

# Gravitational waves from first-order electroweak phase transition in a model with light sgoldstinos (based on arXiv:2112.06083)

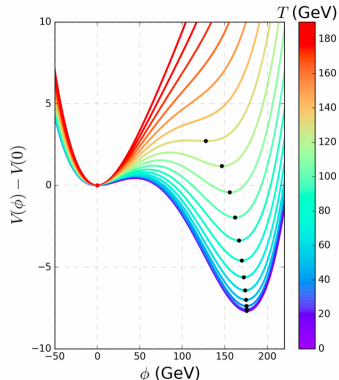
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High-Energy Physics, and Cosmology,  
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# Sgoldstino and the 1st-order EW phase transition

- ▶ Model with the low-energy SUSY breaking (10-100 TeV) in the hidden sector.
- ▶ Supermultiplet of the Goldstone fermion **goldstino** and the scalar **sgoldstino**  
 $\Phi = \phi + \sqrt{2}\theta G + F_\phi \theta^2$ .
- ▶  $\langle F_\phi \rangle = F \neq 0$  — spontaneous SUSY breaking.
- ▶ Can we search for sgoldstino using some **cosmological** signatures?



- ▶ In the extended SM with an additional scalar: 1st-order electroweak phase transition → gravitational waves.
- ▶ **Questions:** Is the 1st-order EWPT possible in this model? What are the expected GW signals?

# Lagrangian of supersymmetric model with sgoldstino

$$\mathcal{L} = \mathcal{L}_K + \mathcal{L}_W + \mathcal{L}_{gauge} + \mathcal{L}_\Phi,$$

$$\mathcal{L}_K = \int d^2\theta d^2\bar{\theta} \sum_k \left( 1 - \frac{m_k^2}{F^2} \Phi^\dagger \Phi \right) \Phi_k^\dagger e^{g_1 V_1 + g_2 V_2 + g_3 V_3} \Phi_k,$$

$$\begin{aligned} \mathcal{L}_W = \int d^2\theta \epsilon_{ij} & \left( \left( \mu - \frac{B}{F} \Phi \right) H_D^i H_U^j + \left( Y_{ab}^L + \frac{A_{ab}^L}{F} \Phi \right) L_a^j E_b^c H_D^i + \right. \\ & \left. + \left( Y_{ab}^D + \frac{A_{ab}^D}{F} \Phi \right) Q_a^j D_b^c H_D^i + \left( Y_{ab}^U + \frac{A_{ab}^U}{F} \Phi \right) Q_a^i U_b^c H_U^j \right) + h.c., \end{aligned}$$

$$\mathcal{L}_{gauge} = \frac{1}{4} \sum_a \int d^2\theta \left( 1 + \frac{2M_a}{F} \Phi \right) \text{Tr} W_\alpha W^\alpha + h.c.,$$

$$\begin{aligned} \mathcal{L}_\Phi = \int d^2\theta d^2\bar{\theta} & \left( \Phi^\dagger \Phi - \frac{\widetilde{m}_s^2 + \widetilde{m}_p^2}{8F^2} (\Phi^\dagger \Phi)^2 - \frac{\widetilde{m}_s^2 - \widetilde{m}_p^2}{12F^2} (\Phi^\dagger \Phi^3 + \Phi^\dagger{}^3 \Phi) - \right. \\ & - \frac{\delta\lambda_2}{4F^2} H_U^\dagger H_U (\Phi^\dagger \Phi)^2 - \frac{\delta\lambda_3}{9F^2} (\Phi^\dagger \Phi)^3 - \frac{\delta\lambda_4}{3F^2} H_U^\dagger H_U (\Phi^\dagger \Phi^3 + \Phi^\dagger{}^3 \Phi) - \\ & - \frac{\delta\lambda_5}{5F^2} (\Phi^\dagger \Phi^5 + \Phi^\dagger{}^5 \Phi) - \frac{\delta\lambda_6}{8F^2} (\Phi^\dagger{}^2 \Phi^4 + \Phi^\dagger{}^4 \Phi^2) - \frac{\delta\mu_1}{2F^2} H_U^\dagger H_U (\Phi^\dagger \Phi^2 + \Phi^\dagger{}^2 \Phi) - \\ & \left. - \frac{\delta\mu_2}{6F^2} (\Phi^\dagger{}^3 \Phi^2 + \Phi^\dagger{}^2 \Phi^3) - \frac{\delta\mu_3}{4F^2} (\Phi^\dagger{}^4 \Phi + \Phi^\dagger \Phi^4) - \frac{\delta C_3}{2F^2} (\Phi^\dagger{}^2 \Phi + \Phi^\dagger \Phi^2) \right) - \left( \int d^2\theta F \Phi + h.c. \right) \end{aligned}$$

