

The influence of external radiation on the Josephson junction + nanomagnet system

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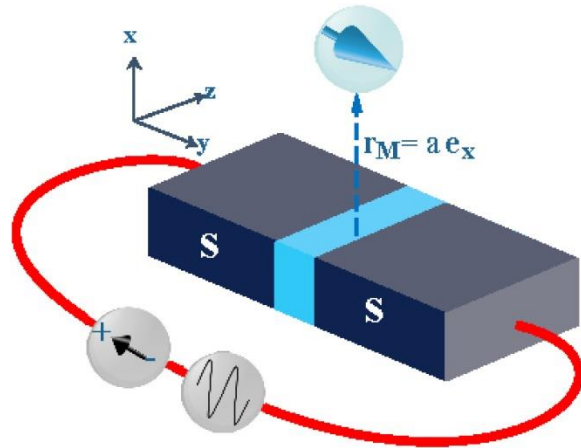
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System of nanomagnet + Josephson junction



The dynamics of magnetic moment can be described by Landau-Lifshitz-Gilbert (LLG) equation

$$\frac{d\mathbf{M}}{dt} = \gamma \mathbf{H}_{eff} \times \mathbf{M} + \frac{\alpha}{M_0} \left(\mathbf{M} \times \frac{d\mathbf{M}}{dt} \right)$$

Effective magnetic field acting on the nanomagnet

$$\begin{aligned} h_x &= 0, && \text{Easy axis} \\ h_y &= m_y, && \text{AC Magnetic field} \\ h_z &= \tilde{h}_z - \delta\epsilon k \dot{m}_z && \text{generated by the} \\ &&& \text{Josephson junction} \end{aligned}$$

where

$$\begin{aligned} \tilde{h}_z &= \epsilon \left[\sin(V\tau - km_z) + \frac{A}{\Omega} \sin(\Omega\tau) \right] \\ &+ (V + A \cos(\Omega\tau) - \beta_c A \Omega \sin(\Omega\tau)) \end{aligned}$$

$$k = (2\pi/\Phi_0) \mu_0 M_s l / a \sqrt{l^2 + a^2} \longrightarrow \text{Coupling parameter}$$

$$l \longrightarrow \text{Length of the JJ}$$

$$a \longrightarrow \text{Distance between nanomagnet and JJ}$$

$$\epsilon = Gk, \quad G = \epsilon_J / K_{an} v \longrightarrow \text{Josephson to magnetic energy ratio}$$

Stability position in the case without periodic drive

LLG equations in spherical form are given by:

$$\dot{\theta} = -\frac{\Omega_F}{(1+\alpha^2)} \frac{\sin\theta}{\left(1 + \frac{\Omega_F \alpha \epsilon k}{1+\alpha^2} \sin^2\theta\right)} \left[\alpha \tilde{h}_z - \sin\phi (\cos\phi + \alpha \cos\theta \sin\phi) \right],$$

$$\dot{\phi} = \frac{\Omega_F}{\alpha^2 + 1} \frac{1}{1 + \frac{\alpha \epsilon k \Omega_F \sin^2\theta}{\alpha^2 + 1}} \left[\tilde{h}_z - \left(-\sin^2\theta \cos\phi \epsilon k \Omega_F + \sin\phi \cos\theta - \alpha \cos\phi \right) \sin(\phi) \right]$$

When V and Ω is much greater than the characteristic times of the variation of θ and ϕ , the Kapitza pendulum forms. To study this, we use the notations

$$\theta \equiv \Theta + \xi \quad \text{and} \quad \phi \equiv \Phi + \zeta$$

Here, Θ and Φ describe the “slower” motion, relevant on longer time scales, whereas the variables ξ and ζ describe the “fast” oscillations of the system. Taking into account that

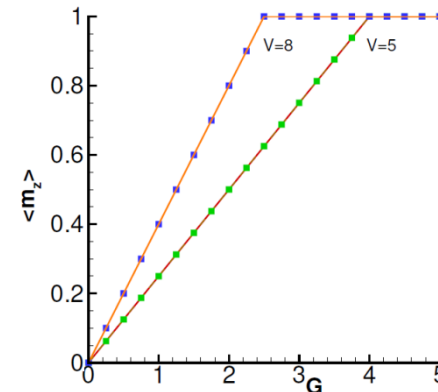
$$\sin[Vt - k \cos\Theta + \frac{A}{\Omega} \sin(\Omega t)] = \sum_{m=-\infty}^{\infty} \text{sign}^m(m) J_{|m|} \left(\frac{A}{\Omega} \right) \sin[(V + m\Omega)t - k \cos\Theta],$$

$$\cos[Vt - k \cos\Theta + \frac{A}{\Omega} \sin(\Omega t)] = \sum_{m=-\infty}^{\infty} \text{sign}^m(m) J_{|m|} \left(\frac{A}{\Omega} \right) \cos[(V + m\Omega)t - k \cos\Theta].$$

We obtain for $A = 0$ the stability position of the magnetic moment which is given by:

$$\Phi = \pi/2 \quad \text{or} \quad \Phi = 3\pi/2.$$

$$\langle m_z \rangle = \cos\Theta = \epsilon \delta V + \frac{\alpha \epsilon^2 k \sin^4 \Theta \Omega_F}{2V(1 + \alpha^2 + \delta \alpha \epsilon k \sin^2 \Theta \Omega_F)^2}$$



