Research in Non-accelerator neutrino physics and astrophysics and proposals for the Seven-Year Plan

Evgeny Yakushev

56th meeting of the PAC for Nuclear Physics Dubna, 2023 The research theme: Non-accelerator neutrino physics and astrophysics

7 main projects: SuperNEMO, GERDA (Legend), Monument, EDELWEISS/Ricochet, vGeN, DANSS, BAIKAL-GVD

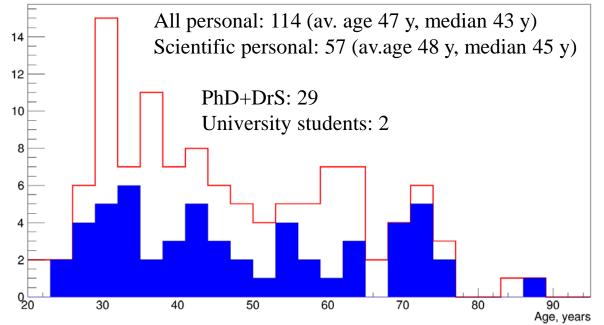
+ Activities

Partners: Azerbaijan, Bulgaria, Germany, Kazakhstan, Russia, Uzbekistan, Czech Republic, France, USA

Non-Accelerator Neutrino Physics and Astrophysics

Base: The Department of Nuclear Spectroscopy and Radiochemistry, Dzhelepov Laboratory of Nuclear Problems

- 50-years experience in high-precision nuclear spectroscopy using semiconductor, scintillator and other types of detectors;
- 30-years experience of rare processes studies in different underground environments;
- 15-years experience with neutrino experiments on nuclear reactors;
- Participation in the BAIKAL neutrino program for 20+ years



The department has the knowledge, personnel and capabilities to create world-class facilities, conduct measurements with them and obtain world-leading results

Scientific directions :

- Double beta decay, clarification of the neutrino nature Majorana or Dirac; ٠
- Nuclear matrix elements for 2β -decays; ۲
- Fundamental neutrino properties (magnetic moment, mixture with a sterile state, coherent scattering on nuclear);
- Monitoring of nuclear reactors with neutrino detectors; ٠
- Direct and indirect search for Dark Matter: ٠
- Investigation of galactic and extragalactic neutrino sources;
- Spectrometry of nuclei far from stability;
- Study of atomic processes accompanying radioactive decay;
- Development and implementation of methods for separation of macro-quantities of a substance from impurities, synthesis of materials from ultrapure precursors, design generators for production of nuclear medical radionuclides [PET diagnostics, radiopharmaceuticals, etc];
- Investigation of hyperfine interactions using the method of perturbed angular correlations on probe nuclei in solid and liquid samples.

Places: JINR, Underground laboratories, Nuclear reactors, Lake Baikal

Nuclear spectrometry methods to study various processes that are interesting at the present stage of the development of science, including and with focus to rare processes. 4 In addition to the scientific personal involved in the theme, the following resources are available to carry out the scientific projects:

➤ the laboratory for the production and repair of semiconductor detectors

➤ laboratory for creation and production of scintillation materials for detectors

 \succ radiochemical sector (creation of calibration radioactive sources, purification of materials designated for low-background measurements from their contamination by natural radioactivities, etc.)

mechanical workshops

➤ a group of computer support

➤ a group of mass separators

Facilities for works/test/repair semiconductor detectors





Infrastructure at JINR for detector R&D and tests





Low radioactive materials Infrastructure at JINR for sample preparation and preliminary tests

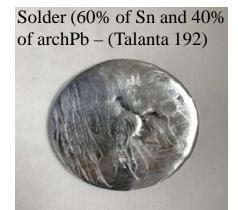














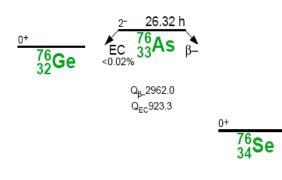
Radiochemistry clean rooms

Photos of the clean room built in JINR for Se and Nd purification On the right: Purifications of Se is in process



We will be glad to have PAC members visit to our department

Connection between neutrino physics and astrophysics (on example of neutrinoless Double Beta Decay)



In the Standard Model this process cannot occur without neutrinos

Observation at any level would imply:

- Lepton number L is not conserved
- Neutrinos have *Majorana masses* masses with a different origin than the quark and charged lepton masses
- Neutrinos are their own antiparticles

Observation of $0\nu\beta\beta$ would make more plausible:

- •The See-Saw model of the origin of neutrino mass
- Leptogenesis, an outgrowth of the See-Saw, which may be the origin of the baryon-antibaryon asymmetry of the Universe

General opinion of the particle physics community:

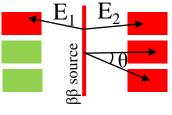
A non-zero signal for $0\nu\beta\beta$ would be a tremendously important discovery

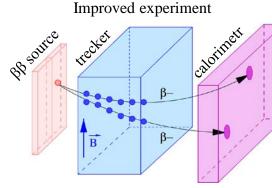
Method to study: nuclear spectrometry

Investigation 2β decay, search for 0v2β Two main approaches:

Track-calorimetric approach







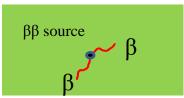
	Isotop	0ν2β	%
	⁴⁸ Ca	4273.7	0.187
 Full information about event (energy of each electrons, particle identification, track reconstruction, angular distribution); Sensitive to non standard modes of 0v2β decays; Possibility to measure several (almost any) isotopes in one detector; Unique information about 2v2β. 	⁷⁶ Ge	2039.1	7.8
	⁸² Se	2995.5	9.2
	¹⁰⁰ Mo	3035.0	9.6
	¹³⁰ Te	2530.3	34.5
	¹³⁶ Xe	2461.9	8.9
	¹⁵⁰ Nd	3367.3	5.6

Isotop

Enge

abudance

Calorimetric approach



- Detector = source;
- Efficiency near 100%;
- Compactness = the possibility of highly effective application of additional active and passive methods of background control, minimization of the amount of materials used near the detectors (background reduction);
 - Several detectors (or selection of a sensitive volume) = systematics, homogeneity;
 - Selection of events (PSD, two measurement channels, no multiplicity (if several detectors ...);
 - Achieving higher mass and low background (i.e. better result for $0v2\beta$ decay) with low cost.

supernemo

From NEMO3 to SuperNEMO through the Demonstrator phase (just started of data takin)

