

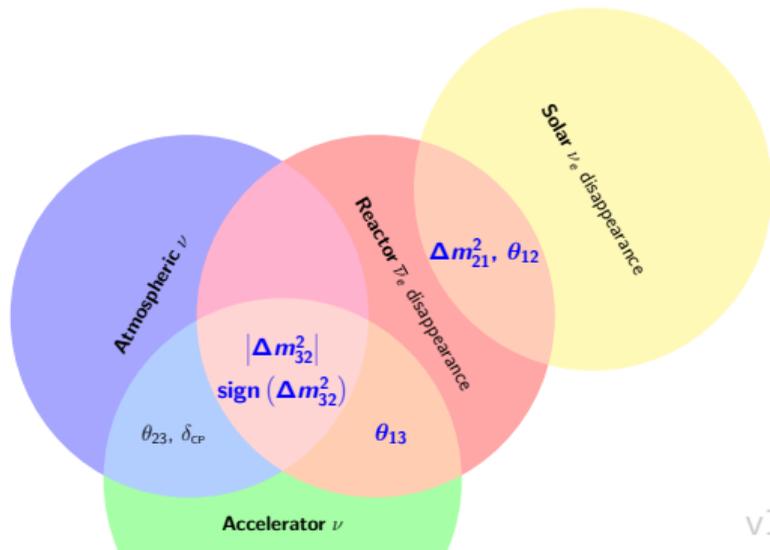


THE JUNO PROJECT

Dmitry Naumov for the JINR group

PAC JINR

June 29, 2020



1 TOPIC

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THE PROJECT UPDATE: JUNO AND DAYA BAY



Topic: 2010-2023

Daya Bay will stop by the end of 2020

- 02-2-1099-2010/2023 Study of Neutrino Oscillations

Past: 2007–2017

- The Daya Bay Project

Present: 2018–2020

- The JUNO/Daya Bay Project

Future: 2021–2023

- The JUNO Project
- The Daya Bay activity

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- The JUNO Project
- The Daya Bay activity

Our team

N. Anfimov, T. Antoshkina, S. Biktemerova, A. Bolshakova, I. Butorov, A. Chetverikov, A. Chukanov, S. Dmitrievsky, D. Dolzhikov, D. Fedoseev, M. Gonchar, Y. Gornushkin, M. Gromov, V. Gromov, D. Korablev, A. Krasnoperov, N. Kutovskiy, K. Kuznetsova, Y. Malyshkin, [D. Naumov](#), E. Naumova, I. Nemchenok, [A. Olshevskiy](#), A. Rybnikov, A. Sadovsky, D. Selivanov, A. Selyunin, V. Sharov, A. Shaydurova, V. Shutov, O. Smirnov, S. Sokolov, A. Sotnikov, M. Strizh, V. Tchalyshev, K. Treskov, N. Tsegelnik, V. Zavadskiy, O. Zaykina

39 from JINR

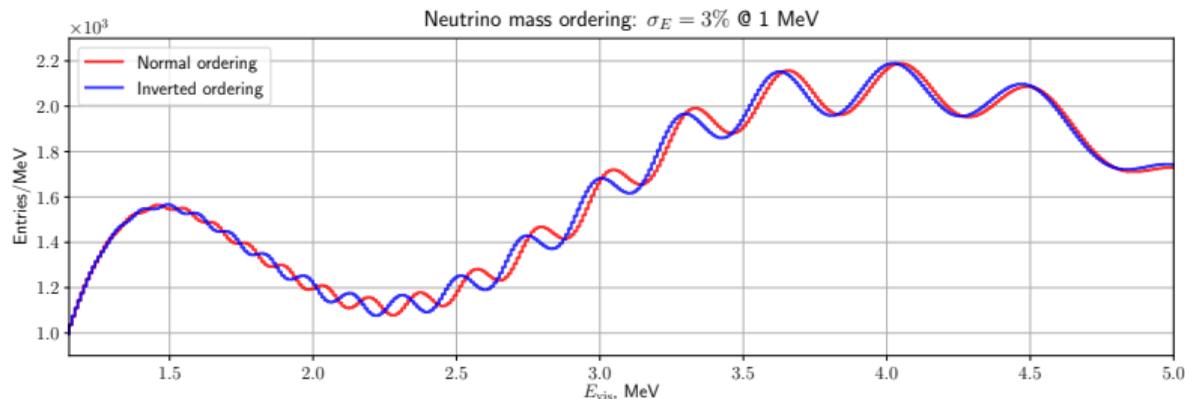


SUMMARY

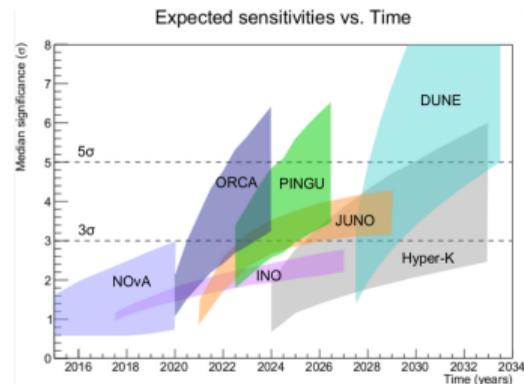
Sensitivity plot: KM3NeT@NOW2016 (outdated)

CDR: 1508.07166

Physics: 1507.05613



- Mass hierarchy determination: $3 - 5 \sigma$
- Measurement of ν mixing parameters: unc. $< 1\%$
- Probe PMNS matrix unitarity: $\sim 1\%$ level
- Precision reactor $\bar{\nu}$ spectrum: $\sigma_E \sim 2\%$ at 1 MeV



- Supernovae ν and DSNB
- Solar and geo- ν
- Atmospheric ν
- Sterile ν and exotic searches
- Proton decay



JUNO SCHEDULE

Complete conceptual design.
International collaboration established.

Bidding of detector components.

PMT mass production and testing.

End of civil construction.
Electronics mass production.

Start of data taking

2014

2015

2016

2017

2018

2019

2020

2021

2022

Start civil construction, PMT production line.

Start PMT mass production.
First electronics prototypes.

Start PMT potting.

PMT installation.
Detector and veto construction.

We are here





JINR GROUP ACTIVITIES

JUNO and TAO: hardware

- PMT:
 - ▶ characterization
 - ▶ mass testing
 - ▶ HV for JUNO
- SiPM for TAO
 - ▶ purchase
 - ▶ acceptance studies
 - ▶ mass characterization
 - ▶ HV for TAO
- Top Tracker Support
- EMF shielding
- Computing

JUNO and TAO: software

- Sensitivity
- Vertex, energy, track reco.
- Machine learning
- Top Tracker DAQ

Analysis frameworks

- GNA

Daya Bay (recent)

- Oscillation analysis
- Wave packets
- IBD selection
- Sterile neutrino
- Neutrino directionality



JINR GROUP PLANS: 2021–2023

- HV for JUNO LPMT and sPMT: assembly, tests, installation. Slow control
- Top Tracker: assembly, installation, commissioning. DAQ software. Slow control
- SiPM for TAO: purchase, mass test, hv units
- HV for TAO SiPM: R&D, production, software
- EMF protection for WP LPMTs: R&D, production
- Reconstruction of primary vertex, energy, muon track: include calibration, production
- Analysis: data analyses, combination with other experiments
- GNA: GPU support, automatic differentiation, development and support, data analyses

PUBLICATIONS 2018–2020 I



1. D.V.Naumov, V.A.Naumov, “Quantum Field Theory of Neutrino Oscillations”,
10.1134/S1063779620010050. Phys.Part.Nucl. 51 (2020) no.1, 1-106.
2. O.Smirnov, “Experimental aspects of geoneutrino detection: Status and perspectives”,
Progress in Particle and Nuclear Physics 109 (2019) 103712.
3. N.Anfimov, A.Rybnikov, and A.Sotnikov. “Optimization of the light intensity for photodetector calibration”.
Nucl. Instrum. Meth., A939:61 – 65, 2019.
4. A.Fatkina et al., “GNA: new framework for statistical data analysis”.
[arXiv:1903.05567 [cs.MS]]. EPJ Web Conf. 2014 (2019)
5. A.Fatkina et al., “CUDA Support in GNA Data Analysis Framework”.
[arXiv:1804.07682 [cs.DC]]. Computational Science and Its Applications Proc IV (2018).
6. N.Anfimov (JUNO). “Large photocathode 20-inch PMT testing methods for the JUNO experiment”
JINST 12 (2017).



PUBLICATIONS 2018–2020 II

1. Daya Bay Collaboration, “Extraction of the ^{235}U and ^{239}Pu Antineutrino Spectra at Daya Bay”,
arXiv:1904.07812 [hep-ex]. Phys.Rev.Lett. 123 (2019) no.11, 111801.
2. Daya Bay Collaboration, “A high precision calibration of the nonlinear energy response at Daya Bay”,
arXiv:1902.08241 [physics.ins-det]. Nucl.Instrum.Meth. A940 (2019) 230-242.
3. Daya Bay Collaboration, “Measurement of the Electron Antineutrino Oscillation with 1958 Days of Operation at Daya Bay”,
arXiv:1809.02261 [hep-ex]. Phys.Rev.Lett. 121 (2018) no.24, 241805.
4. Daya Bay Collaboration, “Improved Measurement of the Reactor Antineutrino Flux at Daya Bay”,
arXiv:1808.10836 [hep-ex]. Phys.Rev. D100 (2019) no.5, 052004.

