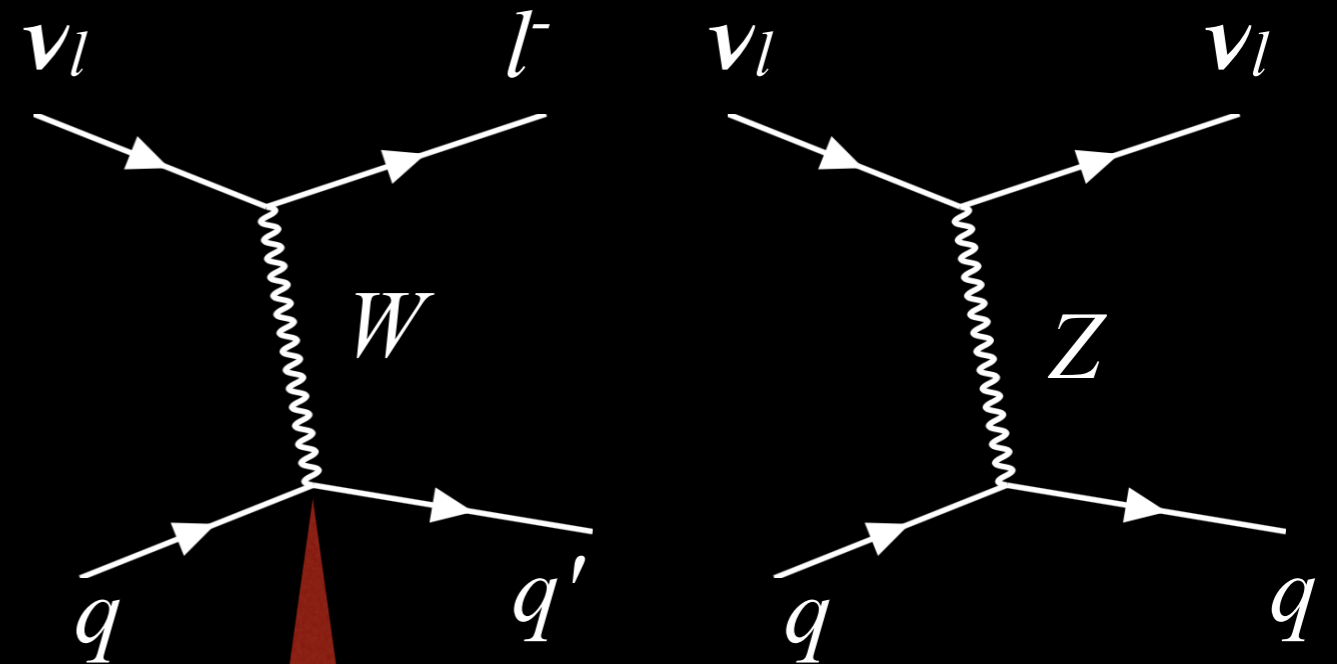
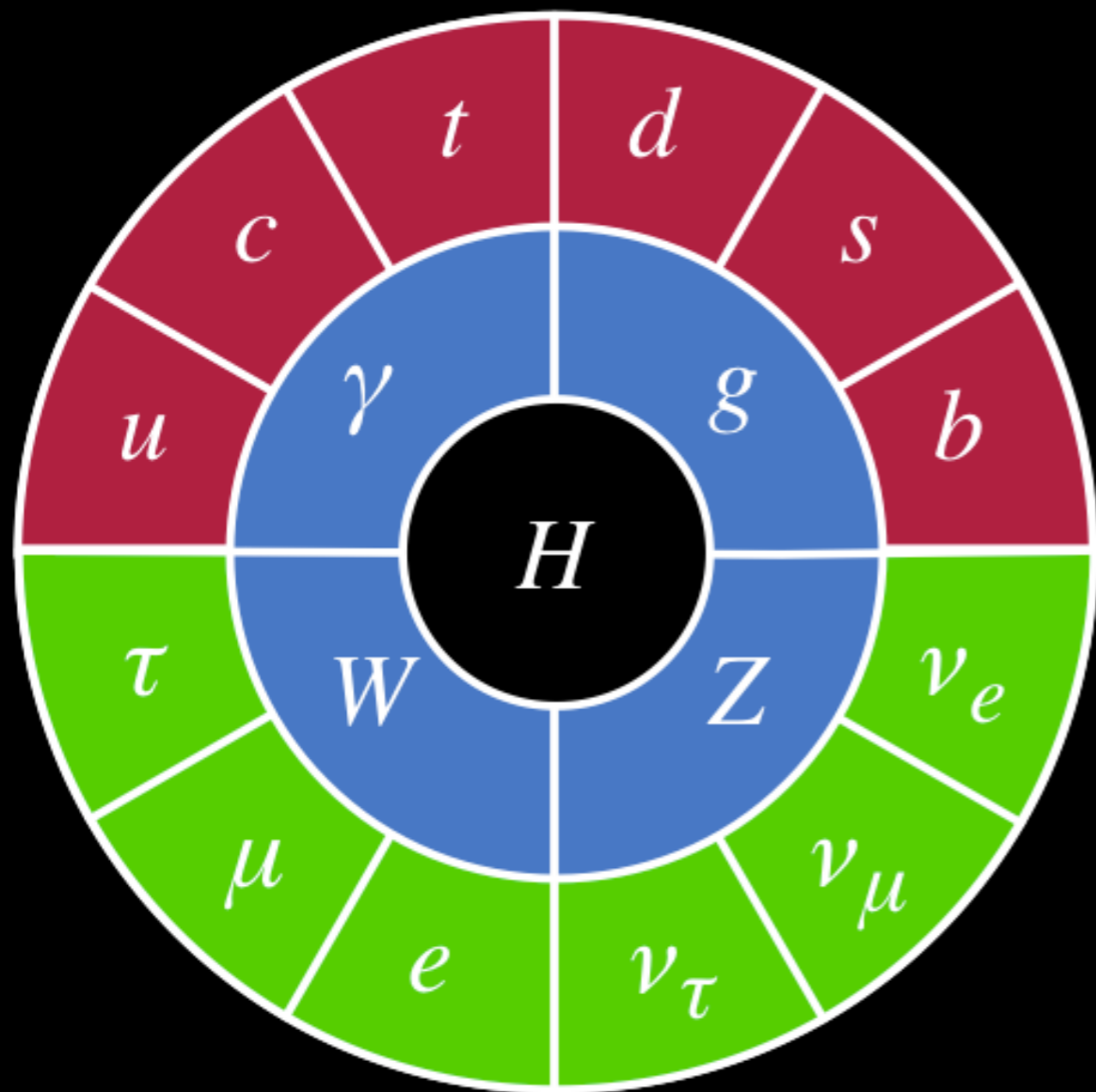
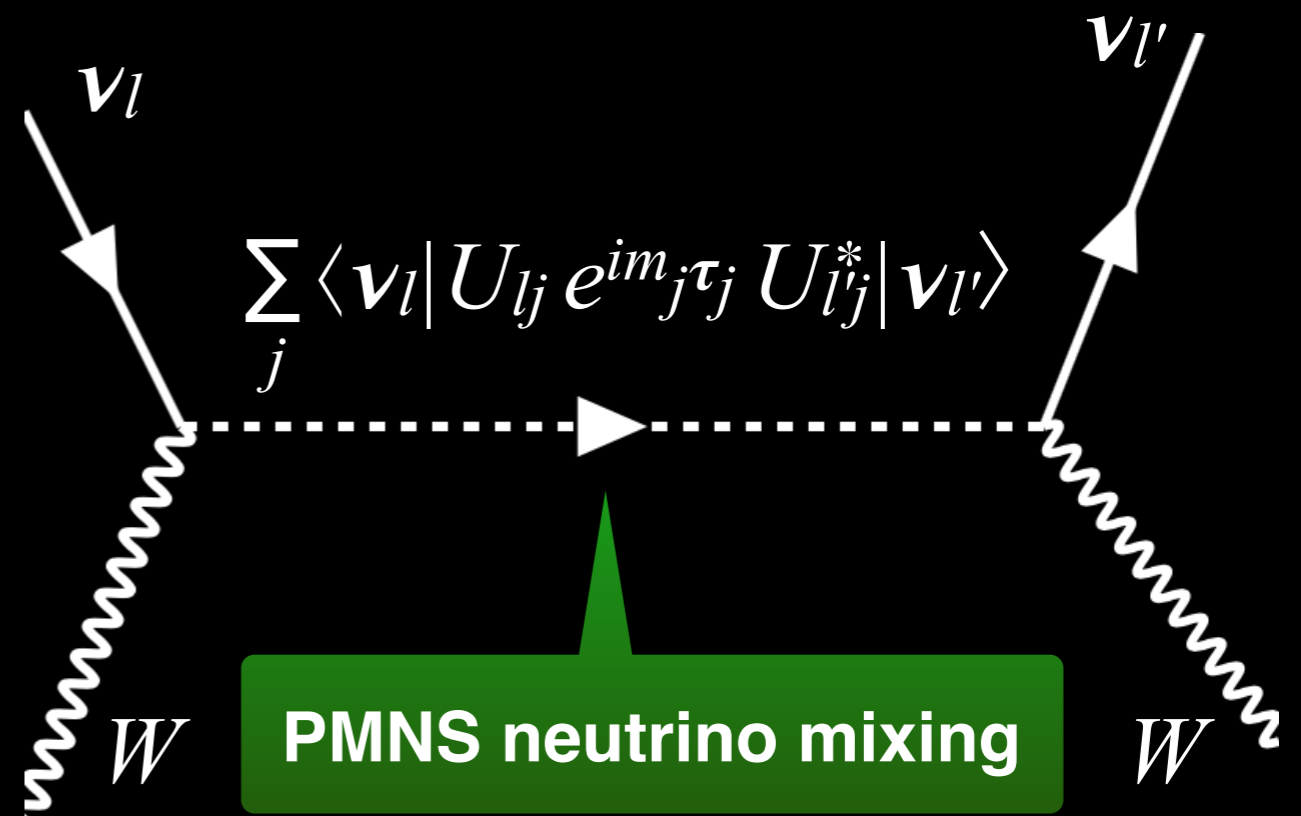


Neutrino



CKM quark mixing



PMNS neutrino mixing

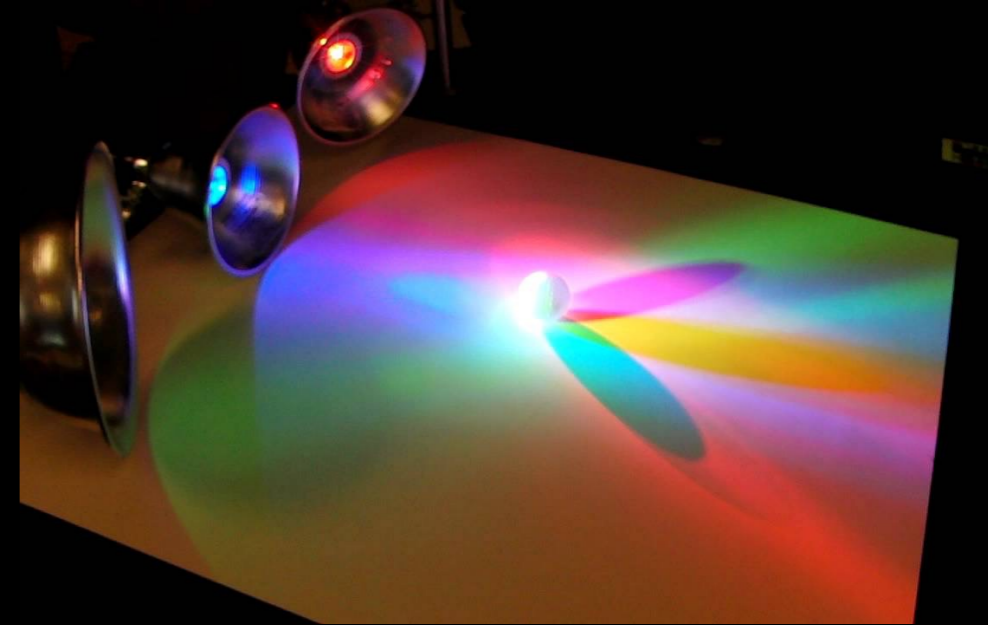
Neutrino

- Neutrinos mix, just like quarks?

$$|\nu_l\rangle = \sum_i U_{li}^* |\nu_i\rangle \quad l=e,\mu,\tau \quad i=1,2,3$$

- PMNS matrix like CKM matrix?

- Unlike the quarks, neutrino mixing are large



	CKM			PMNS		
	d	s	b	ν_1	ν_2	ν_3
u						
c						
t						
ν_e						
ν_μ						
ν_τ						

Neutrino

- Neutrinos mix, just like quarks?

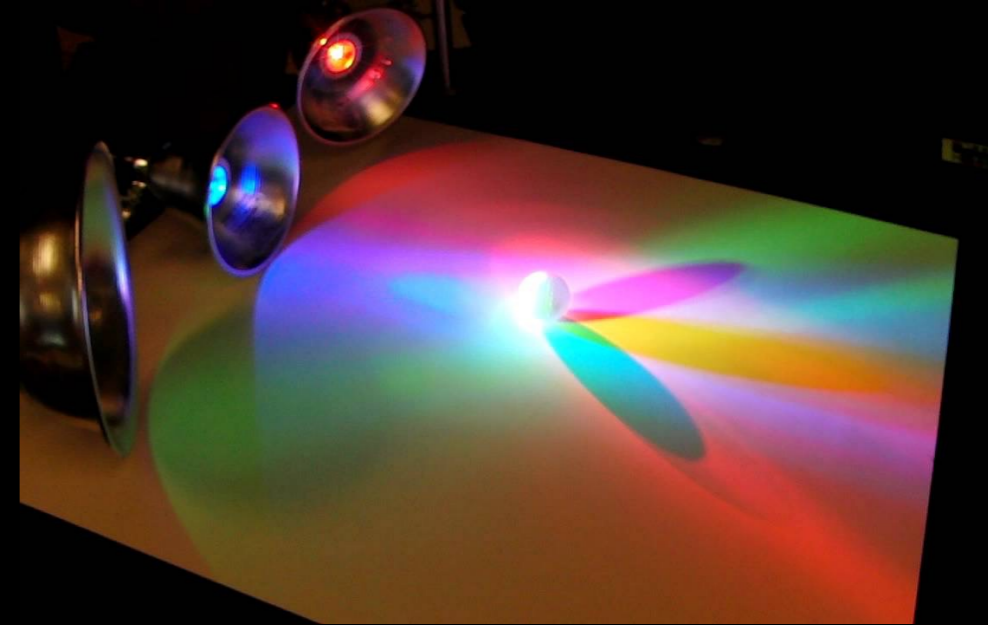
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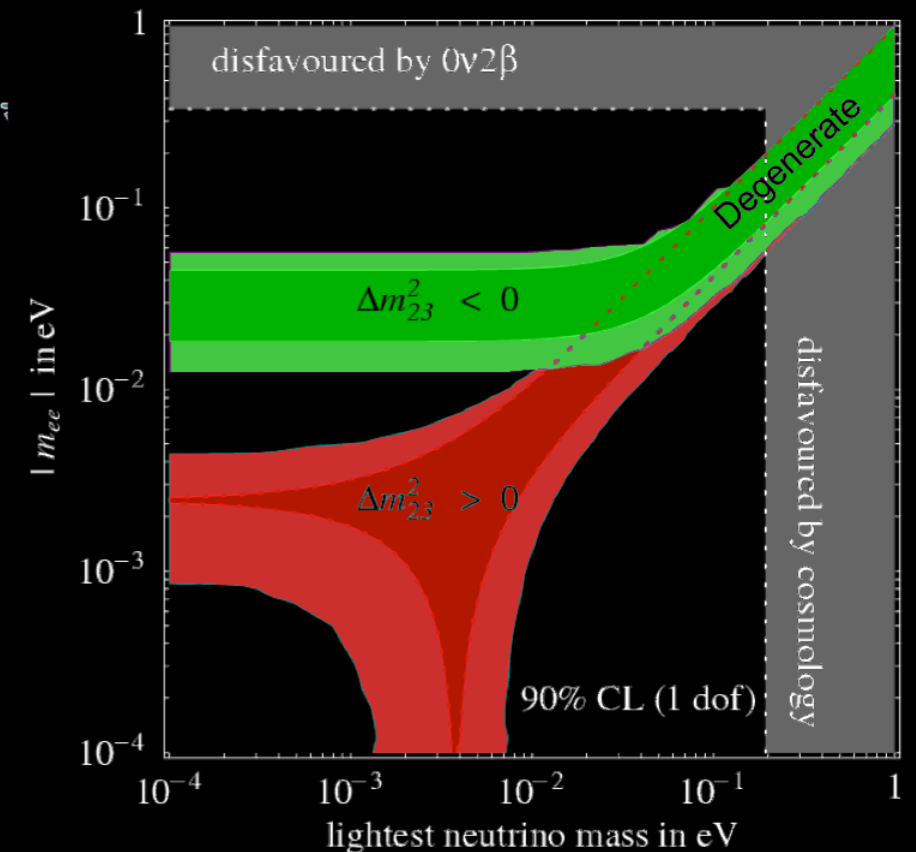
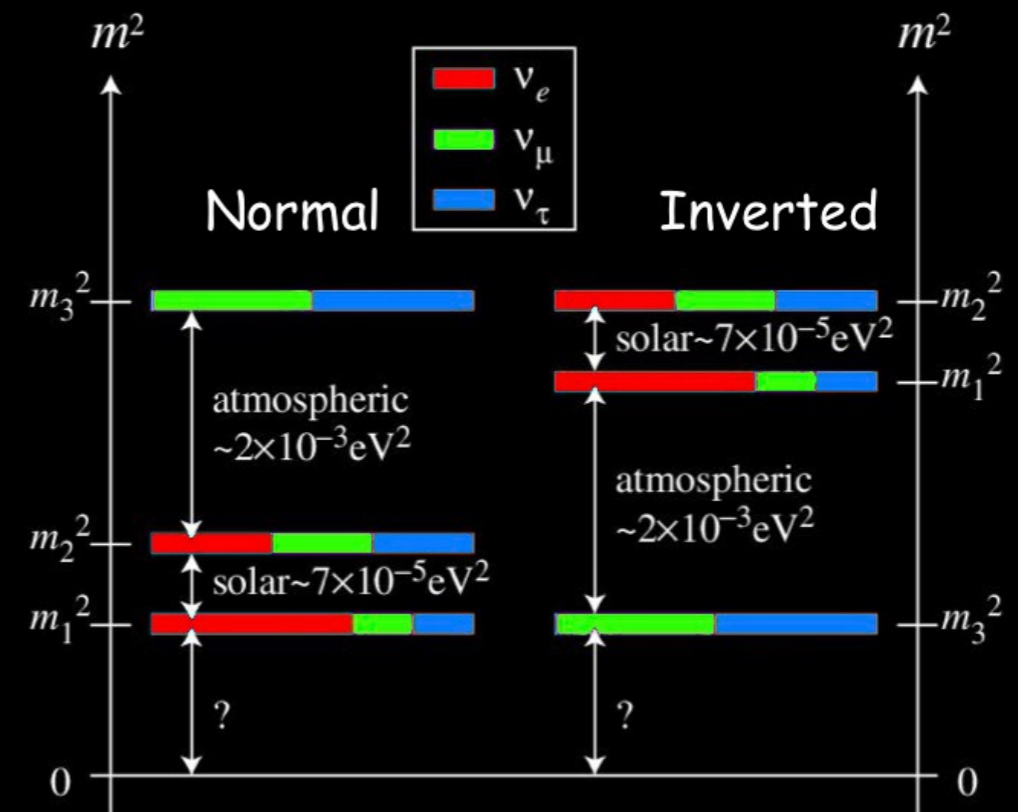
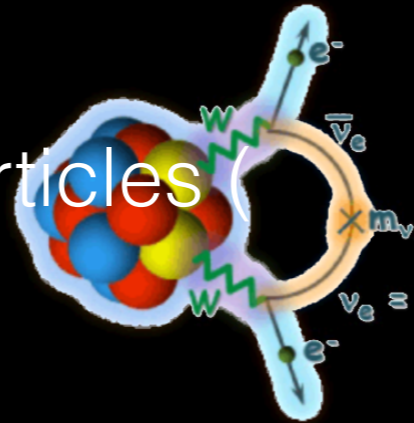
- Open neutrino questions:

Dirac or Majorana? • Absolute masses • Mass ordering • CP-violation • Random mixing parameters or patterns? • Just 3 neutrino types?



Why hierarchy?

- Is the most electron-like state lightest?
- i.e. Does the pattern of the masses match the charged leptons?
- Are neutrinos Majorana particles ($\bar{\nu} = \nu$)? —
- Observation of $0\nu\beta\beta$ would be proof they are
- Impact of IH determination: lack of $0\nu\beta\beta$ implies Dirac nature



Why CP-violation?

- Does e.g. $P(\nu_{\mu} \rightarrow \nu_e) = P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)$?
- Insight into fundamental symmetries of the lepton sector
- Why is the universe not equal parts matter and antimatter?
- Sakharov conditions: Baryon number violation • Out of thermal equilibrium • C and CP violation
- CPV in the Standard Model, eg for K and B mesons, but too small
- “Leptogenesis”: generate asymmetry in neutrinos, transfer to baryons
- Require **neutrino appearance** experiment to discover

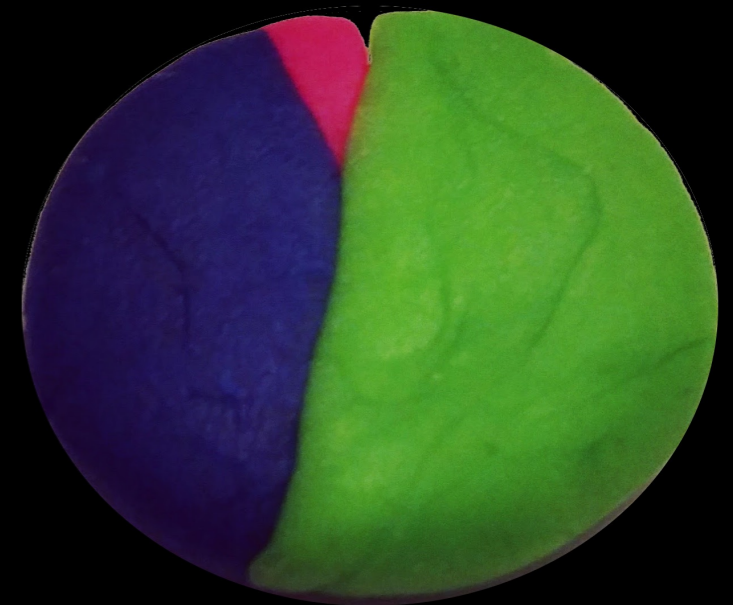
Mixing patterns

- Only a small fraction of ν_e in $|\nu_3\rangle$
(the famous $\text{Sin}^2 2\theta_{13}$)
- The remainder is split about 50/50 ν_μ/ν_τ ($\text{Sin}^2 \theta_{23}$)
- Accident? Or a sign of underlying structure?



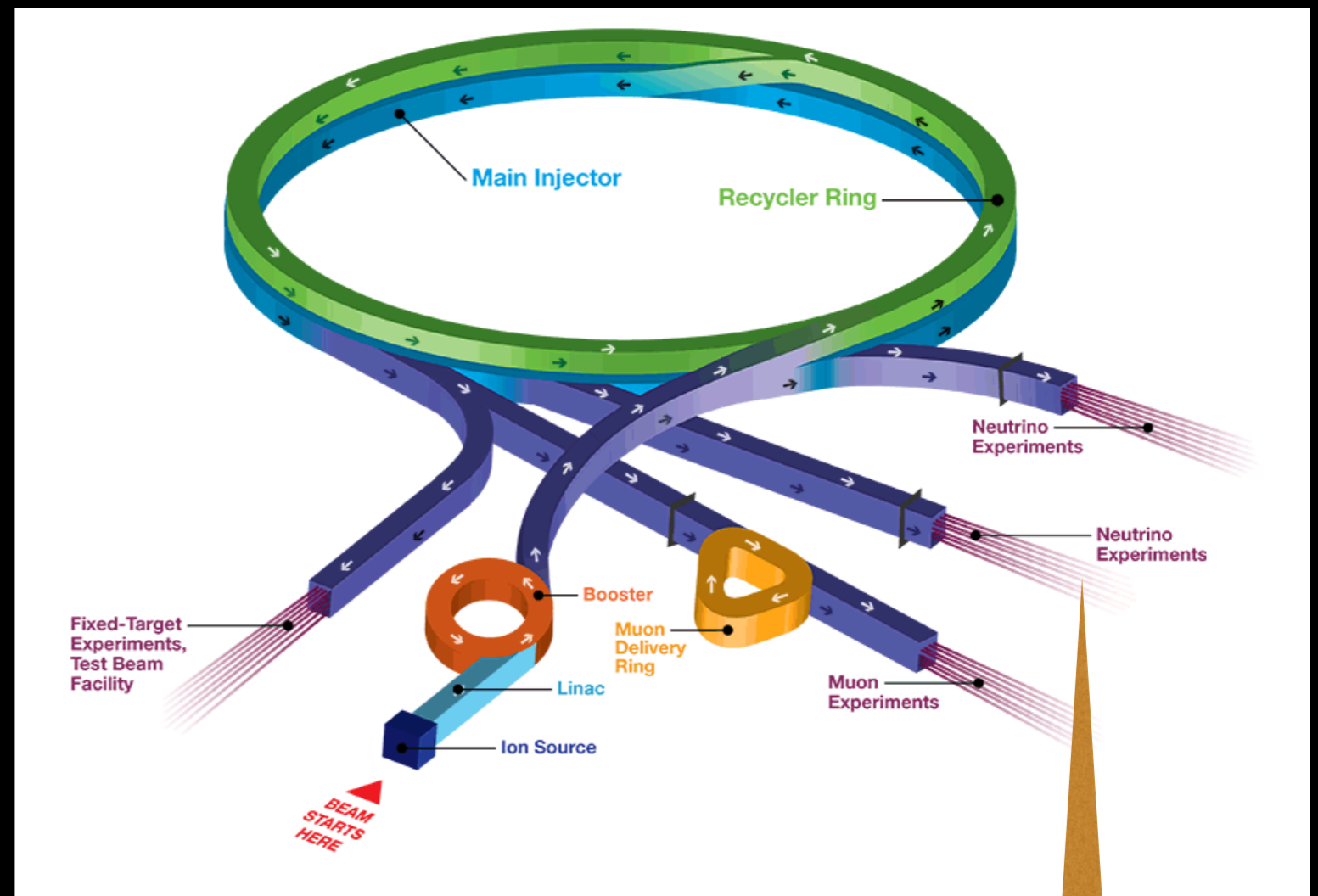
Mixing patterns

- Only a small fraction of ν_e in $|\nu_3\rangle$
(the famous $\text{Sin}^2 2\theta_{13}$)
- The remainder is split about 50/50 ν_μ/ν_τ ($\text{Sin}^2 \theta_{23}$)
- Accident? Or a sign of underlying structure?
- Is θ_{23} exactly 45° ?
- If not, it is
 - $< 45^\circ$ $|\nu_3\rangle$ more ν_τ , like in quarks
 - $> 45^\circ$ $|\nu_3\rangle$ more ν_μ , unlike quarks



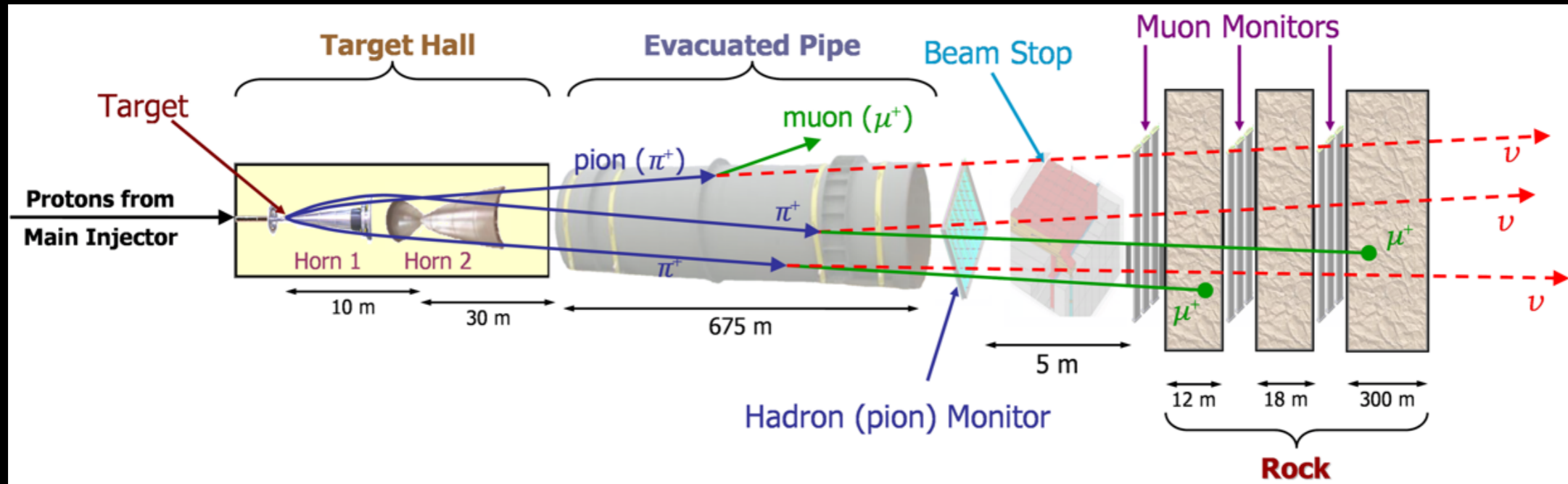
Fermilab accelerator complex

- Neutrinos produced at Main Injector (**NuMI**)
- ➔ Linac 750 keV
- ➔ Booster 400 MeV
- ➔ Recycler 8 GeV
- ➔ NuMI 120 GeV
- ➔ to Carbon target



Line to High Energy Neutrino Experiments

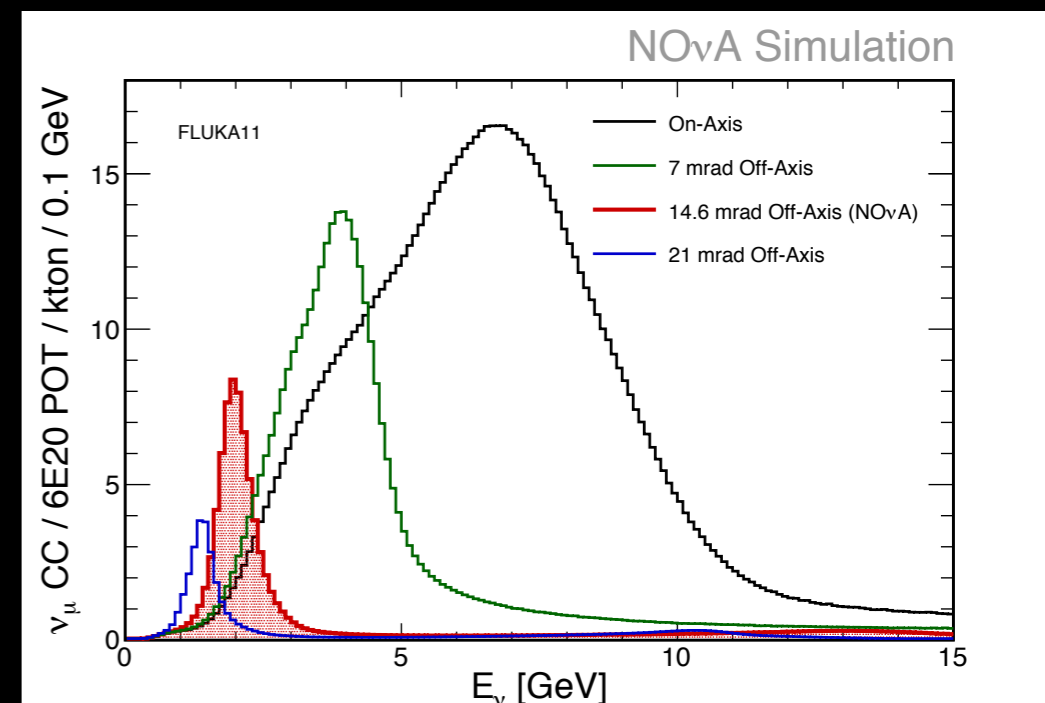
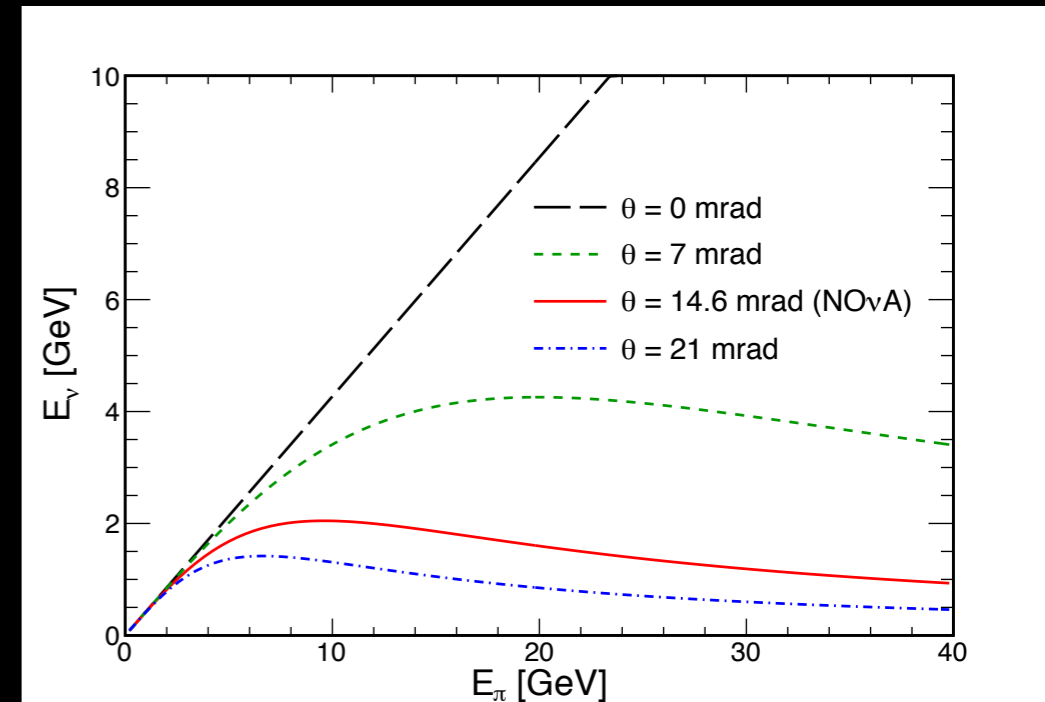
The NuMI beam



