CALIBRATION OF Ge DETECTORS WITH AUTOMATED ALGORITHM

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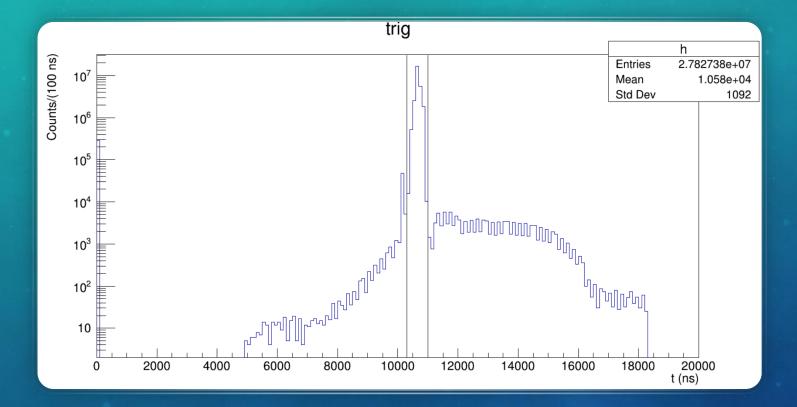
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OUTLINE

- Discuss the updated fit function
 - This work was done during winter holidays and results were included in the progress report, but it was not discussed in the collaboration meeting
 - The results from previous fit function (Gaus + lin) were presented without applying any cuts
 - The new fit function was used after applying the cuts
- Discuss the algorithm for automatically fitting peaks from detectors 2-8
 - Focus is to reduce entering raw numbers by hand
 - Look at the energy resolution of all the detectors

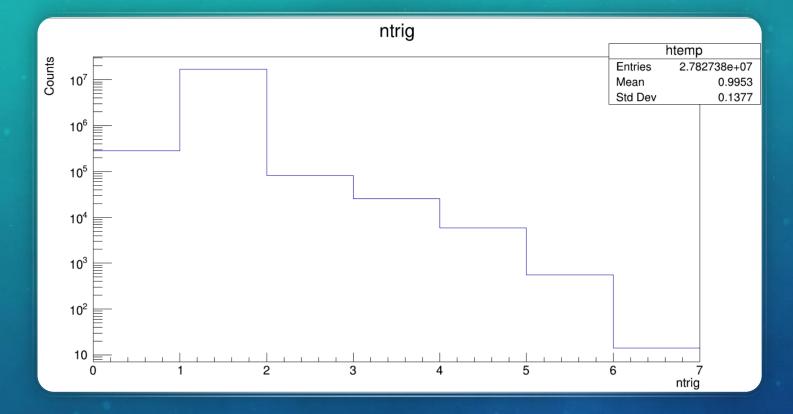
UPDATING THE FIT FUNCTION

- We changed the fit function to include incomplete charge collection and Compton scattering
 - Total function = Gaus + Erfc + Lin + Gaus*Erfc
 - Gaus: to model the peak
 - Erfc+Lin: to model flat background with linear and Compton scattering with erfc
 - Gaus*Erfc: to model incomplete charge collection in Ge detector
- Before using the new fit function, we applied some quality cuts to our data



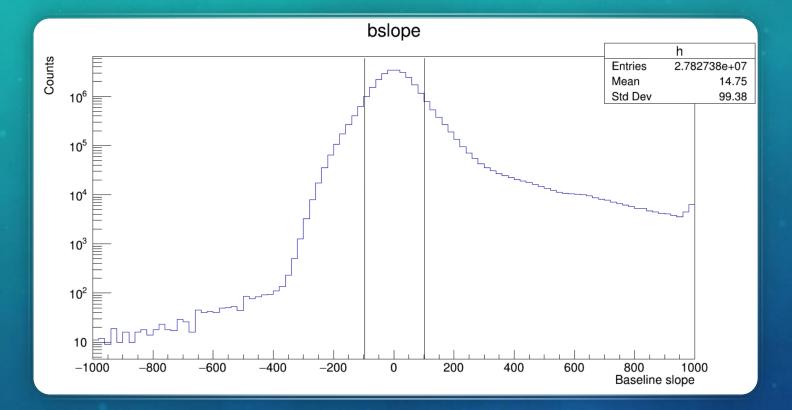
CUT ON TRIGGER TIME

- In LLAMA, the position of reconstructed peak within the waveform is in "GEMDTrigger_trigger"
- If a pulse is reconstructed correctly, the trigger should be approximately in the middle of the waveform
- Trigger timing cut for detector #1 was applied between 10300 ns and 11000 ns
- The boundaries are chosen where first minima (approximately) on both sides of the peak is obtained



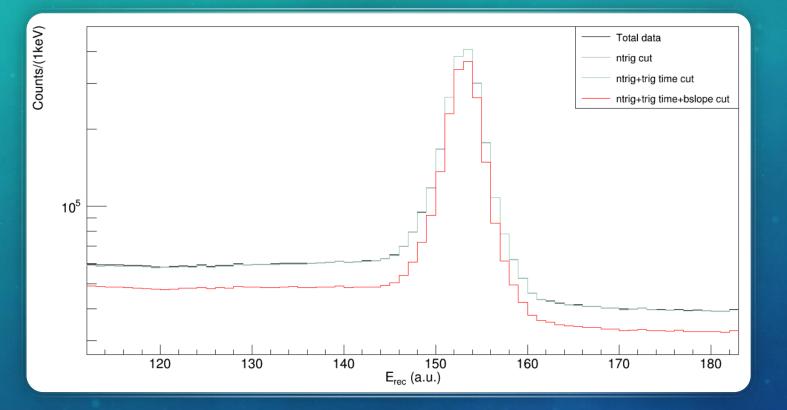
CUT ON TRIGGER NUMBER

- When the digitizer gives V-t distribution within a time window, ideally, there should be only 1 voltage pulse reconstructed within the waveform
- Multiple pulses can be sometimes reconstructed within the waveform, as seen in plot
- LLAMA saves this information in "GEMDTrigger_triggerNumber"
- We applied a cut to confine the number of triggers to 1
- We noticed some discrepancy between total number of events and the integral of this histogram, still investigating the issue