INSPIRE HEP Summary	Citeable	Published
Papers	16	11
Citations	125	121
Citations/paper (avg)	7.8	11



2017	candidate	M. Gonchar	The measurement of neutrino mixing angle θ_{13} and neutrino mass splitting Δm_{32}^2 in the Daya Bay experiment.
2017	doctor	D. Naumov	Measurement θ_{13} and Δm_{32}^2 of and quantum-field theory of neutrino oscillations.
2019	doctor	I. Nemchenok	Development and research of plastic and liquid scintillators for detectors of experiments in the field of neutrino physics.
Soon	candidate	N. Anfimov	Methods for the research of photodetectors and their application.
Soon	doctor	O. Smirnov	Study of the geo- and pp-chain solar neutrino fluxes with the Borexino detector.
...	candidate	K. Treskov	... oscillation analysis in Daya Bay

Students: 2018–2020

✓ Bachelors: 4 students

✓ Masters: 2 students

✓ Summer Student Program: 6 students



STAFF I: AVERAGE FOR 3 YEARS

1	N. Anfimov	staff	0.4	PMT testing group leader
2	T. Antoshkina	PhD student	1	PMT optics response simulation. Formulation of requirements for PMT testing quality
3	S. Biktemerova	PhD student	1	sensitivity estimation, detector simulation, data analysis
4	A. Bolshakova	staff	1	PMT data analysis
5	I. Butorov	engineer	0.5	Designing and technical work PMT testing, analysis
6	A. Chetverikov	engineer	0.3	Designing and technical work PMT testing
7	A. Chukanov	candidate	0.7	Reconstruction, data analysis
8	S. Dmitrievsky	candidate	0.5	Simulation and software development for TT
9	D. Dolzhikov	student	0.33	Selection and analysis
10	D. Fedoseev	engineer	0.3	Designing and technical work
11	M. Gonchar	PhD student	1	sensitivity estimation, detector simulation, data analysis
12	Yu. Gornushkin	candidate	0.5	The TT project coordination
13	M. Gromov	candidate	0.5	Analysis, SuperNOVA
14	V. Gromov	engineer	0.5	Software development for TT/TAO JUNO
15	D. Korablev	staff	0.57	software development for PMT testing and TT, Long term-stability, PMT testing, Analysis



STAFF II: AVERAGE FOR 3 YEARS

15	D. Korablev	staff	0.57	software development for PMT testing and TT, Long term-stability, PMT testing, Analysis
16	A. Krasnoperov	candidate	0.3	Software development for TT JUNO
17	N. Kutovskiy	candidate (LIT)	0.2	IT
18	K. Kuznetsova	engineer	0.3	SiPM testing, Analysis
19	Y. Malyshkin	candidate	0.5	Analysis, reconstruction
20	D. Naumov	doctor	0.6	project management. Reactor spectrum measurement. Oscillation analyses. Global analysis
21	E. Naumova	staff	1	Reactor spectrum measurement.
22	I. Nemchenok	candidate	0.5	Investigation of properties and stability of liquid scintillator
23	A. Olshevskiy	doctor	0.5	analysis preparation, HV and other JINR hardware activities coordination
24	A. Rybnikov	engineer	0.3	PMT testing, SiPM testing
25	A. Sadovsky	candidate	0.5	PMT HV R&D
26	D. Selivanov	student	0.13	Reconstruction
27	A. Selyunin	engineer	0.2	PMT testing



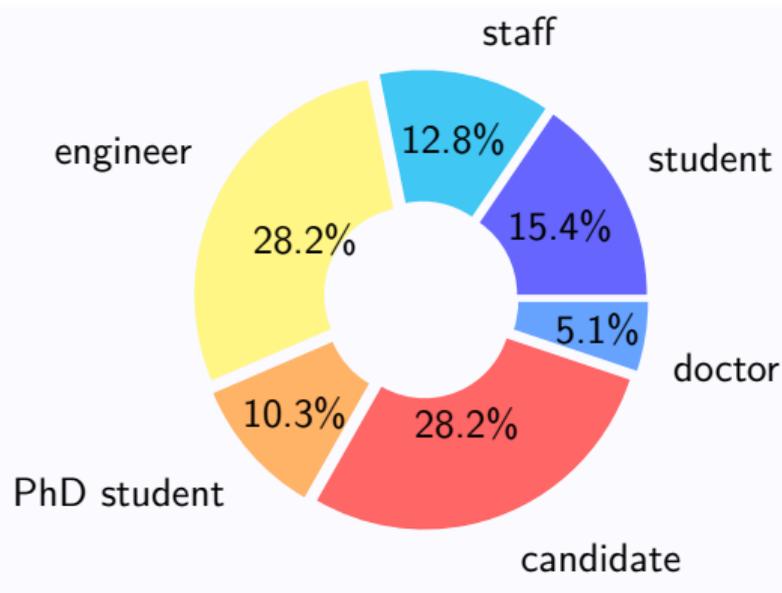
STAFF III: AVERAGE FOR 3 YEARS

27	A. Selyunin	engineer	0.2	PMT testing
28	V. Sharov	engineer	0.5	PMT testing
29	A. Shaydurova	student	0.13	Neutrino oscillations in matter
30	V. Shutov	engineer	0.3	HV unit software and tests
31	O. Smirnov	candidate	0.6	PMTs magnetic shielding, energy resolution studies
32	S. Sokolov	engineer	0.5	Designing and technical work, PMT testing
33	A. Sotnikov	engineer	0.3	PMTs magnetic shielding, PMT tests
34	M. Strizh	student	0.13	reconstruction of neutrino directionality
35	V. Tchalyshev	candidate	0.5	SiPM testing, clean room support
36	K. Treskov	PhD student	1	sensitivity estimation, software development, data analysis
37	N. Tsegelnik	student (LTF)	0.5	fitting software
38	V. Zavadskiyi	student	0.27	Oscillation Analysis
39	O. Zaykina	staff	1	Reconstruction
Total FTE			19.87	
Avg. people			38	Mostly DLNP
FTE/person			0.52	

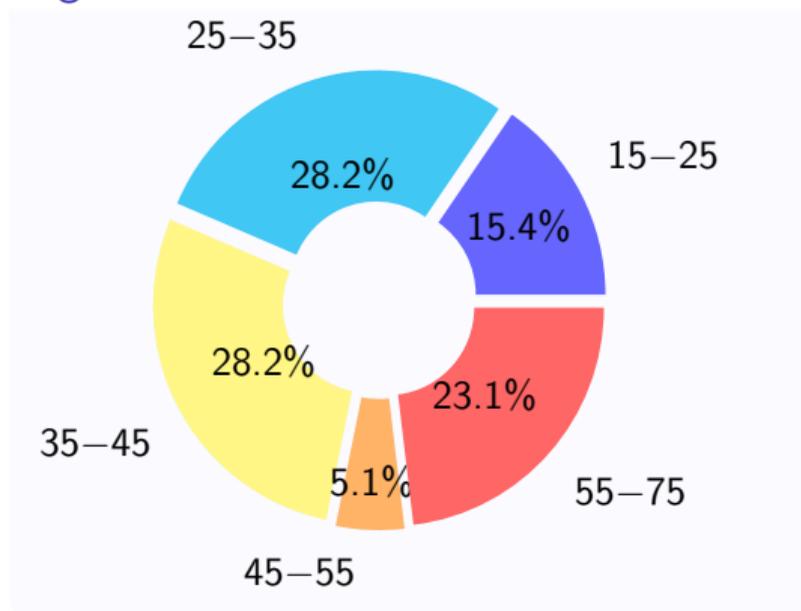


STAFF SUMMARY: 2021–2023

Status distribution



Age distribution



Total FTE 19.87
 Avg. people 38
 FTE/person 0.52

JUNO EXPENDITURES 2018–2020



	HV Units	TOP Tracker	PMT Scanning	JUNO Computing	TAO	Total
JINR contribution, USD	2M	(in-kind) 0.8M + 0.2M	2 stations + personnel on site	2M	1.5M	6.5M
Payment status	100%	80%	Secured	50%	66.6%	74%
In progress, USD		0.2M		1M	0.5M	1.7M

FORM 26



Item		Total	2021	2022	2023
Equipment, k\$	1. JUNO maintenance	1000	—	500	500
	2. SiPM для TAO purchase and tests	500	300	200	—
	3. Top Tracker Support production, assembly, installation	200	200	—	—
	4. HV assembly, installation, tests	200	200	—	—
	5. Computing Infrastructure servers, disks, tape	500	300	100	100
Total		2400	1000	800	600
Other	launch, shifts, analysis, meetings, conferences	k\$ 800	300	250	250
Source	Budget	k\$ 3200	1300	1050	850
	Extra	funds	30	10	10



