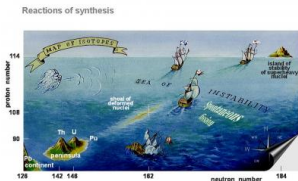


Investigation of the spin-orbit strengths on the prediction of the closed shells for superheavy nuclei based on Two Center Shell Model

A.N.Bezbakh

Joint Institute for Nuclear Research

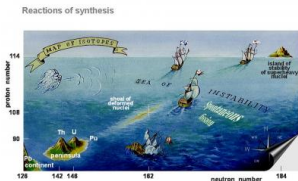
12 October 2021, Almaty, Kazakhstan



- The experimental study of heaviest nuclei can be guided by the theoretical analysis. JINR Superheavy Elements Factory \implies a new era in SHN research.

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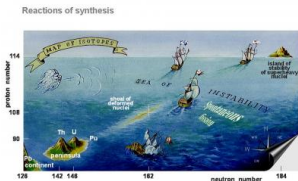
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- The investigation of transfermium elements expands our knowledge of the single-particle structure, location of the shell closures, and decay modes of heaviest nuclei
- Increasing stability of nuclei approaching $N = 184$, and indication quite a large shell effects behind $Z = 114 \implies$ Are valid the predictions of relativistic and nonrelativistic mean-field models in our case ($Z = 120 - 126$, $N = 182 - 184$)? Or the phenomenological model¹ ($Z = 126$)? Note: the mic-mac models² predict $Z = 114$.

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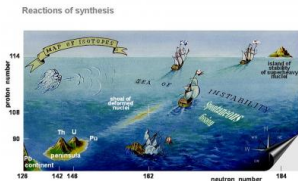
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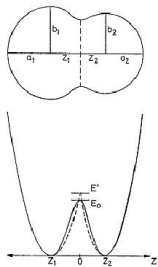
Our Aim

To investigate the role of spin-orbital strengths on the position of the magic shell and how they affect the description of low-lying states

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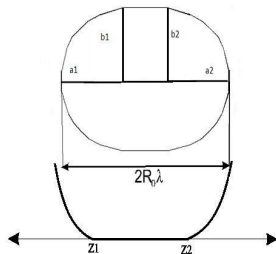
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$$H = (-\hbar/2m)\nabla^2 + V(\rho, z) + V_{l,s} + V_{\rho}^3$$



- $\lambda = L/2R_0$,
- $\beta = a/b = \beta_1 = \beta_2$ the case,
- $\varepsilon = E_0/E' = 0$,
- $\eta = (A_1 - A_2)/(A_1 + A_2) = 0$;

Other variables are fixed.



where the momentum-independent part is $V(\rho, z)$ and the momentum-dependent part consists of

$$V_{ls} = -\frac{2\hbar\kappa}{m\omega'_0} (\nabla V \times \mathbf{p}) \cdot \mathbf{s}$$

$$V_{\rho} = -\kappa\mu\hbar\omega'_0 I^2 + \kappa\mu\hbar\omega'_0 \frac{N(N+3)}{2} \delta_{if}$$

³J. Maruhn and W. Greiner, Z. Phys. A **251**, 431 (1972)

In order to improve the description of spins and parities of the nuclear ground states, we introduce a **weak dependence on $(N - Z)$** in the parameters $\kappa_{n,p}$ and $\mu_{n,p}$. For the actinide and transactinide region we suggest^{4, 5}:

$$\begin{aligned}\kappa_n &= -0.076 + 0.0058(N - Z) - 6.53 \times 10^{-5}(N - Z)^2 + 0.002A^{1/3} \\ \mu_n &= 1.598 - 0.0295(N - Z) + 3.036 \times 10^{-4}(N - Z)^2 - 0.095A^{1/3}\end{aligned}$$

and

$$\begin{aligned}\kappa_p &= 0.0383 + 0.00137(N - Z) - 1.22 \times 10^{-5}(N - Z)^2 - 0.003A^{1/3} \\ \mu_p &= 0.335 + 0.01(N - Z) - 9.367 \times 10^{-5}(N - Z)^2 + 0.003A^{1/3}\end{aligned}$$

With those Eqs. we are able to describe correctly the ground-state spins of many heavy nuclei treated.

Note that the introduced additional dependence on $N - Z$ mainly supplies a better order of the single-particle levels near the Fermi surface.