- 120 GeV protons from MI to Carbon target
- Produce mainly pions and kaons
- Focused by two magnetic horns
- Allow us to select charge sign for (anti)neutrinos
- Neutrinos produced every 1.3 sec in a spill with 6 doubled bunches $10 \mu\text{s}$ time window
- NuMI designed to provide for NOvA 700kW beam, producing 6×10^{20} POT/year

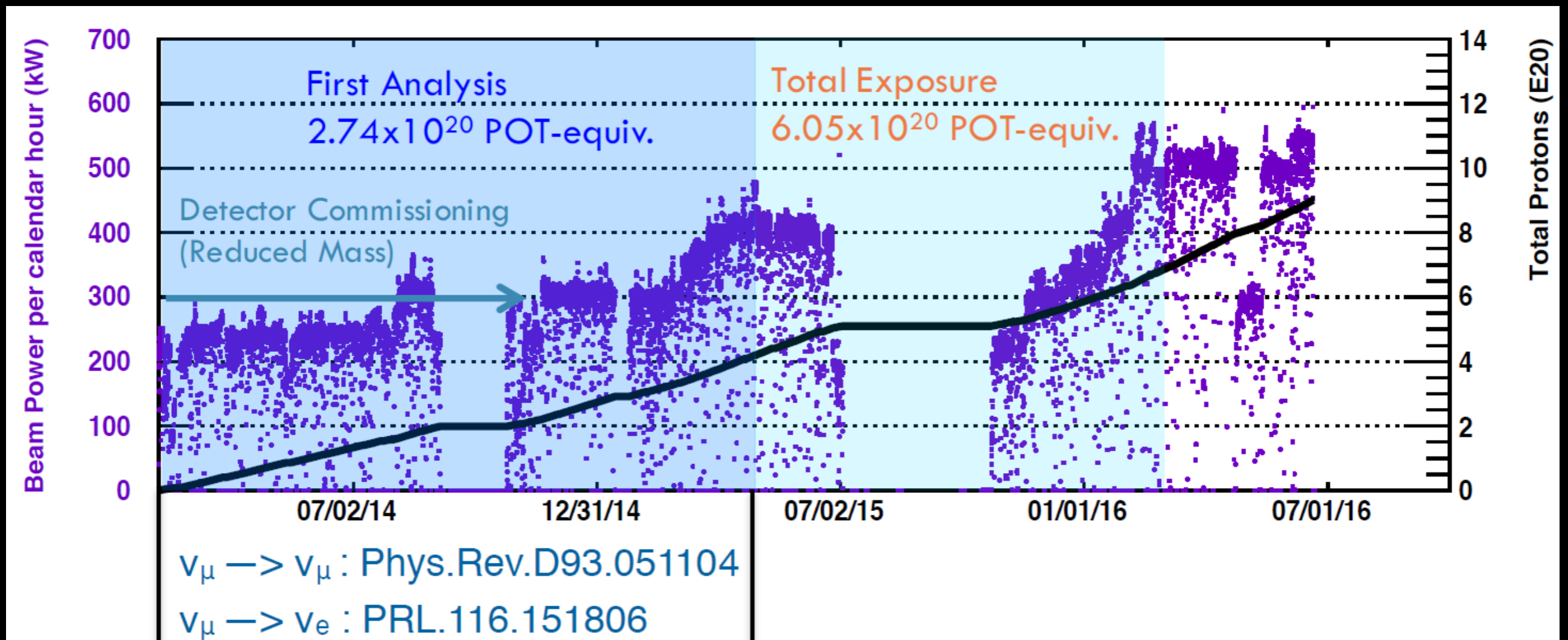
NuMI off-axis beam

- NOvA detectors are sited 14 mrad off the NuMI beam axis
 - With the medium-energy NuMI tune, yields a narrow 2-GeV spectrum at the NOvA detectors
- Reduces NC and ν_e CC backgrounds in the oscillation analyses while maintaining high ν_μ flux at 2 GeV

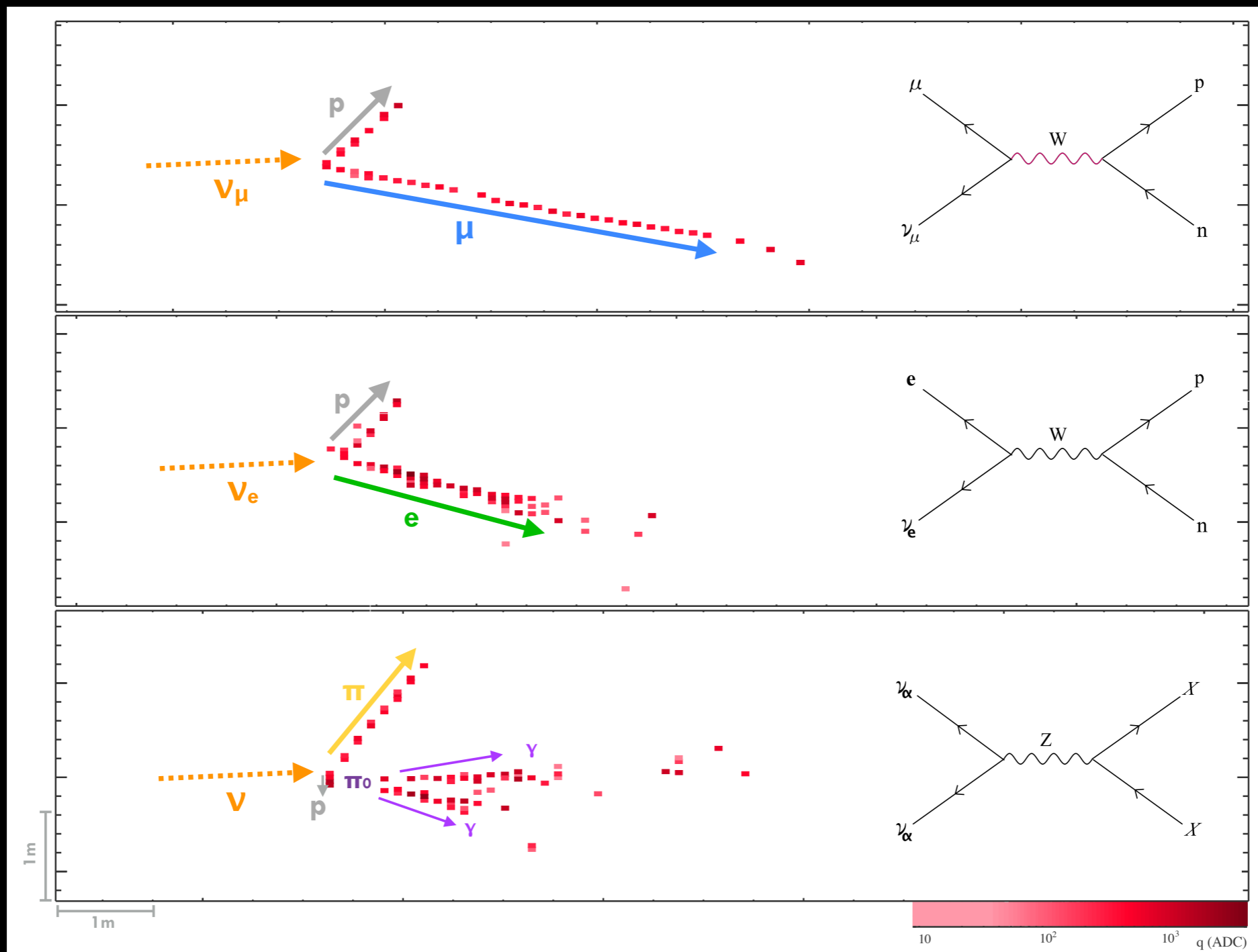


Beam status

- Data from Feb 6, 2014 – May 2, 2016
- Achieved the 700kW design goal
- Have now about $9e20$ POT neutrinos
- Switched to antineutrino mode from Feb 20th, 2017



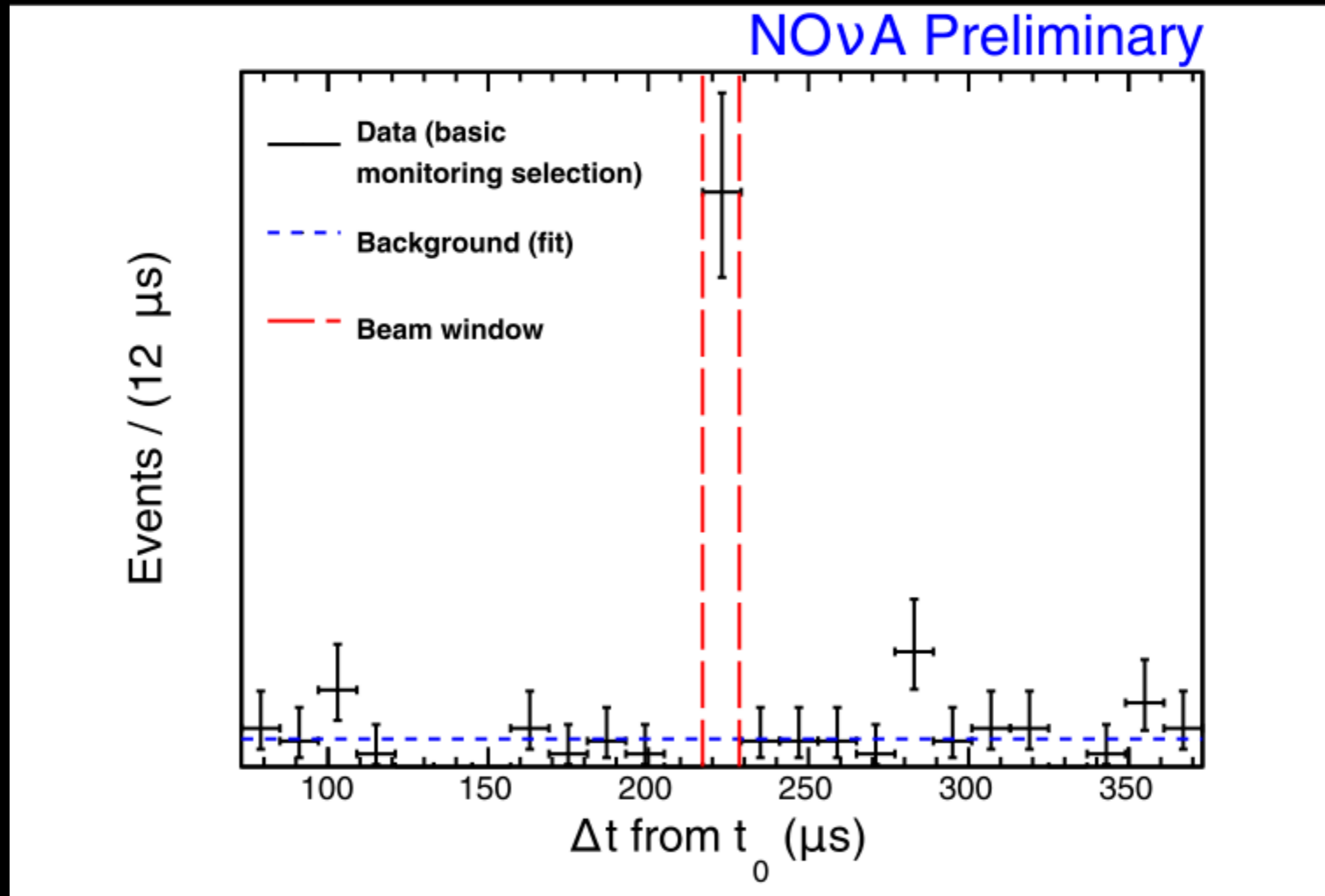
Event topologies



- Very good granularity
- $X_0 = 38$ cm (6 cell depths, 10 cell widths)

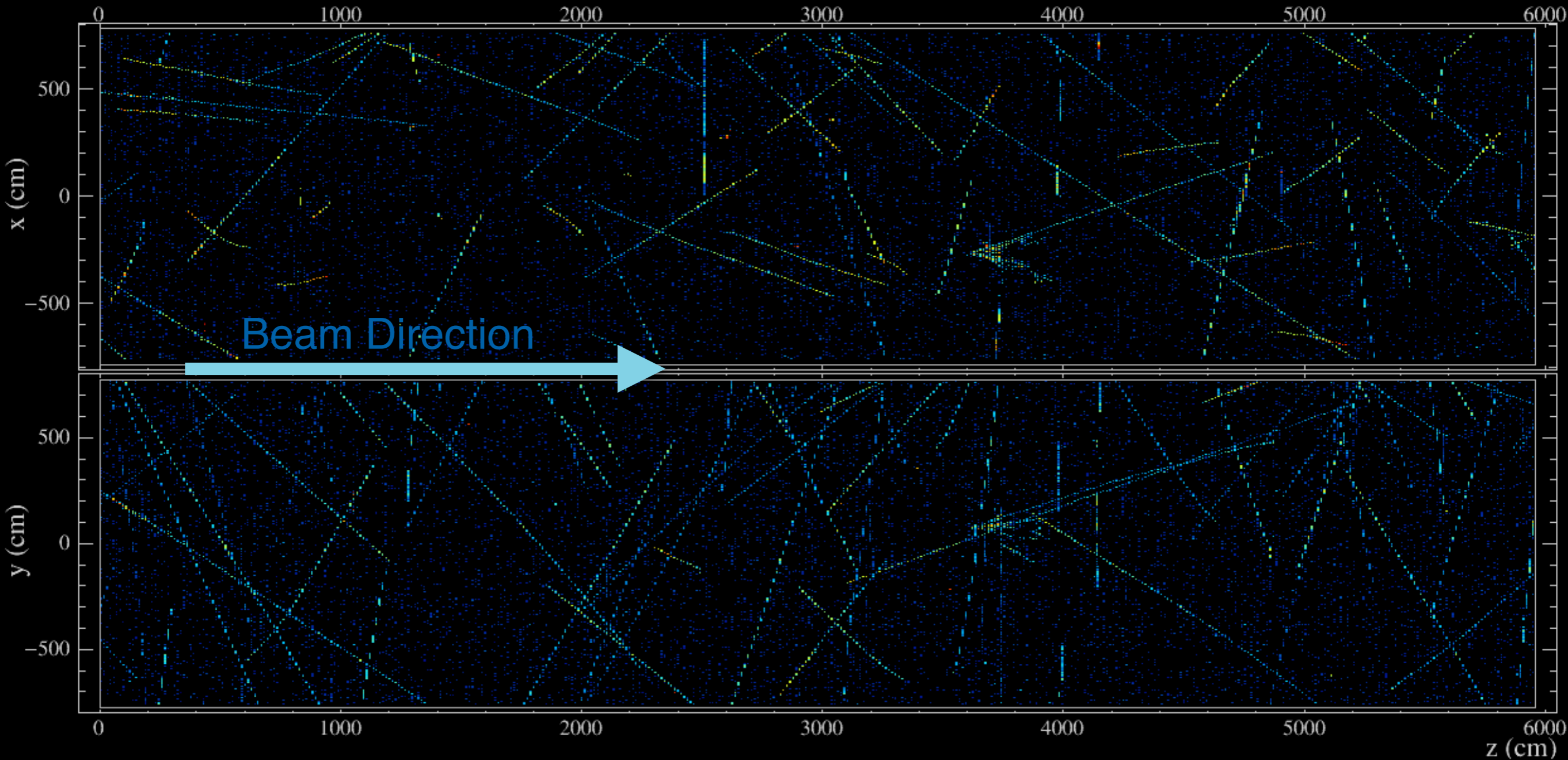
FD Beam Peak

- Trigger structure: 550 μs window, NuMI neutrinos arrive for 10 μs starting at 218 μs



Far Detector 550 μ s Readout Window

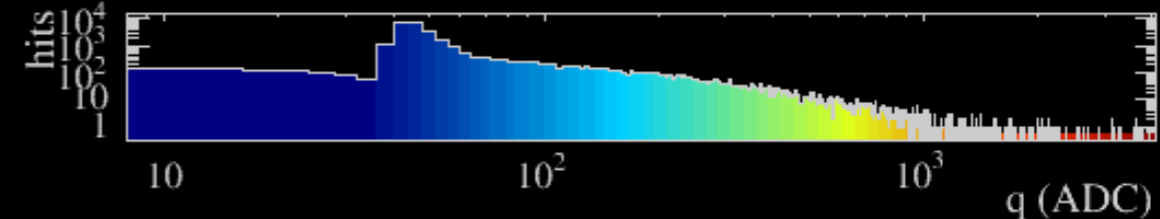
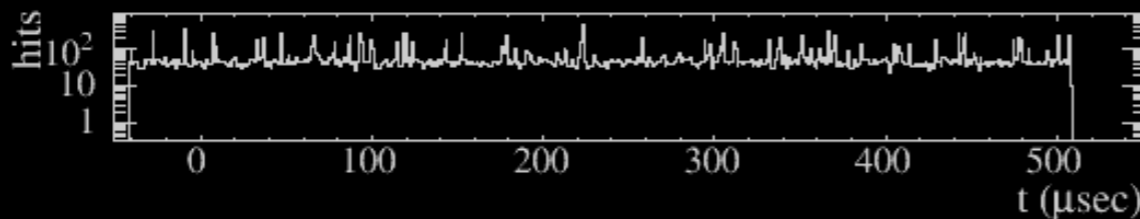
Cell hits colored by charge deposition



NOvA - FNAL E929

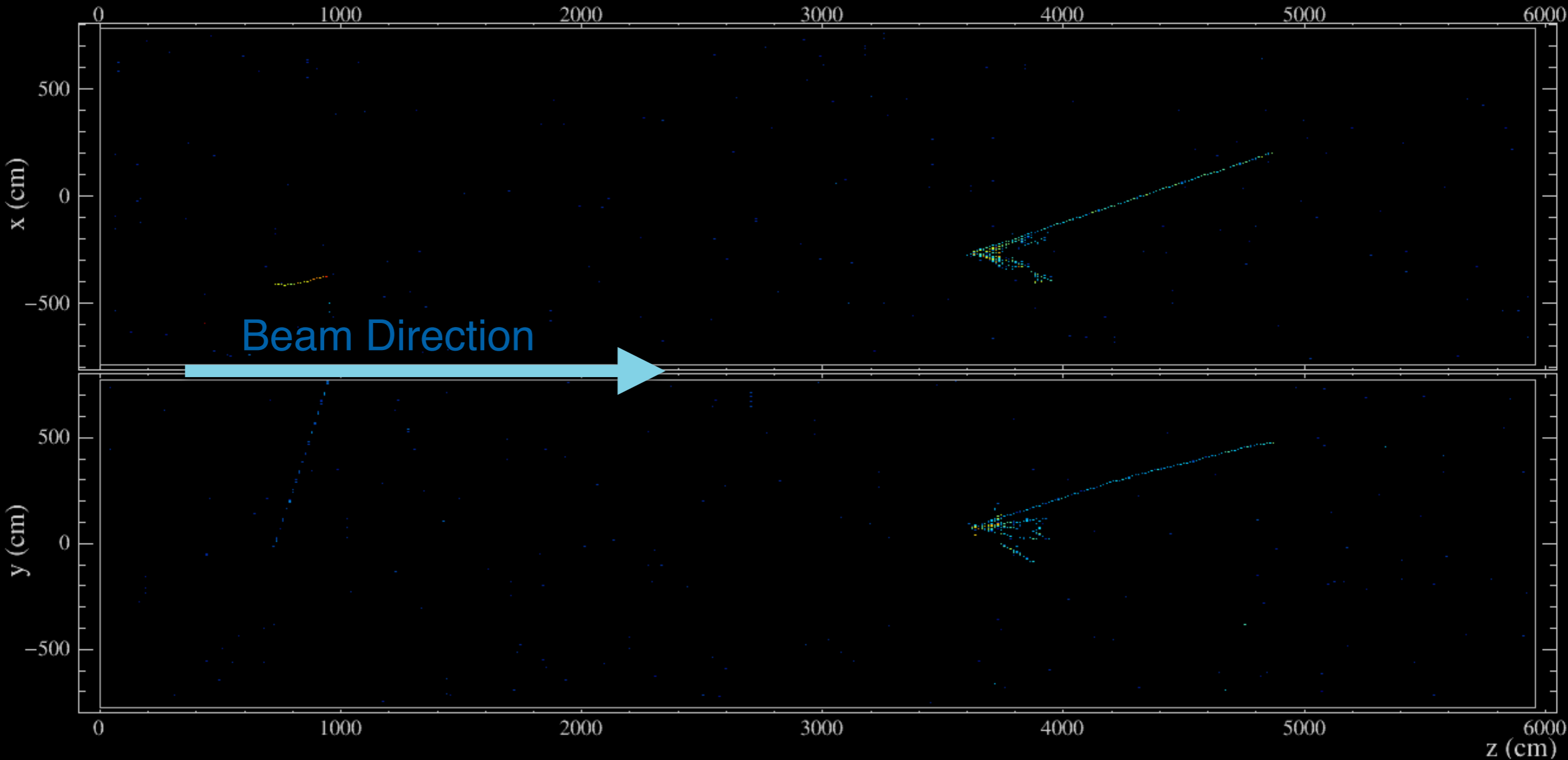
Run: 18620 / 13
Event: 178402 / --

UTC Fri Jan 9, 2015
00:13:53.087341608



Far Detector 10 μs NuMI Beam Window

Cell hits colored by charge deposition



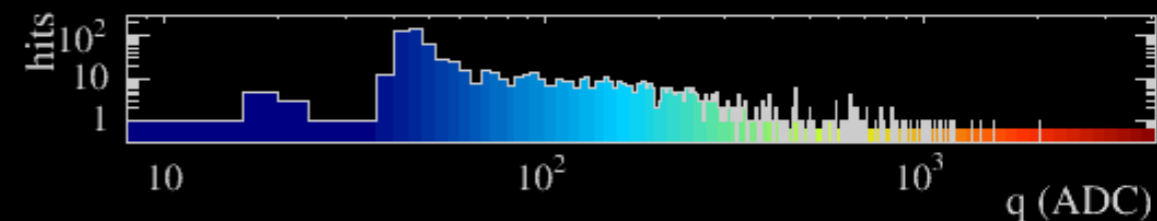
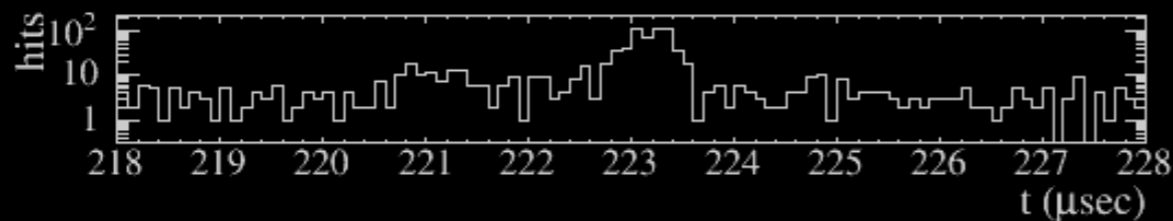
NOvA - FNAL E929

Run: 18620 / 13

Event: 178402 / --

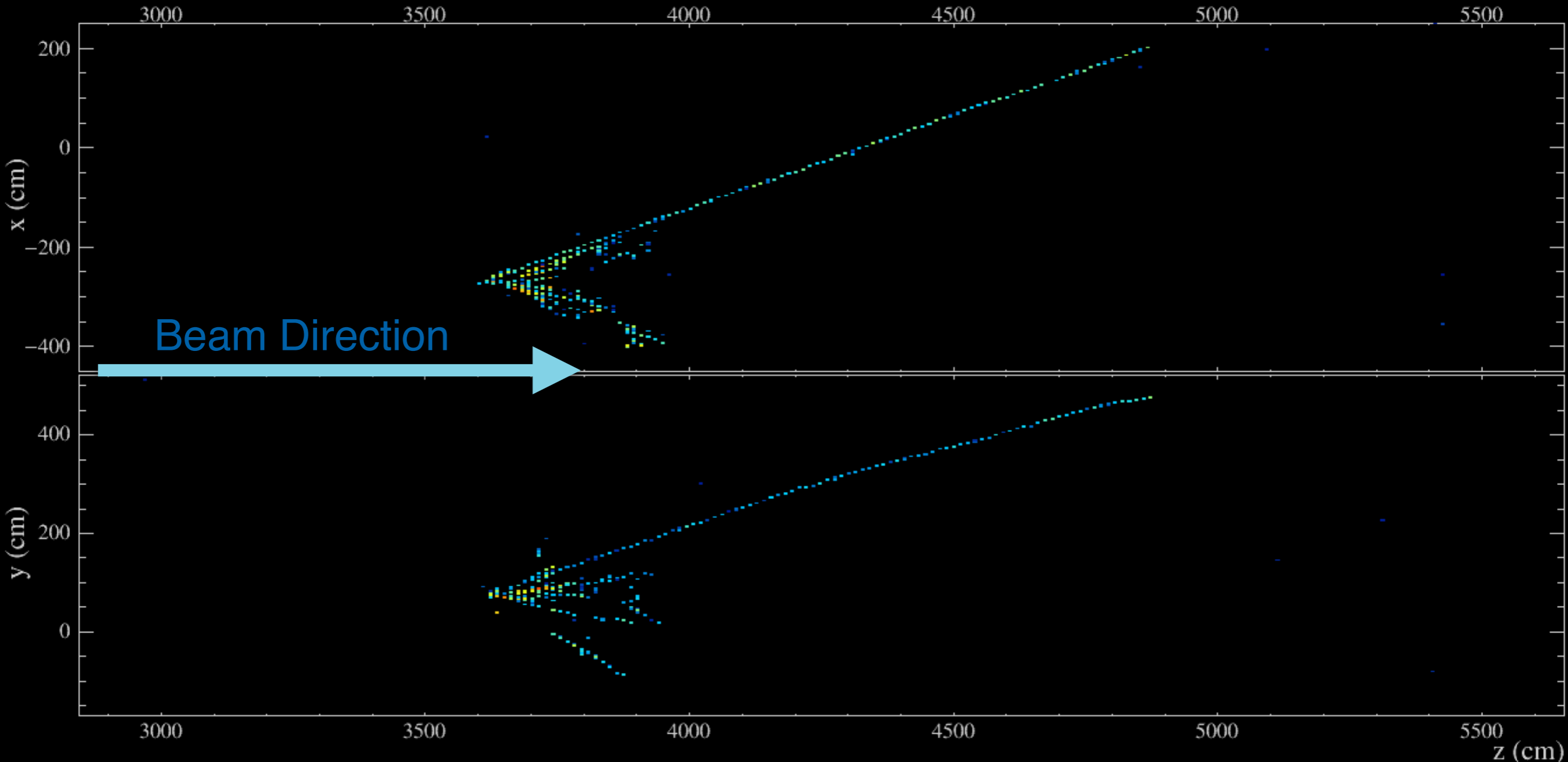
UTC Fri Jan 9, 2015

00:13:53.087341608



Far Detector Neutrino Interaction

Cell hits colored by charge deposition



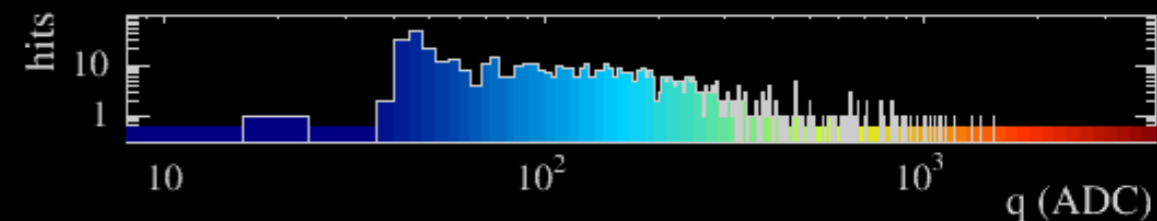
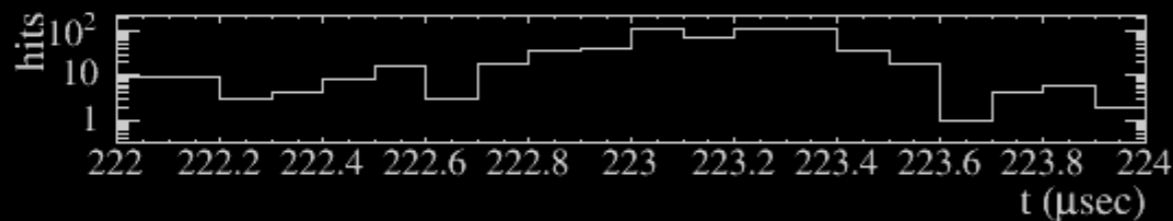
NOvA - FNAL E929

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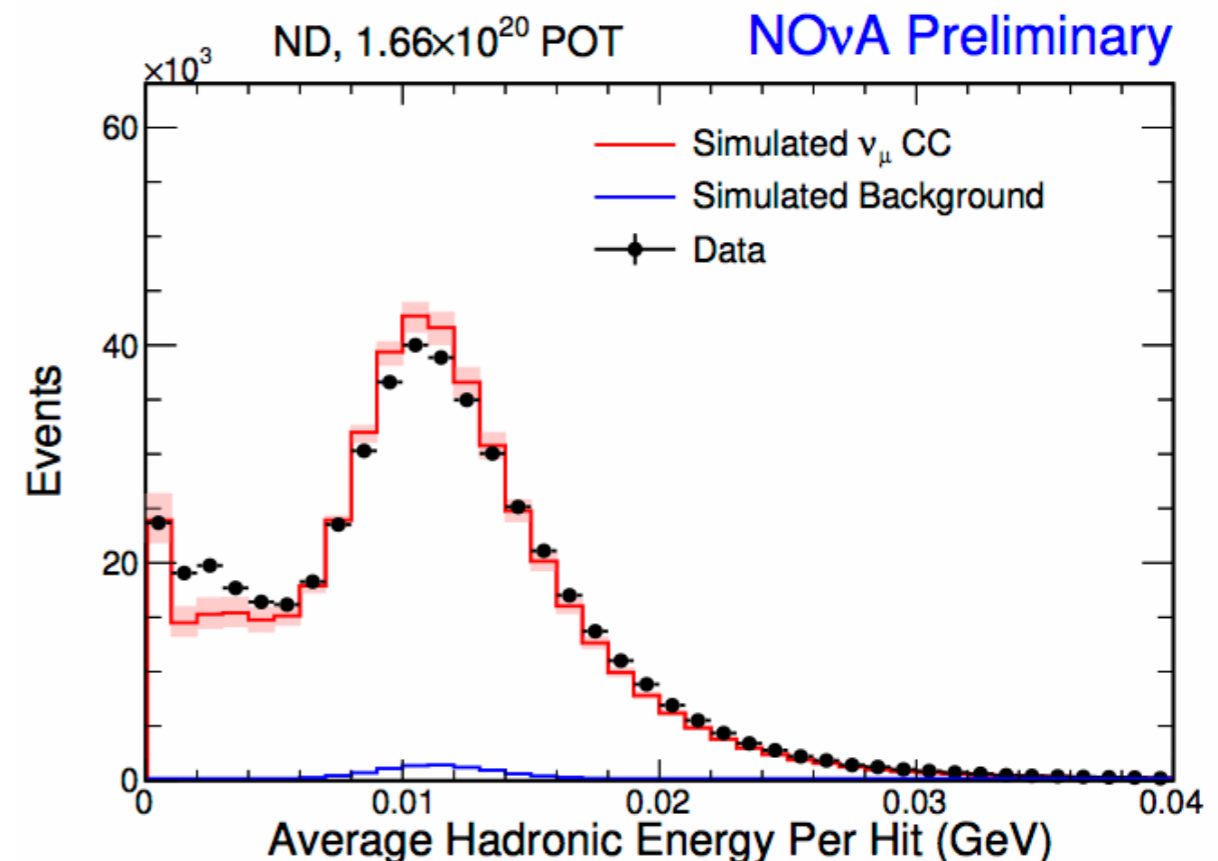
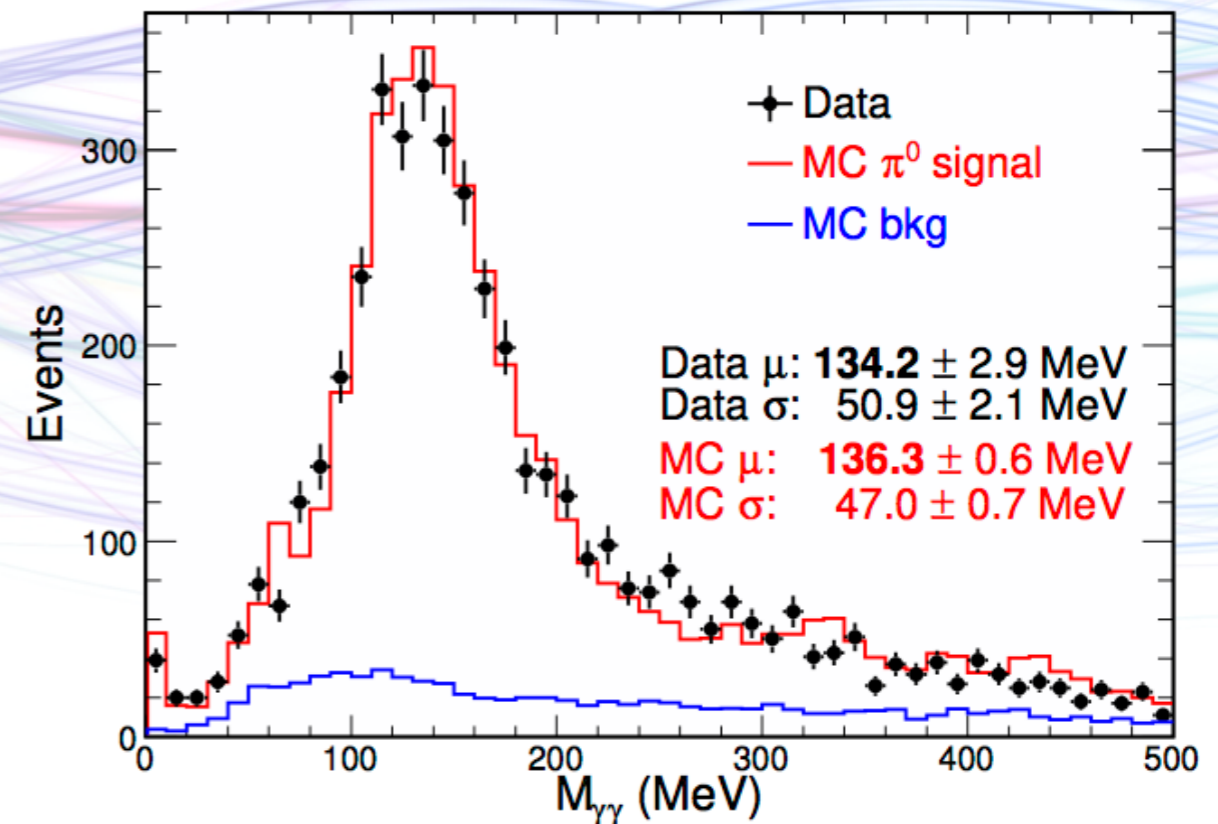
00:13:53.087341608



