

# Calculation of yield in OMC on $^{24}\text{Mg}$

**Yu. Shitov**

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# Data processing and hit types

- $\mu$ -stops are analyzing separately in C0 and C1 counters **inside CW=14  $\mu$ s coincidence window**. For each ge-hit we determine its type:

// mu-ge hits coincidences determined with

// C1 entrance counter while C0 (ring) counter is using as veto

// Hit types:

// 1 - **good**: events with single muon in C1 in CW

// 2 - **multiple**: multiple muons in C1 in CW

// 3 - **flagged**: single mu in C1, but non-zero flag(s)

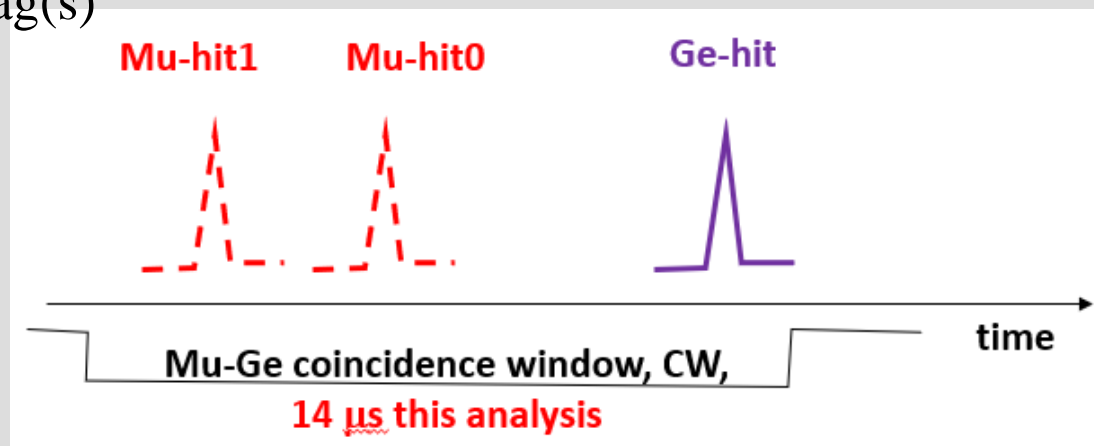
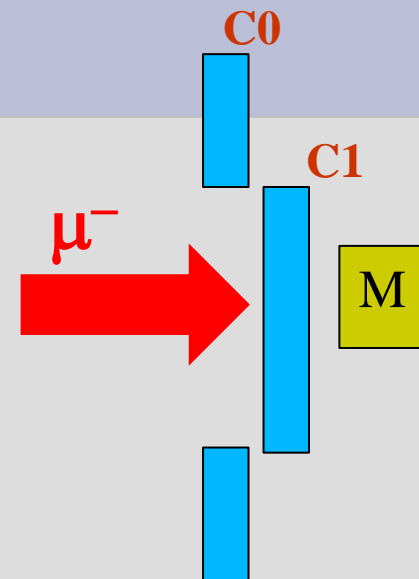
// either in ge- or/and muon hit

// 4 - **uncorrelated**: no muons in C1 in CW

// 5 - **uncorrelated0**: like 4, but was muon in C0

// 6 - **good0**: like 1, but was muon in C0

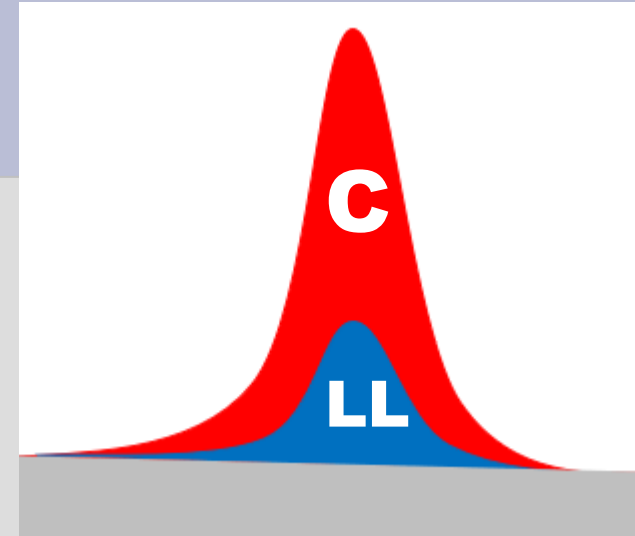
For yield analysis we group **1-3+6-hit as correlated (C)**, and **4+5 as uncorrelated (LL)**.



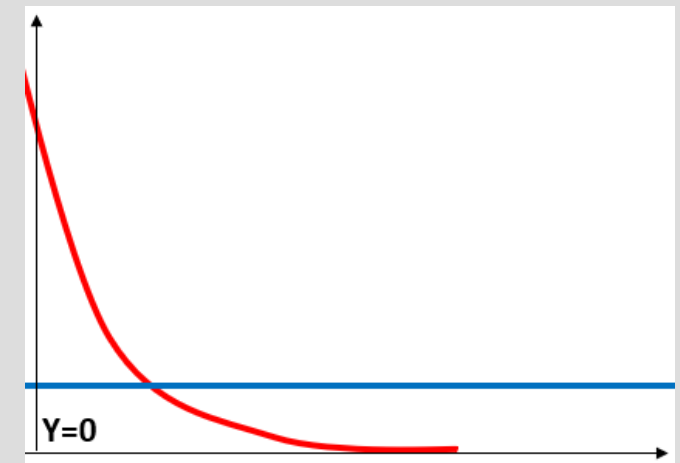
# $\gamma$ -peak: contributions

- In  $\gamma$ -peak we have 3 contributions:
  1.  $\gamma$ -rays from isotope produced in OMC (C)
  2.  $\gamma$ -rays from the same long-lived (LL) (w.r.t.  $CW=14 \mu s$ ) isotope (0 if not exist)
  3. From pedestal, but don't care if count peak intensities.
- Our task is to determine both C and LL production w.r.t. to OMC rate

Energy spectra

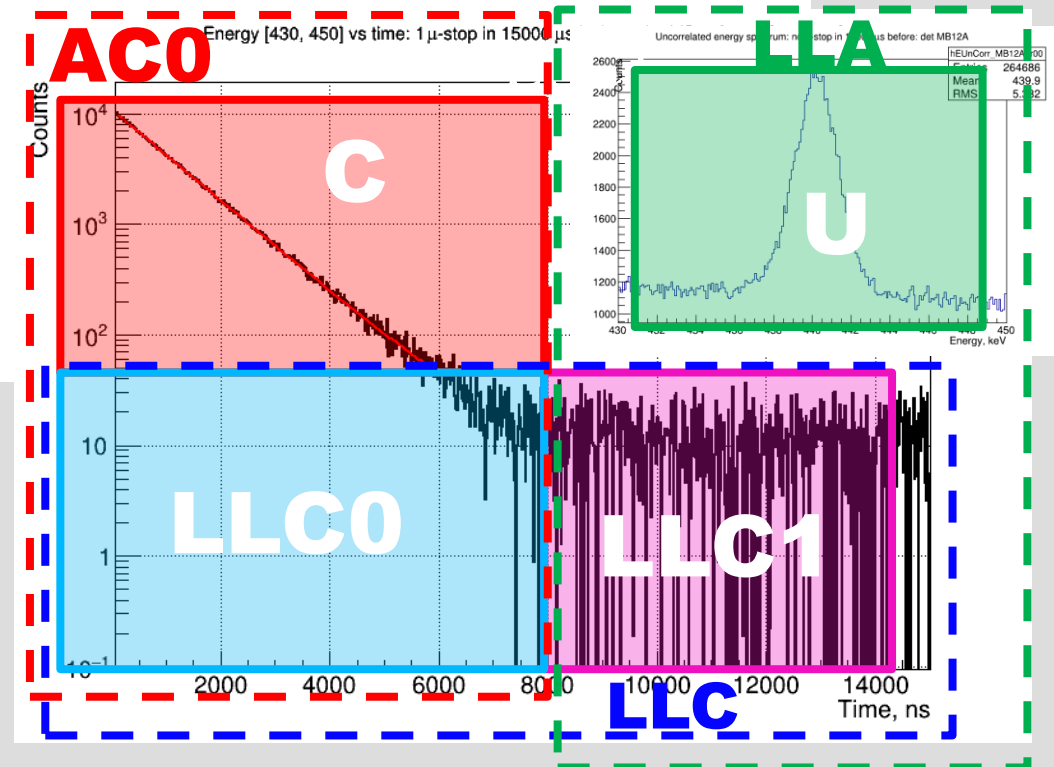
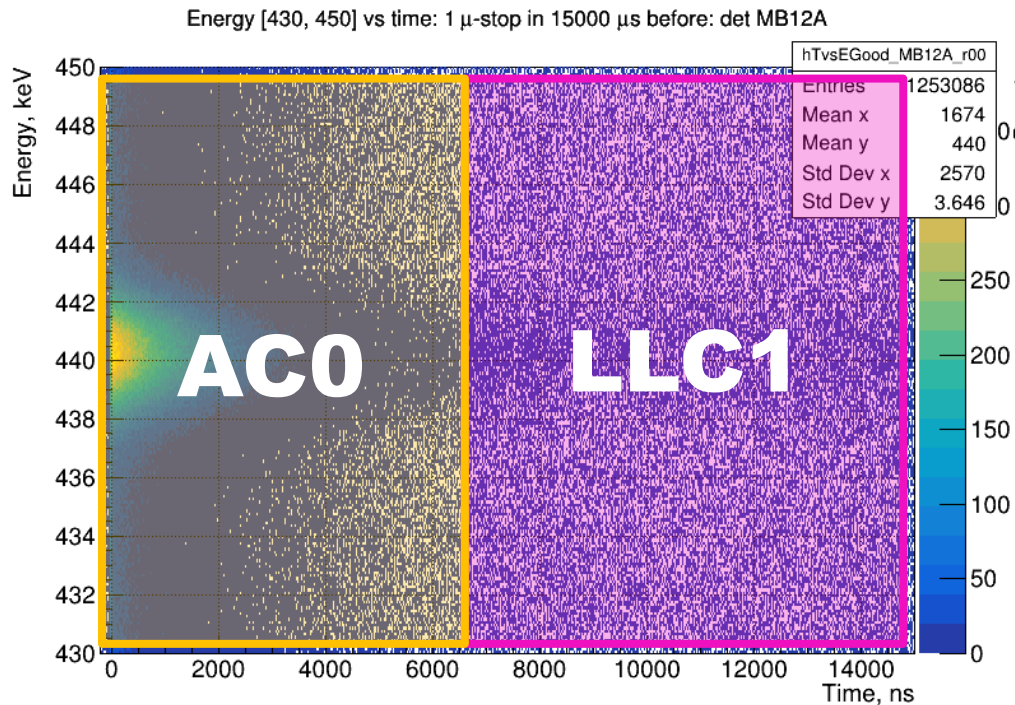


