Долговременные измерения потоков космических лучей в земной атмосфере

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Введение

Регулярные измерения потоков заряженных частиц в атмосфере были начаты в июне 1957 г. (JGY).



Академик С.Н. Вернов



Профессор А.Н. Чарахчьян



(11.07.1910 - 26.09.1982)





India Ocean expedition, 1949



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Table 1. The sites and periods of regular measurements of CR fluxes in the atmosphere

Site of measurements	Geographical coordinates	R _c (GV)	Period of measurements	
Murmansk region	67°33′N; 33°20′E	0.6	07.1957–present time	
Dolgoprudny, Moscow region	55°56'N; 37°31'E	2.4	07.1957–present time	
Alma-Ata, Kazakhstan	43°15′N; 76°55′E	6.7	03.1962-04.1993	
Mirny, Antarctica	66°34′S; 92°55′E	0.03	03.1963–present time	
Simeiz, Crimea	44°00'N; 34°00'E	5.9	03.1958–12.1961	
			03.1964–04.1970	
Voyeikovo, Leningrad region	60°00′N; 30°42′E	1.7	11.1964–03.1970	
Norilsk, Krasnoyarsk region	69°00'N; 88°00'E	0.6	11.1974–06.1982	
Yerevan, Armenia	40°10′N; 44°30′E	7.6	01.1976-05.1989	
Tixie, Yakutiya	71°36′N; 128°54′E	0.5	02.1978-09.1987	
Dalnerechensk,	45°52′N; 133°44′E	7.35	08.1978-05.1982	
Khabarovsk region				
Vostok station, Antarctica	78°47′S; 106°87′ E	0.00	01.1980-02.1980	
Barentzburg, Norway	78°36'N; 16°24'E	0.06	05.1982, 03–07.1983	
Campinas, Brazil	23°00′S; 47°08′W	10.9	01.1988-02.1991	
Sea expeditions	60°00'N-70°00'S	0.1–17	1963–1987	

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Радиозонд для измерения потоков заряженных частиц в атмосфере



Cosmic ray fluxes in the atmosphere vs altitude



count/min





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E, MeV/n



Year



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Solar proton events 9, 14 and 16 July 1959: left panel – altitudinal dependences of charged particle fluxes; right panel – energy spectra of solar protons. The measurements were made at the northern polar latitude ($R_c = 0.5$ GV) in the period of 9 – 17 July 1959.



Intrusion of SEPs into the polar atmosphere. Right axis shows the energy of protons penetrating down to a given altitude. The protons with E < 500 MeV are just absorbed in the atmosphere due to ionization loss.

35th COSPAR Scientific Assembly

Charged particle fluxes in the polar atmosphere (solar and galactic cosmic rays)



Integral energy spectrum of solar protons 20.01.2005, 08 UT





Year

Solar proton events recorded in the stratosphere. Vertical lines mark the time of event and heights of vertical ones give the solar proton intensity (protons with $E \ge 100$ MeV). Red curve – sunspot number.

From 1958 till now 115 solar proton events were recorded in the stratosphere.

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Geomagnetic Locations of Stations



Examples of electron precipitation events observed in the stratosphere



Three minute averages of count rates of single counters during (1) EPEs and (2) quiet conditions, and (3) telescopes versus residual atmospheric depth. Count rates of telescopes are multiplied by 5. On the left panel is EPE of September 26, 1997; on the right panel is EPE of October 9, 1998.

Photon spectra in the atmosphere and

precipitating electron spectra at the top of the atmosphere

DATE	A _{ph,o} (cm ⁻² s ⁻¹ sr ⁻¹)	E _{ph,o} (keV)	$\mathbf{B}_{\mathrm{oe}},(\mathrm{cm}^{-2}\mathrm{s}^{-1}\mathrm{sr}^{-1})$	E _{oe} (keV)
8 March 1994	31	50	2816	150
15 March 1994	798	30	5E5	48
5 April 1994	335	16	1E5	21
7 May 1994	5	34	2015	60
13 May 1994	5.5	30	3093	48

 $dN_{ph}/dE = A_{ph,o} * exp(-E/E_{ph,o})$

 $dN_e/dE = B_{oe}^* \exp(-E/E_{oe})$

EPEs observed at Murmansk polar region in 1958-2010



Yearly numbers of electron precipitation events in the Murmansk region. Sunspot number (blue curve) is given to outline the 11-year solar activity cycles.

