CERN-EP-2023-130

Search for Light Dark Matter with NA64 at CERN

Yu. M. Andreev[®],¹ D. Banerjee[®],² B. Banto Oberhauser[®],³ J. Bernhard[®],² P. Bisio[®],^{4,5} A. Celentano[®],⁴

N. Charitonidis[®],² A. G. Chumakov[®],¹ D. Cooke,⁶ P. Crivelli[®],³ E. Depero[®],³ A. V. Dermenev[®],¹

S. V. Donskov⁰,¹ R. R. Dusaev⁰,¹ T. Enik⁰,⁷ V. N. Frolov,⁷ R. B. Galleguillos Silva,^{8,9} A. Gardikiotis⁰,¹⁰

S. V. Gertsenberger[®],⁷ S. Girod,² S. N. Gninenko[®],^{1,*} M. Hösgen,¹¹ V. A. Kachanov[®],¹ Y. Kambar[®],⁷

A. E. Karneyeu[®],¹ E. A. Kasianova,⁷ G. D. Kekelidze[®],⁷ B. Ketzer[®],¹¹ D. V. Kirpichnikov[®],¹

M. M. Kirsanov[®],¹ V. N. Kolosov,¹ V. A. Kramarenko[®],^{1,7} L. V. Kravchuk[®],¹ N. V. Krasnikov[®],^{1,7}

S. V. Kuleshov[®],^{8,9} V. E. Lyubovitskij[®],^{1,12,9} V. Lysan[®],⁷ A. Marini[®],⁴ L. Marsicano[®],⁴ V. A. Matveev[®],⁷

R. Mena Fredes,^{9,12} R. G. Mena Yanssen,^{9,12} L. Molina Bueno⁽¹⁾,¹³ M. Mongillo⁽⁰⁾,³ D. V. Peshekhonov⁽⁰⁾,⁷

V. A. Polyakov[®],¹ B. Radics[®],¹⁴ K. M. Salamatin[®],⁷ V. D. Samoylenko,¹ H. Sieber[®],³ D. A. Shchukin[®],¹ O. Soto,^{15,9} V. O. Tikhomirov[®],¹ I. V. Tlisova[®],¹ A. N. Toropin[®],¹ M. Tuzi[®],¹³ B. I. Vasilishin,¹

P. V. Volkov[®],⁷ V. Yu. Volkov[®],¹ I. V. Voronchikhin[®],¹ J. Zamora-Saá[®],^{8,9} and A. S. Zhevlakov[®]⁷

(The NA64 Collaboration)

¹Authors affiliated with an institute covered by a cooperation agreement with CERN

²CERN, European Organization for Nuclear Research, CH-1211 Geneva, Switzerland

³ETH Zürich, Institute for Particle Physics and Astrophysics, CH-8093 Zürich, Switzerland

⁴INFN, Sezione di Genova, 16147 Genova, Italia

⁵Università degli Studi di Genova, 16126 Genova, Italia

⁶UCL Departement of Physics and Astronomy, University College London,

Gower St. London WC1E 6BT, United Kingdom

⁷Authors affiliated with an international laboratory covered by a cooperation agreement with CERN

⁸Center for Theoretical and Experimental Particle Physics, Facultad de Ciencias Exactas,

Universidad Andres Bello, Fernandez Concha 700, Santiago, Chile

⁹ Millennium Institute for Subatomic Physics at High-Energy Frontier (SAPHIR), Fernandez Concha 700, Santiago, Chile

¹⁰Physics Department, University of Patras, 265 04 Patras, Greece

¹¹ Universität Bonn, Helmholtz-Institut für Strahlen-und Kernphysik, 53115 Bonn, Germany

¹² Universidad Técnica Federico Santa María and CCTVal, 2390123 Valparaíso, Chile

¹³Instituto de Fisica Corpuscular (CSIC/UV), Carrer del Catedratic Jose Beltran Martinez, 2, 46980 Paterna, Valencia, Spain ¹⁴York University, Toronto, Canada

¹⁵Departamento de Fisica, Facultad de Ciencias,

Universidad de La Serena, Avenida Cisternas 1200, La Serena, Chile

(Dated: July 6, 2023)

