

# Status and plans of the BM@N experiment



#### M.Kapishin



## **Baryonic Matter at Nuclotron (BM@N) Collaboration:**



# 10 Countries, 20 Institutions, 240 participants

- University of Plovdiv, Bulgaria → MoU signed;
- St.Petersburg University → MoU signed;
- Shanghai Institute of Nuclear and Applied
  Physics, CFS, China;
- Nuclear Physics Institute CAS, Czech Republic→ MoU signed;
- CEA, Saclay, France;
- TU Darmstadt, Germany;
- GSI & FAIR, Germany;
- Tubingen University, Germany → MoU signed;
- Tel Aviv University, Israel;
- Joint Institute for Nuclear Research;
- Warsaw University of Technology, Poland→ MoU signed;
- University of Wroclaw, Poland → MoU signed;
- Institute of Nuclear Research RAS, Moscow, Russia → MoU signed; BM@N Experiment

- NRC Kurchatov Institute, Moscow;
- Institute of Theoretical & Experimental
- Physics, NRC KI, Moscow  $\rightarrow$  MoU signed;
  - Moscow Engineer and Physics Institute, Russia  $\rightarrow$  MoU signed;
- Skobeltsin Institute of Nuclear Physics, MSU, Russia → MoU signed;
- Moscow Institute of Physics and Technics, Moscow, Russia → MoU signed;
- Massachusetts Institute of Technology, Cambridge, USA.

# **Heavy Ion Collision Experiments**



Future CBM experiment: Au+Au at  $\sqrt{s_{NN}}$ ~ 2.7 – 4.9 GeV

# EOS of symmetric and asymmetric nuclear matter

**BM@N** experiment

Ch. Fuchs and H.H. Wolter, EPJA 30 (2006) 5



EOS: relation between density, pressure, temperature, energy and isospin asymmetry

$$\mathsf{E}_{\mathsf{A}}(\rho,\delta) = \mathsf{E}_{\mathsf{A}}(\rho,0) + \mathsf{E}_{\mathsf{sym}}(\rho) \cdot \delta^2$$

with  $\delta = (\rho_n - \rho_p)/\rho$  E/A( $\rho_o$ ) = -16 MeV

Curvature defined by nuclear incompressibility:  $K = 9\rho^2 \ \delta^2(E/A)/\delta\rho^2$ 

Study symmetric matter EOS at  $\rho$ =3-5  $\rho_0$   $\rightarrow$  elliptic flow of protons, mesons and hyperons

 $\rightarrow$  sub-threshold production of strange mesons and hyperons

 $\rightarrow$  extract K from data to model predictions

► Constrain symmetry energy E<sub>sym</sub>

 $\rightarrow$  elliptic flow of neutrons vs protons

 $\rightarrow$  sub-threshold production of particles with opposite isospin

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NICA main competitor  $\rightarrow$  STAR experiment: BES Fixed Target program Collected 2.10<sup>9</sup> interactions of Au+Au at  $\sqrt{s}$  = 3 GeV in 2021

# Plan for BM@N Experimental physics run for 800 hours (33 days) in spring 2022

BM@N: Estimated hyperon yields in Xe + Cs collisions

4 A GeV Xe+Cs collisions, multiplicities from PHSD model, Beam intensity 2.5·10<sup>5</sup>/s, DAQ rate 2.5·10<sup>3</sup>/s, accelerator duty factor 0.25 1.8·10<sup>9</sup> interactions

1.8.10<sup>11</sup> beam ions

E <sub>thr</sub> NN	М	3	Yield/s	Yield / 800		
GeV	b<10 fm	%	b<10fm	hours		
1.6	1.5	3	220		0.8·10 <sup>8</sup>	
3.7	2.3·10 <sup>-2</sup>	1	1.1		4·10 <sup>5</sup>	
6.9	2.6·10 <sup>-5</sup>	1	1.3·10 <sup>-3</sup>		470	
7.1	1.5·10 <sup>-5</sup>	3	2.2·10 <sup>-3</sup>		800	
	E <sub>thr</sub> NN GeV 1.6 3.7 6.9 7.1	$E_{thr}NN$ MGeVb<10 fm	$E_{thr}NN$ MεGeVb<10 fm	$E_{thr}NN$ M $\epsilon$ Yield/sGeVb<10 fm	$E_{thr}NN$ MεYield/sYield/sGeVb<10 fm	$\begin{array}{c c c c c c c c c c c c c c c c c c c $