FORM 29

№	Item, k\$	Total	2021	2022	2023
1.	Accelerator	—	—	—	—
2.	Computer	—	—	—	—
3.	Design bureau (DLNP)	2100	700	700	700
4.	Experimental Workshop (JINR)	2400	800	800	800
	<i>working hours</i>				
5.	Computers, connection	30	10	10	10
6.	Materials	900	500	300	100
7.	Equipment	1440	480	480	480
8.	R&D	30	10	10	10
9.	Missions	800	300	250	250
	Total	3200	1300	1050	850



SWOT ANALYSIS

	HELPFUL	HARMFUL
INTERNAL	<p>Strengths</p> <ul style="list-style-type: none"> ✓ Neutrino hierarchy determination ✓ Method different from other experiments ✓ Precision measurement of 3 oscillation parameters ✓ Precision measurement of 3 PMNS elements (+DB) ✓ Geo-neutrinos measurement ✓ Solar, atmospheric ν, proton decay, etc. 	<p>Weaknesses</p> <ul style="list-style-type: none"> ✗ Failure to achieve 3% energy resolution ✗ Insufficient detector/structure integrity ✗ Insufficient electronics/HV reliability ✗ Delay with detector installation
	<p>Opportunities</p> <ul style="list-style-type: none"> ▪ Supernova burst ▪ Diffuse Supernova background ν ▪ New physics 	<p>Threats</p> <ul style="list-style-type: none"> ✗ Underground collapse and flooding
EXTERNAL		

Thank you for your attention!

Spare slides:

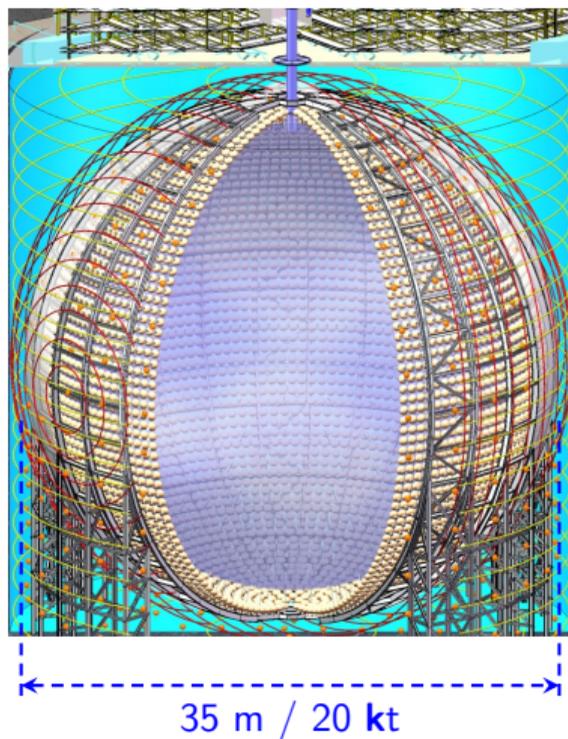
5 FINANCES

6 PHYSICS AT JUNO

7 JINR ACTIVITIES

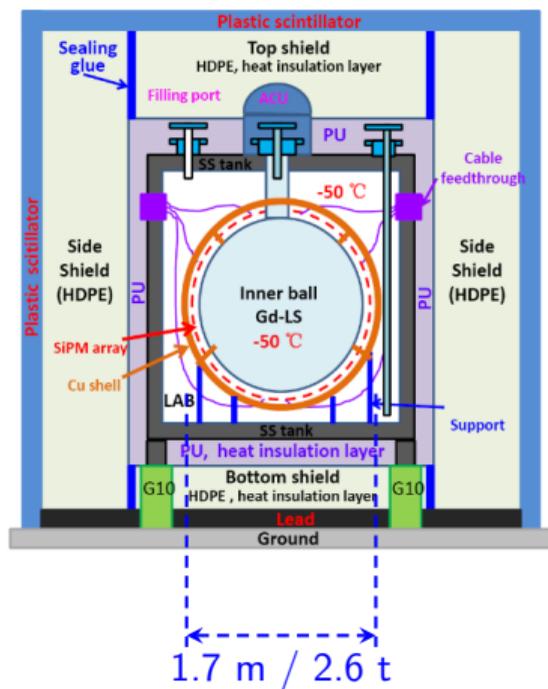
- Top Tracker
- PMT High Voltage
- PMT scanning
- EMF protection
- TAO detector
- Computing
- Reconstruction
- GNA Project

ANTINEUTRINO DETECTORS (AD)



	Daya Bay	JUNO
Attention	Uncorr. ε unc.	Energy resolution
Method	Identical ADs 3 zones	Light collection
Scintillator	GdLS/LS	LS
PMTs	192 8"	18k 20" +26k 3"
Coverage, %	12	78
Light col. p.e./MeV	160	1200 1350
σ_E at 1 MeV, %	8.7	3
Detectors	4/4 ^{far} _{near}	1
Thermal power, GW	17.4	35.8 26.6
Baseline	0.5 km–2 km	52 km
IBD/day/AD	75/635 ^{far} _{near}	60 45

ANTINEUTRINO DETECTORS (AD)



	Daya Bay	TAO	JUNO
Attention	Uncorr. ϵ unc.	Energy resolution	
Method	Identical ADs	Light collection	
	3 zones	Dark noise	
Scintillator	GdLS/LS	GdLS	LS
		@ -50 °C	
PMTs	192 8"	SiPM	18k 20"
		1.5M 5 mm	+26k 3"
Coverage, %	12	94	78
Light col. p.e./MeV	160	4500	1200 1350
σ_E at 1 MeV, %	8.7	2	3
Detectors	4/4 far near	1	1
Thermal power, GW	17.4	4.6	35.8 26.6
Baseline	0.5 km–2 km	30 m	52 km
IBD/day/AD	75/635 far near	2000	60 45



NEUTRINO PHYSICS AT JUNO I

- Neutrino mass ordering (NMO)
 - ▶ 3σ NMO sensitivity within ≈ 8 years.
 - ▶ 4σ with Δm_{32}^2 input from accelerator experiments.
 - ▶ $> 5\sigma$ combined analysis with IceCube within 3–7 years
or PINGU in 2 years.
 - ▶ Combination with accelerator experiments — promising.



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Physics with TAO

- Precision reactor $\bar{\nu}_e$ spectra:
 - ▶ Total spectrum.
 - ▶ $^{235}\text{U}/^{239}\text{Pu}$ spectra.
- Search for sterile neutrino.



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- Neutrino oscillation parameters measurement
 - ▶ ~ 20 oscillation cycles in a single experiment.
 - ▶ Expected precision for $\Delta m_{32/21}^2$ and $\sin^2 2\theta_{12} < 0.7\%$.
 - ▶ $\sin^2 2\theta_{13}$ precision 15%. (Daya Bay: $< 3\%$).
 - ▶ Test U_{PMNS} unitarity on $< 1\%$ level
↔ similar to quark sector.

Physics with TAO

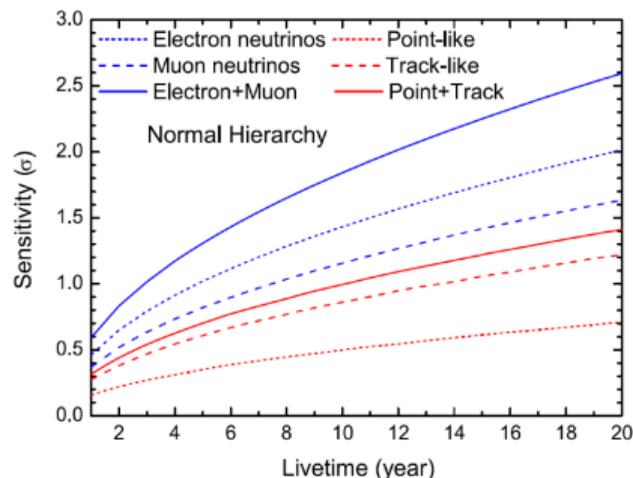
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 - ▶ Test U_{PMNS} unitarity on $< 1\%$ level
 - ↔ similar to quark sector.
- Atmospheric neutrinos ▶
 - ▶ Measure θ_{23} with 6° precision.
 - ▶ Complimentary NMO sensitivity.

