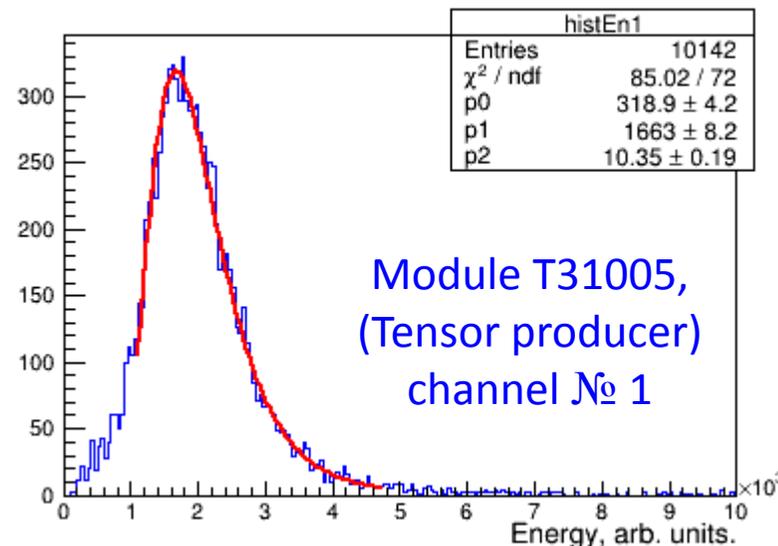
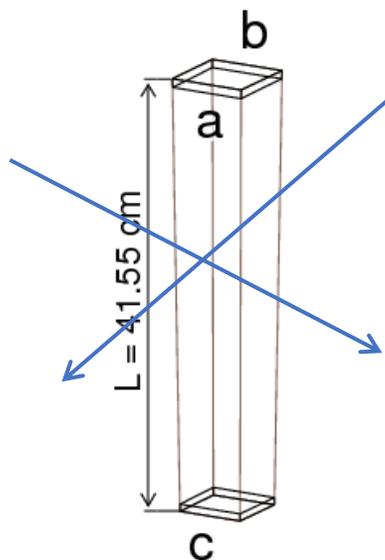


## *Simulation results of two calibration methods for MPD- ECAL*

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M.A. Martemianov, G. E. Fomenko*

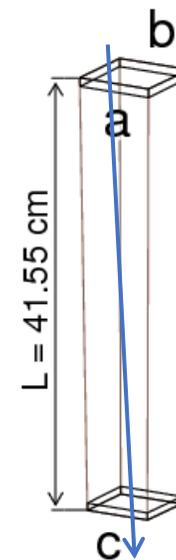
- Currently, the calibration of the entire number of calorimetric modules is being carried out by two different ways. In the first stage, a set of four modules is tested on cosmic muons installed in horizontal position using the same electronics for different modules. A large amount ( $> 1500$ ) of such modules is tested yet.
- In the second stage, one basket (48 modules / 768 towers) with definite electronics is tested in vertical position. These calibration data can be used to align the signals from the towers in the future
- In the final stage, when all baskets will be installed into MPD detector, additional calibration procedure will be necessary
- This report considers the possibility of calibrating the calorimeter after its assembly on cosmic muons based on simulated data
- Simulation based in the frame of mpdroot with standard MPD setup and magnetic field.
- To produce cosmic muons, specific generator (class [MpdCosmicGenerator](#)) was created
- Based on simulated data, two possible calibration methods are considered, as well as algorithms for extracting the peak of deposited energy from cosmic muons in calorimeter towers
- Advantages and disadvantages of these methods are discussed, also dependence of the energy deposition peak from tower position in the ECal setup is studied

## Method 1 : signal from inclined tracks



- ✓ Fast method, applicable for all tower rotation
- ✓ Energy deposition equal around 6 MeV, this value should be slightly depending from tower position
- ✓ Features of the applicability of this method on the full geometry of ECAL can be defined using MC calculations
- ✓ This method is used now for testing produced modules

## Method 2 : signal from longitudinal tracks



- ✓ All muons in tower pass the same length
- ✓ Energy deposition equal around 60 MeV
- ✓ Method doesn't depend from tower position
- ✓ It takes too long to collect the required statistics, possible not be applicable for highly inclined towers

- ✓ Parameterization of the cosmic ray spectra (atmospheric muons) obtained from experimental data and covers wide energy range
- ✓ Assuming a flat Earth is leads to  $\cos^{N-1} \Theta$  dependence, where N – power of the energy distribution

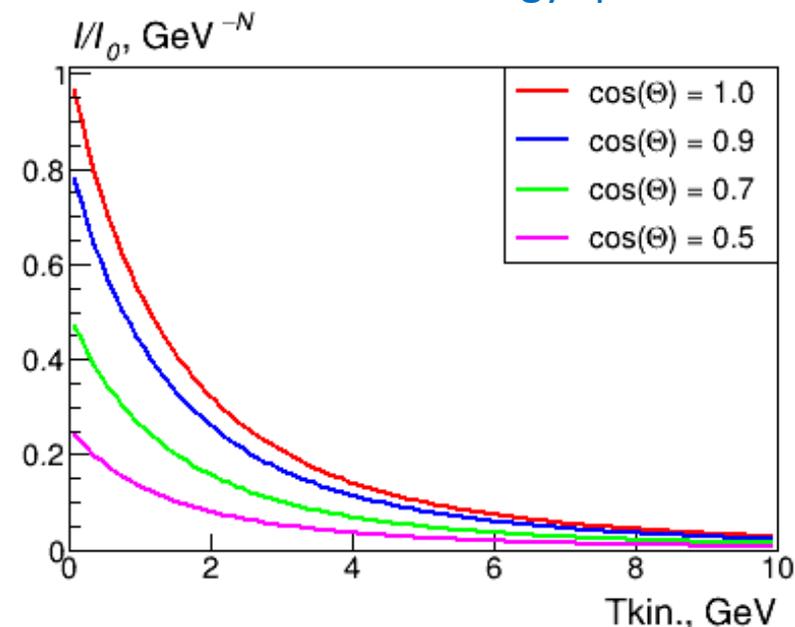
[P. Shukla, Int.J.Mod.Phys. A33 (2018) no.30, 1850175]