Numu disappearance

Calibration and energy scale

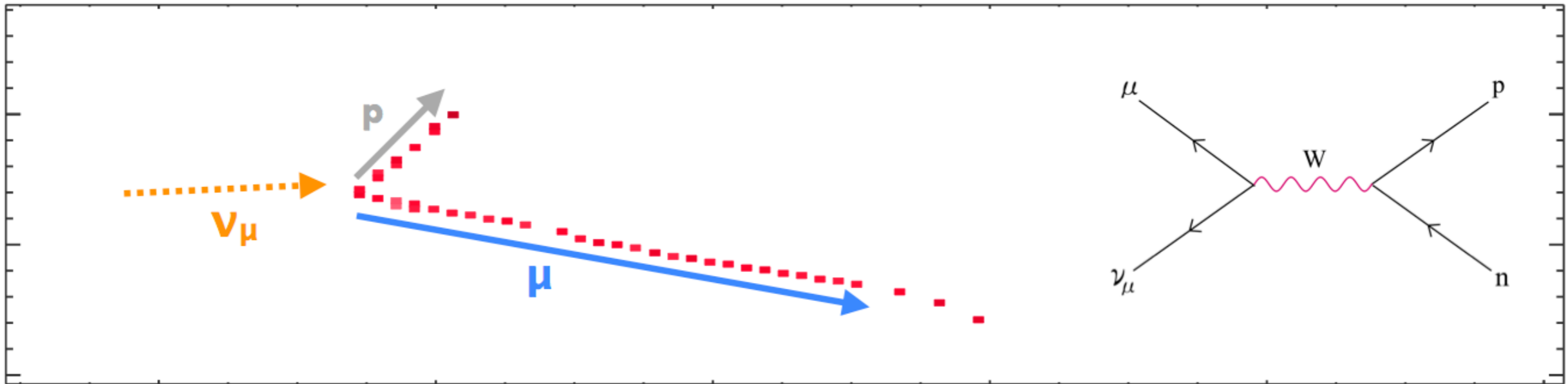
- ▶ Response varies substantially along cell due to light atten.
- ▶ Use cosmic ray muons as a standard candle to calibrate every channel individually
- ▶ Use dE/dx near the end of stopping muon to set abs. scale
- ▶ Multiple calibration x-checks
 - ▶ Beam muon dE/dx
 - ▶ Michel energy spectrum
 - ▶ π^0 mass peak
 - ▶ Hadronic energy/hit
- ▶ Take 5% abs. and rel. errors on energy scale



Principle of the ν_μ measurement



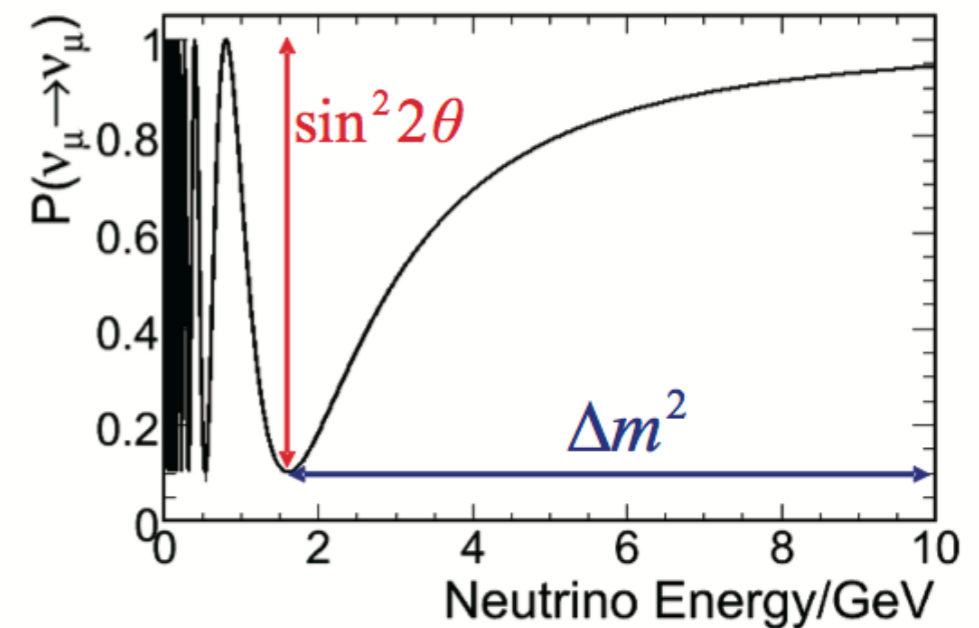
- ▶ Separate ν_μ CC interactions from backgrounds
 - ▶ Long muon track with distinctive dE/dx easy to spot
- ▶ Extrapolate observed ND spectrum to make FD unosc. prediction
- ▶ Measure shape of ν_μ deficit in the FD



Principle of the ν_μ measurement

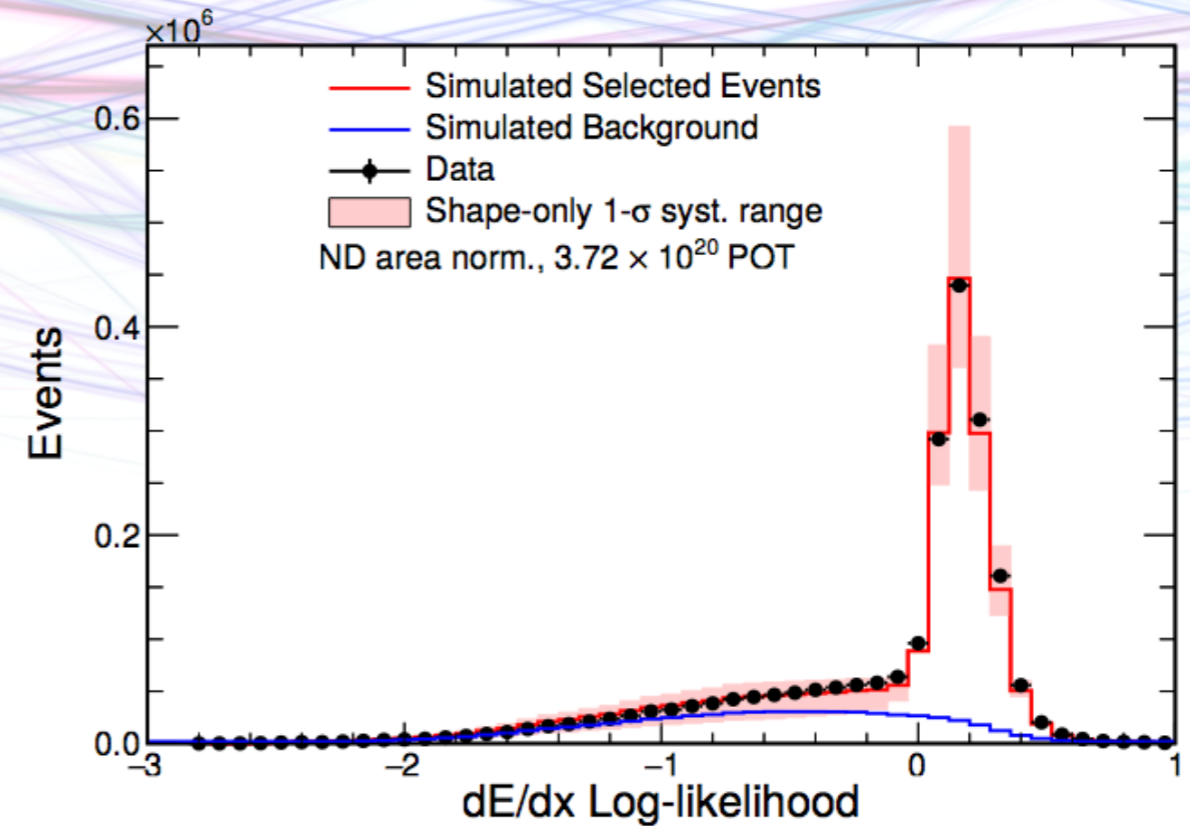


- ▶ Separate ν_μ CC interactions from backgrounds
 - ▶ Long muon track with distinctive dE/dx easy to spot
- ▶ Extrapolate observed ND spectrum to make FD unosc. prediction
- ▶ Measure shape of ν_μ deficit in the FD
- ▶ Two flavor approx. works well here
- ▶ $P_{\mu\mu} \approx 1 - \sin^2 2\theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right)$
- ▶ $\theta_{23} \approx 45^\circ \rightarrow$ almost all ν_μ expected to disappear at oscillation max.

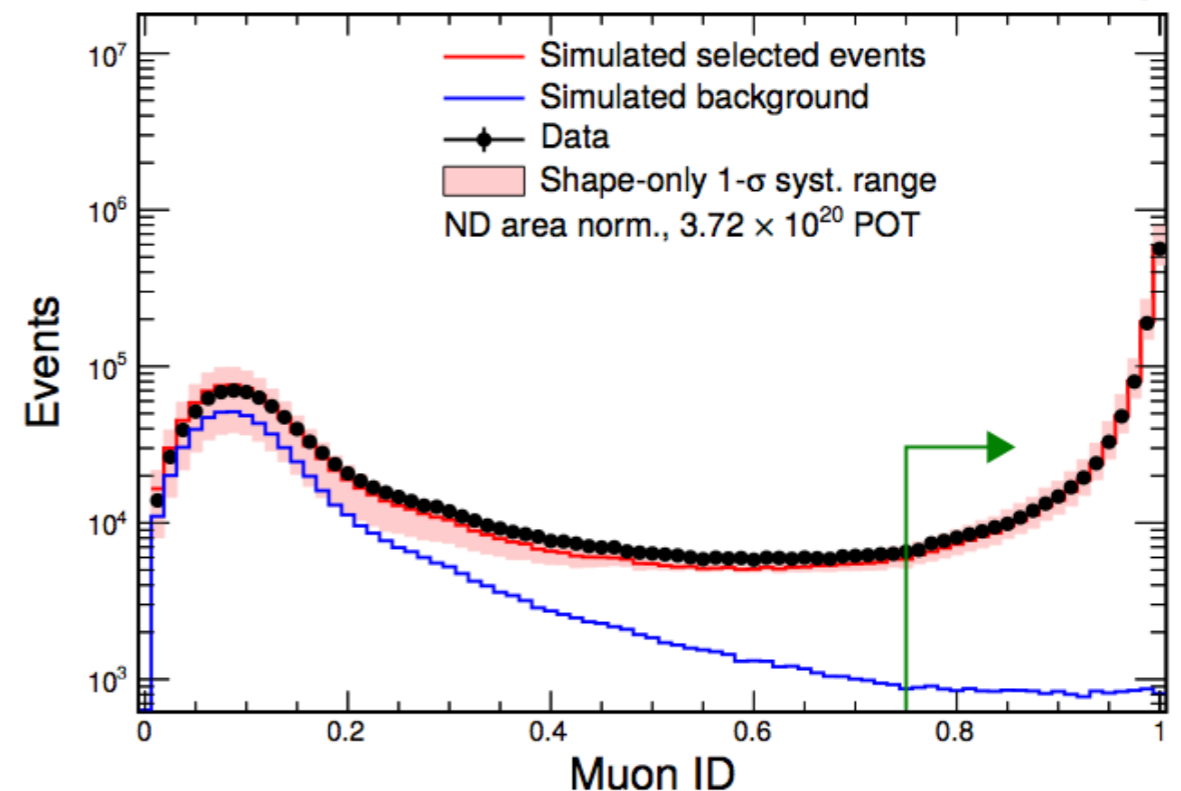


Selecting muon neutrinos

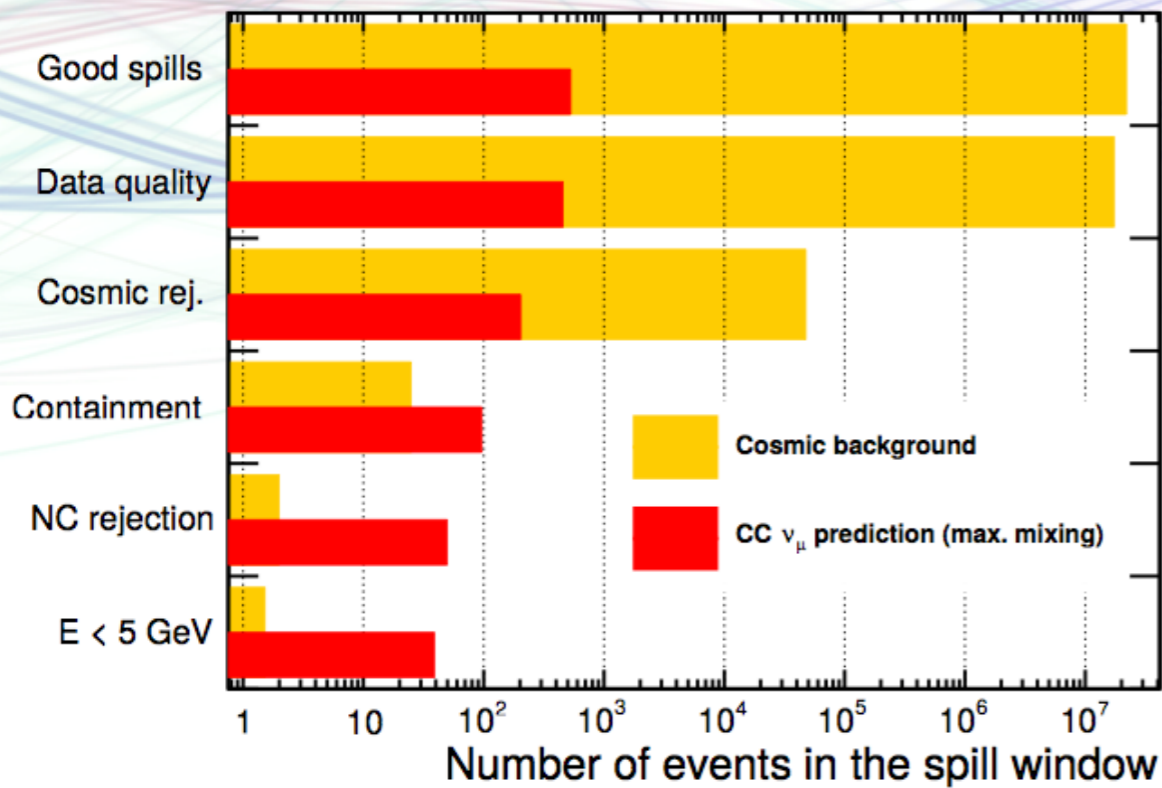
- ▶ Basic containment cuts requiring no activity close to detector walls
- ▶ kNN-based ν_μ classifier using 4 inputs
 - ▶ Track length
 - ▶ dE/dx
 - ▶ Scattering
 - ▶ Fraction of planes that have track-only
- ▶ Selection 81% efficient for ν_μ signal, 95% pure



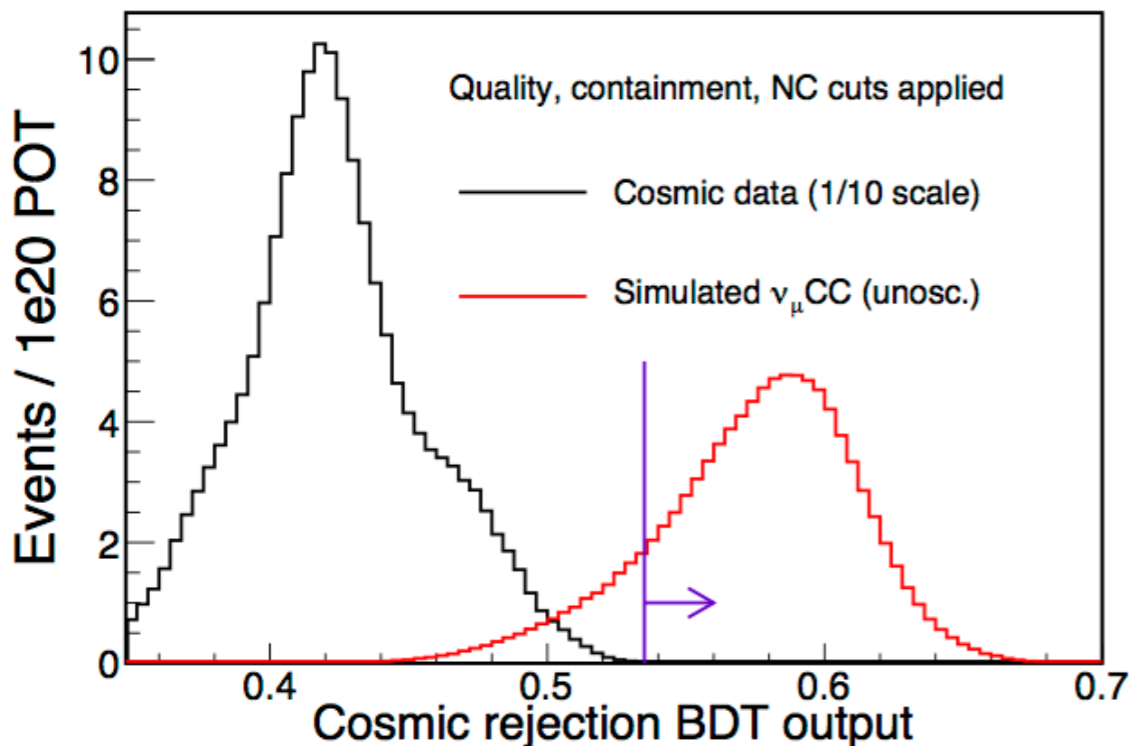
NOvA Preliminary



Cosmic rejection for ν_μ analysis

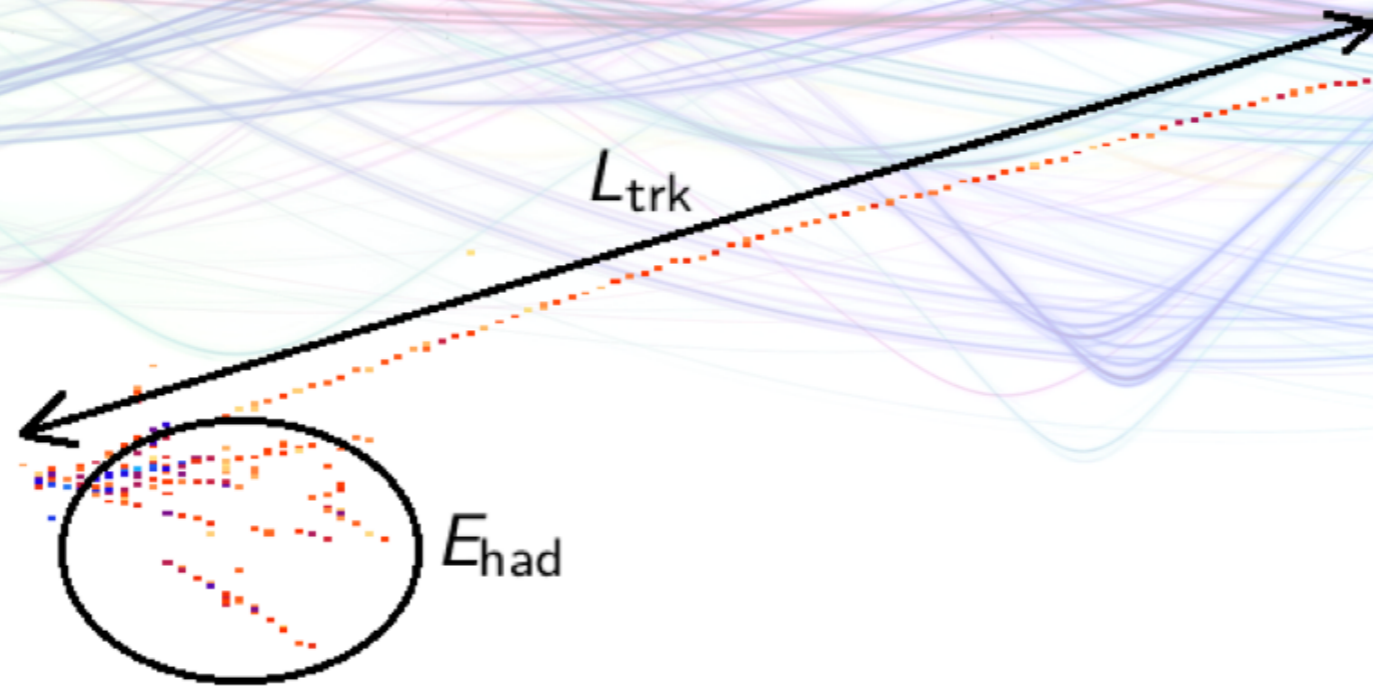


NOvA Preliminary



- ▶ $10\mu\text{s}$ spill window at $\sim 1\text{Hz}$ gives 10^5 rejection
- ▶ Cosmic background rate measured from data adjacent in time to the beam spill window
- ▶ Additional factor 10^7 from event topology plus boosted decision tree based on
 - ▶ Track direction
 - ▶ Track start and end points
 - ▶ Track length
 - ▶ Energy
 - ▶ Number of hits

Muon neutrino energy reconstruction

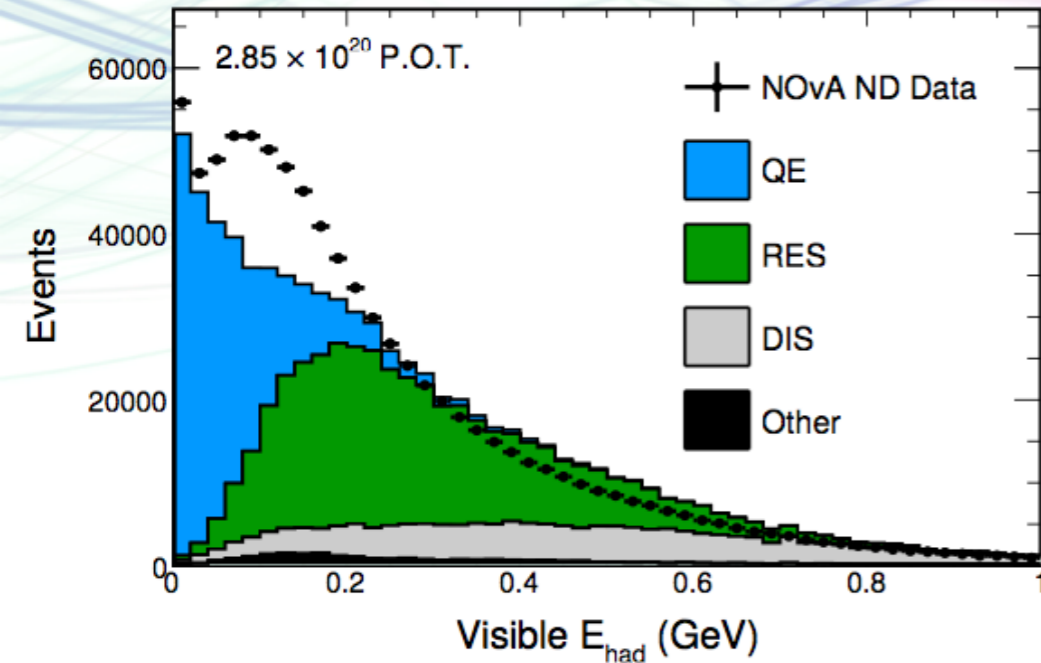


- ▶ Estimate energy of selected events to trace out oscillation structure
- ▶ Known muon $dE/dx \rightarrow E_\mu = f(L_{\text{trk}}) \sim k \times L_{\text{trk}}$
- ▶ Hadronic part of the event estimated calorimetrically

- ▶ $E_\nu = f(L_{\text{trk}}) + E_{\text{had}}$
- ▶ Achieve 7% energy resolution

Nuclear correlations

NOvA Preliminary



- ▶ ND hadE distributions suggest extra process between QE and Δ prod.
- ▶ MINERvA report similar excess in their data¹

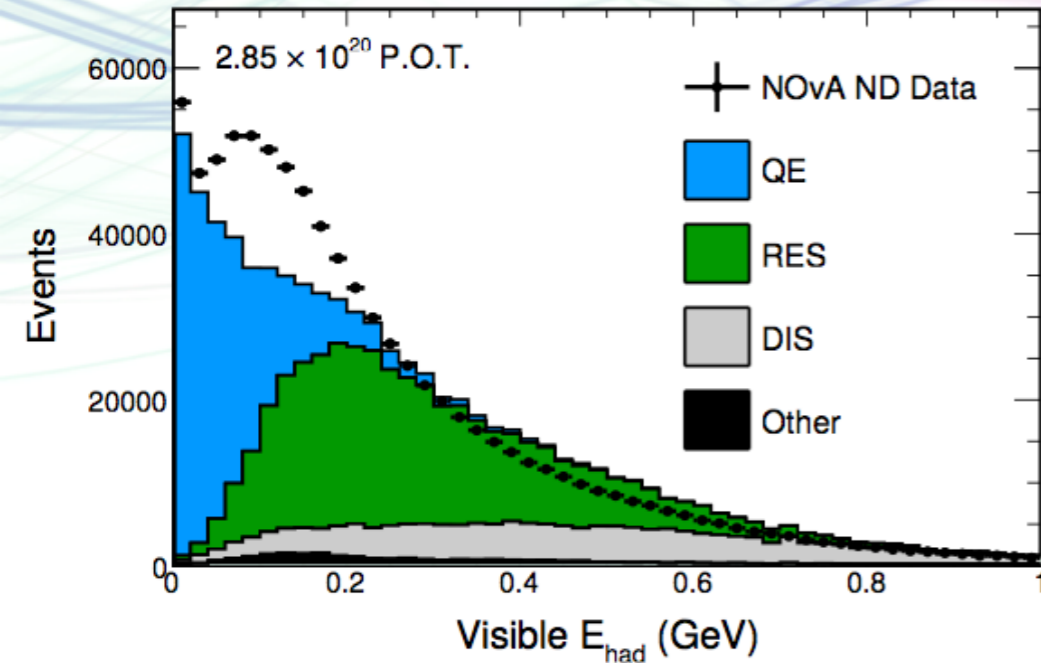
¹P.A. Rodrigues *et al.*, PRL 116 (2016) 071802 (arXiv:1511.05944)

