

Instead of seeing the separation of scales as a drawback, we can embrace it and use it to our advantage \Rightarrow Effective Field Theory.

We give up the requirement that the theory is valid at all scales, as long as it produces accurate quantitative results at the relevant scales of the problem. Corrections need to be known. (see e.g. hep-ph/0410047)

General setup: Physics at a low energy scale described by effective d.o.f. treated explicitly while physics @ higher scales enters via
 ① contact interactions ② low-energy constants (Wilson coefficients in L)

Construct a Lagrangian in low energy d.o.f. that reproduces physical observables.

For heavy quarks: $\frac{\Lambda_{QCD}}{M} \ll 1$ no spontaneous $Q\bar{Q}$ pair production
 $\frac{T}{M} \ll 1$ no thermal $Q\bar{Q}$ pair production $\left. \vphantom{\frac{\Lambda_{QCD}}{M} \ll 1} \right\}$ Non-relativistic quarks.

How to construct the most general Lagrangian for non-rel. quarks compatible with the symmetries of QCD? (No general formula available)

Clever Trick: Foldy-Tani-Wouthuysen transformation + Expand the exponentials

$$\bar{Q}(x) (D_\mu \gamma^\mu + m) Q \quad Q \rightarrow Q' = e^{S'} e^S Q \quad \bar{Q} \rightarrow \bar{Q}' = \bar{Q} e^S e^{S'}$$

$$S = \frac{1}{2M} D_i \gamma_i = -S^\dagger \quad S' = \frac{1}{4M^2} \epsilon_{ijk} \gamma_k F_{ij}$$

$$Q = \begin{pmatrix} \psi \\ \chi \end{pmatrix}$$

From the appearing terms construct Lagrangian $L = \sum_n \frac{C_n(\mu_s(M), M)}{M^n} O_n(\mu, m_0, m_0^2)$

$$L \stackrel{NRQCD}{=} \underbrace{\psi^\dagger (iD_0 + H_4) \psi}_{L_\psi} + \underbrace{\chi^\dagger (-iD_0 + H_2) \chi}_{L_\chi = L_\psi^*} + \text{contact terms} - \frac{1}{4} F^{\mu\nu} F_{\mu\nu} + \bar{\psi}(\dots)\psi$$

light d.o.f @ scale μ

$$iD^0 = i\partial^0 - gA^0 \quad i\vec{D} = i\vec{\partial} + g\vec{A} \quad E = F^{i0} \quad B = -\epsilon_{ijk} F^{jk}/2$$

$$L_\psi = \psi^\dagger \left(iD_0 + \frac{c_1}{2M} D_i^2 + \frac{c_4}{8M^3} D^4 + \frac{c_F}{2M} \vec{\sigma} \cdot \vec{B} + \frac{c_D}{8M} (\vec{D} \cdot g\vec{E} - g\vec{E} \cdot \vec{D}) + \dots \right) \psi$$

* relativistic correction $\sqrt{p^2 + M^2} \approx M + \frac{p^2}{2M} - \frac{p^4}{8M^3}$ $C_k = C_4 = 1$ Constraints on c_i 's as requirement of underlying Lorentz invariance.

In general c_i 's can receive non-trivial contributions from UV physics @ the scale $\mu \sim M$, can even be complex: Need matching

Compute correlation functions both in QCD and NRQCD with the same physics content and set equal at same scale. That fixes the c_i 's. Carry out QCD computations either perturbatively or using lattice QCD. (In the following neglected)

$$j_n^A(x) = b_n(\psi^\dagger \sigma_n \psi) + O\left(\frac{1}{M^2}\right)$$

Note: Explicit mass term $2M$ from FTW removed from L_ψ . All energies are shifted and absolute energy scale needs to be calibrated.

