

Elliptic flow fluctuations at NICA energies

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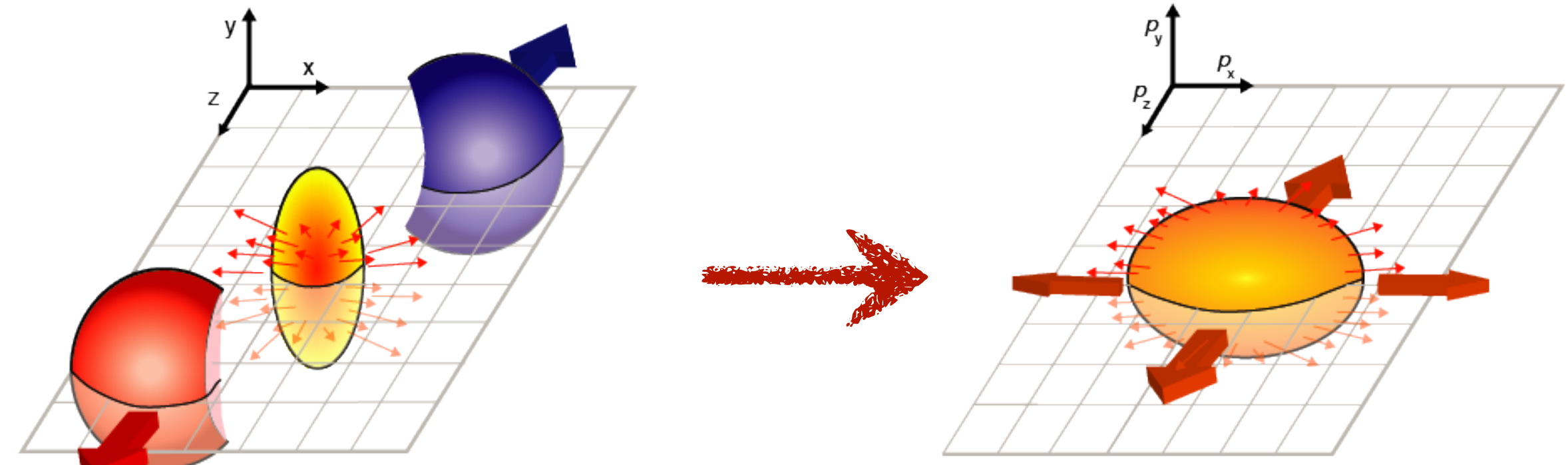
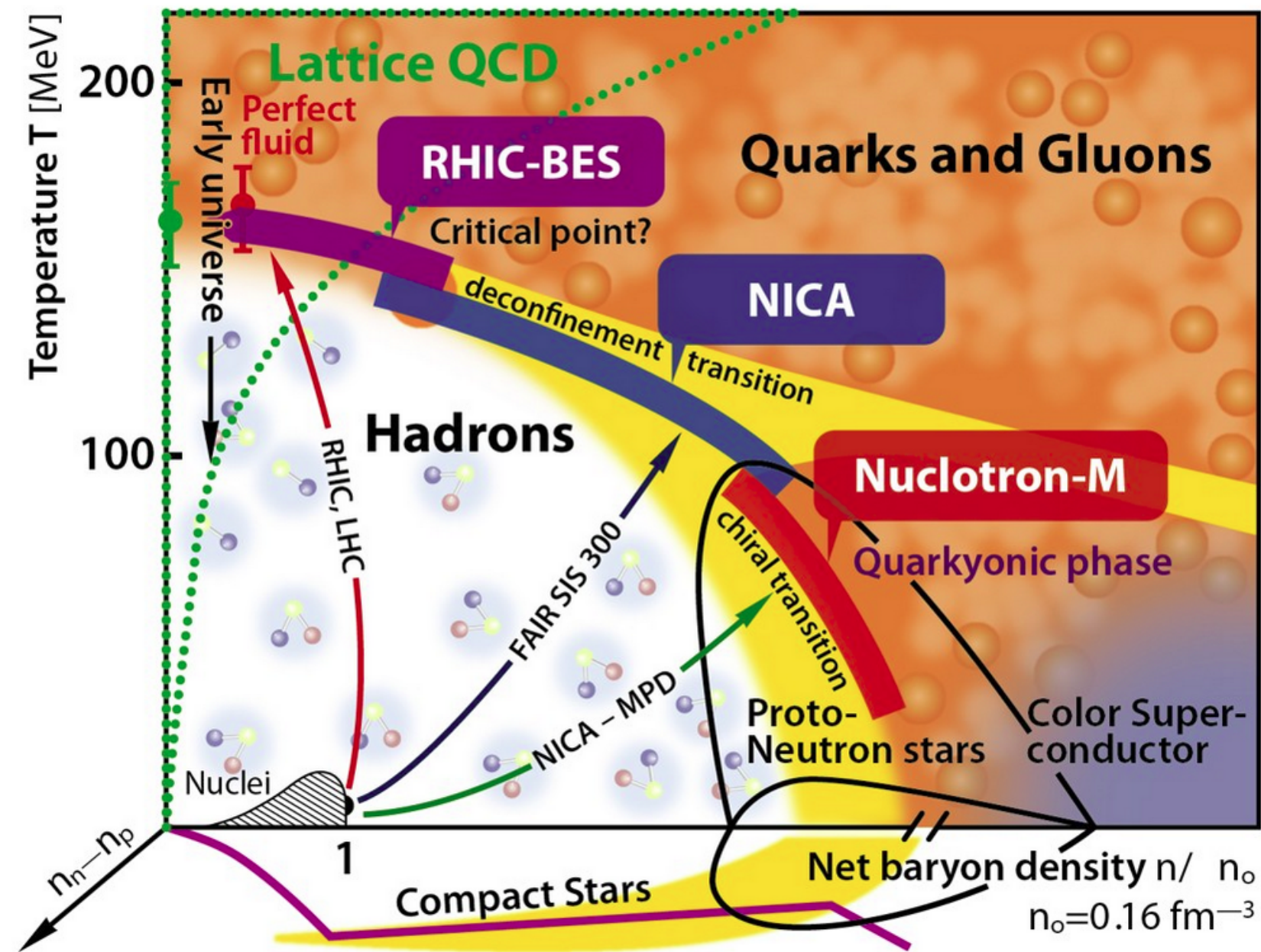
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Outline