Seasonal effect in the EPE occurrence rate.



The monthly occurrence rate (in % per month) of the Electron Precipitation Events observed in the atmosphere at Olenya, Murmansk region during 1970-1987. Total number of EPEs is 240.

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Count rate, particles/min



On 14 October 1970 the power nuclear explosion on the Earth's surface was produced at Lobnor's poligon in China. Radioactive cloud from this explosion was recorded in the atmosphere at Murmansk, Moscow and Alma-Ata duing October and November. The radioactive cloud was observed in the atmosphere more than 1 month. In Murmansk region the increase of count rate was in several times in comparison with natural background. For the period from 10 o'clock 25 October till 16 o'clock 26 October 14 launches of radiosounds were made (here and below the Moscow time is used). At the previous slide the data for 25 October are shown with solid curves and for 26 October - with dashed curves. The maximum of count rate was recorded at 17 o'clock 25 October and it excesses the background in 7.5 times. The particle fluxes were increased till 17 o'clock 25 October and after this time the gradual decrease of particle fluxes was observed. This gradual decrease took place up to 16 o'clock 26 October.

The maximum density of radionuclides was recorded on 25 October at the altitudes of 18 – 24 km (atmosphere pressure is 30–70 g/cm²). During the next day the position of maximum decreased at 20–18 km.

The radioactive cloud had a flat disk form at the altitudes 18 – 24 km. This cloud had strong marked core. The cloud horizontal sizes were evaluated from wind speed in the atmosphere over Murmansk. At the altitudes of 18 – 24 km the wind speed was about 10 m/sec. The cloud was recorded about 30 hours. It gave the horizontal size of cloud along the wind direction (west – east) about 1100 km. As cloud sizes along the wind direction and perpendicular to it are formed the diffusion mechanism, it means that cloud sizes along the wind direction and perpendicular to it have to be equal approximately.

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Ion production rate in the atmosphere (over globe)

Ionization source	Number of ion pairs per second	
Natural radioactivity (h < 3 km)	< 10 ²⁴	
GCRs (all atmosphere)	$\sim (10^{27} - 10^{28})$	
SCRs, precipitation (polar latitudes)	~ $(10^{26} - 10^{27})$	
UV - and X – solar irradiation ($h > 50$ km)	~ 10 ²⁸	
Solar wind (ionosphere)	~ 10 ²⁶	
lightning (regions with thunderstorm activity, <i>h</i> <10 km)	$\sim (10^{27} - 10^{28})$	

Global electric current circuit



surface

Electrical characteristics of the atmosphere: lightning, sprites, electric current and so on.

Electric charge of the Earth $Q \approx -600\ 000\ \mathrm{C}$ **Electric field near the Earth's surface** $E \approx 130 \text{ V/m}$ **Electric current in the atmosphere** $J \approx 10^{-12} \text{ A/m}^2$ $J_{\text{total}} \approx 2000 \,\mathrm{A}$



 $r(N, J) = +0.67 \pm 0.14; r(N, R_z) = -0.32 \pm 0.22$

GCR flux and number of lightning





 $[(n)_{e}/(n)_{m}]^{2} = [(I)_{e}/(I)_{m}] \qquad [(n)_{e}/(n)_{m}] = [(I)_{e}/(I)_{m}]$



Ion balance equation in the atmosphere $\underline{q} = \alpha \cdot n^2$ is wrong.

The right equation is $\underline{q} = \beta \cdot \underline{n}$.

- *q* ion production rate
- α volume recombination coefficient
- β linear coefficient of recombination
- *n* concentration of light ions



Scheme of rise, development and decay of thundercloud: a - initial phase of cloud rise; 6 - mature phase of thundercloud; B - decay phase; J - negative ion current from ionosphereto the cloud top; 1 - front of warm air; 2 - front of cold air; 3 - ascending fluxes of wet air; $4, 5 - extensive air showers produced by particles with energy <math>E > 10^{14}$ eV and $E > 10^{15}$ eV; 6 - intra cloud electric discharges; 7 - cloud - the Earth's surface lightning; 8 - the Earth'ssurface - cloud lightning; 9 - negative electric charge at the top of cloud; 10 - positive electriccharge at the bottom cloud.

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Cosmic ray fluxes in the troposphere at the middle latitude, $R_c = 2.4 \text{ GV}$

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