5							
201-	$\frac{t_{1/2}(10^2)}{10^2}$	1 yr)		ISOTOPE	TRANSITION	METHOD	DOCUMENT
	••• \	We do not	use the fol		ata for avera	ges, fits, limits, e	etc. • • •
	> 0.87			¹³⁴ Xe		EXO-200	¹ ALBERT
<u>ک</u> ^م کر	0.82	± 0.02	± 0.06	130 Te		CUORE-0	² ALDUINO
~	0.00690	± 0.00015	5 ± 0.00037	100 Mo		CUPID	³ ARMENG
	0.0274	±0.0004		¹¹⁶ Cd		NEMO-3	⁴ ARNOLD
collaboration	0.064	$+0.007 \\ -0.006$	$^{+0.012}_{-0.009}$	⁴⁸ Ca		NEMO-3	⁵ ARNOLD
	0.00934	±0.00022	2 + 0.00062 - 0.00060	150 _{Nd}		NEMO-3	⁶ ARNOLD
	1.926	± 0.094		76 Ge		GERDA	⁷ AGOSTIN
and a second second	0.00693	± 0.00004	1	100 _{Mo}		NEMO-3	⁸ ARNOLD
and the second second	2.165	± 0.016	±0.059	136 Xe		EXO-200	⁹ ALBERT
	9.2	$^{+5.5}_{-2.6}$	± 1.3	78 _{Kr}		BAKSAN	10 GAVRILYA
	2.38	± 0.02	± 0.14	136 Xe		KamLAND-Zen	11 GANDO
	0.7	± 0.09	± 0.11	¹³⁰ Te		NEMO-3	¹² ARNOLD
	0.0235		±0.0016	96Zr		NEMO-3	13 ARGYRIA
	0.69	$^{+0.10}_{-0.08}$	± 0.07		$\mathbf{0^+} \rightarrow \mathbf{0^+_1}$		14 BELLI
	0.57	$^{+0.13}_{-0.09}$	± 0.08		$\mathbf{0^+} \rightarrow \mathbf{0_1^+}$	NEMO-3	¹⁵ ARNOLD
	0.096	± 0.003	±0.010	⁸² Se		NEMO-3	¹⁶ ARNOLD
	0.029	+0.004		¹¹⁶ Cd		¹¹⁶ CdWO ₄ scir	17 DANEVIC
NEMO-3	L / r						D0 17
ββ event, as recorded by NEMO-3			50	00 kg× $T_{1/2}^{0 u}$	$> 10^{26}$ y	/T	
				$\langle m_{\nu} \rangle$	< 40 -	$110 \mathrm{meV}$	

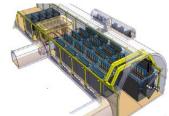
¹⁶ Cd	¹¹⁶ CdWO ₄ scint ¹⁷	DANEVICH 03
-		Demor
		17.5 kç
		T_{1}^{0}
		$\langle m_i$
l	CLOY	

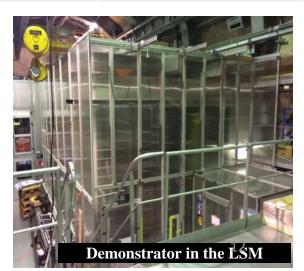
DOCUMENT ID			NEMO3	SuperNEMO
, etc. • • • ¹ ALBERT ² ALDUINO ³ ARMENGAUD ⁴ ARNOLD	17c 17 17 17	Isotope	Mo-100	Se-82 (Nd-150)
⁵ ARNOLD ⁶ ARNOLD ⁷ AGOSTINI ⁸ ARNOLD	16 16A 15A 15	Efficiency	18%	30%
 ⁹ ALBERT ¹⁰ GAVRILYAK ¹¹ GANDO ¹² ARNOLD ¹³ ARGYRIADES ¹⁴ BELLI 	14 13 12A 11 10 10	Energy resolution for 3 MeV elevtrons	8% (FWHM)	4% (FWHM)
¹⁵ ARNOLD ¹⁶ ARNOLD cint ^{1,7} DANEVICH	07 05A 03	Backgrounds		Reduced by factor 10

Demonstrator Module (2.5 year run) 17.5 kg×yr initial exposure :

 $T_{1/2}^{0\nu} > 5.9 \times 10^{24}~{\rm yr}$

 $\langle m_{\nu} \rangle < 0.20 - 0.55 \, \text{eV}$





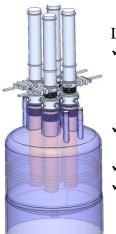
$0\nu\beta\beta$ search with HP⁷⁶Ge detectors immersed in liquid argon GERDA – LEGEND



GERDA phase II	aim	result	
Background	~ 10^{-3} cts/(keV kg yr)	5. $2^{+1.6}_{-1.3} \times 10^{-4}$ cts/(keV kg yr)	
Statistic accumulated	\geq 100 kg yr	103.7 kg yr	
Sensitivity	$T_{1/2}^{0\nu} \ge 10^{26} \text{ yr}$	$T_{1/2}^{0\nu} > 1.8 imes 10^{26} ext{ yr}$	

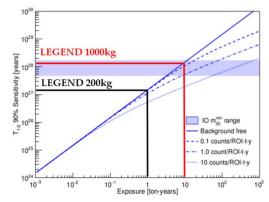


- ✓ First step to tone-scale 76 Ge exp.
- ✓ Improvements:
 - bigger mass of each of the detectors (in average)
 - improved LAr veto
 - less materials, new cables (less radioactive)
 - new electronics with lower noises
- ✓ Background improvement (estimate) : ~ x5 to GERDA
- ✓ 200 kg 76 Ge
- ✓ Target: $T_{1/2} \ge 10^{27} \text{ y}$

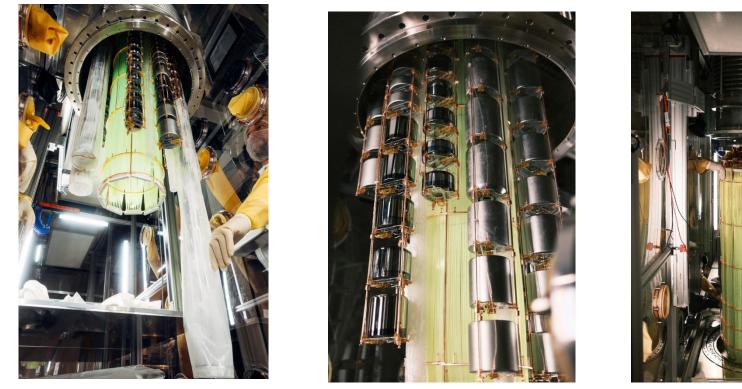


LEGEND-1000

- New infrastructure and design (some developments for Majorana)
- Improvements based on LEGEND-200
- ✓ 1000 kg ⁷⁶Ge
- ✓ Target: $T_{1/2} \ge 10^{28} \text{ y}$

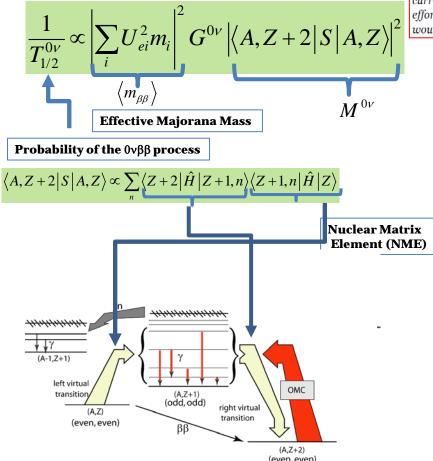


In the fall of 2022, the DLNP participated in the detectors installation. The active argon instrumentation system and all currently available Ge-76 detectors were re-installed. Now 101 germanium detectors are assembled in LEGEND and data collection has begun. The enriched Ge-76 isotope mass exceeded 100 kg for the first time. Now it is ~140 kg.