- ◉ Q-cumulants method
- ◉ Sensitivity of different order of Q-cumulants to flow fluctuations and non-flow
- ◉ Results from models on elliptic flow fluctuations at NICA energy regime
- ◉ Conclusion

Anisotropic flow phenomenon



$$\varepsilon_n = \sqrt{\frac{\langle r^n \cos n\psi \rangle + \langle r^n \sin n\phi \rangle}{\langle r^n \rangle}}$$

$$\frac{dN}{d(\phi - \Psi_{RP})} \propto 1 + 2 \sum_{n=1}^{\infty} v_n \cos [n(\phi - \Psi_{RP})]$$

$$v_n = \langle \cos [n(\phi - \Psi_{RP})] \rangle$$

- LHC/top RHIC: cross-over phase transition to the sQGP
- Beam energy scan programs: RHIC/SPS/FAIR/**NICA**: searching for the critical end point & 1st order phase transition

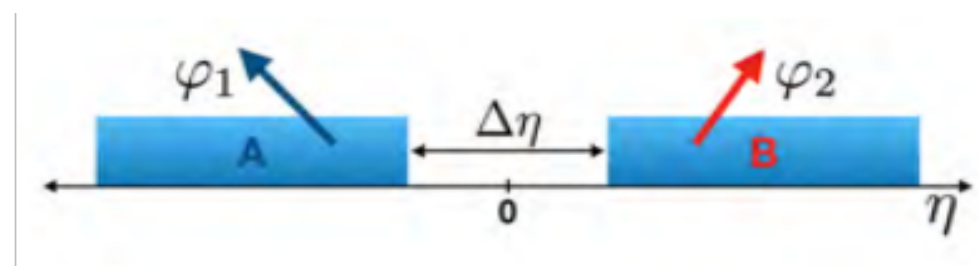
- Transfer of initial anisotropy ε_n in coordinate space to final anisotropy v_n in momentum space via the thermalized medium
- Anisotropic flow: sensitive probe to study the sQGP properties (η/s , ζ/s , EoS, ...)

Q-cumulants methods for flow measurements

- Ψ_{RP} cannot be measured directly \rightarrow Multi-particle azimuthal correlations method is used:

$$\langle\langle 2 \rangle\rangle = \left\langle \left\langle e^{in(\phi_1 - \phi_2)} \right\rangle \right\rangle = \left\langle \left\langle e^{in[(\phi_1 - \Psi_{RP}) - (\phi_2 - \Psi_{RP})]} \right\rangle \right\rangle = \left\langle \left\langle e^{in(\phi_1 - \Psi_{RP})} \right\rangle \left\langle e^{-in(\phi_2 - \Psi_{RP})} \right\rangle + \delta_2 \right\rangle = \langle v_n^2 + \delta_2 \rangle$$

- **2-particle Q-cumulant:** $\Delta\eta = 0.1$ is applied between 2 sub-events **A** and **B** to suppress 2-particle non-flow δ_2

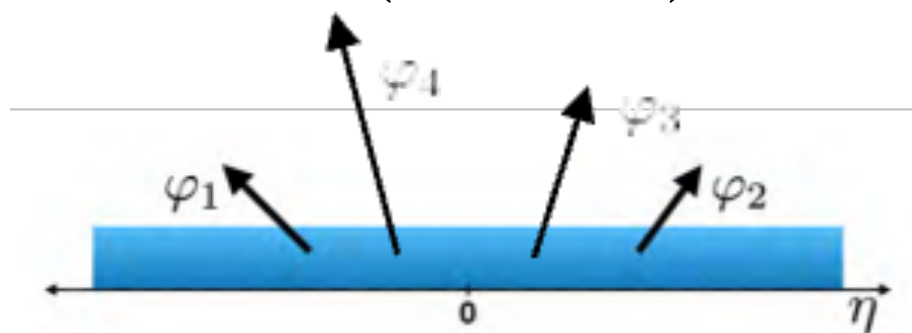


$$Q_n = \sum_{j=1}^M e^{in\phi_j} \quad \langle 2 \rangle_{a|b} = \frac{Q_n^a Q_n^{b*}}{M_a M_b} \quad v_n\{2\} = \sqrt{\langle\langle 2 \rangle\rangle_{a|b}}$$

- **4-, 6-particle Q-cumulants**

$$\langle 2 \rangle = \frac{|Q_n|^2 - M}{M(M-1)}$$

$$\langle 4 \rangle = \frac{|Q_n|^4 + |Q_{2n}|^2 - 2 \operatorname{Re}[Q_{2n} Q_n^* Q_n^*] - 4(M-2)|Q_n|^2 - 2M(M-3)}{M(M-1)(M-2)(M-3)}$$



$$v_n\{4\} = \sqrt[4]{2 \langle\langle 2 \rangle\rangle^2 - \langle\langle 4 \rangle\rangle} \quad v_n\{6\} = \sqrt[6]{1/4 \left(\langle\langle 6 \rangle\rangle - 9 \langle\langle 4 \rangle\rangle \langle\langle 2 \rangle\rangle + 12 \langle\langle 2 \rangle\rangle^3 \right)}$$

Formulae for $\langle 6 \rangle$ can be found in [A. Bilandzic et al, PRC 83 \(2011\), 044913](#)