# Plan for BM@N experimental physics run with Au (Bi) beam for 800 hours (33 days) in spring 2023

### BM@N: Estimated hyperon yields in Au+Au collisions

4 A GeV min. bias Au+Au collisions, multiplicities from statistical model, Beam intensity  $2.5 \cdot 10^{5}$ /s , DAQ rate  $2.5 \cdot 10^{3}$ /s, accelerator duty factor 0.25

Experimental run for 800 hours (33 days)

1.8 $\cdot$ 10<sup>9</sup> interactions 1.8 $\cdot$ 10<sup>11</sup> beam ions

Particle	E <sub>thr</sub> NN	M	M	3	Yield/s	Yield / 800	
	GeV	central	m.bias	%	m. Bias	hours	
						m. Bias	
[1]	3.7	1.10 <sup>-1</sup>	2.5·10 <sup>-2</sup>	1	2.5	4.5·10⁵	
Ω	6.9	2·10 <sup>-3</sup>	5·10 <sup>-4</sup>	1	5·10 <sup>-2</sup>	0.9·10 <sup>4</sup>	
Anti- $\Lambda$	7.1	2.10-4	5·10 <sup>-5</sup>	3	1.5·10 <sup>-2</sup>	2700	
<b>Ξ</b> +	9.0	6·10 <sup>-5</sup>	1.5·10 <sup>-5</sup>	1	1.5·10 <sup>-3</sup>	270	
$\Omega^+$	12.7	1.10 <sup>-5</sup>	2.5·10 <sup>-6</sup>	1	2.5·10 <sup>-4</sup>	45	
To perform main BM@N physics program need					^3H	0.9·10 <sup>5</sup>	

10 times more statistics  $\rightarrow 2.10^{10}$  interactions

#### Comparison HADES, STAR FxT, BM@N

	year	A+A	E <sub>kin</sub> A GeV	# Events	Rare Observables		vables
					e+e-	Ξ <sup>-</sup> , Ω <sup>-</sup>	hypernuclei
HADES	2012	Au+Au	1.23	7·10 <sup>9</sup>	$\checkmark$		
HADES	2019	Ag+Ag	1.58	1.4·10 <sup>10</sup>	$\checkmark$		800 <sup>3</sup> <sub>A</sub> H
STAR FxT	2018	Au+Au	2.9	3·10 <sup>8</sup>		10 <sup>4</sup> Ξ <sup>-</sup>	10 <sup>4</sup> <sup>3</sup> <sub>Λ</sub> H, 6·10 <sup>3</sup> <sup>4</sup> <sub>Λ</sub> H,
STAR FxT	2021 planned	Au+Au	2.9	2·10 <sup>9</sup>		7·10⁴Ξ⁻, Ω⁻?	7·10 <sup>4</sup> <sup>3</sup> <sub>A</sub> H, 4·10 <sup>4</sup> <sup>4</sup> <sub>A</sub> H, <sup>5</sup> <sub>A</sub> He, <sup>7</sup> <sub>A</sub> Li, <sup>7</sup> <sub>A</sub> He, ?
BM@N	simulated	Au+Au	3.8	2·10 <sup>10</sup>		$5 \cdot 10^{6} \equiv^{-1}$ Expected: $10^{5} \Omega^{-1}$ $3 \cdot 10^{4}$ anti-Λ $5 \cdot 10^{2} \Omega^{+1}$	10 <sup>6</sup> ${}^{3}_{\Lambda}$ H, ${}^{4}_{\Lambda}$ H, ${}^{5}_{\Lambda}$ He, ${}^{7}_{\Lambda}$ Li, ${}^{7}_{\Lambda}$ He, Expected: 10 <sup>2</sup> ${}^{5}_{\Lambda\Lambda}$ H

Reaction rates: HADES  $\approx$  20 kHz, BM@N  $\approx$  20 kHz, STAR FxT  $\approx$  2 kHz

Energy Au beams: HADES: 0.2 - 1.25 A GeV, BM@N: 1.5 – 3.8 A GeV, STAR FxT: > 2.9 A GeV Conclusion:

HADES and BM@N are complementary , no cascade hyperons ( $\Xi^-, \Omega^-$ ) at HADES Statistics at BM@N  $\approx$ 70 times higher ( $\Xi^-$ ) than at STAR FxT





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For heavy ion beam intensities of few  $10^6$  Hz  $\rightarrow$  keep 4 STS + 7 GEM