²S. Dytman, based on J. W. Lightbody, J. S. OConnell, Comp. in Phys. 2 (1988) 57

³P.A. Rodrigues *et al.*, arXiv:1601.01888

Nuclear correlations

NOvA Preliminary



1. Leptonic model
(Dytman model)

ν

μ

2. Hadronic model
(Nucleon cluster model)

3. FSI model
(hA model)

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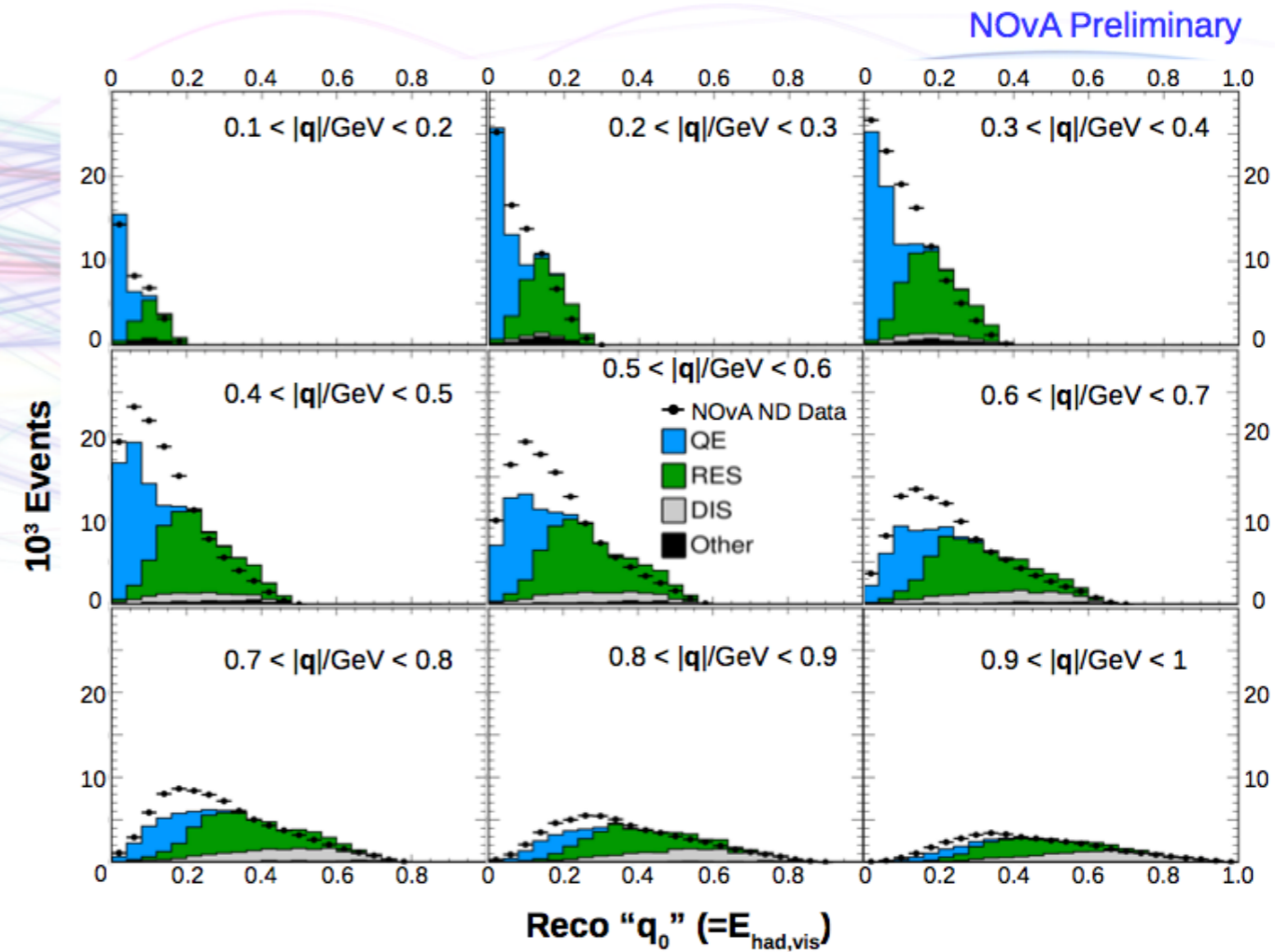
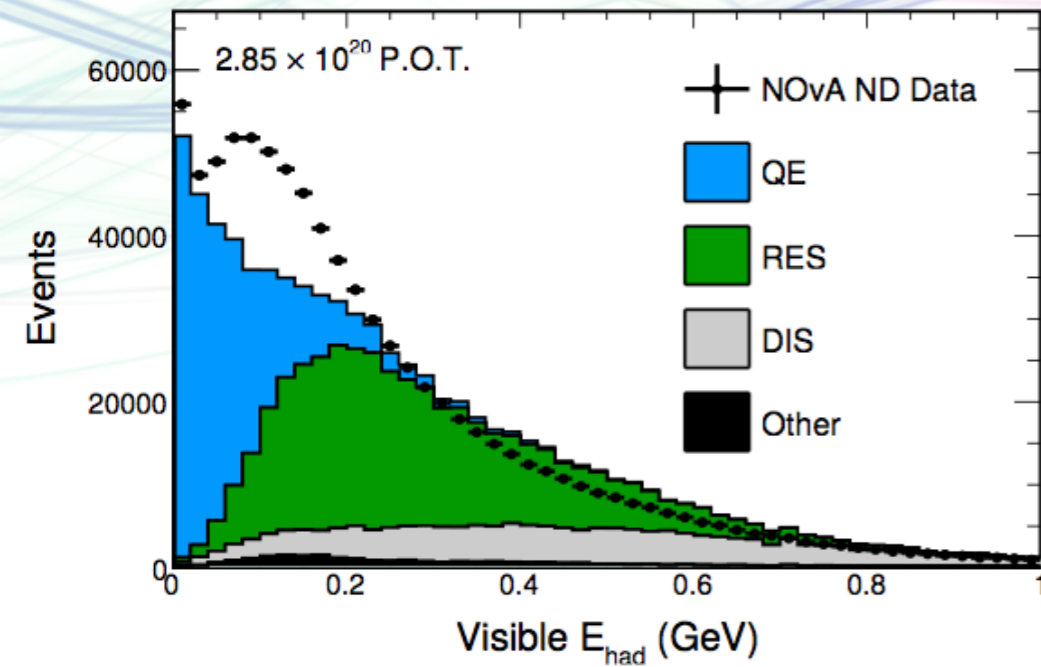
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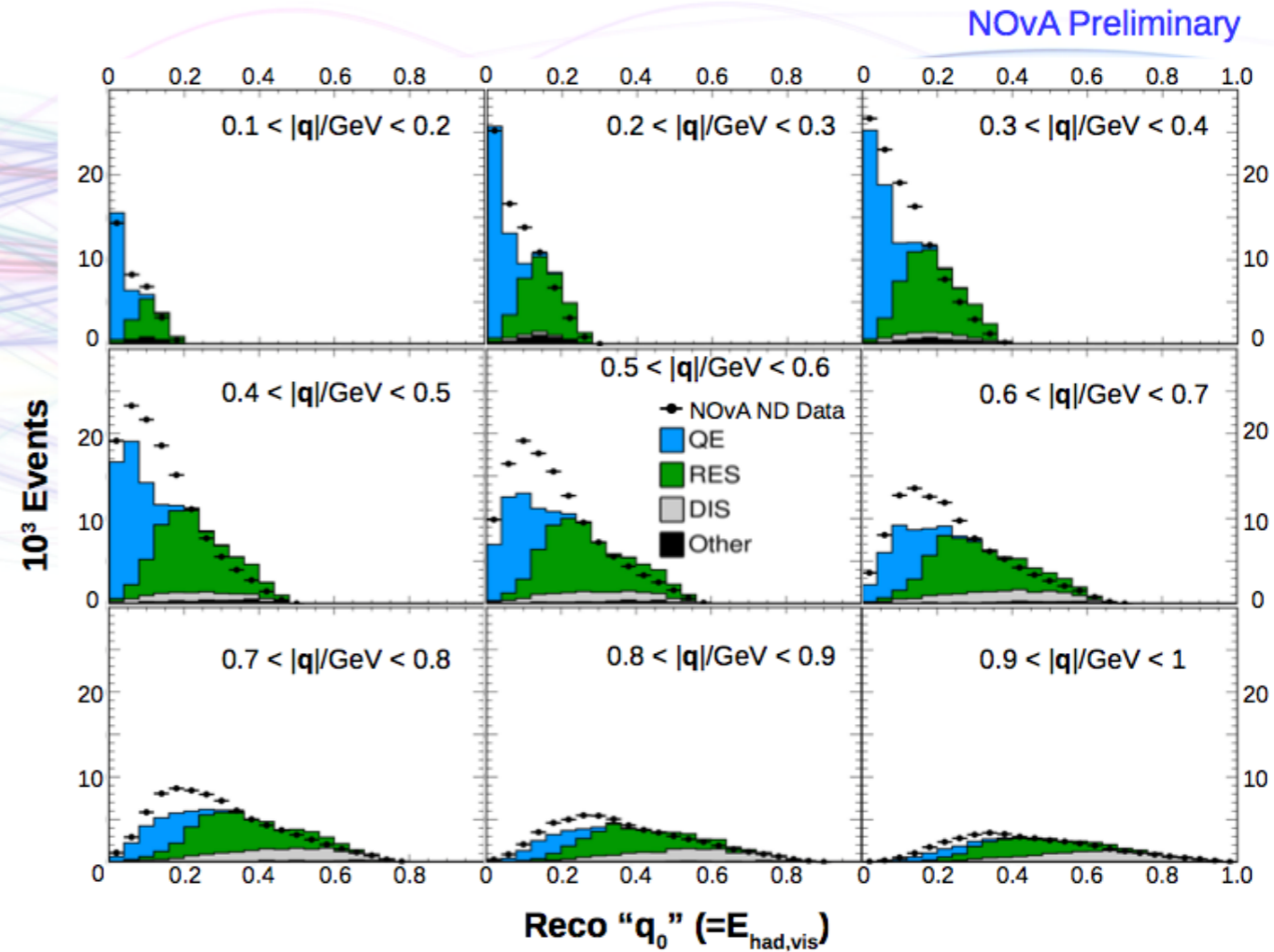
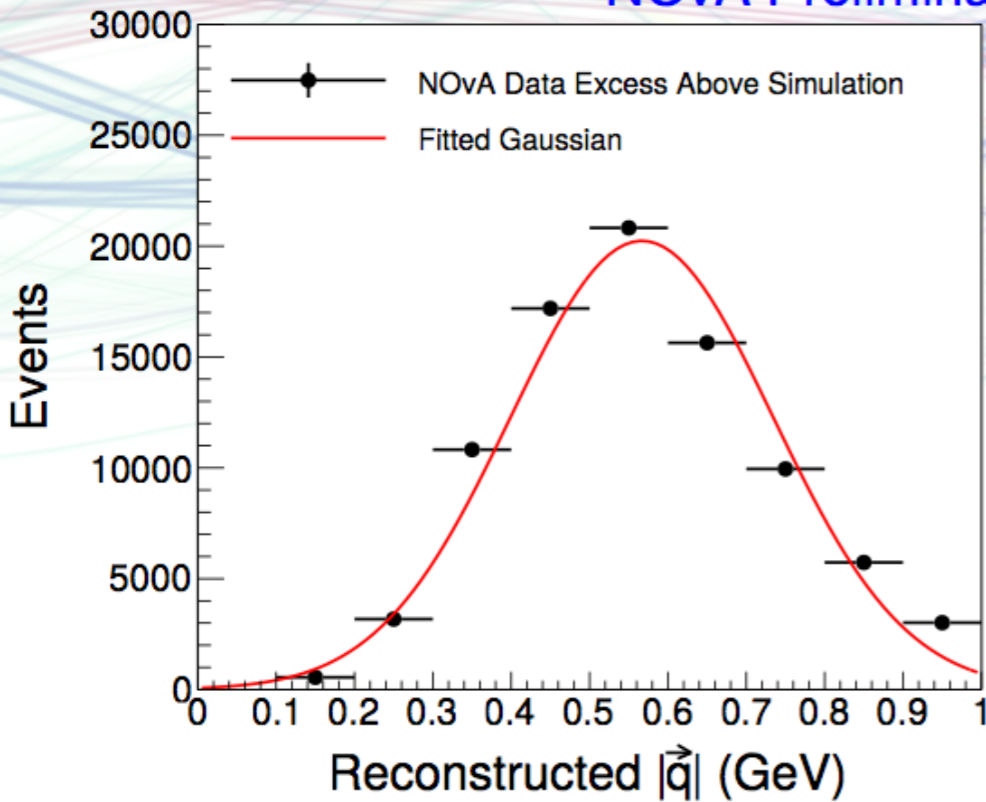
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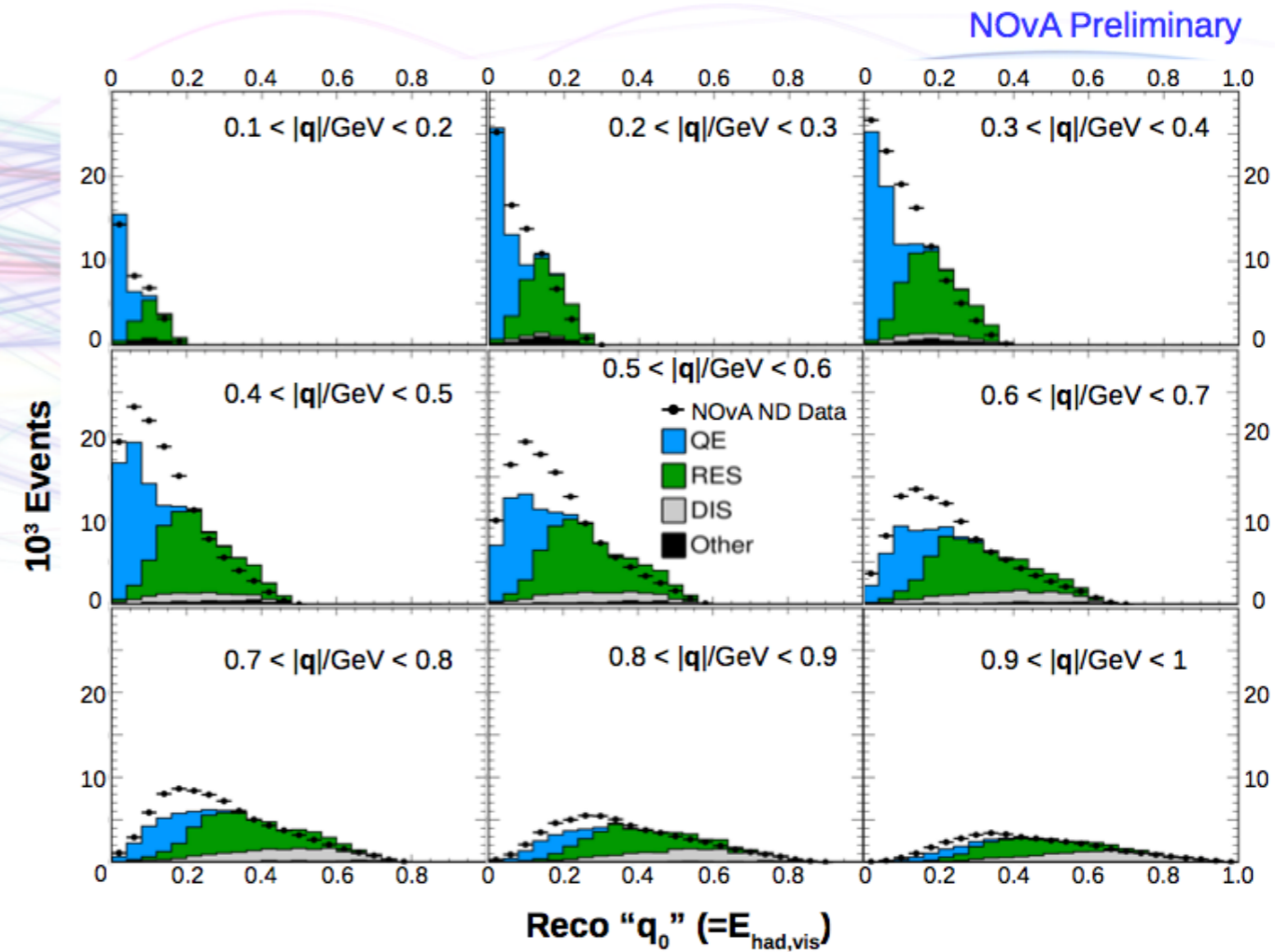
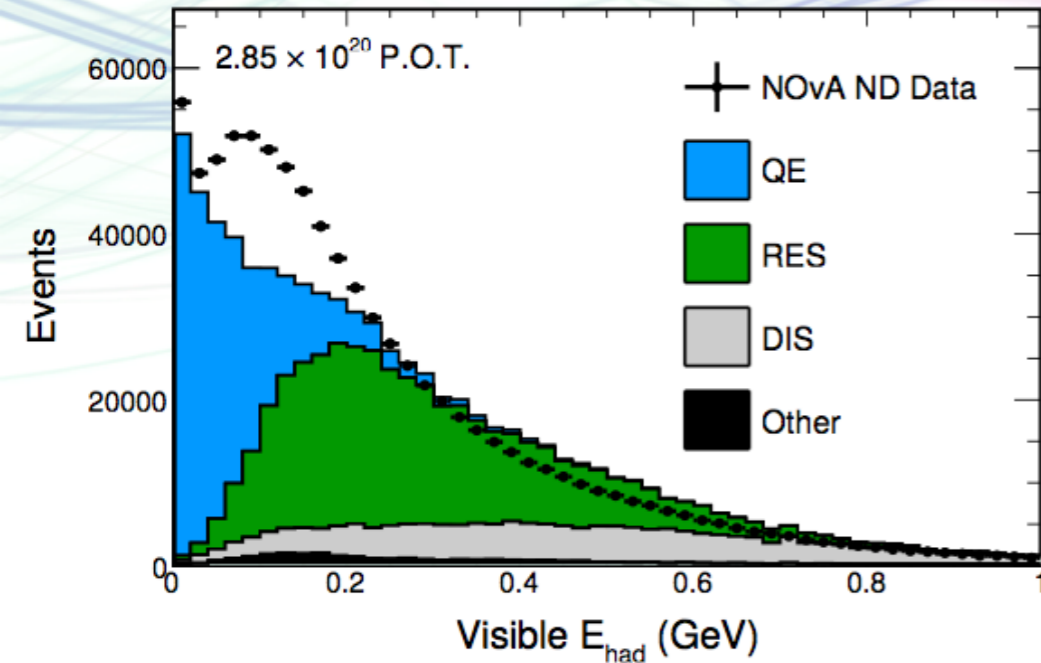
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- ▶ Reweight to match observed excess as a function of \vec{p} transfer
- ▶ Also reduce single non-resonant pion production by 50%³

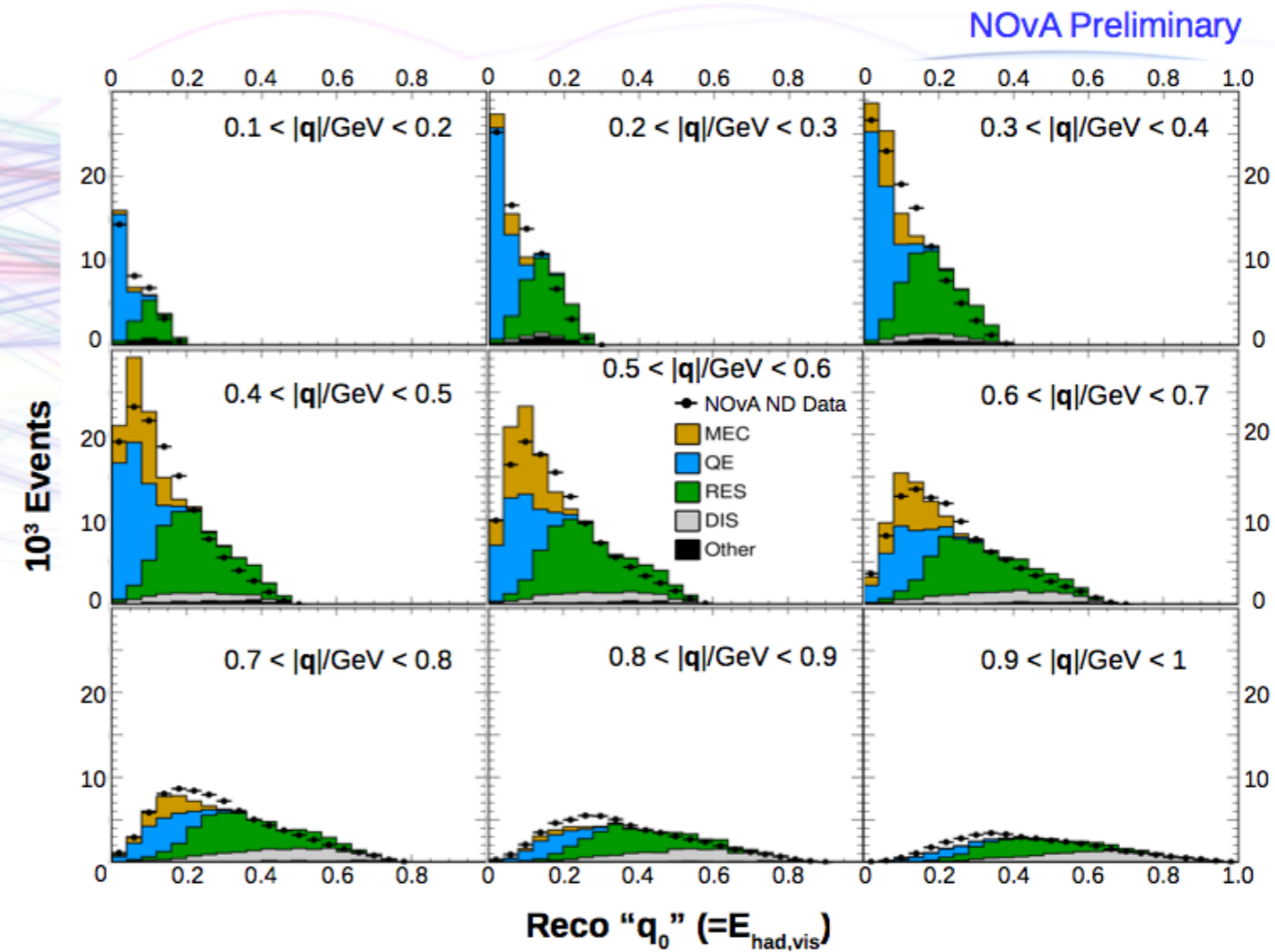
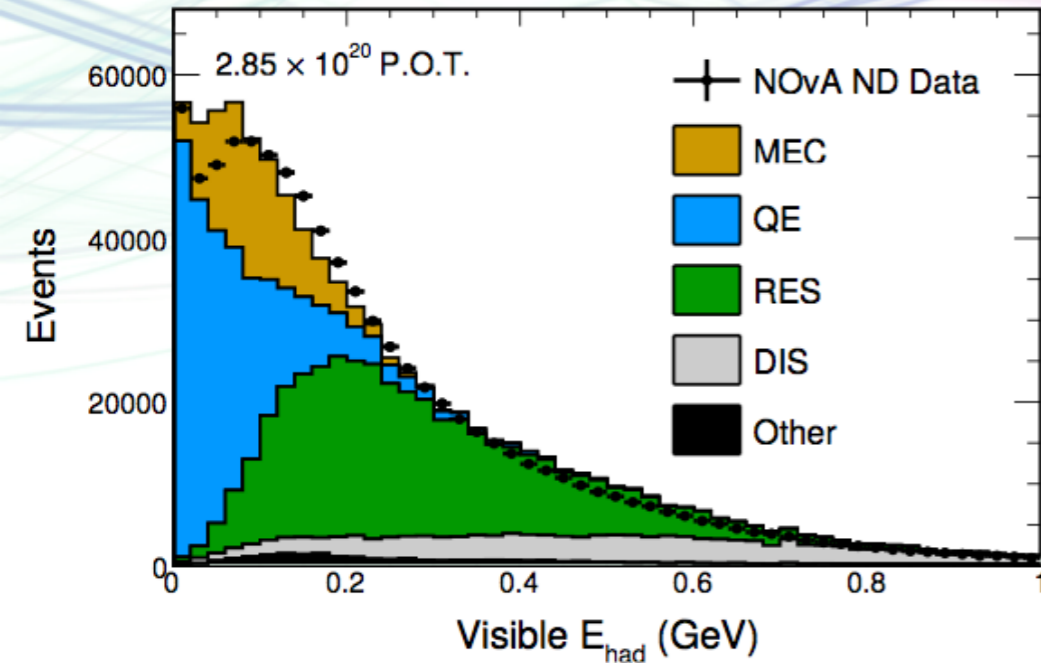
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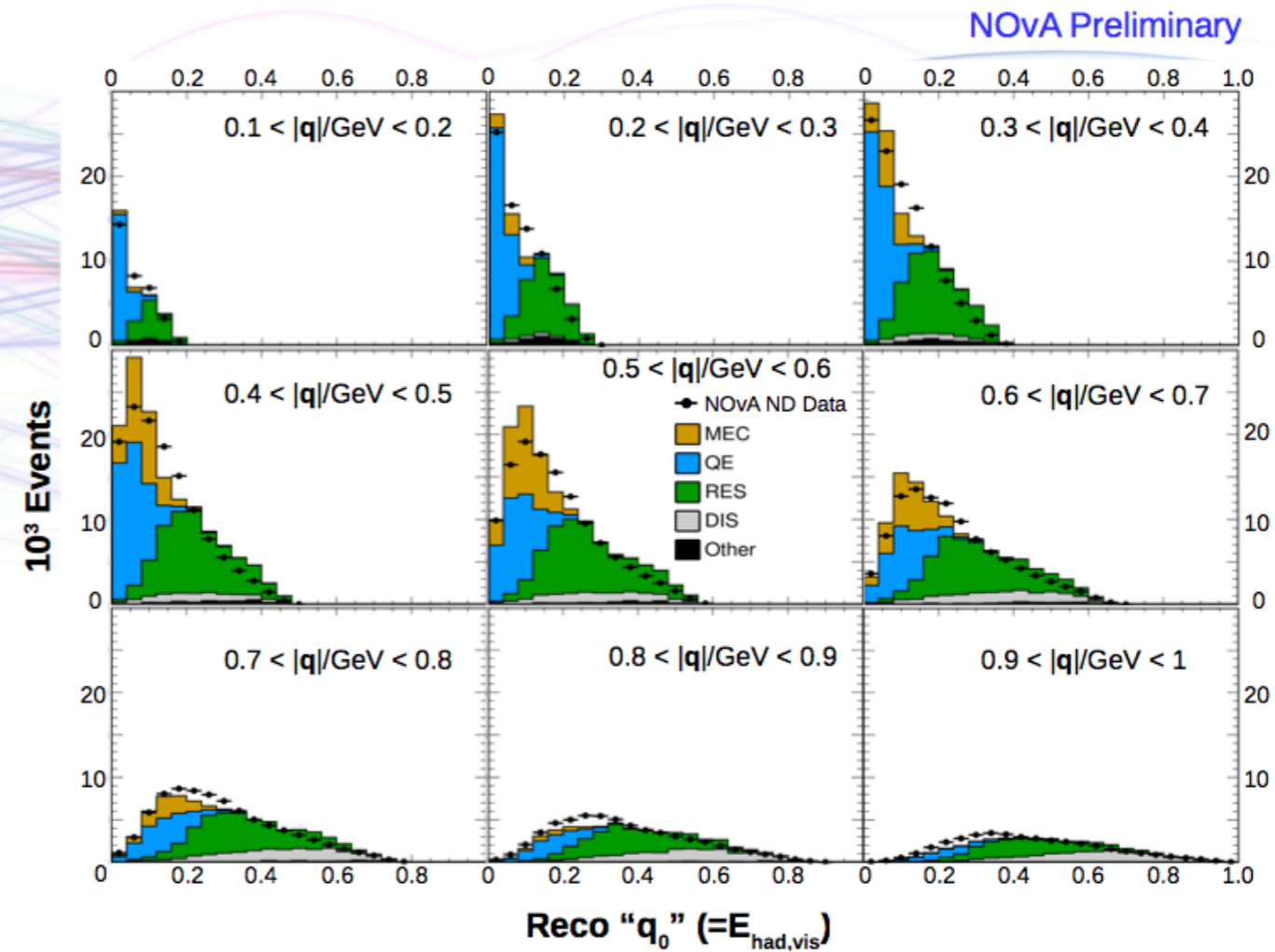
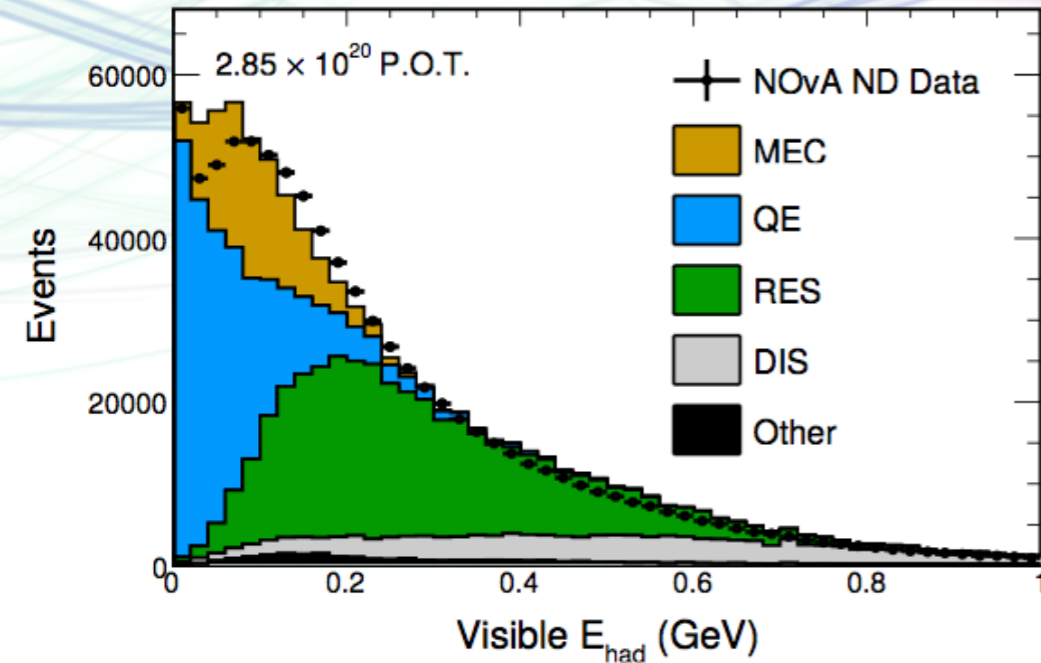
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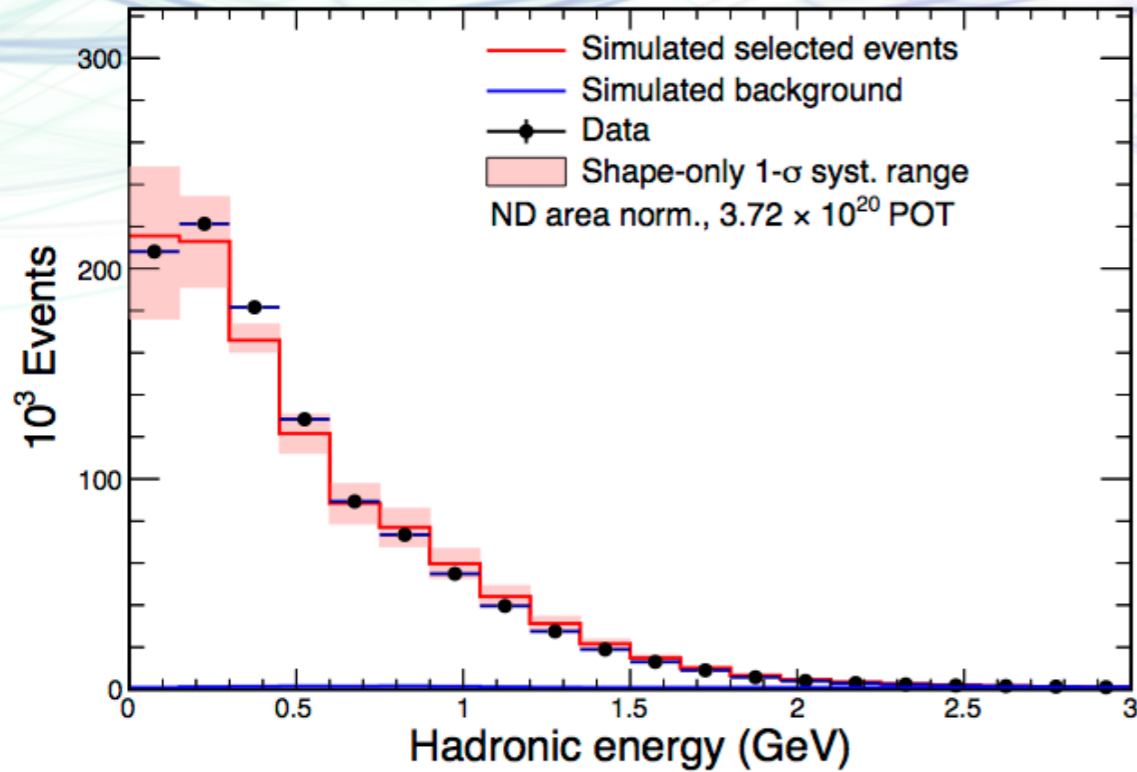
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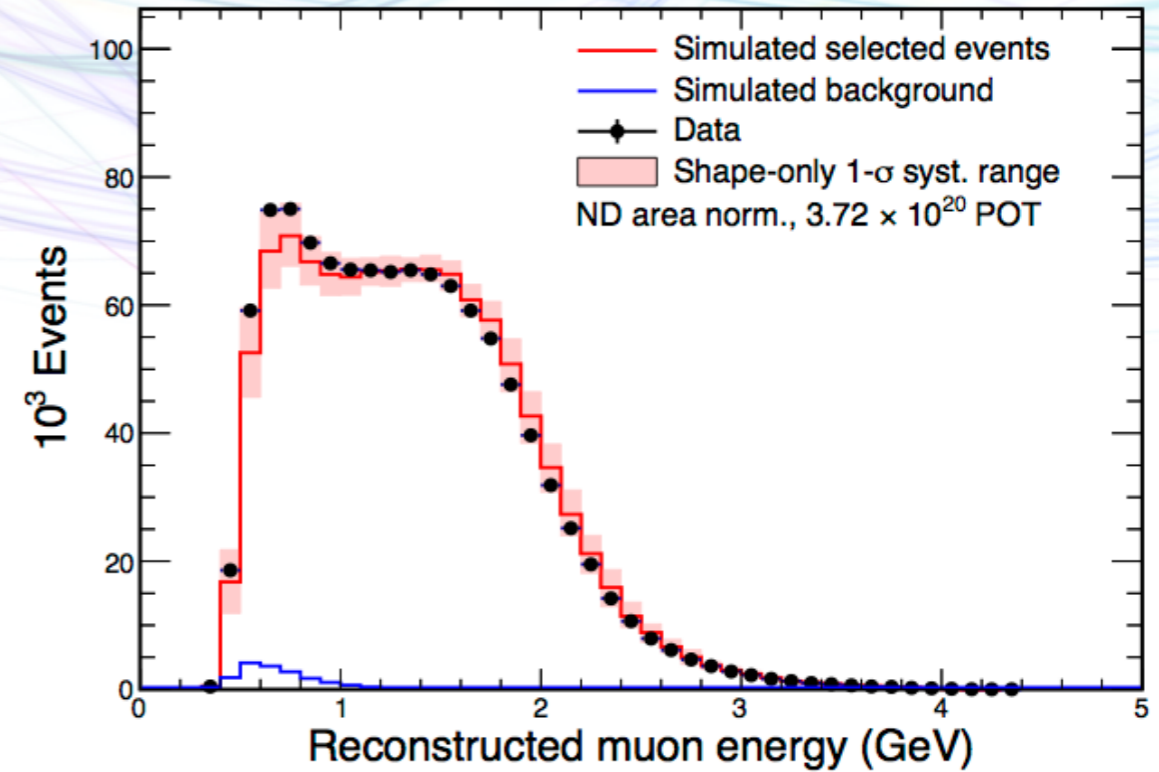
Muon neutrino energy reconstruction

NOvA Preliminary



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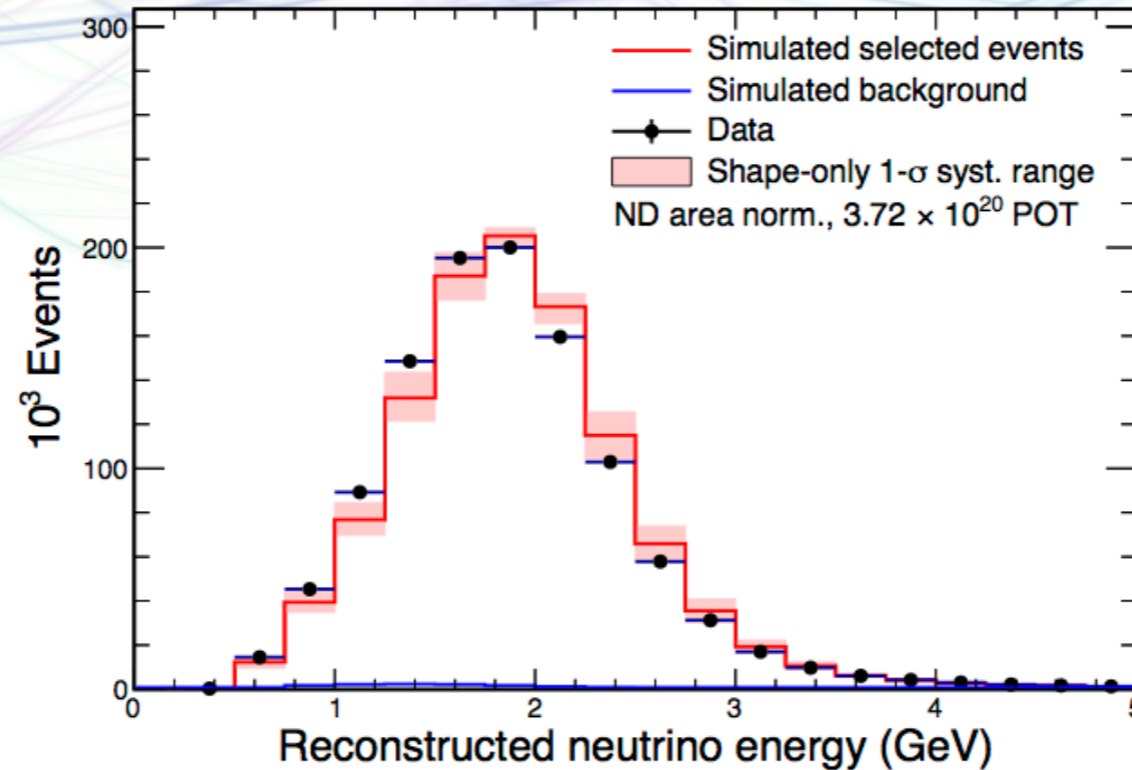
NOvA Preliminary