## Tree-level potential in model with sgoldstino

1. Model with low-energy SUSY breaking: MSSM, goldstino-MSSM interaction, goldstino multiplet

$$\mathcal{L} = \mathcal{L}_{MSSM} + \mathcal{L}_{soft} + \mathcal{L}_\Phi$$

2. Scalar fields tree-level potential

$$V = V(h_d, h_u, \phi), \quad h_d = \begin{pmatrix} h_d^0 \\ H^- \end{pmatrix}, \quad h_u = \begin{pmatrix} H^+ \\ h_u^0 \end{pmatrix}$$

3. Decoupling limit at low energies

$$h_u \rightarrow \mathcal{H} \sin \beta, \quad h_d \rightarrow -\epsilon \mathcal{H}^* \cos \beta, \quad \tan \beta \equiv v_u/v_d$$

G. F. Giudice and A. Romanino, *Nucl. Phys. B* **699** (2004), 65-89

[arXiv:hep-ph/0406088].

4. Extract Higgs field, Goldstone bosons, scalar and pseudoscalar sgoldstino

$$\mathcal{H} = \frac{1}{\sqrt{2}} \begin{pmatrix} G^+ \\ h + iG^0 \end{pmatrix}, \quad \phi = \frac{1}{\sqrt{2}}(s + ip).$$

5. Potential of three scalar fields  $V_0 = V_0(h, s, p)$

## Potential at zero temperature

$$V_{T=0}(h, s, p) = V_0 + V_{CW} + V_{CT}$$

### Tree-level potential

$$V_0(h, s, p) = \frac{\lambda_1}{4} h^4 + \frac{\lambda_{hs}}{4} h^2 s^2 + \frac{\lambda_{hp}}{4} h^2 p^2 + \frac{\lambda_s}{4} s^4 + \frac{\lambda_p}{4} p^4 + \frac{\lambda_{sp}}{4} s^2 p^2 + \\ + \frac{\mu_1}{2} s h^2 + \frac{\mu_s}{6} s^3 + \frac{\mu_{sp}}{2} s p^2 - \frac{\tilde{M}_1^2}{2} h^2 + \frac{M_s^2}{2} s^2 + \frac{M_p^2}{2} p^2 + C_3 s.$$

### One-loop corrections at $T = 0$

$$V_{CW}(h, s, p) = \frac{1}{64\pi^2} \sum_i (-1)^{s_i} n_i m_i^4(h, s, p) \left( \log \frac{m_i^2(h, s, p)}{Q^2} - c_i \right).$$

### Counterterms: potential minimum at $T = 0$ in $(v, 0, 0)$

$$V_{CT}(h, s, p) = \frac{\delta\lambda_1}{4} h^4 + \frac{\delta\lambda_{hs}}{4} h^2 s^2 + \frac{\delta\lambda_{hp}}{4} h^2 p^2 + \frac{\delta\mu_1}{2} s h^2 - \frac{\delta\tilde{M}_1^2}{2} h^2 + \delta C_3 s.$$

# Effective potential at finite temperature

$$V_{\text{eff}}(T, h, s, p) = V_{T=0} + V_T + V_d$$

One-loop **thermal** corrections

$$V_T(T, h, s, p) = \frac{T^4}{2\pi^2} \sum_i n_i J_{B/F} \left( \frac{m_i(h, s, p)}{T} \right),$$

L. Dolan and R. Jackiw, *Phys. Rev. D* **9** (1974) 3320.

Corrections from **daisy diagrams** resummation

$$V_d(T, h, s, p) = -\frac{T}{12\pi} \sum_i a_i n_i \left( (m_{Ti}^2(T, h, s, p))^{3/2} - (m_i^2(h, s, p))^{3/2} \right),$$