Thermal dark matter models with particle χ masses below the electroweak scale can provide an explanation for the observed relic dark matter density. This would imply the existence of a new feeble interaction between the dark and ordinary matter. We report on a new search for the sub-GeV χ production through the interaction mediated by a new vector boson, called the dark photon A', in collisions of 100 GeV electrons with the active target of the NA64 experiment at the CERN SPS. With 9.37×10^{11} electrons on target collected during 2016-2022 runs NA64 probes for the first time the well-motivated region of parameter space of benchmark thermal scalar and fermionic dark matter models. No evidence for dark matter production has been found. This allows us to set the most sensitive limits on the A' couplings to photons for masses $m_{A'} \lesssim 0.35$ GeV, and to exclude scalar and Majorana dark matter with the $\chi - A'$ coupling $\alpha_D \leq 0.1$ for masses $0.001 \lesssim m_\chi \lesssim 0.1$ GeV and $3m_{\chi} \leq m_{A'}$.

Thermal light dark matter (LDM) with DM particles (χ) masses below the electroweak scale, $m_{\chi} \ll 100$ GeV, is one of the most popular candidates to explain the origin of DM. In this model, dark and ordinary matter were initially in thermal equilibrium and annihilate to each other at equal rates. However, as the Universe was expanding and cooling down, the annihilation rates fell out of equilibrium and the DM number density froze out at the value when equilibrium was lost. In this scenario, the existence of a new interaction between the χ and the Standard Model (SM) is required to accomodate the

^{*} Corresponding author; sergei.gninenko@cern.ch

relic DM density. It is remarkable that the strength of this interaction is such, that it allows one to search for its manifestations at present accelerators [1, 2].

In one of the most interesting cases, this interaction could be transmitted by a new vector boson, called dark photon (A'). The A' could have a mass in the sub-GeV mass range, and couple to the SM via kinetic mixing with the SM photon field described by the term $\frac{\epsilon}{2}F'_{\mu\nu}F^{\mu\nu}$, here $F'_{\mu\nu} = \partial_{\mu}A'_{\nu} - \partial_{\nu}A'_{\mu}$, where ϵ is the mixing strength [3–6]. The massive A'_{μ} field, associated with the spontaneously broken $U_D(1)$ gauge group, has the dark coupling strength e_D of the $U_D(1)$ gauge interactions to χ 's, which is given by $\mathcal{L}_{int} = -e_D A'_{\mu} J^{\mu}_D$, where J_D is the dark matter current. The mixing term results in the interaction $\mathcal{L}_{int} = \epsilon e A'_{\mu} J^{\mu}_{em}$ of the A' with the electromagnetic (e-m) current J_{em}^{μ} with a strength ϵe , where e is the e-m coupling and $\epsilon \ll 1$ [7–9]. Small values of ϵ can be obtained in grand unified theories from loop effects of particles having both the dark $U_D(1)$ and SM U(1) charges with a typical one-loop value $\epsilon = ee_D/16\pi^2 \simeq 10^{-2} - 10^{-4}$ or from two-loop contributions resulting in $\epsilon \simeq 10^{-3} - 10^{-5}$ [9].

If the A' is the lightest state in the dark sector, then it would decay mostly to a visible state, i.e. to SM leptons or hadrons [10–18]. However, if the decay $A' \to \chi \chi$ is kinematically allowed, the A' would dominantly decay invisibly into χ 's provided $m_{\chi} < m_{A'}/2$ and $e_D > \epsilon e$. Various LDM models motivate the existence of sub-GeV χ 's which could be either scalar, Majorana, or pseudo-Dirac particles coupled to the A' [19–25]. Models introducing the invisible A', i.e. that invisible decay mode is predominant, $\Gamma(A' \to \bar{\chi}\chi)/\Gamma_{tot} \simeq 1$, are subject to various experimental probes leaving, however, a large parameter area still to be explored.