- ▶ Good data/MC agreement for muon neutrino selected events
- ▶ Hadronic energy scale uncertainty improved to 5%

Muon neutrino energy reconstruction

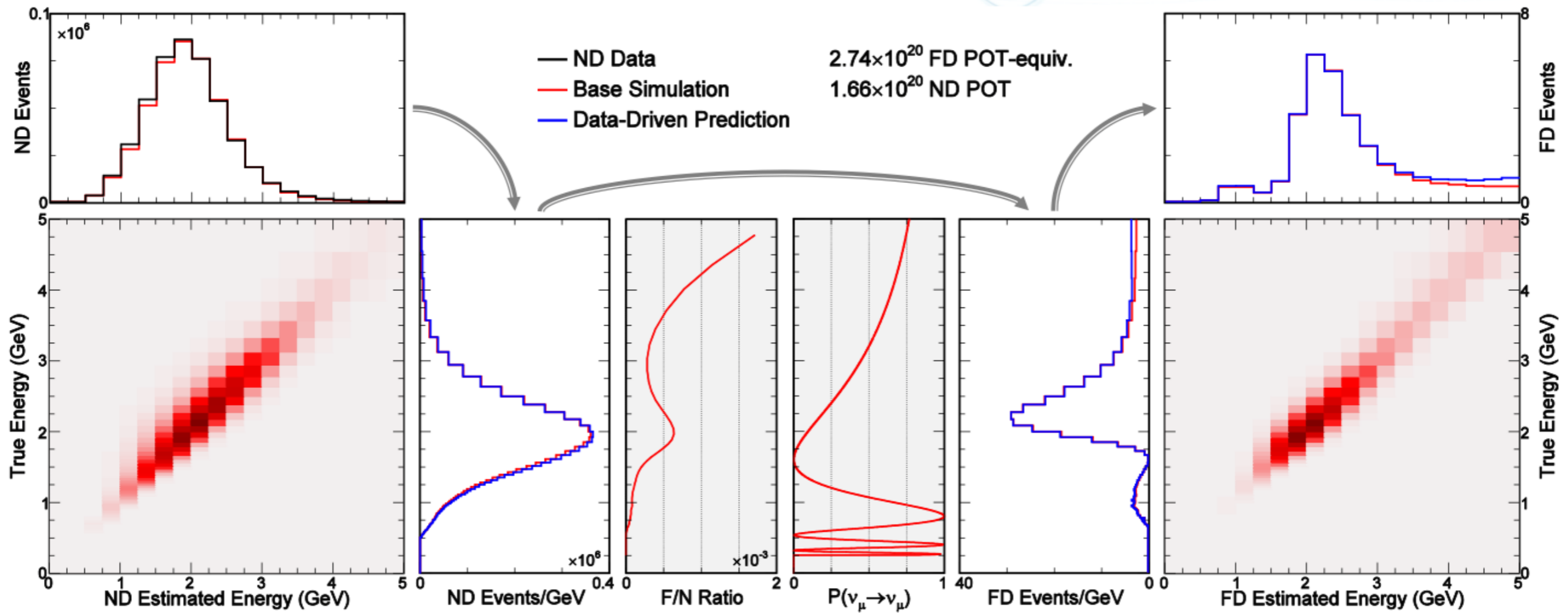
NOvA Preliminary



- ▶ Good data/MC agreement for muon neutrino selected events
- ▶ Hadronic energy scale uncertainty improved to 5%
- ▶ Use ND data to predict FD neutrino spectrum

Extrapolation procedure

- ▶ Translate ND observations to true energy
- ▶ Transport to far detector and oscillate
- ▶ Smear back to reco energy



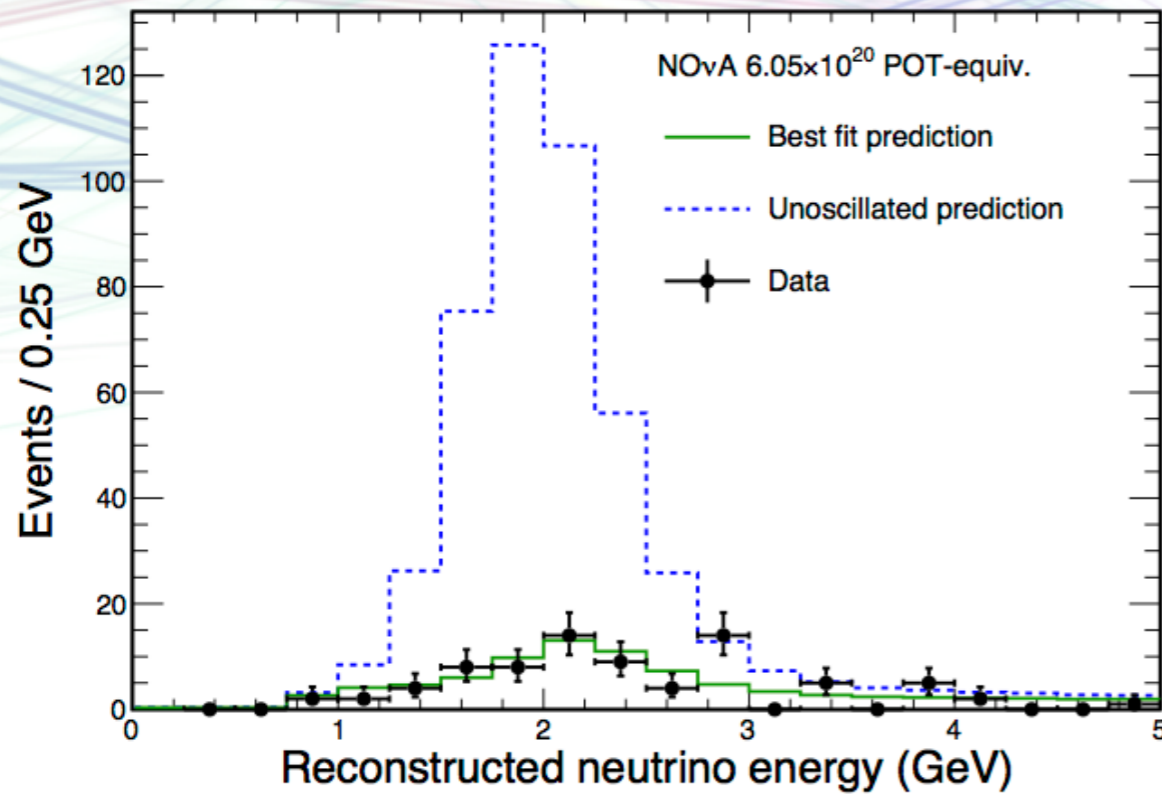
Systematic uncertainties

Source of uncertainty	Fractional uncertainty $\sin^2 \theta_{23}$ ($\pm\%$)	Fractional uncertainty Δm_{32}^2 ($\pm\%$)
Normalization	1.0	0.2
Muon E scale	2.2	0.8
Calibration	2.0	0.2
Relative E scale	2.0	0.9
Cross sections + FSI	0.6	0.5
Osc. parameters	0.7	1.5
Beam backgrounds	0.9	0.5
Scintillation model	0.7	0.1
Total systematic	3.4	2.4
Statistical uncertainty	4.1	3.5

- ▶ Consider multiple possible sources of systematic error
- ▶ Propagate effect of each through extrapolation
- ▶ Include as pull terms in fit
- ▶ Quoting increase (in quadrature) of measurement error

ν_μ disappearance results

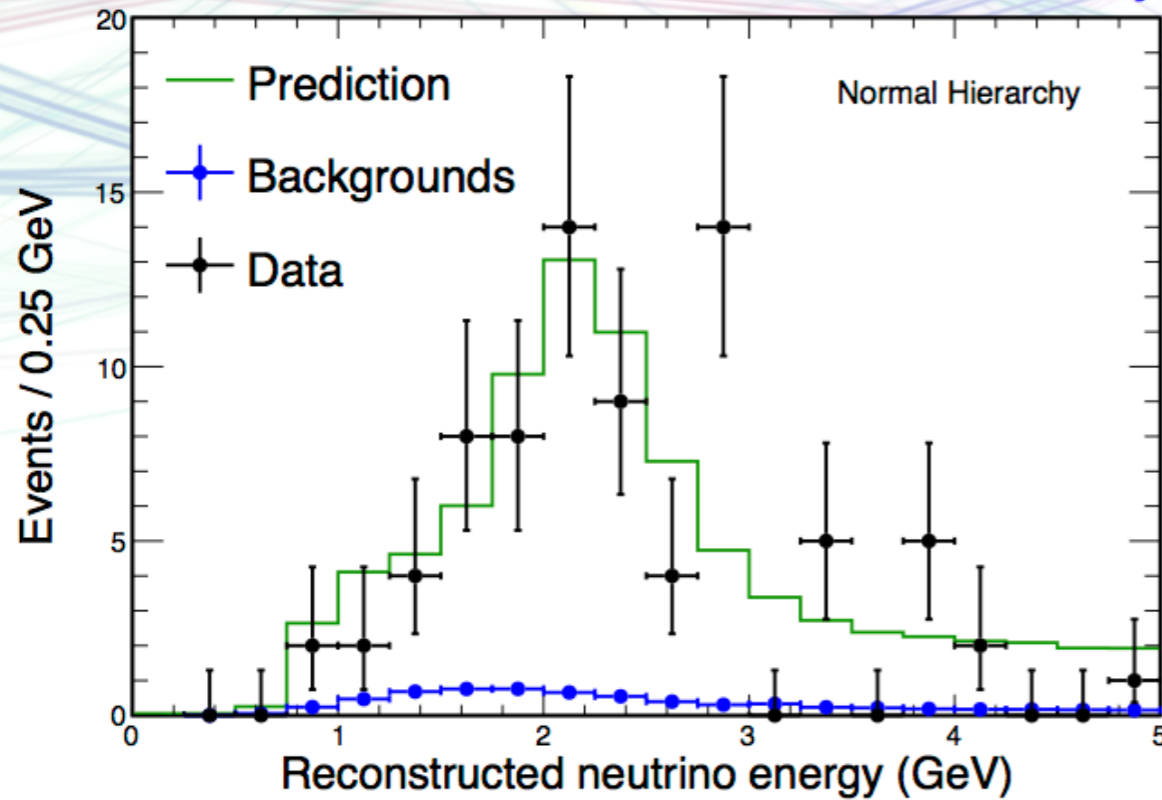
NOvA Preliminary



- ▶ Expect 473 FD ν_μ CC events with no oscillation
- ▶ Observe 82 (inc. 3.7 beam bkg. and 2.9 cosmic)

ν_μ disappearance results

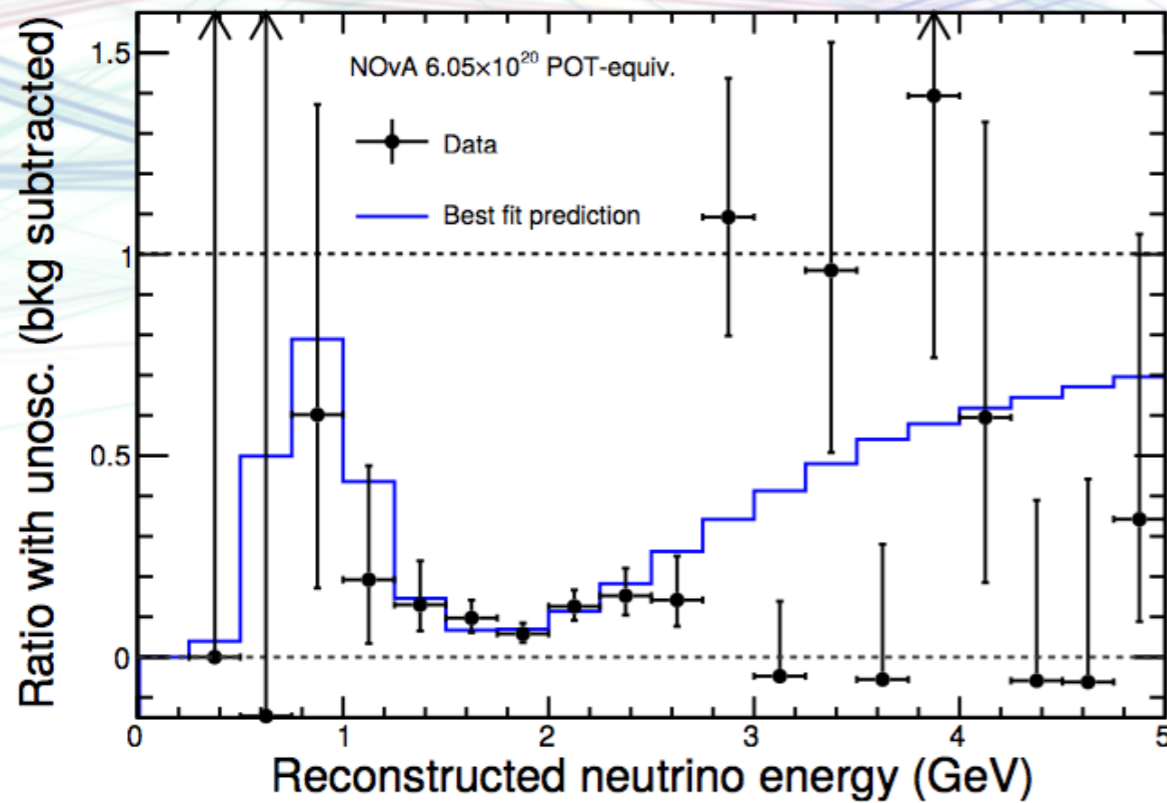
NOvA Preliminary



- ▶ Expect 473 FD ν_μ CC events with no oscillation
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ν_μ disappearance results

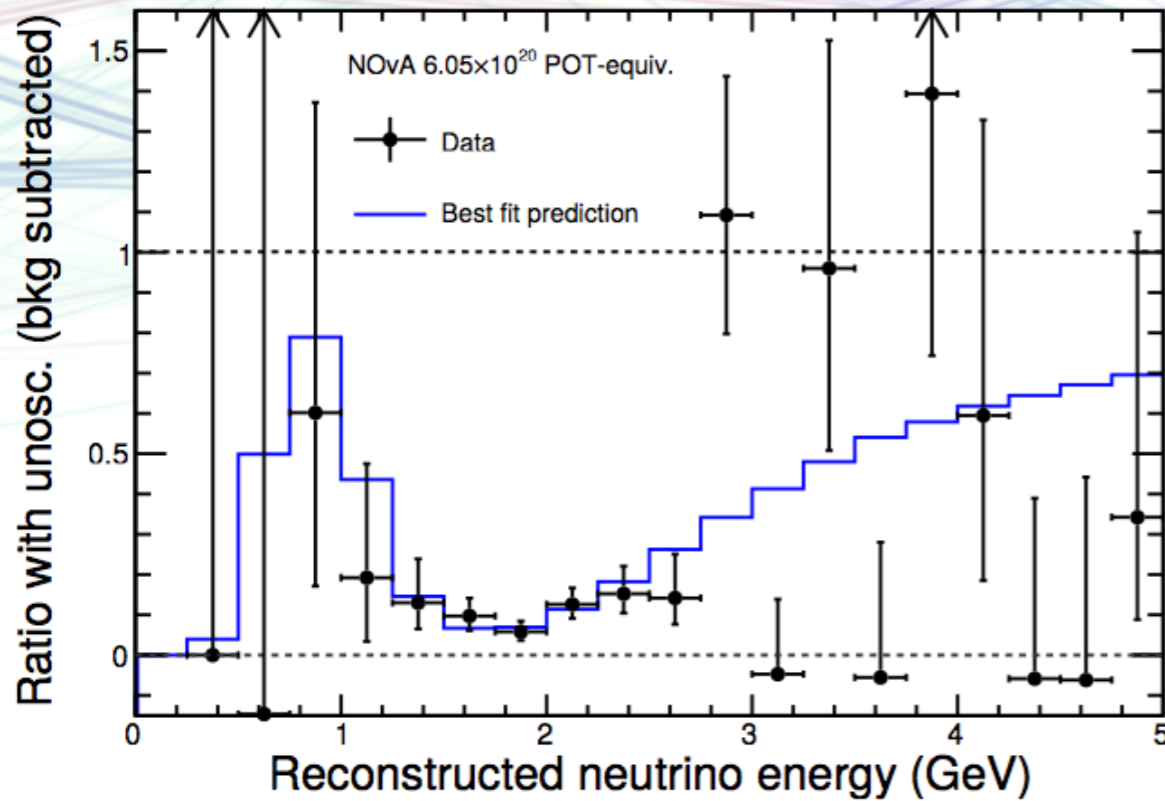
NOvA Preliminary



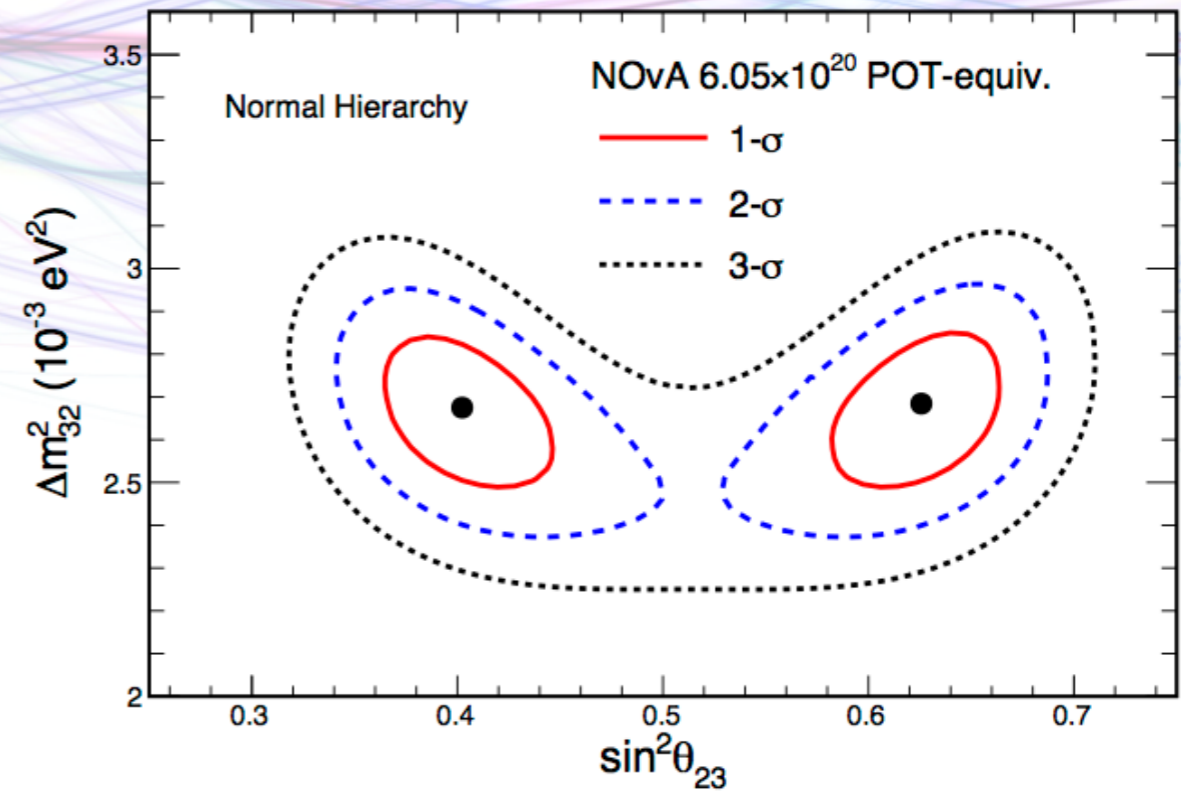
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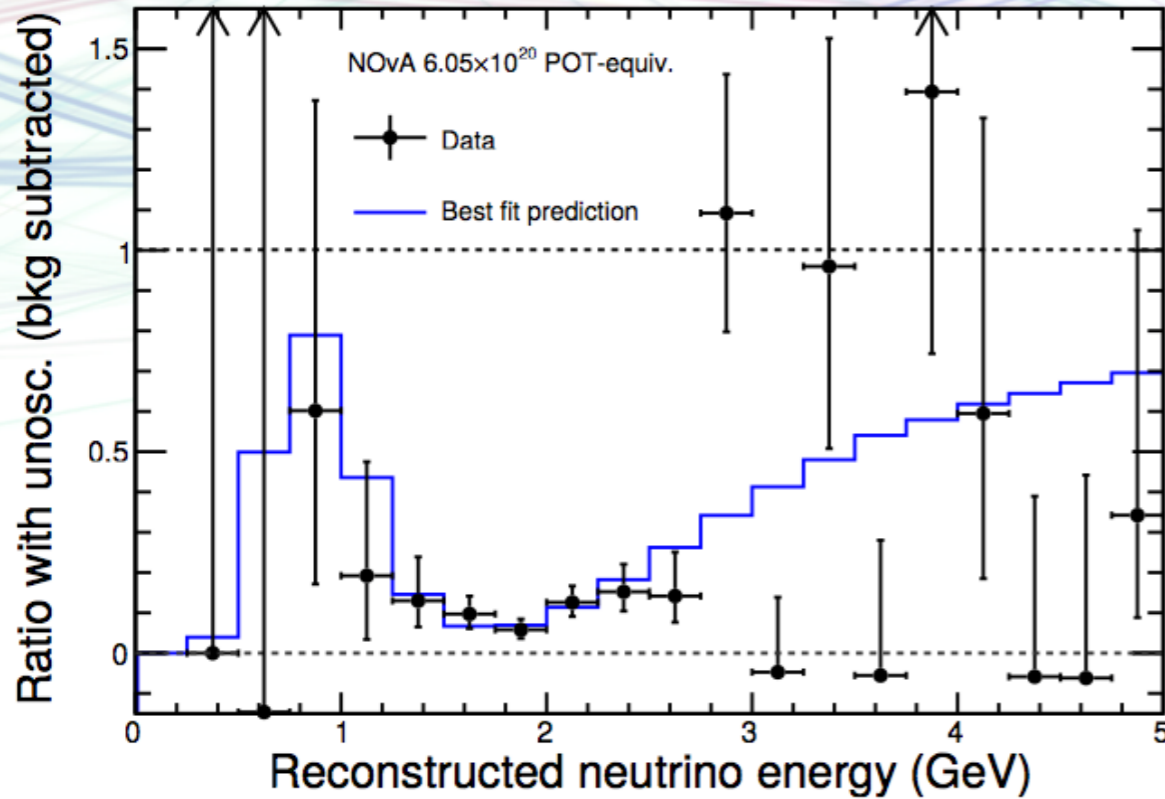
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$$\sin^2 \theta_{23} = 0.40_{-0.02}^{+0.03} \text{ (} 0.63_{-0.03}^{+0.02} \text{)}$$

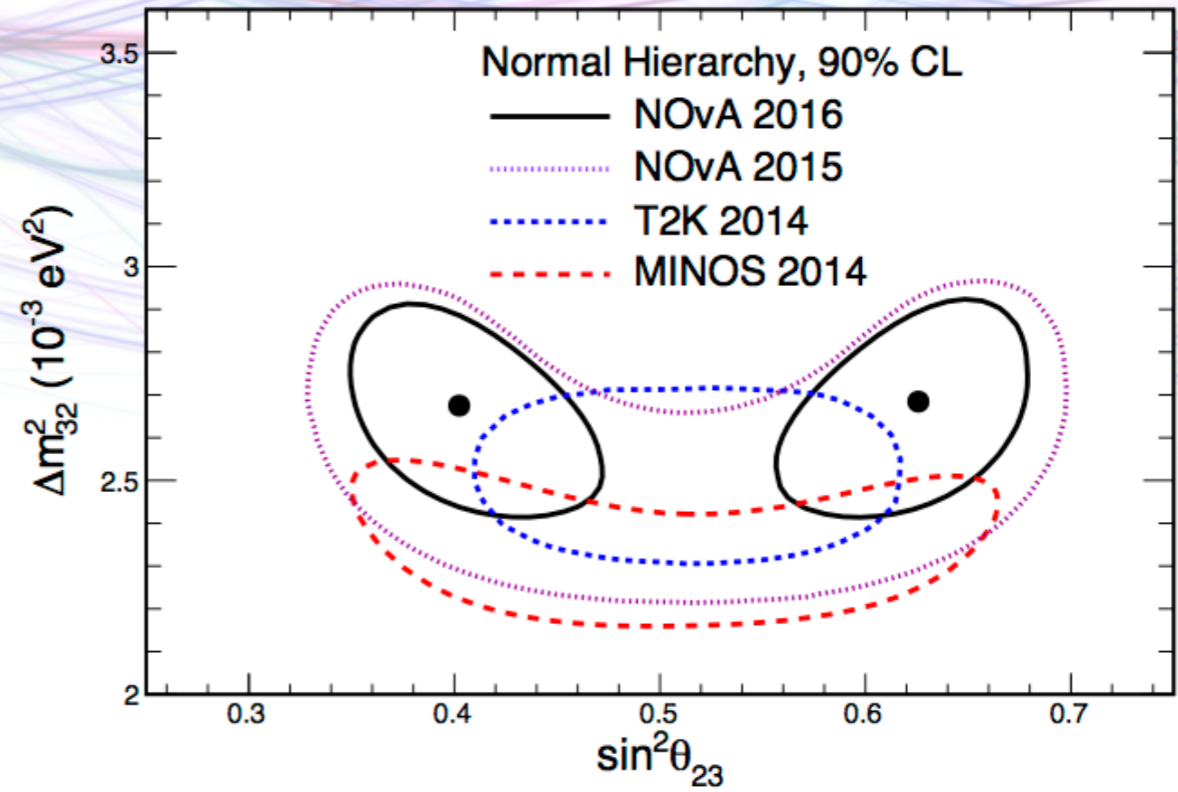
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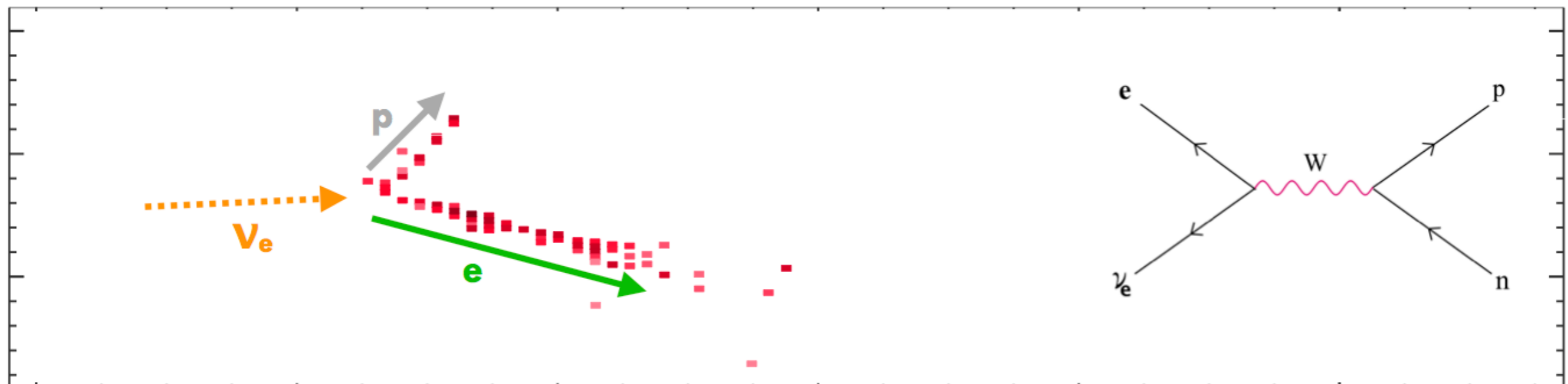
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Nue appearance

Principle of the ν_e measurement



- ▶ Separate ν_e CC interactions from beam backgrounds
 - ▶ Harder problem than ν_μ CC selection
- ▶ Evaluate remaining backgrounds in ND
 - ▶ Intrinsic beam ν_e
 - ▶ Neutral currents
 - ▶ ν_μ CC – mostly oscillates away
- ▶ An excess in the FD is the sign of $\nu_\mu \rightarrow \nu_e$ oscillations



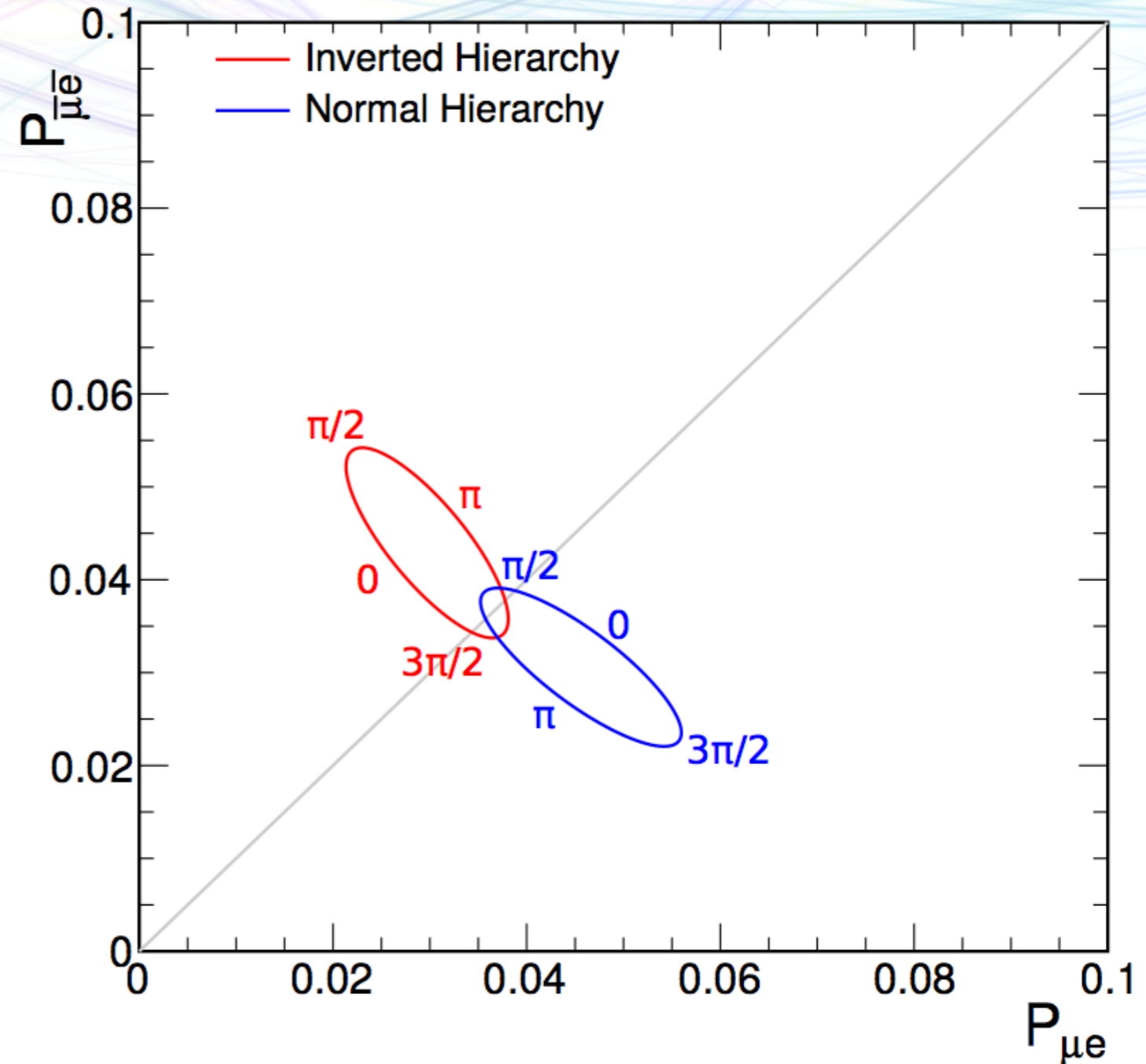
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- ▶ An excess in the FD is the sign of $\nu_\mu \rightarrow \nu_e$ oscillations
- ▶ $P_{\mu e} \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left(\frac{\Delta m_{32}^2 L}{4E} \right) + f(\text{sign}(\Delta m_{32}^2)) + f(\delta_{CP})$
- ▶ θ_{13} only 8.5° degrees, most ν_μ go to ν_τ instead
- ▶ Look for deviations due to hierarchy (matter effects) and CP-violation