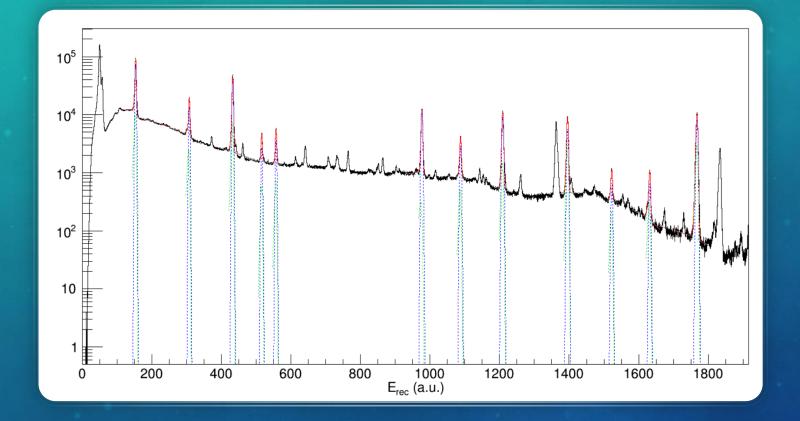
CUT ON BASELINE SLOPE

- Baseline slope cut within ±100 for flat/regular events (det #1)
- Region chosen close to ±2 sigma (sigma ~60) from the mean of baseline slope
- Need to discuss this in detail with TUM group for better understanding and selecting the range



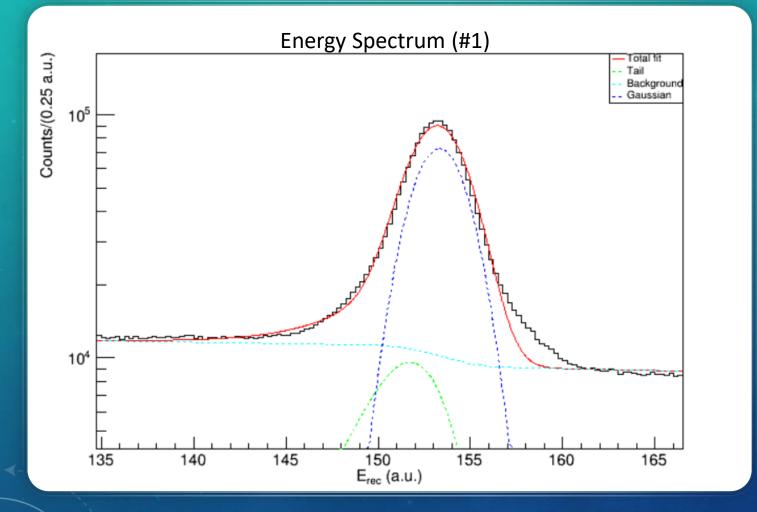
EFFECT OF INDIVIDUAL CUTS ON DATA

- Change in energy spectrum with each consecutive cut is shown
- Trigger number and trigger timing cuts do not have huge impact on the counts
- Baseline slope reduces the counts by some fraction
- Efficiency of all cuts = 82% accepted



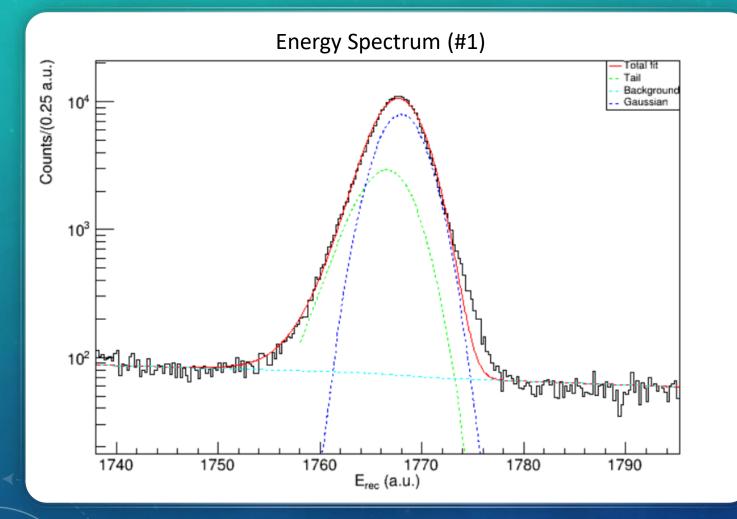
DETECTOR #1- FIT OF ALL THE PEAKS

- All 12 peaks were fitted to the same function for detector #1
- Status "CONVERGED" and Error Matrix "ACCURATE" in all the peaks
- The range of fitting (left/right width) is a "hyperparameter" and needs to be tuned by hand to give a good fit.