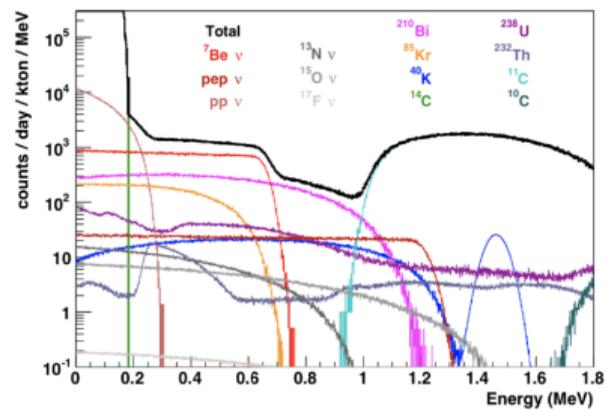
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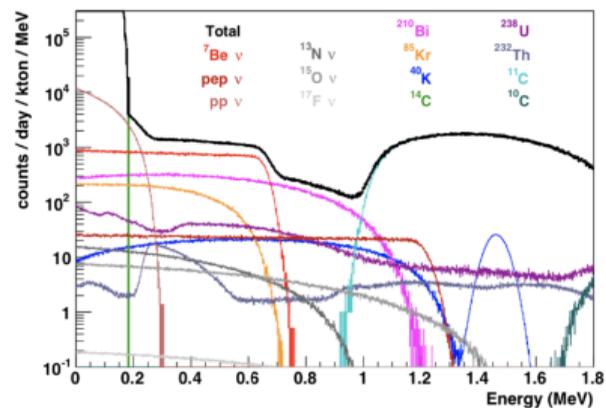
NEUTRINO PHYSICS AT JUNO II

- Solar neutrino ▶
 - ▶ 1000 ${}^7\text{Be}$ and 10 ${}^8\text{B}$ neutrino interactions per day.



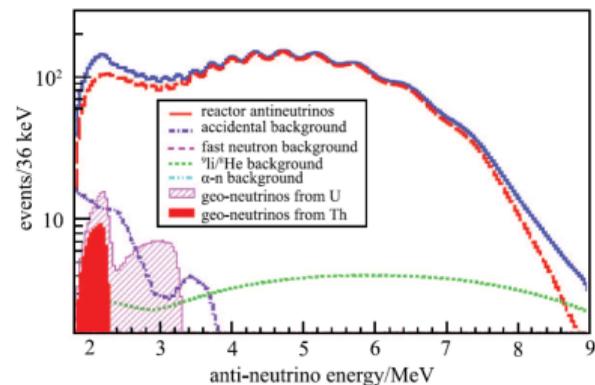
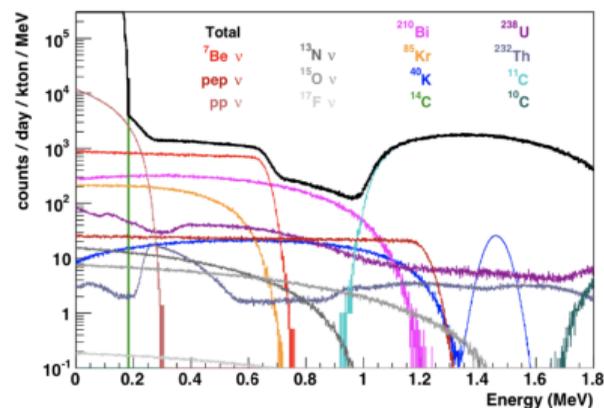
NEUTRINO PHYSICS AT JUNO II

- Solar neutrino ▶
 - ▶ 1000 ^7Be and 10^8 ^8B neutrino interactions per day.
- SuperNOVA
 - ▶ Sensitivity: flavor content, energy spectrum, time evolution.
 - ▶ 10k events (5k via IBD) for SN @ 10kpc.
- Diffuse SuperNOVA background (DSNB)
 - ▶ 3σ sensitivity in 10 years or strongest constraint.



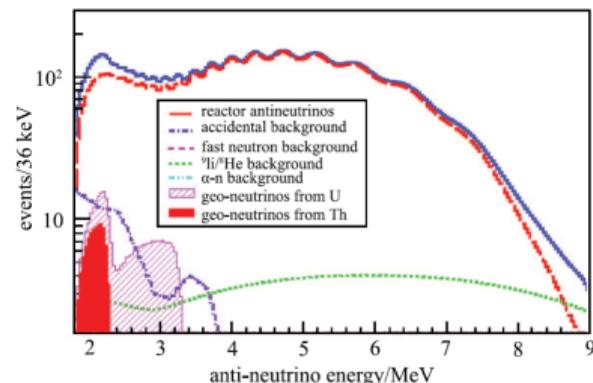
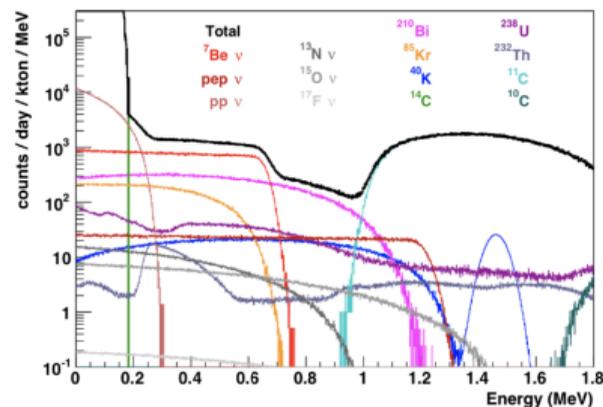
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 - ▶ Largest statistics. Precision 5% in 10 years.



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- Proton decay
 - ▶ Competitive sensitivity via $p \rightarrow \bar{\nu} + K^+$.
 - ▶ Triple coincidence signal.

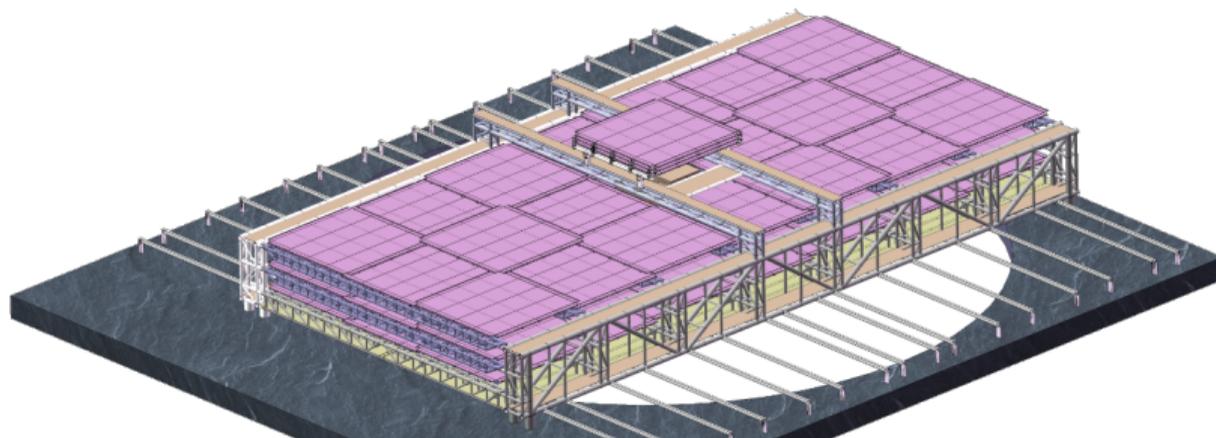




VETO: TOP MUON TRACKER (TT)

Motivation

- Precision muon tracking: $0.2^\circ/0.5^\circ$
- Layered plastic scintillator detector
- Partial coverage: $\sim 63\%$
- 3 layers x 21 “walls” x 8 modules
- Wall: $7 \times 7 \text{ m}^2$, 1 t / Layer: $\sim 1000 \text{ m}^2$





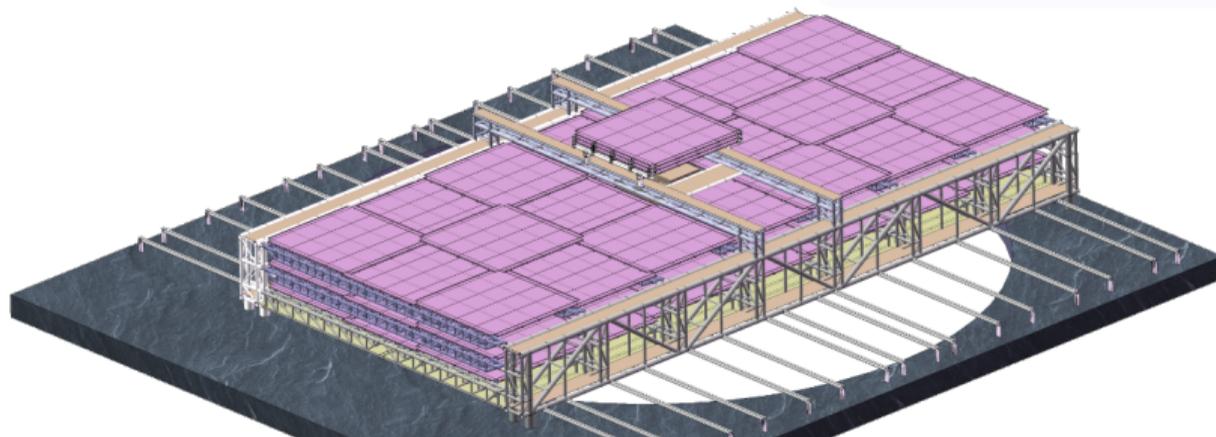
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Status: JINR