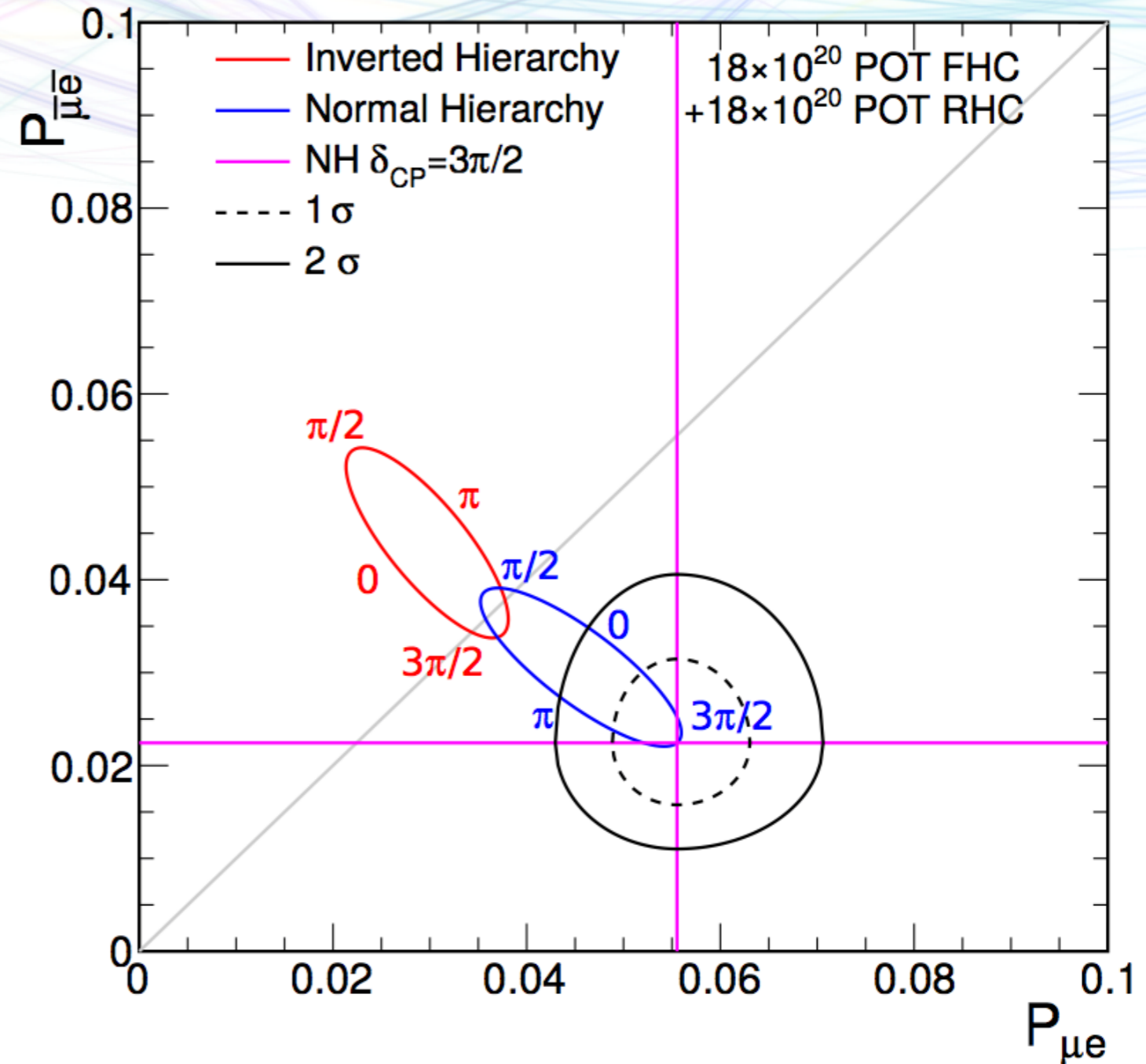
Principle of the ν_e measurement

- ▶ To first order, NOvA measures $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ evaluated at 2GeV
- ▶ These depend differently on $\text{sign}(\Delta m_{32}^2)$ and δ_{CP}



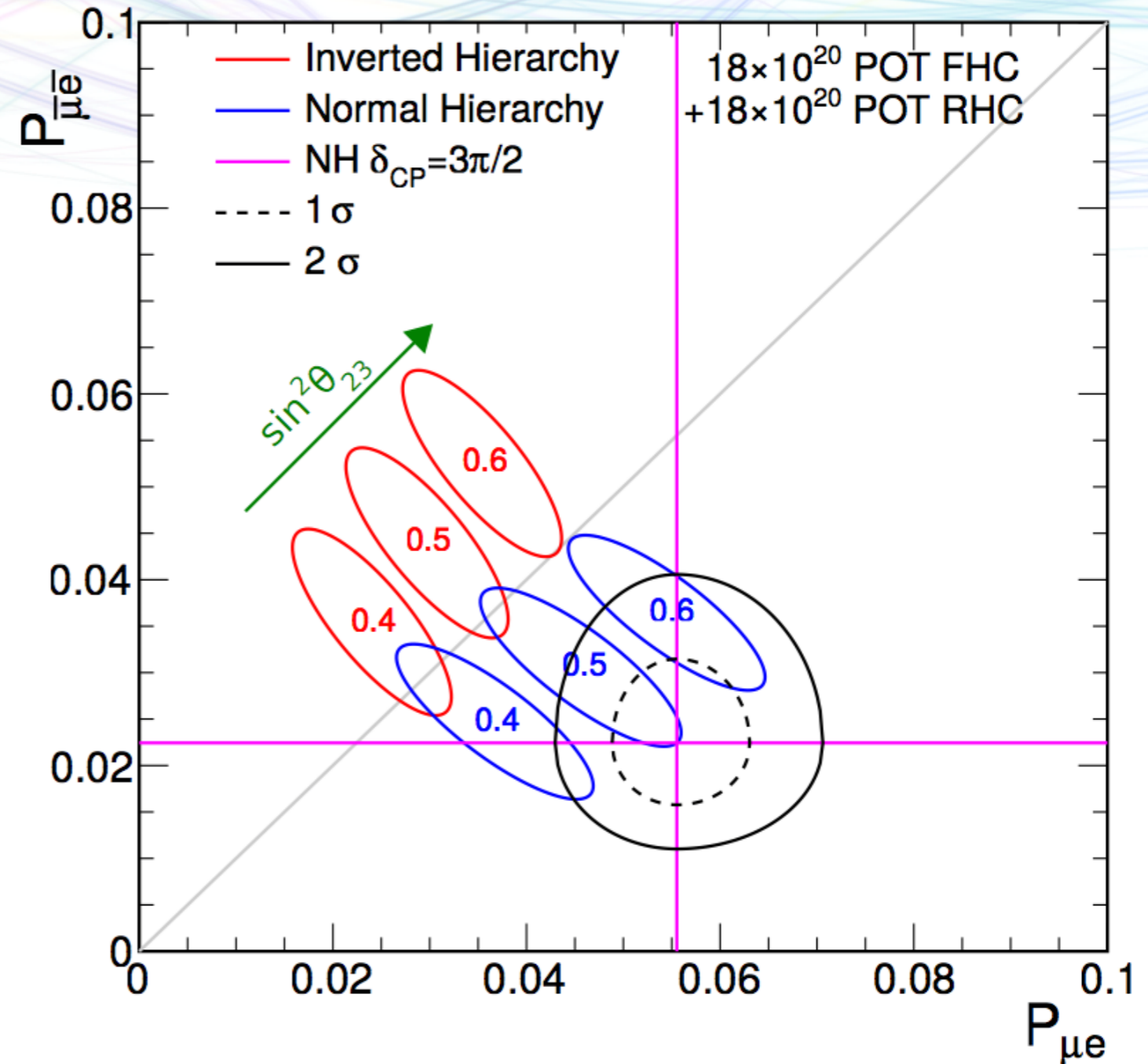
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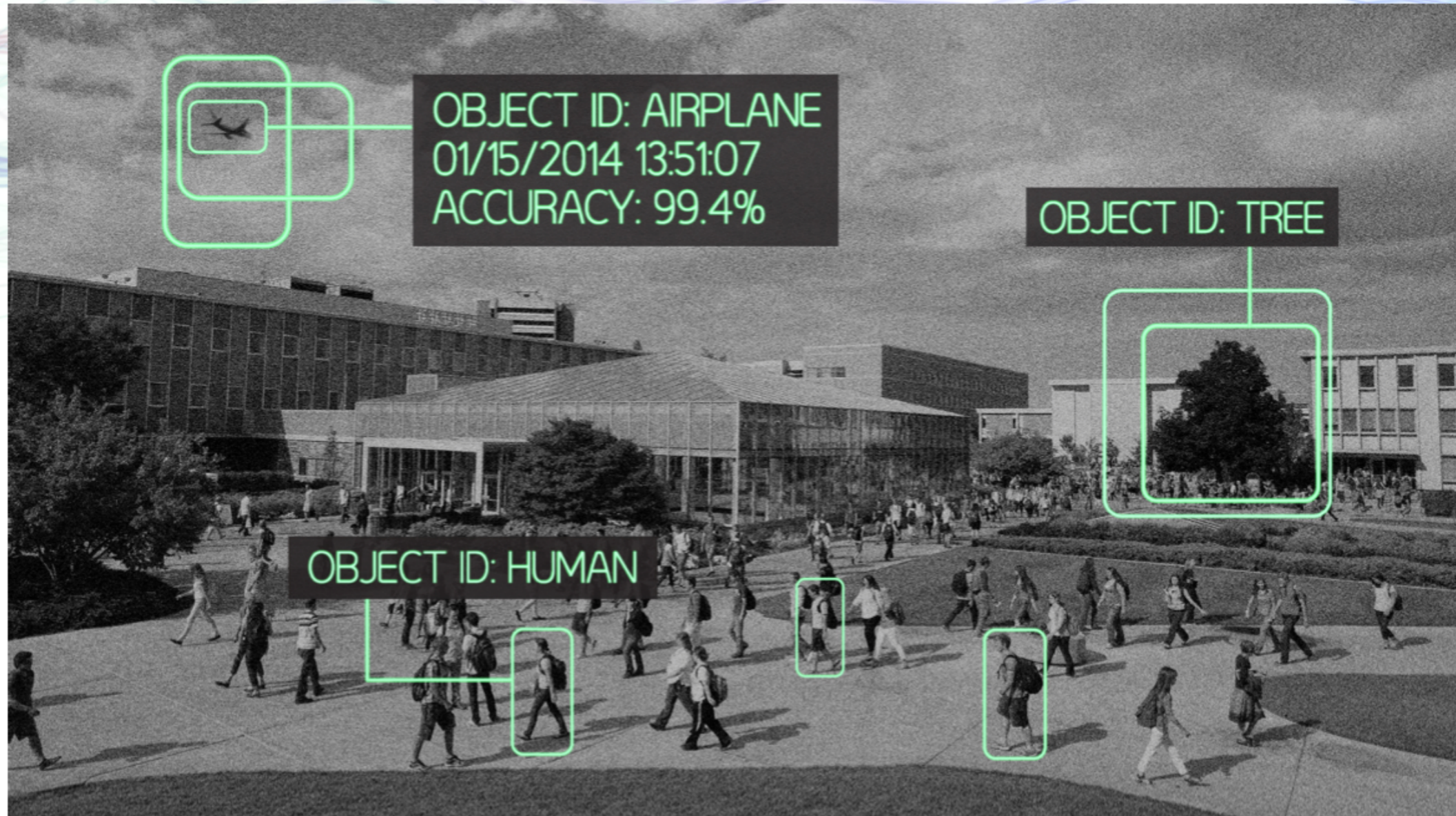


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- ▶ P also $\propto \sin^2 \theta_{23}$



Convolution Neural Networks



- ▶ Recent advances in machine learning/computer vision
- ▶ Achieving near-human performance on image classification tasks
- ▶ Why not classify event-displays?

Convolution Neural Networks

$$\frac{1}{8} \begin{bmatrix} -1 & -1 & -1 \\ -1 & +8 & -1 \\ -1 & -1 & -1 \end{bmatrix}$$

Edge-detection kernel



- ▶ Recent advances in machine learning/computer vision
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- ▶ **CNN** – deep neural network, inputs are the pixels of the image
- ▶ Take advantage of translational invariance → convolutions

Convolution Neural Networks



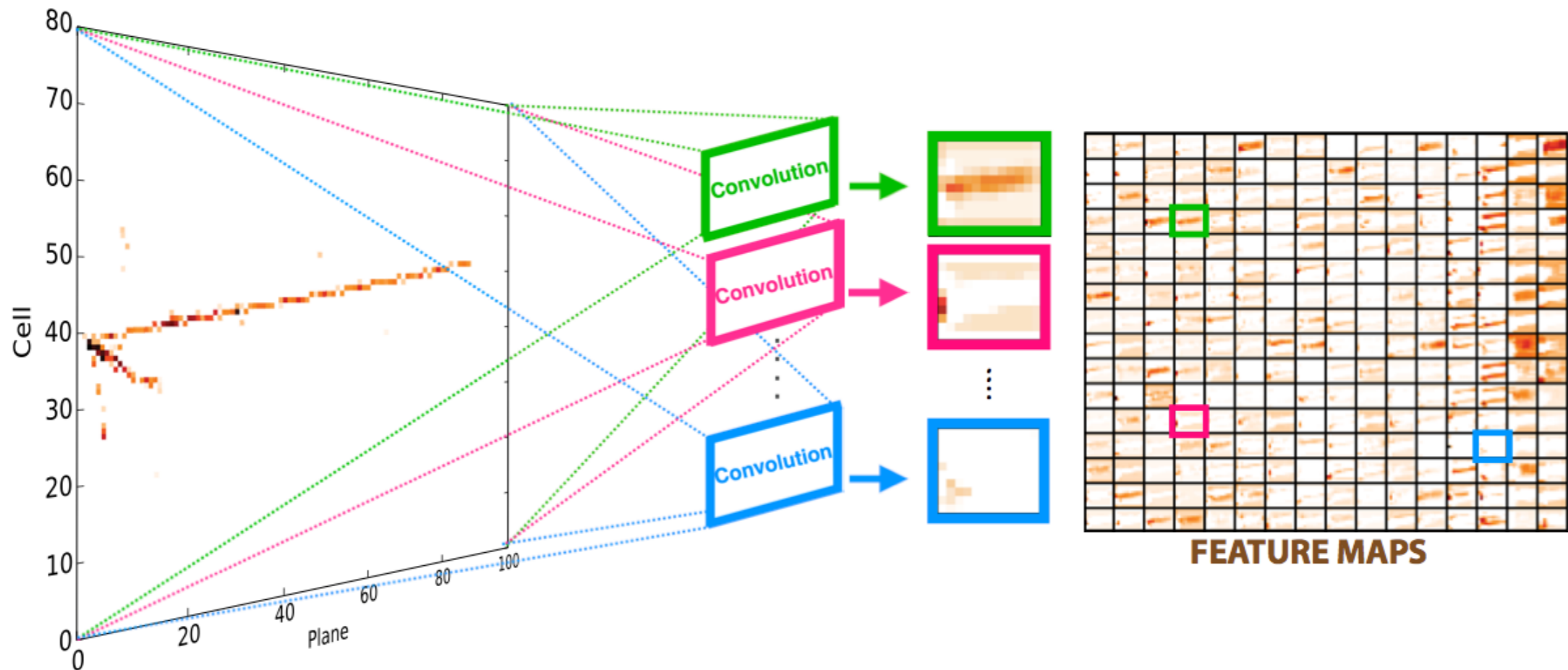
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Selecting electron neutrinos – CVN

- ▶ **C**onvolutional **V**isual **N**etwork (CVN)
- ▶ Early layers perform convolutions to pick out abstract features
- ▶ Fully-connected final layers
- ▶ Trained using FNAL's Wilson Cluster GPUs
- ▶ Statistical power equivalent to 30% more exposure than previous IDs



arXiv:1604.01444

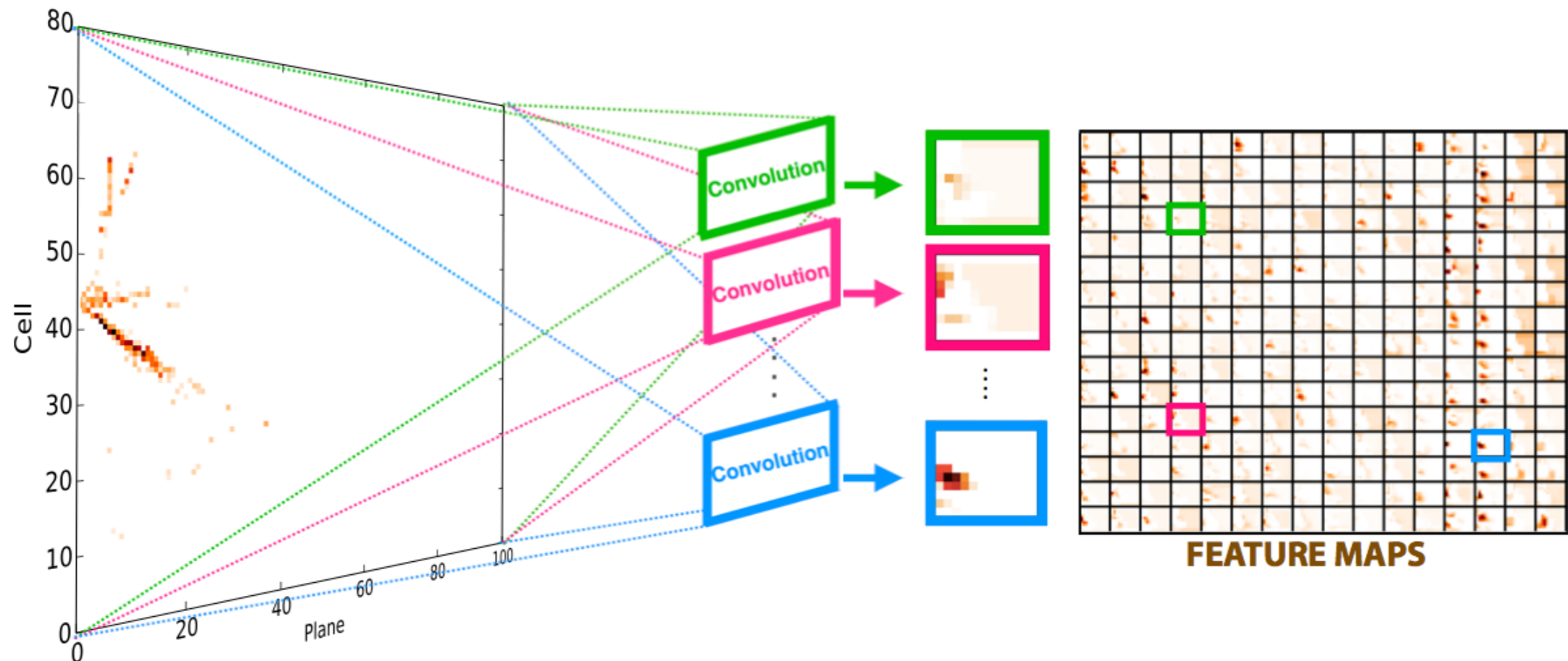
C. Backhouse (Caltech)

NOvA

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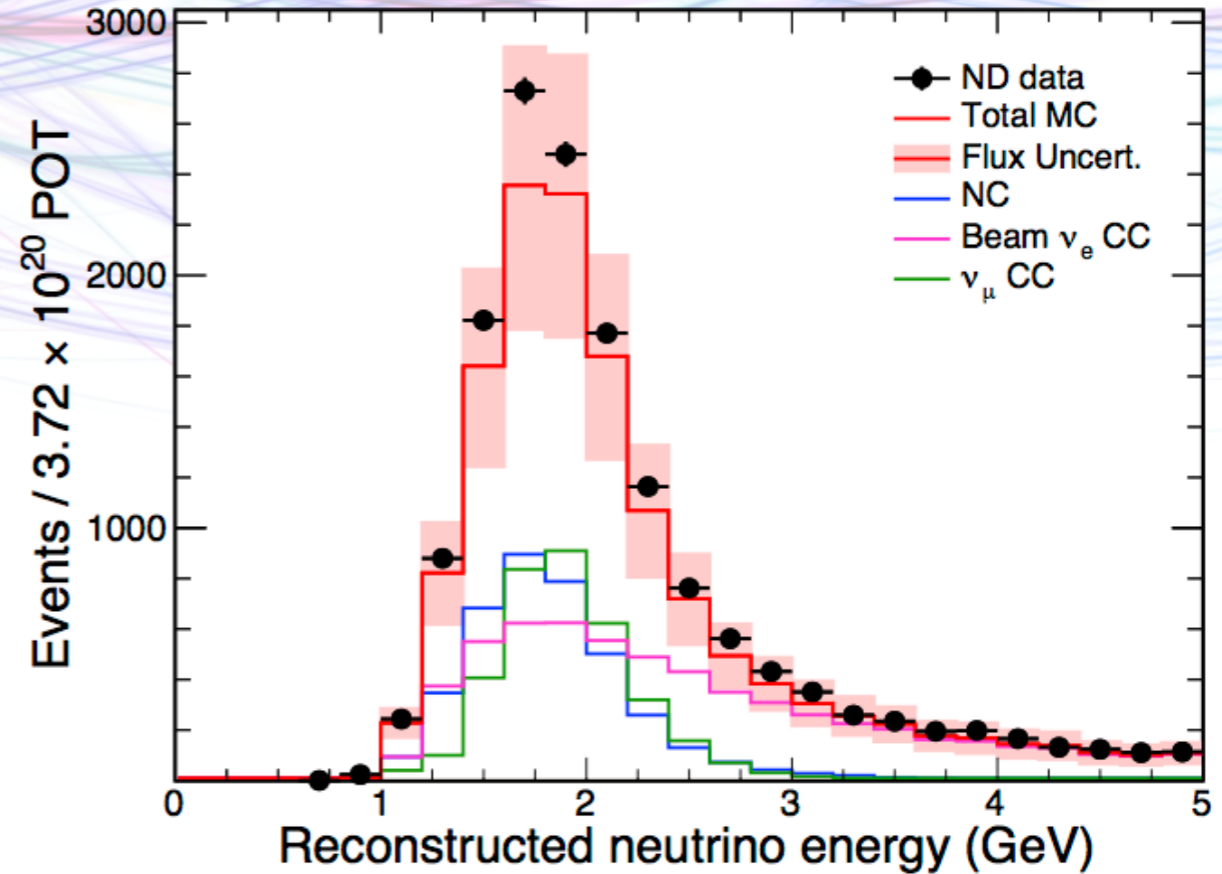
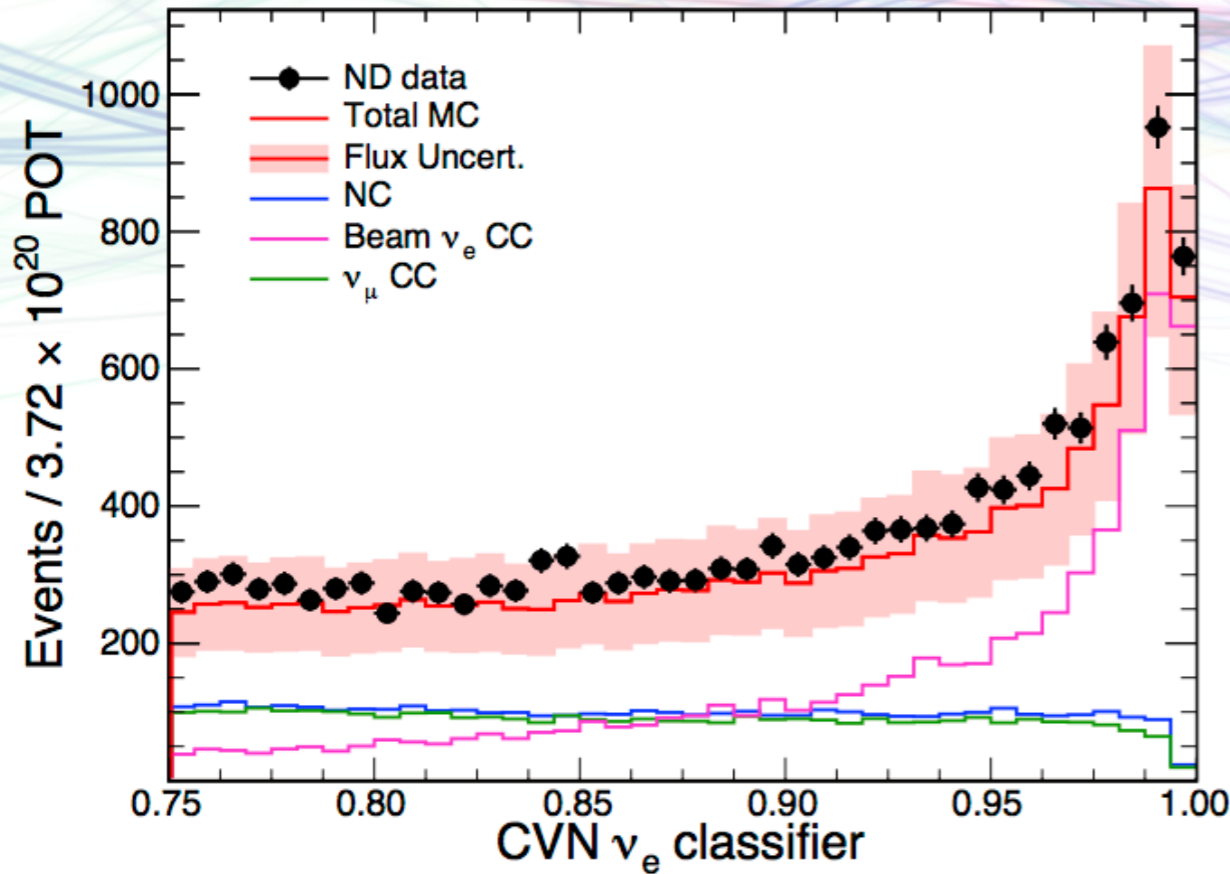
NOvA

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Selecting electron neutrinos – CVN

NOvA Preliminary

NOvA Preliminary

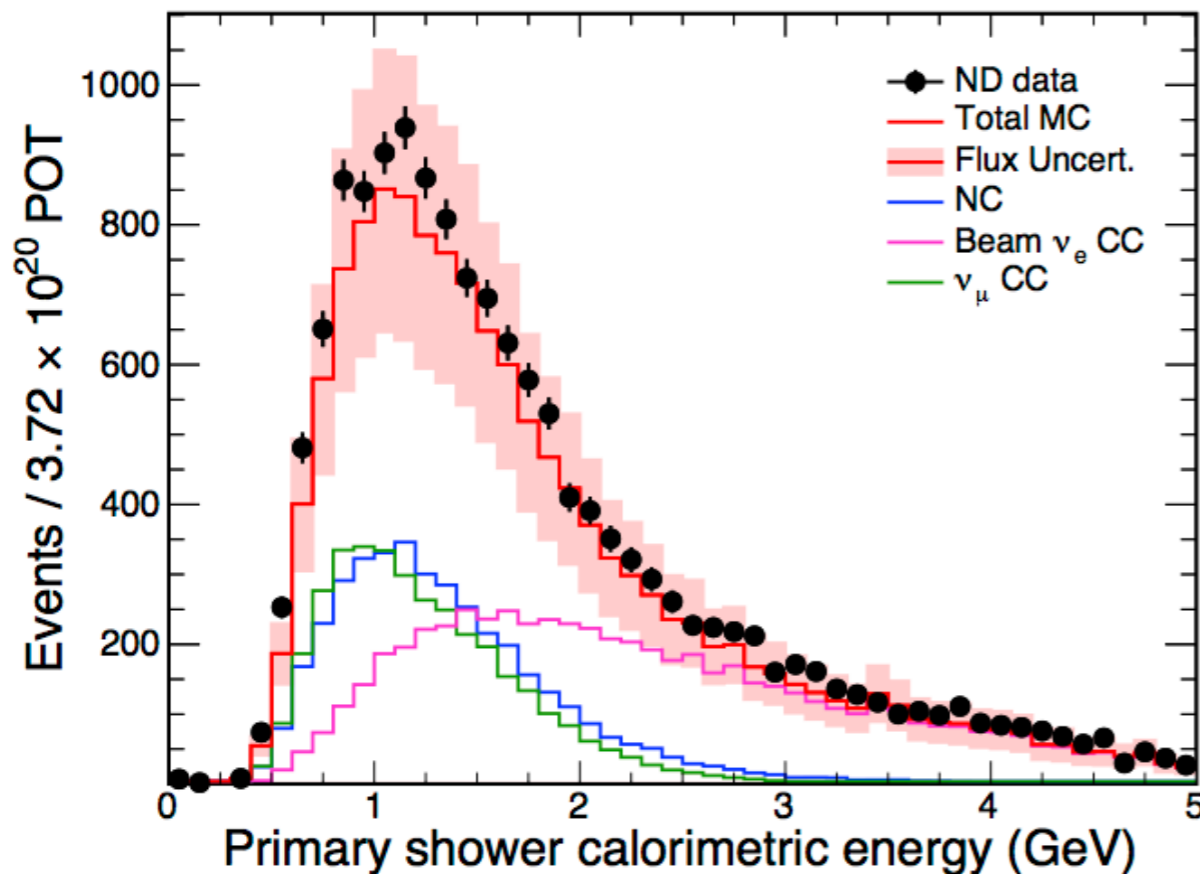


- ▶ 73% ν_e signal efficiency, 76% purity
- ▶ Loosen PID cut to maximize $s/\sqrt{s+b}$
- ▶ Analyze in 3 PID \times 4 energy bins

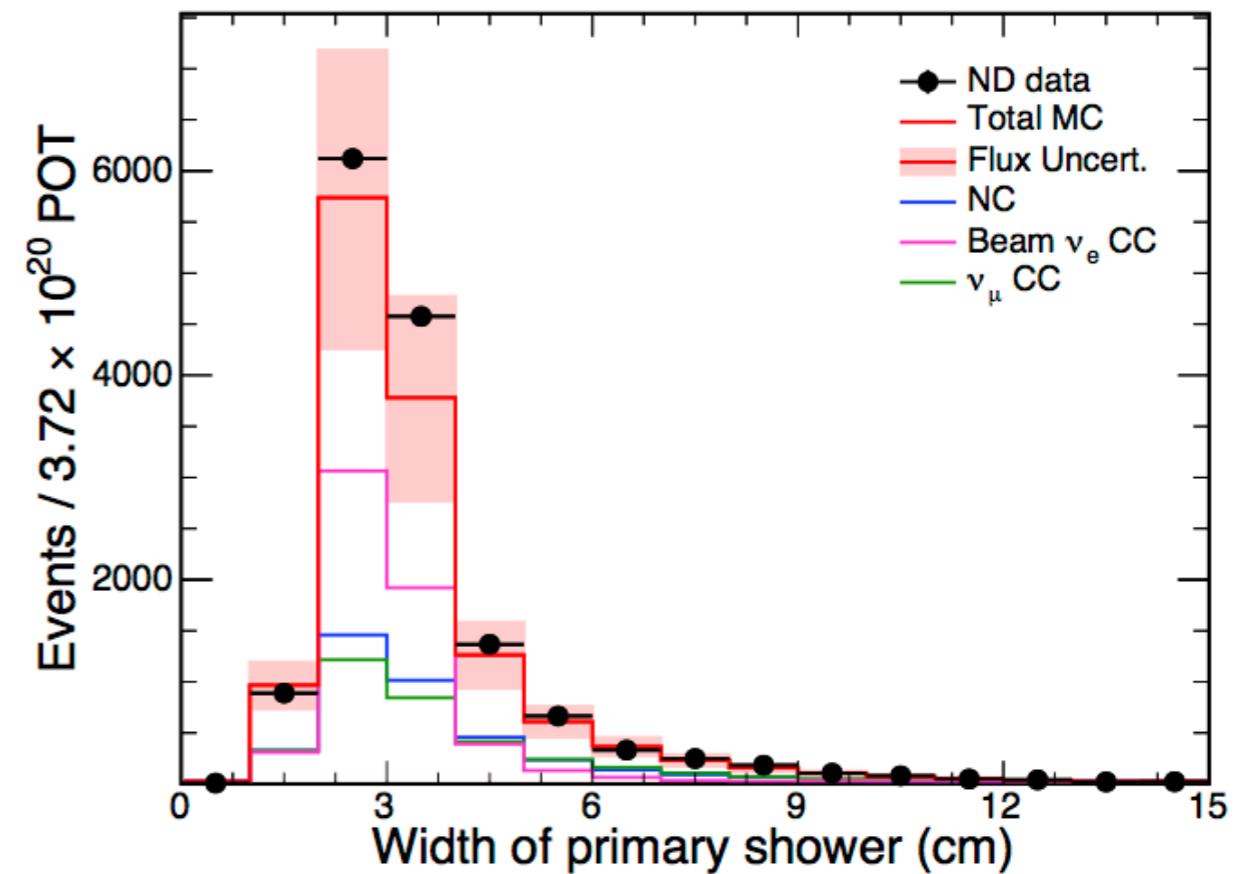
Selector checks

- ▶ CVN subjected to the same (or greater) scrutiny than past PIDs
- ▶ ND data/MC is good
- ▶ Better cosmic rejection and similar systematics to other PID options

NOvA Preliminary

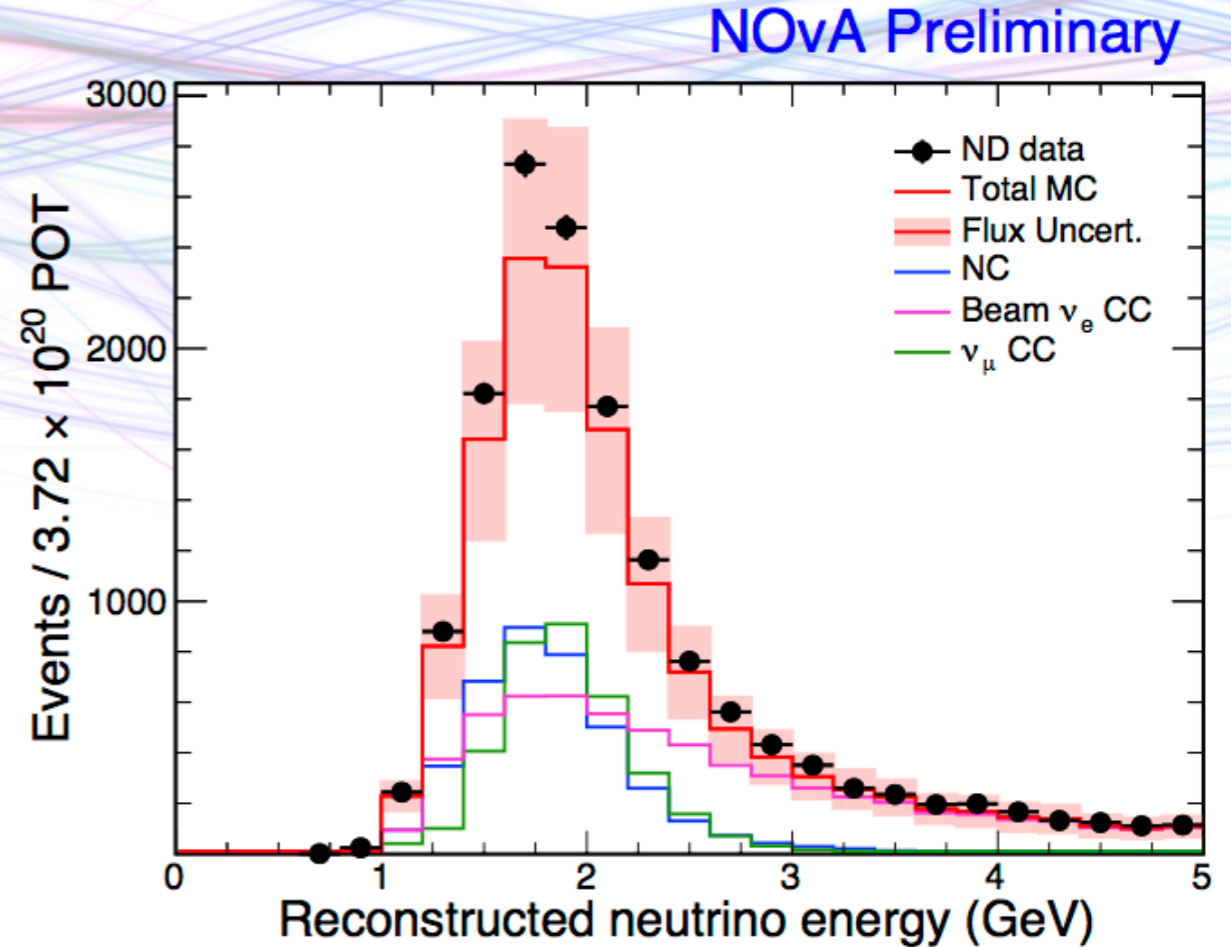


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ND decomposition

- ▶ Use ND data to predict FD backgrounds
 - ▶ Beam ν_e CC
 - ▶ NC
 - ▶ ν_μ CC
- ▶ $\sim 10\%$ excess of data over MC

