# **Critical effects on the evolution** of Quark Gluon Plasma

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## Motivation

- Hydro equations and the EOS with a critical point
- Propagation of perturbations in the QGP
  - Linear perturbations
  - Nonlinear perturbations
- Results & Discussion
- Summary & conclusion

# OUTLINE







- An anticipated critical end point (CEP) at the end of the first-order phase transition line at finite  $\mu_R$ .
- Location of the CEP is predicted by effective models, depends on the parameters of the models.
- Ab-initio lattice calculations are not valid at finite  $\mu_R$ .
- Relevant for RHIC-BES program, GSI-FAIR, NICA energies.



- Medium response is recorded to study the propagation of perturbations.
- Response could be imprinted in the hadron spectra.
- Isentropic trajectories near the CEP and away from the CEP will get influenced separately.
- Trajectories could provide information on the collision energies.
- Separating the collision energies by identifying hadron spectra.
- Dealing with small (linear) as well as large (nonlinear) perturbations.

# WHY PERTURBATIONS?





Promising observables: Flow harmonics are attributed to the hydrodynamic response of QGP to the initial geometry.

Formation of Mach cone due to shock wave propagation.

Aim to observe the effects of the CEP on such observables.

# **OBSERVABLES**

### spatial anisotropy — Flow







Müller-Israel-Stewart theory for describing the fluid.

An equation of state (EoS) to close the hydro equations.

Used an EoS embedding a critical point: constructed on the basis of universality hypothesis.

QCD critical point belongs to the same universality class as of the critical point of Ising model.

# EOS IN HYDRO EQUATIONS

### $(T_c, \mu_c) = (154, 367) MeV$ Entropy density







## To check the reliability of the EoS, we compare our results with the available lattice results S. Borsanyi et al, JHEP (2012), PLB (2014)



### Entropy density









S. Weinberg, Astrophys. J. 168, 175 (1971)

**\*** Separated out the real part and the imaginary part from dispersion relation.

**\*** Fate of a wave is dictated by magnitude of the real and the imaginary part of the dispersion relation.

★ A threshold wavelength is calculated via 
$$k_{th}$$
 by putting  $\left| \frac{\omega_{\Re e}}{\omega_{\mathcal{J}m}} \right| = 1$  Liao and Koch, PRC (2010)
$$k_{th} = \sqrt{\frac{a_0 d_0^2}{b_0 c_0^2}} - \frac{1}{2} \left( \frac{d_1 c_2^2}{b_0 d_0^2} + \frac{a_1 d_0}{a_0 b_0} - \frac{a_1 c_2^2 d_1}{a_0 b_0 c_0 d_0} + \frac{b_1^2 c_2^2 d_1}{a_0 b_0 d_0^3} - \frac{a_1 d_0}{a_0^2 b_0} \right) \left[ 1 + \frac{1}{2} \left( \frac{b_1 d_0^2}{b_0 c_0^2} + \frac{c_2 d_0^2}{c_0^3} - \frac{a_1 b_1 d_0^3}{a_0 b_0^2} + \frac{a_1 b_1 d_0^3}{a_0 b_0^3 c_0^2} - \frac{a_1 c_2 d_0^3}{b_0^3 c_0^3} \right) \right]$$
First-order correction

# LINEAR PERTURBATIONS

\* As strategy, a space-time dependent perturbation,  $e^{i(kx-\omega t)}$  is placed into the fluid via hydro variables.



## \* A wave with wavelength $\lambda \geq \lambda_{th}$ , can only propagate and rest will dissipate in the medium. Staig and Shuryak, PRC (2011)

are and are not dissipated by the viscosity effects.

highest order of surviving flow harmonics  $(n_v) \sim 1/R_v$  $R_v$ 

 $\diamond$  At the CEP, we found  $\lambda_{th}$  to be very large, implies survival of no wave.

# FATE OF THE PERTURBATION

# Viscous horizon is a length scale $(R_v) \sim \lambda_{th}$ : separates the wavelengths of the sound which







## **As consequences, Mach cone formation is prevented.**

## The viscous horizon scale is also divergent, means no wave can survive from viscous damping.

**All the possible flow harmonics collapse at the CEP.** 

## Ideally fully collapsed harmonics may be the signature of the CEP, but in experiments from the formation of the QGP to hadron is superposition of all the T, $\mu$ states, one expects suppression.





**Propagation of jets in QGP creates disturbances**  $\sim 1$  GeV/ fm<sup>3</sup>. Linear analysis is not valid, as they are nonlinear perturbations. **\*** The nonlinear perturbations are treated as nonlinear waves.



