

# CHIRAL EFFECTS: NEW TRENDS

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# Outline of the talk

Introduction (Quantum anomalies and hydrodynamics)

Part I Gauge anomaly ( $\partial_\alpha J_5^\alpha = e^2 C_5^{el} \vec{E} \cdot \vec{B}$ )

- a) Changing the symmetry behind
- b) Anomalies as infrared effects

Part II Gravitational anomaly ( $\partial_\alpha J_5^\alpha = C_5^{grav} R\tilde{R}$ )

- a) Basic chiral eddect (kinematical chiral effect)
- b) Applications

Part II: collaboration with G. Yu. Prokhorov, O.V. Teryaev  
(talk here by G. Prokhorov)

Part I: more of review-type + earlier papers of ITEP group

# Chiral magnetic effect: a reminder

The best known effect is Chiral Magnetic (CME) Effect:

$$\vec{j}^{el} = e^2 \cdot C_5^{el} \mu_5 \vec{B}$$

where  $\mu_5$  is axial chemical potential  $\vec{B}$  is magnetic field,  
As indicated by the  $e^2$  factor, it is a Hall-type current.

The guess turns true if one keeps in mind  
specific hydrodynamic interaction

$$\hat{H}_{effective} = \hat{H}_{field\ theory} - \mu_{el} \hat{Q}_{el} - \mu_5 \hat{Q}_5$$

# Chiral vortical effect

using analogy between  $\vec{H}$  and  $\vec{\Omega}$   
brings in chiral vortical effect (CVE)

$$\vec{J}_5 = \mu_{el}^2 C_5^{el} \vec{\Omega}$$

which is most interesting  
since it survives in absence of  $\vec{E}, \vec{B}$  and is to be conserved

No place for a new Noether current (no free symmetry).

The way out: conservation is specific  
for absence of dissipation or for ideal fluid

as observed in 2016 but consequences emerge only recently

# Upgrading symmetry of the problem

For ideal fluid there is, indeed,  
another conserved axial charge (not chiral!):

$$Q_{\text{helicity}} = (\text{const}) \int d^3x \vec{v} \cdot (\vec{\nabla} \times \vec{v}) ,$$

(or  $J_{\text{helicity}}^\alpha \sim \epsilon^{\alpha\beta\gamma\delta} u_\beta \partial_\gamma u_\delta$  where  $u_\alpha$  is fluid 4-velocity)

as consequence of diffeomorphism symmetry of ideal fluid

Thus chiral symmetry is embedded into diffeomorphism, –  
fundamental change of symmetry

# Regenerating e-m interactions

To get CME back from CVE replace  $\partial_\alpha \rightarrow \nabla_\alpha$

Amusing **non-relativistic** analogs to chiral anomalies

$$\partial_\alpha \mathbf{J}_5^\alpha \sim \vec{E} \cdot \vec{B} \quad (\text{as "usual"})$$

$$\partial_\alpha \mathbf{J}_{el}^\alpha \sim \vec{E}_5 \cdot \vec{B} \equiv \vec{\nabla} \mu_{5,non-rel} \cdot \vec{B}$$

where  $\vec{E} \equiv \vec{\nabla} \mu$ ,  $\vec{E}_5 \equiv \vec{\nabla} \mu_5$

- no ultraviolet divergences
- non-relativistic anomalies match UV anomalies only up to a factor
- non-conservation of electric charge is declared

## A few references on ideal fluid and “anomalies”

“On consistency of hydrodynamic approximation for chiral media”, A. Avdoshkin , V.P. Kirilin , A.V. Sadofyev , V.I. Zakharov, e-Print: 1402.3587 [hep-th].

A. G. Abanov and P. B. Wiegmann, “Anomalies in fluid dynamics: flows in a chiral background via variational principle,” [arXiv:2207.10195 [hep-th]] + 3 other papers (2021)

“Divergence anomaly and Schwinger terms: Towards a consistent theory of anomalous classical fluids”, Arpan Krishna Mitra and Subir Ghosh, 2111.00473 [hep-th].