 $\rightarrow$  fast FEE and readout electronics



# BM@N main detector activities towards heavy ion run BM@N

#### **Central tracking system**

#### **GEM detectors**



#### Forward Silicon Tracker



Carbon fiber vacuum beam pipe



Forward Hadron Calorimeter and Hodoscope

#### Outer tracker: Cathode Strip Chambers



#### Silicon Tracking System





# Beam parameters and setup at different BM@N stages of the BM@N experiment

Year	2016	2017 spring	2018 spring	2022 spring	2023	After 2023
Beam	d(↑)	С	Ar,Kr, C(SRC)	Xe	Au (Bi)	Au (Bi)
Max.inten sity, Hz	0.5M	0.5M	0.5M	0.5M	0.5M	2M
Trigger rate, Hz	5k	5k	10k	10k	10k	up to 50k
Central tracker status	6 GEM half planes	6 GEM half planes	6 GEM half planes + 3 forward Si planes	7 GEM full planes + 3 forward Si planes	7 GEM full planes + 4 forward Si + 2 large STS planes	7 GEM full planes + 4 large STS planes
Experimen tal status	technical run	technical run	technical run+physics	stage 1 physics	stage1 physics	High rate stage 2 physics

SRC physics run with C12 beam (4 weeks of data taking)

• Only Nuclotron with laser source is sufficient

# Limitations / requirements for BM@N physics run with Xe beam in spring 2022 (800 hours of physics data taking to collect 2.10<sup>9</sup> Xe + Csl interactions)

- Need Booster Nuclotron accelerator system
- Need 2 months for transition from SRC set-up to heavy ion setup + 0.5 month for magnetic field map measurement
- Full vacuum transport channel from Nuclotron to BM@N
- Xe beam of maximal possible energy (up to 3.9 AGeV)
- Need few days for technical run before physics run to prove beam quality and detector response, in case of problems → postpose physics run
- If SRC run extends to January 2022:
  → only chance to shift BM@N physics run to April May 2022

Requirements for BM@N physics run with Bi beam in spring 2023 (800 hours of physics data taking to collect 2.10<sup>9</sup> Bi + Bi interactions)

- Full vacuum transport channel from Nuclotron to BM@N
- Bi beam energy of maximal possible energy (up to 3.8 AGeV)

► To perform main BM@N physics program need 10 times more statistics  $\rightarrow$  2·10<sup>10</sup> Bi+Bi interactions with beam energies from 1.5 AGeV up to 3.8 AGeV

Need also C + C and Xe + Csl interactions at these energies for reference

# Thank you for attention!

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## Beam tracing through BMN beam pipe and profile monitoring



First task of the next run  $\rightarrow$  trace beam and monitor its profile in the end of the setup (try to find optimal trajectory to reduce background)





# Nuclotron and BM@N beam line





26 elements of magnetic optics:

- $\rightarrow$  8 dipole magnets
- $\rightarrow$  18 quadruple lenses

Requirements for Au beam:

Minimum dead material

 $\rightarrow$  need to replace air intervals / foils with



## Feasibility studies for first physics run: $\Xi^{-}$ and $_{\Lambda}H^{3}$ reconstruction in Xe +A interactions: 3 Forward Si + GEM

200 Entries / 1 MeV/c<sup>2</sup>  $\Xi^{-} \rightarrow \Lambda + \pi^{-}$ **5M** interactions **AGeV** 150 Mass = 1.3213Sigma = 0.0010100 S/B = 3.0 $S/\sqrt{S+B} = 15.8$ Eff. = 0.6%50 1.28 1.32 1.3 1.34 1.36 1.38  $M_{inv}$ , GeV/c<sup>2</sup> Entries / 1 MeV/c<sup>2</sup>  $_{\Lambda}H^{3} \rightarrow He^{3} + \pi^{-}$ Mass = 2.9916300 for 140M Sigma = 0.0010Generated E GeV/c S/B = 1.1interactions 3.5  $S/\sqrt{S+B} = 16.4$ 200 ۰, Eff. = 0.3%2.5 100 1.5 2.942.96 2.98 3.02 3.04 0.5 3  $M_{inv}$ , GeV/c<sup>2</sup> 0<u>-</u>1 -0.50 0.5 1.5 2 1

DCM-SMM model: Xe + Sn , T<sub>0</sub>= 3.9 AGeV



BM@N



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