# NONLINEAR PERTURBATIONS

G. Sarwar, MH, M. Rahaman, A. Bhattacharyya and J. Alam, PLB (2021)

**Double hump structure at away** side jet in particle correlation at  $\Delta \phi = \pi \pm 1.2 rad$ STAR collaboration, PRL (2005) **Formation of** Mach cone

Mach cone formation is prevented in linear analysis. 

**We aim to observe if such phenomena occurs in** nonlinear waves too.



Hydro equations from MIS theory

Adopted Reductive Perturbative Method (R)

$$\hat{\epsilon} = \frac{\epsilon}{\epsilon_0} = 1 + \sigma\epsilon_1 + \sigma^2\epsilon_2 + \sigma^3\epsilon_3 + \dots$$

• Different order of  $\sigma$  results in Breaking wave, Burger's, KdV equations.

EOM for First-order perturbation

$$\frac{\partial \hat{\epsilon}_1}{\partial t} + [1 + \frac{3}{2}q_1\hat{\epsilon}_1]c_s\frac{\partial \hat{\epsilon}_1}{\partial x} + [\frac{1}{2(\epsilon_0 + p_0)}(\zeta + \frac{4}{3}\eta) - \kappa q_2]$$

EOM for Second-order perturbation

$$\frac{\partial \hat{\epsilon}_2}{\partial t} + \mathcal{S}_1 \frac{\partial \hat{\epsilon}_2}{\partial x} + \mathcal{S}_2 \frac{\partial \hat{\epsilon}_1}{\partial x} + \mathcal{S}_3 (\frac{\partial \hat{\epsilon}_1}{\partial x})^2 + \mathcal{S}_4 \frac{\partial^2 \hat{\epsilon}_1}{\partial x^2} + \mathcal{S}_5 \frac{1}{\hat{\epsilon}_1} \frac{\partial \hat{\epsilon}_1}{\partial x} \frac{\partial^2 \hat{\epsilon}_1}{\partial x^2} + \mathcal{S}_6 \frac{\partial^3 \hat{\epsilon}_1}{\partial x^3} + \mathcal{S}_7 \frac{\partial^2 \hat{\epsilon}_2}{\partial x^2} = 0$$

# **EOM: NONLINEAR PERTURBATIONS**

**PM) technique:** 
$$X = \frac{\sigma^{1/2}}{L}(x - c_s t), \quad Y = \frac{\sigma^{3/2}}{L}(c_s t)$$

Fogaça et. al., PRC (2010)

## **Initial profiles**

$$\frac{\partial^2 \hat{\epsilon}_1}{\partial x^2} = 0$$

$$\hat{\epsilon}_1 = A_1 \left[ \operatorname{sech} \frac{(x - x_0)}{B_1} \right]$$
$$\hat{\epsilon}_2 = A_2 \left[ \operatorname{sech} \frac{(x - x_0)}{B_2} \right]$$









 $\checkmark$  Viscous coefficients  $\eta, \zeta, \kappa$  are present.

Waves are attenuated, but not dissipated completely.

# FATE OF NONLINEAR PERTURBATIONS

## **NEAR CRITICAL POINT**

- $\checkmark$  Viscous coefficients  $\eta, \zeta, \kappa$  are present.
- Waves dissipate rapidly in presence of the critical point.





Like the linear disturbances, the nonlinear disturbances too created by jets are highly suppressed. The Mach cone can not form, thus it will affect the particle correlation. Flow harmonics will be suppressed.

The trajectory of the system in the QCD phase diagram will be event-dependent.

### Nonaka et. al, NPA (2009)



In event-by-event analysis, isentropic trajectories form a few events may go through the CEP and the others away from the CEP, cause large event-by-event fluctuation of flow harmonics.





# **Critical effects are observed through an EoS containing the critical point.** implying forbiddance of propagation of perturbations. the critical point.



- \* Studied linear perturbations by imparting a perturbation  $e^{i(kx-\omega t)}$  through hydro variables.
- $\diamond \lambda_{th}$  is calculated via dispersion relation, and it is found to be diverging near the critical point,
- **Consequently, the speed of sound in the QGP vanishes, insinuates the formation of Mach cone.**
- **\*** Viscous horizon scale also diverges and we find no harmonics to survive when the system hits



- - dissipated near the CEP.
- \* As a result, the Mach cone effect will disappear in the particle correlation of jets.
  - **We argue that the disappearance of Mach cone effects in particle**

**Along** with the linear perturbations, we have also studied the nonlinear perturbations in the fluid. \* The nonlinear waves survives the damping, when the system is away from the CEP, but highly

correlation and the enhancement of fluctuations in flow harmonics in the event-by-event analysis may be considered as signals of the critical point.