Imposing the thermal freeze-out condition of DM annihilation into visible sector through $\gamma - A'$ mixing allows one to predict values of the parameter

$$y = \alpha_D \epsilon^2 \left(\frac{m_\chi}{m_{A'}}\right)^4 \tag{1}$$

which defines the annihilation cross section and hence the relic DM density, and also relates the dark coupling $\alpha_D = e_D^2/4\pi$ and mixing ϵ by

$$\alpha_D \simeq 0.02 f \left(\frac{10^{-3}}{\epsilon}\right)^2 \left(\frac{m_{A'}}{100 \text{ MeV}}\right)^4 \left(\frac{10 \text{ MeV}}{m_{\chi}}\right)^2 \quad (2)$$

where the parameter f depends on $m_{A'}$ and m_{χ} [1]. For $\frac{m_{A'}}{m_{\chi}} = 3, f \lesssim 10$ for a scalar [21], and $f \lesssim 1$ for a fermion case [22]. The accessibility of the predicted y and α_D values in direct searches and at accelerator experiments provide an important complementarity between different techniques[26]. Thus, motivating a worldwide effort towards dark forces and other portals between the visible and dark sectors; see, e.g. Refs. [21, 27–50]. In particular, the predicted (ϵ , y, $m_{A'}$, m_{χ}) parameter space which can be probed at the beam energies of the CERN SPS.

In this Letter, we report new results on the search for the invisible A' mediator and light-dark matter in the fixed-target experiment NA64 at the CERN SPS [55, 56], obtained from the combined statistics of 2016-2022 runs. The method we chose for the search was proposed in Refs. [57, 58]. It is based on the detection of the missing energy, arising from prompt decays $A' \to \chi \chi$ of the hard bremsstrahlung A' produced in the process $e^-Z \rightarrow e^-ZA'$ of high-energy electrons scattering in the active beam dump target. Another A' production mechanism considered in this work is through the resonant annihilation of secondary positrons from the e-m shower developed in the target with its atomic electrons, $e^+e^- \to A' \to \chi\chi$ [59]. Due to the resonant cross-section enhancement, it provides a strong increase to the signal yield, and thus to the sensitivity of the search, in particular for the high-mass region, $m_{A'} \gtrsim m_{\mu}$ [60]. The advantage of the NA64 approach compared to the classical beam dump ones [21, 22, 40, 43, 61, 62] is that its sensitivity is proportional to ϵ^2 , associated with the A' production and its subsequent prompt invisible decay, while in the latter case, it is proportional to $\epsilon^4 \alpha_D$, with ϵ^2 associated with the A' production in the dump and $\epsilon^2 \alpha_D$ coming from the χ particle scattering in the far detector [57].

The NA64 detector upgraded for a more sensitive LDM search in 2021-2022 runs is schematically shown in Fig. 1. It employs the optimized H4 100 GeV electron beam at CERN SPS, which has a maximal intensity $\simeq 10^7$ electrons per SPS spill of 4.8 s [63]. The beam is defined by the scintillator (Sc) counters S_{1-3} and a veto counter V_1 . A magnetic spectrometer is used to reconstruct the momentum of the incoming e^{-} 's with the precision $\delta p/p \simeq 1\%$ [64]. The spectrometer consists of two consecutive dipole magnets $MBPL_{1,2}$ with the total magnetic field of $\simeq 7$ T·m and a low-material-budget tracker composed of a set of two upstream Micromegas $(MM_{1,2})$ and two straw-tube chambers $(ST_{1,2})$ and two downstream MM_{3-4} , $ST_{3,4}$, and $GEM_{1,2}$ stations. The synchrotron radiation (SR) emitted in the MBPL magnetic field and detected with an SR detector (SRD) was used for electron identification. The SRD is an array of a PbSc sandwich calorimeter of a fine segmentation [57, 65]. With this technique the initial admixture of the hadron contamination in the beam $\pi/e^- \lesssim 10^{-2}$ was further suppressed to $\simeq 2 \times 10^{-5}$ [66]. Downstream the setup was equipped with an active dump target, an e-m calorimeter (ECAL), for measurement of the recoil electron energy E_{ECAL} and the transverse and longitudinal shape of the corresponding e-m shower. The ECAL was a matrix of 5×6 Shashlik-type modules assembled from Pb and Sc plates of $\simeq 40$ radiation lengths (X_0) , with the first $4X_0$ serving as a preshower detector. Finally, the ECAL followed by a high-efficiency veto counter VETO, and a massive, hermetic hadronic calorimeter (HCAL) of three modules $HCAL_{1-3}$, $\simeq 30$ nuclear interaction lengths in total to veto muons or hadronic secondaries produced in the e^- nuclei interactions in the target. Zero-degree



FIG. 1. Schematic illustration of the setup to search for $A' \rightarrow invisible$ decays of the bremsstrahlung A's produced in the reaction $eZ \rightarrow eZA'$ of 100 GeV e^- incident on the active ECAL target in 2021-2022 runs.