- Mechanical support structure R&D, prototyping and validation: **done**
- Assembly procedure, tools: **done, reviewed**
- Bidding: **done**
- Manufacturing (140 t): **2021**
- Assembly on site: **2021.09 \rightarrow 2022.03**

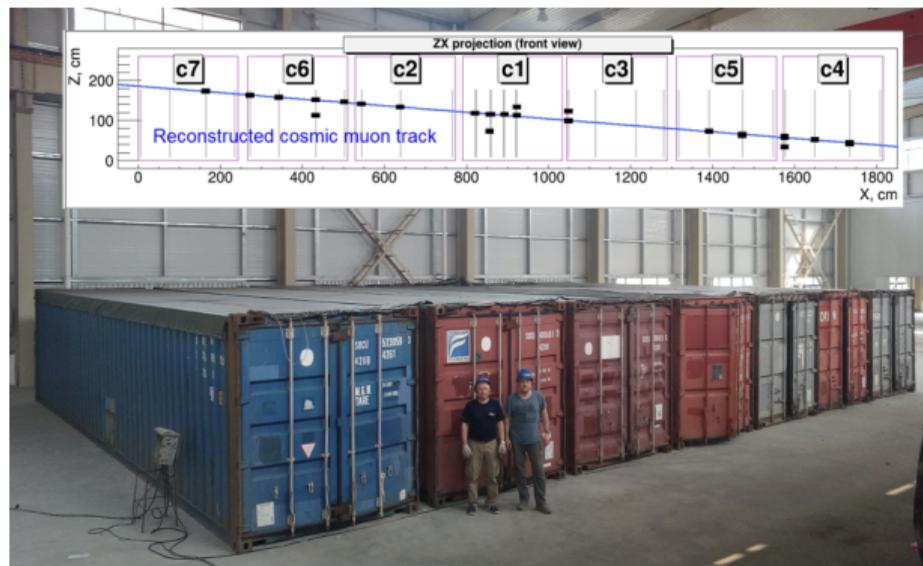
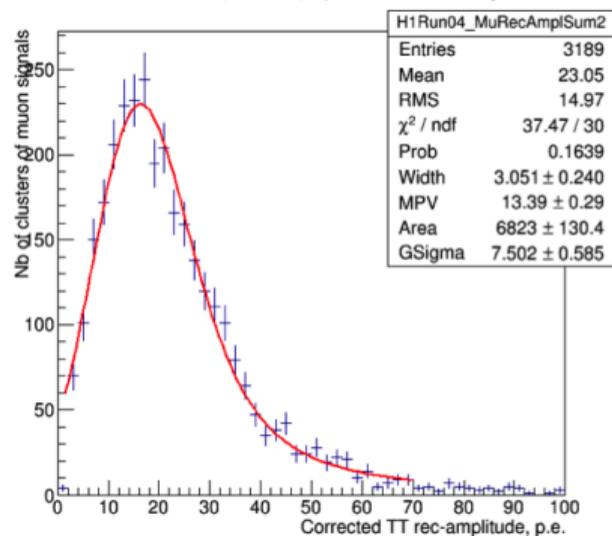


TT: PLASTIC SCINTILLATOR MONITORING

Plastic scintillator for TT

- Re-used OPERA Target Tracker
- ✓ Delivered on-site and stored in 7 containers
- ✓ Active DAQ to study aging

Sum of reconstructed (and corrected) amplitudes for clusters of muon signals





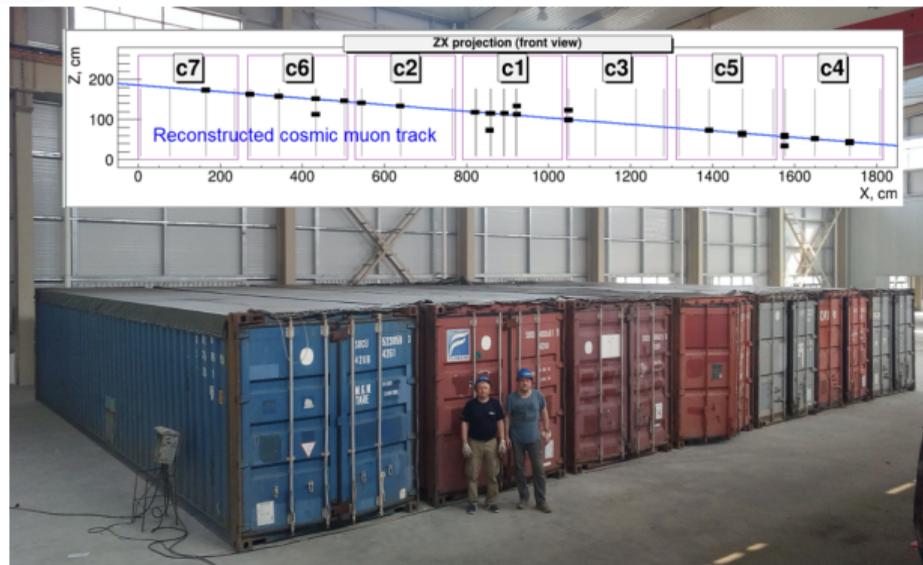
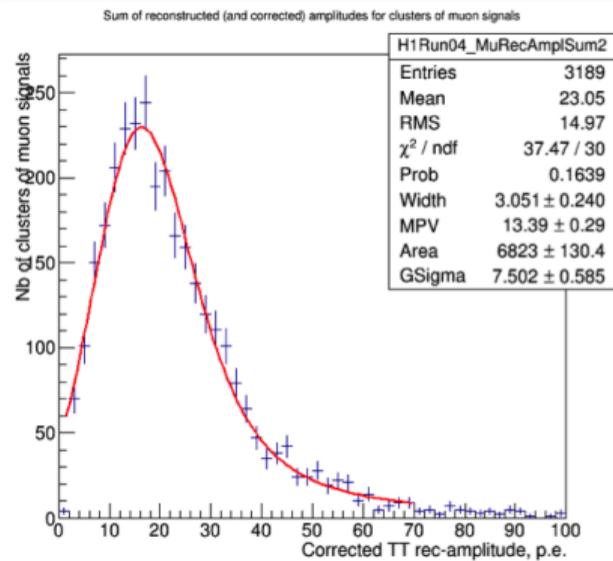
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JINR

- ✓ Mobile DAQ and software (storage): operating
- DAQ software for TT: in progress



HV UNITS

High Voltage supply provided by JINR

- ~18'000 large PMTs central detector
- ~ 2'000 large PMTs veto
- ~25'600 small PMTs central detector
- ~**25'000** underwater HV units **required**
1 unit per 8 sPMTs



HV UNITS

High Voltage supply provided by JINR

- ~18'000 large PMTs central detector
- ~ 2'000 large PMTs veto
- ~25'600 small PMTs central detector
- ~**25'000** underwater HV units **required**
1 unit per 8 sPMTs



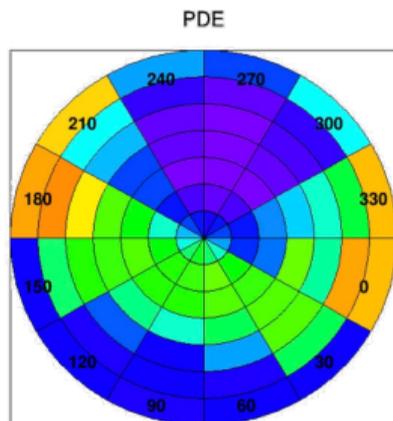
Status

- ✓ R&D, prototyping
- ✓ Testing: materials, ageing, thermo cycling
- ✓ Factory setup, procedures: [Shenzhen](#)
- ✓ Test batch: 500 items
- Production via single batch: → [2020](#)

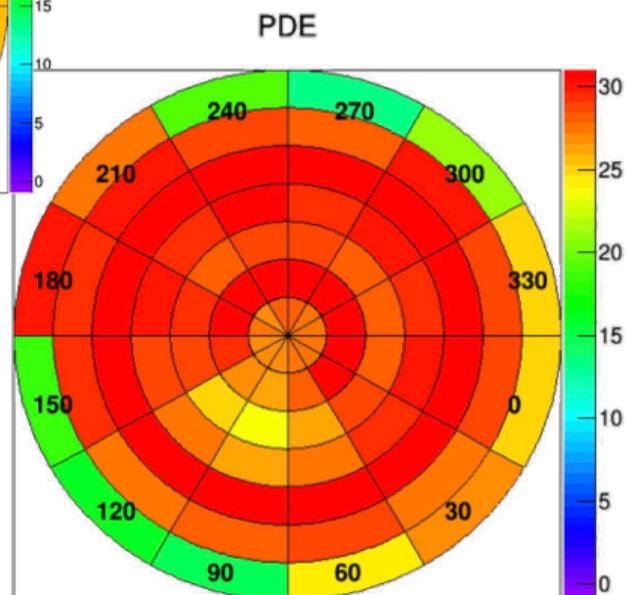


PMT SCANNING

- ✓ 3 Scanning stations produced @JINR:
 - ▶ 1 in DLNP / 2 in China
 - ▶ All in individual dark rooms
 - ▶ Dedicated software
- ✓ Scanning: 3-4 PMTs/day
- ✓ ~ 2500 PMTs scanned
- ✓ Maintain database, web accessible
- ✓ Study Earth Magnetic Field impact ▶
 - Complementary to mass testing