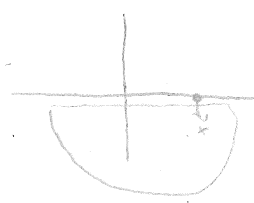
Let us have a look at the non-interacting theory ($A=0$)

$$\int D\psi D\psi^\dagger \psi^\dagger \psi e^{iS_{\text{free}}} = \int D\psi D\psi^\dagger \psi^\dagger \psi e^{i\psi^\dagger \left(i\partial_0 + \frac{p^2}{2M} \right) \psi}$$

$$= K \det(K^{-1}) \approx K \quad K^{-1} K_\psi = \mathbb{1} \Rightarrow (i\partial_0 + \frac{p^2}{2M}) K_\psi(x,x') = \delta(x-x')$$

Solve for the propagator explicitly: $(i\partial_0 + \frac{p^2}{2M}) K_\psi = \delta$ $(-i\partial_0 + \frac{p^2}{2M}) K_{\bar{\psi}} = \delta$

$$K(x_1, x_2) = \frac{1}{(2\pi)^4} \int d\omega \int d^3p e^{-i\omega t} e^{i\vec{p}\vec{x}} \frac{-1}{\omega - \frac{p^2}{2M} + i\epsilon}$$



$$= \frac{i\Theta(t)}{(2\pi)^3} \int d^3p e^{-i\frac{p^2}{2M}t} e^{i\vec{p}\vec{x}}$$

$$= \frac{i\Theta(t)}{(2\pi)^2} \int dp \int_{-1}^1 dq e^{-i\frac{p^2}{2M}t} e^{ipr} p^2$$

$$= -\frac{i\epsilon}{2M} \left(p^2 - pr \frac{2M}{\epsilon} + \frac{M^2 r^2}{\epsilon^2} - \frac{M^2 r^2}{\epsilon^2} \right)$$

$$= \frac{i\Theta(t)}{(2\pi)^2} \int dp \frac{p}{ir} (e^{ipr} - e^{-ipr}) e^{+i\frac{p^2}{2M}t}$$

$$= -\frac{i\epsilon}{2M} \left[\left(p - \frac{Mv}{\epsilon} \right)^2 - \frac{M^2 v^2}{\epsilon^2} \right]$$

$$= \frac{\Theta(t)}{(2\pi)^2} \int_{-\infty}^{\infty} dp \frac{p}{ir} e^{ipr} e^{+i\frac{p^2}{2M}t}$$

$$= \frac{\Theta(t)}{(2\pi)^2} e^{i\frac{M}{2t}r^2} \int_{-\infty}^{\infty} dp' \frac{p' + \frac{M}{\epsilon}r}{r} e^{-i\frac{t}{2M}p'^2}$$

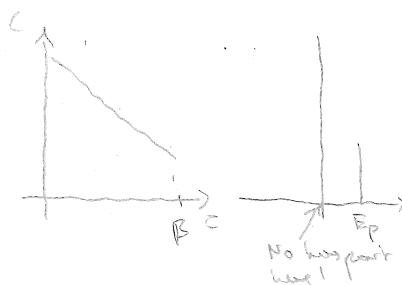
$$= \frac{\Theta(t)}{(2\pi)^2} \frac{\sqrt{\pi}}{2} \left(\frac{2M}{t} \right)^{3/2} e^{i\frac{M}{t}r^2}$$

$$= \frac{1}{2} \left(\frac{2M}{t} \right) \left(\frac{2M}{t} \right)^{1/2} \sqrt{\pi}$$

Quick check for $A \neq 0$ but $M \rightarrow \infty$ $(i\partial_0 - gA^0) K = \delta \Rightarrow K = \Theta(t) e^{i\int_{t'}^t A(s) ds} \delta(x)$

\Rightarrow There is only a single forward propagating mode.

Time-slice correlator $C(\epsilon) = K(x_1, \vec{p}) = e^{-\frac{p^2}{2M}\epsilon} \Theta(\epsilon)$
(single heavy quark) $\delta(\omega) = \Theta(\omega) \delta(\omega - \frac{p^2}{2M})$



Now put NRQCD on the lattice: $t \rightarrow i\tau$ $D_t \rightarrow iD_\tau$

$$a \Delta^+ \psi(x) = U_{x,x+a\hat{p}} \psi(x+a\hat{p}) - \psi(x) \quad \Delta^\pm = \frac{1}{2} (\Delta^+ + \Delta^-) \quad \Delta^2 = \sum_i \Delta_i^+ \Delta_i^-$$

