Complete Covered Robininy I. B. Pestov BLTP, JINR	Complete General Relativity
Who?	I. B. Pestov BLTP, JINR
When?	August 11, 2022, SQS'22

- 《 ロ ト 《 國 ト 《 国 ト 《 国 ト 《 国 - の �)

General Relativity

I. B. Pestov BLTP, JINR

Einstein Equation

$$R_{ij}-\frac{1}{2}g_{ij}R=T_{ij}$$

I. B. Pestov BLTP, JINR The presentation is organized as follows.

1 Parallel displacement and Einstein equation

2 General Gauge Relativity

3 Mechanism of mass generation of General Gauge Field

Equations of Complete General Relativity

Physical aspects of Complete General Relativity. 🗉 🔊

I. B. Pestov BLTP, JINR

4

5

I. B. Pestov BLTP, JINR

Parallel displacement and the gravitational field

$$\delta V^i = dV^i + \Gamma^i_{jk} dx^j V^k = 0,$$

where

$$\Gamma^{i}_{jk} = rac{1}{2}g^{il}(\partial_{i}g_{jl} + \partial_{j}g_{il} - \partial_{l}g_{ij}).$$

The commutator

$$(\nabla_i \nabla_j - \nabla_j \nabla_i) V^k = R_{ijl}{}^k V^l$$

defines

$$R_{ijl}{}^{k} = \partial_{i}\Gamma_{jl}^{k} - \partial_{j}\Gamma_{il}^{k} + \Gamma_{in}^{k}\Gamma_{jl}^{n} - \Gamma_{jn}^{k}\Gamma_{il}^{n}$$

The antisymmetrical tensor (trace) $R_{ijk}^{\ k}$ is trivial. But the symmetrical tensor $R_{jl} = R_{kjl}^{\ k}$ and the scalar $R = R_{jl}g^{jl}$ define absolutely the l. h. s. of the Einstein equation.

I. B. Pestov BLTP, JINR

Complete General Relativity

I. B. Pestov BLTP, JINR

The general parallel displacement

$$\delta V^i = dV^i + P^i_{jk} dx^j V^k = 0$$

where the connection P_{jk}^{i} is considered as a primary entity. To define the Principle of General Gauge Relativity, we start from the consideration of the linear operators in the space of vector fields V^{i} . Non-degenerate linear transformation has the form

$$\overline{V}^i=S^i_jV^j,\quad {
m Det}(S^i_j)
eq {\sf 0},$$

where S_j^i is a tensor field of the second rank. These local internal transformations form a group of General Gauge Relativity with an associative binary operation $P_i^i = S_k^i T_i^k$. This is the general gauge group.

I. B. Pestov BLTP, JINR

From the law of parallel displacement we have

$$\overline{P}^i_{jk} = S^i_m P^m_{jn} S^{-1}{}^n_k + S^i_m \partial_j S^{-1}{}^m_k$$

where $S^{-1}{}_{j}^{k}$ are the components of the operator S^{-1} inverse to the operator S, $S_{k}^{i}S^{-1}{}_{j}^{k} = \delta_{j}^{i}$. For the gauge covariant derivative $\delta V^{i} = D_{j}V^{i}dx^{j}$

$$D_j V^i = \partial_j V^i + P^i_{jk} V^k$$

we have

$$\overline{\mathrm{D}}_{j}\overline{V}^{i}=S_{k}^{i}\mathrm{D}_{j}V_{\mathrm{CP}}^{k}$$
 , is the set of \mathcal{O}

I. B. Pestov BLTP, JINR The commutator of the gauge covariant derivatives $(D_i D_j - D_j D_i)V^k = [D_i, D_j]V^k = H_{ijl}{}^k V^l,$ gives the curvature tensor of the general gauge field $H_{ijl}{}^k = \partial_i P_{jl}^k - \partial_j P_{il}^k + P_{in}^k P_{jl}^n - P_{jn}^k P_{il}^n.$ The extinct metric tensor (trace of the curvature)

The antisymmetric tensor (trace of the curvature tensor)

$$F_{ij} = H_{ijk}{}^k = \partial_i P_{jk}^k - \partial_j P_{ik}^k$$

is non trivial here and should be considered separately from the irreducible (traceless) tensor of curvature

$$I_{ijl}^{\ k} = H_{ijl}^{\ k} - \frac{1}{4} H_{ijn}^{\ n} \delta_l^k, \quad \exists I_{ijl}^{\ l} = 0 \quad \exists \quad n \in \mathbb{N}$$

