Evaluation of the sensitivity of DarkSide-50 experiment to two neutrino double K-capture on ³⁶Ar

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Introduction

2e-capture

- $(2\text{EC}2\nu)$ is found for ¹²⁴Xe in XENON1T[1] $T_{1/2}^{2\text{EC}2\nu} = (1.8 \pm 0.5 \pm 0.1) \times 10^{22}$ yr.;
- limits determined for 85 Kr in the Baksan Laboratory[2] $T_{1/2}^{2EC2\nu} > 1.9 \times 10^{22}$ yr., CL = 90%;
- $(2EC0\nu)$ not found



2β -decay

- (2β2ν) experimentally validated for more than 10 nuclei
- $(2\beta 0\nu)$ not found



Registration of $2\beta 0\nu$ - or $2EC0\nu$ -decay will confirm the Majorana nature of neutrino.

 $2\text{EC}2\nu$ is possible for ³⁶Ar (argon is the active medium in DarkSide-50 experiment).

Purpose and tasks of the work

Purpose

Evaluation of the sensitivity of the DarkSide-50 experiment to two-neutrino double electron capture on the $^{36}\rm{Ar}.$

Tasks

- development of a software code for simulating the energy spectrum of double electron capture (2EC2ν) on ³⁶Ar;
- application of the detector response of the DarkSide-50 detector to the simulated spectrum;
- statistical data analysis using a model spectrum;
- derivation of a lower limit for the half-life of ³⁶Ar.

Novelty

Studies of $2\text{EC}2\nu$ have not previously been carried out on argon isotopes.

DarkSide-50 experiment

DarkSide-50 – experiment to search for dark matter particles; two-phase time-projection chamber filled with liquid ultra-pure argon.





Detector scheme

TPC scheme

The principle of particle detection by the DarkSide-50 detector



- The particle causes a scintillation flash (S1) and ionization of the medium;
- Electroluminescence occurs in a gaseous medium (S2).
- The time interval between S1 and S2 allows to determine the Z coordinate;
- According to S2, you can restore the coordinates (X,Y);
- The ratio of amplitudes S1 and S2 is used to discriminate events from an electron and a recoil nucleus.

Underground argon — the active medium in DarkSide-50

• 40 Ar was form from decay of 40 K:

$$e^- + {}^{40}K \rightarrow {}^{40}Ar +
u_e + \gamma$$

• ³⁹Ar was formed in the atmosphere by cosmic rays:

$$n + {}^{40}\text{Ar} \rightarrow {}^{39}\text{Ar} + 2n + \gamma;$$

The use of argon from underground deposits will reduce the contribution of this component;

- ³⁹Ar is the background source for the experiment and limits the sensitivity at low energies;
- The vast majority of primordial argon consists of isotopes ³⁶Ar and ³⁸Ar, because potassium is distributed in space about 50,000 times less than argon (on Earth, potassium prevails over argon by 660 times).
- ³⁶Ar was formed during nucleosynthesis in massive stars;
- the isotopic abundance of ³⁶Ar in atmosphere greatly exceeds the abundance ³⁶Ar underground.

Double electron capture

• Two protons in the nucleus simultaneously capture two electrons from the K-shell, and two neutrinos are emitted.

(Transitions from the L- and M-shells can also be considered as a second-order effect.)

- The problem of detecting double electron capture is reduced to detecting the characteristic fluorescence quanta that appear when two vacancies are filled.
- $E_{2\text{EC}2\nu} \approx 4.8 \text{ keV}.$

$$^{36}\mathrm{Ar} + 2e^- \longrightarrow ~^{36}\mathrm{S} + 2
u_e$$



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Modeling the spectral component of 2K-capture

The simplified model of response is used here, an exact model is being developed.



Estimated half-life of ³⁶Ar

- a($^{36}\text{Ar})\approx$ (0.0025 \div 0.334)% ^{36}Ar abundance limits in underground argon[3][4]
- $E_{2\nu 2EC} \approx 4.8 \ \text{keV} \sim 60 e^-$ of ionization,
- The background at this energy is $\sim 2100 \text{ events}/1e^-$ [5].
- Probability of two-photon fluorescence $\epsilon = 50\%$.

$$T_{1/2} = \ln(2) \frac{N(^{36}Ar)}{\sqrt{N_{\text{bg}}}} M_{\text{LAr}} \cdot T \cdot \epsilon \sim (10^{19} \div 10^{21}) \text{ years}$$
(1)

Data analysis Best background fit



Best fit of the experimental data by the background model (brown line) and partial contributions of the background components to this model in range [30,166] N_e .

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- $\Delta y_i = data_i model_i$ • $\chi^2 = \sum_{i=1}^n (\frac{\Delta y_i}{\sigma_i})^2$; *i*-bin number • $\Delta \chi^2_{signal} = \chi^2_{signal} - \chi^2_0$; • $L = e^{-\sum_i \frac{\Delta y_i^2}{2\sigma_i^2}} = e^{-\frac{1}{2}\chi^2}$;
- Top: Change in $\Delta \chi^2$ with increasing signal amplitude
- Bottom: Dependence of the maximum likelihood function on $\Delta\chi^2$

CL=90%;



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Fitting experimental data with a background and signal model for CL = 90%. Fitting was performed in the range $[30,166]N_e$.

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DS-50 sensitivity evaluation

Lifetime estimation of ³⁶Ar for 2K-capture

- $M(^{36}Ar) = 39.948 \text{ g/mol}$
- 36 Ar isotope abundance in UAr: $\eta = (0.0025 \div 0.334)\%$
- probability of two-photon fluorescence $\epsilon \approx 50\%$

$$T_{1/2} = \ln(2) rac{\eta N_A M T \epsilon}{M(^{36}Ar) N_{2
u 2EC}} > (0.6 imes 10^{19} \div 0.7 imes 10^{21})$$
 years, CL = 90%

Results

Preliminary result for ³⁶Ar in DarkSide-50

$$T_{1/2} > (0.6 imes 10^{19} \div 0.7 imes 10^{21})$$
 years, CL $= 90\%$

Expected result for ³⁶Ar in DarkSide-20k

 $M^{\text{DS-20k}} = 400 M^{\text{DS-50}}; N_{bg}^{\text{DS-20k}} \sim 0.01 N_{bg}^{\text{DS-50}}$

$$T_{1/2}^{ ext{DS-20k}} \sim 200 \, T_{1/2}^{ ext{DS-50}}$$
 years

Result in XENON1T for ¹²⁴Xe

$${\cal T}_{1/2} = (1.8\pm0.5\pm0.1) imes10^{22}$$
 years.

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DS-50 sensitivity evaluation

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Conclusion

- In the future, it is planned to carry out more thorough calculations of the 2EC2*v* probabilities on ³⁶Ar and a more accurate calculation of the energy release in this process.
 It is also possible to measure the isotopic abundance of argon in the DS-50 using mass spectroscopy.
- It would be a good option to check the calculation on a target enriched in ³⁶Ar.
- The analysis performed in this work is planned to be carried out during the experiment on the next generation DarkSide detector.

Thank you for attention!

Sourses I

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