(see e.g. PRD 43 (1991) 1967; hep-lat/9205007)

discretized Laplacian becomes series in powers of $\frac{1}{Ma}$

$$S_{\text{lat}, \psi}^{\text{NRQCD}} = \psi^\dagger K^{-1} \psi = \sum_{x,z} \psi^\dagger \left(\Delta_x^+ - \sum_{i=1}^3 \frac{\Delta_i^+ \Delta_i^-}{2M} + \frac{\vec{\nabla}^2 \vec{\nabla}^2}{2M} + \dots \right) \psi_{x,z}$$

$B = \begin{pmatrix} \square & \square \\ \square & \square \end{pmatrix}$ clover leaf approximation

The heavy propagator from $\int d^4p \psi^\dagger(p) \left[i\tilde{p}_4 a - \frac{p_4^2 a^2}{2} + \frac{1}{2M} \sum_i (\hat{p}_i^2) \right] \psi(p)$

$$\text{since } e^{i\tilde{p}_4 a} - 1 = i\tilde{p}_4 a - \frac{a^2 \tilde{p}_4^2}{2}$$

This expression only vanishes once in the Brillouin zone \rightarrow no doublers.

How to compute heavy quark propagation on the lattice?

$$K^{-1}K = \mathbb{1} \quad \left(\Delta_4^+ - \sum_{j=1}^3 \frac{\Delta_j^+ \Delta_j^-}{2M} \right) K_{x,\tau} = \delta_{x,\tau}$$

This is a discretized stochastic diffusion equation (U_n fluctuate)

① Initial conditions with: $K_{x,0} = \eta$ with $\langle \eta \rangle = 0$ $\langle \eta \eta' \rangle = \delta_{x,x'}$
use different noise for each configuration.

② $K_{x,\tau+n} = \underbrace{U_{x,\tau}^+ (1 - aH)}_{\neq e^{-aH\tau}}$ $K_{x,\tau}$ for better stability usually define lepton parameter η .

$$= \left(1 - \frac{a}{2n} H\right)^n U_{x,\tau}^+ \left(1 - \frac{a}{2n} H\right)^n K_{x,\tau}$$

⇒ Much simpler than relativistic theory, no inversion of Dirac operators.

Now we are ready to compute heavy quarkonium current-current correlators:

Channel	Relativistic	NRQCD
$3S_1$	$\bar{Q} \gamma^\mu Q$	$z^+ \sigma_i z$
$3P_1$	$\bar{Q} \gamma^i \gamma^j Q$	$z^+ (\delta_i^j \sigma_j - \delta_j^i \sigma_i) z$

$$D(\tau, \vec{x}) = \int D(\psi, \psi^\dagger) \int D(z, z^\dagger) \int D\psi \quad (z^+ \sigma_i z) (z^+ \sigma_i z)^\dagger e^{-S[\psi, \psi^\dagger, z, z^\dagger, \psi]} e^{-S[\psi, \psi^\dagger, z, z^\dagger, \psi]}$$

In large ω limit the fermion gives $D(\tau, \vec{x}) = \int D\psi \bar{\psi} [K_2^+(\tau, \vec{x}) \sigma_i K_2(\tau, \vec{x}) \sigma_i] \psi e^{-S[\psi]}$

Quarkonium at rest $D(\tau) = \int d^3x D(\tau, \vec{x})$

Strategy: compute the evolution of single heavy quark then combine into quarkonium correlator. (in Euclidean time $K_1 = K_2$)

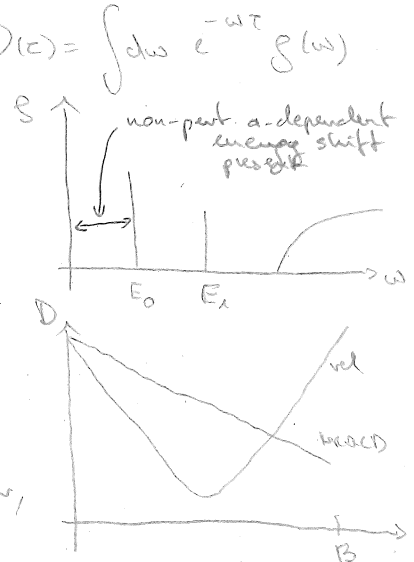
At $T=0$ $D = \sum_n \langle 0 | J(x,0) J^\dagger(x,0) | 0 \rangle \xrightarrow[\text{evol}]{\text{NRQCD time}} \sum_n \sum_x \langle 0 | J(x,0) e^{-H\tau} | n \rangle \langle n | J^\dagger(x,0) | 0 \rangle$ state with @ @