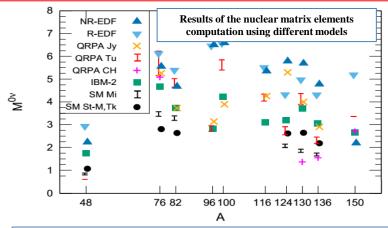


In 2023, when the remaining detectors are made, they will be added to the setup to reach the final Ge-76 mass of 200 kg. Then, after 5 years of data taking, the LEGEND-200 aims to reach the desired neutrinoless double beta decay sensitivity of 10^{27} years.

From half-life to neutrino mass: the key information is in nuclear matrix elementsMONUMENT project (Measurement of ordinary muon capture (OMC) for verification of nuclear matrix elements of2β-decays)APPEC-2019, Recommendation 6: The computation of nuclear matrix elements is challenging and



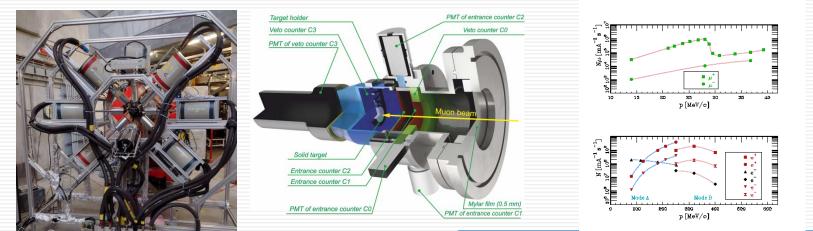
APPEC-2019, Recommendation 6: The computation of nuclear matrix elements is challenging and currently is affected by an uncertainty which is typically quantified in a factor of 2-3... An enhanced effort is required and a stronger interactions between the particle physics and nuclear community would be highly beneficial. Dedicated experiments may be required.



> In case of observed $0\nu\beta\beta$ decay the experimental results could help to improve NME calculations to define the effective Majorana neutrino mass

- > Right virtual transitions in DBD will be tested by Ordinary Muon Capture -> OMC and Ονββ operate in the q ≈ 100 MeV momentum-exchange region – high-lying states will be populated
- ≻ g_A suppression probing -- via capture rates calculations
 15

Status of the MONUMENT project (2020-2022)



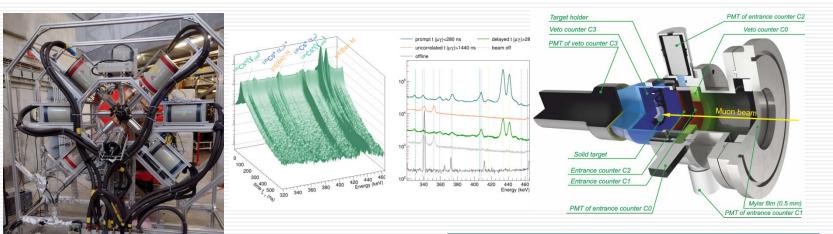
Purpose	OMC targets (enrichment)	Year/Status		
experimental input for DBD NME calculations	⁷⁶ Se (99.97%)	2021 / analysis and publication		
experimental input for DBD NME calculations	¹³⁶ Ba (95.27%)	2021 / analysis and publication		
experimental input for astrophysics investigations with SN	¹⁰⁰ Mo (97.3%)	2022 / started data analysis		
Nuclear spectroscopy, total cap. rates, yields	^{nat} Mo	2022 / started data analysis		
experimental input for DBD NME calculations	⁹⁶ Mo (exp. 95.3%)	2023 / in production		
testing nuclear shell model (SM) calculations	⁵⁶ Fe (exp. 99.9%)	2023 / in preparation		

> **OMC is the sensitive tool to probe properties of DBD process**. It is based on mature experimental technique successfully developed during many years, which demonstrates satisfactory agreement between experimental and theoretical data;

> The unique information obtained at OMC provide a significant experimental contribution to the nuclear spectroscopy, which is very actual in the nowadays;

➢It is planned to continue the intensive multiyear research program (at PSI beamline).
¹⁶

Status of the MONUMENT project (2020-2022)

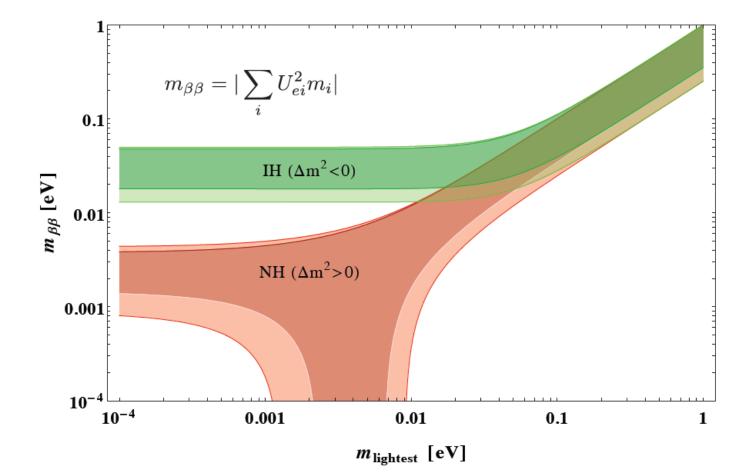


Purpose	OMC targets (enrichment)	Year/Status		
experimental input for DBD NME calculations	⁷⁶ Se (99.97%)	2021 / analysis and publication		
experimental input for DBD NME calculations	¹³⁶ Ba (95.27%)	2021 / analysis and publication		
experimental input for astrophysics investigations with SN	¹⁰⁰ Mo (97.3%)	2022 / started data analysis		
Nuclear spectroscopy, total cap. rates, yields	^{nat} Mo	2022 / started data analysis		
experimental input for DBD NME calculations	⁹⁶ Mo (exp. 95.3%)	2023 / in production		
testing nuclear shell model (SM) calculations	⁵⁶ Fe (exp. 99.9%)	2023 / in preparation		

>OMC is the sensitive tool to probe properties of DBD process. It is based on mature experimental technique successfully developed during many years, which demonstrates satisfactory agreement between experimental and theoretical data;

The unique information obtained at OMC provide a significant experimental contribution to the nuclear spectroscopy, which is very actual in the nowadays;

>It is planned to continue the intensive multiyear research program (at PSI beamline). ¹⁷



EDELWEISS/RICOCHET: Joint project for Direct Dark Matter search and precision study of CEvNS with new cryogenic detectors





EDELWEISS: Direct search for Dark Matter, HPGe detectors at ~20 mK, low background environment in deep underground laboratory (LSM).