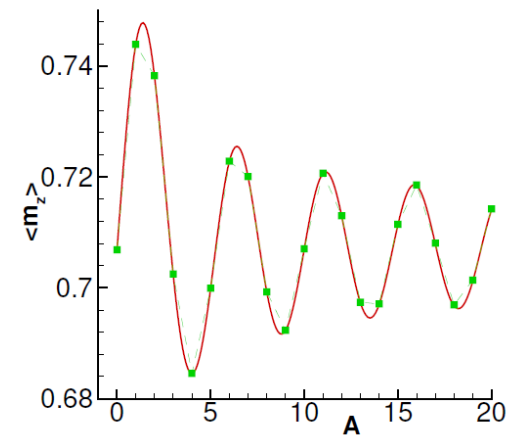
Stability position under external drive

In the 0th order of approximation we have terms that oscillate with frequencies $V + m\Omega$. If there is a negative integer m_0 , such that $V + m_0\Omega = 0$, then we obtain

$$\begin{aligned}\dot{\Theta}_0 &= \frac{\Omega_F}{1 + \alpha^2 + \alpha\delta\epsilon k\Omega_F \sin^2 \Theta} \left\{ (\alpha \sin \Phi \cos \Theta + \cos \Phi) \sin \Theta \sin \Phi \right. \\ &\quad \left. - \alpha\epsilon \sin \Theta \left[\delta V - \text{sign}^{m_0}(m_0) J_{m_0} \left(\frac{A}{\Omega} \right) \sin(km_z) \right] \right\}, \\ \dot{\Phi}_0 &= \frac{\Omega_F}{1 + \alpha^2 + \alpha\delta\epsilon k\Omega_F \sin^2 \Theta} \left\{ (\delta\epsilon k\Omega_F \sin^2 \Theta \cos \Phi - \cos \Theta \sin \Phi + \alpha \cos \Phi) \sin \Phi \right. \\ &\quad \left. + \epsilon \left[\delta V - \text{sign}^{m_0}(m_0) J_{m_0} \left(\frac{A}{\Omega} \right) \sin(km_z) \right] \right\},\end{aligned}$$

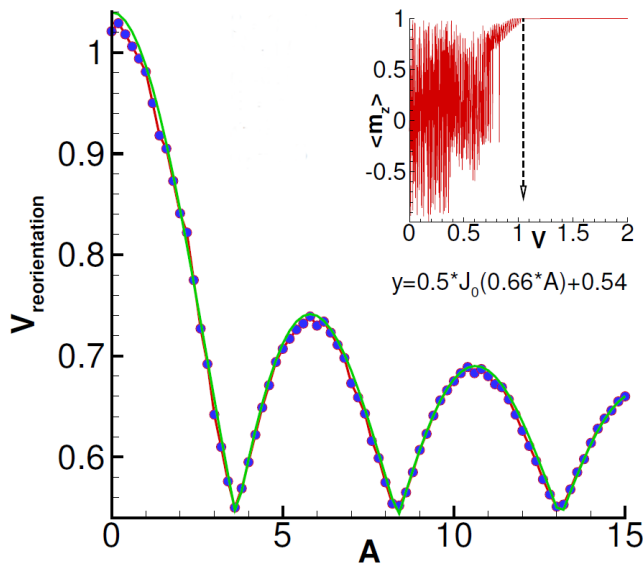
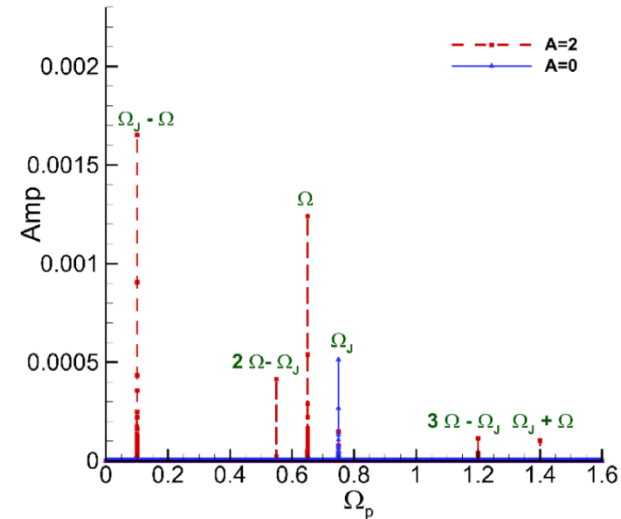
As in the case $A = 0$, the equation $\dot{\Theta} = \dot{\Phi} = 0$ lead to $\Phi = \pi/2$ or $3\pi/2$ and the equation for Θ :

$$\cos \Theta = \epsilon\delta V - \epsilon \text{sign}^{m_0}(m_0) J_{m_0} \left(\frac{A}{\Omega} \right) \sin(k \cos \Theta)$$



Influence of external radiation

The presence of the terms that oscillate in time may be emphasized by the FFT of $m_z(t)$. The oscillating frequencies should be $|n\dot{V} + m\Omega|$, where n, m are integers and $n \geq 0$.



The applied external periodic drive also affects the voltage of complete reorientation V_r , which indicates the stabilization of the magnetic moment dynamics (see inset). The numerical data is well fitted by the Bessel function $0.5J(A/\Omega)+0.54$ shown by green solid line

Summary

- We obtained an analytical expressions to describe the movement of the stability position in the yz plane and had a very good agreement with the direct numerical simulations.
- It was shown that the average value of magnetic moment component m_z as a function of amplitude of external drive demonstrate Bessel behavior.
- We emphasis that influence of external periodic drive on the Kapitza-like pendulum effects on the orders of magnitude more pronounced when the Josephson frequency V is equal to integer number of external drive frequencies Ω ($V + m_0\Omega = 0$). Otherwise, it is very small and the Kapitza-like pendulum effects come mainly from the Josephson oscillations.
- We also shown that the voltage of complete reorientation depend as the Bessel function on the external drive amplitude.

Thank you for attention.