$$I / I_0 = (E_0 + E)^{-N} \times \cos^{N-1} \Theta / (1 + E / 854);$$

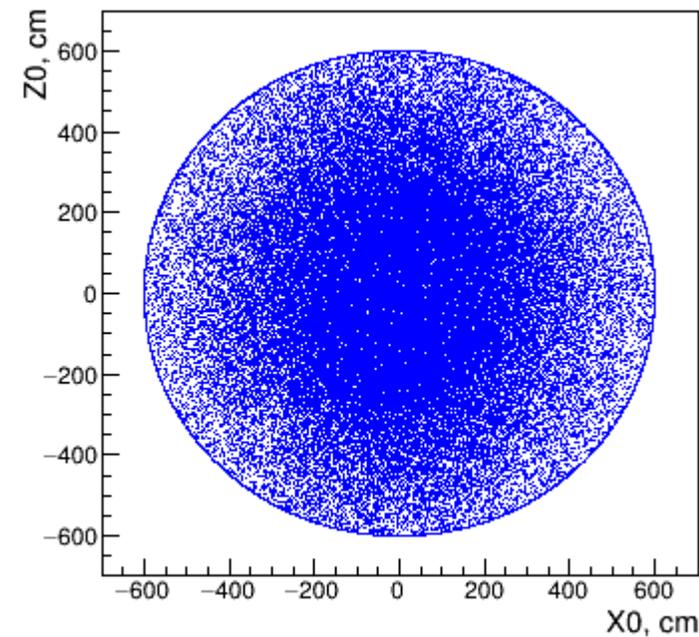
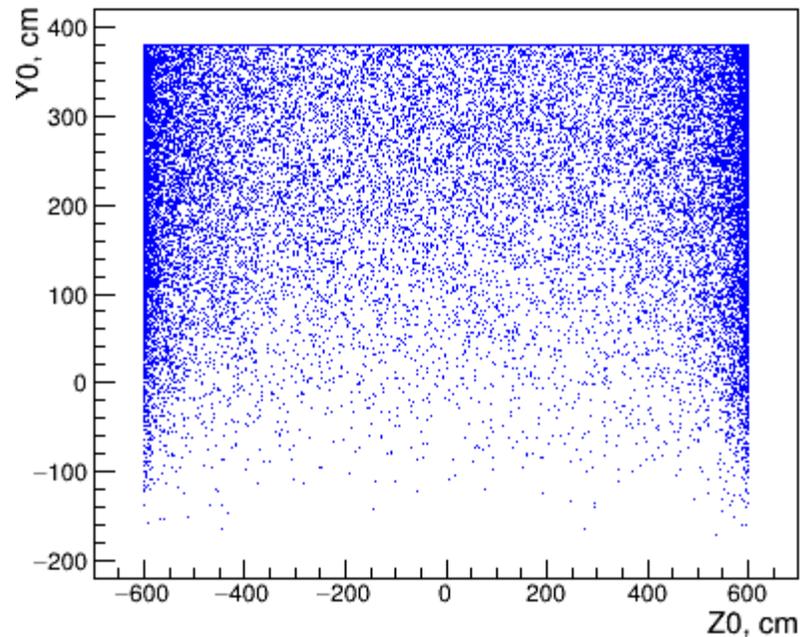
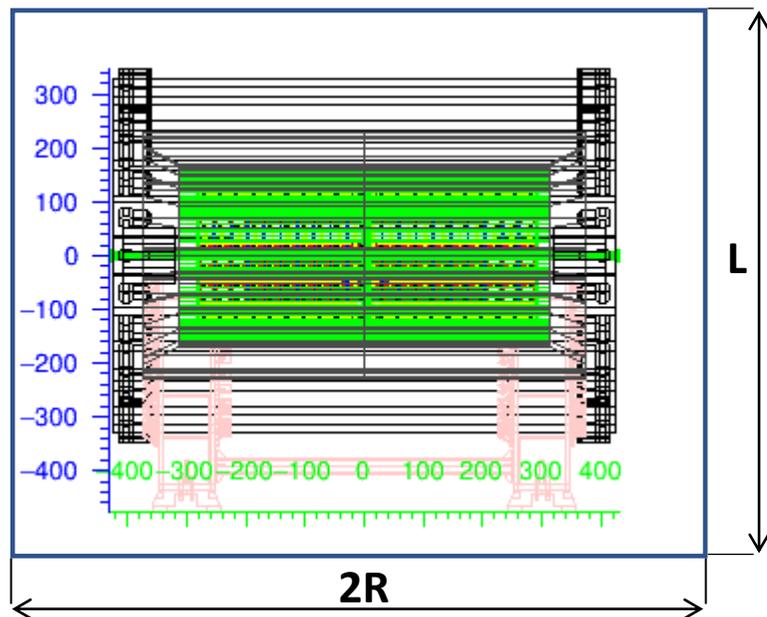
$$E_0 = 4.29, N = 3.01$$