For the last 25 years EDELWEISS has been the leading experiment in the direct detection of Dark Matter with HPGe bolometer detectors.

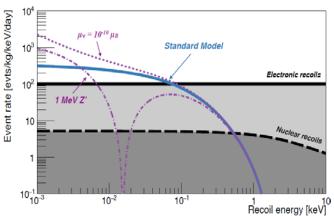
New detectors developed by EDELWEISS, have unique characteristics for detecting ultralowenergy nuclei recoils.

The same technology and detectors will be applied for precision measurements of CEvNS in the region of full coherency in the Ricochet experiment (reactor neutrinos). The main goal: precise (1% level) study of CEvNS, and target other New physics phenomena.

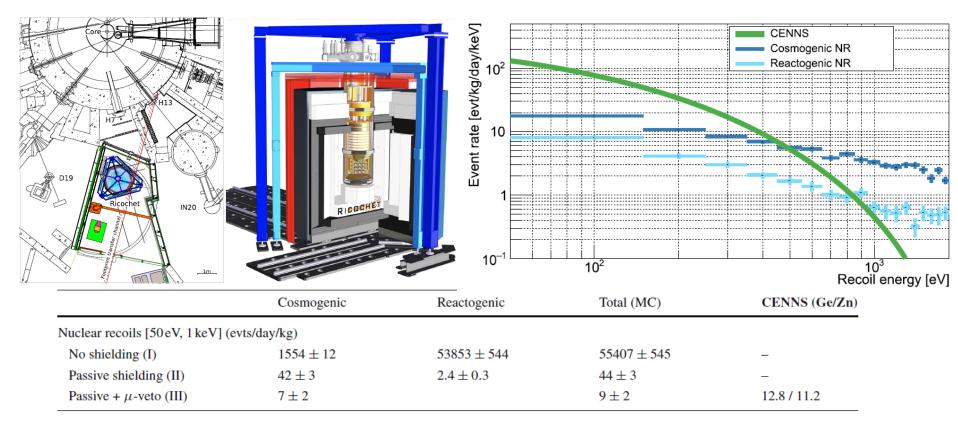


Thanks to the latest developments, the experiment remains competitive in the energy regions (low mass WIMPs, bosonic DM, etc, inaccessible to large Ar / Xe experiments.





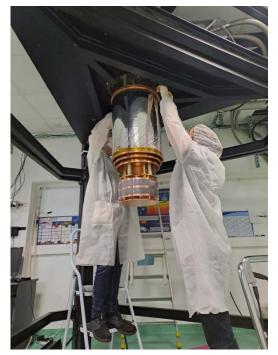
The experiment will deploy a kg-scale low-energy-threshold detector array combining Ge and Zn target crystals 8.8 m away from the 58MW research nuclear reactor core of the ILL

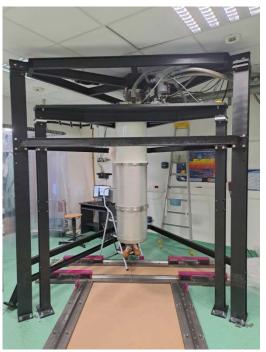


The Ricochet experiment should reach a statistical significance of 4.6 to 13.6 σ for the detection of CENNS after one reactor cycle

Dry, low-background, low-noise ³He-⁴He dilution cryostat produced in cooperation between JINR, IP2I and CryoConcept

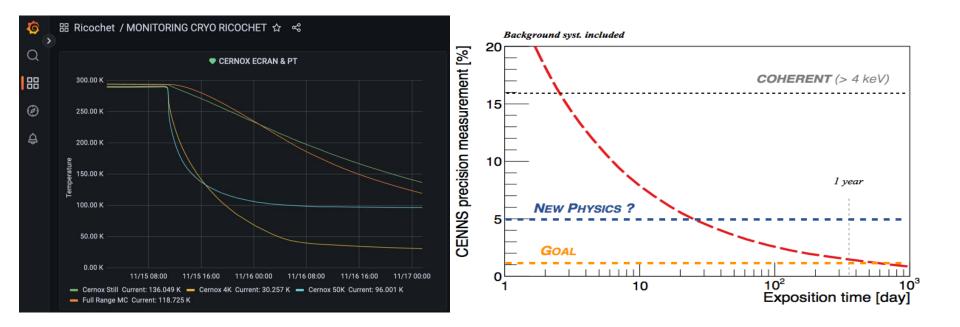






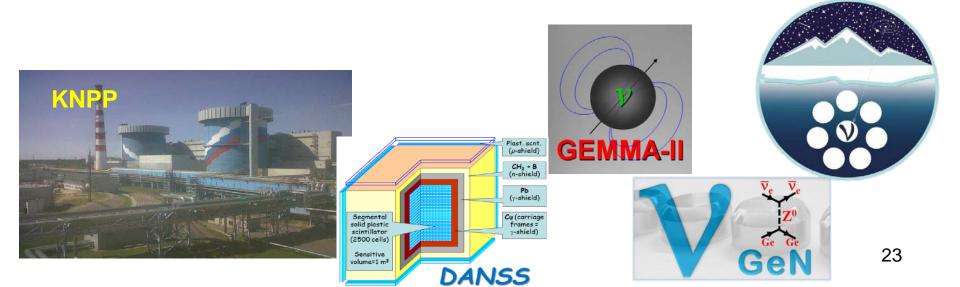
Our other involvement: muon veto, supplementary detectors, etc.

After some delay with the construction, the cryostat is now working and is ready to be installed at ILL !





Our participation in experiments with world leading sensitivities to ultra-rare processes provides the solid base for home neutrino experiments



DANSS

Segmental solid plastic

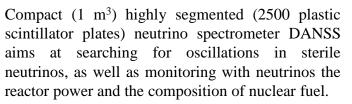
scintillator

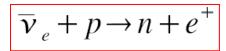
(2500 cells)

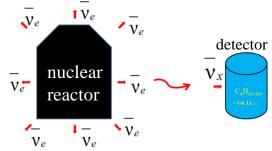
Sensitive

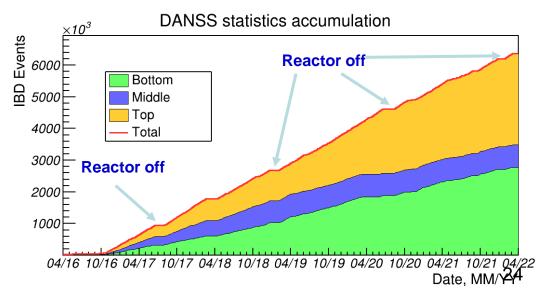
volume=1 m³

Investigations of reactor antineutrinos at the KNPP with an inverse beta decay detector









• DANSS is located under unit #4 KNPP (VVER-1000, $6 \times 10^{20} \ \overline{v_e}$ / sec);

Plast. scnt. (µ-shield)