 HCAL_4 was used to reject beam electrons accompanied by neural secondaries.



FIG. 2. The measured distribution of events in the $(E_{ECAL}; E_{HCAL})$ plane after applying all selection criteria. The shaded area is the signal box, with the size along the E_{HCAL} axis increased for illustration purposes. The side bands A and C are the ones used for the background estimate inside the signal region.

Our data were collected in several runs, during two periods with the trigger requiring the ECAL energy $E_{ECAL} \lesssim 90$ GeV. The first period had 2.83×10^{11} electrons on target (EOT) accumulated during 2016-2018 runs ((hereafter called respectively runs I-III) [47, 48]. The second, with 2021 run (run IV) [67] and 2022 run (run V), had 6.54×10^{11} EOT collected with the beam intensity in the range $\simeq (5-7) \times 10^6$ e⁻ per spill. Data with a total of 9.37×10^{11} EOT from these five runs were processed with selection criteria and combined as described below.

A Geant4 [68, 69] based Monte Carlo (MC) simulation package DMG4 [70] is used to study the performance of the detector, signal acceptance, and background level, as well as the analysis procedure including selection of cuts and sensitivity estimate. To maximize the signal accep-



FIG. 3. The NA64 90% C.L. exclusion region in the $(m_{A'}, \epsilon)$ plane. Constraints from the E787 and E949 [41, 42], *BABAR* [49] and NA62 [50] experiments, from the consideration of the anomalous magnetic moment of electron α_e [51–54], as well as the favored area explaining the α_{μ} anomaly with the A' contribution [10] are also shown. For more limits from indirect searches and planned measurements; see, e.g., Refs. [28–30].

tance and to minimize background, the following selection criteria were used: (i) The incoming track should have the momentum 100 ± 10 GeV. (ii) The track angle with respect to the deflected beam axis should be within 3 mrad to reject large angle events from the upstream e^- interactions. (iii) The detected SR energy should be

within the range $\simeq 1 - 100$ MeV emitted by e^{-s} and in time with the trigger. (iv) The longitudinal and lateral shape of the ECAL cluster should be consistent with the one expected for the signal event [71]. (v) There should be no multiple hits in the $ST_{3,4}$ chambers and no activity in VETO. This was an effective cut against the electroproduction of charged secondaries in the upstream beam material. The measured distribution of $\simeq 9.6 \times 10^5$ events in the $(E_{ECAL}; E_{HCAL})$ plane that passed these criteria from combined runs IV and V is shown in Fig. 2. The candidate events were requested to have the missing energy $E_{miss} \gtrsim 50$ GeV. The optimal signal box was defined based on the energy spectrum calculations for A's emitted by e^{\pm} from the e-m shower generated by the primary e^{-s} in the target [71, 72] and the expected background level for each particular run. Events originated from the rare QED dimuon production in the target were used as a benchmark reaction allowing us to verify the reliability of the MC simulation, correct the signal acceptance vs the A' energy, cross-check systematic uncertainties and background estimate [47, 48]. The A'acceptance was evaluated from simulations and e^- data sample taking into account the selection efficiency for the longitudinal and transverse e-m shower shape in the target from signal events [71]. The energy corrections were extracted using the measured spectra of recoil e^- in the ECAL from dimuon events. The A' production cross section for the bremsstrahlung reaction was obtained with exact tree-level calculations as described in Ref. [72] with uncertainty $\simeq 10\%$ [47, 48], while the A' yield from the resonant process was calculated as described in Ref. [60].