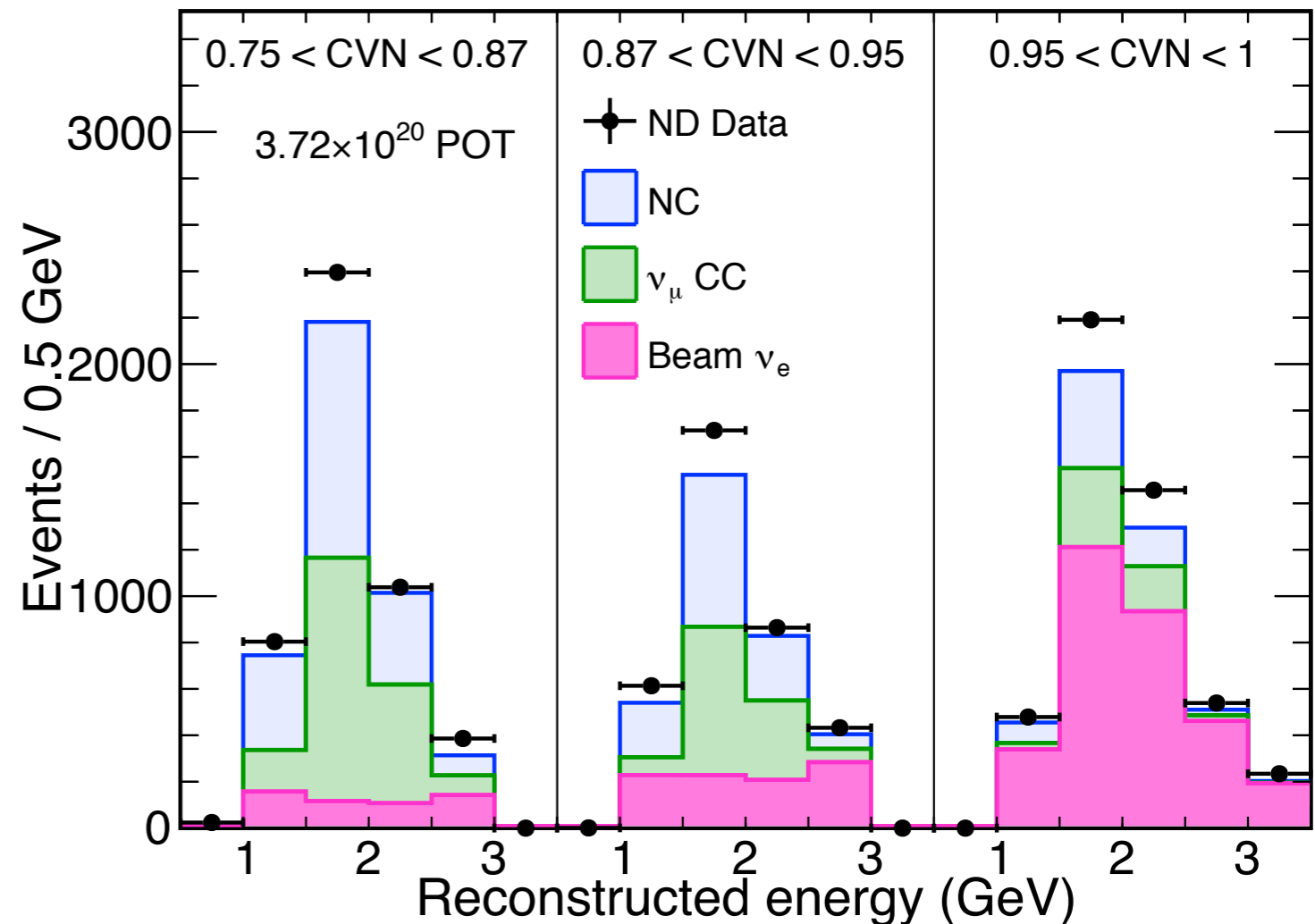


- ▶ How to divide between the components?
- ▶ e.g. most ν_μ CC oscillate away before FD



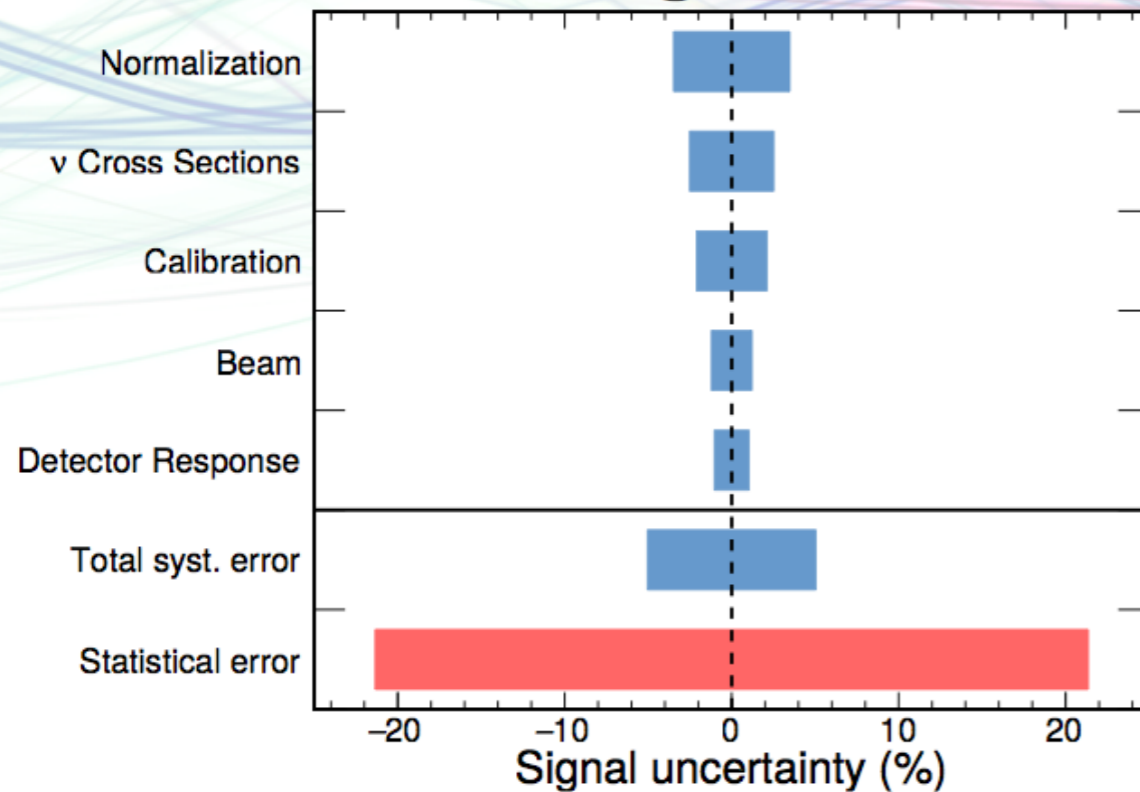
Data Driven Background Corrections

- ν_e CC selection in the ND picks out FD backgrounds
 - beam ν_e CC
 - ν_μ CC
 - NC
- $\sim 10\%$ excess of data over MC in the ND
- Extrapolate data/MC differences to adjust FD prediction
- Each component oscillates differently
- Must decompose the data into constituent components

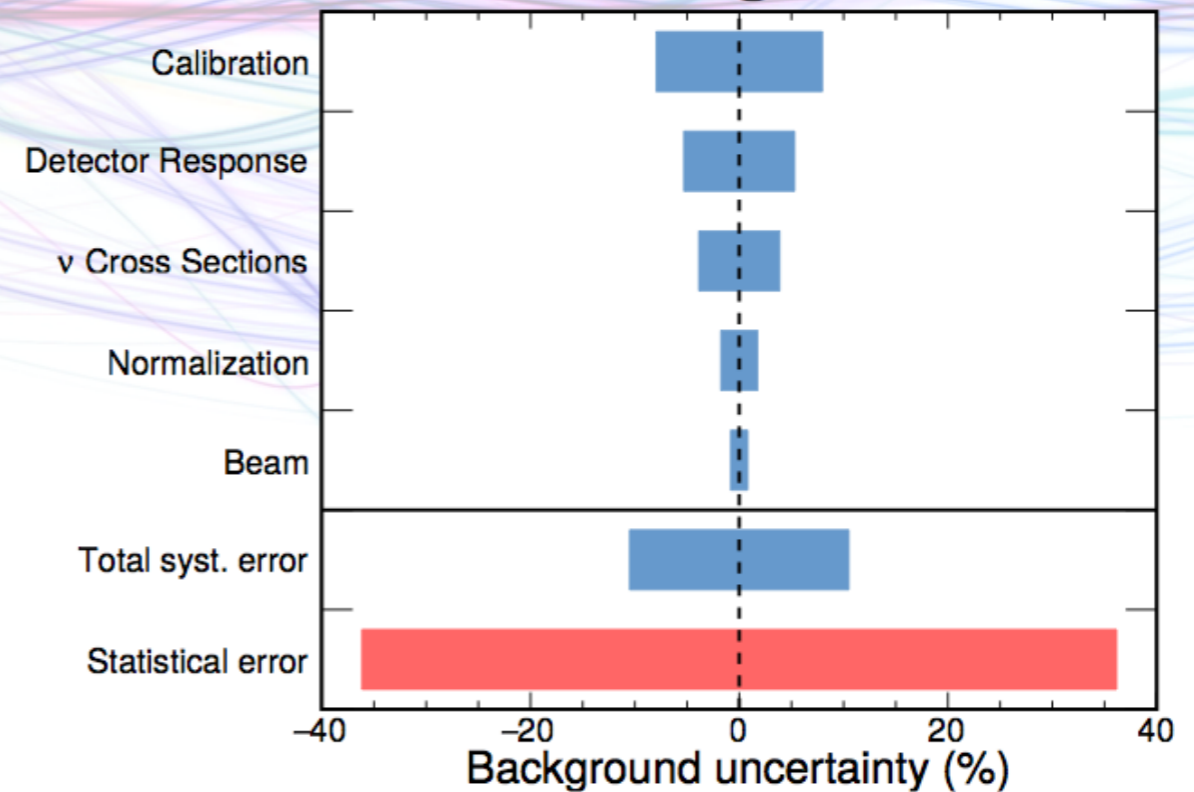


Systematics

Signal



Background



- ▶ Considered multiple possible sources of systematic error
- ▶ Propagate shifts through to update FD prediction
- ▶ Total $\sim 5\%$ error on signal, 10% on bkg.
- ▶ Fit nuisance parameters as pull terms
- ▶ Dominated by statistical error

Event count expectations

Total prediction

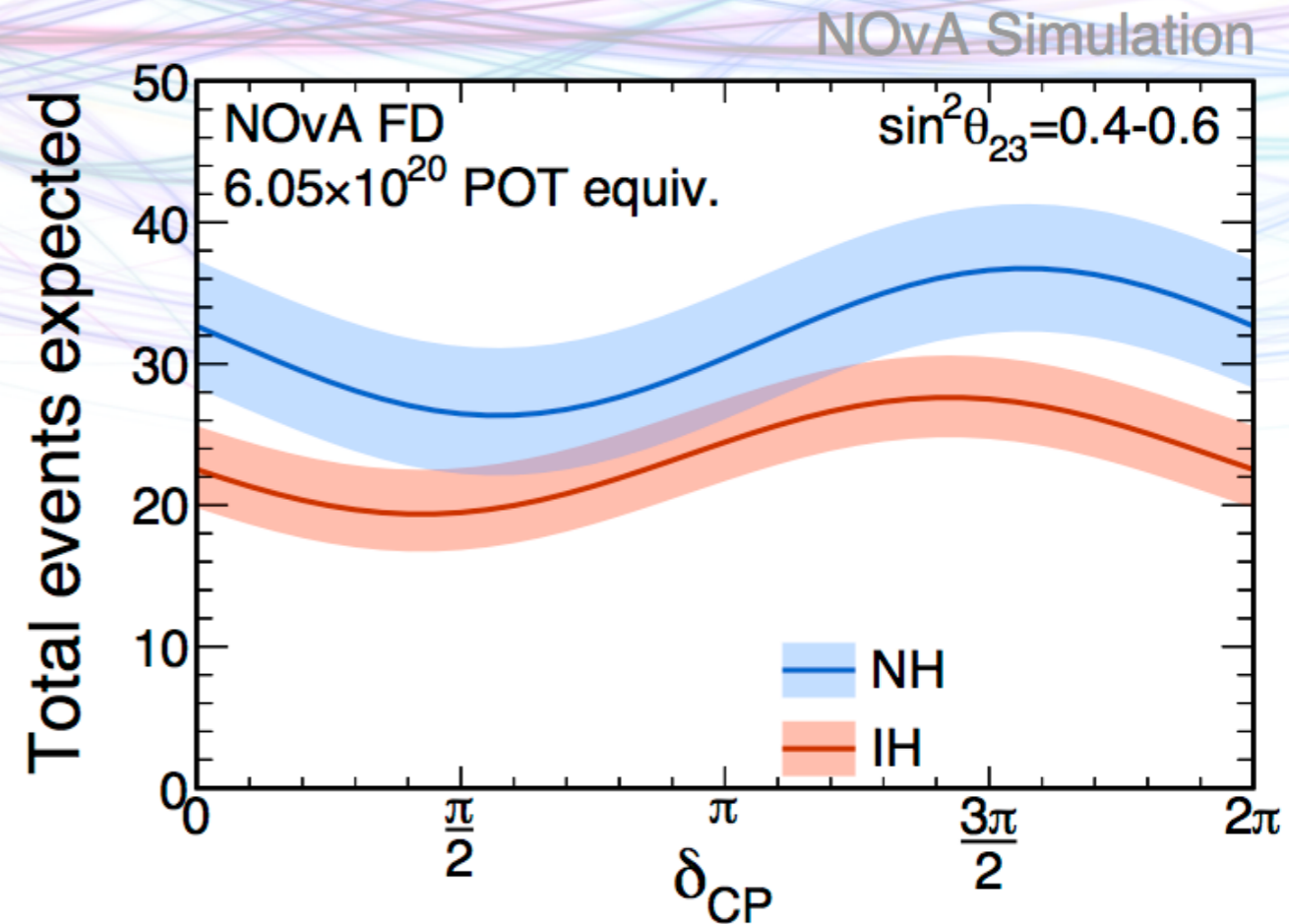
NH $\frac{3\pi}{2}$	IH $\frac{\pi}{2}$	$\sin^2 \theta_{23} = 0.5$
36.4	19.4	$\pm 5\%$ syst.

$$P_{\nu_{\mu} \rightarrow \nu_e} \propto \sin^2 \theta_{23} \sin^2 2\theta_{13}$$

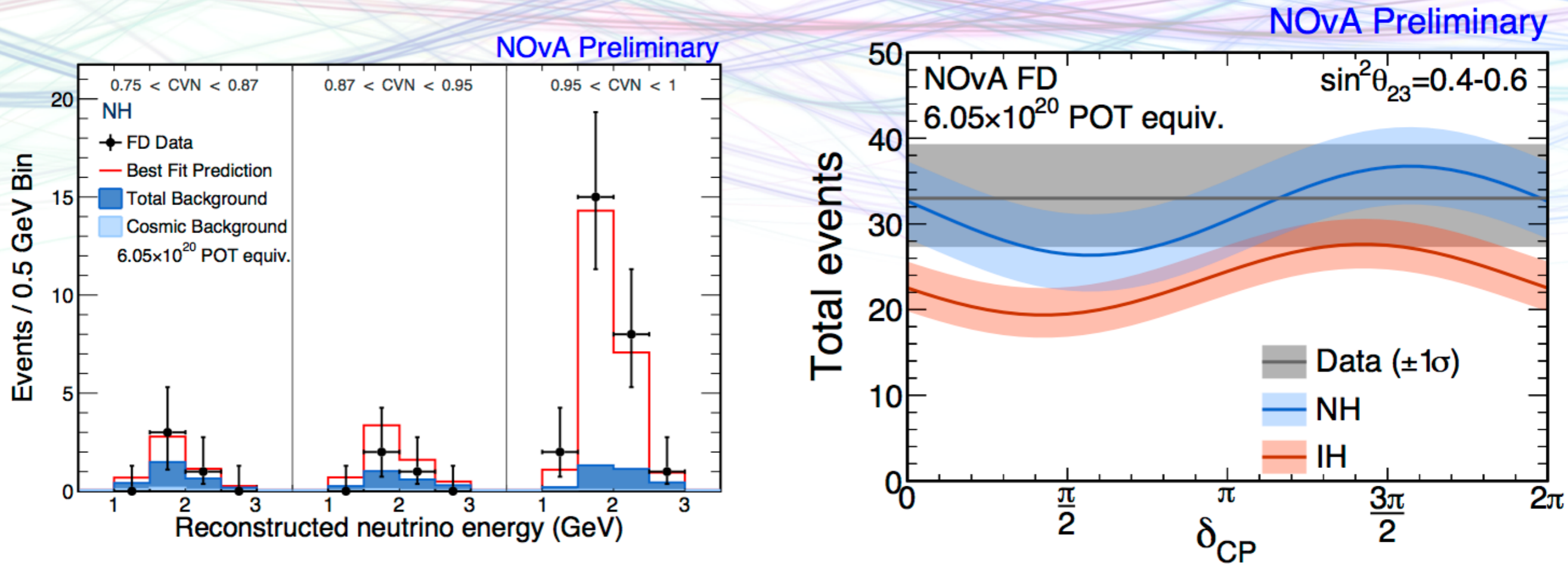
Background components

Total bkg	NC	beam ν_e	ν_{μ} CC	ν_{τ} CC	cosmics	
8.2	3.7	3.1	0.7	0.1	0.5	$\pm 10\%$ syst.

Essentially independent of oscillation parameters



ν_e appearance results

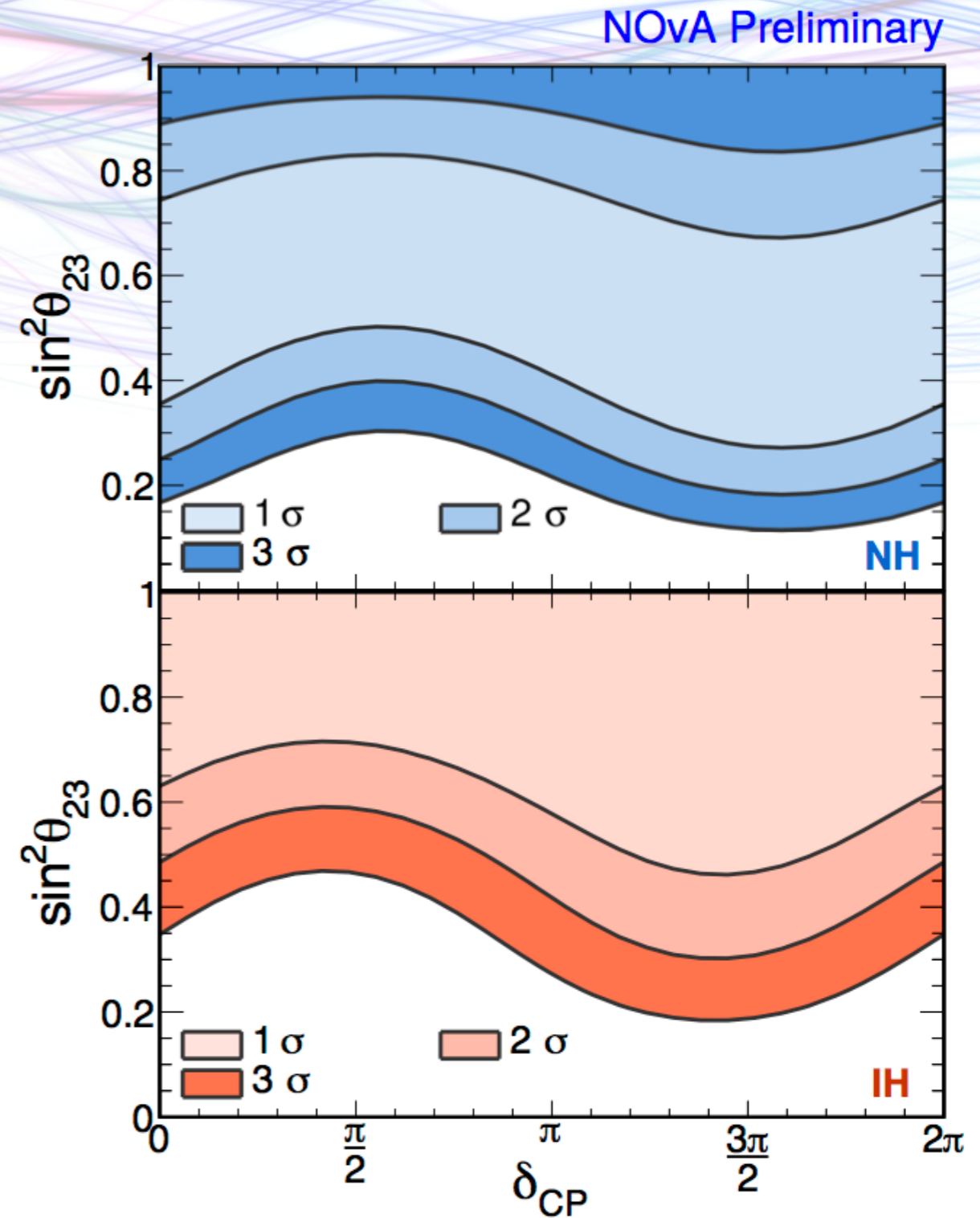


- ▶ Observe **33** events passing ν_e selection
- ▶ On 8.2 background
- ▶ Towards the higher end of expectations

Previous result PIDs: LID(LEM) sees 34(33) events on bkg. of 12.2(10.3)

ν_e fit results

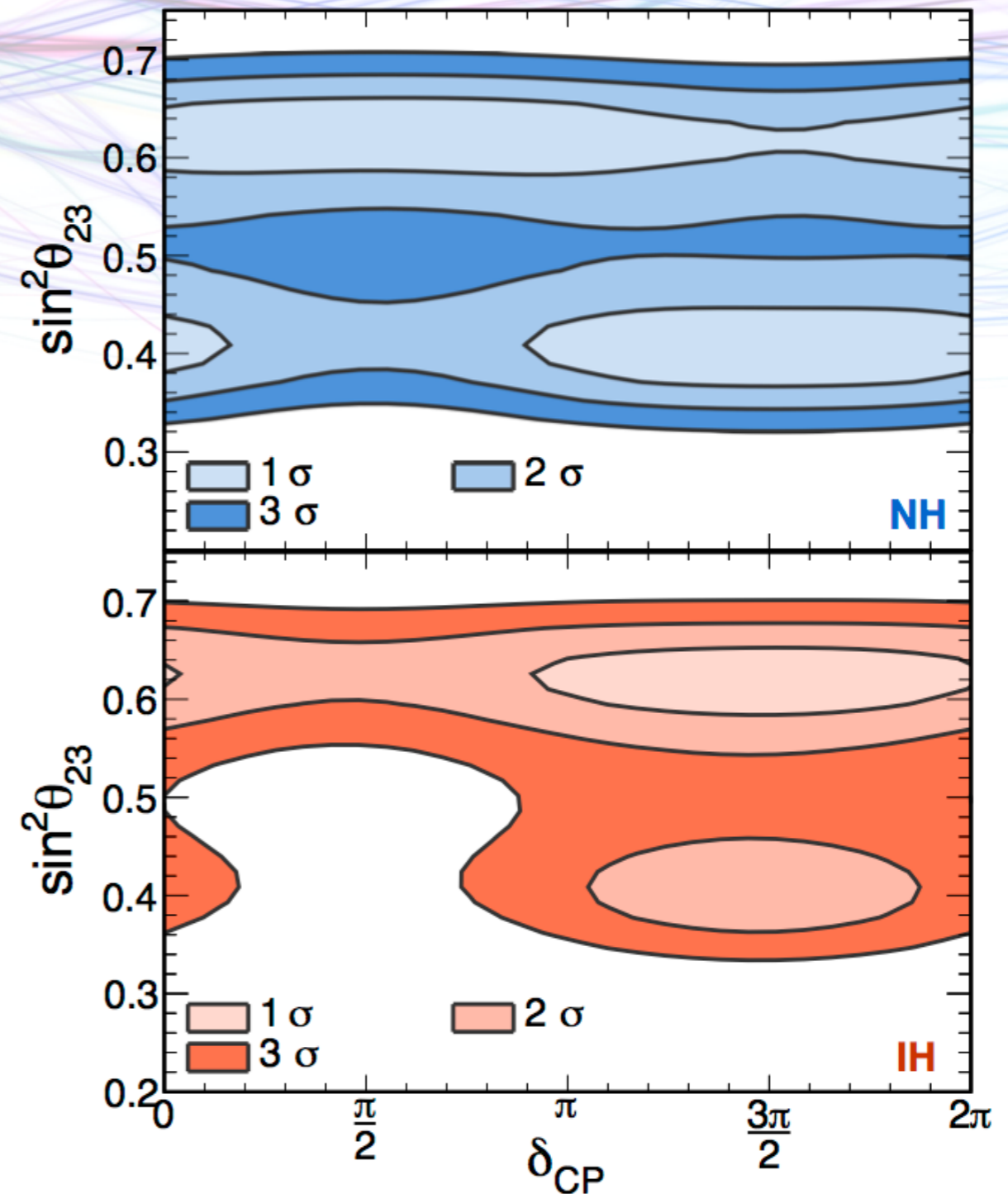
- ▶ Constrain θ_{13} to reactor average
 $\sin^2 2\theta_{13} = 0.085 \pm 0.005$



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- ▶ Not a full joint fit. No syst./osc. param correlations. No FC.

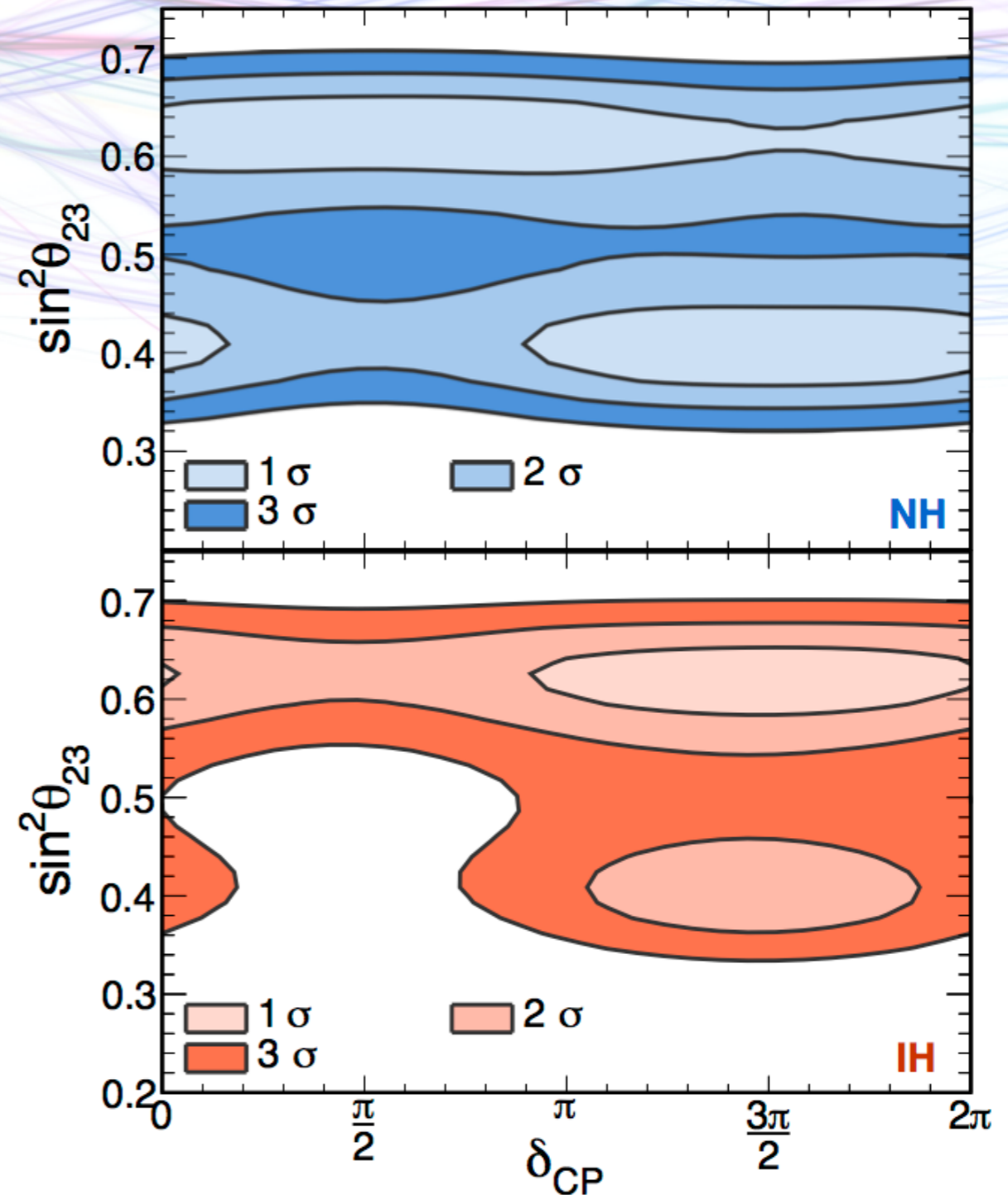
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ν_e fit results