- ✓  $\cos \Theta$  – zenith angle ( $\Theta$  – angle between particle and perpendicular to the ground), E – energy; this formula is valid for muons, not for protons part
- ✓ This physical event generator was implemented in the frame of the mpdroot to produce  $20 \times 10^6$  initial events in the standard root format

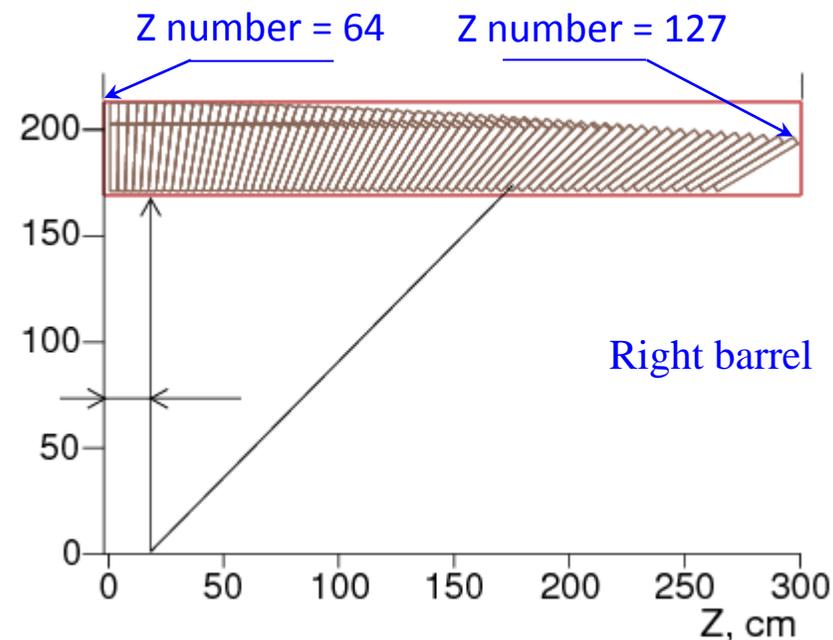
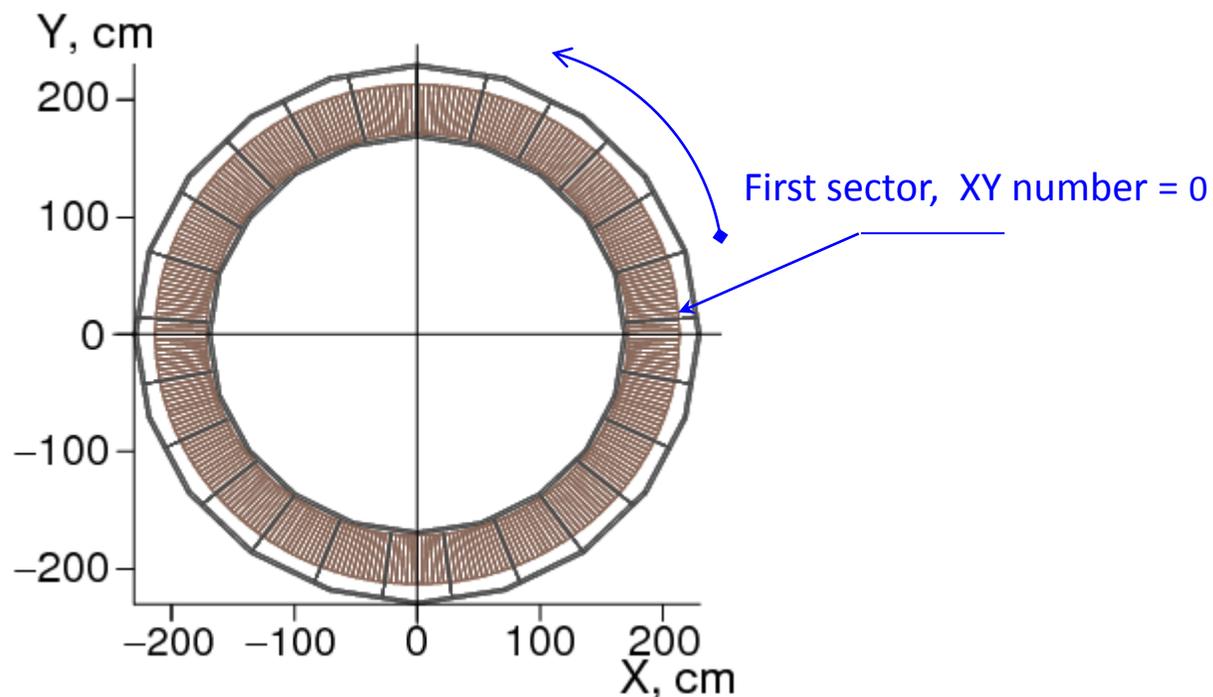
Cosmic muons energy spectra



