



Reconstruction of neutrino direction in the Baikal-GVD experiment by neural networks

The study was supported by a grant from the Russian Science Foundation 22-22-20063

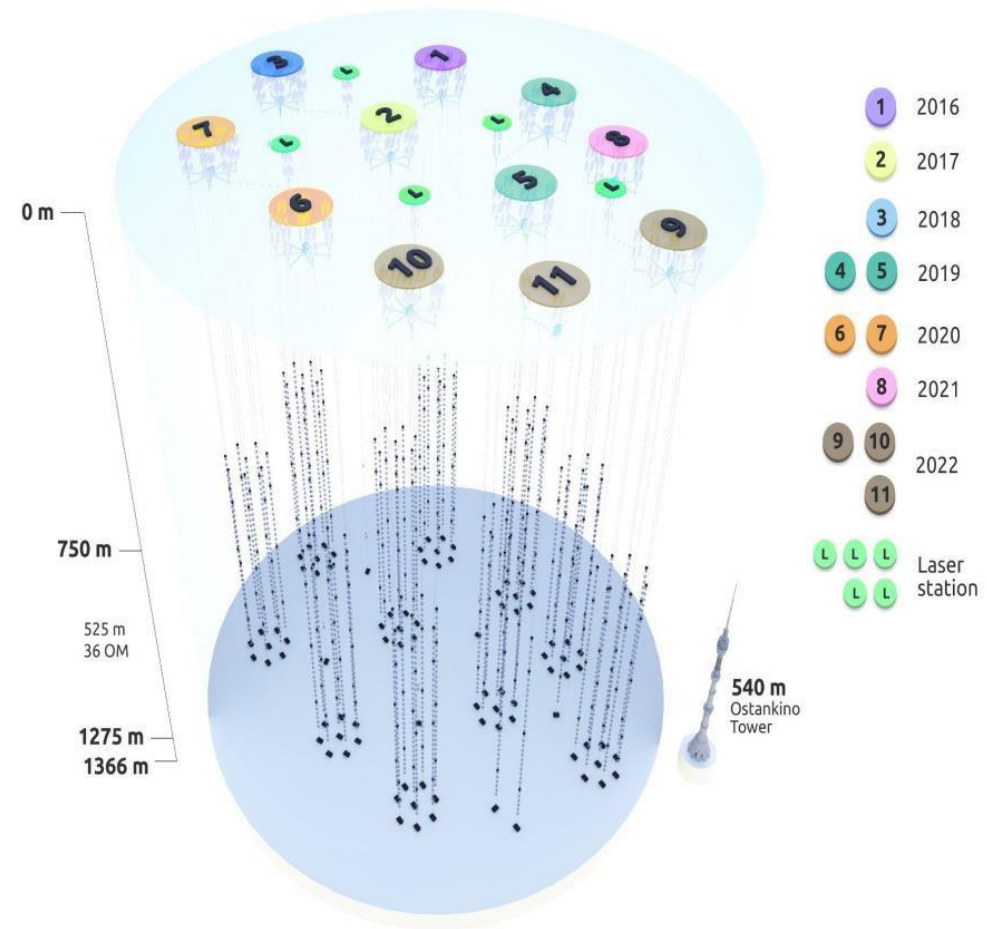
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Dubna 2023

Baikal-GVD Experiment

Baikal-GVD - neutrino underwater telescope

- The telescope consists of 12 clusters (winter 2023)
- Cluster - 8 strings, in the form of a regular heptagon with a center
- Each string has 36 evenly spaced detectors



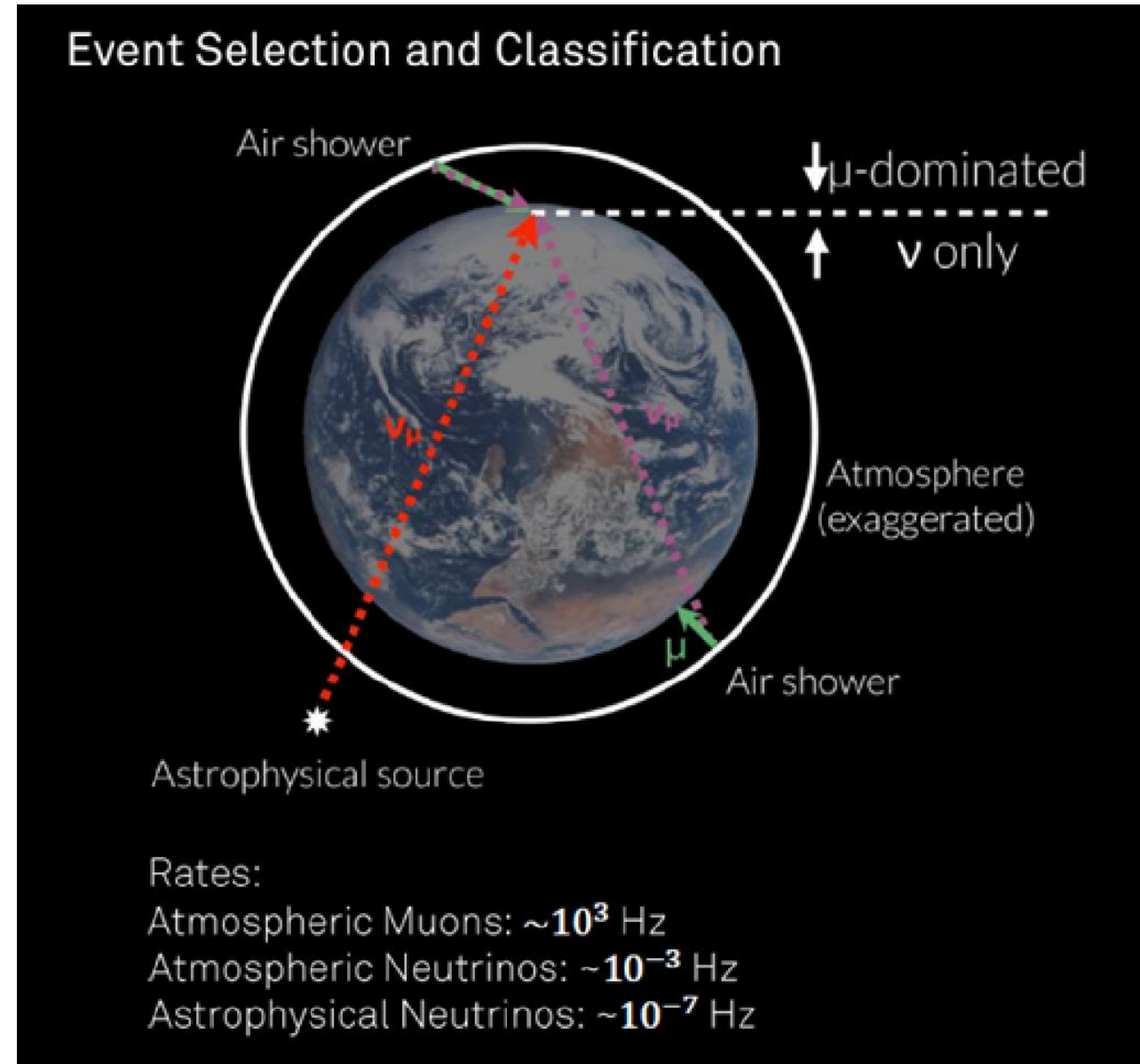
Status for 2021 [1]

[1] <https://arxiv.org/abs/2109.14344>

Registered particles

1) Data - atmospheric muons (top) and neutrinos (bottom)

2) Goal - astrophysical neutrinos



Relevance

- Neutrinos are not deflected by a magnetic field and have a small interaction cross section.

- The angular distribution of atmospheric neutrinos is successfully modeled.

- Astrophysical neutrinos are an important source of information about high-energy astrophysical objects.

- Improving the angular resolution will make it possible to reduce the background of atmospheric events when studying astrophysical neutrinos.