Time spectra



# Fractions of spectra (applied for 440 keV for example)

- **LLC1** – Long-lived part of spectra in uncorrelated part of correlated spectra
- **LLC0** – long-lived part of spectra in OMC-correlated part of spectra
- **C** – OMC-correlated (directly produced in OMC) events
- **U** – long-lived part in uncorrelated spectra.

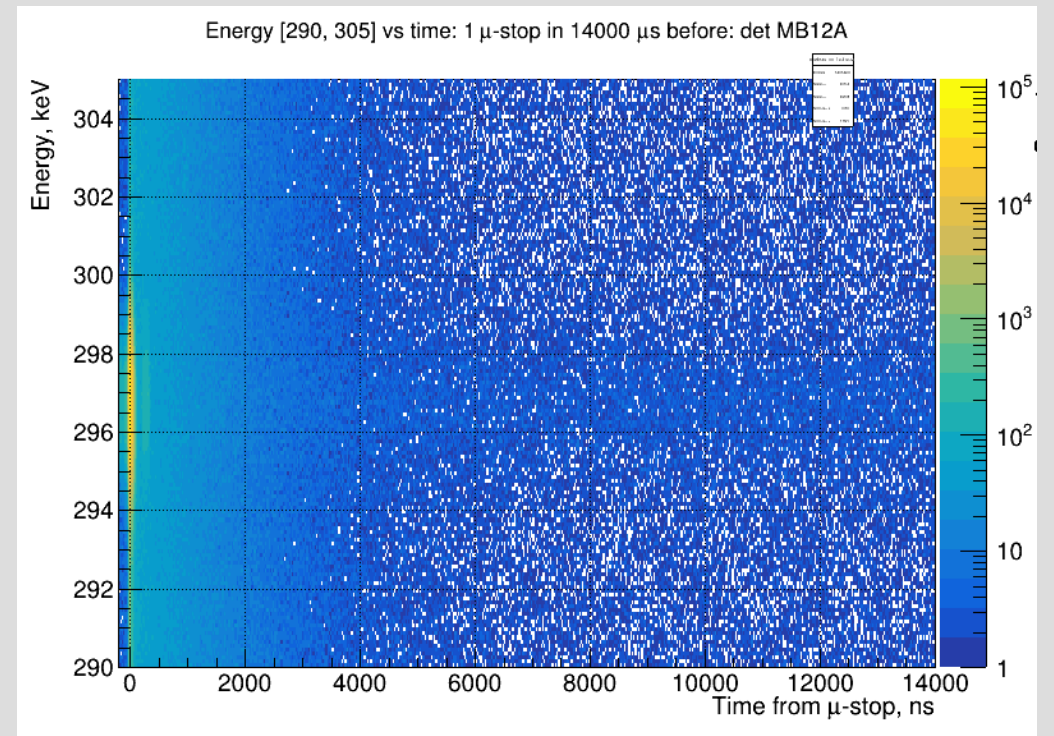
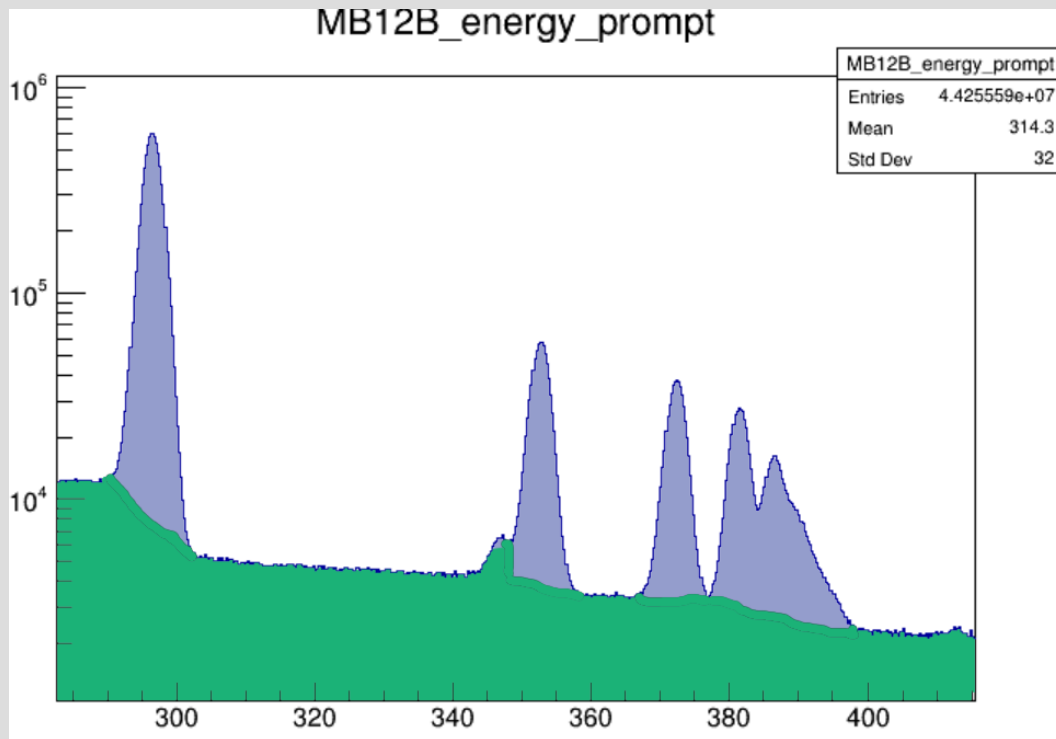


# How to calculate **C** and **LLA**?

1. Taking 2D histos for all correlated event ( Good+Mult + Good0),
2. Making **hAC0** 1D histo as TH2::ProectionY(), for correlated part of spectra [0, Tmin=8] mks fit counting area (no fit) and **AC0 = C + LLC0**.
3. Making **hLLC1** 1D histo as TH2::ProjectionY() for OMC-free tail in [Tmin=8, Tmax=14] mks
4. Make aggregated histo **hLLC1U = hU** (histo of all uncorrelated events **U + U0**) + **hLLC1**, fit it with gauss + pol1 model and take gaus area as **LLC1U = LLC1 + LLU**. **Fix position&sigma of gauss for all next fits** as it is natural peak w/o Doppler broadening with maximal intensity.
5. Fit **hLLC1** with gauss + pol1 model (fixed position & sigma from 4.) and take gauss area as **LLC1**.
6. Assuming flat LL distribution in correlated spectra one can calculate **LLC0 = LLC1 \* Tmin / (Tmax - Tmin)=LLC1\*4/3**.
7. Now one can count **total intensity of OMC-correlated events** as **C = AC0 - LLC0**
8. Finally calculate **total number of LL events** as **LLA = LLC1U + LLC0**.