MPD setup

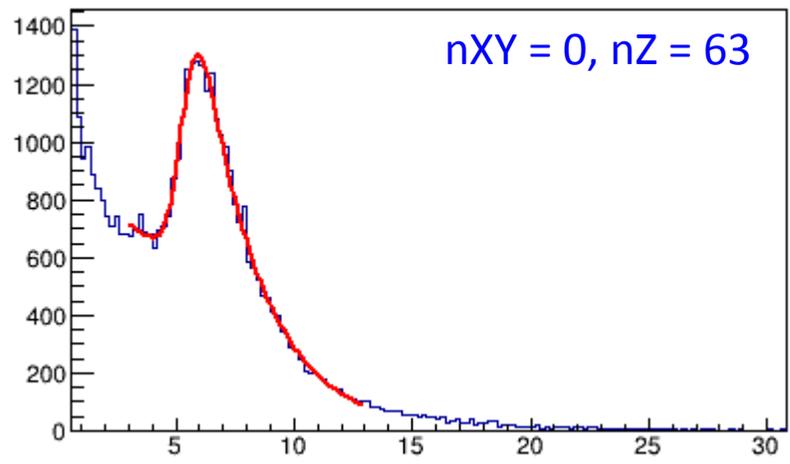


- Initial coordinates are defined by the surface of the cylinder ( $R = 600$  cm,  $L = 760$  cm) covering the experimental setup with some space margins
- To accelerate generation of cosmic tracks inside ECal, only those tracks are accepted for consideration that are directed to the ECal area
- This cut suppress tracks produced in the bottom cylinder region with  $Y_0 < 0$

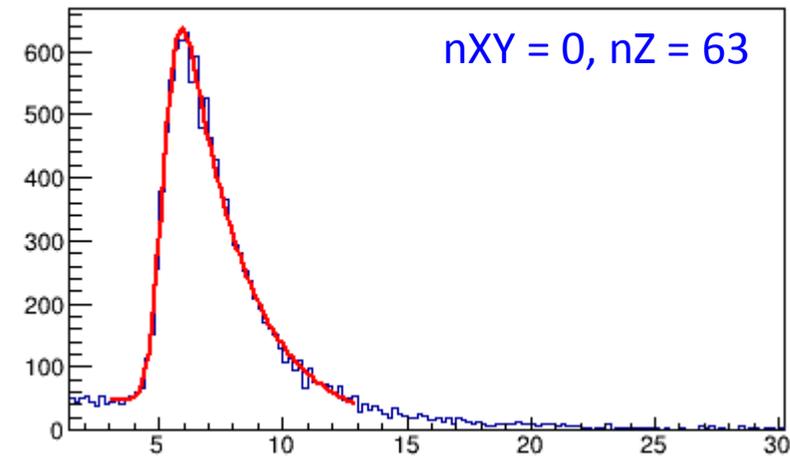


- Total number of towers : 38400
- Number of towers in XY plane (cross section transverse to Z – axis) : 300; step by  $\phi - 1.2$  deg. ( $\phi$  start = 4.2 deg.)
- Number of towers in one line along Z – axis: 128; enumeration is starting from left barrel
- MC data were collected with total ECal geometry. To accelerate calculation signals from left/right barrels and from two adjacent towers in XY (same module) were summarized (number of such towers : 9600)

No selection

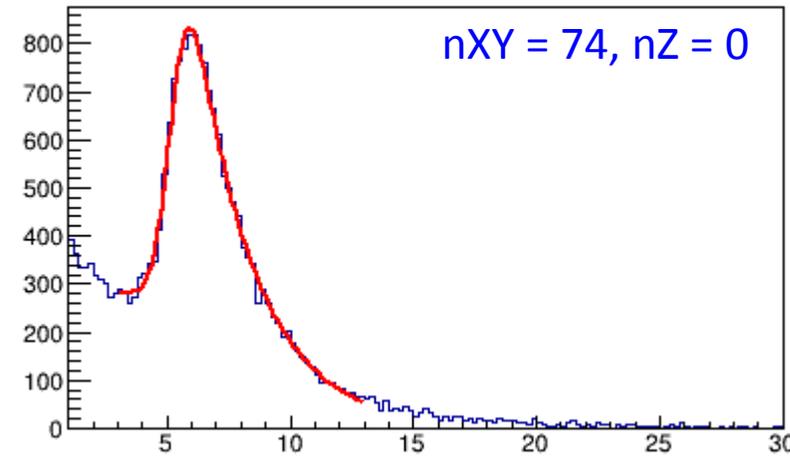
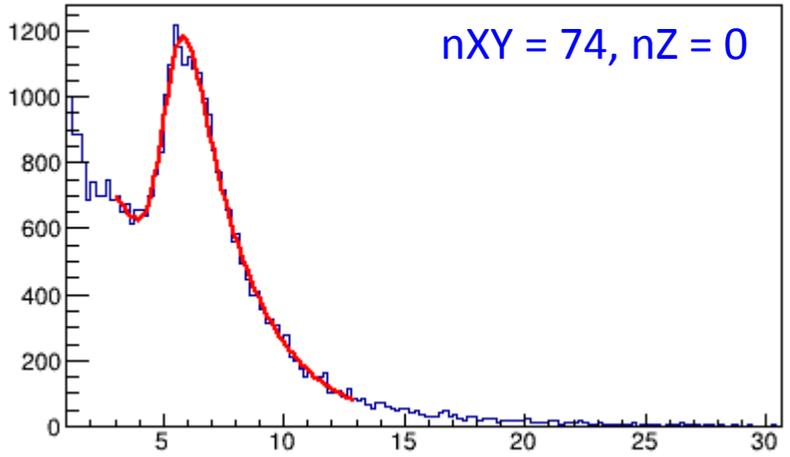