# Conclusions (preliminary) on Part I

What is **chiral fluid of massless quarks** at short distances  
might look as **ideal non-relativistic fluid** at large distances  
since anomalies are similar  
(variation of 't Hooft consistency condition)



# Gravitational anomaly and hydrodynamics

General Relativity is built on non-trivial metric tensor

$$g_{\mu\nu} \approx \eta_{\mu\nu} + h_{\mu\nu}$$

Gauge transformation, analog of

$$\delta A_\mu = \partial_\mu \Lambda$$

is

$$\delta h_{\mu\nu} = \partial_\mu \epsilon_\nu + \partial_\nu \epsilon_\mu$$

Gauge invariant field, analog of  $\vec{E}, \vec{B}$ , is Riemann tensor  $R_{\alpha\beta\gamma\delta}$  which involves  $(\partial_\alpha \partial_\beta h_{\mu\nu})$  **and**  $(\partial_\alpha h_{\mu\nu})^2$

Gravitational anomaly is built on the curvature  $R_{\alpha\beta\gamma\delta}$

In absence of sizable curvature we are interested in  
what is left of anomaly in absence of curvature

# What is left of GR in absence of curvature

**Equivalence principle:** physics in **non-inertial frame** is imitated by a nontrivial gravitational field

Two basic non-inertial frames, considered by Einstein:  
**accelerated and rotated frames**

(corresponding  $h_{\mu\nu}$  are easy to identify)

In applications to quark-gluon plasma acceleration and rotation 4-vectors are given by:

$$\mathbf{a}_\mu = u^\alpha \partial_\alpha \mathbf{u}_\mu, \quad \omega_\mu = \epsilon_{\mu\nu\rho\sigma} u^\nu \partial^\rho u^\sigma \quad (u^\nu \text{ is 4-velocity})$$

Plasma produced in heavy-ion collisions  
both accelerated and rotated

# Chiral Kinematical Effect (2022)

In presence of gravitational field the axial current  $J_5^\alpha$  defined as

$$\nabla_\alpha J_5^\alpha = C_5^{grav} R\tilde{R}$$

where  $C_5^{grav}$  depends on spin of fermions and are known  
In absence of curvature  $R_{\alpha\beta\gamma\delta}$  the current is still not vanishing and fixed uniquely in terms of kinematical quantities  $\mathbf{a}_\mu, \mathbf{w}_\mu$ .

In case of spin 1/2

$$J_5^{\alpha,kinematic} = -\frac{1}{24\pi^2} \left( 3\mathbf{a}_\mu^2 + \mathbf{w}_\mu^2 \right) w^\alpha$$

Idea and details of derivation in George's Prokhorov talk  
We add interpretation of the effect

## Unruh effect (a reminder)

Observer moving with acceleration  $\mathbf{a}$  with respect to Minkowskian vacuum sees thermal distribution of particles with temperature

$$T_{\text{Unruh}} = \frac{a}{2\pi} \quad (\text{quantum effect})$$

while an observer at rest sees no particles

For many, it sounds disappointing that the Unruh effect is observer-dependent. However, it is dynamical: By virtue of the equivalence principle, accelerated frame equivalent to vacuum in strong gravitational field resulting in the same acceleration  $\mathbf{a}$ . Naturally, such a field produces particles

Kinematical effects refer to noninertial frames describing results of measurements on the Unruh sample of particles

# Statistical vs gravitational approaches

Properties of fluids in equilibrium are evaluated statistically in terms of density operator, or effective interaction

$$\hat{H}_{\text{eff}} = \vec{\Omega} \cdot \hat{\vec{M}} + \vec{a} \cdot \hat{\vec{K}}$$

where  $\vec{M}$  is angular momentum and  $\vec{K}$  is the boost (for details see papers of F. Becattini)

In field theory, gravitational interaction is described by

$$\hat{H}_{\text{fund}} = \frac{1}{2} \hat{\Theta}^{\alpha\beta} h_{\alpha\beta}$$

where  $\Theta^{\alpha\beta}$  is the energy momentum tensor,  $h_{\alpha\beta}$  is the grav. potentials accommodating the same  $\vec{\Omega}$ ,  $\vec{a}$

Evaluate “external probes”,  $\langle \Theta^{\alpha\beta} \rangle$ ,  $\langle \mathbf{J}_5^\alpha \rangle$  Expect results to be the same (**duality**)

# More on duality

We have classical set up for duality:

- Two theories with same symmetry pattern: both ideal fluid and gravity are diffeomorphic invariant
- Infrared vs ultraviolet sensitive. Statistics not valid at short distances Field theory needs UV regularization
- Both theories are valid to evaluate a common set of quantities

As a result, evaluation of the kinematical effect can be rewritten as regularization, via acceleration, of gravitational chiral anomaly

## Conclusions on part II

- Working in non-inertial frames, like hydrodynamic gradient expansion valid in non-inertial frames, becomes a trend. Non-inertial frames is a part of gravity and in this sense universal. Specific feature: nonconservation of Noether currents. The new universality is not fully explored yet.