FITTED PEAK (121 keV)

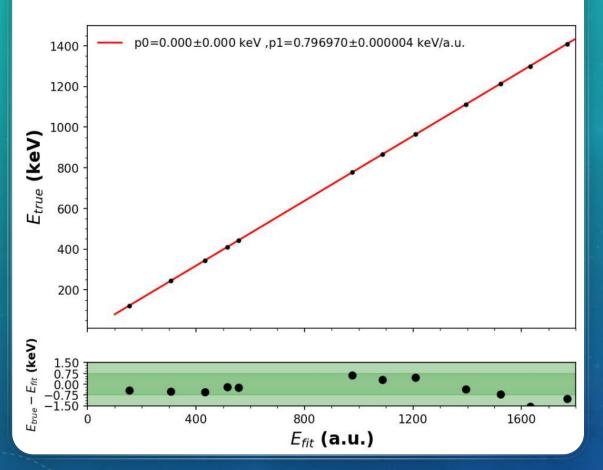
- Shows fit of highest intensity peak
- Individual functions such as Gaussian peak (blue), background (cyan) and tail portion (green) are shown along with complete fit (red)
- The more complicated fit functions still does not describe the peaks very well, as there are obvious right tails
- These right tails are seen with trapezoidal filter in Co-60 spectrum, so we should investigate it further



Peak at 1408 keV

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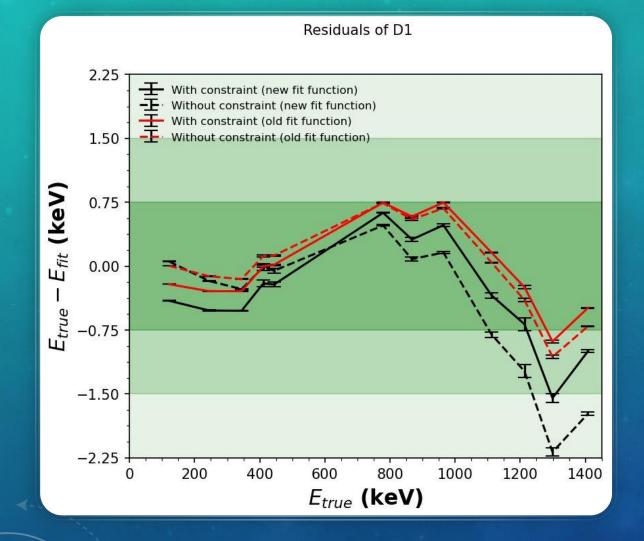
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Energy calibration (#1)

ENERGY CALIBRATION CURVE

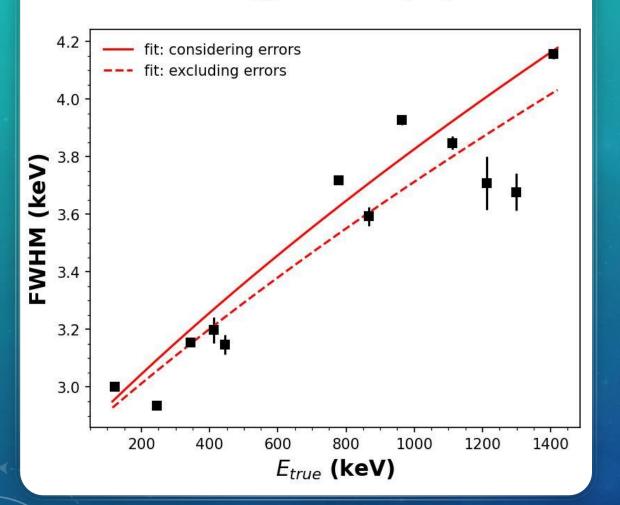
- Calibration constant has changed by a very small value (of the order of ~1E-4) compared to previous fit function
- In the previous fit, no limit was put on intercept value, but the parameter limit in this new fit was set between 0 and 1E5
- Intercept is very small (~1E-11) but positive (negative in earlier case, which makes no sense)



RESIDUALS

- With constraint on intercept, the residuals of the new peak fitting function are within -1.5 keV to +0.6 keV (black solid line)
- With same constraint, the residuals from previous peak fitting function are within -0.9 keV to +0.75 keV (red solid line)
- The residuals become worse with the new fit function
- Error bars show statistical errors only

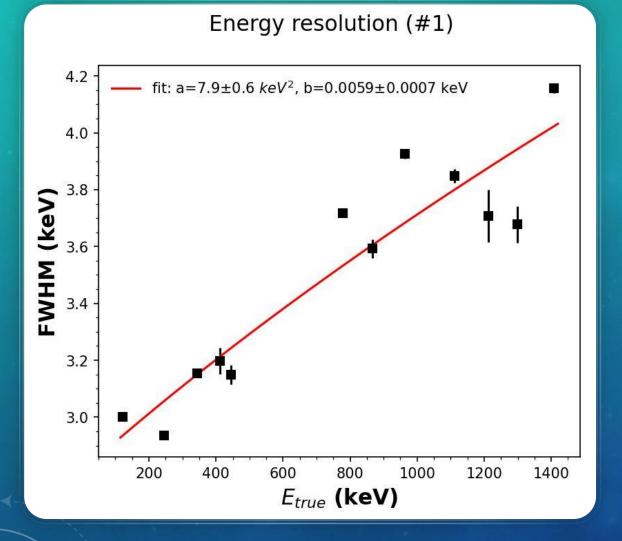
Energy resolution (#1)



ENERGY RESOLUTION

•When including errors, the fit ignores points at higher energies (solid line)

This means that these points have much larger errors, but here we include only statistical error
We should consider systematic errors in future
For now, we fit the data without taking errors into account (dashed line)