TABLE I. Expected background for 2021-2022 runs

Background source	Background, n_b
(i) dimuons losses or decays in the target	0.04 ± 0.01
(ii) $\mu, \pi, K \to e + \dots$ decays in the beam line	0.3 ± 0.05
(iii) lost γ, n, K^0 from upstream interactions	0.16 ± 0.12
(iv) Punch-through leading n, K_L^0	< 0.01
Total n_b (conservatively)	0.51 ± 0.13

Several processes shown in Table I contribute to background: (i) dimuons losses due to their inefficient detection or decays in the target. It was estimated from the measurements of the single muon detection efficiency and the number of observed dimuons; (ii) decays in flight of mistagged μ, π , K. It was evaluated from the simulations and measurements of the beam composition [66]; (iii) escaping large-angle neutrals from the electroproduction in the beam line due to the insufficient HCAL coverage. Compared to the first period of data taking, this background was significantly reduced in 2021-2022 runs by increasing the HCAL acceptance by moving it $\simeq 3$ m upstream. Additionally, to minimize interactions in the beam line the amount of dead material was further reduced; the last background source is (iv) punch-through of leading neutral hadrons (n, K_L^0) from the e^- interactions in the target. It was evaluated from the direct measurements of punch-through events [73]. After applying the selection cuts, we expected mostly background events of type (iii) to remain in the data. Their number was evaluated from the data itself by the extrapolation of events from the sideband C $(E_{ECAL} > 50 \text{ GeV}; E_{HCAL} < 1 \text{ GeV})$ shown in Fig. 2 into the signal region and assessing the systematic errors by varying the fit functions. The shape of the extrapolation functions was taken from the analysis of the data and cross-checked with simulations of the e^- hadronic interactions in the beam. Finally, the estimated background inside the signal region was 0.51 ± 0.13 events. Compared to the 2016-2018 runs [48], background was further rejected by a factor $\simeq 6$. After determining all the selection criteria and background levels, no event is found in the signal region.

To obtain upper limits on the mixing strength, runs I-V were analyzed simultaneously using the technique of multi-bin limit setting based on the RooStats package [74]. First, the background levels, efficiencies, their corrections and uncertainties were used to optimize the ECAL energy cut for the signal box, by comparing sensitivities, defined as an average expected bound calculated using the profile likelihood ratio method for each run. For this procedure, the expected number of background events for each run, obtained from the data extrapolation to the signal region with errors estimated from the variation of the extrapolation functions, were the most important inputs. For the data samples from the 2016-2021 runs we used the previously optimised value of the ECAL energy cut of 50 GeV [47, 48, 67]. For the long 2022 run, the optimal cut was selected in the range $E_{ECAL} \lesssim 47-50$ GeV, slightly depending on the run conditions and detector performance during the data taking. The overall signal efficiency, given by the product of efficiencies accounting for the geometrical acceptance, the track, SRD, VETO and HCAL signal reconstruction, was in the range 0.4-0.5. The VETO (0.94) and HCAL(0.95)efficiencies were defined as a fraction of events below the zero-energy thresholds, with the loss mostly due to pileup in high-intensity runs.

The combined 90% confidence level (C.L.) upper limits for mixing strength ϵ is obtained by applying the modified frequentist approach for confidence levels, considering the profile likelihood ratio as a test statistic in the asymptotic approximation [75–77]. The number of events in the signal box is the sum of expected events from all five runs:

$$N_{A'} = \sum_{i=1}^{5} N_{A'}^{i} = \sum_{i=1}^{5} n_{EOT}^{i} \epsilon_{A'}^{i} n_{A'}^{i} (\epsilon, m_{A'}, \Delta E_{e}) \quad (3)$$

where $\epsilon_{A'}^i$ is the signal efficiency in run *i*, and $n_{A'}^i(\epsilon, m_{A'}, \Delta E_e)$ is the number of signal events per EOT produced in the energy interval ΔE_e . Signal events for



FIG. 4. The top row shows the NA64 limits in the $(y;m_{\chi})$ plane obtained for $\alpha_D = 0.5$ (left panel) and $\alpha_D = 0.1$ (right panel) assuming $m_{A'} = 3m_{\chi}$, from the full 2016-2022 data set. The bottom row shows the NA64 constraints in the $(\alpha_D;m_{\chi})$ plane on the pseudo-Dirac (left panel) and Majorana (right panel) DM. The limits are shown in comparison with bounds obtained in Refs.[22–24, 28, 29] from the results of the LSND [21, 43], E137 [44], MiniBooNE [46], BABAR [49], COHERENT [78], and direct detection [79] experiments. The favored parameters to account for the observed relic DM density for the scalar, pseudo-Dirac and Majorana type of light DM are shown as the lowest solid line in top plots; see, e.g. [38].