Data

Monte Carlo simulation[2]

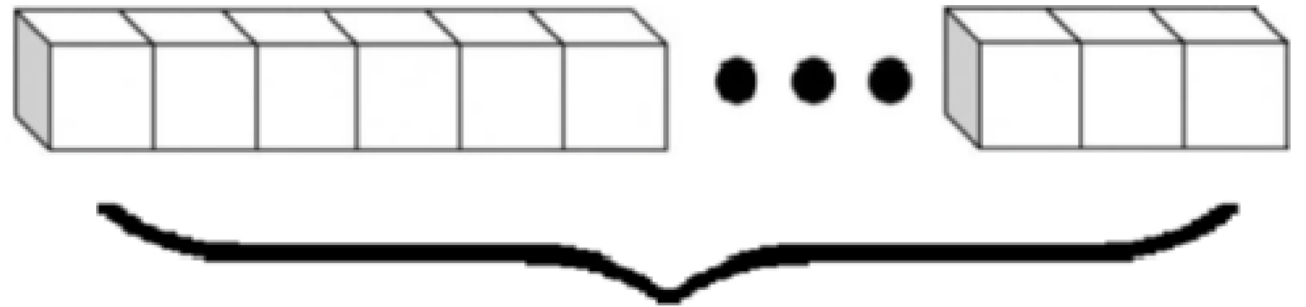
- $\sim 1.6 \cdot 10^6$ events
- Single-cluster events ($\approx 90\%$ of all events)
- String number > 1

Detector signal:

- 1) Charge
- 2) Response time
- 3-5) Coordinates

- Signals are ordered by time and cleared of noise

$$\text{input} = \{x, y, z, t, Q\}$$



$$N = 32$$

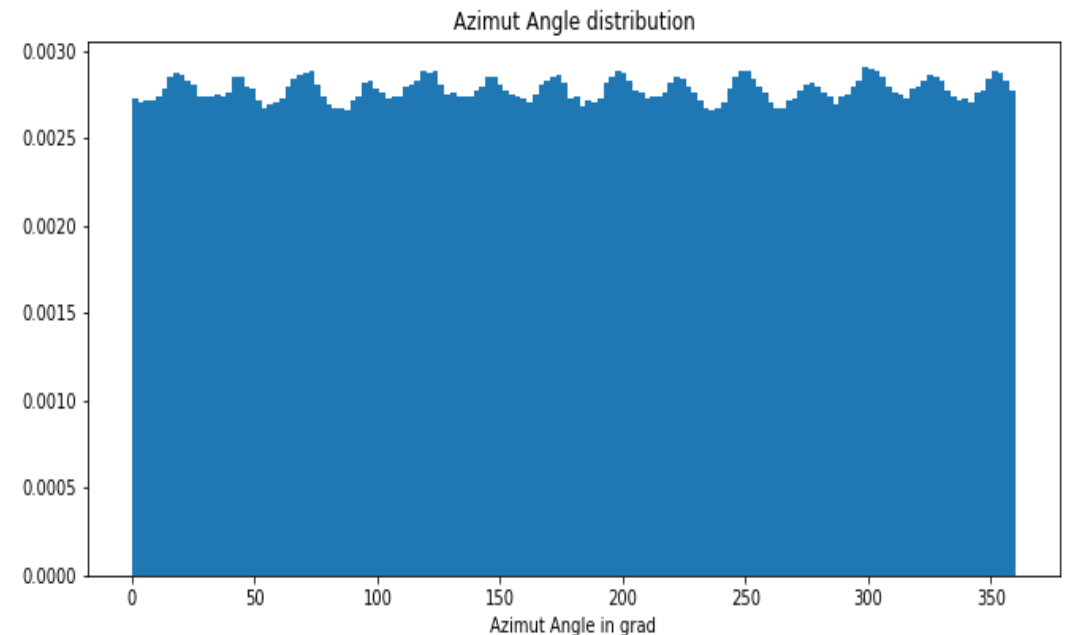
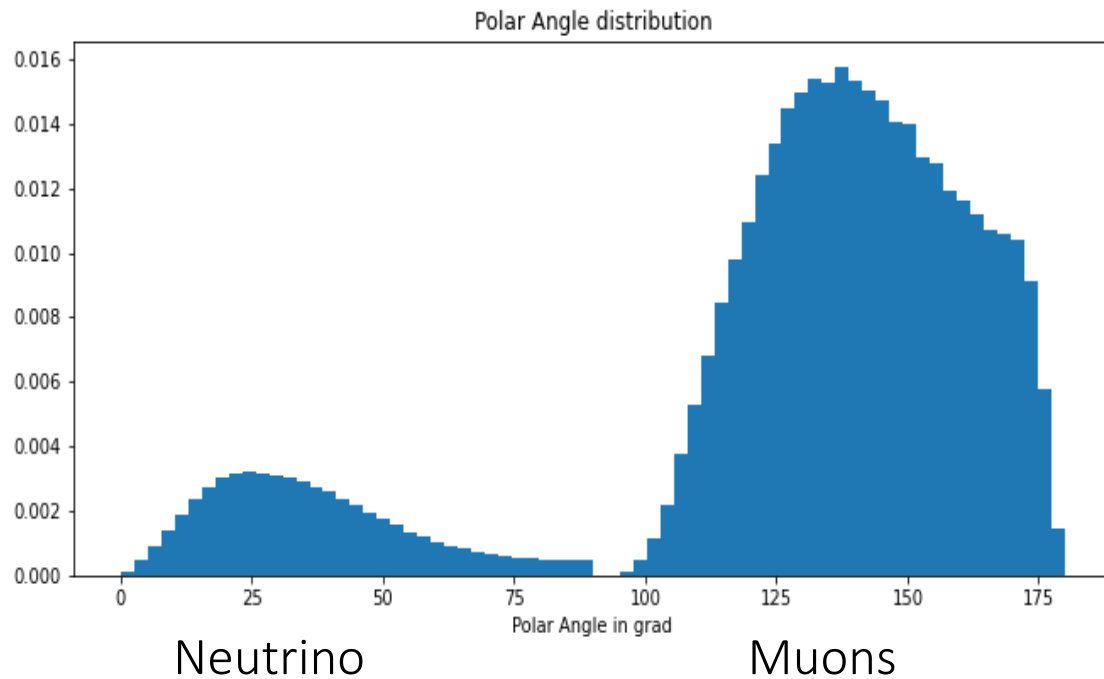
[2] arxiv.org/abs/2106.06288

Reconstruction

Direction => 3-dimensional polar vector
Direction => polar/azimuth angles

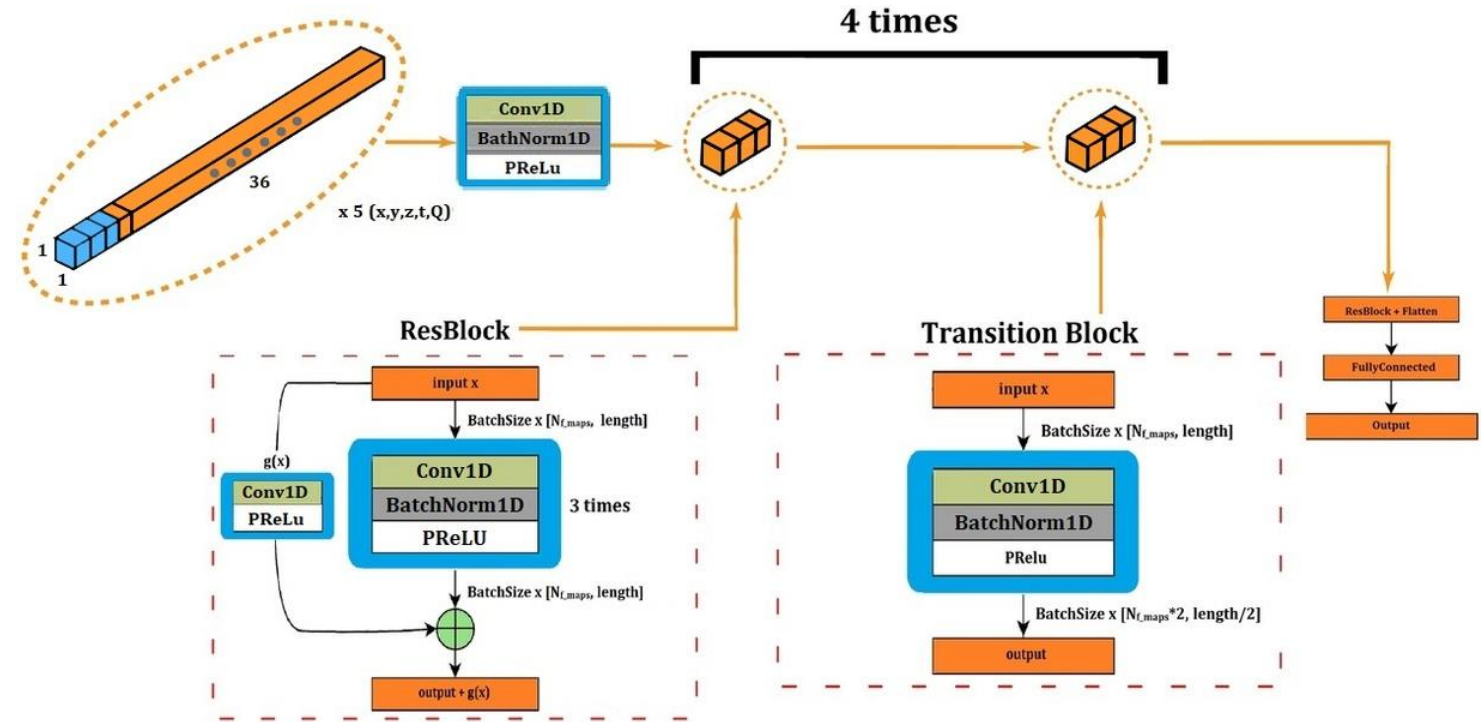
$$\vec{R} = (\cos \phi \sin \theta, \sin \phi \sin \theta, \cos \theta)$$