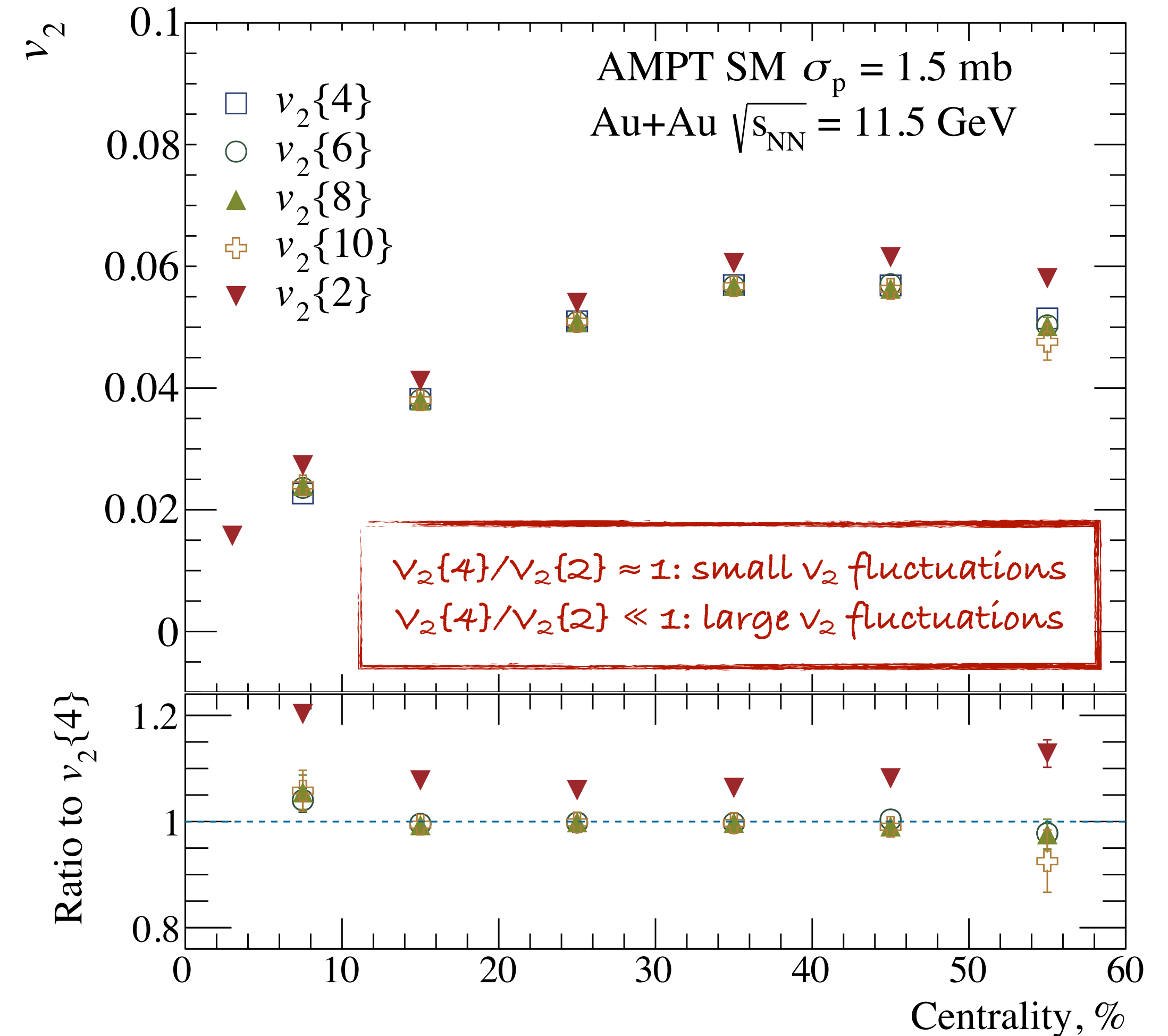
Sensitivity of $v_2\{2k\}$ to flow fluctuations and non-flow

- Non-flow contribution for k -particle cumulants:
 $\delta_k \sim 1/M^{k-1}$
- Elliptic flow fluctuations: $\sigma_{v_2}^2 = \langle v_2^2 \rangle - \langle v_2 \rangle^2$
- Assuming $\sigma_{v_2} \ll \langle v_2 \rangle$, fluctuations enhance $v_2\{2\}$ and suppress $v_2\{2k, k > 1\}$ compared to $\langle v_2 \rangle$

$$\triangleright v_2\{2\} \approx \langle v_2 \rangle + \frac{1}{2} \frac{\sigma_{v_2}^2}{\langle v_2 \rangle}$$

$$\triangleright v_2\{4\} \approx \langle v_2 \rangle - \frac{1}{2} \frac{\sigma_{v_2}^2}{\langle v_2 \rangle}$$