Typical view



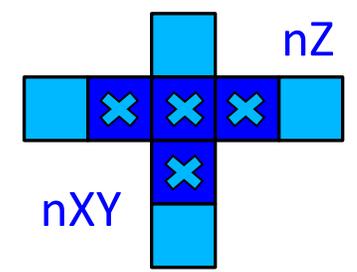
Selection

Border view

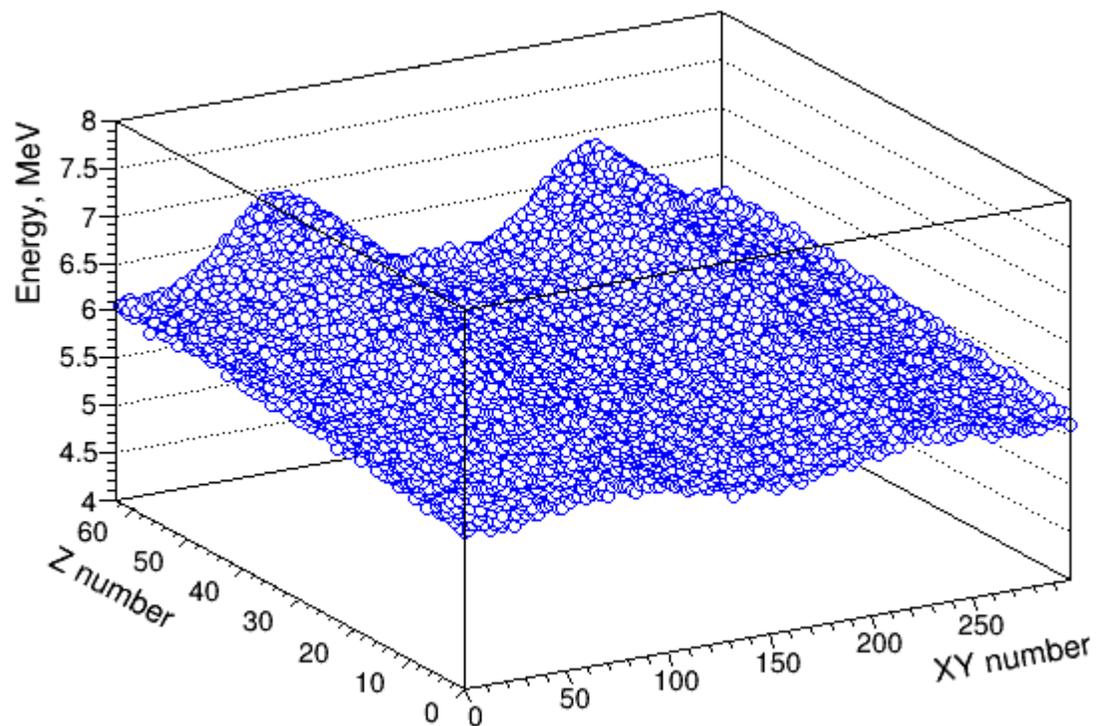


Energy, MeV

✓ Selection : triggered more than three, considered central (XY line and Z line are based on OR logic)