ENERGY RESOLUTION

•The energy resolution obtained from the fit is 3.9 keV at 1.3 MeV

•Comparable to the value obtained from MIDAS (~ 5.7 keV Co-60 at 1.1 MeV)

•Fit function used is $FWHM = 2.355 \cdot \sqrt{a + bE}$

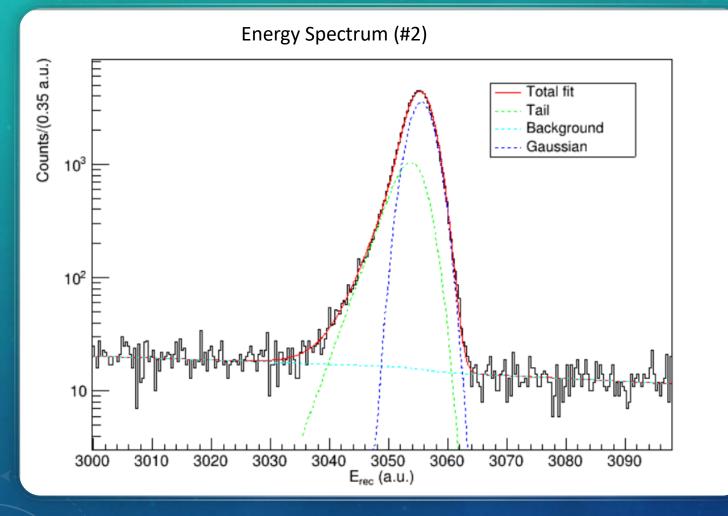
•Need to try a different fit function for resolution to include more physics

CALIBRATION OF OTHER DETECTORS

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AUTOMATED ALGORITHM

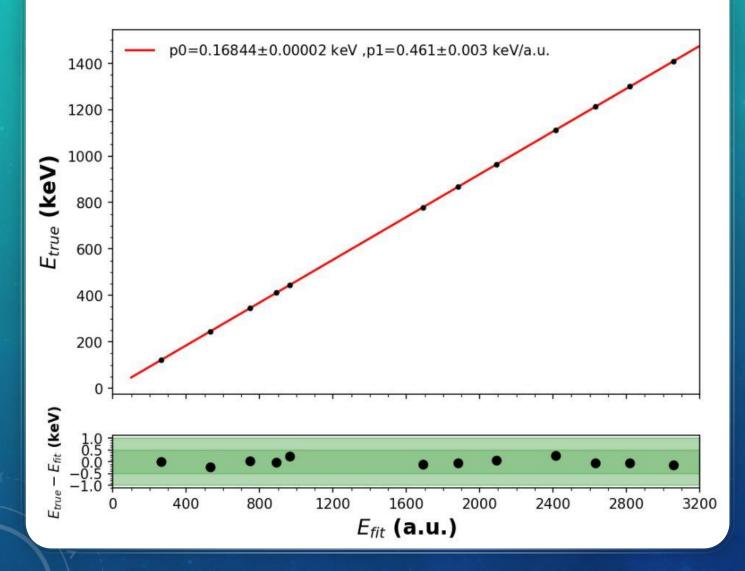
- Determine 2 peaks corresponding to true energies 121 keV and 244 keV
- Determine an approximate width of the peaks, width0 (usually 1.5-2 keV)
- Calculate the preliminary slope of the spectrum
 - Prelim. Slope = difference in true energy of first two peaks /difference between location of first two peaks in a.u.
- Calculate the approximate location of other 10 peaks, we call it E_{trv}
 - $E_{try} = E_{true} / prelim. slope$
- To find the exact location of peaks, find the left/right bins using E_{try}-10*width0 and E_{try}+10*width0, respectively
- Set the axis range within this region and get the maximum bin within this range
- Obtain E_{refined} using maximum bin and bin width, E_{refined} = maximum bin* bin width
- This E_{refined} becomes the energy parameter we set for our fit function



Peak at 1408 keV

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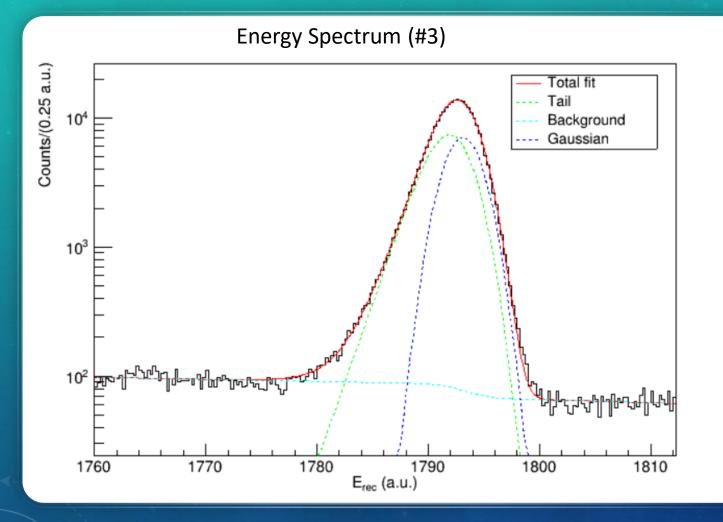
Energy calibration (#2)