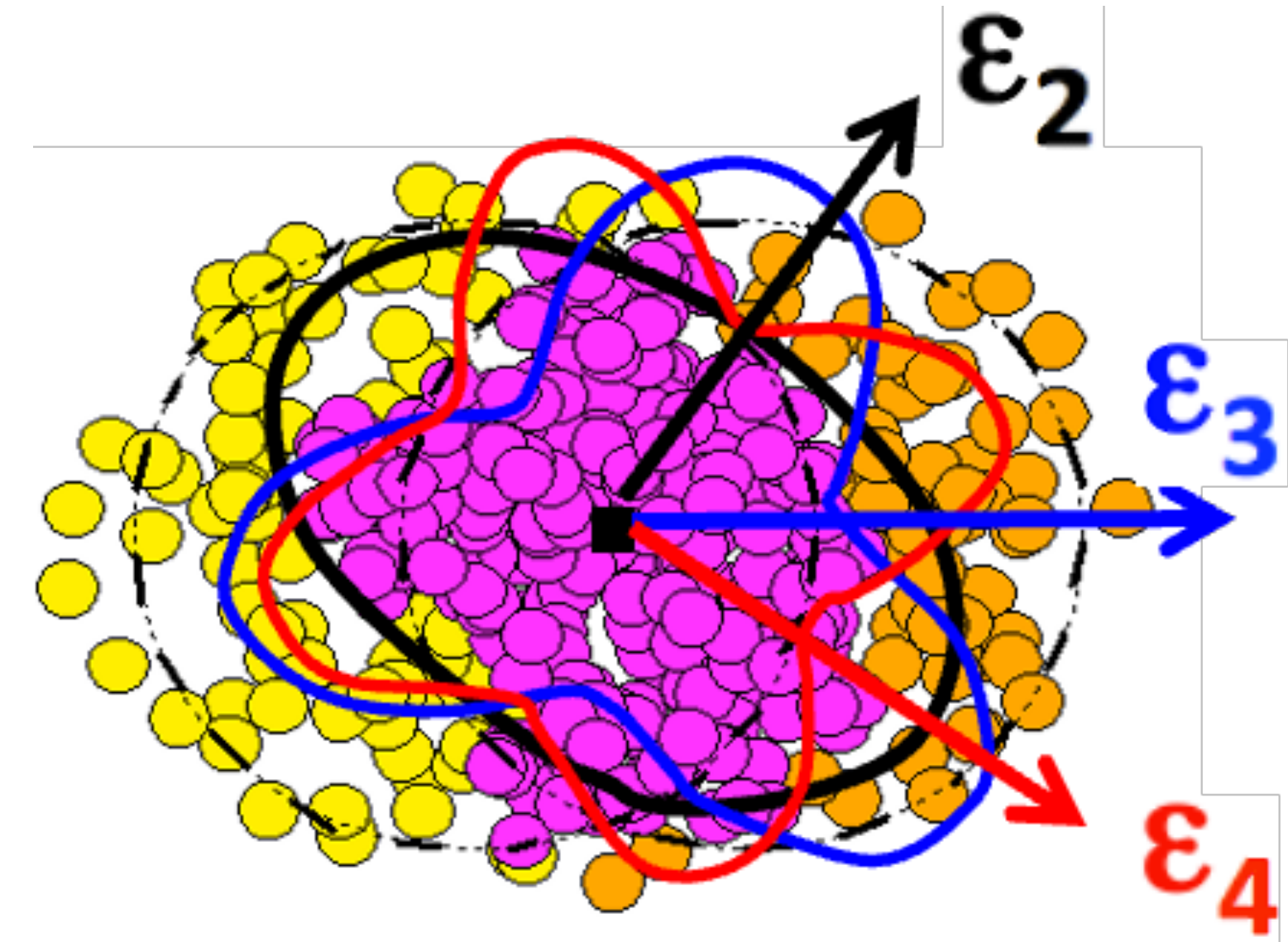
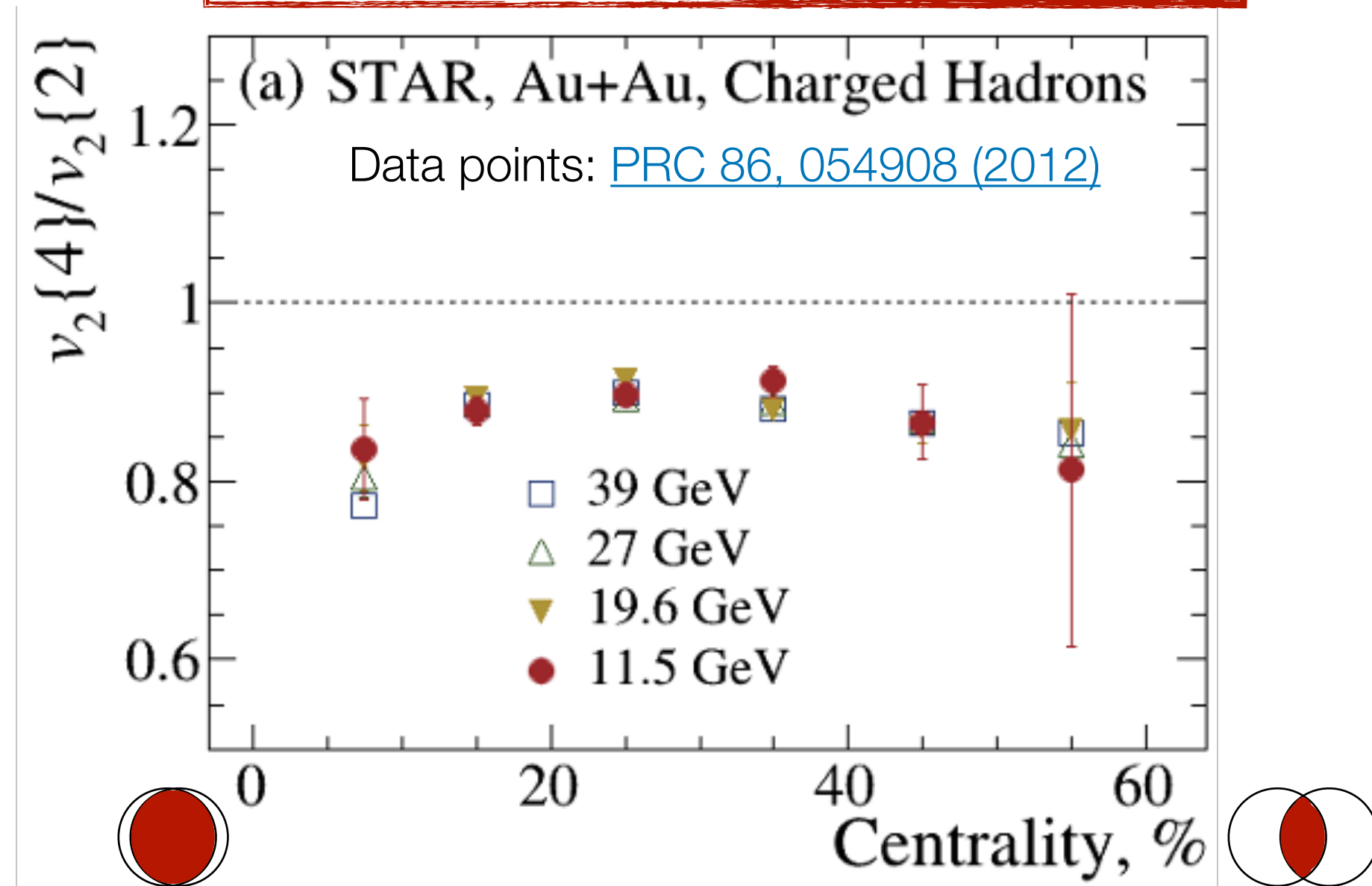
- Assuming a Gaussian form of fluctuations
 - $\triangleright v_2\{4\} \approx v_2\{6\} \approx v_2\{8\} \approx v_2\{10\}$



$$v_2\{2\} > v_2\{4\} \approx v_2\{6\} \approx v_2\{8\} \approx v_2\{10\}$$

Motivation of elliptic flow fluctuation study

$v_2\{4\}/v_2\{2\} \approx 1$: small v_2 fluctuations
 $v_2\{4\}/v_2\{2\} \ll 1$: large v_2 fluctuations