$$= \sum_n \sum_x \underbrace{|K_0 J(x,0) | n \rangle|^2}_{A_n} e^{-E_n \tau}$$

In general even at $T > 0$ $D(\tau) = \int d\omega e^{-\omega \tau} g(\omega)$

- Because E_0 still contains an energy shift, since we subtracted $2M$ in L_{NRQCD} . Need to calibrate for each lattice spacing by fixing e.g. $3S_1$ mass to PDG value.

- Positive: - Absence of backward travelling mode \rightarrow no symmetry in $\tau \rightarrow$ all data points usable.
- Due to energy shift, exponential decay weaker, better S/N ratio.



What is the meaning of A_n ?

$$\psi_n(x) = \langle 0 | Z^\dagger(x/2) \psi(-x/2) | n \rangle \quad \text{Bethe-Salpeter wavefunction}$$

Related to the decay of heavy quarkonium, e.g. into dileptons.

$$\Gamma(Q\bar{Q} \rightarrow \ell\bar{\ell}) = \sum_n \frac{2 \operatorname{Im}(f_n(\Lambda))}{M^{d_n-4}} \langle Q\bar{Q} | O_n(\Lambda) | Q\bar{Q} \rangle \quad (\text{see e.g. hep-ph/9407339})$$

which in NRQCD is mediated via contact interactions, since it describes physics at the scale M which was integrated out (not explicit in the EFT)

$$\Gamma(\Upsilon \rightarrow e^+e^-) = \frac{2 \operatorname{Im} f_{\Upsilon}(\Lambda)}{M^2} |\langle 0 | \chi^\dagger \psi | \Upsilon \rangle|^2 + \frac{2 \operatorname{Im} g_{\Upsilon}(\Lambda)}{M^4} \operatorname{Re} \left[\langle \Upsilon | \bar{\chi} \psi | 0 \rangle \langle 0 | \chi^\dagger \psi | \Upsilon \rangle \right]$$

Other channels can also contain color octet contributions.

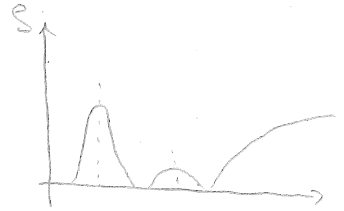
Decay rates can be systematically related to non-relativistic wavefunctions at the origin $A_n(1S_0) = |\psi_n(0)|^2$ via Bethe-Salpeter wavefunction.

(current research interest: use BSW to derive unbar force from lattice QCD)

What is the situation at finite temperature?

Spectrum gives information about dilepton emission of heavy

$$Q\bar{Q} \text{ in a medium: } R_{\Upsilon e^+e^-} \propto \int d^3p_0 \int d^3p \frac{\rho(p)}{p^2} v_B(p_0)$$



reduces to the $T=0$ formula if one assumes $\rho(p) = \rho(\omega = \sqrt{p^2 + M^2})$ and $g(\omega) = \delta(\omega - E_0)$

Lattice NRQCD at finite temperature:

The derivation of the lattice action proceeds equally if $T \neq 0$, i.e. for a finite Euclidean time extent. As long as $T \ll M$ the Wilson coefficients only receive small radiative corrections, which is neglected in current $T=0$ NRQCD studies. Solve diffusion equation along finite τ -axis.