DETECTOR #2

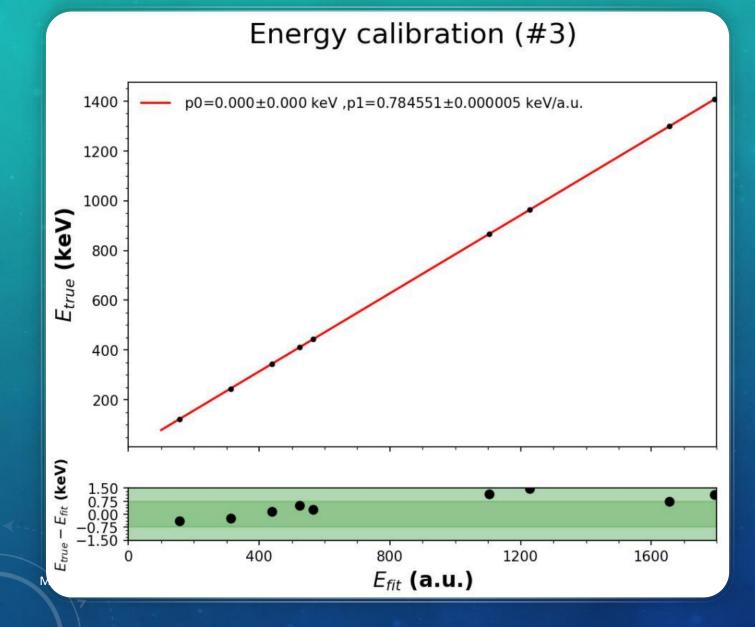
 Status "CONVERGED" and Error Matrix "ACCURATE" in all the peaks



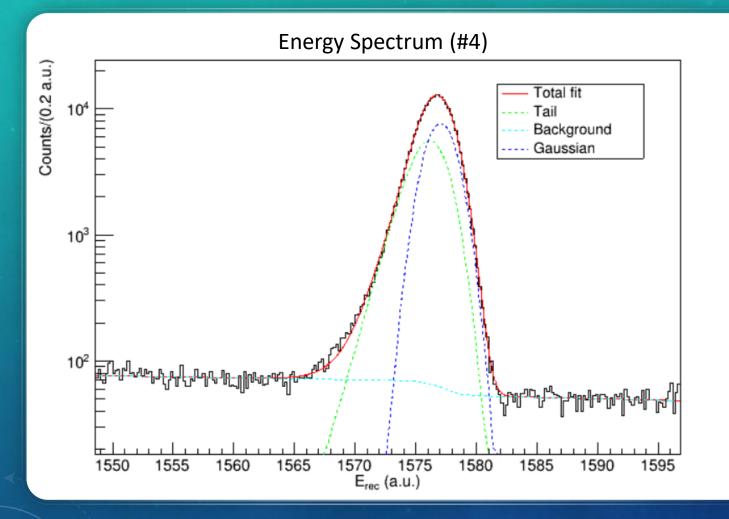


Peak at 1408 keV

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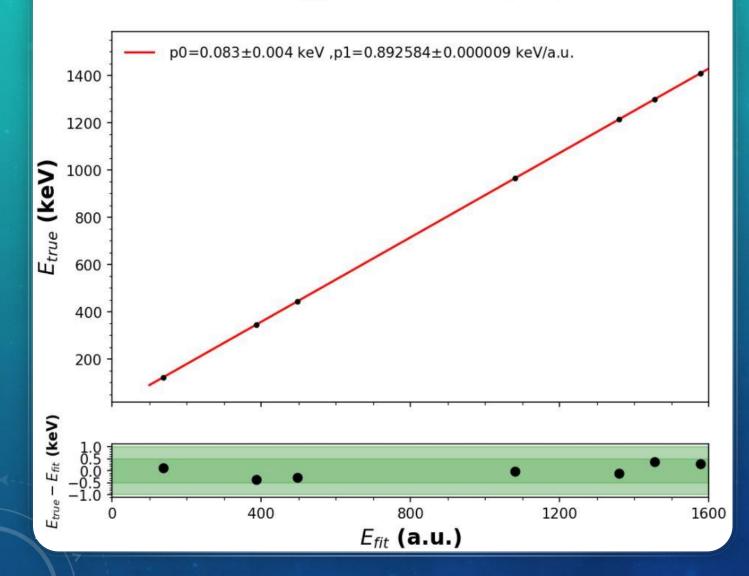
- A few peaks do not give a good fit, Error Matrix has some uncertainty, or it is not positive-definite
- Look at every fit and exclude those lines
- Lines excluded: 778.904 keV, 1112.074 keV, 1212.948 keV
- Fit rest of the lines, status "CONVERGED" and Error Matrix "ACCURATE"
- Intercept is very small (~1E-11) $_{1/31/2022}$ (~1E-11)



• Peak at 1408 keV

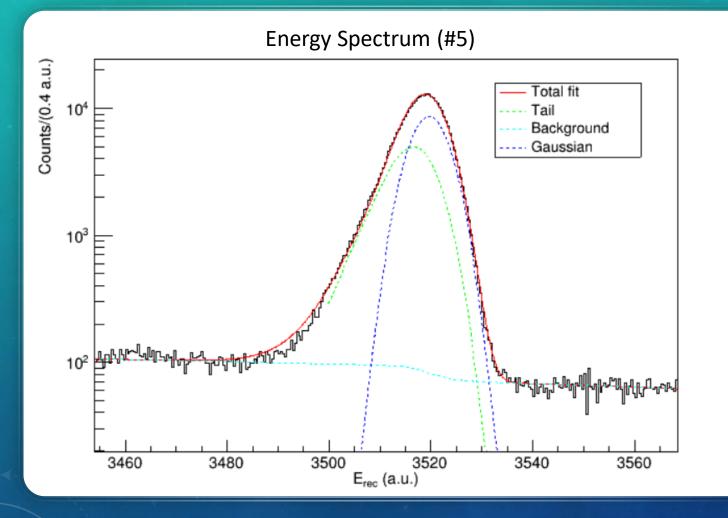
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Energy calibration (#4)