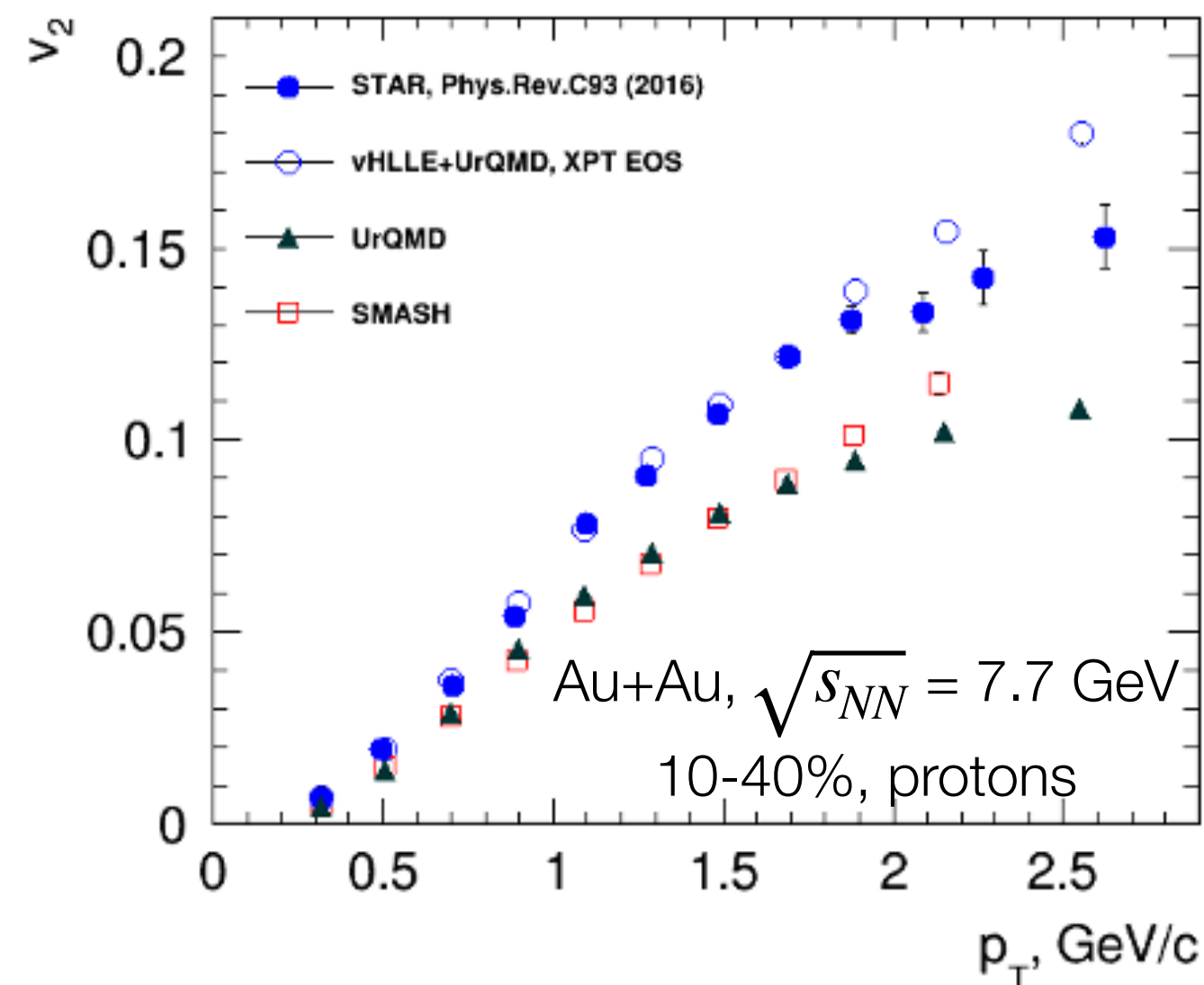
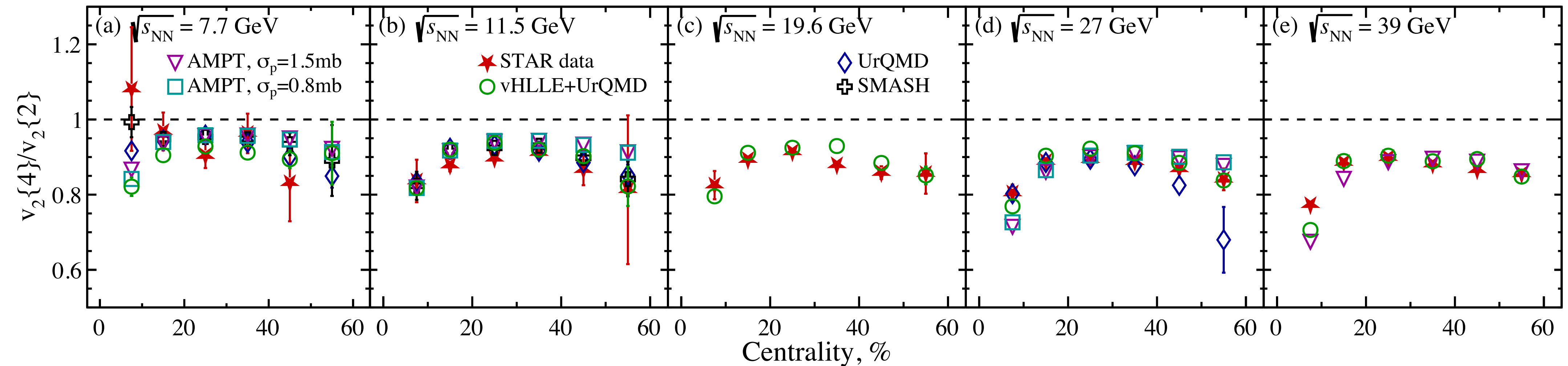
- v_2 fluctuations at $\sqrt{s_{NN}} = 11.5 - 39$ GeV observed in STAR:
 - ▶ Weak dependence on collision energy

- Indicate a dominated initial state driven fluctuations σ_{ϵ_2}
- Provide constraints for IS models and shear viscosity $\eta(T/s)$

How about v_2 fluctuations at NICA energies?