⁴G.G. Adamian, N.V. Antonenko, and W. Scheid, Phys. Rev. C **81**, 024320 (2010)

⁵A.N. Kuzmina, G.G. Adamian, N.V. Antonenko, and W. Scheid, Phys. Rev. C **85**, 014319 (2012)

$$E = E_{LDM} + \delta E_{mic}$$

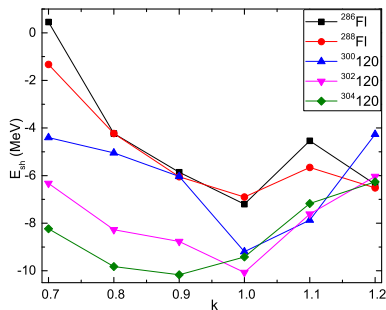
- The Coulomb and surface energies
- The shell E_{sh} and pairing corrections

Note

The stability of SHN correlates with the shell correction energy E_{sh} in the ground state. The larger $|E_{sh}|$, the greater the stability of SHN with respect to spontaneous fission and α -decay.

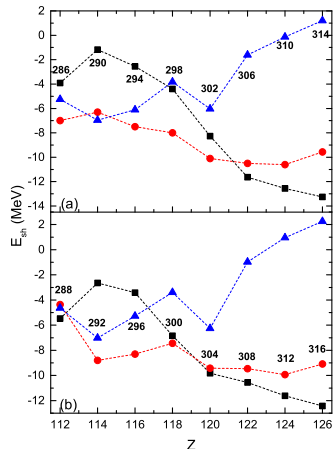
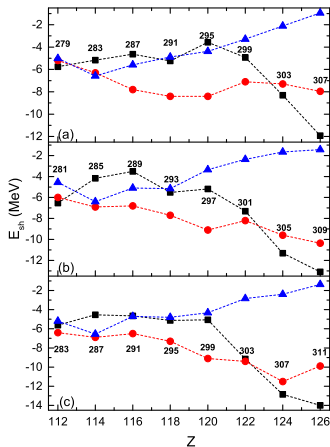
$$H = (-\hbar/2m)\nabla^2 + V(\rho, z) + kV_{l,s} + V_{l^2}$$

In order to study the influence of spin-orbit (sl) strengths in the region of SHN with the modified TCSM, we take the sl -term as $kV_{l,s}$ and study how the results depend on the coefficient k varying from 0.8 to 1.2. The value $k = 1$ corresponds to the parameters defined in Eqs. for κ and μ .



Calculated results: Position of shell closure

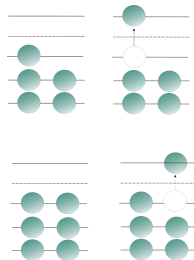
$k = 0.8$ (squares), 1.0 (circles), and 1.2 (triangles)



The stability of the nuclei with $Z > 120$ decreases with increasing k

The strength of spin-orbit interaction is crucial to define the position of the shell closures in nuclei beyond Pb. The 20% variation of the spin-orbit strength can strongly shift the position of the minimum of E_{sh} .

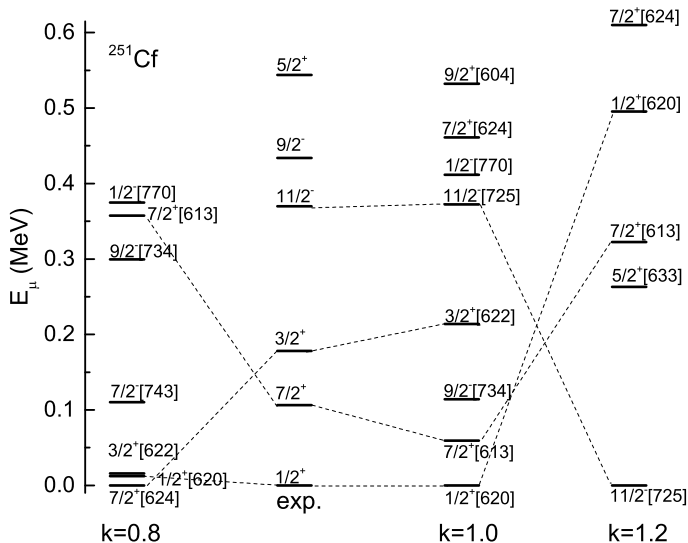
Calculating the potential energy surface as a function of collective coordinates with the TCSM, we find the ground-state potential minimum in which the energies of the low-lying one-quasiparticle states are obtained.



$$E_\mu = \sqrt{(e_\mu - e_F)^2 + \Delta^2} - \sqrt{(e'_\mu - e_F)^2 + \Delta^2}$$