Thank you



• The EOM is obtained by the conservation equations:

# **BACKUP: MIS THEORY**

## $+\Pi)\Delta^{\mu\nu}+h^{\mu}u^{\nu}+h^{\nu}u^{\mu}+\pi^{\mu\nu}$

 $nu^{\mu} - nq^{\mu}/(\epsilon + p)$ 



![](_page_17_Picture_6.jpeg)

- Constructed on the basis of universality hypothesis: says that CEP of the QCD belongs to the same universality class of 3D Ising model.
- The universality hypothesis permits the linear mapping between the CEP of the Ising model and the CEP of the QCD phase diagram as:
- Magnetization, M(r, h), plays as an order j critical entropy density s<sub>c</sub>(T, µ) in QCD

$$s_c(T,\mu) = \frac{M(r,h)}{\Delta T_c}$$

$$\frac{-\mu_c}{\mu_c}, \ h = \frac{T - T_c}{\Delta T_c}$$

Magnetization, M(r, h), plays as an order parameter in Ising model and that is mapped as

$$= M\left(\frac{\mu - \mu_c}{\Delta \mu_c}, \frac{T - T_c}{\Delta T_c}\right) \frac{1}{\Delta T_c}$$

![](_page_18_Picture_8.jpeg)

## A dimensionless entropy density is calculated as:

$$S_c = \left[ D \sqrt{\Delta T_c^2 + \Delta \mu_c^2} \right] S_c$$

• The entropy density of  $s_Q$  and  $s_H$  is connected by using  $S_c$  as switching function as:

$$s(T,\mu) = \frac{1}{2} [1 - \tanh S_c(T,\mu)] s_Q(T,\mu) + \frac{1}{2} [1 + \tanh S_c(T,\mu)] s_H(T,\mu)$$

**Once entropy density is known, other thermodynamic variables can be calculated as:** 

$$n(T,\mu) = \int_0^T \frac{\partial s(T',\mu)}{\partial \mu} dT' \qquad p(T,\mu) = \int_0^T s(T',\mu) dT' \qquad \varepsilon(T,\mu) = Ts(T,\mu) - p(T,\mu) + \mu r$$

To get the first order phase boundary, the discontinuity in the entropy density along the transition line

$$\frac{\partial T_c(\mu)}{\partial \mu} \left[ \left[ s(T_c(\mu) + \Delta, \mu) - s(T_c(\mu) - \Delta, \mu) \right] \right]$$

![](_page_19_Picture_12.jpeg)

### • The EOMs for different components of perturbation are obtained by the conservation equations. **Dispersion relation** Imaginary part Real part

$$\omega_{\text{Re}} = \sqrt{\frac{a_0}{b_0}} \left[ k - \frac{1}{2} \frac{a_1}{a_0} k^2 + \left( \frac{1}{2} \frac{a_2}{a_0} - \frac{1}{8} a_1 a_0^2 + \frac{b_1}{b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} a_1^2 a_0^3 - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_2}{a_0^2} + \frac{1}{16} \frac{a_1 a_0^2}{a_0^2} - \frac{1}{2} \frac{a_1 b_1}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_1 a_0^2}{a_0^2} - \frac{1}{2} \frac{a_1 a_0^2}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_1 a_0^2}{a_0^2} - \frac{1}{2} \frac{a_0 a_0^2}{a_0 b_0} \right) k^3 + \left( \frac{1}{4} \frac{a_1 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_1 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} \right) k^3 + \left( \frac{1}{4} \frac{a_0 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} \right) k^3 + \left( \frac{1}{4} \frac{a_0 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} \right) k^3 + \left( \frac{1}{4} \frac{a_0 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} \right) k^3 + \left( \frac{1}{4} \frac{a_0 a_0^2}{a_0^2} + \frac{1}{16} \frac{a_0 a_0^2}{a_0^2} \right) k^3 + \left( \frac{1}{4}$$