Angle distributions



Models

- Convolutional neural network (CNN) based on the ResNet network [3]
- Graph neural network (GNN), based on EdgeCNN[4]
- Network performance is assessed using median angular resolutions

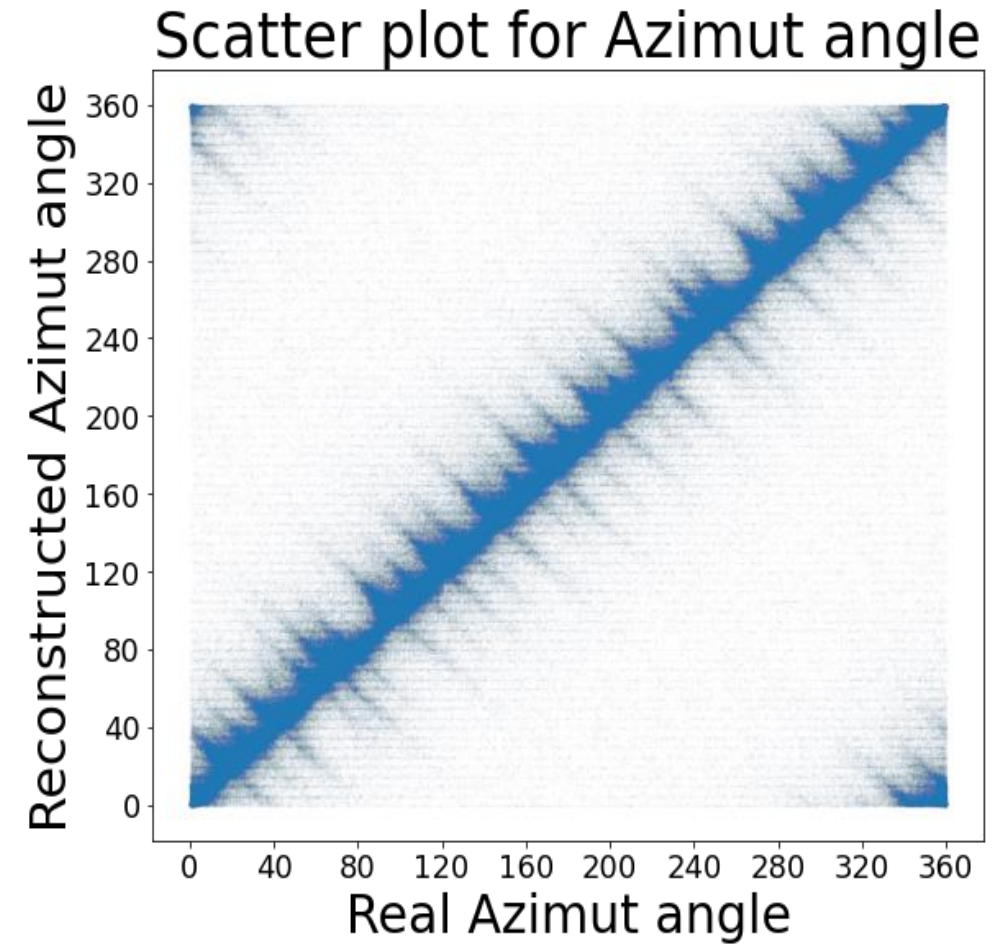
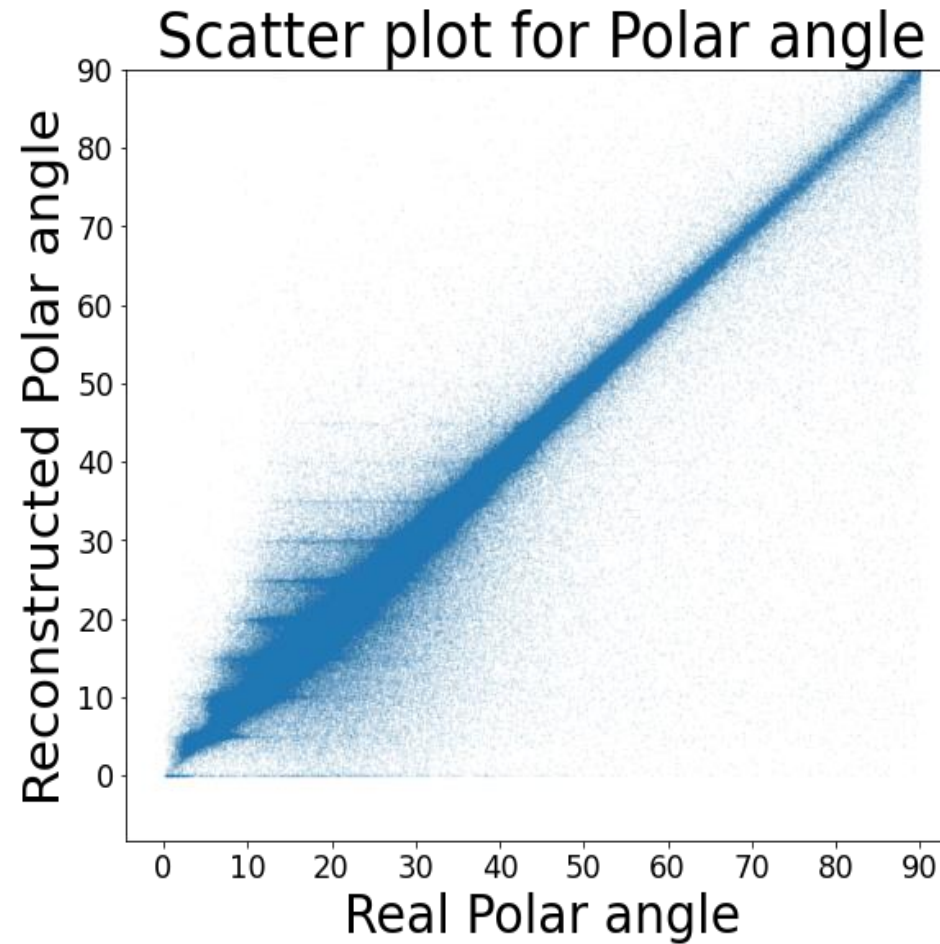


CNN architecture

[3] arxiv.org/abs/1512.03385

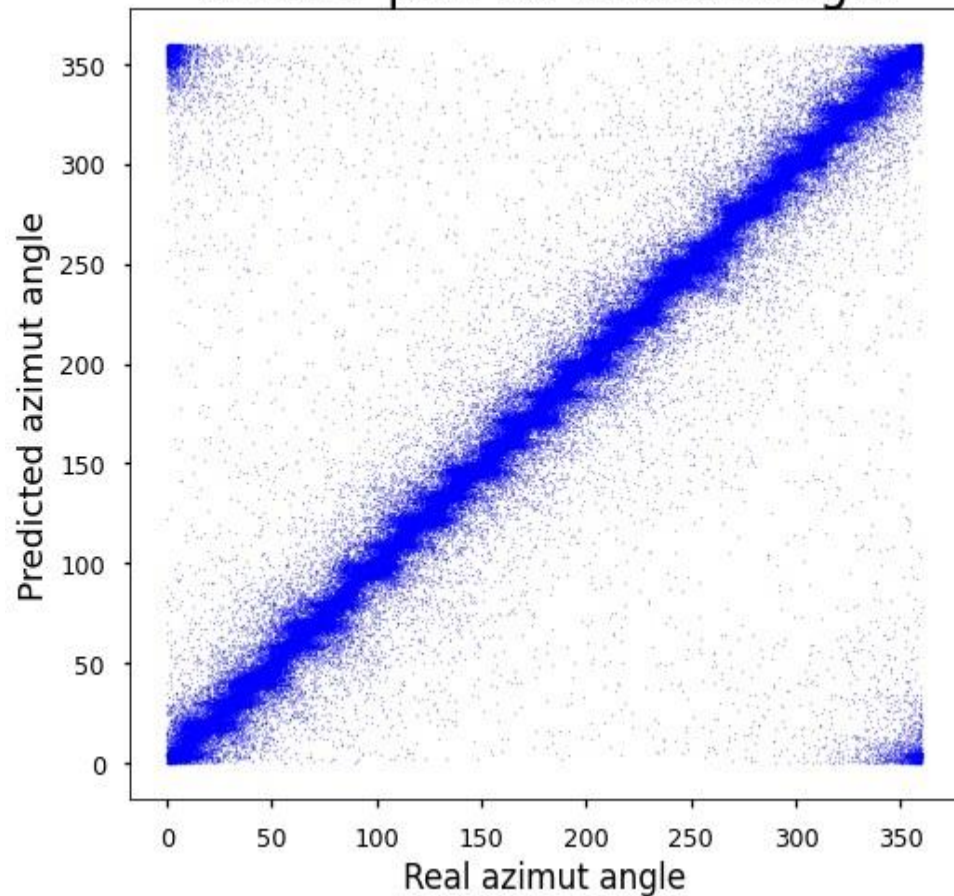
[4] arxiv.org/abs/1801.07829

Standard reconstruction[4]