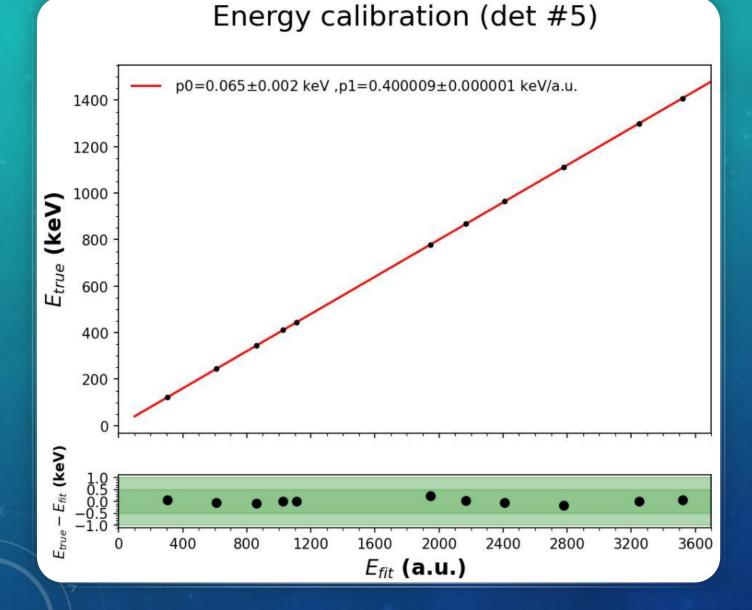
DETECTOR #4

- Among all the detectors, had to exclude 5 lines (maximum)
- Lines excluded: 244.6975 keV, 411.1163 keV, 778.904 keV, 867.378 keV, 1112.074 keV

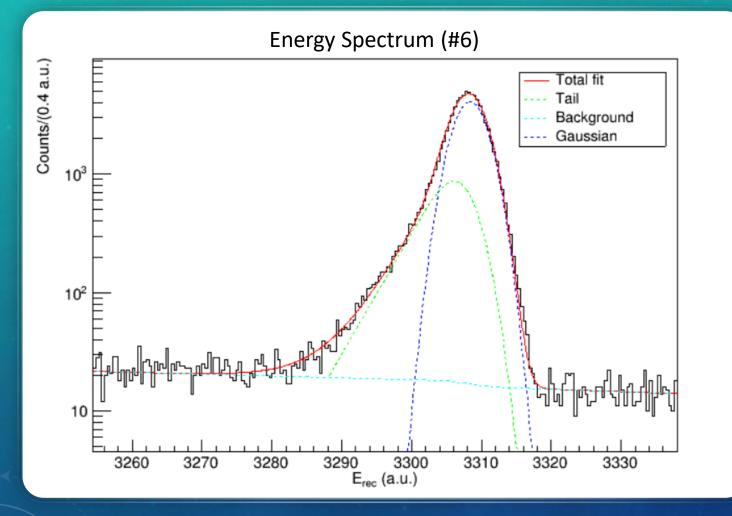


• Peak at 1408 keV

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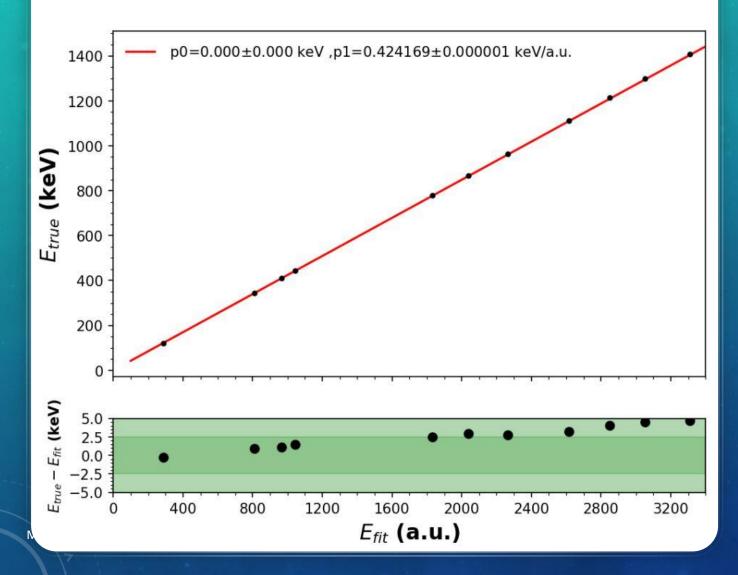
• Lines excluded: 1212.948 keV