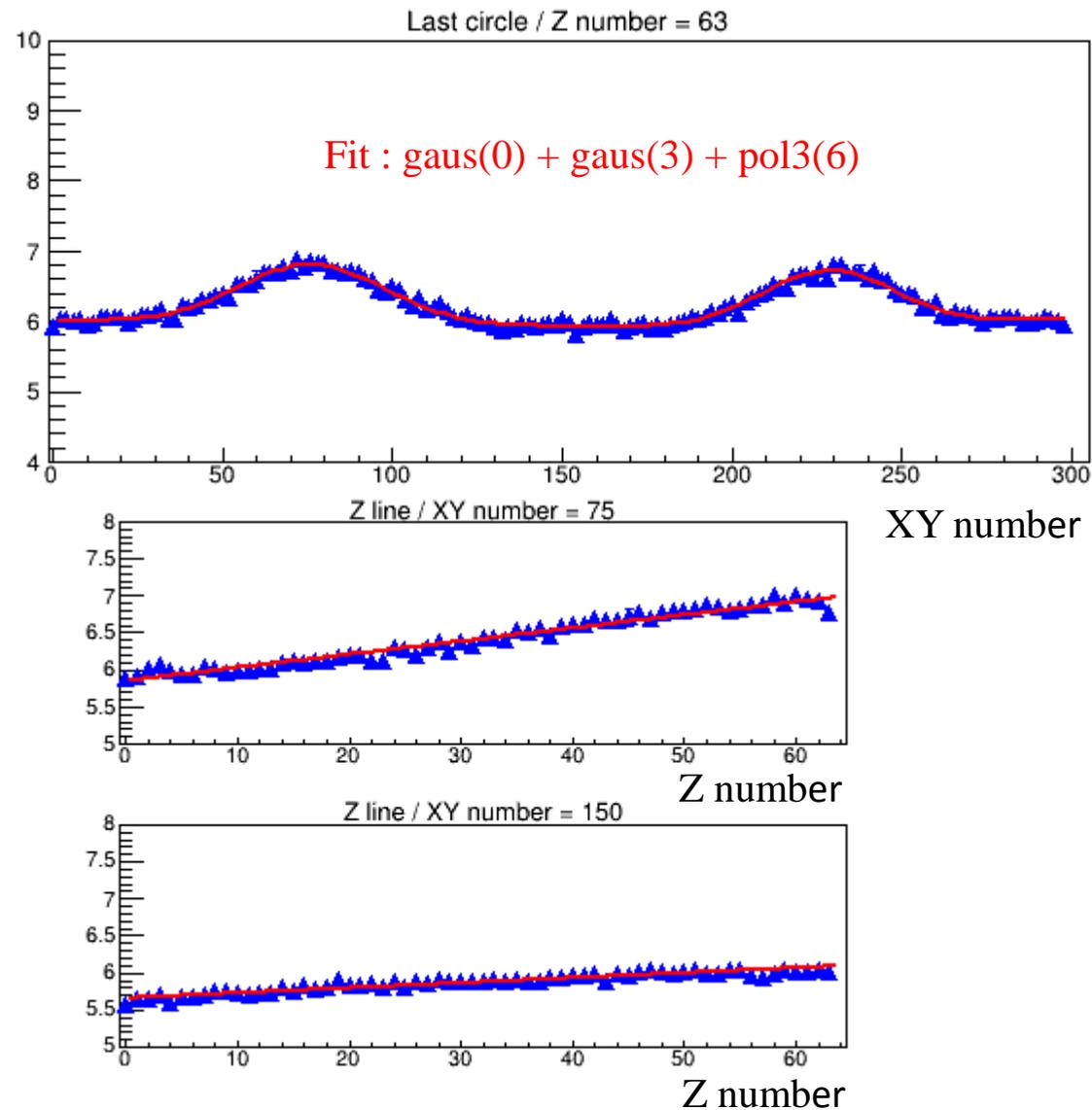


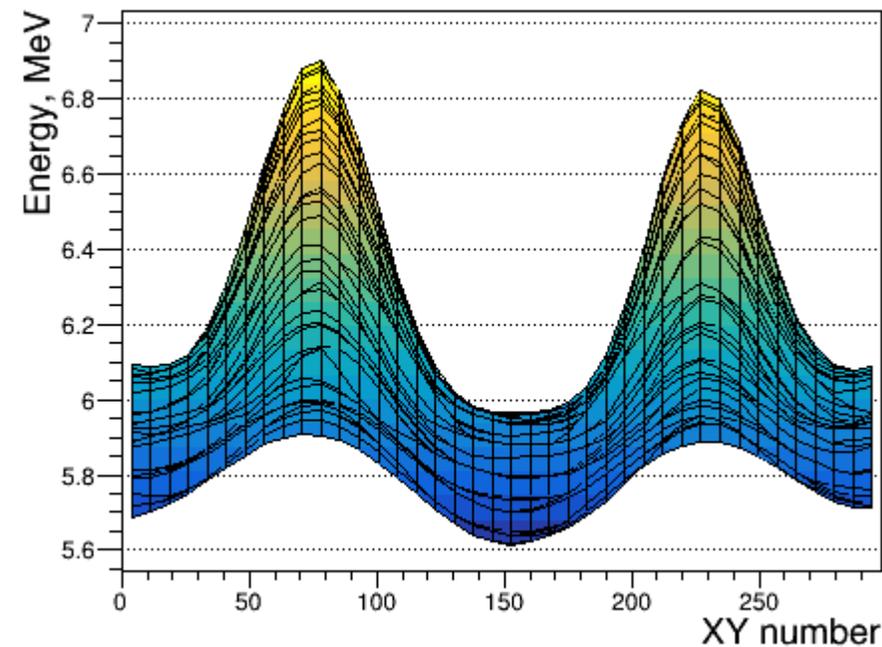
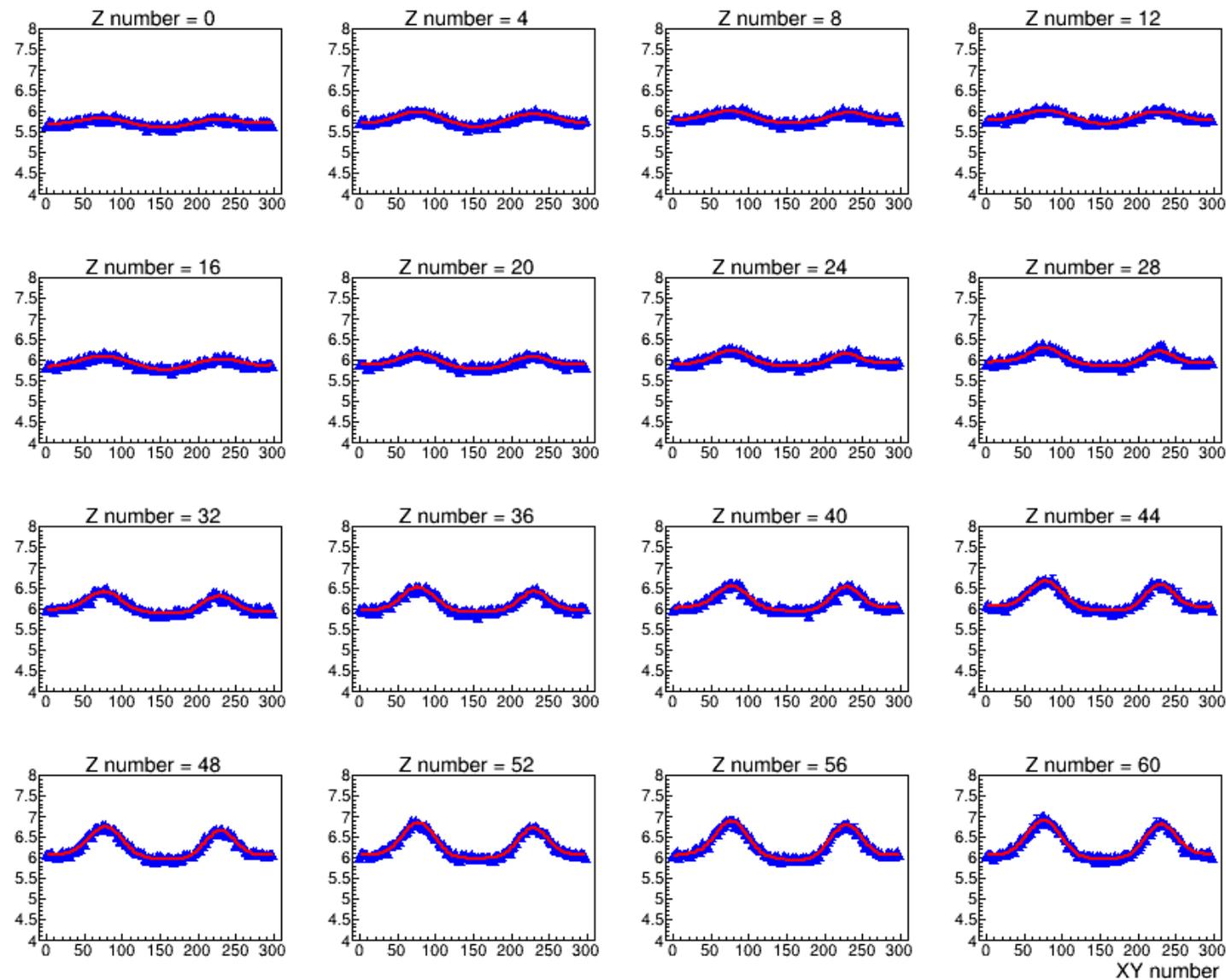
- ✓ It gives a good suppression of the low-energy background
- ✓ Border line of towers shows less background suppression, but the main peak is clearly visible due to the tower incline
- ✓ For all towers: total number of selected events :  $0.9 \times 10^8$