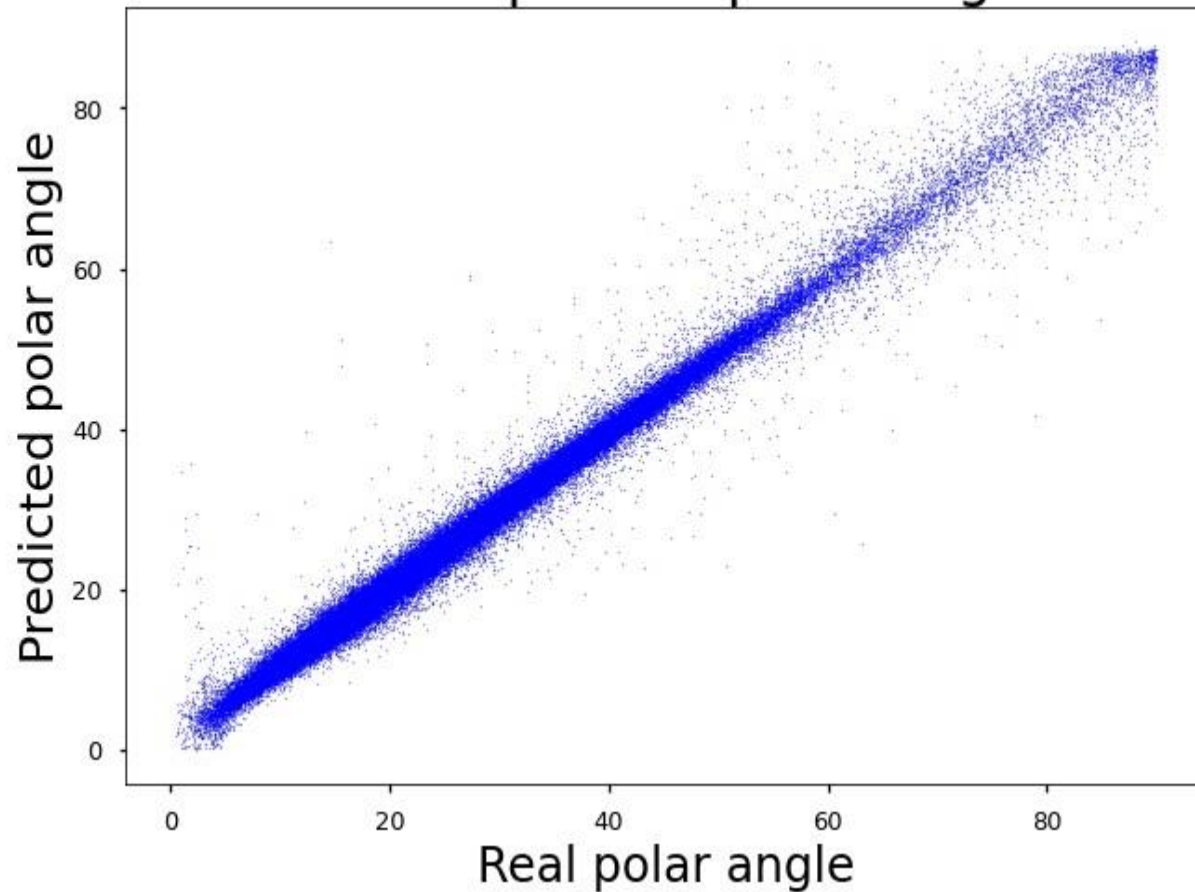


Reconstructed Angles

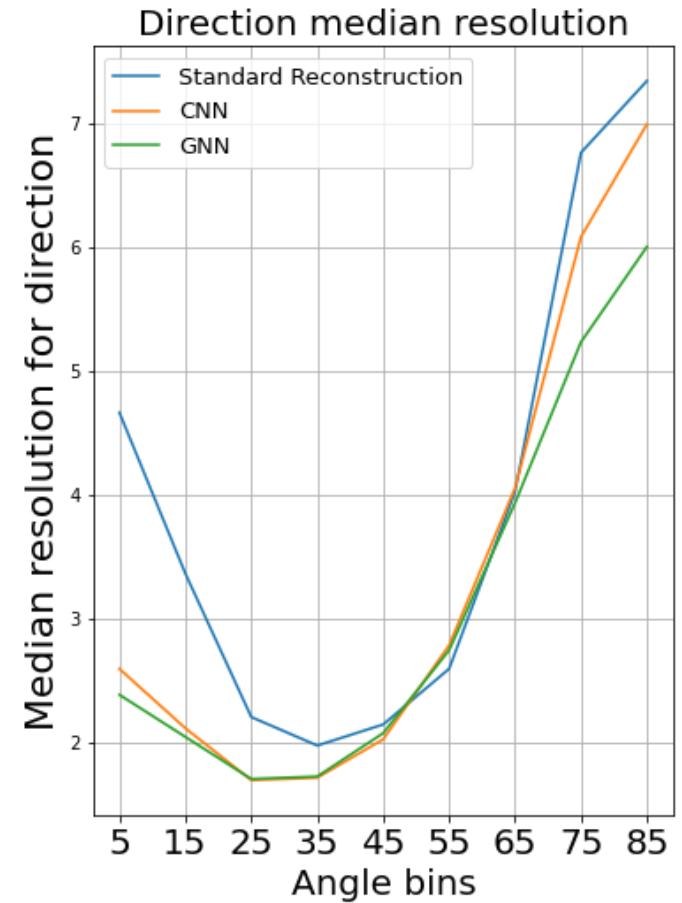
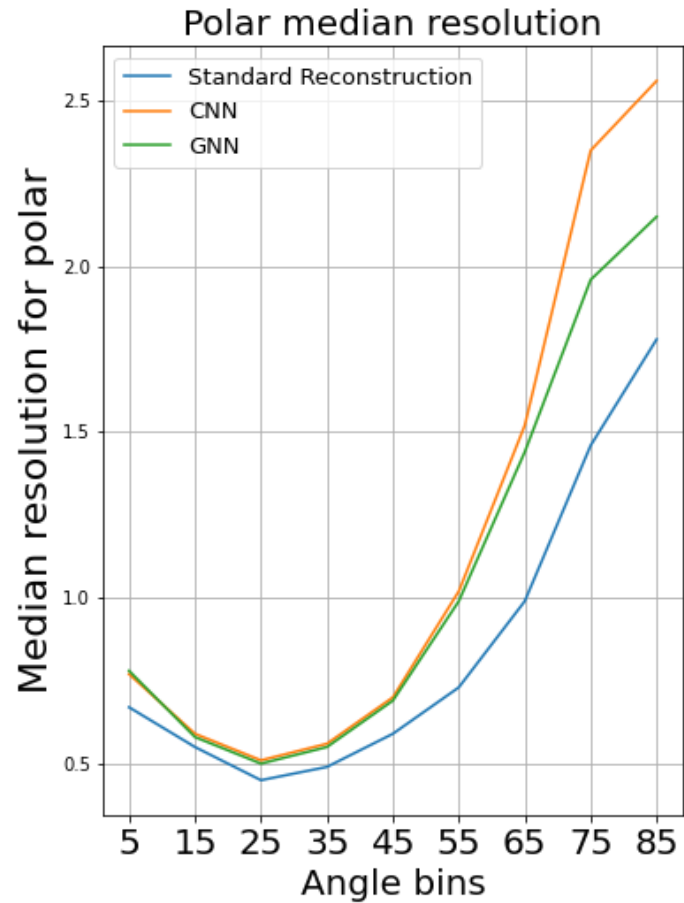
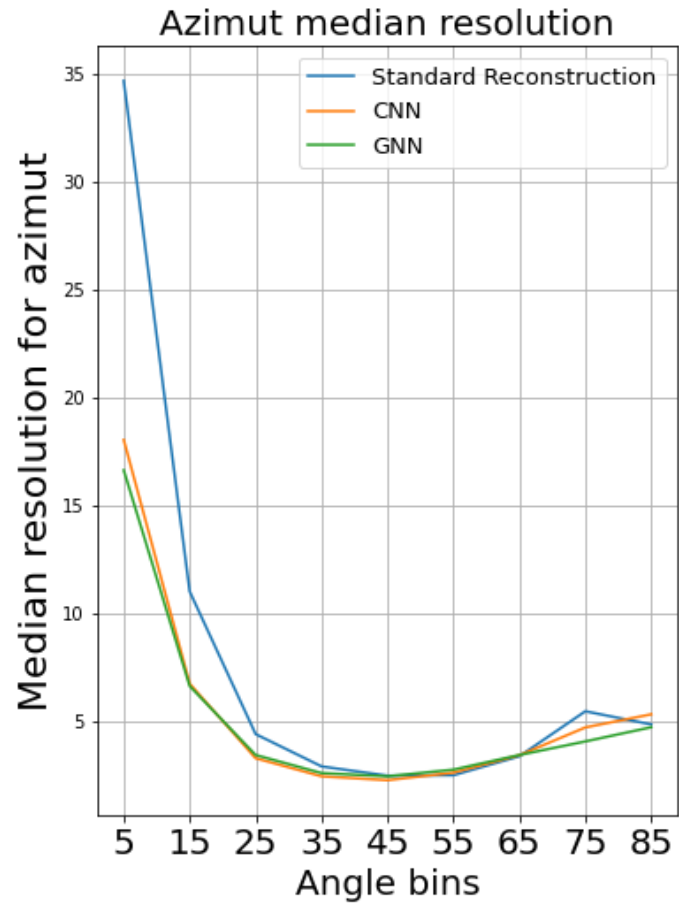
Scatter plot for azimuth angle



Scatter plot for polar angle



Angular resolutions CNN/GNN



Models Metrics

Median Resolution	Azimuth Angle	Polar Angle	Direction
Standard Reconstruction	5.44°	0.52°	2.64°
GCN	3.90°	0.58°	1.93°
CNN	3.82°	0.59°	1.95°

- At small polar angles, the direction is restored more accurately
- The poorer recovery of the azimuthal angle can be explained by the specific structure of the clusters.
- The GCN and CNN model show similar performance, which outperforms the standard algorithm.