• Peak at 1408 keV

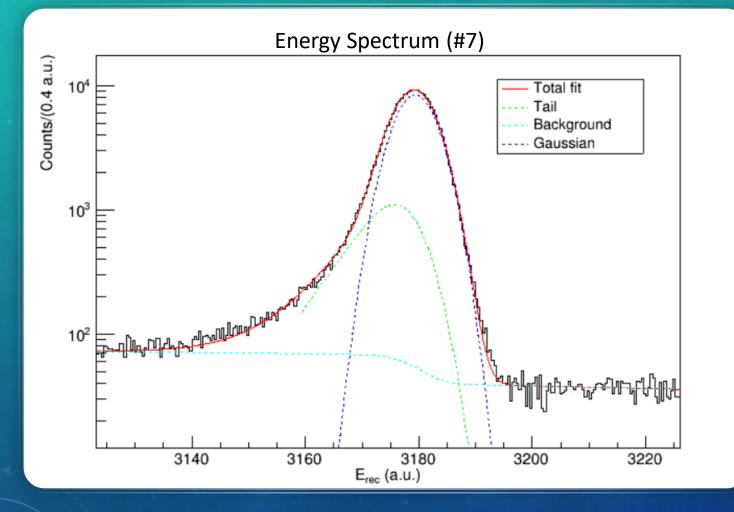
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Energy calibration (det #6)



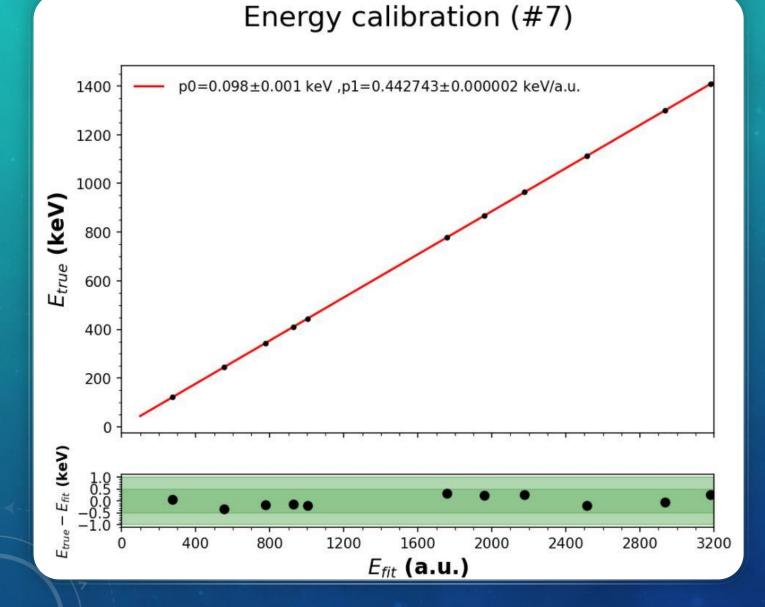
DETECTOR #6

- Lines excluded: 244.6975 keV
- Intercept is very small (~1E-11)

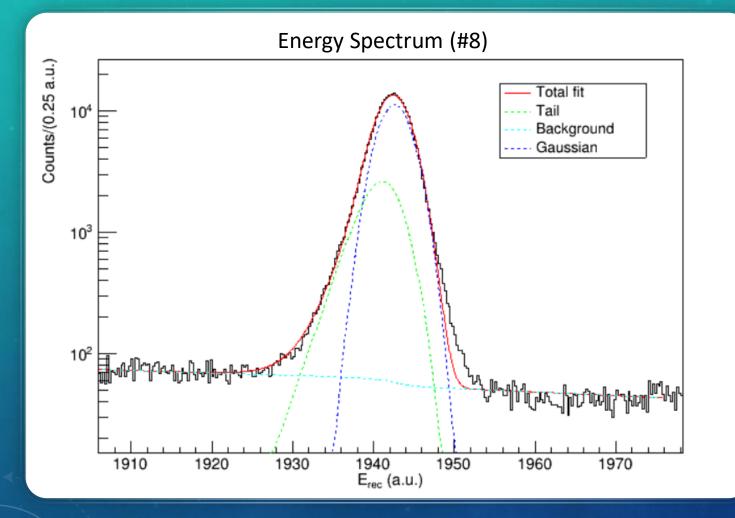


Peak at 1408 keV

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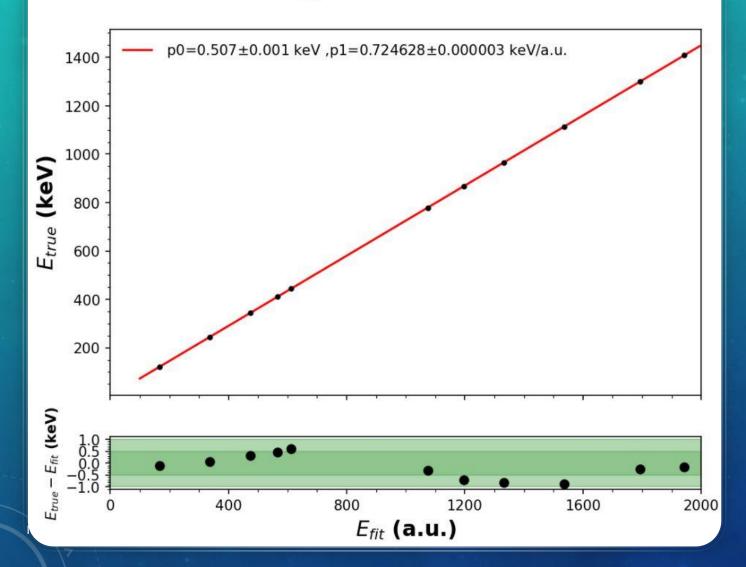
• Lines excluded: 1212.948 keV