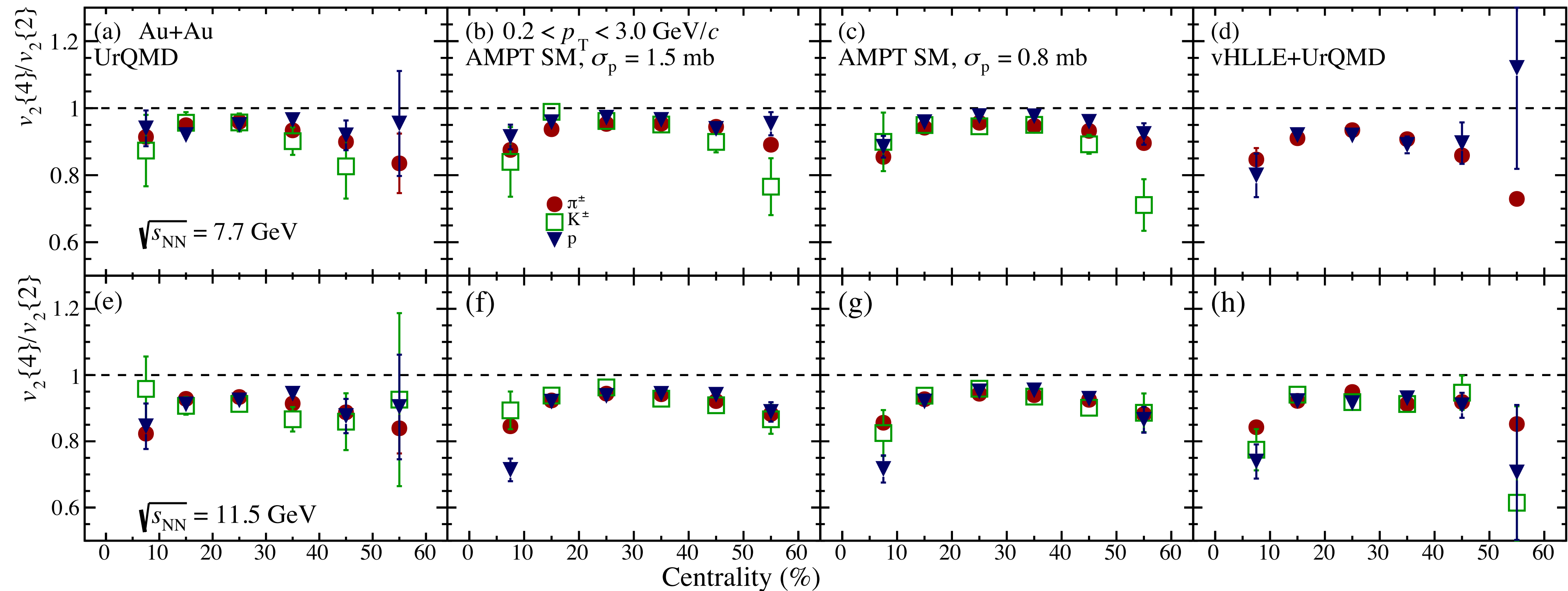
v_2 fluctuations at $\sqrt{s_{NN}} = 7.7 - 39$ GeV

Au+Au, charged hadrons, $|\eta| < 1$



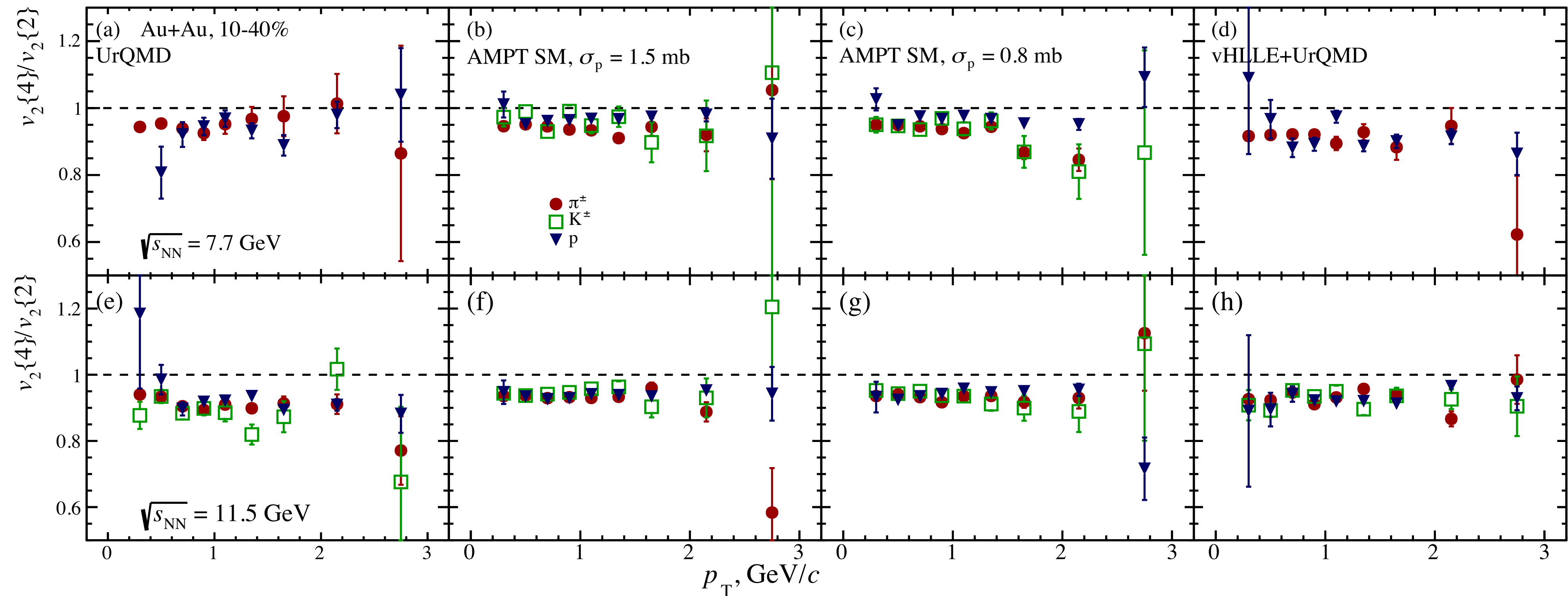
- At $\sqrt{s_{NN}} \geq 7.7$ GeV pure string/cascade models without QGP phase underestimate v_2
- v_2 fluctuations observed in STAR can be reproduced by model either with or without partonic phase description
 - ▶ v_2 fluctuations dominated by ε_2 fluctuations
 - ▶ $v_2\{4\}/v_2\{2\}$: direct probe for the initial state conditions

