







Study of strongly interacting matter properties at the energies of the NICA collider using the methods of femtoscopy and factorial moments

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People:

- Ludmila Malinina (SINP MSU, JINR),
- Konstantin Mikhaylov (ITEP & JINR), convener
- Grigory Nigmatkulov (NRNU MEPhl),
- Olga Kodolova (SINP MSU),
- Igor Lokhtin (SINP MSU),
- Gleb Romanenko (student, MSU),
- Marya Cheremnova (student, MSU)
- Yevheniia Khyzniak (PhD student, NRNU⁴ MEPhI)
- Chernishov Alexey (student MSU)



Motivation: Phase diagram QCD

- Crossover transition to QGP occurs at RHIC & LHC
- 1st order phase transition to QGP occurs at lower energies (?)



 Beam Energy Scan (BES) program: search for location of transition point from Crossover (XPT) to 1st order (1PT) phase transitions

- BES RHIC (√s=3-60 GeV)
- NA61@SPS (E_{lab}=10-158 GeV);

projects: CBM@FAIR (GSI) и MPD@NICA (JINR)





Hybrid (hydro+hadron gas) vHLLE+UrQMD model



Iu. Karpenko, P. Huovinen, H.Petersen, M. Bleicher, Phys.Rev. C.91, 2015, 064901

Pre-thermal phase

UrQMD

The transition to hydrodynamics occurs at a hyper-surface of constant longitudinal proper time T_0 Minimal value of the starting time T0 is taken to be equal to the average time for the two colliding nuclei to completely pass through each other:

Parameters τ_0 , R_{\perp} , R_{η} and η/s adjusted using basic observables in the RHIC BES-I region.

Model tuned by matching with existing experimental data from SPS and BES-I RHIC Hydrodynamic phase

vHLLE (3+1)-D viscous hydrodynamics

EoS to be used in the model

- Chiral EoS crossover transition J. Steinheimer et al., J. Phys. G 38, 035001 (2011)
- Hadron Gas + Bag Model 1st-order phase transition
 P. F. Kolb et al., Phys.Rev. C 62, 054909 (2000)

Hydrodynamic phase lasts longer with 1PT, especially at lower energies but cascade smears this difference.

Hadronic cascade

UrQMD

Fluid to particle transition at energy density = 0.5 GeV/fm³

Pion emission time

after hydrodynamic phase



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Examples of vHLEE+UrQMD with XPT calculations for bulk observables





Adjustment to experimental data suggests $\eta/s = 0.2 \rightarrow 0.08$ when $\sqrt{s} = 7.7 \rightarrow 200$ GeV, modulo initial state (UrQMD) and EoS (Chiral model) used.

Iu. Karpenko, P. Huovinen, H.Petersen, M. Bleicher, Phys.Rev. C 91, 064901 (2015), arXiv:1502.01978,1509.3751 , talk QM2015



Femtoscopy



<u>Correlation femtoscopy :</u>

Measurement of space-time characteristics \mathbf{R} , $\mathbf{c\tau}$ of particle production using particle correlations due to the effects of quantum statistics (QS) and final state interactions (FSI)

Two-particle correlation function:

theory:

 q_{side}

 q_{long}

 $k_{\rm T} = |p_{\rm T,1} + p_{\rm T,2}|/2$

experiment:

 $C(q) = \frac{N_{2}(p_{1}, p_{2})}{N_{1}(p_{1}) \cdot N_{2}(p_{1})}, C(\infty) = 1$ $C(q) = \frac{S(q)}{B(q)}, q = p_{1} - p_{2}$

S(q) – distribution of pair momentum difference from same event B(q) – reference distribution built by mixing different events

Parametrizations used:

1D CF: $C(q_{inv}) = 1 + \lambda e^{-R^2 q_{inv}^2}$ *R* – Gaussian radius in PRF, λ – correlation strength parameter

3D CF: $C(q_{out}, q_{side}, q_{long}) = 1 + \lambda e^{-R_{out}^2 q_{out}^2 - R_{side}^2 q_{side}^2 - R_{long}^2 q_{long}^2}$

R and *q* are in Longitudinally Co-Moving Frame (**LCMS**) long || beam; out || transverse pair velocity v_{T} ; side normal to out,long

Rside

<u>3D analysis:</u>

 R_{side} sensitive to geometrical transverse size. R_{long} sensitive to time of freeze-

out.