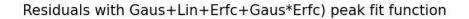
Peak at 1408 keV

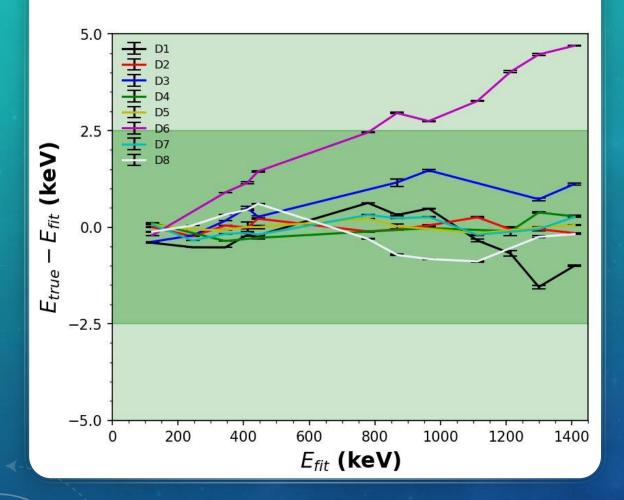
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• Lines excluded: 1212.948 keV



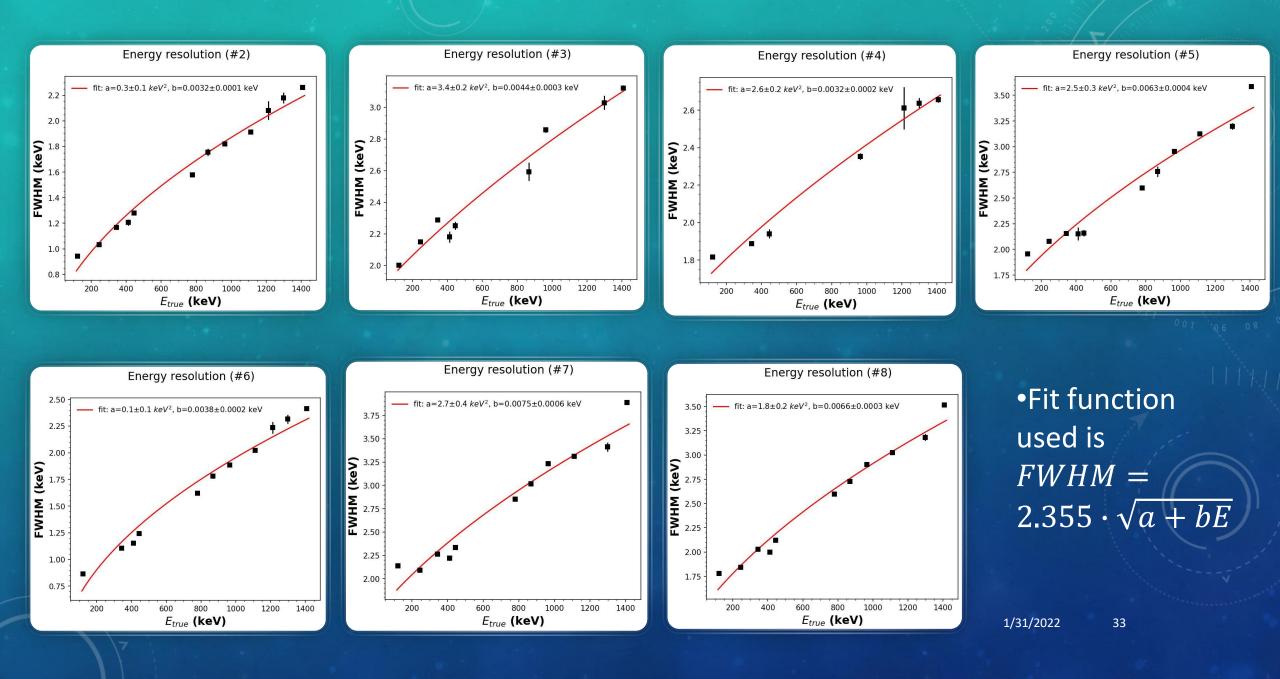


RESIDUALS

- For different detectors, range of residuals varies in different ranges
- We need to look in detail at det #6 and understand why the residuals are worse than usual here

ENERGY RESOLUTION

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ALL PARAMETERS IN TABLE

Detector #	Energy calibration		Residuals	Energy resolution		
						FWHM @1.3 MeV
	Fit func. y = m*x+c		E _{true} - E _{fit}	Fit func. y= (a+b*x)^0.5		
	Slope (m) (keV/a.u.)	Intercept (c) (keV)		a (keV)^2	b (keV)	(keV)
1	0.8	0	(-1.5)- (0.6)	7.93	0.007	3.9
2	0.46	0.17	(-0.24)- (0.26)	0.56	0.003	2.1
3	0.78	0	(-0.38)- (1.47)	3.45	0.005	3
4	0.89	0.08	(-0.36)- (0.38)	2.91	0.003	2.6
5	0.4	0.07	(-0.16)- (0.22)	2.91	0.006	3.3
6	0.42	0	(-0.21)- (4.69)	0.38	0.003	2.2
7	0.44	0.1	(-0.35)- (0.32)	3.49	0.007	3.5
8	0.72	0.51	(-0.89)- (0.61)	2.27	0.006	3.2

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SUMMARY

- We updated the fit function and applied some quality cuts suggested by TUM
- Using the updated method, FWHM at 1.3 MeV is 3.9 keV for detector #1 and <4keV for other detectors
- A new algorithm to fit peaks in the energy spectrum was applied which seems to work very well to a first approximation
- Not all peaks were included for energy calibration
- In the final study, this technique could help us understand how the results change when we change our approach



- Analyze the change in calibration constants using data from calibration runs on different days
- Apply other cuts to data after further discussion in upcoming LLAMA analysis meetings
- Redo the entire process with trapezoidal filter