Relative v_2 fluctuations of identified hadrons



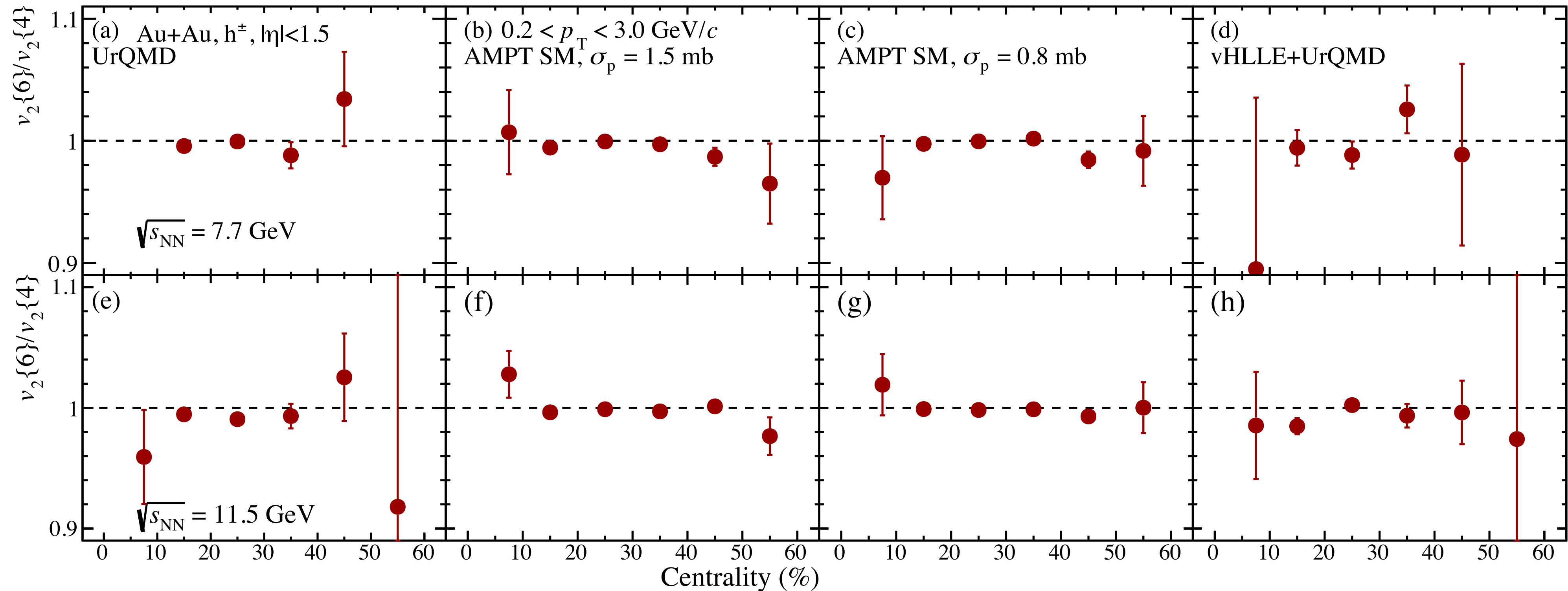
Weak dependence on particle species (pions, kaons, protons)

Relative v_2 fluctuations of identified hadrons



Weak dependence on p_T and particle species (pions, kaons, protons)

Skewness of $P(v_2)$

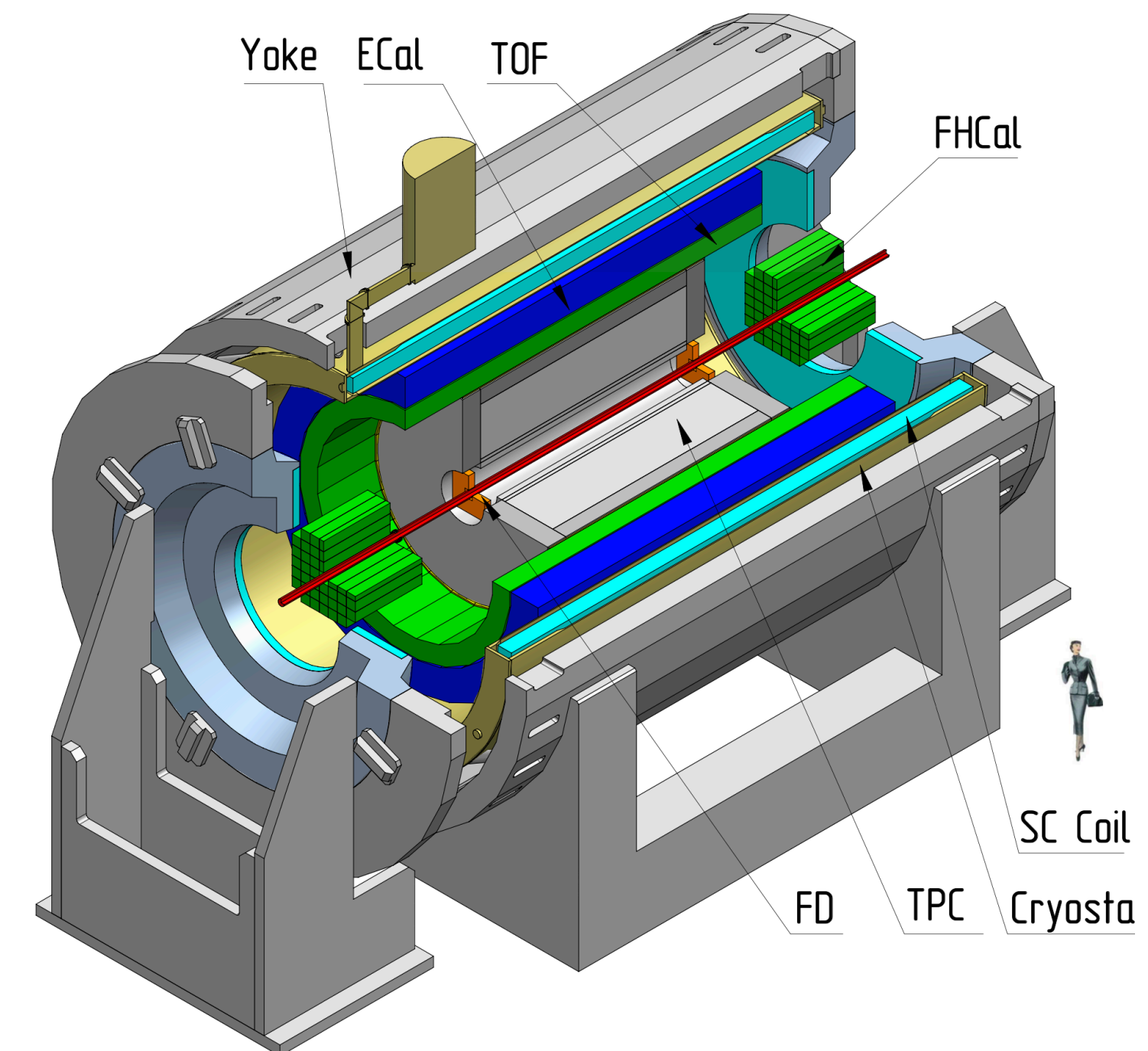


$v_2\{6\}/v_2\{4\} \approx 1 \rightarrow P(v_2)$ is likely to be Gaussian. Higher statistic is needed