CH2 + B

(n-shield)

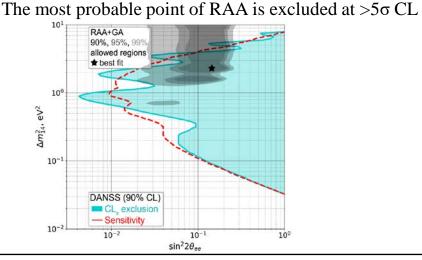
Pb (v-shield)

Cu (carriage

frames =

y-shield)

- Shield from cosmic ~ 50 m.w.e.;
- Lifting platform: distance to reactor from 10.9 to 12.9 m;
- ~5000 events/day;
- World leading results between all reactor based experiments for sterile neutrino search;
- No indications on a significant level for sterile neutrino were found

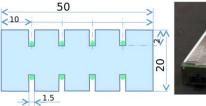


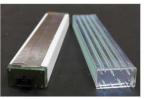
Upgrade to DANSS-2, the main goal is to reach the resolution $13\%/\sqrt{E}$ (now $34\%/\sqrt{E}$).

New geometry: **Strips:** 2x5x120 cm, 2-side 4SiPM readout **Structure:** 60 layers x 24 strips: 1.7 m³ **Setup** uses <u>the same shielding</u> <u>and moving platform.</u> **Gd** is in foils between layers.

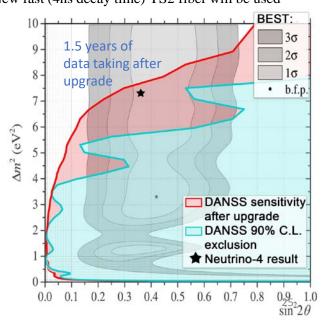


New scintillator strips





WLS fiber positions were optimized for better uniformity of response New fast (4ns decay time) YS2 fiber will be used

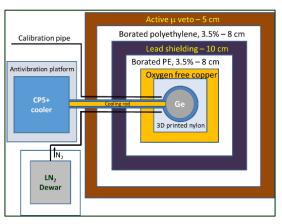


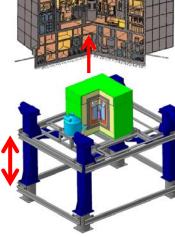
experiment at Kalinin nuclear power plant

- ✓ The vGeN project studies neutrino scattering at Kalinin Nuclear Power Plant (KNPP, Russia). Main interests: coherent elastic neutrino-nucleus scattering (CEvNS), the search for the magnetic moment of neutrino (MMN), search for New Physics beyond the SM, and many other applications, including reactor monitoring.
- ✓ The experimental setup is constructed under reactor unit #3 of KNPP at a distance of 11.1-12.3 m from the center of the 3.1 GW_{th} core under enormous antineutrino flux (4.4-3.6)×10¹³ v/cm²/s at 50 m w.e. overburden.
- ✓ The low-threshold HP Ge detector with a mass of 1.4 kg is used to detect desired signals. The installation of setup started in 2019. Since that time we performed the data taking, several improvements were performed since that time.



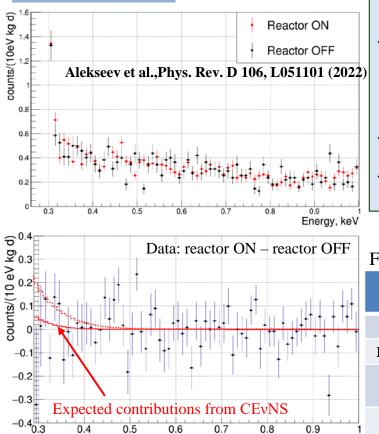
- ✓ The passive and active shielding covers detectors from all sides suppressing backgrounds.
- \checkmark Active antivibration platform suppresses microphonic noises.
- ✓ Nitrogen flushing decrease background from radon.
- Movable platform (completed in 2022) allows to change the distance to the reactor, thus changing the v flux.
- Special acquisition system was developed to suppress noise and achieve energy resolution of 101.6(5) eV (FWHM).





Ge





Energy, keV

First results on CEvNS by vGeN					
	Days	Counts per kgd			
Reactor on	94.50	2.32 ± 0.15			
Reactor off	47.09	$\textbf{2.34} \pm \textbf{0.21}$			
On-off	-	< 0.4 (90% CL)			
CEvNS (expect)	-	~ 0.058-0.66			

- ✓ No significant difference between regimes with reactor ON and OFF has been observed in first data taken in 2021. The new limit on the quenching parameter in germanium has been obtained (k<0.26).</p>
- ✓ In 2021-2023 we continued data taking, until 2023 we accumulated raw data: reactor ON 928 kgd, OFF: 175 kgd. Similar statistics with improved threshold and background level at low energy suggests that we already reached a better sensitivity for the magnetic moment of neutrino than GEMMA collaboration (set up the current world best limit for MMN).
- ✓ The new measurements at a closer distance from the reactor (11.1 m instead of 12.2 m) with higher flux started in September 2022.
- ✓ New results with more statistics and optimized measurement modes are expected soon. Search for evidence of other effects in data are ongoing.
 - ✓ The experiment mostly performed by JINR, more manpower needed.
 - ✓ In 2022 the vGeN collaboration were extended by groups from CTU (Prague, Czech Republic) and LPI RAS (Moscow, Russia).
 - ✓ The internal shielding and new better detectors are planned to be used to improve sensitivity of experiment.



North hemisphere biggest neutrino detector BAIKAL-GVD

Neutrino telescopes brings an important information complementary to the traditional optic and radio telescopes.

Main principle : determination of direction and energy of charged particles (appearing as result of neutrino interaction) with the help of Cherenkov radiation

Baikal project: From 1980 tests and R&D, started in 1993.

Now – building BAIKAL-GVD (**Gigaton Volume Detector**) has huge progress and will be continued

Physics:

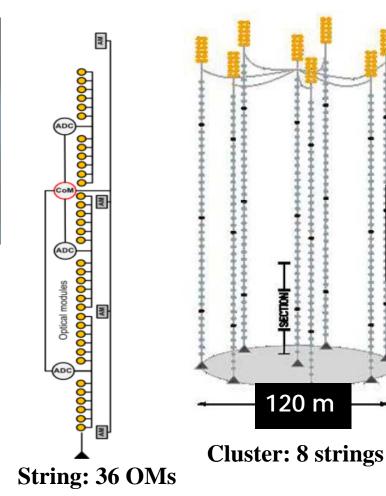
- Investigate Galactic and extragalactic neutrino "point sources";
- Diffuse neutrino flux energy spectrum, local and global anisotropy, flavor content;
- Indirect search for Dark Matter;
- Exotic particles monopoles, Q-balls, nuclearites, ...