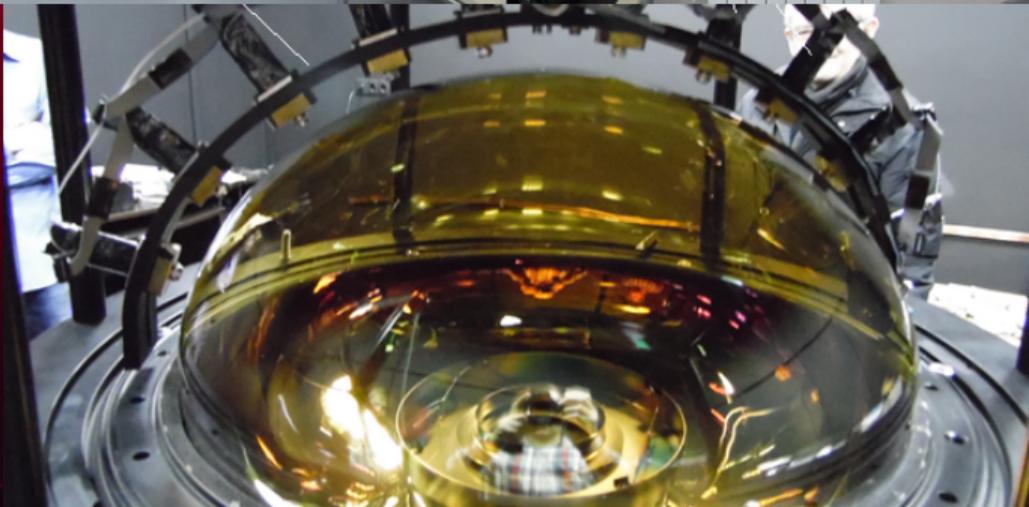


◀ Magnetic field: $42 \mu\text{T}$



Magnetic field: compensated ▶

JINR PMT TESTING LABORATORY

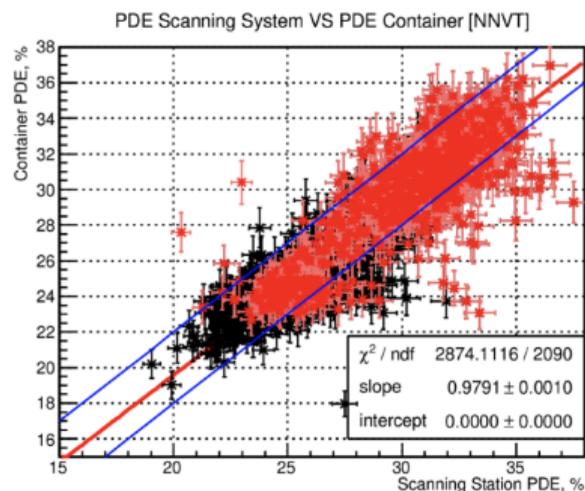




MASS PMT TESTING

Scanning and mass testing

- ✓ Almost 17k PMTs tested
- ✗ ~3% rejected
- ✓ 3'110 PMTs tested after potting
- ✓ Complementary and consistent performance



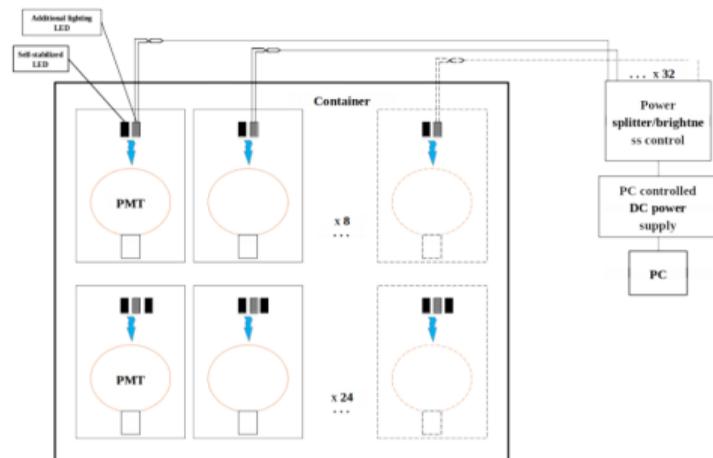
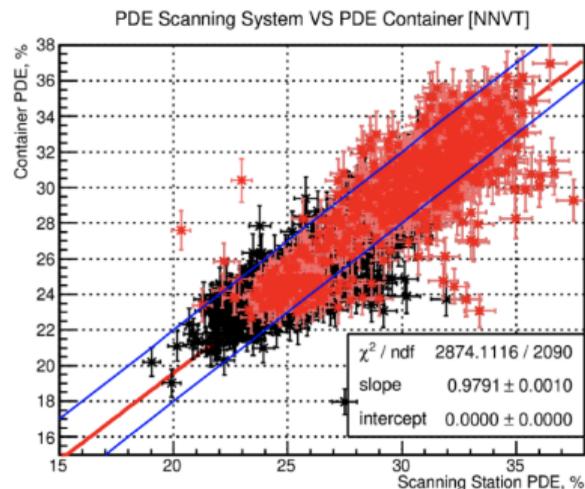
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Long term stability

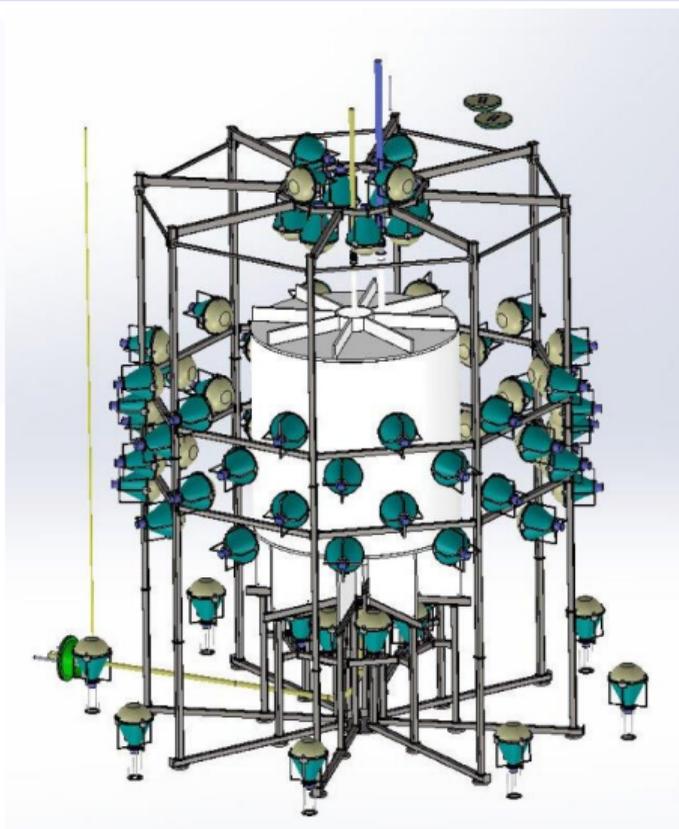
- ✓ 1 container equipped by JINR
- ✓ Operating since February 2020
 - 32 PMTs for 1 year
- ✓ DAQ software by JINR group