each *i*th entry in Eq.(3) are simulated and reconstructed with the same selection criteria and efficiency corrections as for the data sample from run *i*. The combined 90% C.L. exclusion limits on ϵ as a function of the A' mass, calculated by taking into account the estimated backgrounds and systematic errors ~ 15% for the $\epsilon^i_{A'}$ dominated by the ~ 10% uncertainty in the A' yield [47] can be seen in Fig. 3.

Using obtained limits, Eqs.(1) and (2), one can get constraints on the LDM models, which are shown in the $(y;m_{\chi})$ and $(\alpha_D;m_{\chi})$ planes in Fig. 4 for $m_{\chi} \leq 1$ GeV. The favored y parameter curves for scalar, pseudo-Dirac (with a small splitting) and Majorana scenario of LDM obtained by taking into account the observed relic DM density [38] are also shown on the same plot. One can see that our results are already starting to probe the $y; m_{\chi}$ parameter space predicted for the benchmark values $\alpha_D = 0.1$ and $m_{A'} = 3m_{\chi}$ [29, 30] providing the best limits in comparison with bounds from other experiments. Note, that choice of $\alpha_D = 0.5$ value is still compatible with the constraints obtained from consideration of the running α_D [72, 80]. The limits on α_D for the case of pseudo-Dirac fermions shown in Fig. 4 (left panel in the bottom row) were calculated by taking the value f = 0.25, while for the Majorana case (right panel) the value f = 3 in Eq.(2) was used, see Ref.[48].

In summary, with the combined statistics of the 2016-2022 runs, NA64 started probing the very exciting region of parameter space predicted by benchmark LDM scenarios. Using the missing energy technique, NA64 places stringent bounds on ϵ , y, α_D which are one or more orders more sensitive than the other searches in the mass range $0.001 \leq m_{A'} \leq 0.35$ GeV [10]. The scalar and Majorana models for the coupling $\alpha_D \leq 0.1$, the mass range $0.001 \leq m_{\chi} \leq 0.1$ GeV and $3m_{\chi} \leq m_{A'}$ are excluded. Further detector upgrade will enable improvements in sensitivity and coverage of the LDM parameter space.

ACKNOWLEDGMENTS

We gratefully acknowledge the support of the CERN management and staff, as well as contributions from HISKP, University of Bonn (Germany), ETH Zurich