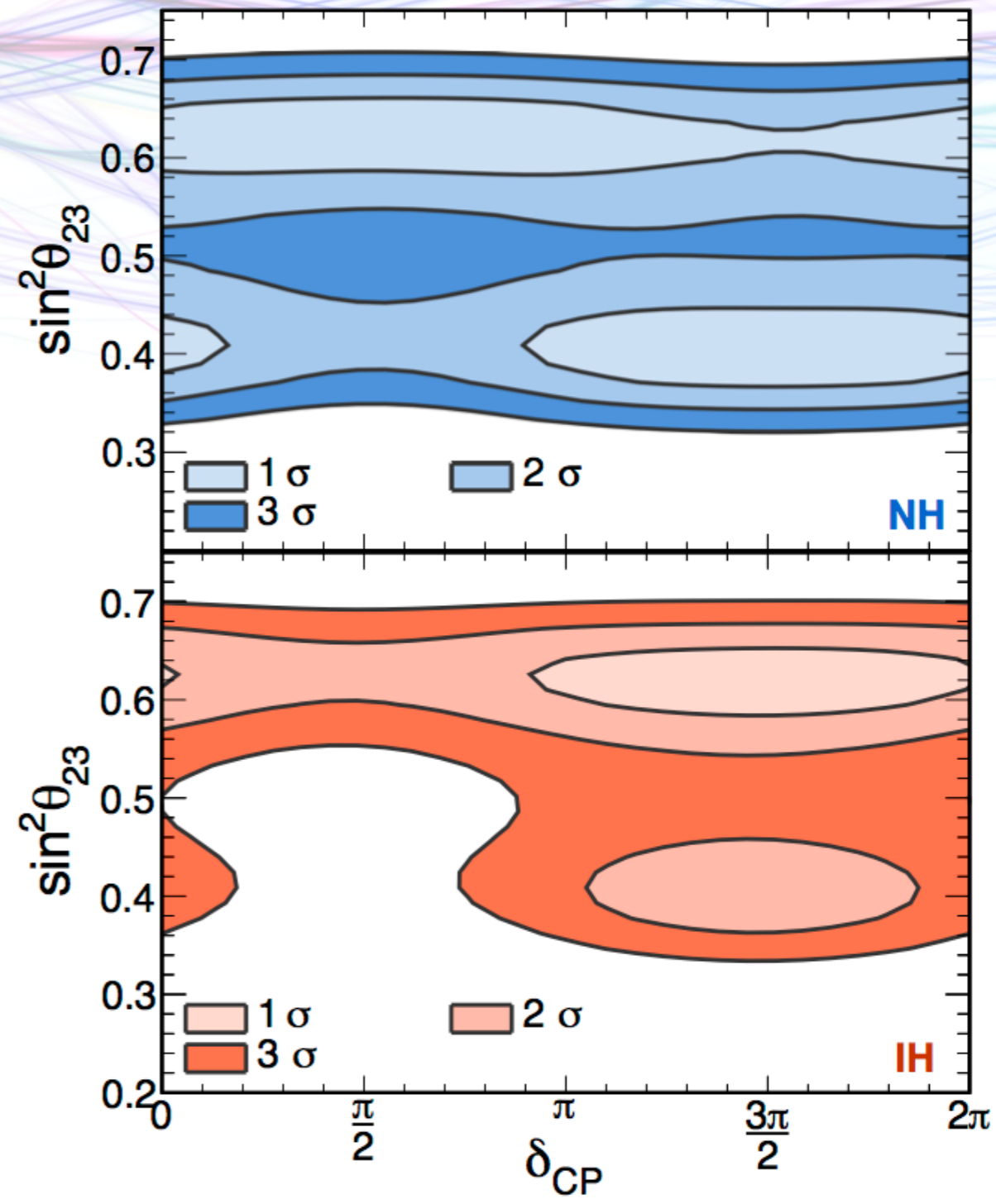
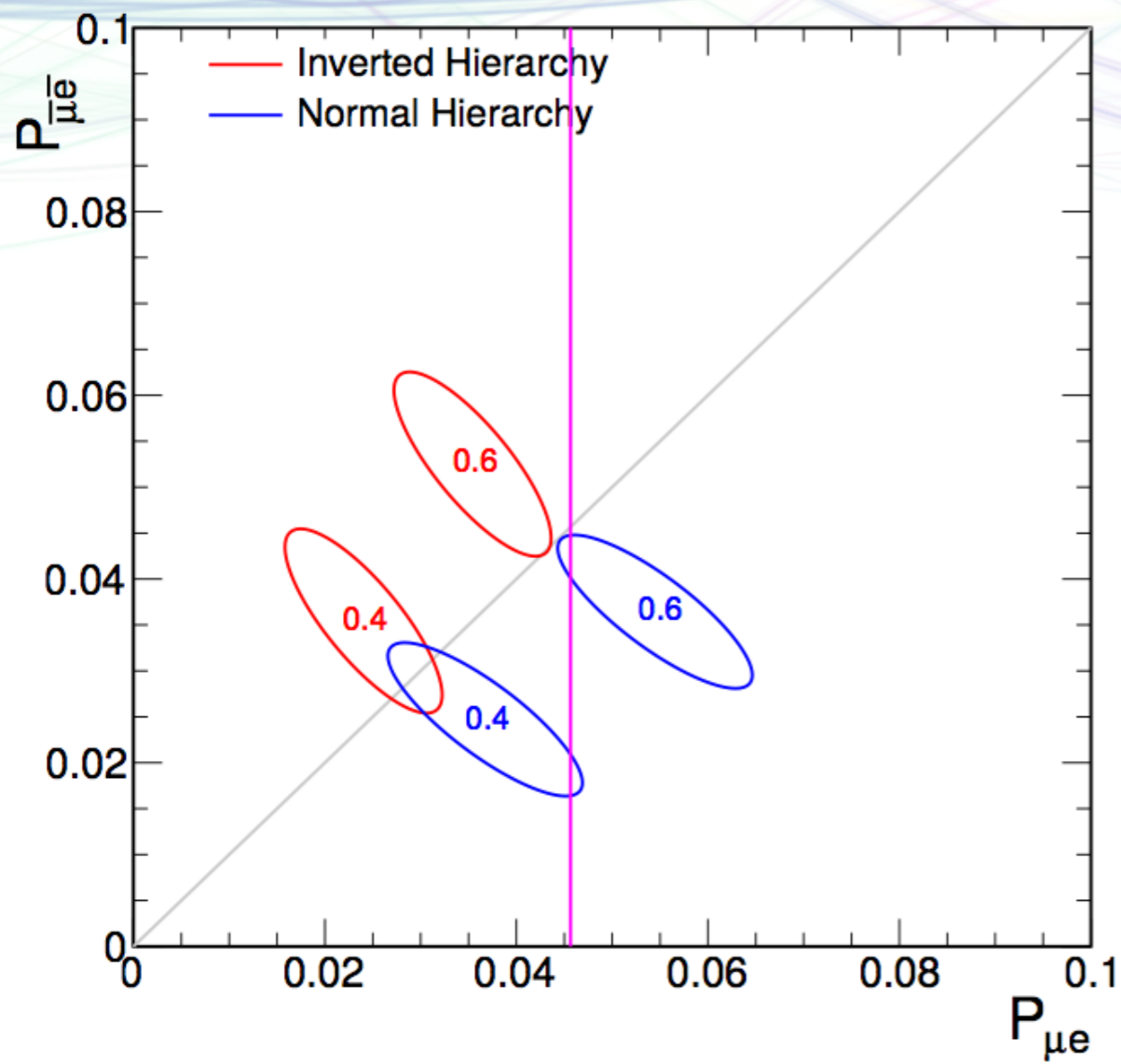
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- ▶ Add $\Delta m_{32}^2/\theta_{23}$ results from ν_μ analysis
- ▶ Not a full joint fit. No syst./osc. param correlations. No FC.
- ▶ Prefer NH, not statistically significant $\Delta\chi^2 = 0.46$
- ▶ Exclude region in IH, lower octant, around $\delta_{CP} = \pi/2$ at 3σ

NOvA Preliminary



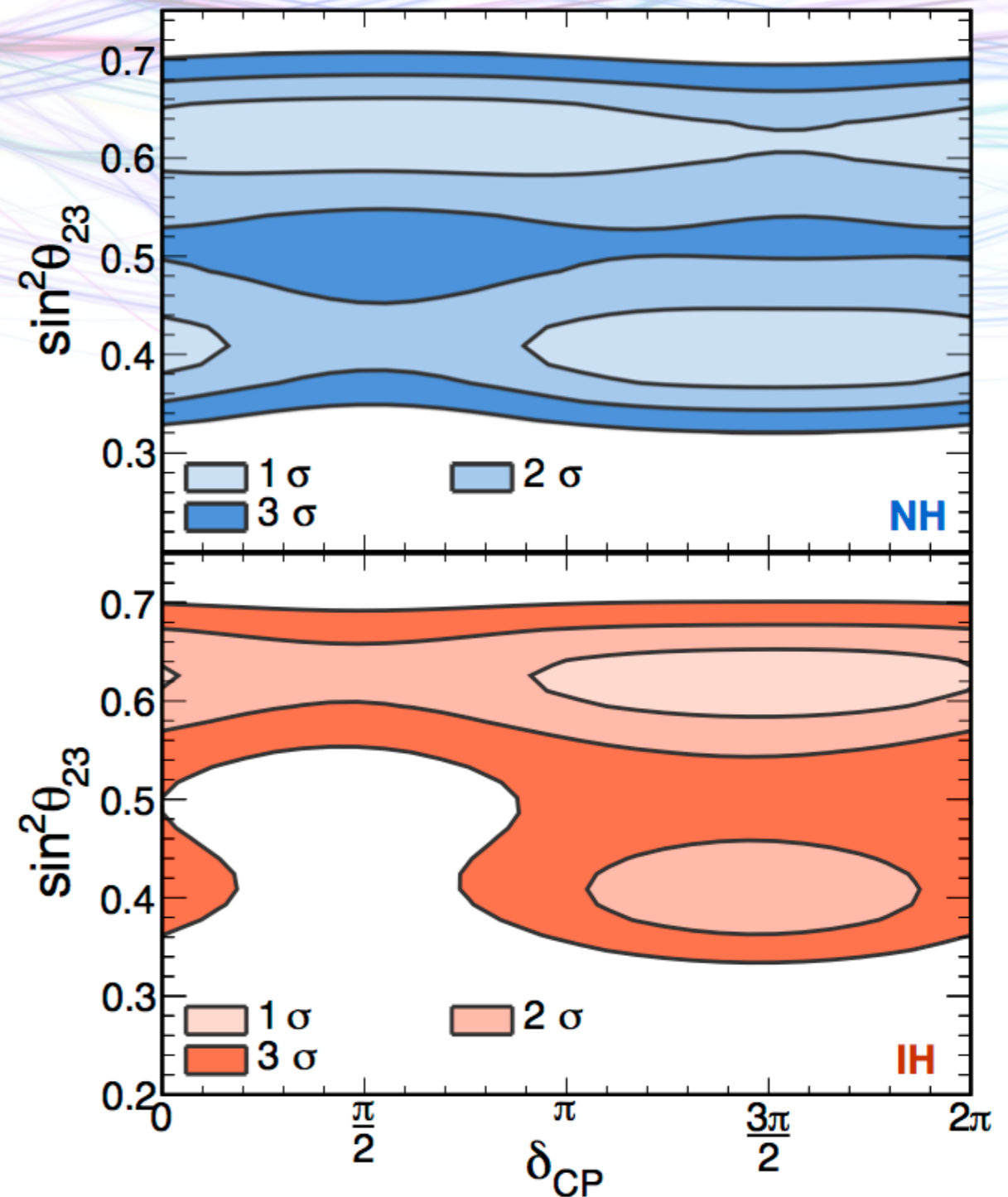
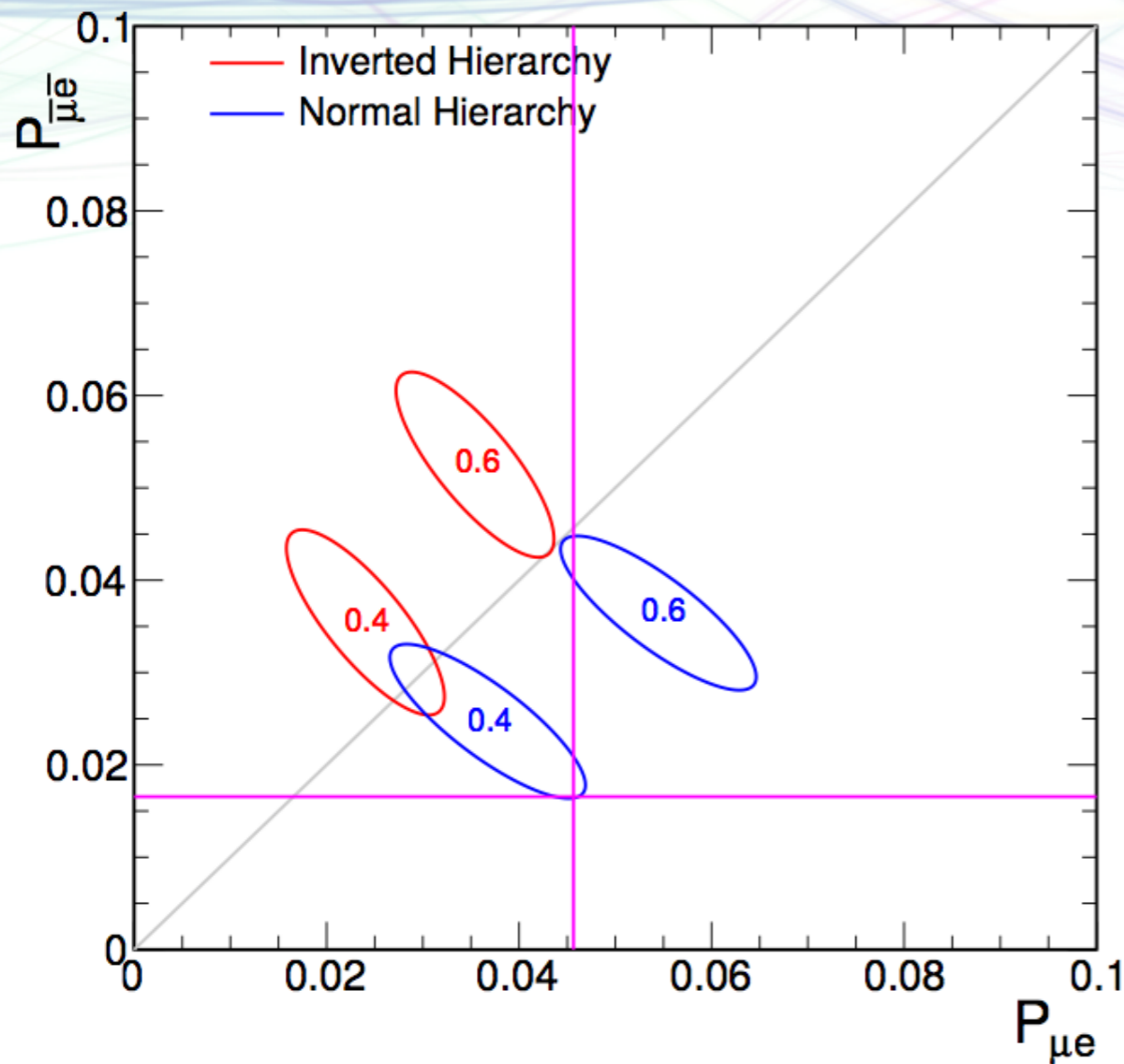
ν_e fit results

NOvA Preliminary



ν_e fit results

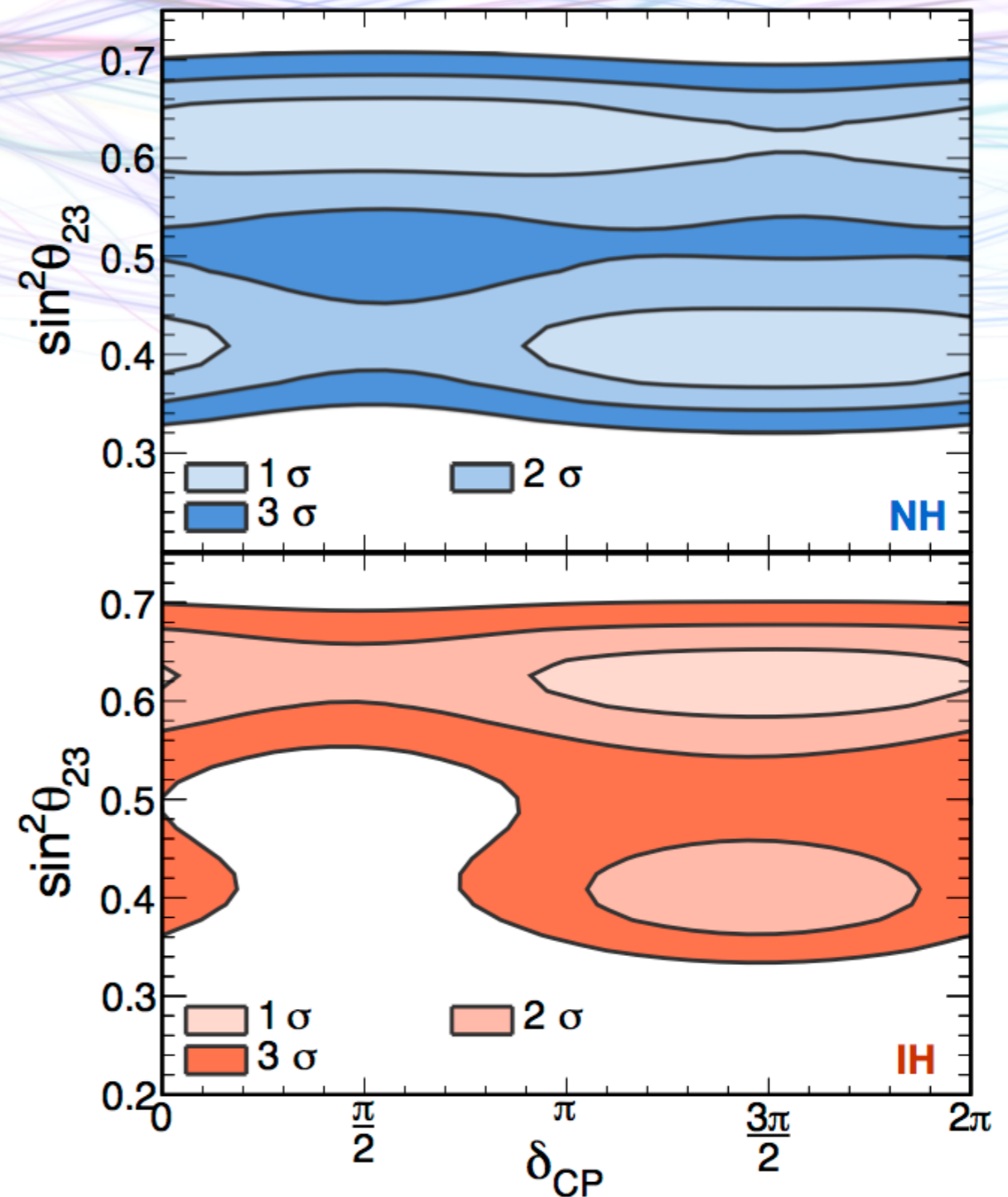
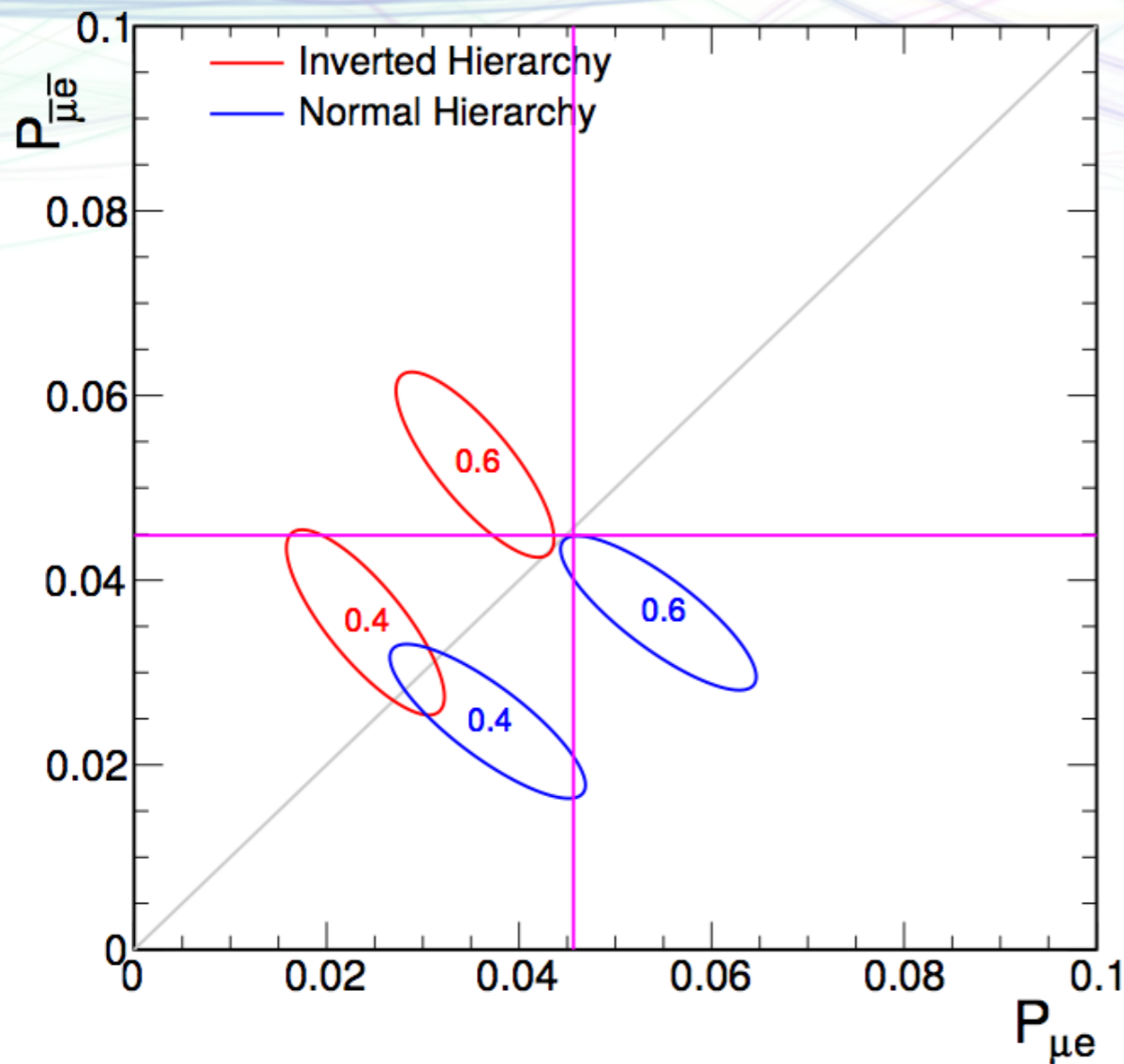
NOvA Preliminary



- ▶ Antineutrino data will help to resolve degeneracies
- ▶ $> 2\times$ difference in $\bar{\nu}_e$ rate between solutions

ν_e fit results

NOvA Preliminary

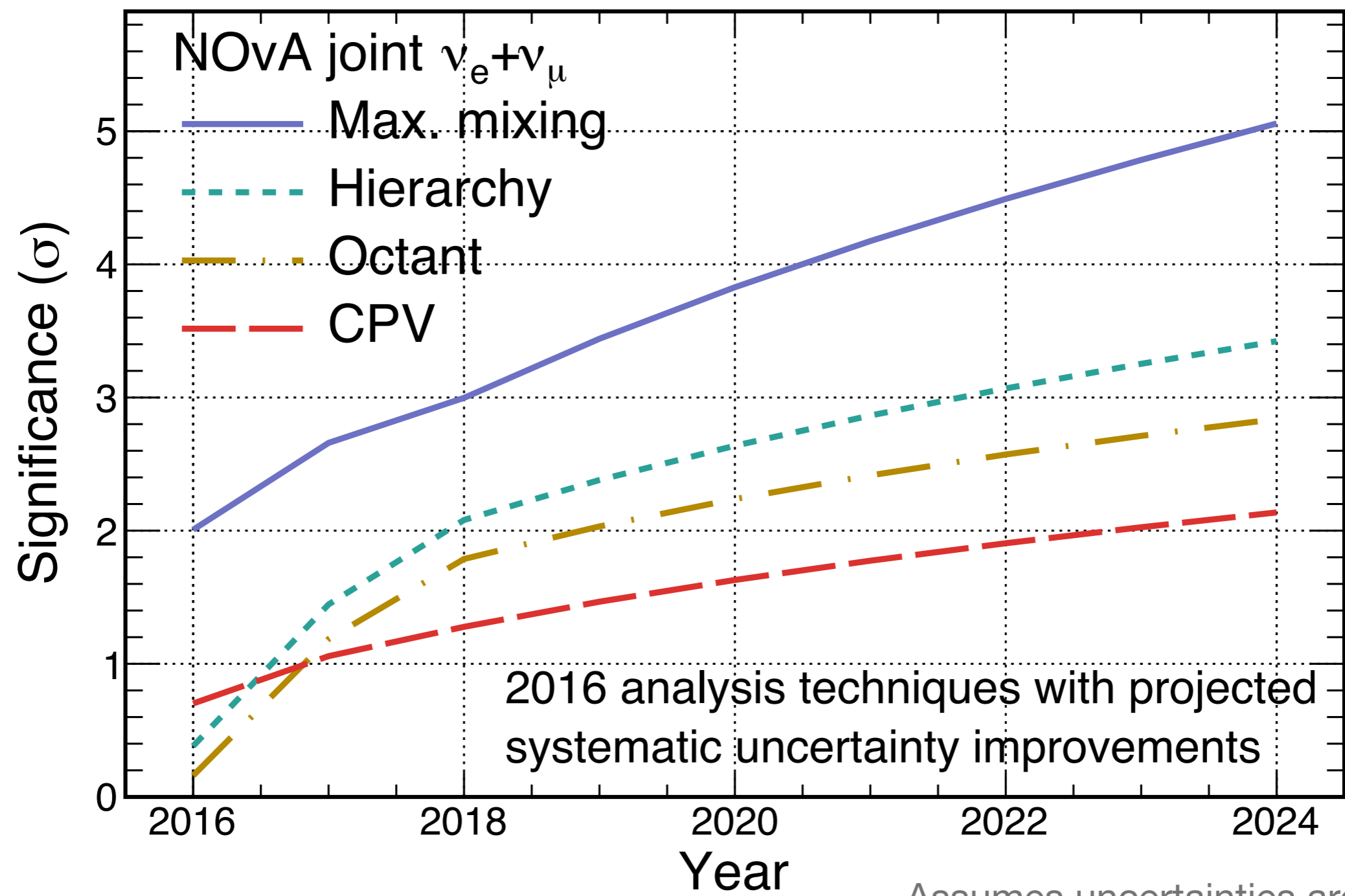


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Looking ahead

NOvA Simulation

Normal $\delta_{CP}=3\pi/2$, $\sin^2\theta_{23}=0.403$
 $\Delta m_{32}^2=2.5\times 10^{-3} \text{eV}^2$, $\sin^2\theta_{13}=0.022$



Projected NOvA physics reach
50/50 run plan for normal hierarchy
lower octant

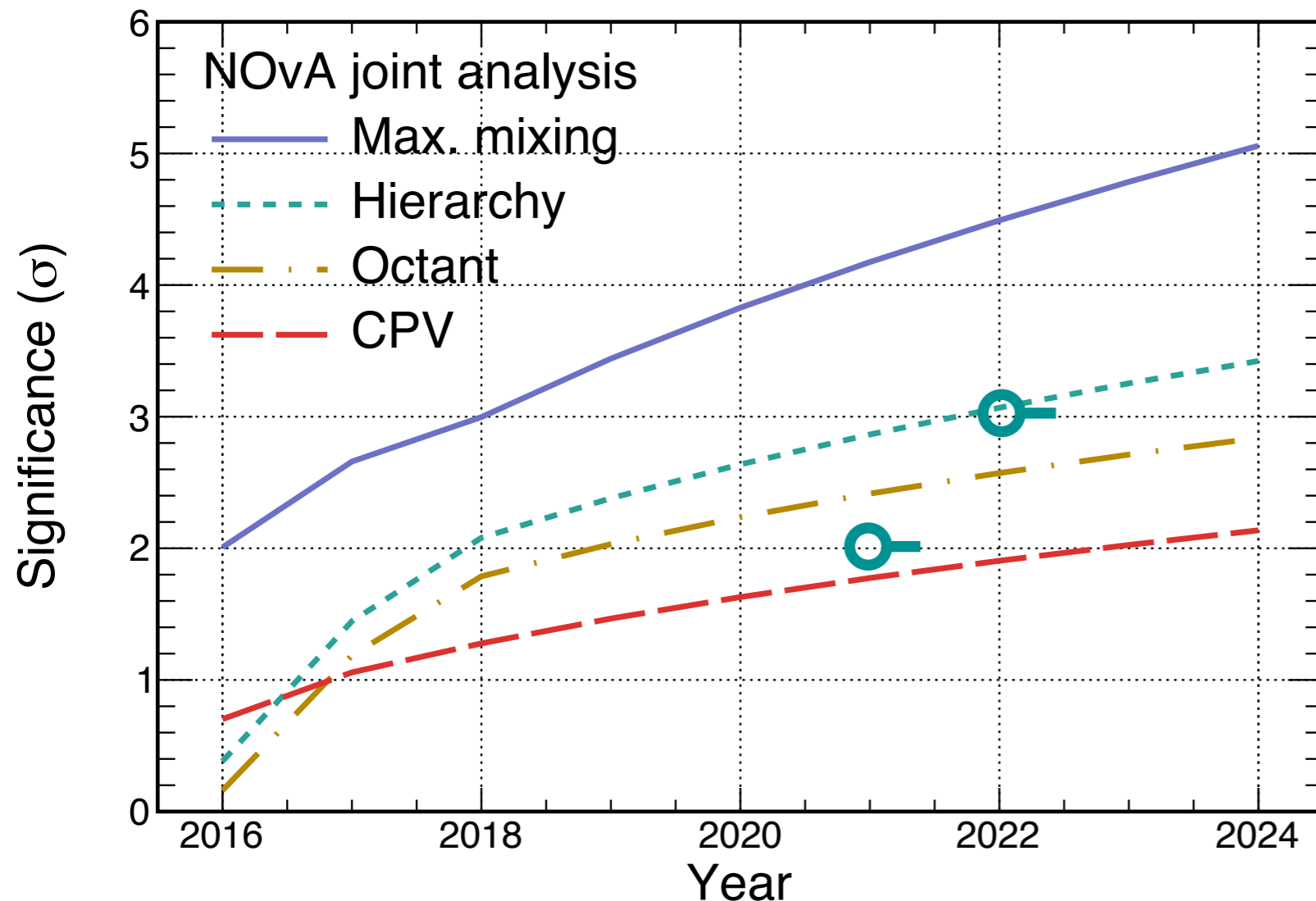
Assumes uncertainties are reduced to

- 2% signal / 5% background for electron neutrinos
- 2% muon energy scale, 3% hadronic energy, ~0% NC backgrounds for muon neutrinos

Competition

This opportunity is not unique to Fermilab. There are several projects hoping to capitalize on this opportunity.

$$\text{NH } 3\pi/2, \sin^2\theta_{23}=0.403, \Delta m_{32}^2=2.5\times 10^{-3}\text{eV}^2, \sin^2\theta_{13}=0.022$$



Both **JUNO** and **ORCA** have construction underway.

Nearly identical schedules for mass hierarchy reach:

- 2σ in 2021
- 3σ in 2022

A Super-K + T2K combination gives roughly 2σ

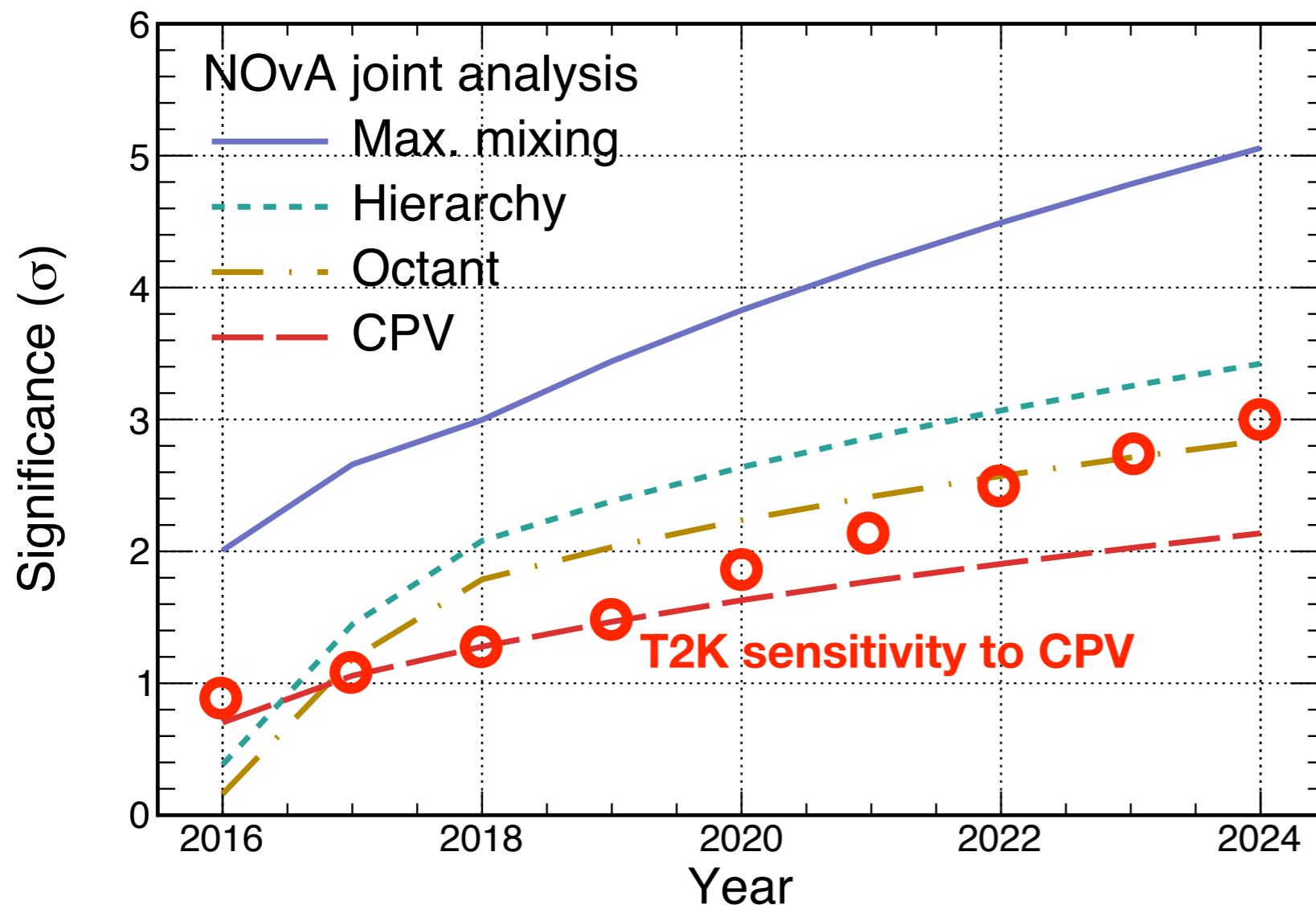
Other competition from, global fits, and cosmology fits.

Competition

T2K has proposed an extended run to get 3σ sigma evidence for CPV

(arXiv:1607.08004v1 [hep-ex] 27 Jul 2016)

NH $3\pi/2$, $\sin^2\theta_{23}=0.403$, $\Delta m_{32}^2=2.5\times 10^{-3}\text{eV}^2$, $\sin^2\theta_{13}=0.022$



Until 2020 NOvA running flat-out and T2K have same CPV reach.

T2K beam power ramps from current 420 kW to 770 kW by 2020 (surpassing NuMI power) and then to 1.1+ MW by 2023. Assumes 5 months / year beam allocation for T2K

This plus analysis improvements drives the CPV reach of T2K to 3 sigma in 2024.