Conclusions

- In this problem, deep learning algorithms are applicable and comparable in metrics to standard reconstruction.
- Neural networks perform better in the region of small polar angles

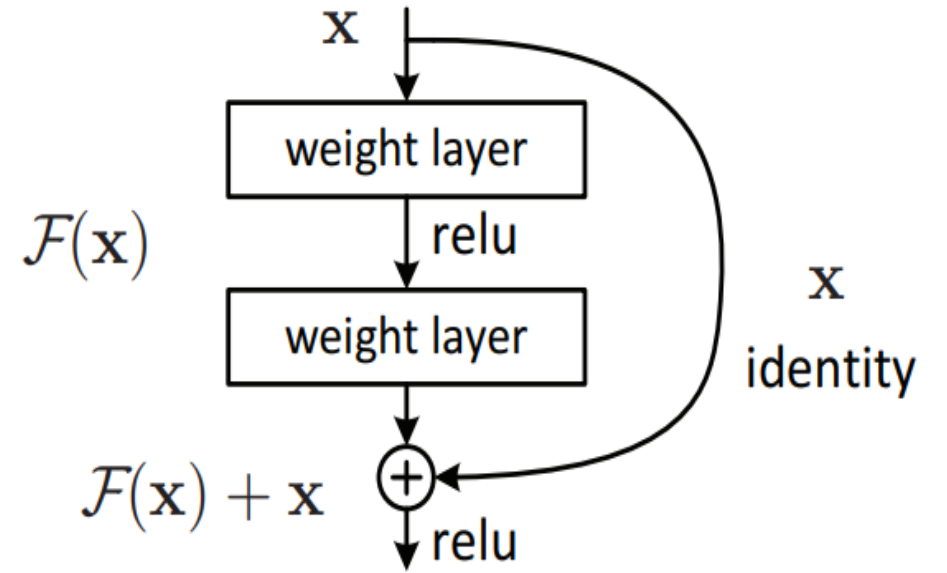
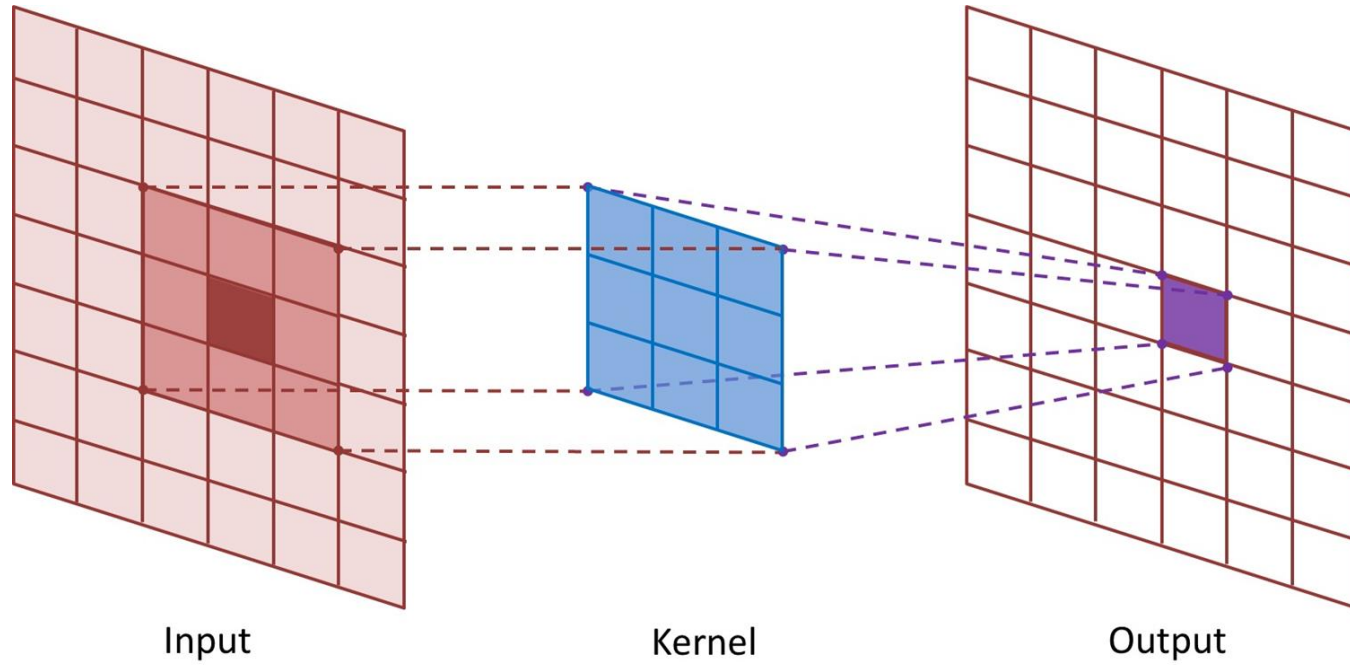
Further development:

- Optimizing the architectures of the models
- Use neural networks to predict confidence in prediction
- Transformers

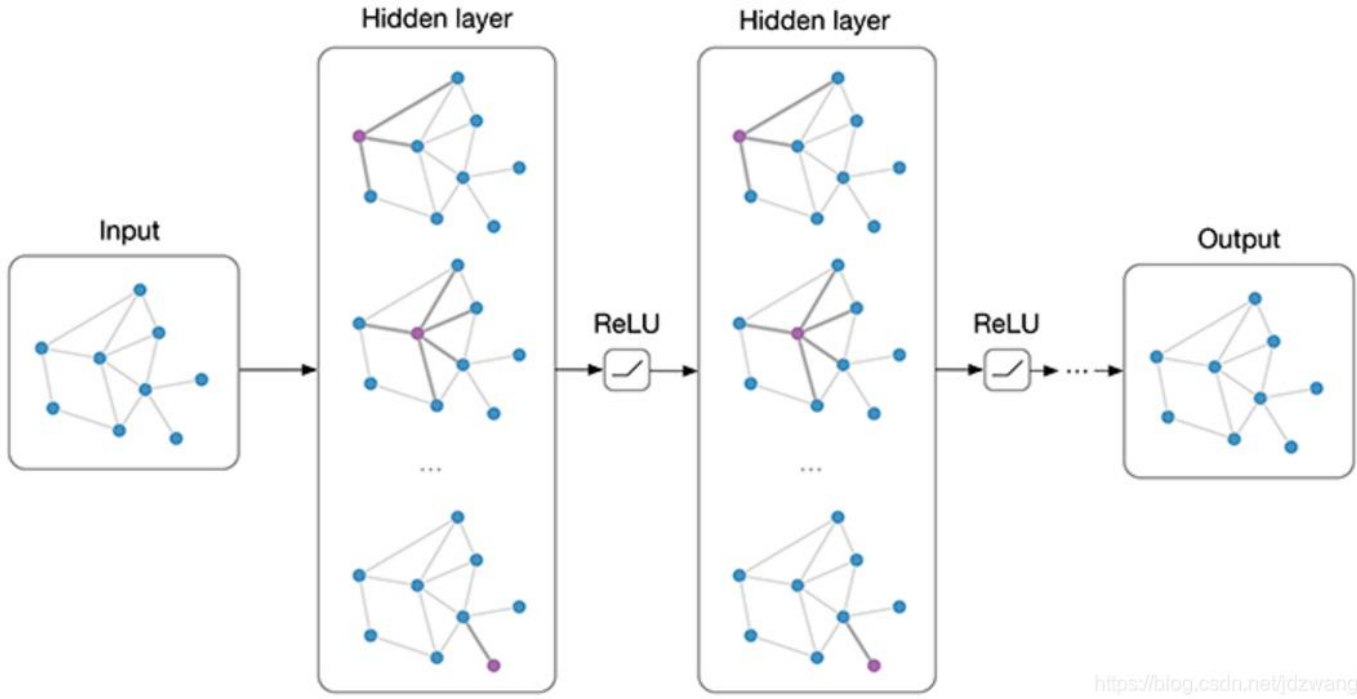
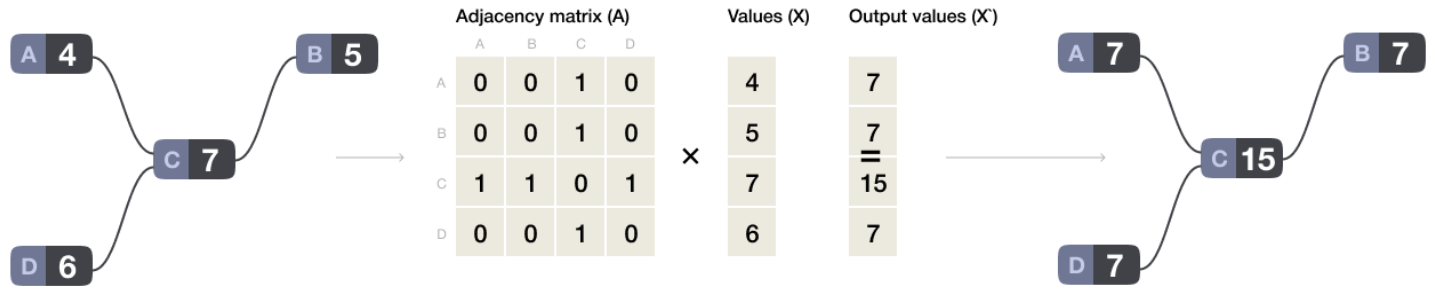
THANK YOU FOR YOUR ATTENTION!