DLNP created facilities to produce and tests optical modules (the main element of BAIKAL-GVD)

Our staff is also important force to install the modules, commissioning, running, data taking and final analysis





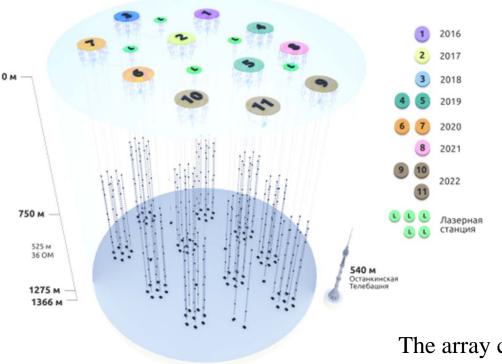
Ξ

525



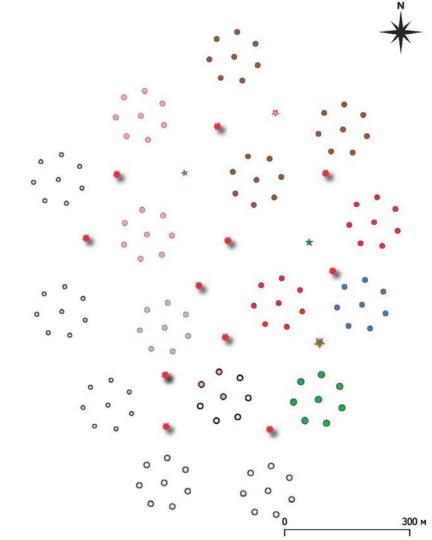
Baikal-GVD construction status

Status 2022: 10 clusters, 5 laser stations, experimental cluster prototype with new DAQ system



About 700 optical modules are assembled for deployment in 2023.
The collaboration is planning to install additional 2 new clusters in case a good external conditions (weather and ice).

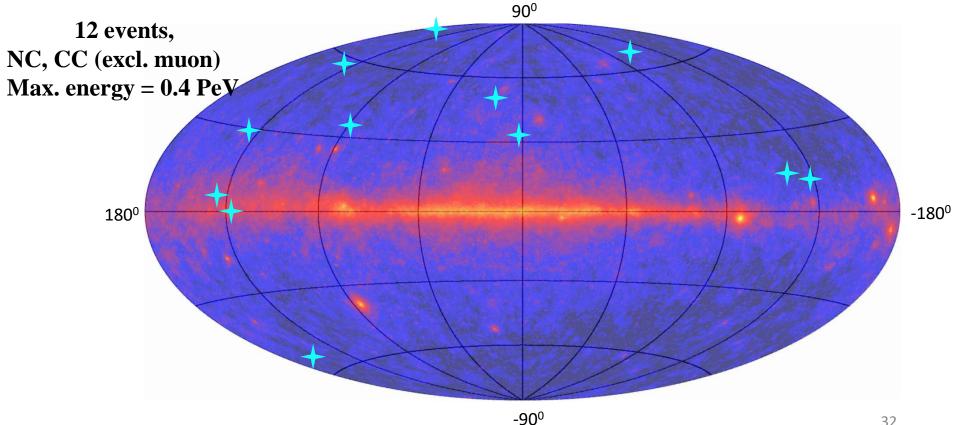
The array contains 2916 optical modules.



Deployment schedule

Year	Number of clusters	Number of OMs
2016	1	288
2017	2	576
2018	3	864
2019	5	1440
2022	10	2916
2023	12	3564
2024	14	4212
2025	16	4860

Some Baikal-GVD cascade events >100 TeV



To improve the efficiency and to reflect real involvement of personal and resources we propose to reorganize the structure of the theme.

If supported, the new structure will have 4 main projects:

➢ BAIKAL-GVD

> A short baseline reactor neutrino investigation (Ricochet, DANSS-2 and vGeN experiments)

> Nuclear spectrometry for search and investigation of rare phenomena (all double beta decay related experiments and activities, search for Dark Matter by nuclear spectrometry methods, etc)

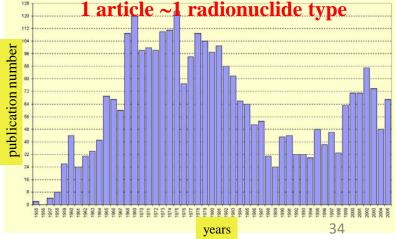
> Radiochemistry plus spectroscopy for astrophysics and nuclear medicine

Radiochemistry plus spectroscopy for astrophysics and nuclear medicine

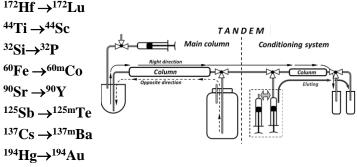
Huge accumulated experience since 1950: radiochemistry and spectroscopy / target irradiated on the DLNP Phasotron

- The radiochemical experience in the production of radionuclides of almost all elements for spectroscopy + the sources preparation
- In cooperation with foreign colleagues, production methods of the most important radiopharmaceuticals for nuclear medicine were developed: ¹⁷⁷Lu, ⁹⁰Y, ²¹¹At, etc.
- ★ In recent years, the development of the production methods of the relevant radiopharmaceuticals for nuclear medicine has also continued. It can be highlighted: production method of ⁶⁸Ga radiopharmaceuticals for PET diagnostics from the ⁶⁸Ge→ ⁶⁸Ga generator, the reverse-tandem scheme of the ⁴⁴Ti→ ⁴⁴Sc generator, the development of the production methods of ²²⁵Ac, etc.
- * Low-background samples for research in the field of neutrino physics and the search for Dark Matter particles has been actively developed. For example, in order to search of $0v2\beta$, ⁸²Se purification was carried out and kilogram quantities of the elemental highly dispersed form with a unique low content of radioactive impurities were produced.





- Study of sorption processes for various solution-sorbent systems as a chemical basis for methods of purification and analysis of low-background materials.
- Development of radionuclides production and analysis methods for various nuclear spectrometric and chemical tasks, including calibrations of low-background installations and various spectroscopic R&D. It is planned to create about 30 (50) generators with long-lived radionuclides using reverse-tandem schemes.
- Development of low-background samples production methods (^{82}Se , ^{96}Zr , shielding materials, solder, etc.) for astrophysical problems at a new ultra-low level of impurity content ($\mu Bq/kg$ for Th and U) through the use of countercurrent chromatography, low-boiling and other prepared or selected reagents, as well as selected and prepared reactor materials.
- Development of low-background samples analysis methods at an ultra-low sensitivity level (µBq/kg for Th and U) using ICP-MS and neutron activation analysis. To achieve this sensitivity level it is necessary to develop methods of the preliminary chemical concentration (separation) of impurities from materials.