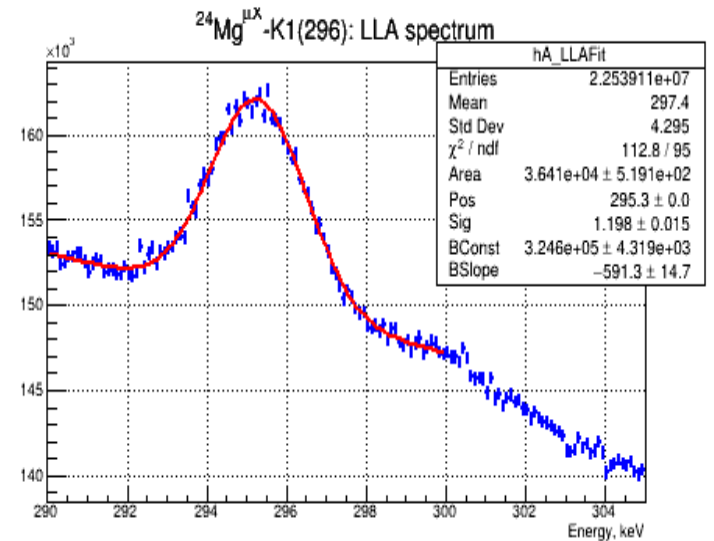
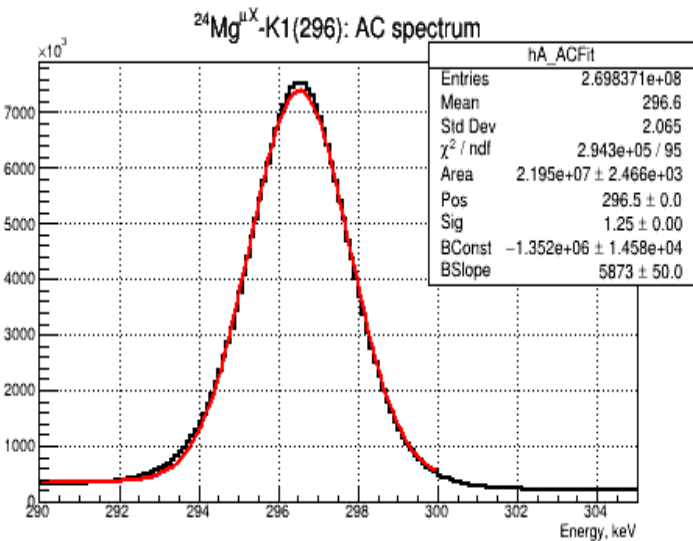
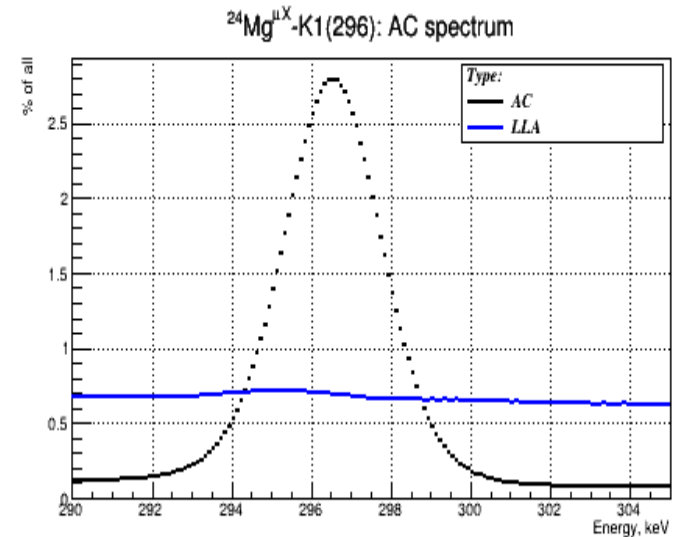
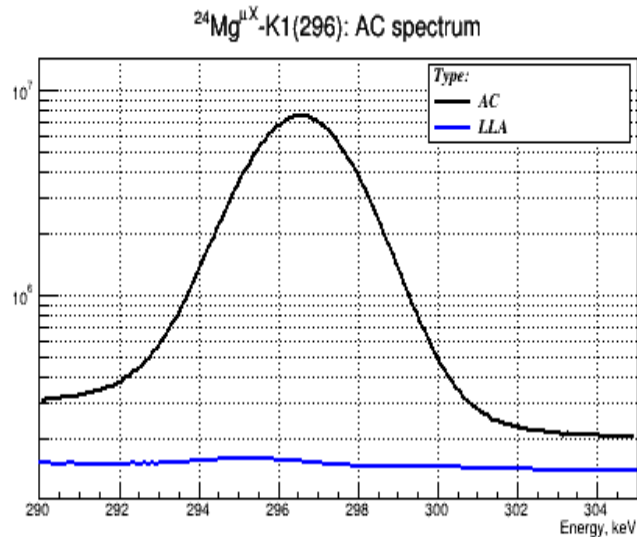
# Defining number of mu-stops

- Determined using full set of Mg-24  $\mu$ X-ray K-series lines. Peak intensities has been counted as all counts (light blue) above linear background (shown by green color).



# Check random coincidences on $\mu\text{X}$ K- $\alpha$

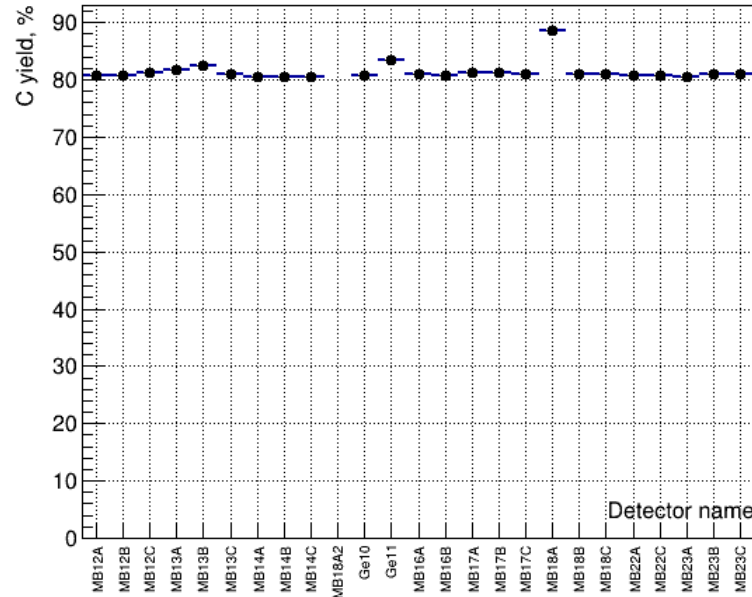
- It is obvious that events correlated with OMS can randomly fall into the uncorrelated part of the spectrum for various reasons.
- The proportion of such events can be estimated by repeating the described analysis for the  $\mu\text{X}$ -ray K- $\alpha$  line, calculating the number of **correlated AC** and **all long-lived LLA**.
- The analysis shows that the proportion of random **LLA/AC = 0.17%**, while for some reason **there is a shift in the peak position by 1.23 keV** for correlated and random LLA. The reason for the shift is not clear.