Applications

CNN, Residual connection

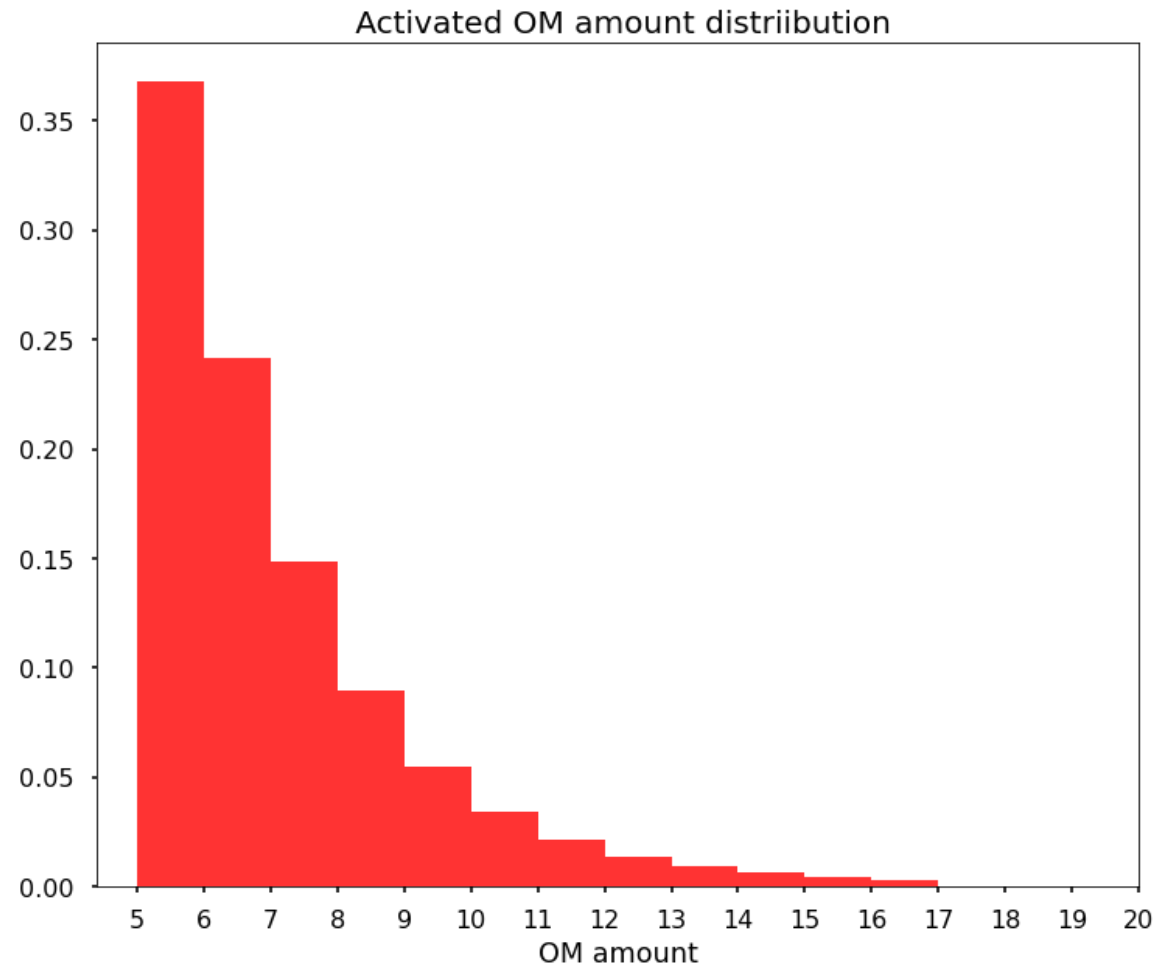


Graph Convolution Layer

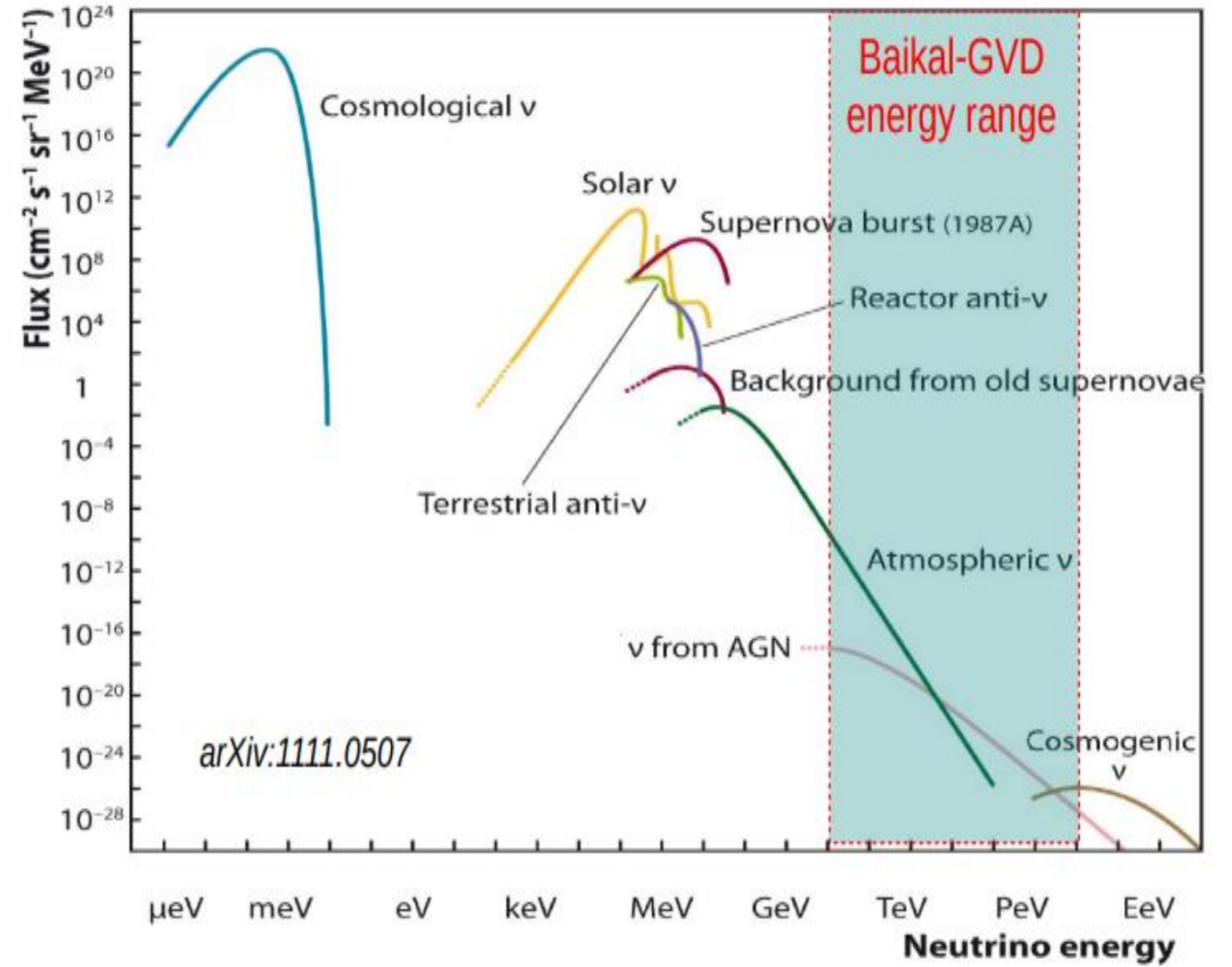
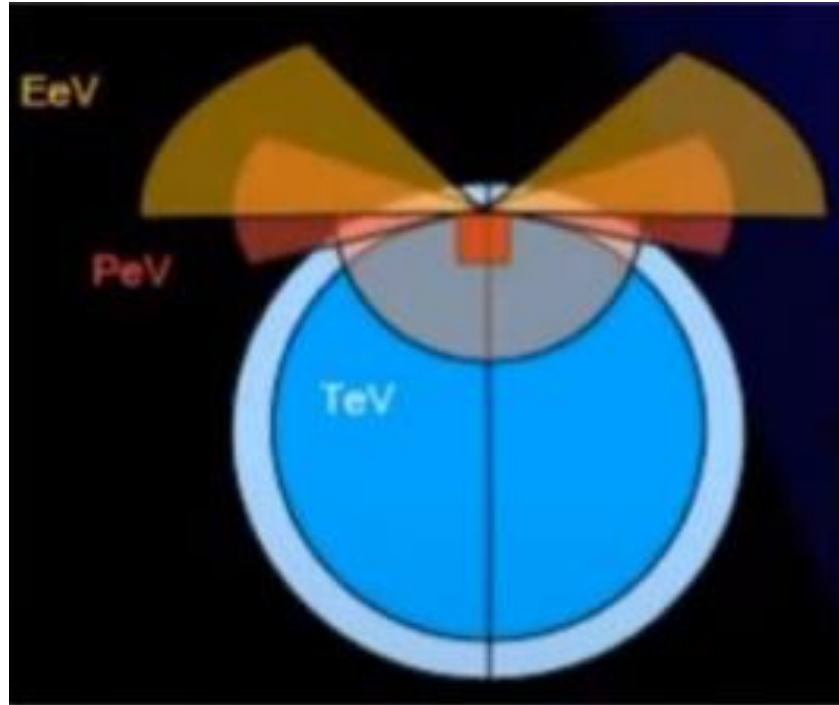


