Total OMC rate analysis in ⁷⁶Se

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Steps:

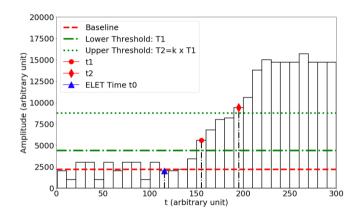
- Muon- & gamma-event timestamps calculation
- Main Trigger (OMC) definition and filling correlated spectra from Ge-event after Main Trigger
- Analysis of correlated spectra evolution for the list of lines (183, 199, 265, 280 keV) for each detector
- Getting average τ_{total} from all of the lines and detectors
- Calculation of systematics uncertanties

Muon- & gamma-event timestamp's calculation

• MIDAS writes hardware Struck SIS3316 timestamps with 4 ns resolution for muon events. (hardware defined, unchangeable)

• For Ge-events we use a special ELET (*Extrapolated Leading Edge Threshold*) algorithm for time calculation: timestamp is a crossection of flat background before pulse and line which is going trough 2 threshold's points (software defined, changeable)

ELET



Idea : We tune the coefficient k and the lower threshold to implement such conditions

• $\Delta t = t_1 - t_0$

•
$$\Delta t = t_2 - t_1$$

And than we calculate t_0 like this:

$$t_0 = 2t_1 - t_2$$

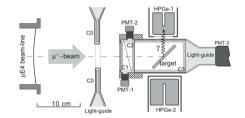
Muon- & gamma-event timestamp's calculation. Systematics

• Muon-event timestamp slightly depends on energy (constant hardware threshold).

• ELET (Ge-event timestamp) depends on energy, it works good for average pulses and bad for small or big ones. That is why we have a bad time binding for BEGe detectors #2 and #6 with MIDAS data.

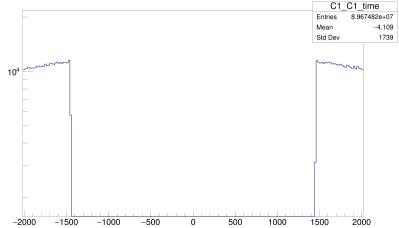
Main trigger variations

- C1&C2 at 100 ns.
- C1&C2 at 100 ns with hardware pileup protection
- C1&C2 at 100 ns + $\overline{C0}$ + $\overline{C3}$ at 1440 ns
- C1&C2 at 100 ns + C0 + C3 at 1440 ns with hardware pileup protection
- C1&C2 at 100 ns + $\overline{C0}$ at 1440 ns
- C1&C2 at 100 ns + $\overline{C0}$ at 1440 ns with hardware pileup protection



Why 1440 ns ?

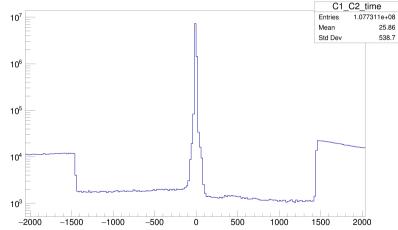
We have 1440 dead time for C# counters applying hardware trapezoid filter



C1_C1_time

Why 100 ns ?

C1_C2_time



Additional cuts (in question)

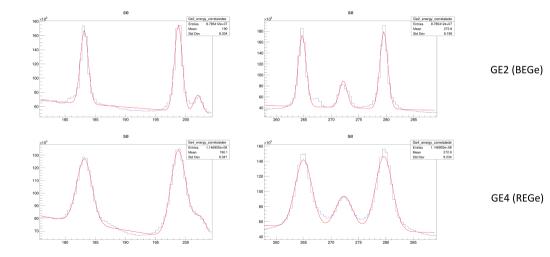
• No C0 and/or C3 event during special time window : 9120 ns = 12000 - 2*1440 ns (looks not necessary)

 No any C# event at 2*1440ns before main trigger (special cut for missed C# events)

Analysis of gamma-line evolution

- Fill correlated spectra from Ge-events (2d-histogram, E vs Time) after Main Trigger with 12 ns steps
- Fit selected gamma lines at integral correlated spectra by "gaus+pol1(3)" function (gaus + linear background). Fix sigma, energy and fit lines with steps every 12 ns after Main Trigger.
- Plot histograms for intensities time evolution
- Fit left tail of the time evolution data with "expo(0)+pol0(2)" model (exponent + flat background)