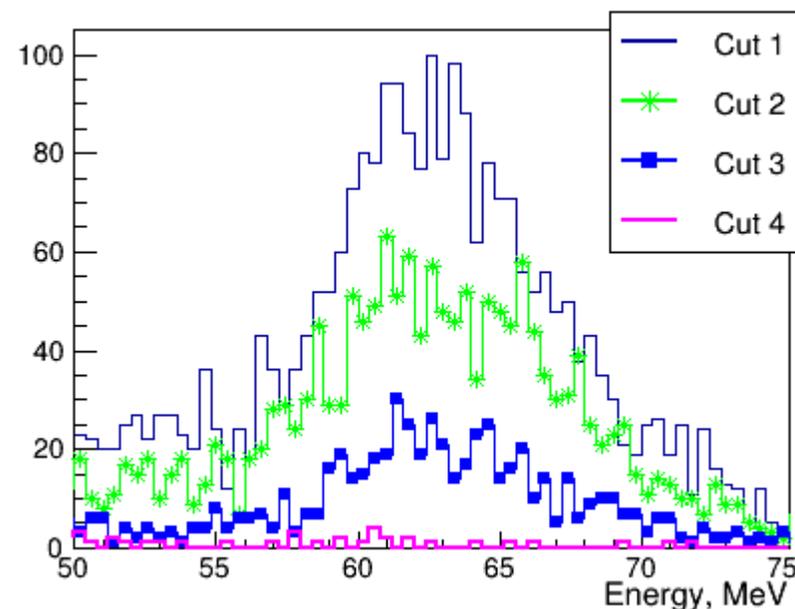
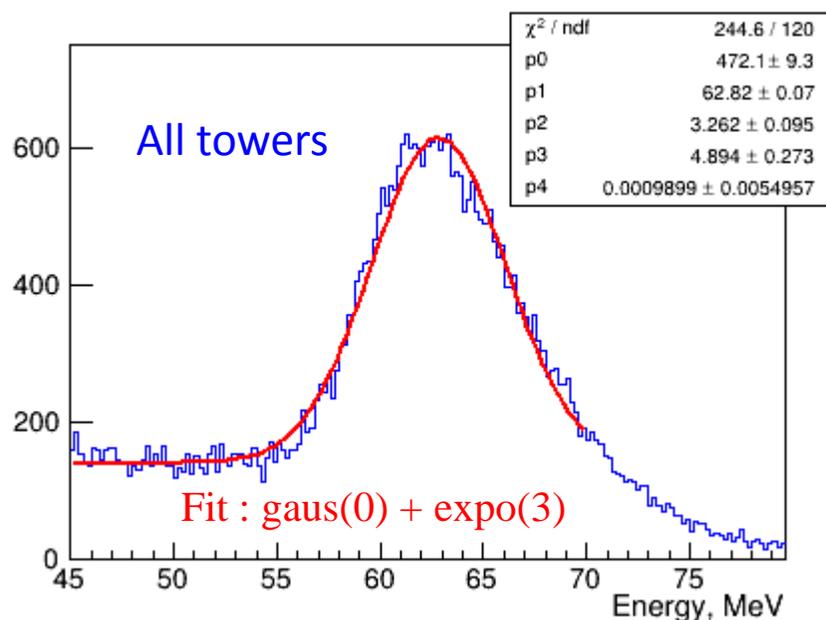
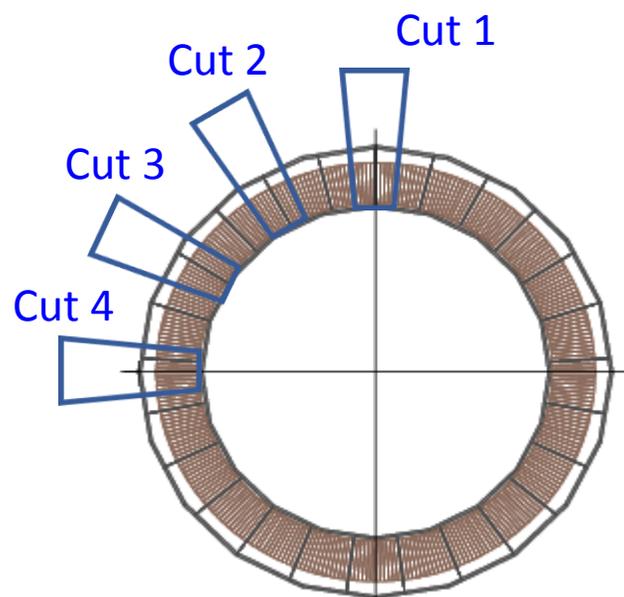
- 3Dim shows mean deposited energy versus tower position
- Dependence is not flat and can be fitted by a set of functions
- At fixed Z position every slice is described by :