<https://blog.csdn.net/jdzwanghao>

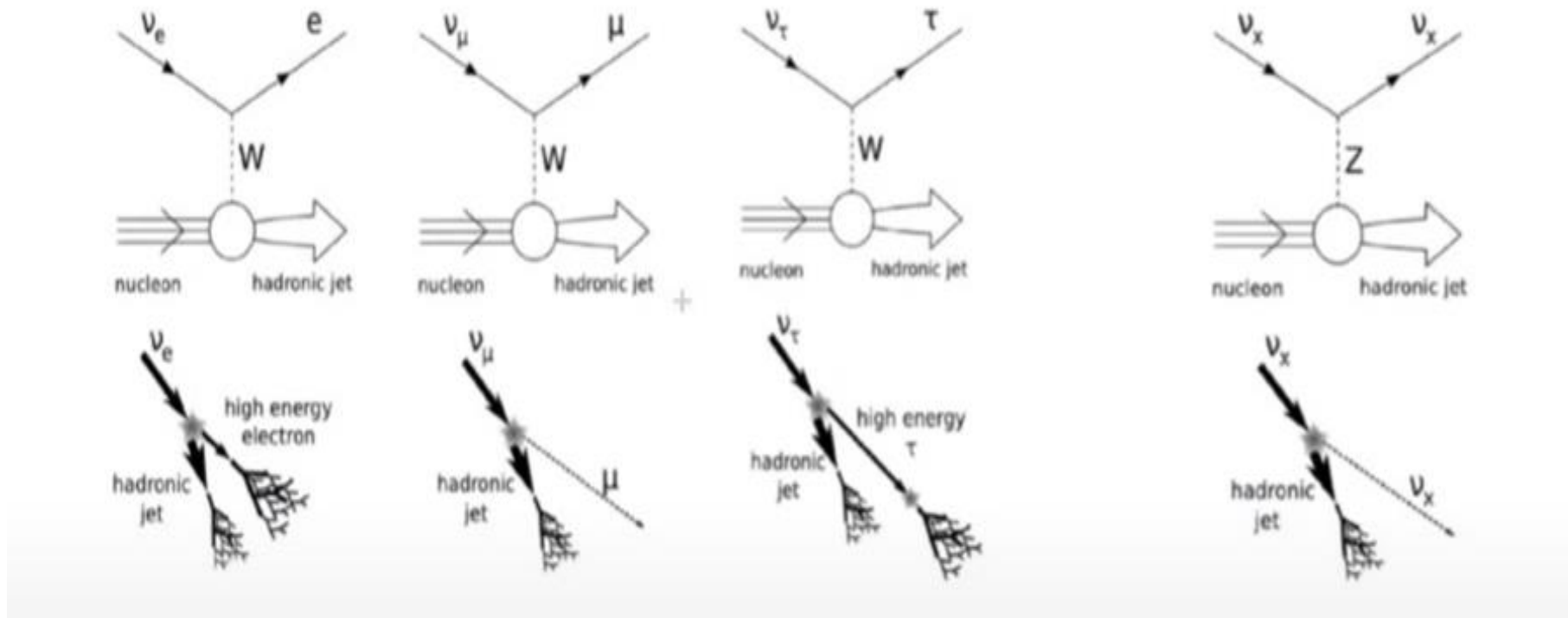
Distribution by the number of triggered detectors