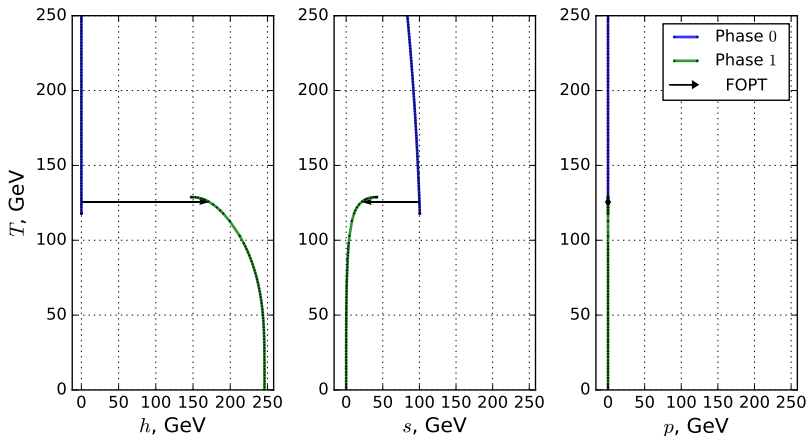
M. E. Carrington, *Phys. Rev. D* **45** (1992) 2933,

P. B. Arnold and O. Espinosa, *Phys. Rev. D* **47** (1993) 3546.

# Benchmark model with the first-order phase transition

Minima of effective potential are plotted using the package  
PhaseTracer

P. Athron et al., *Eur. Phys. J. C* **80** (2020) no.6, 567 [arXiv:2003.02859]



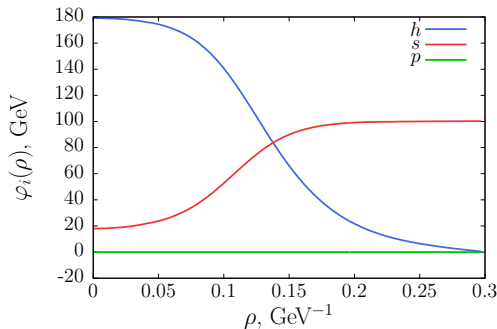
# Bubbles of new phase: nucleation and percolation

Bubble production rate per unit volume

$$P \sim \mathcal{A}(T) \exp\left\{-\frac{S_3}{T}\right\}, \quad \left.\frac{S_3}{T}\right|_{T=T_{\text{nuc}}} = 140.$$

Phase transition characteristics:  $T_{\text{nuc}}$ ,  $T_{\text{perc}}$ ,  $\alpha$ ,  $\beta/H_C$ .

$$\alpha \equiv \left.\frac{\text{latent heat}}{\text{radiation energy}}\right|_{T=T_{\text{perc}}}, \quad \frac{\beta}{H_C} \equiv T \left.\frac{d}{dT}\left(\frac{S_3}{T}\right)\right|_{T_{\text{perc}}}.$$



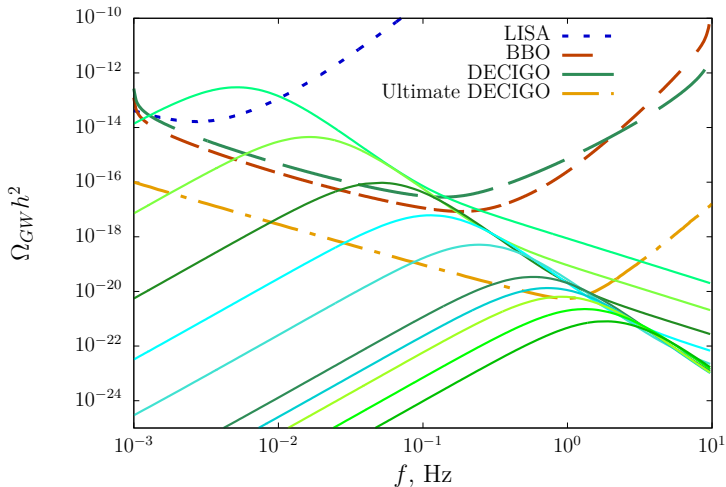
Bounce solutions are found using the package FindBounce V. Guada, M. Nemevšek and M. Pintar, *Comput. Phys. Commun.* **256** (2020), 107480 [arXiv:2002.00881]