I. B. Pestov BLTP, JINR For brevity, in what follows we will use the matrix notation

$$\mathbf{S} = (S_I^k), \quad \mathbf{P}_i = (P_{il}^k), \quad \mathbf{E} = (\delta_I^k),$$

$$\mathbf{H}_{ij} = (H_{ijl}{}^k), \quad \mathrm{Tr}\,\mathbf{S} = S_k^k.$$

The transformations of gauge symmetry in question take the form

 $\overline{\mathbf{P}}_i = \mathbf{S}\mathbf{P}_i\mathbf{S}^{-1} + \mathbf{S}\partial_i\mathbf{S}^{-1} = \mathbf{P}_i + \mathbf{S}\mathbf{D}_i\mathbf{S}^{-1},$

where D_i is the gauge covariant derivative

 $D_i \mathbf{S} = \partial_i \mathbf{S} + \mathbf{P}_i \mathbf{S} - \mathbf{S} \mathbf{P}_i = \partial_i \mathbf{S} + [\mathbf{P}_i, \mathbf{S}]. = Oac$

I. B. Pestov BLTP, JINR

For the curvature tensor we have

$$D_i \mathbf{H}_{jk} = \partial_i \mathbf{H}_{jk} + [\mathbf{P}_i, \mathbf{H}_{jk}]$$

and the general covariant identity

$$D_i \mathbf{H}_{jk} + D_j \mathbf{H}_{ki} + D_k \mathbf{H}_{ij} = 0$$

holds.

Under the transformations of the general gauge group the curvature tensor of the connection P^i_{ik}

$$\mathsf{H}_{ij} = \partial_i \mathsf{P}_j - \partial_j \mathsf{P}_i + [\mathsf{P}_i, \mathsf{P}_j]$$

transforms homogeneously

$$\overline{\mathbf{H}}_{ij} = \mathbf{S}\mathbf{H}_{ij}\mathbf{S}^{-1}.$$

I. B. Pestov BLTP, JINR

The ground state of the general gauge field is defined by the equation

$$\mathbf{H}_{ij}=0,$$

which has nontrivial solution

$$P^i_{jk} = L^i_{jk} = E^i_\mu \partial_j E^\mu_k$$

(a linear connection of the ground state).

I. B. Pestov BLTP, JINR A transition from the ground state to the exited one is characterized by the tensor of transition

$$T^i_{jk}=P^i_{jk}-L^i_{jk},$$

with a simple (homogeneous) law of transformation

$$\overline{T}_{jk}^{i} = S_{m}^{i} Q_{jn}^{m} S^{-1}{}_{j}^{k}, \quad \overline{\mathbf{T}}_{i} = \mathbf{S} \mathbf{T}_{i} \mathbf{S}^{-1}$$

and the traceless tensor

$$Q_{jk}^i = T_{jk}^i - rac{1}{4}T_{jl}^l \delta_k^i, \quad \mathbf{Q}_j = \mathbf{T}_j - rac{1}{4}\mathrm{Tr}(\mathbf{T}_j)\mathbf{E}.$$

I. B. Pestov BLTP, JINR

We put

$$\mathcal{L}_{P} = -rac{1}{4} \mathrm{Tr}(\mathbf{I}_{ij}\mathbf{I}^{ij}) + rac{\mu^{2}}{2} \mathrm{Tr}(\mathbf{Q}_{i}\mathbf{Q}^{i}), \quad \mathcal{L}_{em} = -rac{1}{4}F_{ij}F^{ij}.$$

The general covariant and general gauge covariant Lagrangian $\mathcal L$ of the general gauge field takes the form

$$\mathcal{L} = \alpha \mathcal{L}_{P} + \beta \mathcal{L}_{em} =$$

$$\alpha[-\frac{1}{4}\mathrm{Tr}(\mathbf{I}_{ij}\mathbf{I}^{ij})+\frac{\mu^2}{2}\mathrm{Tr}(\mathbf{Q}_i\mathbf{Q}^i)]+\beta[-\frac{1}{4}F_{ij}F^{ij}],$$

where μ is a constant of dimension of cm^{-1} , and α and β are dimensionless parameters, dimension of the P_{jk}^i is equal to cm^{-1} and the action is dimensionless, $\mathbf{I}^{ij} = g^{ik}g^{jl}\mathbf{I}_{kl}, \quad \mathbf{Q}^i = g^{ik}\mathbf{Q}_k.$