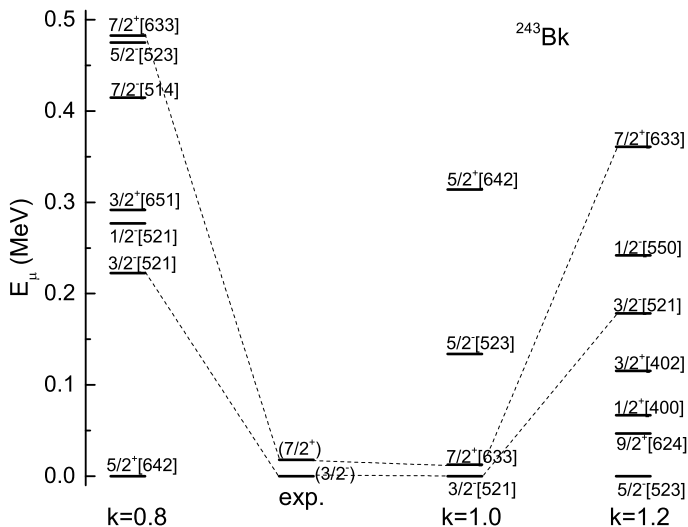
Single-particle states μ with energies e_μ ,
the pairing-energy gap parameter Δ ,
the Fermi energy e_F .

Calculated results: Dependence of one-quasiparticle spectra on spin-orbit strength



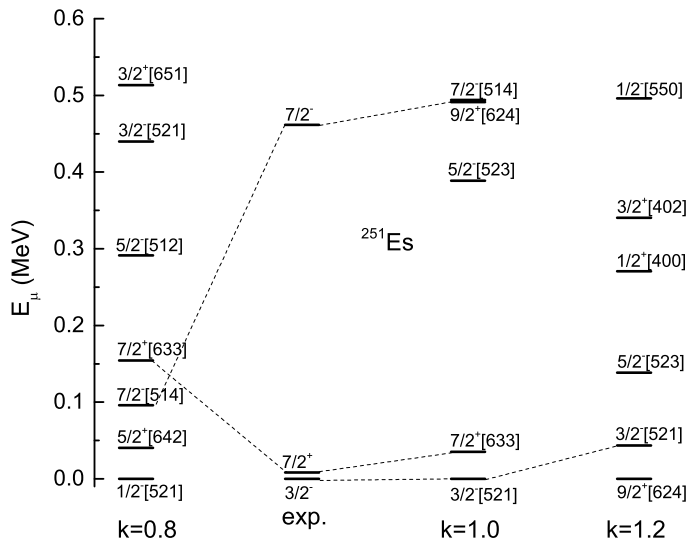
exp.: <https://www.nndc.bnl.gov>

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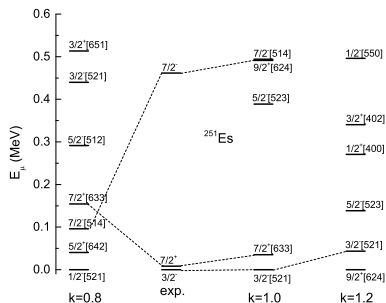
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$k = 0.8$ (squares), 1.0 (circles), and 1.2 (triangles)



Briefly

- The experimental energies, spins, and parities are well described (within 250 keV) with $k = 1.0$.
- The calculated results obtained at $k = 0.8$ and 1.2 are less consistent with the experimental data – the ground-state spins and parities can not be reproduced.
- In most cases, the one-quasiparticle spectra become denser with $k = 0.8$ or 1.2 .
- At $k = 1$ we have the best description of low-lying one-quasiparticle states.

- As shown, the quality of the description of low-lying one-quasiparticle states crucially depends on the spin-orbit strength. The spin-orbit strength taken in the modified TCSM at $k = 1$ allows us to describe well the low-lying one-quasiparticle spectra in heavy nuclei.

⁶A.N. Kuzmina, G.G. Adamian, N.V. Antonenko, and W. Scheid, Phys. Rev. C **85**, 014319 (2012)

⁷G.G. Adamian, I.A. Malov, N.V. Antonenko, H. Lenske, K. Wang, and S.-G. Zhou, Eur. Phys. J. A **54**, 170

SUMMARY

- As shown, the quality of the description of low-lying one-quasiparticle states crucially depends on the spin-orbit strength. The spin-orbit strength taken in the modified TCSM at $k = 1$ allows us to describe well the low-lying one-quasiparticle spectra in heavy nuclei.
- At $k = 0.8$ and 1.2 the calculated spectra are less consistent with the experimental data. So the choice of the TCSM parameters in ⁶ was optimal.

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- At $k = 1$ the strongest shell effects are found for the nuclei with $Z = 120$ or 124 and 126 at N approaching 184 . However, the variation of the value of E_{sh} in the isospin chains is relatively small, which confirms the results of self-consistent calculations ⁷

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- The shape of the island of stability is also defined by the spin-orbit strength. At $k > 1$, it is located between $Z = 112$ and 120 , while at $k \leq 1$ it is extended to $Z = 126$. The experiments on production of the $Z = 120$ nucleus could help us to answer the question of whether there is a shelf of stability beyond $Z = 120$.

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