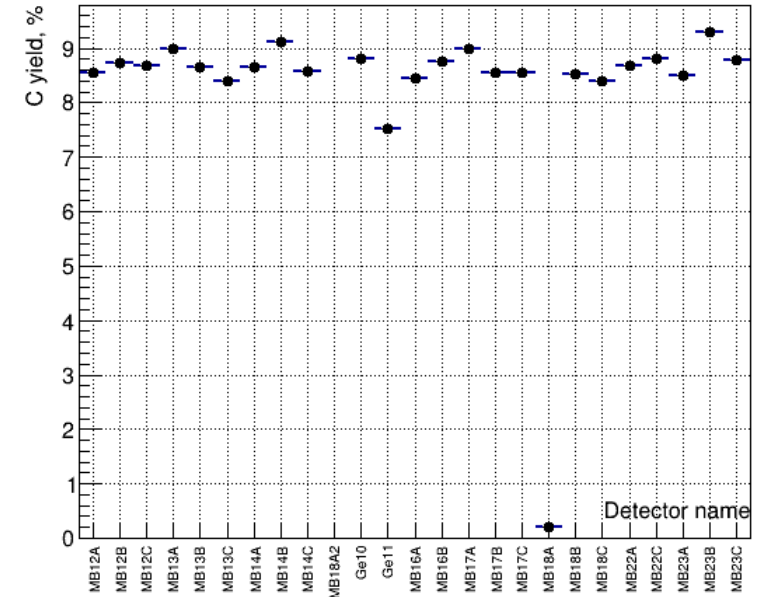
# Test on $\mu X$ K-series

- Determined ratios of lines from K-series of  $\mu X$ .
- No correction on efficiencies has been applied here

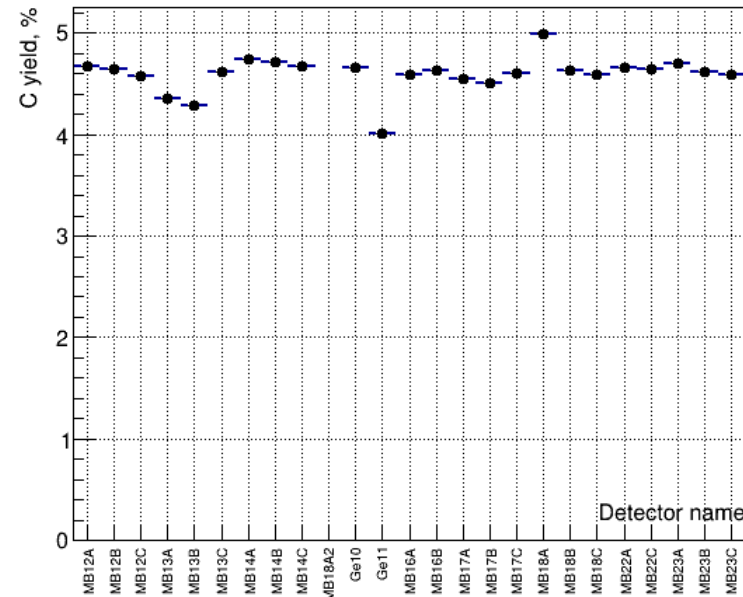
$^{24}\text{Mg}^{\mu X}$ -K1(296): OMC-correlated (C) yield



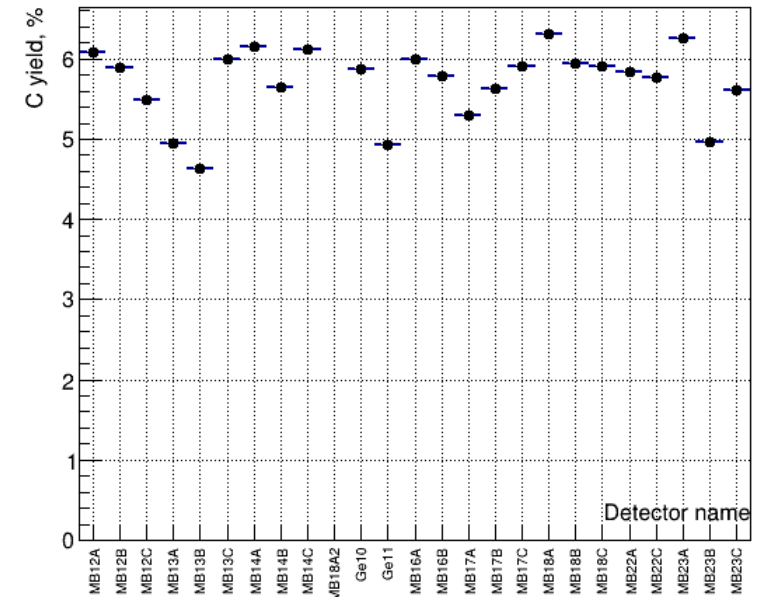
$^{24}\text{Mg}^{\mu X}$ -K2(353): OMC-correlated (C) yield



$^{24}\text{Mg}^{\mu X}$ -K3(373): OMC-correlated (C) yield



$^{24}\text{Mg}^{\mu X}$ -K4-N: OMC-correlated (C) yield





# Results for different lines

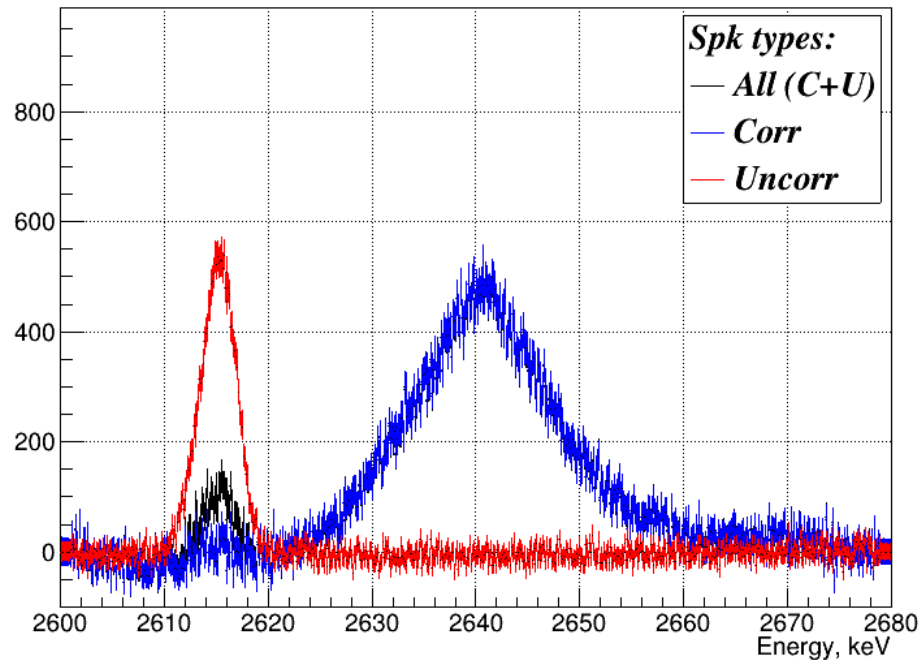
- **NOTE!** Next some results were **NOT** corrected on detector efficiency factor  $\text{eff}(\sum \mu X(K))/\text{eff}(E\gamma)$ .
- Just show some interesting lines to demonstrate different features

# Example of lines

- Left picture - center - clean correlated line with clean doppler broadening – perfect triangle.
- Left wing - standard uncorrelated TI-line
- Right picture - K-40 line

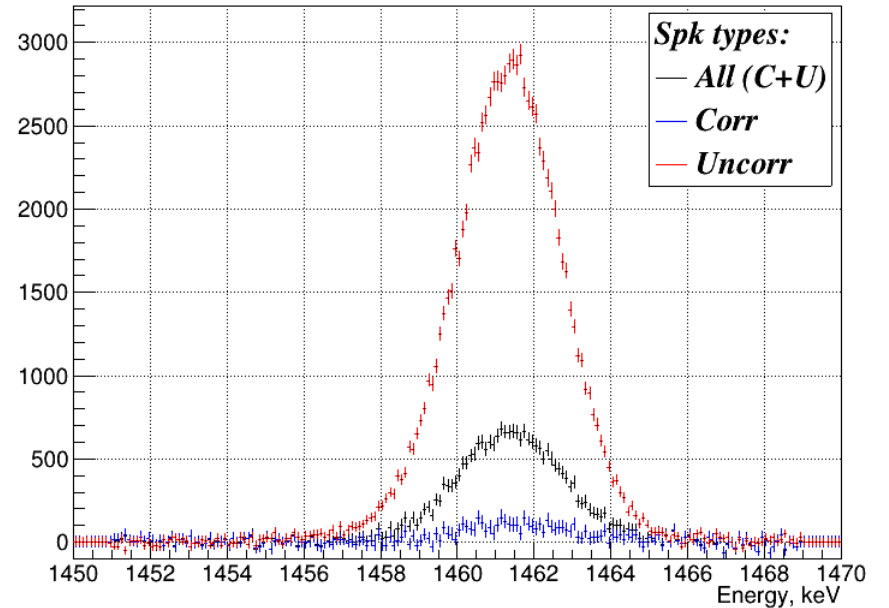
$^{23}\text{Na}(2640 \text{ keV})$  &  $^{208}\text{Tl}(2615 \text{ keV})$

$^{23}\text{Na}(2640)$ : different type of spectra



$^{40}\text{K}$  (1491 keV)