I. B. Pestov BLTP, JINR

By varying the Lagrangian \mathcal{L} with respect to \mathbf{P}_i , the following equation holds $\alpha[\frac{1}{\sqrt{g}}\mathrm{D}_{i}(\sqrt{g}\mathsf{I}^{ij}) + \mu^{2}\mathsf{Q}^{j}] + \beta[\frac{1}{\sqrt{g}}\partial_{i}(\sqrt{g}F^{ij})\mathsf{E}] = 0, (1)$ where $g = -\text{Det}(g_{ii})$. Taking trace from equation (1), we find that $rac{1}{\sqrt{g}}\partial_i(\sqrt{g}F^{ij})=0,$ (2)since $\operatorname{Tr}(\mathbf{I}^{ij}) = \operatorname{Tr}(\mathbf{Q}^{j}) = 0$. Hence, $\frac{1}{\sqrt{g}}\mathrm{D}_i(\sqrt{g}\mathbf{I}^{ij}) + \mu^2 \mathbf{Q}^j = 0.$ (3)From (3) it follows that \mathbf{Q}^i has to satisfy the equation $\frac{1}{\sqrt{\sigma}} \mathbf{D}_i(\sqrt{g} \mathbf{Q}^i) = 0, \qquad (4)$

I. B. Pestov BLTP, JINR because $D_i D_j (\sqrt{g} \mathbf{I}^{ij}) = 0$. From the Lagrangian \mathcal{L} it follows that in Complete General Relativity the r. h. s. of the Einstein equation (energy-momentum tensor) is defined by the curvature tensor of the general gauge field as follows:

$$T_{ij} = \alpha \left[-\operatorname{Tr}\left(\mathbf{I}_{ik}\mathbf{I}_{j}^{k}\right) - g_{ij}\mathcal{L}_{P} + \mu^{2}\operatorname{Tr}\left(\mathbf{Q}_{i}\mathbf{Q}_{j}\right)\right] + \beta \left[-F_{il}F_{j}^{l} + \frac{1}{4}F_{kl}F^{kl}g_{ij}\right],$$

where $\mathbf{I}_{j}^{k} = \mathbf{I}_{jl}g^{kl}$.

One can show that the energy-momentum tensor in question satisfies the equation

$$\nabla^i T_{ij} = 0,$$

where ∇_i denotes the covariant derivative with respect to the connection belonging to metric g_{ij} , where $\nabla_i = -\infty$

I. B. Pestov BLTP, JINR

I. B. Pestov BLTP, JINR

At last, we write the Lagrangian

$$\mathcal{L}_g = -\frac{l_g}{2}R,$$

where l_g is a constant of dimension cm^{-2} and the Einstein equation takes the form

$$I_{g}G_{ij} = \alpha [-\mathrm{Tr} \left(\mathbf{I}_{ik}\mathbf{I}_{j}^{k}\right) - g_{ij}\mathcal{L}_{P} + \mu^{2}\mathrm{Tr} \left(\mathbf{Q}_{i}\mathbf{Q}_{j}\right)] +$$

$$\beta[-F_{il}F_j^{\ l}+\frac{1}{4}F_{kl}F^{kl}g_{ij}].$$

$$\mathcal{L}_P = -rac{1}{4} \mathrm{Tr}(\mathbf{I}_{ij}\mathbf{I}^{ij}) + rac{\mu^2}{2} \mathrm{Tr}(\mathbf{Q}_i\mathbf{Q}^i)$$
.

I. B. Pestov BLTP, JINR From the Einstein equation it follows that the interactions between the different states of the general gauge field are realized by a graviton exchange. This interaction can be characterized by an angle of mixing. In accordance with the Einstein equation, we can put

$$\sin \varphi = \frac{\alpha}{\sqrt{\alpha^2 + \beta^2}}, \quad \cos \varphi = \frac{\beta}{\sqrt{\alpha^2 + \beta^2}}$$

and redefine l_g .