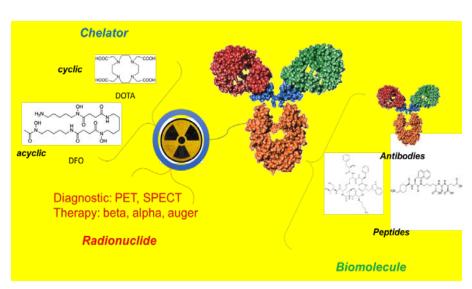


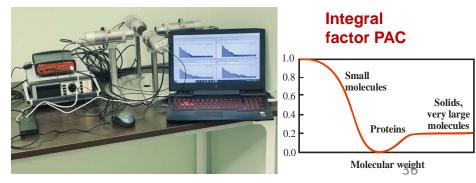
- $^{243}\text{Am} \rightarrow ^{239}\text{Np}$
- Etc.

.



- Study of sorption processes for various solution-sorbent systems as a chemical basis for methods of radiopreparations purification and radionuclide generators design for the radiopharmaceuticals production.
- Development of methods of the radionuclides production and isolation from targets irradiated with protons, neutrons and gamma-rays (bremsstrahlung) for the radiopharmaceuticals production.
- To expand the nuclear medical radionuclides production possibilities, the development of a wide range of radionuclide generators (${}^{44}Ti \rightarrow {}^{44}Sc$, ${}^{68}Ge \rightarrow {}^{68}Ga$, ${}^{90}Sr \rightarrow {}^{90}Y$, ${}^{238}U \rightarrow {}^{234}Th$, ${}^{237}Np \rightarrow {}^{233}Pa$, ${}^{229}Th \rightarrow {}^{225}Ac$, ${}^{227}Ac \rightarrow {}^{227}Th \rightarrow {}^{223}Ra$, ${}^{202}Pb \rightarrow$ ${}^{202}Tl$, ${}^{194}Hg \rightarrow {}^{194}Au$, ${}^{32}Si \rightarrow {}^{32}P$, etc.) based on the reverse-tandem methods will continue. The 1-2 generators design possibility of significant activity for external users will be considered.
- Development of methods for labeling radiopharmaceuticals with radionuclides.
- Application of nuclear spectrometric and radiochemical methods for the modern radiopharmaceuticals and their precursors study. The Perturbed Angular Correlation (PAC) method of Mössbauer spectroscopy will be used. It is also supposed to use electrophoresis methods in free electrolyte to determine the kinetic and thermodynamic constants of radiopharmaceuticals with radionuclides.





Summary of the major aims for new seven years plan:

• DANSS experiment: Checking the extended region of the phase space of oscillations into a sterile neutrino with reaching the signal region of the NEUTRINO-4 experiment; precision measurement of the spectrum of reactor antineutrinos; measuring the evolution of nuclear fuel.

• RICOCHET experiment: New physics with precision measurements of CEvNS at reactors.

• vGeN: Observation of coherent elastic neutrino-nuclear scattering with antineutrino from reactor, search for magnetic moment of neutrino up to the world leading levels.

• LEGEND: achievement the discovery sensitivity for neutrinoless double-beta decay at 10^{28} years.

• BAIKAL-GVD: Observation of ultra-high energy astrophysical neutrinos; identification of their sources and nature.

• Radiochemistry plus spectroscopy for astrophysics and nuclear medicine: Production and analysis methods of low background samples with purity level of μ Bq/kg for experiments of astrophysics and neutrino physics; production and analysis of novel radiopharmaceuticals; new types of radionuclide generators; investigation of new materials and mechanisms of local interactions of radioactive probe nuclei in substance.

➤ Comprehensive scientific program reflecting name of the theme "Non-accelerator neutrino physics and astrophysics";

Participation in world leading experiments;

Brings modern know-how and culture of experiment for "home" projects;

> Creation of infrastructure at JINR to ensure current and future projects at a modern level;

➤ In average 25 publications per year (plus many conference proceedings, preprints, etc).

We would like to ask the Programme Advisory Committee to support our researches conducted in frame of the theme and to support continuation of all individual projects

We propose to present the written projects for 2024-2028 yy during the next/summer 2023 PAC meeting

Only slides1-4, 10-31, 33, 34, 37, 38 will be shown during the PAC presentation

Additional slides

JINR BAIKAL-GVD optical module assembly line

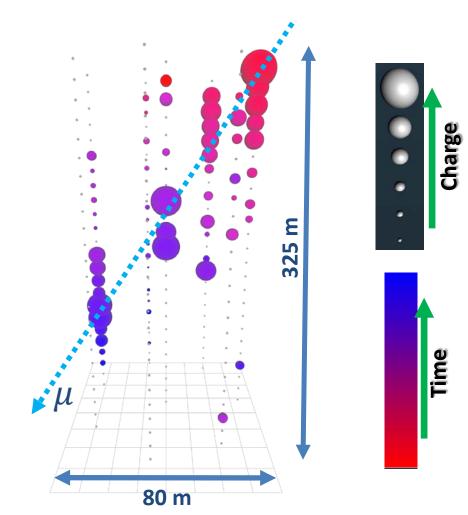




Assembly and tests 12 OM per day

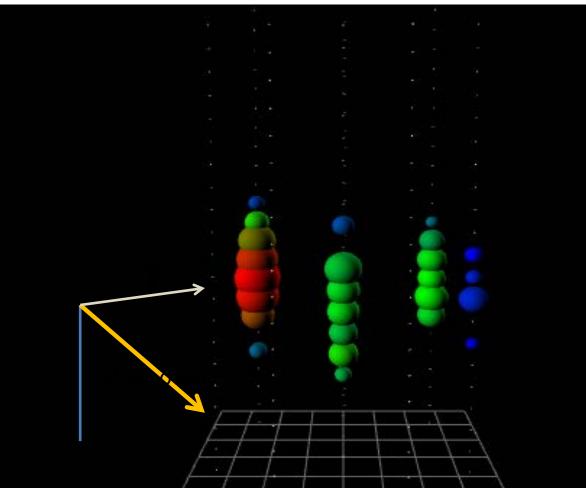


An event in the BAIKAL-GVD



43

$E = 158 \text{ TeV}, \ \theta = 59^{\circ}, \rho = 73 \text{ m}, \ z = -62 \text{ m}$

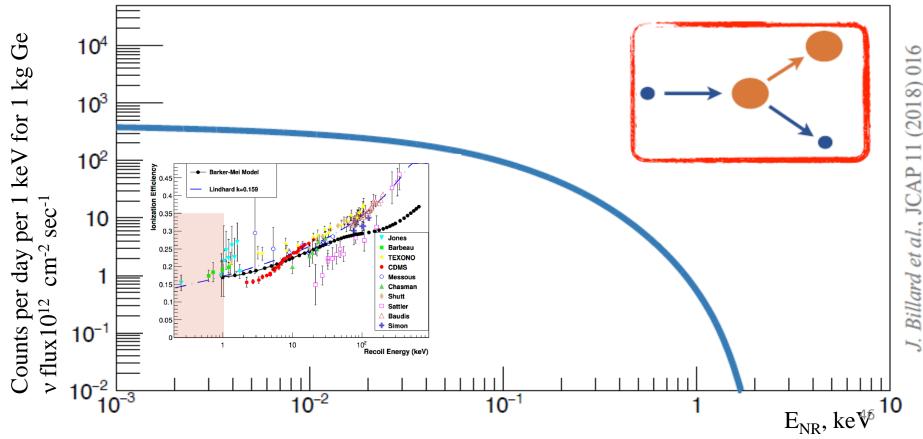




Coherent elastic neutrino-nucleus scattering (CEvNS)

The use of bolometers makes it possible to measure the energy of the nucleus directly (heat signal), in contrast to semiconductor detectors that measure ionization.