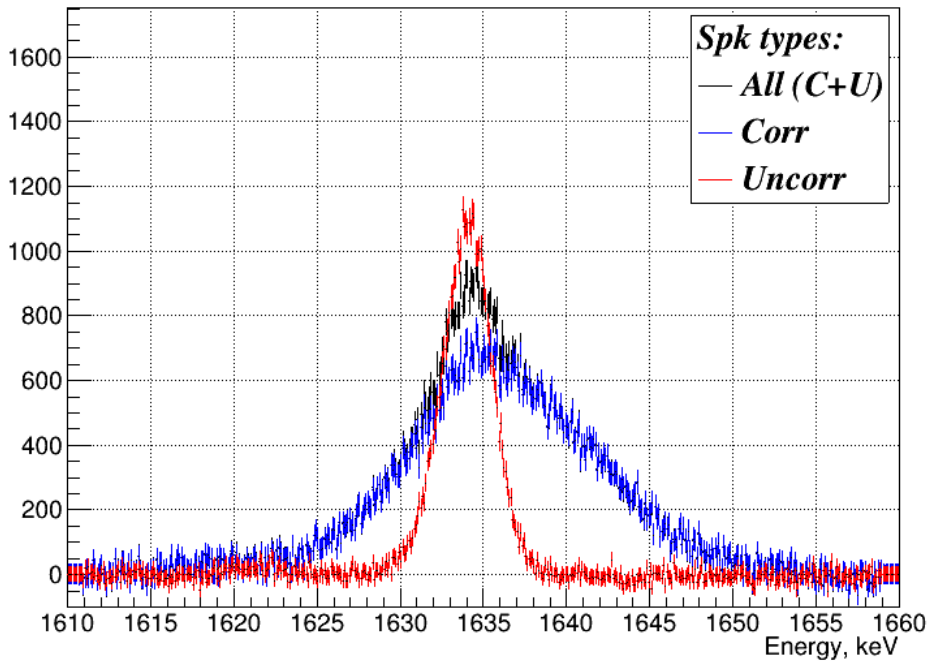
$^{40}\text{K}(1461)$ : different type of spectra



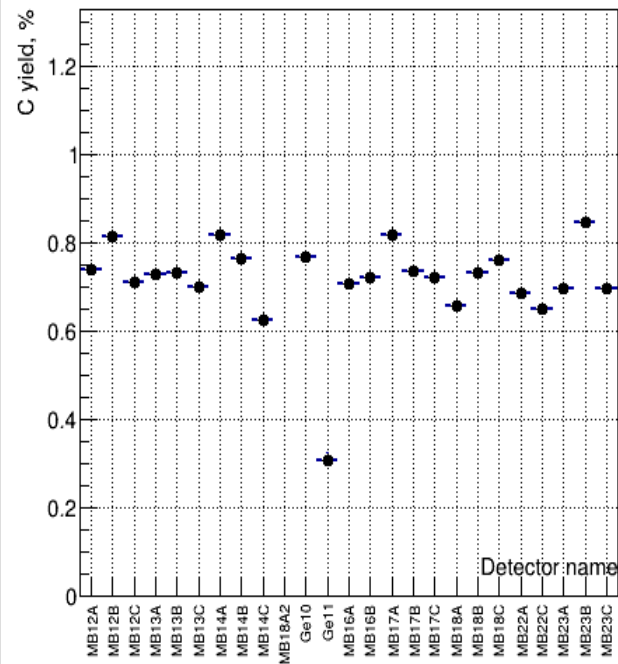
# $^{23}\text{Na}(1636 \text{ keV})$

- Beautiful example of mixture doppler broadened and asymmetric correlated & uncorrelated lines.

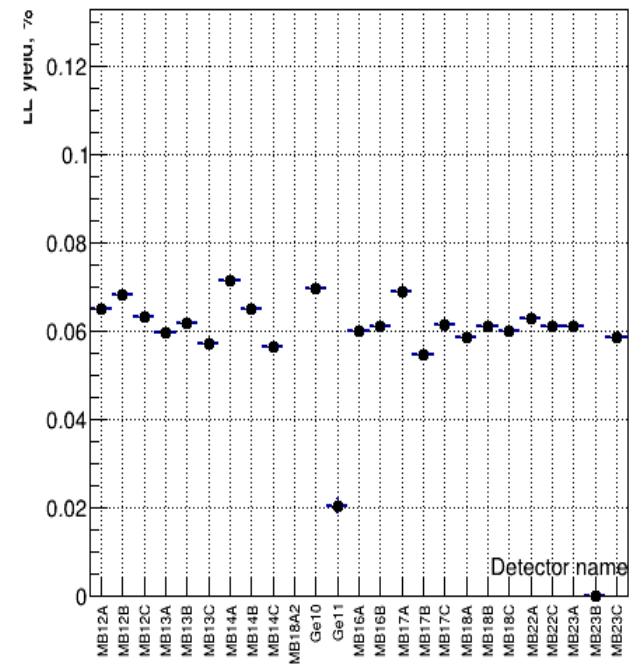
$^{23}\text{Na}(1636)$ : different type of spectra



$^{23}\text{Na}(1636)$ : OMC-correlated (C) yield



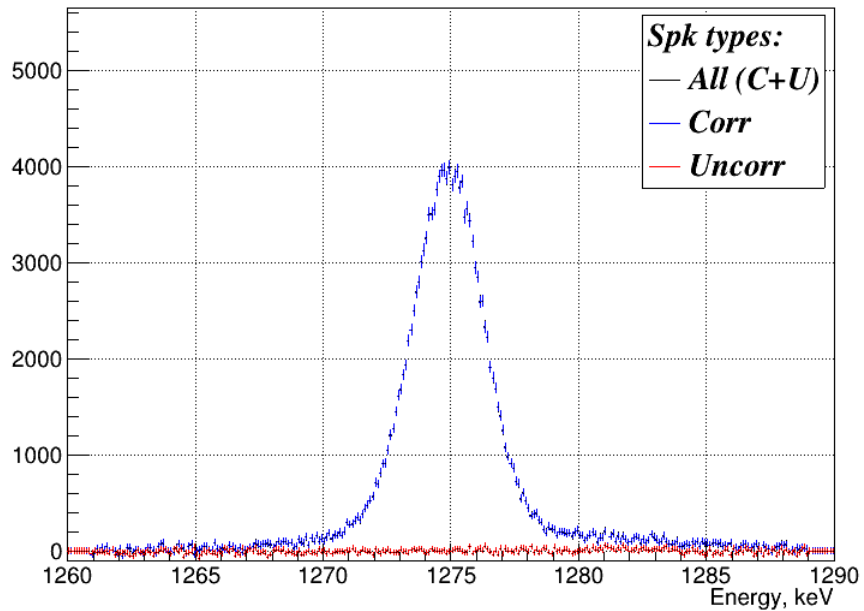
$^{23}\text{Na}(1636)$ : Long-lived (LL) yield



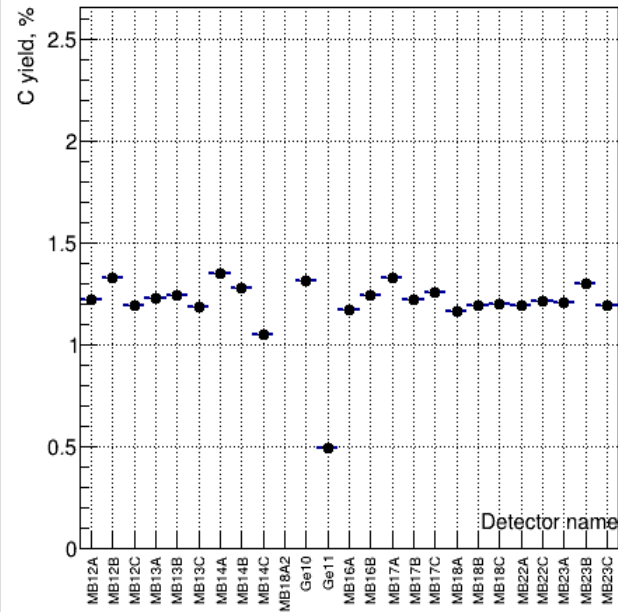
# $^{20}\text{Ne}(1274 \text{ keV})$

- Clean correlated line without visible doppler broadening

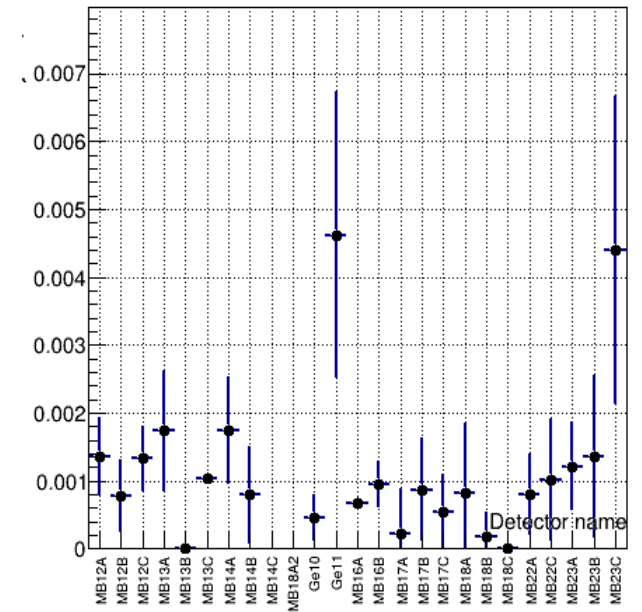
$^{22}\text{Ne}(1274)$ : different type of spectra