I. B. Pestov BLTP, JINR Thus, the general gauge field has two states: the familiar electromagnetic field

 $F_{ij} = \mathrm{Tr}\mathbf{H}_{ij}$

(which should be considered as its singlet state) and more general state

$$\mathsf{I}_{ij} = \mathsf{H}_{ij} - rac{1}{4} \mathrm{Tr} \mathsf{H}_{ij} \mathsf{E}$$

that can be called the general electromagnetic field. On this reason we will call by mphoton a particle that corresponds to a definite state of the general electromagnetic field. We can treat μ as a mass of the mphoton(scaling invariance is broken). These particles are the only source of the gravitational field. $\Xi = - \Xi$

I. B. Pestov BLTP, JINR To make a clear the spontaneous broken symmetry we first of all mention that the strength tensor I_{ij} can be written in terms of the irreducible tensor Q_i only, since

$$\mathbf{I}_{ij} = \stackrel{o}{\mathrm{D}}_i \, \mathbf{Q}_j - \stackrel{o}{\mathrm{D}}_j \, \mathbf{Q}_i + \mathbf{Q}_i \mathbf{Q}_j - \mathbf{Q}_j \mathbf{Q}_i$$

Here $\overset{\circ}{\mathbf{D}}_i$ denotes the gauge covariant derivative with respect to the connection \mathbf{L}_i of the ground state and, hence, $[\overset{\circ}{\mathbf{D}}_i, \overset{\circ}{\mathbf{D}}_j] = 0$. For the antisymmetric tensor F_{ij} we obtain

 $F_{ij} = \partial_i P_{jk}^k - \partial_j P_{ik}^k = \partial_i (L_{jk}^k + T_{jk}^k) - \partial_j (L_{ik}^k + T_{ik}^k) = \partial_i T_{jk}^k - Q_{ik}^k$ since $\operatorname{Tr}(\mathbf{L}_i) = \partial_i \ln |p|$, $p = \operatorname{Det}(E_i^{\mu})$. Thus, we can consider the tensor field Q_{ij}^k with the constraints $Q_{ik}^k = 0$ and covariant vector field $A_i = T_{ik}^k$ as

I. B. Pestov BLTP, JINR

independent quantities, which obey the equations

$$\frac{1}{\sqrt{g}}\partial_i(\sqrt{g}F^{ij}) = 0, \quad F_{ij} = \partial_i A_j - \partial_j A_i, \quad (5)$$

$$\frac{1}{\sqrt{g}} \overset{o}{\mathbf{D}}_{i} \left(\sqrt{g} \mathbf{I}^{ij} \right) + \left[\mathbf{Q}_{i}, \mathbf{I}^{ij} \right] + \mu^{2} \mathbf{Q}^{j} = 0, \qquad (6)$$

and

$$\frac{1}{\sqrt{g}} \overset{o}{\mathbf{D}}_i \left(\sqrt{g} \mathbf{Q}^i \right) + \left[\mathbf{Q}_i, \mathbf{Q}^i \right] = 0.$$
 (7)

Since the trace of \mathbf{I}_{ij} is equal to zero, it is clear why we need to consider an traceless tensor \mathbf{Q}^i . In our case, the trace of \mathbf{Q}^i is trivial and equation (7) is compatible.

I. B. Pestov BLTP, JINR

Complete General Relativity

I. B. Pestov BLTP, JINR

It is interesting to pay attention to the following analogy between gravity and electromagnetism. In the electron theory of Lorentz, the right- hand side of the Maxwell equations was presented with continues phenomenological distributions of charge and current. With the discovery of quantum mechanics or more exactly the Schrodinger and Dirac equations the details of the right- hand side in this case were clarified. But the physical content of the Maxwell–Dirac equations was disclosed only in the framework of quantum electrodynamics. We see the same situation in the Einstein theory of the gravitational field. It is clear that the investigation of Complete General Relativity as a closed system in the framework of quantum field theory is an urgent problem. Hence we need to look for hidden possibilities to solve the renormalization problem.

I. B. Pestov BLTP, JINR

Whilst the need for invisible matter was established almost a century ago, only its gravitational interaction has been conformed so far. Hence, it is natural to suppose that mphotons can represent the invisible matter and there is no reason for a plethora of models for this matter often called dark matter. In the framework of Complete General Relativity the photons represent the Cosmic Microwave Background and the mphotons correspond to the so called Weakly Interacting Massive Particles (WIMPs). But from the Einstein equation it follows that the Cosmic Microwave Background and invisible matter are tightly connected and, hence, the investigations of CMB can provide the discovery of hidden properties of WIMPs (or mphotons).

Complete General 5

I. B. Pestov BLTP, JINR We believe that main equations of Complete General Relativity provide a justified basis for discovering realistic cosmological models and new insight into the problem of the gravitational waves.

THANK YOU FOR ATTENTION