$$\begin{split} a_{0} &= 9h \bigg[ \bigg( \frac{\partial P}{\partial T} \bigg)_{n} + \alpha_{1} n \bigg\{ \bigg( \frac{\partial \epsilon}{\partial n} \bigg)_{T} \bigg( \frac{\partial P}{\partial T} \bigg)_{n} - \bigg( \frac{\partial \epsilon}{\partial T} \bigg)_{n} \bigg( \frac{\partial P}{\partial n} \bigg)_{T} \bigg\} \bigg], \\ a_{1} &= \frac{9\alpha\beta_{1}n^{2}T^{2}\kappa^{2}}{h} + 12\alpha_{1}\eta\kappa nT \bigg[ \alpha + \frac{\alpha n}{h} \bigg( \frac{\partial \epsilon}{\partial n} \bigg)_{T} - \frac{T}{h} \bigg( \frac{\partial P}{\partial T} \bigg)_{n} + \frac{T}{h} \bigg( \frac{\partial \epsilon}{\partial T} \bigg)_{n} \bigg], \\ a_{2} &= \frac{9\beta_{1}\kappa^{2}n^{2}T}{h} \bigg[ \bigg( \frac{\partial n}{\partial T} \bigg)_{\mu} \bigg( \frac{\partial \mu}{\partial n} \bigg)_{T} + \bigg( \frac{\partial P}{\partial T} \bigg)_{n} \bigg( \frac{\partial P}{\partial n} \bigg)_{T} \bigg] \\ &+ \frac{12\alpha_{1}\eta\kappa nT}{h} \bigg[ \bigg( \frac{\partial P}{\partial T} \bigg)_{n} + n \bigg( \frac{\partial n}{\partial T} \bigg)_{\mu} \bigg( \frac{\partial \mu}{\partial n} \bigg)_{T} + \frac{n}{h} \bigg( \frac{\partial P}{\partial T} \bigg)_{n} \bigg( \frac{\partial \epsilon}{\partial n} \bigg)_{T} \bigg], \\ b_{0} &= 9h \bigg( \frac{\partial \epsilon}{\partial T} \bigg)_{n}, \\ b_{1} &= 24\beta_{2}\eta^{2} \bigg( \frac{\partial \epsilon}{\partial T} \bigg)_{n} + \frac{9\beta_{1}\kappa^{2}n^{2}}{h} \bigg[ T \bigg( \frac{\partial \mu}{\partial n} \bigg)_{T} \bigg( \frac{\partial n}{\partial T} \bigg)_{\mu} - T^{2} \bigg( \frac{\partial \epsilon}{\partial n} \bigg)_{T} \bigg( \frac{\partial \mu}{\partial n} \bigg)_{n} + \alpha \bigg( \frac{\partial \epsilon}{\partial n} \bigg)_{T} \bigg)_{n} \end{split}$$

# **DISPERSION RELATION**

![](_page_20_Picture_4.jpeg)

$$\omega_{\rm Im} = -\frac{c_0}{d_0} \left[ k^2 - \frac{c_1}{c_0} k^3 - \left( \frac{d_1}{d_0} + \frac{c_2}{c_0} \right) k^4 \right]$$

$$\begin{split} c_{0} &= 2\eta h^{2} \left( \frac{\partial \epsilon}{\partial T} \right)_{n} - 3hn^{2} \kappa T \beta_{1} \bigg[ \alpha \kappa \left( \frac{\partial \epsilon}{\partial n} \right)_{T} + h \bigg( \frac{\partial n}{\partial T} \bigg)_{\mu} \bigg( \frac{\partial \mu}{\partial n} \bigg)_{T} + \frac{\alpha_{1}}{\beta_{1}} \bigg( \frac{\partial P}{\partial T} \bigg)_{n} \bigg( c_{1} &= 2\alpha \beta_{1} \eta n^{2} T \bigg( T \kappa^{2} + 4h\eta \frac{\beta - 2}{\beta_{1}} \bigg), \\ c_{2} &= 8\beta_{2} \eta \kappa n^{2} T \bigg[ \bigg( \frac{\partial \epsilon}{\partial T} \bigg)_{n} - \bigg( \frac{\partial P}{\partial n} \bigg)_{T}^{2} \bigg( \frac{\partial \epsilon}{\partial n} \bigg)_{T} \bigg], \\ d_{0} &= 3h^{3} \bigg( \frac{\partial \epsilon}{\partial T} \bigg)_{n}, \\ d_{1} &= 3h\beta_{1} n^{2} \kappa \bigg[ \alpha \kappa \bigg( \frac{\partial \epsilon}{\partial n} \bigg)_{T} + 4T^{2} \kappa \bigg( \frac{\partial \epsilon}{\partial T} \bigg)_{n} \bigg( \frac{\partial \mu}{\partial n} \bigg)_{T}^{2} - T \bigg( \frac{\partial n}{\partial T} \bigg)_{n} \bigg( \frac{\partial \epsilon}{\partial n} \bigg)_{T} \bigg]. \end{split}$$

![](_page_20_Picture_7.jpeg)

 $\partial \epsilon$  $\overline{\partial n}$