# Gravitational wave spectra

Bubbles nucleate and collide  $\rightarrow$  sound waves and MHD turbulence in plasma  $\rightarrow$  gravitational waves

$$\Omega_{GW} h^2 = \Omega_{sw} h^2 + \Omega_m h^2$$

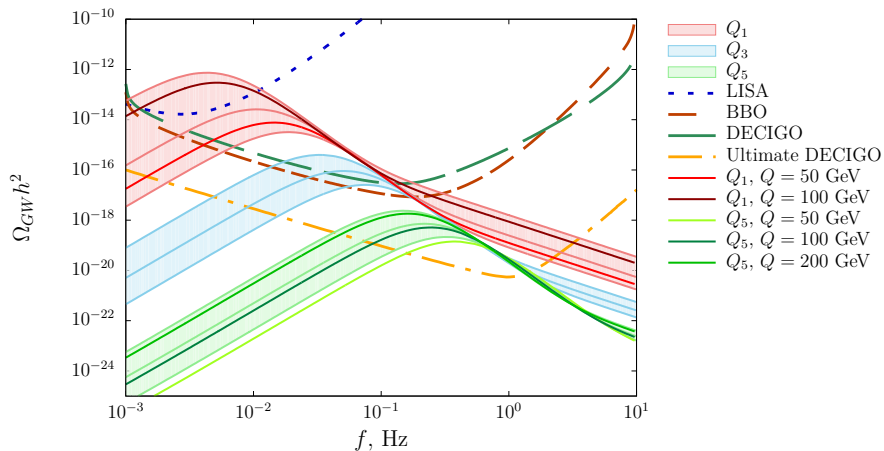


# Theoretical uncertainties in GW spectra predictions

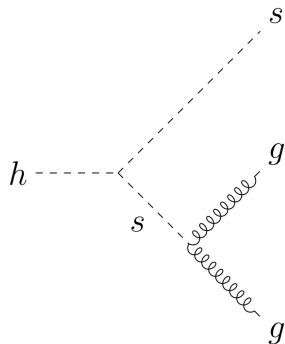
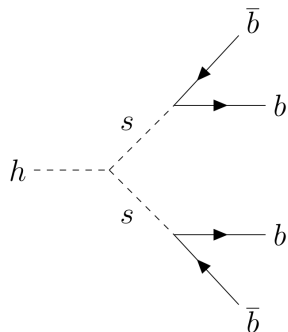
BM	Q, GeV	$T_{\text{nucl}}$ , GeV	$T_{\text{perc}}$ , GeV	$\alpha \cdot 10^3$	$\beta/H_c$
$Q_1$	109.7	60.71	54.88	168	228
	68.4	72.59	68.49	62.5	442
	39.3	81.75	78.63	35.9	683
$Q_3$	181.4	93.17	90.71	20.8	1040
	97.9	99.80	97.86	15.0	1490
	51.9	105.28	103.74	11.6	2060
$Q_5$	231.9	116.94	115.93	7.06	3670
	120.3	121.02	120.27	5.77	5190
	62.0	124.53	123.98	4.80	7530

**Table:** The renormalization scale  $Q$ , nucleation and percolation temperatures, parameters  $\alpha$  and  $\beta/H_c$  for BM  $Q_1$ ,  $Q_3$ ,  $Q_5$ . For each point we calculate the first order electroweak phase transition parameters taking the renormalization scale  $2T_{\text{perc}}$ ,  $T_{\text{perc}}$  and  $T_{\text{perc}}/2$ .

# Theoretical uncertainties in GW spectra predictions



# Possible exotic Higgs boson decays with sgoldstino



# Results

- ▶ We have shown that the first-order electroweak phase transition is possible in the model with low-scale supersymmetry breaking.
- ▶ In benchmark models the 1st-order PT takes place at  $T \approx 60 - 140$  GeV.
- ▶ It may be possible for LISA and the suggested observatories BBO, DECIGO to search for gravitational waves produced during the PT in benchmark models.

Thank you for your attention!