This it the way to the precision measurements.



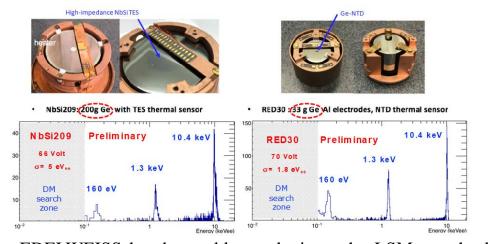


The current phase results

- Continuous data taking in the LSM underground laboratory from January 2019 July 2020;
- Almost continuous data taking in the test dry cryostat in Lyon (R&D);
- 11 different Ge detectors;
 - Rest of the cryostat used for joint physics run with CUPID-Mo 0v2β search
- Compare detector physics in 32 g, 200 g and 800 g detectors;
- Compare performance of NTD and NbSi-TES heat sensors;
- Obtaining near single-electron sensitivity on 33 and 200 g detectors: exploration of DM interactions with electrons and nuclei.
- Study of low-energy backgrounds in Ge detectors operated with large Neganov-Trofimov-Luke amplifications.

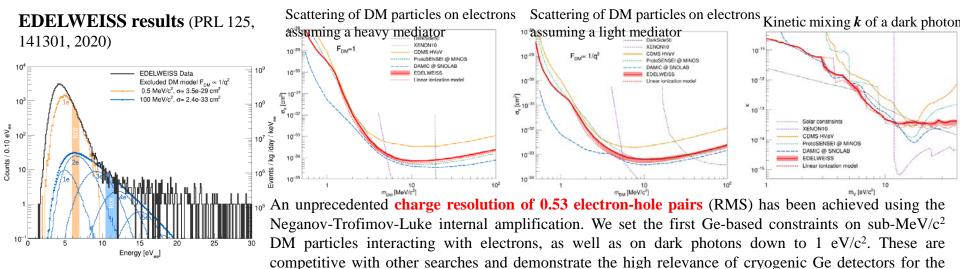






EDELWEISS has been able to obtain at the LSM run the lowest radioactive background levels below 10 keV in massive Ge detectors (~0.1 evt/kg/day/keV)

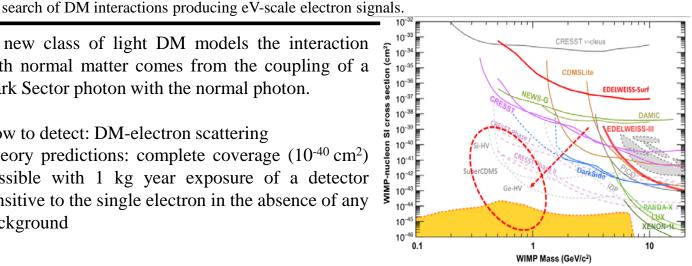
World leading results for DM search (a lot of data treatment is still in process)

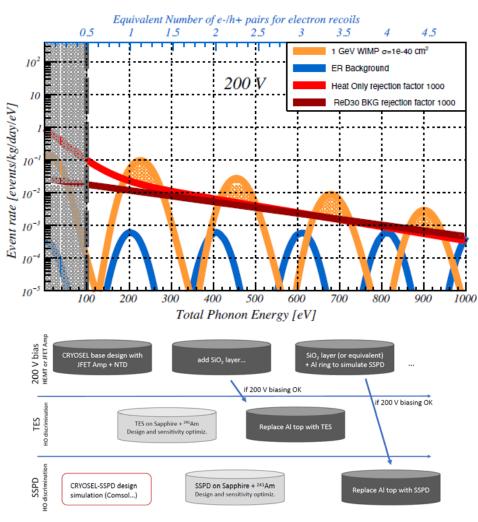


HPGe crystal y

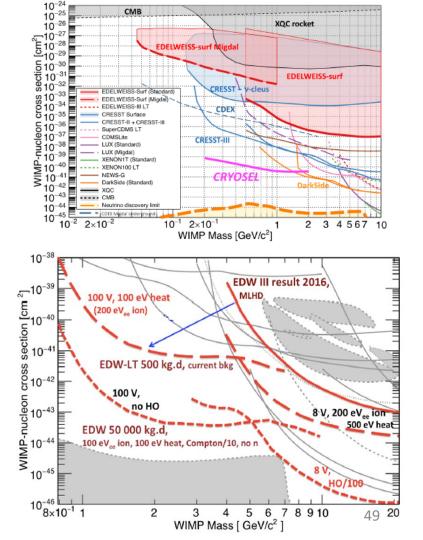
In new class of light DM models the interaction with normal matter comes from the coupling of a Dark Sector photon with the normal photon.

How to detect: DM-electron scattering Theory predictions: complete coverage (10^{-40} cm^2) possible with 1 kg year exposure of a detector sensitive to the single electron in the absence of any background





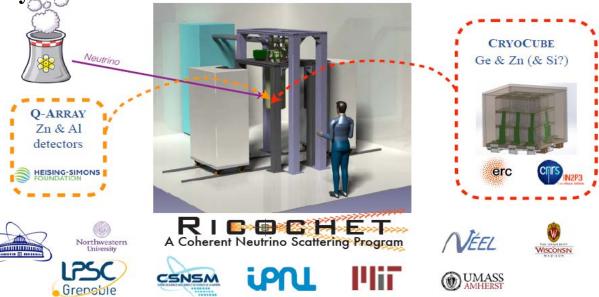
Lab and LSM physics runs to be planned upon R&D results



EDELWEISS bolometers - not only for DM

RICOCHET aims at building the ultra low-energy CEvNS neutrino observatory dedicated to physics beyond the Standard Model

50 eV energy threshold with a 10³ background rejection down to the threshold



The first key feature of the RICOCHET program, compared to other planned or ongoing CEvNS projects, is to aim for a kg-scale experiment with significant background rejection down to the O(10) eV energy threshold. The CRYOCUBE: a compact tabletop size setup

27 x 33 g detectors 8 x 8 x 8 cm³ radio-pure infrared-tight copper_box

Neutrino source

RICOCHET: first phase at ILL (Grenoble)

- 58 MW research reactor;
- Total neutrino flux: 1x10¹⁹ v/sec;
- 20 events of CEvNS per day for RICOCHET (1 kg) for 7 m distance;
- 3-4 ON/OFF per year;
- Cosmic shield ~15 m.w.e.;
- STEREO data about neutrino spectrum and backgrounds.



STEREO Coll., JINST 2018