OSIRIS: PMT PROTECTION VS. EARTH MAGNETIC FIELD



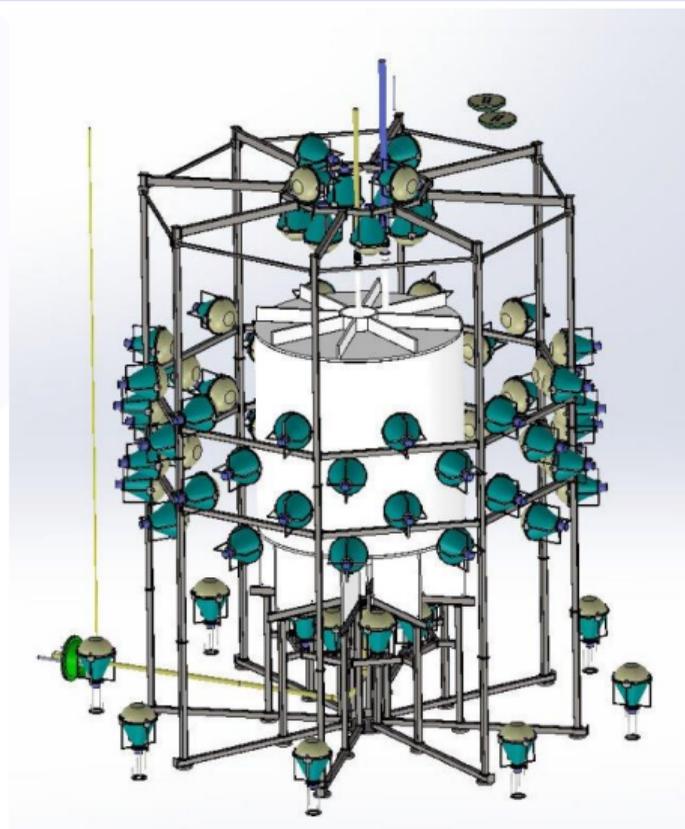
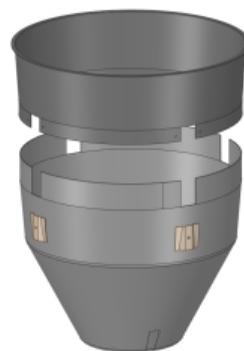
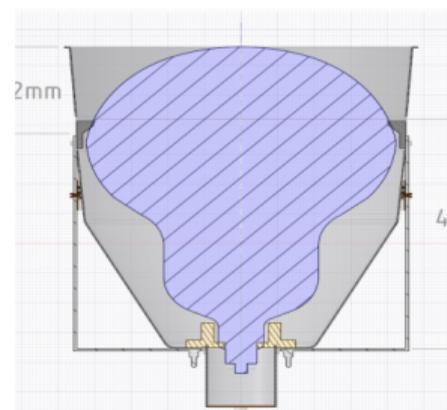
- Online Scintillator Internal Radioactivity Investigation System
- 76 20" PMTs: 64 detector + 12 veto
- Individual EMF protection: Metglas+Al cones:
 - ▶ detector: carbon fiber composite
 - ▶ veto: fiberglass composite
- EMF reduction factor: $\times 2$ \parallel and $\times 10$ \perp



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PROTOTYPES AND RADIOACTIVITY BALANCE

- 3 prototypes produced
- White coating for additional light collection ▶
- Table: carbon fiber option
- Fiberglass: more radioactive, but acceptable



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Material	U, ppb	Fraction		U	Mass, μg	
		Th, ppb	K, ppm		Th	^{40}K
AMAG-170 (0.2 mm)	3	<5	0.84	4.7	<7.8	0.154
Epoxy	<0.1	0.9	0.78	<0.12	0.11	0.1
Carbon fiber	1	<6	15	0.25	<4.9	1.4
Gelcoat white	7	7	4.33	2.5	2.5	0.2
Cu foil	<0.3	<0.2	<0.127	<0.024	0.016	<0.001
Al foil	170	26	<0.96	15.3	2.4	<0.01
Total				22.8	<18	1.9
PMT glass	400	400	60	3600	3600	63

TAO — TAISHAN ANTINEUTRINO OBSERVATORY



Objective

- Precision antineutrino spectrum measurement
- High statistics, no oscillations
- Unprecedented precision: $\sigma_E \sim 2\%$ at 1 MeV

TAO — TAISHAN ANTINEUTRINO OBSERVATORY

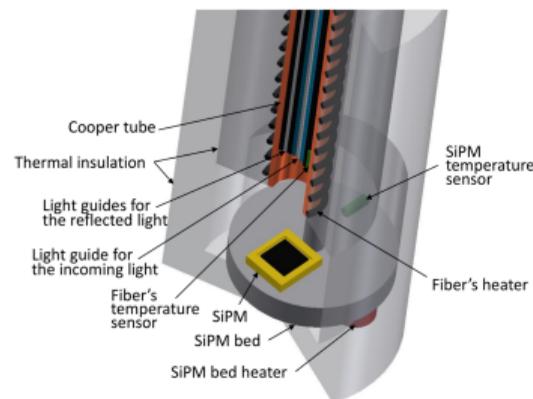


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- ✓ TAO CDR preparation
 - SiPM mass characterization



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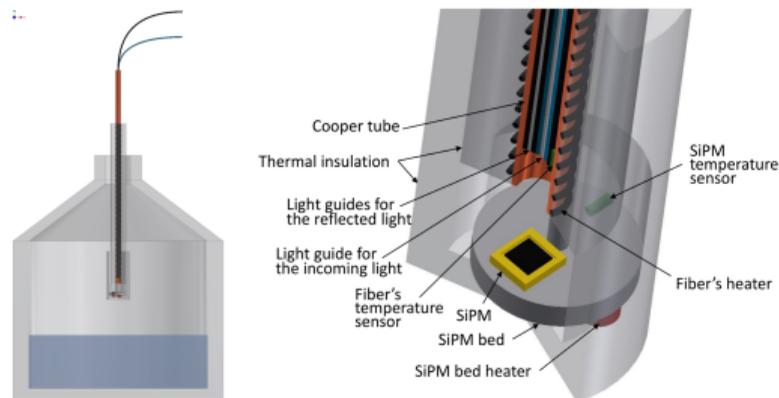


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- ✓ SiPM acceptance studies at -50°C
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- SiPM mass characterization
- SiPM high voltage supply
- 2/3 SiPM purchase funding: 1.5M\$



a) - Custom made HV unit by JINR



b) - DAC81416EVM by TI



COMPUTING: LIT&DLNP

Total requirements before 2040 (JINR)

- 4000 cores / 4 PB disk / 40 PB tape + 4 PB disk cache
- To store complete copy of JUNO data





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CPU and storage

- ✓ Current: 300 cores / 25 TB (dCache) + 500 TB (EOS)
- New: HP servers with 2880 cores with increased RAM 16 GB ▶
 - ✓ purchased and delivered in 2019, installed recently
 - ▶ to be powered
 - ▶ Part of Neutrino Computing Platform: [shared](#), quota





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Network and GRID

- ✓ GRID: CVMFS repository / Secondary VOMS server: [deployed](#)
- ✓ Network [Gbps]: 2x100 (local) / 3x100 (wide) / 2x10 (↔ China)

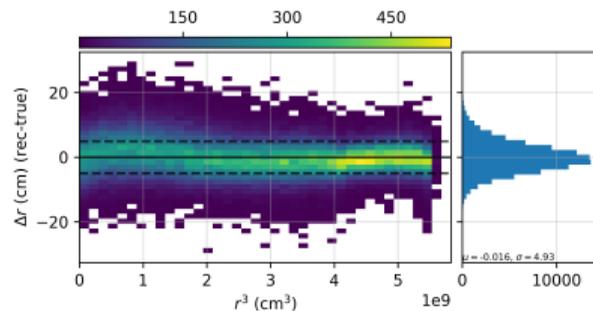




JUNO RECONSTRUCTION AT JINR: PROGRESS

Neural Networks: vertex/energy reco ▶

- ✓ Exceptional speed, industry support
- ✓ Consistent with traditional methods
- ✗ Long training
 - TODO: develop calibration methods
 - Sphere projection? or Graph Neural Networks
- ✓ Many thanks to HybriLIT for GPU computing support!



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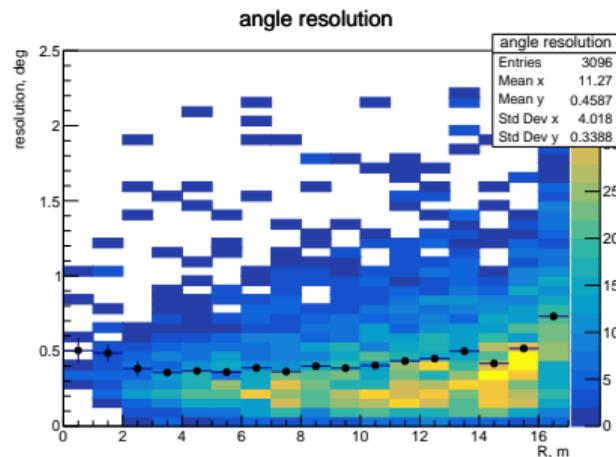
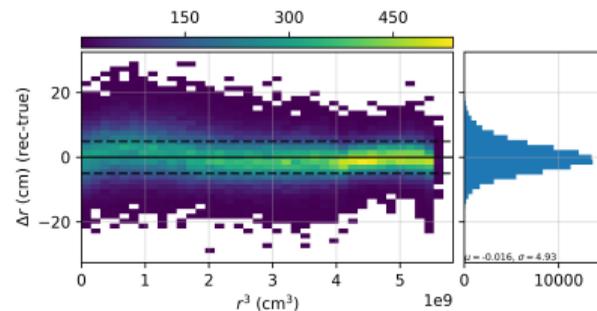