2 Gaussians + 3rd degree polynomial function

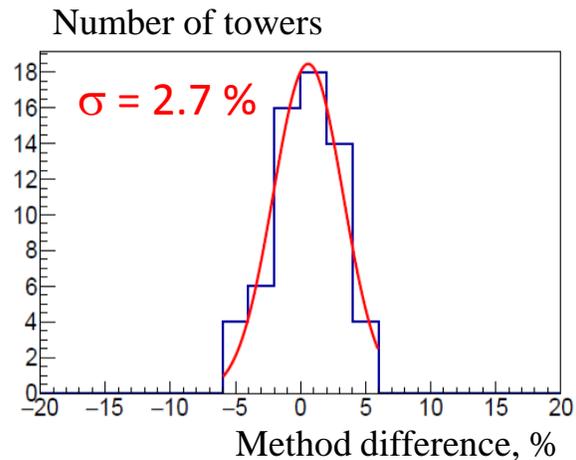
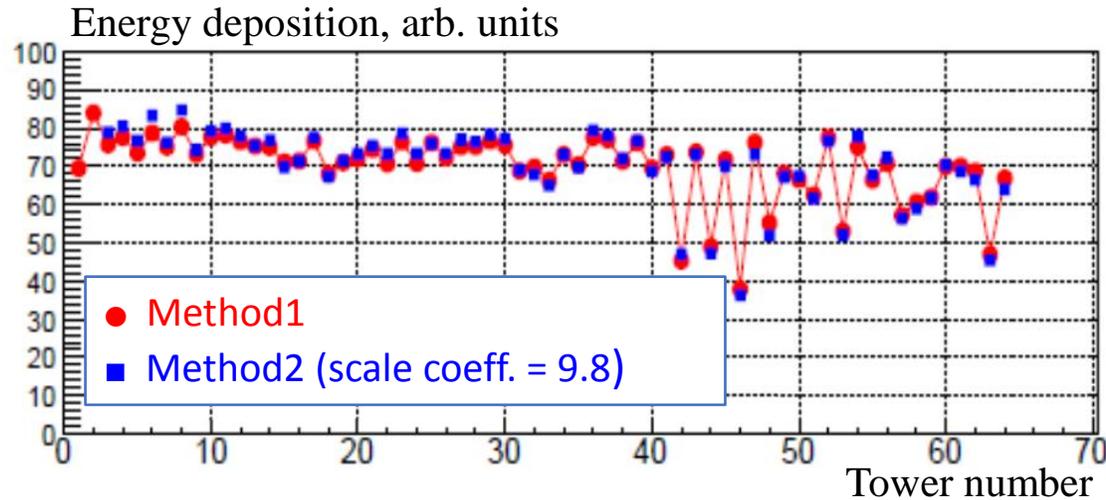




- Fitting procedure were done in XY plane for each Z – number
- Final result – array of the resulting functions with size equal to 64
- 3Dim profile of array looks smooth



- The central tower is triggered, there is no signal in the neighboring towers
- In this case, an energy deposition should be the same for all towers in the complete ECal assembly
- Tower signal is taking into account in that case the deposited energy  $> 1$  MeV
- For all towers: total number of selected events (sum under energy deposition peak)  $\sim 20 \times 10^3$
- Cut 1 – Cut 4 means different limit on iXY number and all iZ numbers (width equals to 10 line along iZ)
- Signal is absolutely suppressed for Cut 4, that corresponds to the horizontal component of cosmic muons



- ✓ Data were collected on cosmic data measurements during a long time data taking
- ✓ Experimental setup was represented by one cluster (one line : 2×64 towers) located horizontally
- ✓ Plots shows data from first 4 modules (64 towers) in the beginning of cluster
- ✓ Two peak were found and fitted by the same functions
- ✓ Positions of these two peaks show the same behavior and can be compared using definite scale factor
- ✓ Methods are in good agreement with accuracy of 2.7 %

- Two methods of the ECal calibration were tested on MC data. For this task, cosmic generator were done to collect data in the frame of the mpdroot. Total number of the generated events is  $20 \times 10^6$
- For inclined muon tracks, where energy deposition energy is  $\sim 6$  MeV, simple selection was implemented to obtain a good view of the energy deposition distribution. This method depends from tower position in the total ECal setup. Peak location as a function of the tower position is defined and described by a set of the fitted functions
- Next method, based on the registration of the longitudinal tracks, was also tested. Selection by the absence of signals in neighboring towers made it possible to identify such a peak with an average energy release in the region  $\sim 60$  MeV
- Both methods are applicable for vertical position of the towers. For inclined towers, especially at angles around  $90^\circ$ , second method cannot be applied



Thank you for attention