS_4 \times S^7$

AdS₄ Killing vectors

$$P_{a} = \partial_{a}, \qquad K_{a} = x^{2}\partial_{a} + 2x_{a}D, D = -x^{m}\partial_{m}, \qquad M_{ab} = x_{a}\partial_{b} - x_{b}\partial_{a},$$
(57)

где a, b = 0, 1, 2 и $m, n = 0, 1, 2, z, x^2 = \eta_{mn} x^m x^n, x_a = \eta_{ab} x^b.$ deformation along $D \land M \land M$:

$$\Omega = \frac{4}{R^3} \, \rho_a \epsilon^{abc} \, D \wedge M_{bd} \wedge M_c{}^d$$

gives deformed background

$$ds^{2} = \frac{R^{2}}{4z^{2}} K^{\frac{2}{3}} \left\{ dx_{a} dx^{a} + \frac{1}{z^{2}} \rho_{a} x^{a} x^{b} dx_{b} dz + \left(1 - \frac{x_{a} x^{a} \rho_{b} x^{b}}{z^{3}}\right) dz^{2} \right\} + R^{2} K^{-\frac{1}{3}} d\Omega_{(7)}^{2},$$

$$F = -\frac{3}{8} \frac{R^{3}}{z^{4}} K^{2} \left(1 + \frac{1}{12} \frac{x_{a} x^{a} \rho_{b} \rho_{c} x^{b} x^{c}}{z^{4}}\right) dx^{0} \wedge dx^{1} \wedge dx^{2} \wedge dz.$$
(58)

solving generalized equations if

$$GYBE = \rho^a \rho_a = 0. \tag{59}$$

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THANK YOU FOR YOUR ATTENTION!



Puc.: "Deformations open the way to the world of new knowledge"



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Equations of motion

From C-frame lagrangian

$$\bar{e}^{-1}h^{-\frac{1}{2}}\mathcal{L} = e^{-5\phi}\mathcal{R}[\bar{g}_{(7)}] + e^{-7\phi}\left(\mathcal{R}[h_{(4)}] + 42h^{mn}\partial_m\phi\partial_n\phi + \frac{1}{2}\nabla_m V^m \nabla_n V^n\right),\tag{60}$$

we obtain the equations of motion for dynamic fields ϕ, h_{mn} и V_m

$$\begin{split} \delta\phi &: \quad \frac{5}{7}e^{2\phi} \mathcal{R}[\bar{g}_{(7)}] + \mathcal{R}[h_{(4)}] + 12 \nabla_m \nabla_n \phi \ h^{mn} - 42 \nabla_m \phi \nabla_n \phi \ h^{mn} + \frac{1}{2} (\nabla V)^2 = 0, \\ \delta V^m &: \quad \partial_m (\nabla V) - 7 (\nabla V) \partial_m \phi = 0, \\ \delta h^{mn} &: \quad \mathcal{R}_{mn}[h_{(4)}] - 7 \partial_m \phi \ \partial_n \phi + 7 \nabla_m \nabla_n \phi \\ &+ h_{mn} \left(-\frac{1}{2}e^{2\phi} \mathcal{R}[\bar{g}_{(7)}] - \frac{1}{2} \mathcal{R}[h_{(4)}] + 28 \ \partial_k \phi \ \partial_l \phi \ h^{kl} - 7 \nabla_k \nabla_l \phi \ h^{kl} + \frac{1}{4} (\nabla V)^2 \right) = 0, \end{split}$$

$$(61)$$



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Generalized Lie derivatives

DFT

$$\mathcal{L}_{\Lambda}V^{M} = \Lambda^{N}\partial_{N}V^{M} - V^{N}\partial_{N}\Lambda^{M} + \eta^{MN}\eta_{KL}\partial_{N}\Lambda^{K}V^{L},$$

$$\mathcal{L}_{\Lambda}d = \Lambda^{M}\partial_{M}d - \frac{1}{2}\partial_{M}\Lambda^{M},$$
 (62)

ExFT

$$\mathcal{L}_{\Lambda}V^{M} = \frac{1}{2}\Lambda^{KL}\partial_{KL}V^{M} - V^{L}\partial_{LK}\Lambda^{MK} + \frac{1}{4}V^{M}\partial_{KL}\Lambda^{KL} + \lambda_{V}\partial_{KL}\Lambda^{KL}V^{M},$$
(63)



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QYBE

$$R_{12}(u-v)R_{13}(u)R_{23}(v) = R_{23}(v)R_{13}(u)R_{12}(u-v).$$
(64)

In classical limit

$$R_{ij} = \mathbb{1} + \epsilon r_{ij} + \mathcal{O}(\epsilon^2).$$
(65)

CYBE

$$[r_{12}, r_{13}] + [r_{12}, r_{23}] + [r_{13}, r_{23}] = 0.$$
(66)

using

$$[\mathbf{r}_{12},\mathbf{r}_{13}] = \mathbf{r}^{\alpha_1\beta_1}\mathbf{r}^{\alpha_1\beta_1} \cdot [\mathbf{e}_{\alpha_1},\mathbf{e}_{\alpha_2}] \otimes \mathbf{e}_{\beta_1} \otimes \mathbf{e}_{\beta_2}, \tag{67}$$

we obtain

$$r^{\alpha_{1}[\beta_{1}]}r^{\alpha_{2}[\beta_{2}}f_{\alpha_{1}\alpha_{2}}^{\gamma]} = 0.$$
(68)

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Tetrahedron equation

$$R_{123}R_{124}R_{134}R_{234} = R_{234}R_{134}R_{124}R_{123}.$$
 (69)

In classical limit

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$$R_{ijk} = \mathbb{1} + \epsilon r_{ijk} + \mathcal{O}(\epsilon^2).$$
(70)

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classical tetrahedron equation

$$\epsilon^{2} \cdot | \qquad [r_{123}, r_{124}] + [r_{123}, r_{134}] + [r_{123}, r_{234}] + [r_{124}, r_{134}] + [r_{124}, r_{234}] + [r_{134}, r_{234}] = 0.$$
(71)

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