Fit of multiplets in ⁷⁶Se (2-3-4 etc)



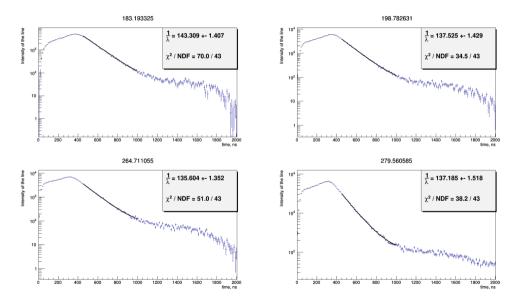
Fit of multiplets in ⁷⁶Se. Systematics

• Missed lines

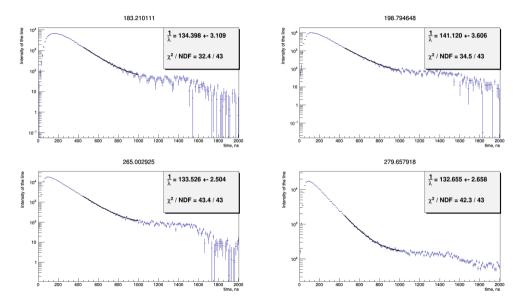
• Gain shift during the beam time (not yet checked)

• Very simple model (without left/right tails and/maybe steps)

⁷⁶Se lines at Ge2. Correlated spectra.

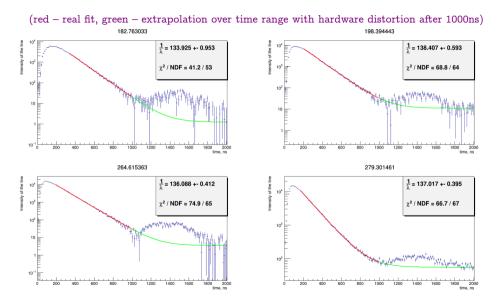


⁷⁶Se lines at Ge4. Correlated spectra.



14 / 22

Time evolution fit



Time evolution fit. Systematics

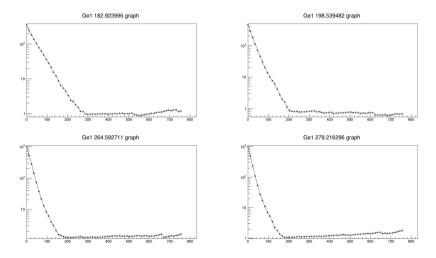
• Why expo+pol0(2) ?

• Maybe $\exp(0) + \exp(2) + \operatorname{pol}(4)$?

• Maybe something else ?

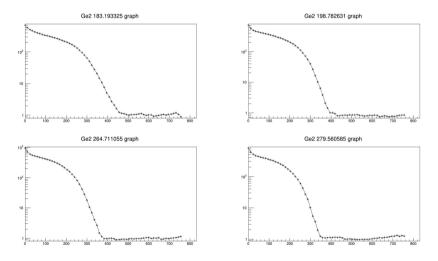
χ^2/NDF vs left border of the fit range

(Ge1, right border = 1000ns)



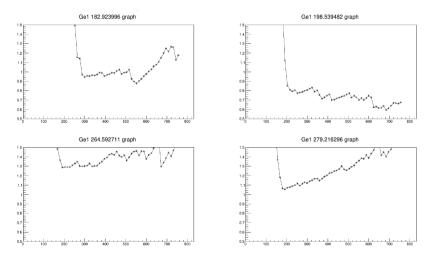
χ^2/NDF vs left border of the fit range

(Ge2, right border = 1000ns)



χ^2/NDF vs left border of the fit range

(Ge1, right border = 1000ns, zoom)



Getting average from all lines and detectors

$$rac{1}{\lambda_{total}} = au_{total} = rac{\sum_{lines} \sum_{detectors} au_{ij}}{N_{lines} N_{detectors}}$$

$$au_{total} = 137.030 \pm 0.046_{stat} ns$$

Calculation of systematics incertanties

In progress

THE END