 R_{out} / R_{side} sensitive to emission duration.

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3D Pion radii with vHLLE model

Phys.Rev. C96 (2017) no.2, 024911



- Femtoscopic radii are sensitive to the type of the phase transition
- Crossover EoS describes better R(mT) dependencies
- $R_{out,long}$ (1PT) > $R_{out,long}$ (XPT) by value of ~1-2 fm.
- R_{out}/R_{side} (XPT) agrees with STAR data points at 7.7 and 11.5 GeV, but then increases with increasing collision energy while ratio in data is independent with collision energy; R_{out}/R_{side} for 1PT systematically overestimate data and is independent on collision energy.

3D Pion Correlation Functions with vHLLE model

Analysis

- MDP femto package (see Grigory Nigmatkulov's talk)
- We performed analysis of pion and kaon correlations for vHLLE and UrQMD for different centrality ranges
- Au+Au, √s_{NN} = 11.5 GeV
 central events vHLLE
- Standard 3D Gaussian fit used
- Projections of 3D pion correlation functions on out-side-long
- XPT CF projections on long direction ^{1.4} are wider than 1PT ones
- CF become wider with increasing $k_{_{\rm T}}$



1PT -green dots; XPT – red dots

Pion radii with vHLLE and UrQMD models



Pion and kaon radii with vHLLE model



• Approximate m_{τ} scaling for R_{side}

Similarly to pions : kaon radii decrease with $m_{\tau} \rightarrow$ radial flow ;

• for 1PT EoS almost flat dependence R_{out} (m_T) is observed \rightarrow weaker flow

• $R_{out,long}$ (1PT) > $R_{out,long}$ (XPT)

• R_{long} kaon radii for XPT > R_{long} pion similarly to experiment (LHC & RHIC)

 Very different predictions of vHLLE model for different $EoS \rightarrow importance$ to study heavier than pions particles \rightarrow kaons

The similar trend is observed for AuAu 11.5 GeV

Pion and kaon radii with UrQMD model



 kaon radii demonstrate almost flat behavior similarly to vHLEE with the 1PT EoS →weak flow

R_{long} kaon radii are larger
 than pion ones similarly to
 experiment (LHC & RHIC)

 The similar trend is observed for AuAu 11.5 GeV

Energy dependence of pion radii



• NICA energy range: $\sqrt{s_{NN}} = 4 - 11 \text{ GeV}$

 We performed femtoscopic analysis of pions in this energy range using vHLEE and UrQMD models

• There is a lack of systematical high accuracy and high statistics studies of QGP phase transition for this region in the world !

Pion radii with vHLLE and UrQMD models



• Au+Au, √s_{NN} = 11.5 GeV

- It is important to study

 k_T (m_T) dependence
 in larger interval than STAR
- Radii decrease with $m_{\tau} \rightarrow$ radial flow
- Increase size with increasing centrality →simple geometric picture of the collisions.
- Crossover EoS describes better R_{out}(m_T)
- $R_{out,long}$ (1PT) > $R_{out,long}$ (XPT)
- URQMD is closed to 1PT
- The similar trends is observed for AuAu 11.5 GeV

Factorial Moments (see Olga Kodolova's talk)

It was proposed by A. Bialas and R. Peschanski (Nucl. Phys. B 273 (1986) 703) to study the dependence of the normalized factorial moments

$$F_{i} = M^{i-1} \times \left\langle \frac{\sum_{j=1}^{M} k_{j} \times (k_{j}-1) \times \dots \times (k_{j}-i+1)}{N \times (N-1) \times \dots \times (N-i+1)} \right\rangle$$

Note: there is a set of definitions of moments and cumulants.

of the rapidity distribution on the size $\delta y (\Delta y/M, M$ is the number of bins, Δy is the size of the mid rapidity window):

1. if fluctuations are purely statistical no variation of moments as a function of δy is expected

2. Observation of variations indicates the presence of physics origin fluctuations

Intermittency (fluctuations of various different sizes in 1D, 2D and 3D phase space) have been studied at LEP, Tevatron, Protvino in ee, hh, hA, AA interactions at the various energies.

Au-Au, UrQMD+vHLLE