Gain some intuition via the free theory:

$$\text{remember: } K_{\tau}^{\pm} = \frac{\Theta(\tau)}{(2\pi)^2} \frac{\Gamma}{2} \left(\frac{2M}{\tau}\right)^{3/2} e^{-\frac{M}{\tau} \tau^2} \quad K_{\tau}^{\pm} = \frac{\Theta(\tau)}{(2\pi)^2} \frac{\Gamma}{2} \left(\frac{2M}{\tau}\right)^{3/2} e^{-\frac{M}{\tau} \tau}$$

$$D(\tau) = \int d^3x K_{\tau}^+ K_{\tau}^- \propto \left(\frac{2M}{\tau}\right)^3 \int d^3x e^{-\frac{2M}{\tau} \tau^2} \propto \left(\frac{2M}{\tau}\right)^3 \left(\frac{2M}{\tau}\right)^{-3/2} \propto \left(\frac{2M}{\tau}\right)^{3/2}$$

$$\underline{D(\tau) \propto \int d^3p e^{-2E_p \tau}} \quad \text{which with } \underline{g(\omega) = \int d\omega e^{-\omega \tau} g(\omega)}$$

$$\text{leads to } g(\omega) = \int d^3p \delta(\omega - 2E_p) = \frac{N_c}{\pi} M^{3/2} \omega^{1/2} \Theta(\omega)$$

Comparison to relativistic free spectrum

(see hep-lat/0509004)

$$\rho(\omega) = \Theta(\omega^2 - 4M^2) \frac{N_c}{8\pi\omega} \sqrt{\omega^2 - 4M^2} \left(1 - 2u_F(\omega/2)\right) (\omega^2 1_c + 4M^2 1_c) + 2\pi\omega \delta(\omega) N_c \gamma$$

$\omega \rightarrow \omega$
 $\omega \rightarrow \omega$

Now the lattice NROCD version: We get the dispersion relation from our time evolution operator $K_{x,t+1} = \left(1 - a \frac{p^2}{2M}\right) K_{x,t} \Rightarrow a E_p = -\log\left(1 - \frac{p^2}{2M}\right)$

C++ Exercise: Compute the NROCD free spectral function for $M_0 = 5$

Let us have a look at some actual real-world NROCD data:

Hot QCD $N_f = 2+1$ HISQ ensembles $48^3 \times 17$ ($T > 0$) $48^3 \times 32, 48, 64$ ($T \approx 0$)

$M_0 a \in [2.8 \dots 0.9]$ $n = 4$ Fig 8 Fig 9 (see also: 1409.3630; 1402.6210; 1109.4496)

\Rightarrow Notice mass difference between different lattice spacings \rightarrow NROCD energy shift

\Rightarrow First hint at in-medium modification from correlator ratios: Fig 10 Fig 11

Correlators only give us information about global in-medium changes but we are interested in changes of individual states. Need spectral information (mass & width @ $T > 0$). One needs acquire thermal width effective mass analysis (is fails (no plateau!))

Spectral function reconstruction using Bayesian inference

In general we have the relation $D(\tau) = \int d\omega K(\tau, \omega) \rho(\omega)$

which needs to be inverted. Since ρ can contain many features needs to be well resolved along ω .

$$D_i = \sum_{\ell=1}^{N_\omega} \delta\omega_\ell K_{i\ell} \rho_\ell$$

Relativistic $\rho(\omega)$
 $\rho(\omega) = \frac{\cosh(\omega(\tau - \beta/2))}{\sinh(\omega\beta/2)}$

NROCD $\rho(\omega)$
 $\rho(\omega) = e^{-\omega\tau}$

Inversion is ill-posed and ill-conditioned:
 $N_\tau \ll N_\omega$
 ∞ solutions to 2^2 bit.
 $D_i = D_i^{ideal} + \eta$

Naive inversion in the case $N_\tau = N_\omega$ results to exponential increase of noise.

Question: How to systematically regularize the problem?

Personal answer: Bayesian inference that allows to incorporate additional prior information on the spectrum. Subjective view of probability, assign $P[X]$ even if X is not random variable.

Multiplication law for probs. \downarrow

$$P(S, D, I) = P(S|D, I) \cdot P(D|I) \\ = P(D|S, I) P(S|I)$$

Posterior
 $P[S|D, I] =$
 \uparrow
test function

$$\frac{\text{Likelihood} \cdot \text{Prior}}{\text{Evidence}} \\ P[D|S, I] P[S|I] \\ \hline P[D|I]$$

see also: Phys. Rev. 269 (1996) 1733