$^{22}\text{Ne}(1274)$ : OMC-correlated (C) yield



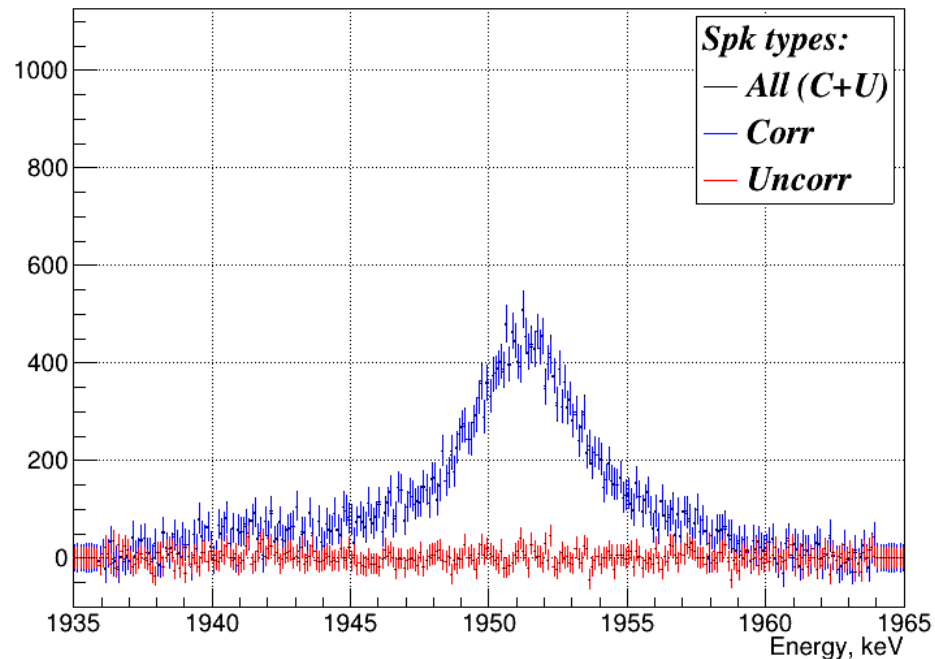
$^{22}\text{Ne}(1274)$ : Long-lived (LL) yield



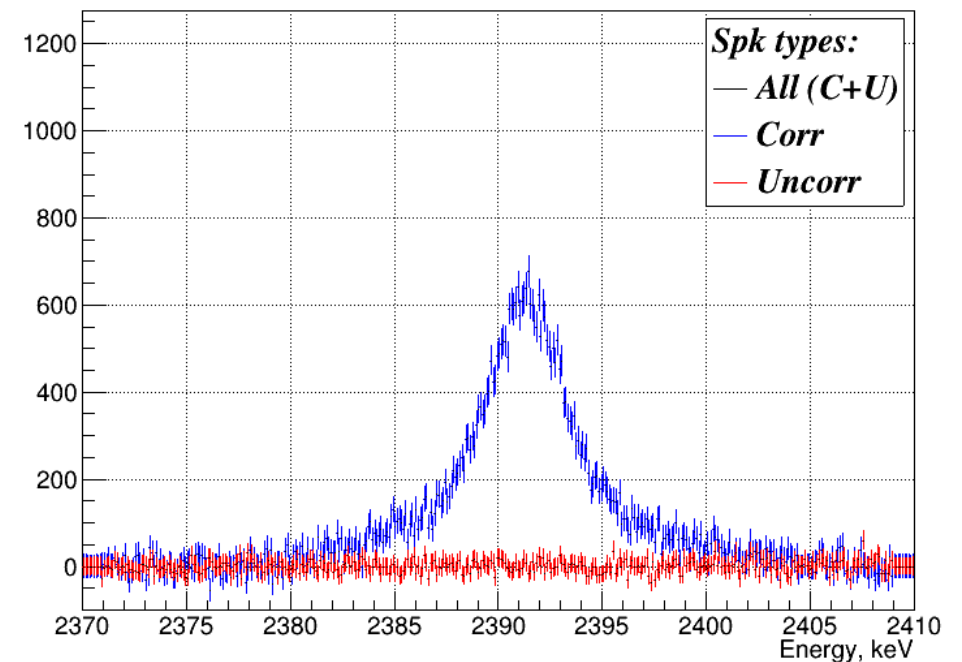
# $^{23}\text{Na}$ (1950 & 2391 keV)

- Clean correlated lines with visible doppler broadening – right left tails

$^{23}\text{Na}$ (1950): different type of spectra



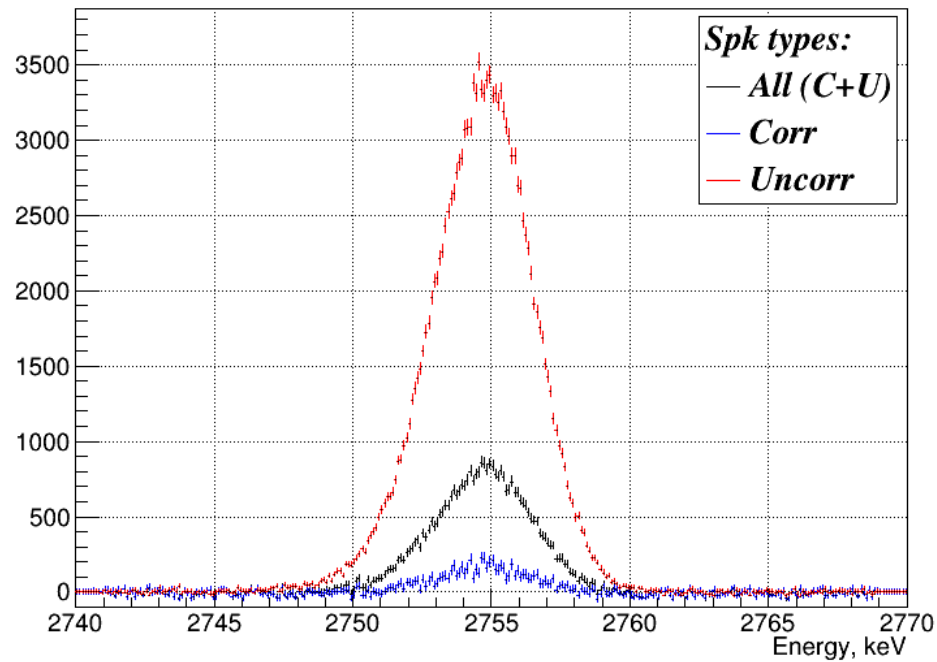
$^{23}\text{Na}$ (2391): different type of spectra



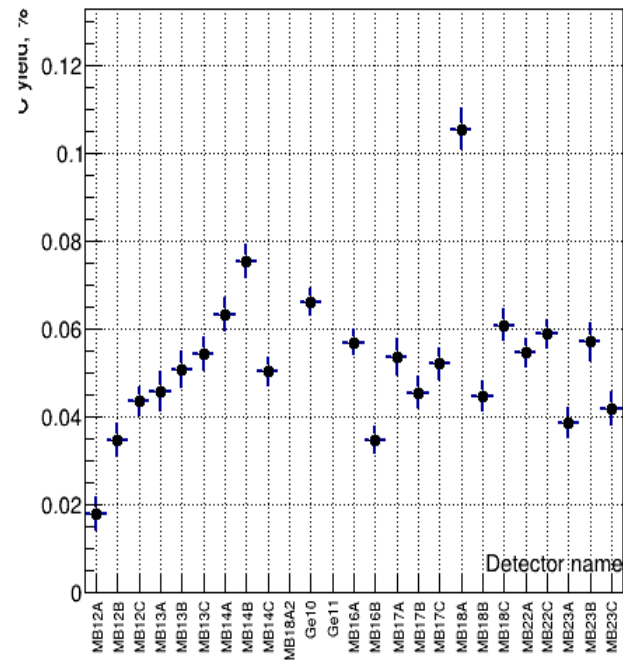
# $^{24}\text{Na}(2754 \text{ keV})$

- Mixture of correlated (no visible doppler broadening) & uncorrelated lines.