## Counterterms potential

The field values in the minimum of zero-temperature potential  $V_0 + V_{CW} + V_{CT}$  are  $h = v$ ,  $s = 0$ ,  $p = 0$  if we choose

$$V_{CT} = \frac{\delta\lambda_1}{4} h^4 + \frac{\delta\lambda_{hs}}{4} h^2 s^2 + \frac{\delta\lambda_{hp}}{4} h^2 p^2 + \frac{\delta\mu_1}{2} s h^2 - \frac{\delta M_1^2}{2} h^2 + \delta C^3 s,$$

$$\delta\lambda_1 = -\frac{1}{2v^2} \left. \frac{\partial^2 V_{CW}}{\partial h^2} \right|_{(v,0,0)} + \frac{1}{2v^3} \left. \frac{\partial V_{CW}}{\partial h} \right|_{(v,0,0)},$$

$$\delta\lambda_{hs} = -\frac{2}{v^2} \left. \frac{\partial^2 V_{CW}}{\partial s^2} \right|_{(v,0,0)}, \quad \delta\lambda_{hp} = -\frac{2}{v^2} \left. \frac{\partial^2 V_{CW}}{\partial p^2} \right|_{(v,0,0)},$$

$$\delta\mu_1 = -\frac{1}{v} \left. \frac{\partial^2 V_{CW}}{\partial h \partial s} \right|_{(v,0,0)}, \quad \delta C^3 = -\left. \frac{\partial V_{CW}}{\partial s} \right|_{(v,0,0)} + \frac{v}{2} \left. \frac{\partial^2 V_{CW}}{\partial h \partial s} \right|_{(v,0,0)},$$

$$\delta M_1^2 = -\frac{1}{2} \left. \frac{\partial^2 V_{CW}}{\partial h^2} \right|_{(v,0,0)} + \frac{3}{2v} \left. \frac{\partial V_{CW}}{\partial h} \right|_{(v,0,0)}.$$

## Benchmark models

BM	$\lambda_{hs}$	$\lambda_{hp}$	$\lambda_s$	$\lambda_p$	$\lambda_{sp}$	$T_{\text{nuc}}$	$T_{\text{perc}}$	$\alpha \cdot 10^3$	$\beta/H_c$
$Q_1$	0.88	0.78	0.70	0.60	0.70	63.58	58.30	119	254
$Q_2$	0.85	0.75	0.70	0.60	0.70	81.51	78.22	37.4	602
$Q_3$	0.80	0.70	0.70	0.60	0.70	99.60	97.64	15.2	1470
$Q_4$	0.75	0.65	0.70	0.60	0.70	112.25	111.05	8.61	2920
$Q_5$	0.70	0.65	0.70	0.60	0.70	122.06	121.36	5.48	5760
$Q_6$	0.65	0.60	0.70	0.60	0.70	130.07	129.72	3.52	12700
$Q_7$	0.60	0.50	0.55	0.45	0.60	133.45	133.13	2.88	15700
$Q_8$	0.60	0.50	0.60	0.50	0.70	134.84	134.61	2.56	19400
$Q_9$	0.60	0.50	0.65	0.55	0.70	135.85	135.66	2.26	27800
$Q_{10}$	0.60	0.40	0.70	0.60	0.70	136.74	136.61	1.95	38100

**Table:** Dimensionless parameters of selected benchmark models (BM) with the first order electroweak phase transition as well as nucleation and percolation temperatures  $T_{\text{nuc}}$ ,  $T_{\text{perc}}$  (in GeV), strength  $\alpha$  and ratio  $\beta/H_c$ .



## Bounce solution

$$\frac{d^2\varphi_i}{d\rho^2} + \frac{2}{\rho} \frac{d\varphi_i}{d\rho} = \frac{\partial V_{\text{eff}}}{\partial \varphi_i},$$

boundary conditions:

1)  $\frac{d\varphi_i(\rho)}{d\rho} = 0$  at  $\rho = 0$ ; 2)  $\varphi_i(\rho) \rightarrow \varphi_i^{\text{false}}$  at  $\rho \rightarrow \infty$

Euclidean action calculated on bounce solution

$$S_3 = \int_0^\infty 4\pi\rho^2 d\rho (V_{\text{eff}}(T, h, s, \rho) + \frac{1}{2} \left(\frac{dh}{d\rho}\right)^2 + \frac{1}{2} \left(\frac{ds}{d\rho}\right)^2 + \frac{1}{2} \left(\frac{d\rho}{d\rho}\right)^2)$$