- [1] E.W. Kolb and M.S. Turner, Front. Phys. 69, 1 (1990).
- [2] V. A. Rubakov and D. S. Gorbunov, "Introduction to the Theory of the Early Universe: Hot Big Bang Theory," World Scientific, 2017, 596 p.
- [3] P. Fayet, Phys. Lett. **95B**, 285 (1980).
- [4] M. Pospelov, A. Ritz, and M. B. Voloshin, Phys. Lett. B 662, 53 (2008).
- [5] N. Arkani-Hamed, D. P. Finkbeiner, T. R. Slatyer, and N. Weiner, Phys. Rev. D 79, 015014 (2009).
- [6] J. Jaeckel and A. Ringwald, Annu. Rev. Nucl. Part. Sci. 60, 405 (2010).
- [7] L. B. Okun, Sov. Phys. JETP 56, 502 (1982) [Zh. Eksp. Teor. Fiz. 83 892 (1982)].
- [8] P. Galison and A. Manohar, Phys. Lett. **136B**, 279 (1984).
- [9] B. Holdom, Phys. Lett. **166B**, 196 (1986).
- [10] R. L. Workman *et al.* (Particle Data Group), Prog. Theor. Exp. Phys. **2022**, 083C01 (2022).
- [11] S. Abrahamyan et al. (APEX Collaboration), Phys. Rev. Lett. **107**, 191804 (2011).
- [12] H. Merkel et al., Phys. Rev. Lett. **112**, 221802 2014).
- [13] J. P. Lees et al. (BABAR Collaboration), Phys. Rev. Lett. 113, 201801 (2014).
- [14] A. Adare et al. (PHENIX Collaboration), Phys. Rev. C 91, 031901 (2015).
- [15] J. R. Batley et al. (NA48/2 Collaboration), Phys. Lett. B 746, 178 (2015).
- [16] A. Anastasi et al. (KLOE-2 Collaboration), Phys. Lett. B 757, 356 (2016).
- [17] D. Banerjee *et al.* (NA64 Collaboration), Phys. Rev. Lett. 120, 231802 (2018).
- [18] D. Banerjee *et al.* (NA64 Collaboration), Phys. Rev. D 101, 071101 (2020).
- [19] P. Fayet, Phys. Rev. D 75, 115017 (2007).
- [20] M. Pospelov, Phys. Rev. D 80, 095002 (2009).
- [21] P. deNiverville, M. Pospelov, and A. Ritz, Phys. Rev. D 84, 075020 (2011).
- [22] E. Izaguirre, G. Krnjaic, P. Schuster, and N. Toro, Phys. Rev. D 91, 094026 (2015).
- [23] E. Izaguirre, G. Krnjaic, P. Schuster, and N. Toro, Phys. Rev. Lett. **115**, 251301 (2015).
- [24] E. Izaguirre, Y. Kahn, G. Krnjaic, and M. Moschella, Phys. Rev. D 96, 055007 (2017).
- [25] P. J. Fitzpatrick, H. Liu, T. R. Slatyer, and Y.-D. Tsai, Phys. Rev. D 106, 083517 (2022).
- [26] A. Boveia et al., arXiv:2210.01770.
- [27] R. Essig *et al.*, arXiv:1311.0029.
- [28] J. Alexander et al., arXiv:1608.08632.
- [29] M. Battaglieri et al., arXiv:1707.04591.
- [30] J. Beacham et al., J. Phys. G 47, 010501 (2020).

I and PID2021-123955NA-100 funded by MCIN/AEI/ 10.13039/501100011033/FEDER, UE (Spain). This result is part of a project that has received funding from the European Research Council (ERC) under the European Union's Horizon 2020 research and innovation programme, Grant agreement No. 947715 (POKER).

- [31] R. Alemany *et al.*, arXiv:1902.00260.
- [32] P. Agraval et al., Eur. Phys. J. C 81, 1015 (2021).
- [33] R. Essig, Y. Kahn, S. Knapen, A. Ringwald, and N. Toro, arXiv:2203.10089.
- [34] B. Batell, N. Blinov, Ch. Hearty, and R. McGehee, arXiv:2207.06905.
- [35] S. Gori et al., arXiv:2209.04671.
- [36] G. Lanfranchi, M. Pospelov, and Ph. Schuster, Ann. Rev. Nucl. Part. Sci. 71, 279 (2021).
- [37] M. Fabbrichesi, E. Gabrielli, and G. Lanfranchi, arXiv:2005.01515.
- [38] A. Berlin, N Blinov, G. Krnjaic, P. Schuster, and N. Toro, Phys. Rev. D 99, 075001 (2019).
- [39] H. S. Lee, Phys. Rev. D 90, 091702(R) (2014).
- [40] M. D. Diamond and P. Schuster, Phys. Rev. Lett. 111, 221803 (2013).
- [41] H. Davoudiasl, H. S. Lee, and W. J. Marciano, Phys. Rev. D 89, 095006 (2014).
- [42] R. Essig, J. Mardon, M. Papucci, T. Volansky, and Y. M. Zhong, J. High Energy Phys. 1311 (2013) 167.
- [43] B. Batell, M. Pospelov, and A. Ritz, Phys. Rev. D 80, 095024 (2009).
- [44] B. Batell, R. Essig, and Z. Surujon, Phys. Rev. Lett. 113, 171802 (2014).
- [45] D. Banerjee *et al.* (NA64 Collaboration), Phys. Rev. Lett. 118, 011802 (2017).
- [46] A. A. Aguilar-Arevalo *et al.* (MiniBooNE Collaboration), Phys. Rev. D 98, 112004 (2018).
- [47] D. Banerjee *et al.* (NA64 Collaboration), Phys. Rev. D 97, 072002 (2018).
- [48] D. Banerjee *et al.* (NA64 Collaboration), Phys. Rev. Lett. 123, 121801 (2019).
- [49] J. P. Lees *et al.* (*BABAR* Collaboration), Phys. Rev. Lett. **119**, 131804 (2017).
- [50] E. Cortina Gil *et al.*, (NA62 Collaboration), J. High Energy Phys. **05** (2019) 182.
- [51] D. Hanneke, S. Fogwell and G. Gabrielse, Phys. Rev. Lett. 100, 120801 (2008).
- [52] R. Bouchendira, P. Clade, S. Guellati-Khelifa, F. Nez, and F. Biraben, Phys. Rev. Lett. 106, 080801 (2011).
- [53] L. Morel, Z. Yao, P. Clade, and S. Guellati-Kh?lifa, Nature 588, 61 (2020).
- [54] T. Aoyama, M. Hayakawa, T. Kinoshita, and M. Nio, Phys. Rev. Lett. **109**, 111807 (2012).
- [55] S.N. Gninenko, N.V. Krasnikov, and V.A. Matveev, Usp. Fiz. Nauk **191**, 12, 1361 (2021).
- [56] P. Crivelli, arXiv:2301.09905.
- [57] S. N. Gninenko, Phys. Rev. D 89, 075008 (2014).
- [58] S. Andreas et al., arXiv:1312.3309.
- [59] L. Marsicano, M. Battaglieri, M. Bondi, C.D.R. Carva-