MPD experiment at NICA

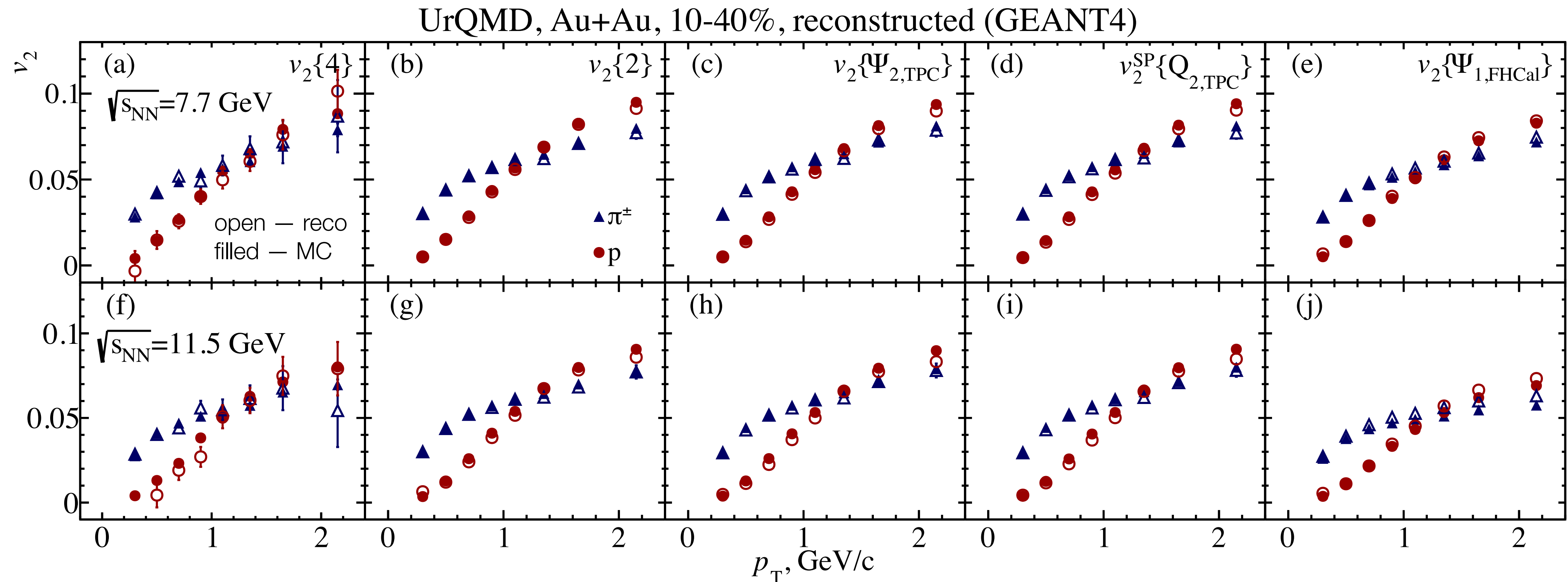


- Colliding system: Au-Au
- Colliding energy: $\sqrt{s_{NN}} = 7.7, 11.5$ GeV
- Centrality determination: b -based
- Event plane: $\Psi_{1,\text{FHCa1}}$ and $\Psi_{2,\text{TPC}}$
- Track selection:
 - ▶ $N_{\text{hits}}^{\text{TPC}} > 16$
 - ▶ $|DCA| < 2\sigma$
 - ▶ $0.2 < p_T < 3.0$ GeV/ c
 - ▶ $|\eta| < 1.5$
 - ▶ PID based on [MpdPid](#)



Multi-Purpose Detector at stage 1

Performance of proton, pion v_2 in MPD



Good agreement of v_2 from reconstructed and generated data for all particle species and methods

Conclusion

- Models reproduce the similar magnitude of v_2 fluctuations at $\sqrt{s_{NN}} = 7.7$, 11.5 GeV observed in STAR experiment at RHIC
 - ▶ v_2 fluctuations are mainly driven from ε_2 fluctuations
- High order cumulant ratio $v_2\{6\}/v_2\{4\} \approx 1$
 - ▶ $P(v_2)$ is likely to be Gaussian \rightarrow more statistic needed
- Outlook: increase the statistic for the $v_2(p_T, \mathbf{PID})$ fluctuations study

Thanks for your attention!

Back-up slides

Models & statistics

Au + Au min. bias

- ◉ **Without partonic phase**
 - ▶ UrQMD v3.4 (cascade)
 - $\sqrt{s_{NN}} = 7.7$ GeV: 88M
 - $\sqrt{s_{NN}} = 11.5$ GeV: 50M
 - ▶ SMASH v1.8
 - $\sqrt{s_{NN}} = 7.7 - 11.5$ GeV: 64M
- ◉ **With partonic phase**
 - ▶ vHLLE+UrQMD
 - $\sqrt{s_{NN}} = 7.7, 11.5$ GeV: 27M
 - ▶ AMPT SM $\sigma_p = 0.8$ mb
 - $\sqrt{s_{NN}} = 7.7$ GeV: 72M
 - $\sqrt{s_{NN}} = 11.5$ GeV: 35M
 - ▶ AMPT SM $\sigma_p = 1.5$ mb
 - $\sqrt{s_{NN}} = 7.7$ GeV: 42M
 - $\sqrt{s_{NN}} = 11.5$ GeV: 60M

Effect of centrality bin width

- 0-80% min bias:
 - 9 bins: (0,5,10,20,30,40,50,60,70,80)
 - 16 bins: 5%-bin width
- Similar results for $v_2\{4\}/v_2\{2\}$ cumulant ratio