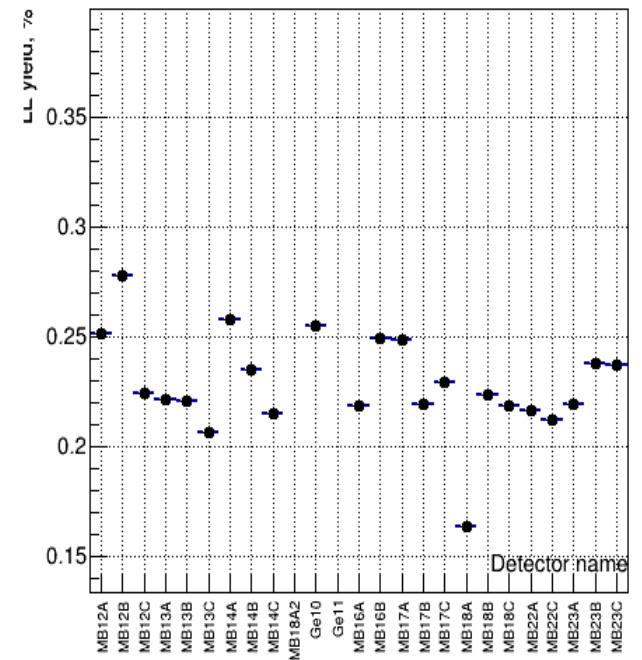
$^{24}\text{Na}(2754)$ : different type of spectra



$^{24}\text{Na}(2754)$ : OMC-correlated (C) yield



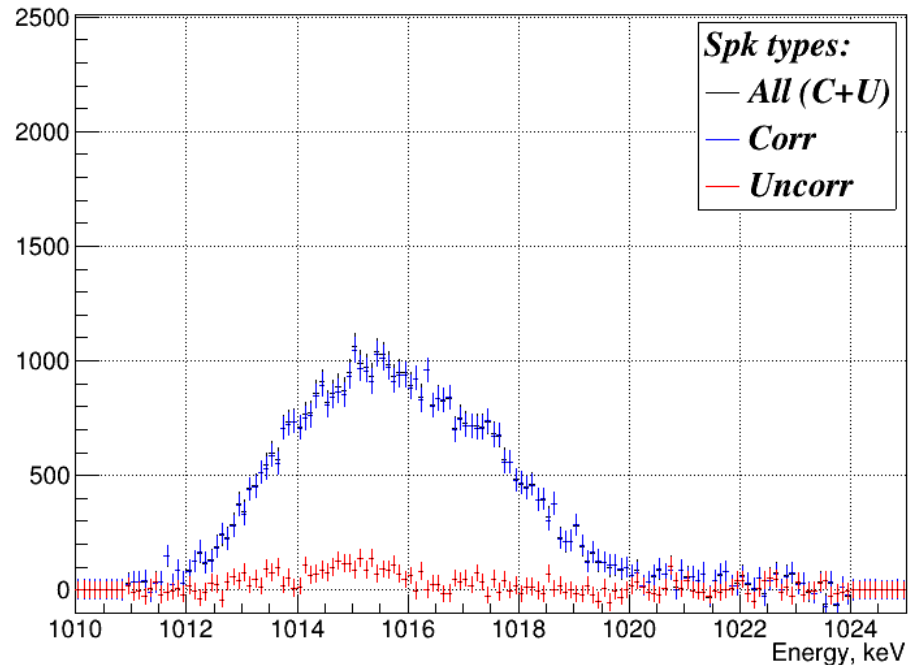
$^{24}\text{Na}(2754)$ : Long-lived (LL) yield



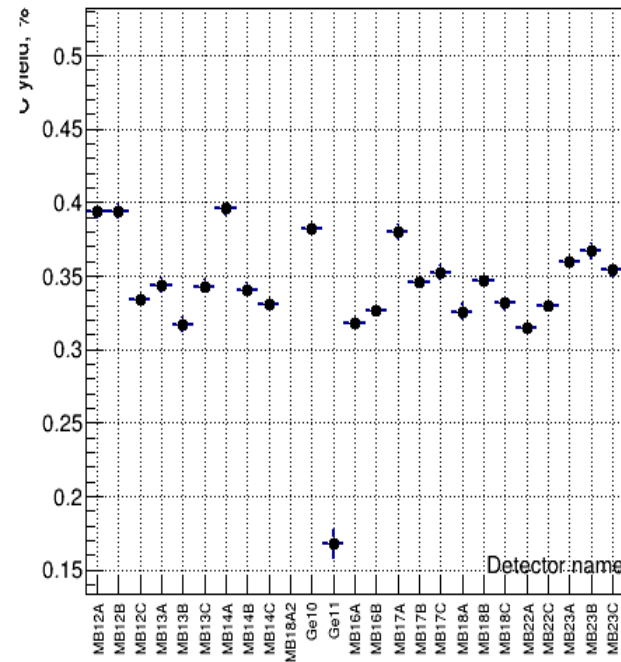
# $^{27}\text{Al}(1015 \text{ keV})$

- Pure correlated, but strong doppler broadening. Could be mixture of OMC from Mg-24 and Al-27 – to be checked with the fit of time evolution...

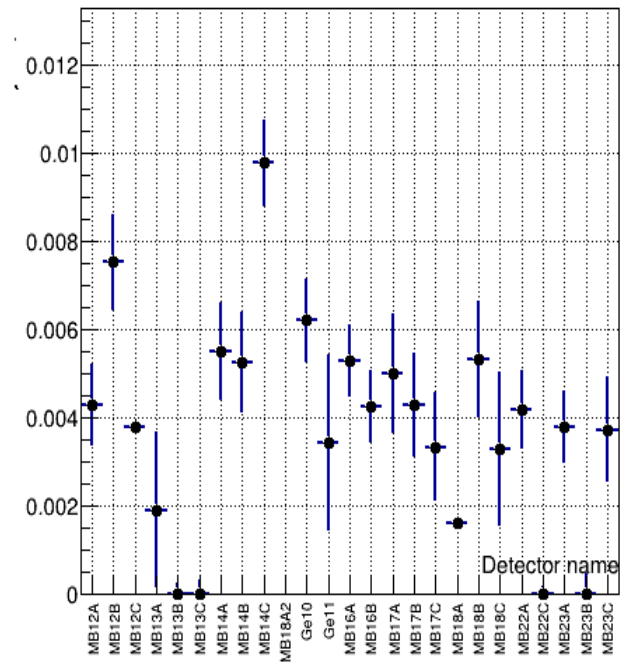
$^{27}\text{Al}(1011)$ : different type of spectra