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Muon track reconstruction ▶

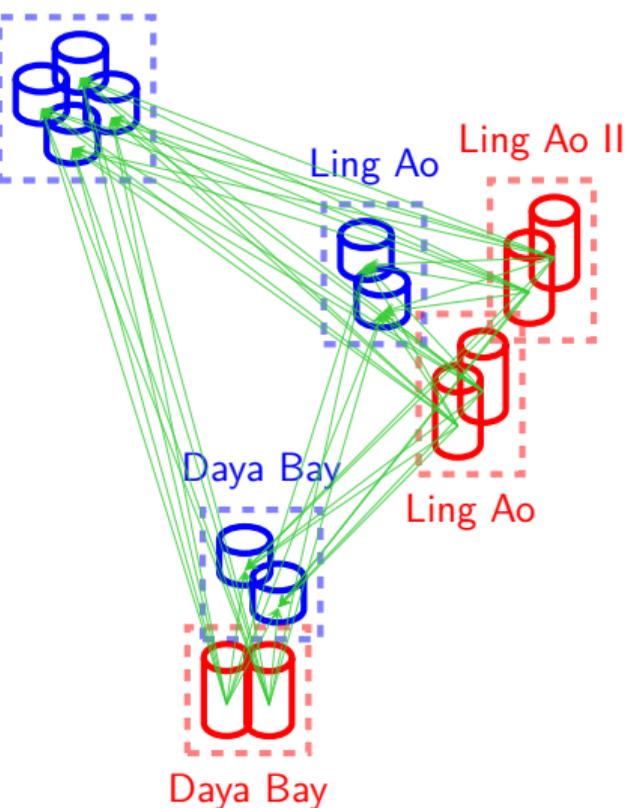
- For background rejection: $^8\text{He}/^9\text{Li}$, fast- n
- ✓ Decomposition into spherical functions
- ✓ Single tracks, muon bundles
- ✓ Precision: 0.5° and 10 cm



WHY GNA: DAY BAY EXPERIMENT

Far site

Count: 48



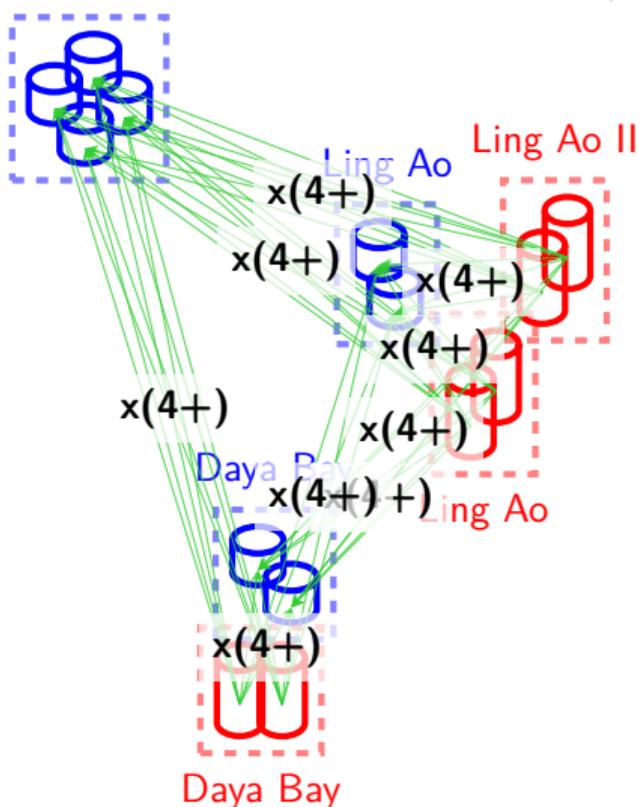
$$\vec{N}^d = \vec{B}^d + \left(\prod_m C_m \right) \times$$

$$\times \sum_t \dots \sum_r \dots \int \int \dots \sum_i \dots \sum_c P_c(L_r^d, \dots) \dots$$

WHY GNA: DAY BAY EXPERIMENT

Far site

Count: 192+



Requirements

- Efficient scalable models, lots of parameters
 - ▶ Laziness, caching
 - ▶ Flexibility
 - ▶ Portability (GPU)
- Long term support and compatibility
 - ▶ Readability, hand-over-ability
 - ▶ Minimize boilerplate code

$$\vec{N}^d = \vec{B}^d + \left(\prod_m C_m \right) \times$$

$$\times \sum_t \dots \sum_r \dots \int \int \dots \sum_i \dots \sum_c P_c(L_r^d, \dots) \dots$$

GNA: JUNO/DAYA BAY IMPLEMENTATION



- GNA framework — scalable high performance fitting.

```

eres[d] |
lsnl[d] |
iav[d] |
  integral2d |
    sum[r] |
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      ibd_xsec(enu(), ctheta())*
jacobian(enu(), ee(), ctheta())
  sum[i] (
power_livetime_factor[d,r,i])*
  anuspec[i](enu() )*
  sum[c] |
pmns[c]*oscprob[c,d,r](enu())

```

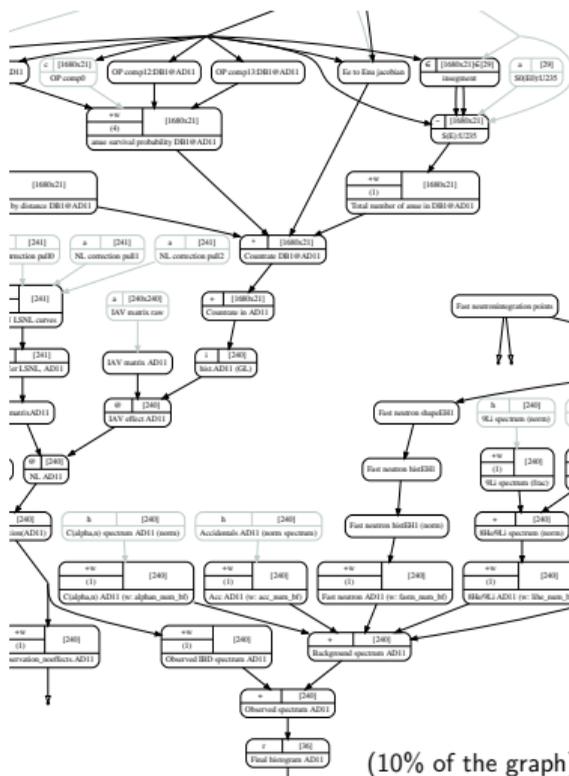


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<http://gna.pages.jinr.ru/gna/gallery.html>

Dmitry Naumov (DLNP)

Daya Bay & JUNO

June 29, 2020

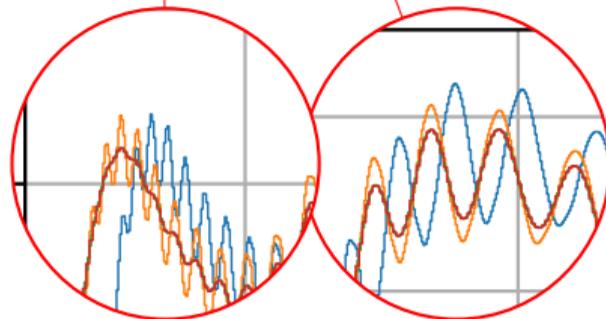
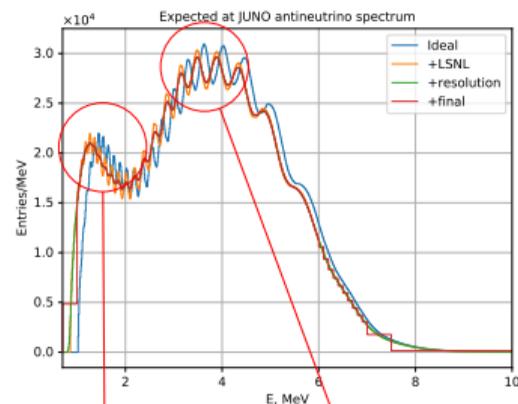
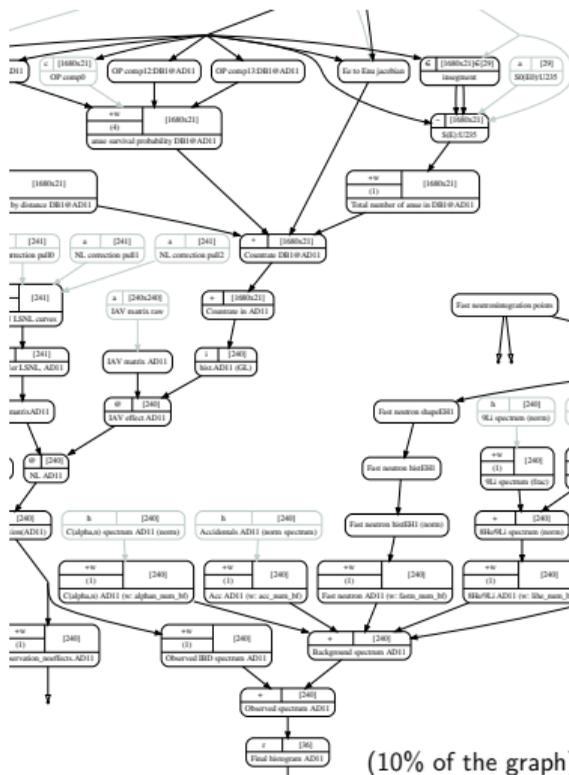
32b / 15

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Daya Bay & JUNO

(10% of the graph)

June 29, 2020

32c / 15

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- First version: 2014→2015
- ✓ Current version: 2017→

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Plans

- GPU support
- TensorFlow support?
- Documentation, tutorial, etc.