jal, A. Celentano, M. De Napoli, R. De Vita, E. Nardi, M. Raggi, and P. Valente, Phys. Rev. Lett. **121**, 041802 (2018).

- [60] Yu. M. Andreev *et al.* (NA64 Collaboration), Phys. Rev. D **104**, L091701 (2021).
- [61] S. N. Gninenko, Phys. Rev. D 85, 055027 (2012).
- [62] S. N. Gninenko, Phys. Lett. B **713**, 244 (2012).
- [63] See, for example, http://sba.web.cern.ch/sba/
- [64] D. Banerjee, P. Crivelli, and A. Rubbia, Adv. High Energy Phys. 2015, 105730 (2015).
- [65] E. Depero *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A 866, 196 (2017).
- [66] Yu.M. Andreev et al. (NA64 Collaboration), arXiv:2305.19411[hep-ex].
- [67] Yu.M. Andreev *et al.* (NA64 Collaboration), Phys. Rev. Lett. **129**, 161801 (2022).
- [68] S. Agostinelli *et al.* (GEANT4 Collaboration), Nucl. Instrum. Methods Phys. Res., Sect. A **506**, 250 (2003).
- [69] J. Allison *et al.*, IEEE Trans. Nucl. Sci. **53**, 270 (2006).
- [70] M. Bondi, A. Celentano, R. R. Dusaev, D. V. Kirpichnikov, M. M. Kirsanov, N. V. Krasnikov, L. Marsicano, and D. Shchukin, Comput. Phys. Commun. 269, 108129

(2021).

- [71] S. N. Gninenko, N. V. Krasnikov, M. M. Kirsanov, and D. V. Kirpichnikov, Phys. Rev. D 94, 095025 (2016).
- [72] S. N. Gninenko, D. V. Kirpichnikov, M. M. Kirsanov, and N. V. Krasnikov, Phys. Lett. B 782, 406 (2018).
- [73] D. Banerjee *et al.* (NA64 Collaboration), Phys. Rev. Lett. 125, 081801 (2020).
- [74] I. Antcheva *et al.*, Comput. Phys. Commun. **180**, 2499 (2009).
- [75] T. Junk, Nucl. Instrum. Methods Phys. Res., Sect. A 434, 435 (1999).
- [76] G. Cowan, K. Cranmer, E. Gross, and O. Vitells, Eur. Phys. J. C 71, 1 (2011).
- [77] A. L. Read, J. Phys. G 28, 2693 (2002).
- [78] D. Akimov *et al.* (COHERENT Collaboration), Phys. Rev. Lett. **130**, 051803 (2023).
- [79] R. Essig, A. Manalaysay, J. Mardon, P. Sorensen, and T. Volansky, Phys. Rev. Lett. 109, 021301 (2012).
- [80] H. Davoudiasl and W. J. Marciano, Phys. Rev. D 92, 035008 (2015).