$^{27}\text{Al}(1011)$ : OMC-correlated (C) yield

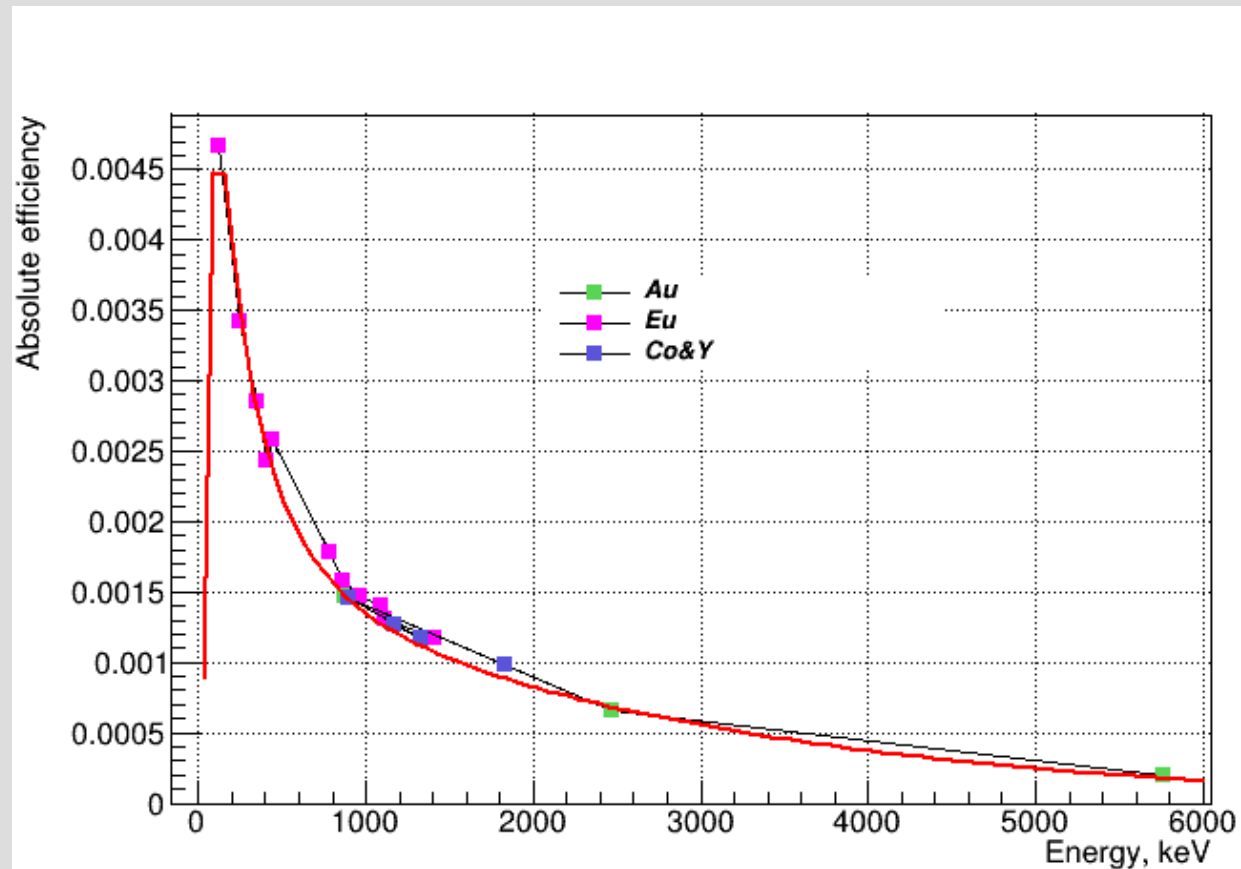


$^{27}\text{Al}(1011)$ : Long-lived (LL) yield



# Efficiencies are taken into account...

We are adding efficiencies calculated by Nadya **Efficiency factor  $\text{eff}(\sum \mu X(K))/\text{eff}(E_\gamma)$ .**





# And branching ratios

ID in code (gr==2)	E_g	Nature	Branching		
0	350.7	21Na / 22.49 s	0.0507		
1	439.9	23Ne / 37.24 s	0.33		
2	472.2	24mNa/ 20.18ms	0.9995		
	996.6	??		не видно в спектре	
3	1014.52	27Mg / 9.458 m	0.282		
	843.52	27Mg / 9.458 m	0.718	слияние с другой линией	
4	1274.5	22Na / 2.6018 y	0.9994		
5	1368.6	24Na /14.997h	0.999936		
6	1635.6	23Ne / 37.24 s	0.01		
7	2075.9	23Ne / 37.24 s	0.00101		
8	2754.007	24Na /14.997h	0.99855		
	1395.1	21F / 4.158 s	0.153	не видно в спектре	
	1633.6	20F / 11.163 s	0.991	слияние с 23Ne	
	197.1	??			
	1356.8	??			
	659.2	???			
	1041.5	??			
	1700.7	27Mg / OMC			
	1809	27Al(n,n')			

Normalization on branching ratios was added for **long-lived lines**

# Preliminary results of the RI yields produced in OMC in $^{24}\text{Mg}$ ( $W=-0,2-6-12$ us) (MB23A)

Uncorr-d_th	Energy_all	Uncor_d	Isotope/Life-time	%_BR_theor	eff-cy_exp	Int_ener_all	Int-ty_Uncorr-d	Int-ty_Uncor/BR_theor	Ycap_Uncord	Pcap_Uncord=Ycap*Lkoef	Int-ty_Eall/BR_theor	Ycap_Eall	Pcap_Eall
472,2	50812,4	29805,1	$^{24}\text{Mg}/20,18\text{ms}$	0,9995	0,002232	70641,074	41436,03284	41411,15878	0,080362925	0,14051779	70598,66817	0,08090816	0,14147115
350,7	70092,1	2000,6	$^{21}\text{Na}/22,49$ s	0,0507	0,0027615	78760,0168	2248,003549	44290,54724	0,085950697	0,150288231	1551743,21	1,77834353	3,10950479
439,9	111287	1489,45	$^{23}\text{Ne}/37,24$ s	0,33	0,0023485	147040,052	1967,963956	5956,967259	0,011560153	0,020213389	445085,7802	0,51008144	0,89189781
1274,5			$^{22}\text{Na}/2,6018$ y	0,9994	0,0011567								
1368,6	23940	13359	$^{24}\text{Na}/14,997\text{h}$	0,999936	0,00110622	67152,8448	37472,63383	37433,8097	0,07264444	0,12702171	67083,27002	0,07687941	0,13442672
2754,007	13460	7895	$^{24}\text{Na}/14,997\text{h}$	0,99855	0,000656	63668,2622	37344,79421	37357,88387	0,072497098	0,126764076	63690,57845	0,07299128	0,12762818
1635,6			$^{23}\text{Ne}/37,24$ s	0,01	0,0009865								
2075,9			$^{23}\text{Ne}/37,24$ s	0,00101	0,0008345								
1633,6	6840,71	1997,76	$^{20}\text{F}/11,6$ s	0,991	0,0009873	21499,7702	6278,789912	6328,842828	0,012281818	0,02147525	21671,16091	0,02483579	0,04342637

$$\tau = 945 \text{ ns}$$

$$\Lambda_{\text{cap}} = 0,605 \times 10^6$$

$$\Lambda_{\text{tot}} = 1,058 \times 10^6$$

$$\Lambda_{\text{koef}} = 1,74864$$

Yield of  $^{24}\text{Na}$  in OMC with  $^{24}\text{Mg}$  is: 12,7 %

Yield of  $^{24\text{m}}\text{Na}$  in OMC with  $^{24}\text{Mg}$  is: 14,05 %

Yield of  $^{23}\text{Ne}$  in OMC with  $^{24}\text{Mg}$  is: 2,02 %

Yield of  $^{20}\text{F}$  in OMC with  $^{24}\text{Mg}$  is: 2,14 %

From  
Daniya

# Preliminary results of the RI yields produced in OMC in $^{24}\text{Mg}$ ( $W=-0,2-6-12$ us) (MB18B)

Uncorr-d_th	Energy_all	Uncor_d	Isotope/Life-time	%_BR_theor	eff-cy_exp	Int_ener_all	Int-ty_Uncorr-d	Int-ty_Uncor/BR_theor	Ycap_Uncord	Pcap_Uncord=Ycap*Lkoef	Int-ty_Eall/BR_theor	Ycap_Eall	Pcap_Eall
472,2	47749,7	28049,1	$^{24}\text{mNa} / 20,18\text{ms}$	0,9995	0,00225642	66332,4302	38964,95617	38941,5655	0,079149248	0,138395626	66292,61082	0,078657449	0,1375357
350,7	70092,1	2000,6	$^{21}\text{Na} / 22,49$ s	0,0507	0,0027918	78697,325	2246,214174	44255,29266	0,089949469	0,157280244	1550508,047	1,839707405	3,21680199
439,9	105965	1407,49	$^{23}\text{Ne} / 37,24$ s	0,33	0,0023745	139883,172	1858,011278	5624,143835	0,011431147	0,019987818	423422,1213	0,502398432	0,87846375
874,4													
996,6													
1274,5			$^{22}\text{Na} / 2,6018$ y	0,9994	0,00114093								
1368,6	22912,6	13021,9	$^{24}\text{Na} / 14,997\text{h}$	0,999936	0,001087	66072,3922	37550,87088	37511,96569	0,076243567	0,133314927	66003,93683	0,078314931	0,13693679
2754,007	12988,7	7641,54	$^{24}\text{Na} / 14,997\text{h}$	0,99855	0,00060791	66973,2848	39401,86739	39415,67807	0,080112888	0,140080589	66996,75952	0,079492935	0,13899658
1635,6			$^{23}\text{Ne} / 37,24$ s	0,01	0,00095869								
2075,9			$^{23}\text{Ne} / 37,24$ s	0,00101	0,0007959								
1395,1													
1633,6	4369,9	1840,46		0,991	0,00095954	14275,2465	6012,270351	6060,198642	0,012317434	0,021537526	14389,04515	0,017072877	0,02985261

$$\tau = 945 \text{ ns}$$

$$\Lambda_{\text{cap}} = 0,605 \times 10^6$$

$$\Lambda_{\text{tot}} = 1,058 \times 10^6$$

$$\Lambda_{\text{koef}} = 1,74864$$

Yield of  $^{24}\text{Na}$  in OMC with  $^{24}\text{Mg}$  is: 13,3 %

Yield of  $^{24\text{m}}\text{Na}$  in OMC with  $^{24}\text{Mg}$  is: 13,8 %

Yield of  $^{23}\text{Ne}$  in OMC with  $^{24}\text{Mg}$  is: 1,99 %

Yield of  $^{20}\text{F}$  in OMC with  $^{24}\text{Mg}$  is: 2,15 %

From  
Daniya

# Conclusion

- **Analysis of LL-isotopes is in progress**
- **Two approaches (me and Daniya) are developing.**
- **Comparison of results and joint analysis & result will be soon.**