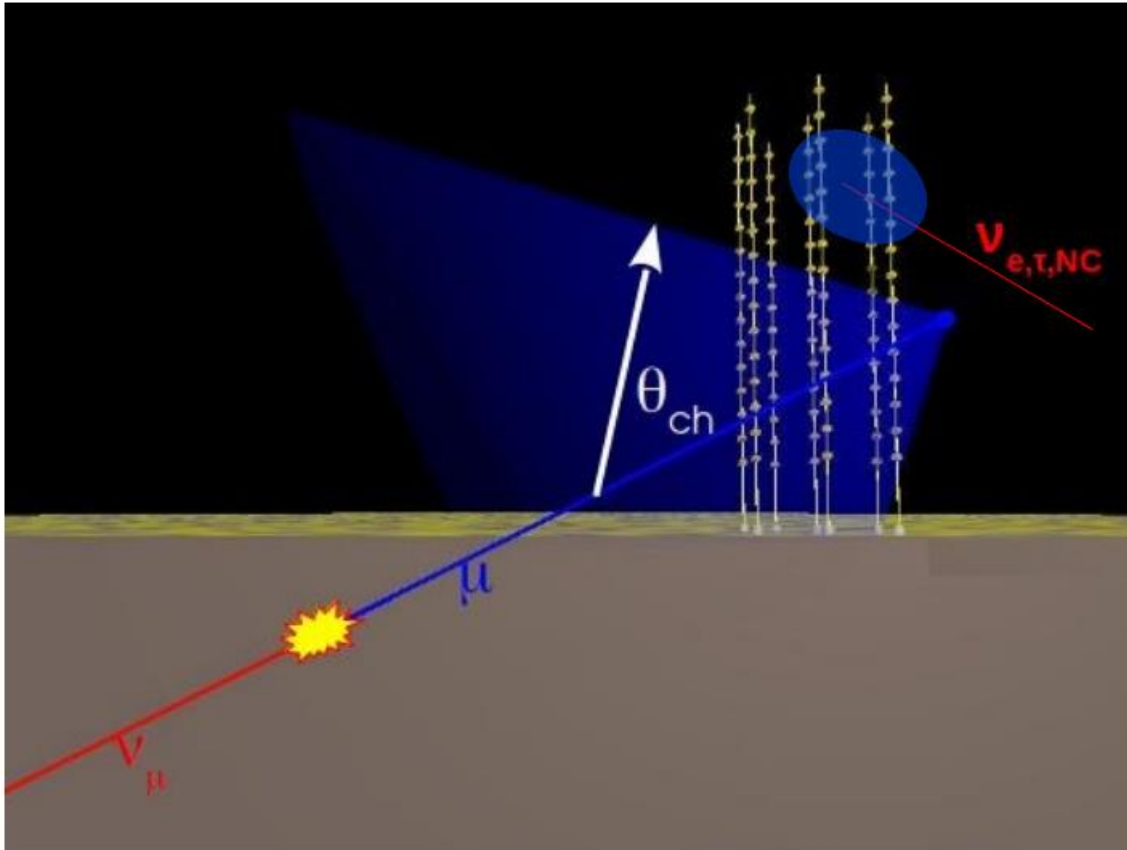


Neutrino classification



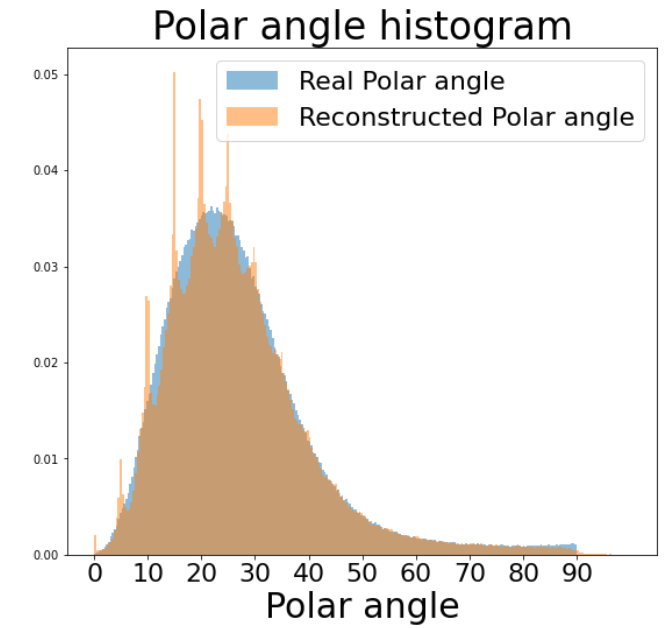
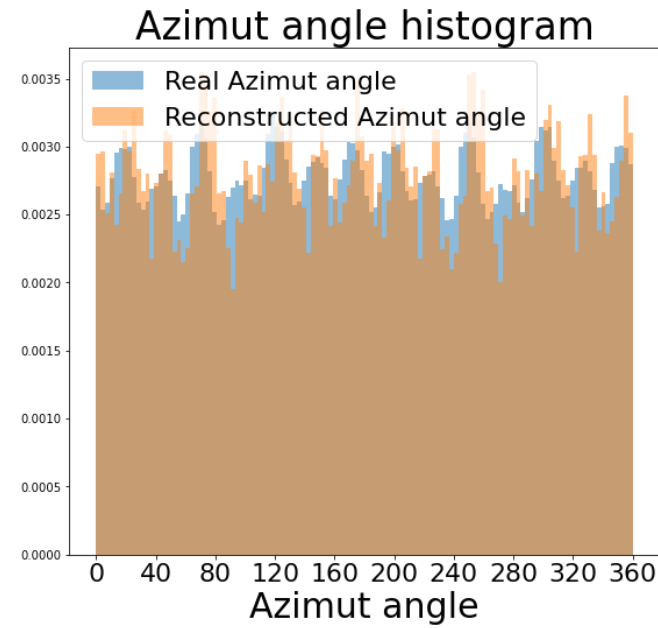
Interaction of neutrinos with matter



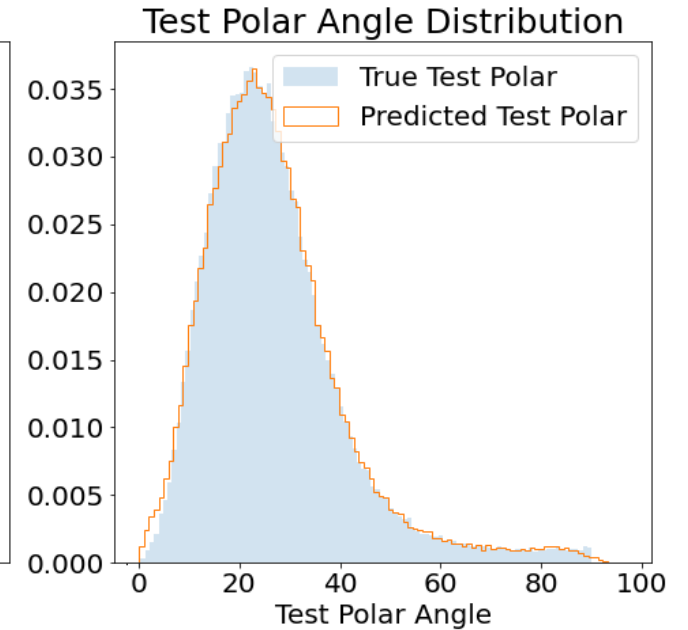
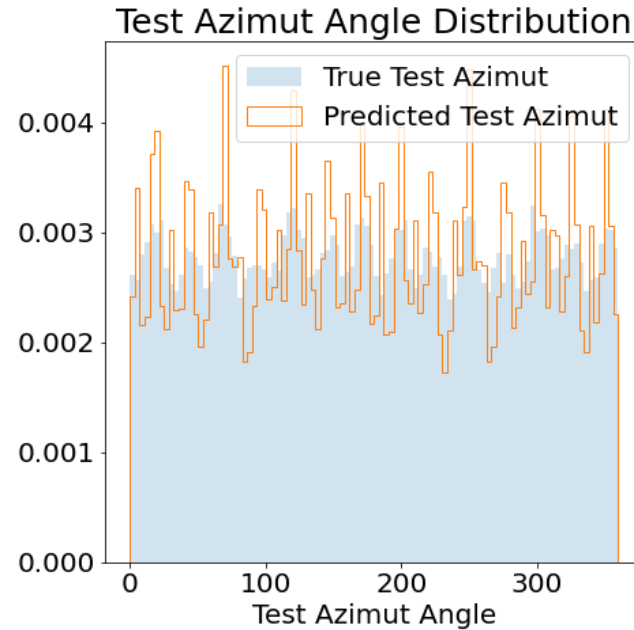


- Large arrays of PMTs in water or ice
- Cherenkov light detected by PMTs
- “Tracks”: ν_μ CC
- “Cascades”: ν_e & ν_τ CC + NC
- Direction reconstructed from hit positions and times
- Energy reconstructed from hit charges

Standard Reconstruction



Our Reconstruction



$$\lambda_{scatterina}^{eff} \approx 480\text{м при } 475 \text{ nm} \quad \lambda_{absorption}^{max} \approx 24\text{м}$$

Tracking events: arrival angle accuracy $\approx 0, 25^\circ$

Cascade events: resolution $\approx 2^\circ$

Technical data

Monte Carlo:

Interaction of neutrinos with nuclei: CTEQ4M
(neutrinos with energies 10 GeV - 100 TeV)

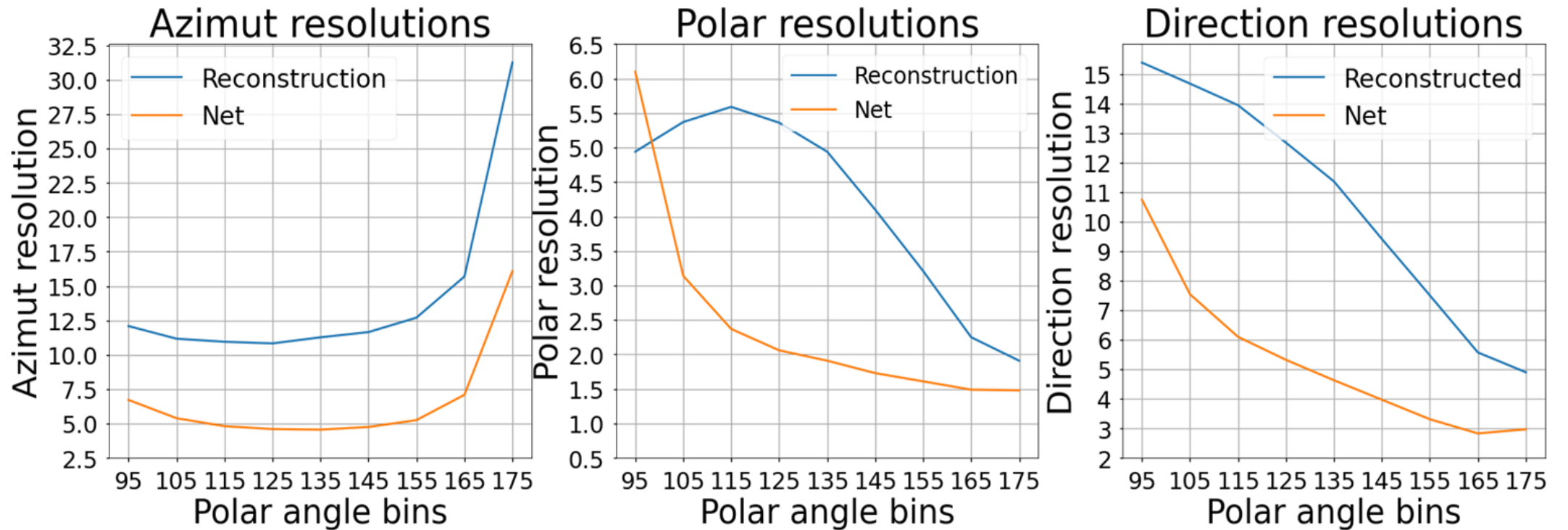
Arrival of muons: CORSIKA 5.7 program on the QGSJET
hadronic impact model

Propagation of muons to Baikal: MUM v1.3u

Cosmic rays: KASCADE based model (240 GeV – 20 PeV)

Time error 5 ns; 30% on charge

Muons: Angular resolutions



Muons: metrics

1D Net			
Metrics	Azimuth Angle	Polar Angle	Direction
50% Resolution	5.34°	1.9°	4.35°
68% Resolution	8.26°	3.09°	6.0°
Standard Reconstruction			
Metrics	Azimuth Angle	Polar Angle	Direction
50% Resolution	12.51°	3.82°	9.11°
68% Resolution	20.73°	6.53°	13.68°