# **MOJIVIRG**



# Gravitational wave astronomy with Virgo and the GW detectors network



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Virgo document: VIR-0650A-18

#### Primer on gravitational waves

- Wave-like perturbation of space-time that travel at the speed of light
  - Predicted by Einstein in 1916
  - Result from linearization of General Relativity field equations
- Produce a change in proper distance between geodesics
  - Act in plane orthogonal to direction of propagation
  - Two independent polarization
  - Relative distance change (strain)
- Generated by systems with non-zero second derivative of the mass quadrupole moment
  - accelerated, non-symmetric mass distributions (i.e. binary systems)
- Typical effect from astronomical sources is very small!



$$Q_{jk} = \int \rho x_i x_k d^3 x \qquad h_{jk} = \frac{2}{r} \ddot{Q}_{jk}$$
$$h_{ab} = \begin{pmatrix} h_+ & h_\times \\ h_\times & -h_+ \end{pmatrix}$$

 $\Delta L/L \sim 10^{-21}$ 



#### Broadband GW astrophysics



September 27th, 2018 - NTIHEP, Becici, Montenegro

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#### Laser interferometers



#### Initial detectors: not just Michelson interferometers...







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- Showed that such instruments can be successfully operated
- Set upper limits on GW emission from several sources
- Paved the (scientific, political and financial) way for Advanced Detectors

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#### From Initial to Advanced







September 27th, 2018 - NTIHEP, Becici, Montenegro

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#### From Initial to Advanced (design values)



#### Also improved:

- Beam size
- mirror polishing and coatings
- compensation of thermal effects
- Vacuum
- Stray light control



#### 14 September 2015: the birth of GW astronomy



#### LIGO: O1 and O2... up to August 2017

- 4 Binary Black Hole detections and one likely candidate
- 12 (15) new Black Holes observed
- Unexpected mass range
- LIGO-Virgo Collaboration effort, despite only LIGO instruments being online



LVT151012 ~~~~~~

GW150914 70 60 GW170104 **Solar Masses** <sup>20</sup>
<sup>20</sup>
<sup>20</sup> LVT151012 20 **X-Ray Studies** 10 -......... GW170608 GW151226 0

### Tests of General Relativity

- GW exists! No big surprise here...
- Signal vs GR models:
  - Deviation in residuals from random noise: <4%
  - Parameter inferred from pre- and post-merger phase are consistent
- No evidence of deviation in PN parameters
- Test of modified GW dispersion relation:

 $E^2 = p^2 c^2 + A p^\alpha c^\alpha$ 

• Upper bounds on graviton's mass

We didn't see new physics either! But we are happy about it... 🙂



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0.8

0.6

0.2

1.0

0.5

-0 4

0.5PN

Abbott et al. PRL 118, 221101 (2017)

OPN

 $10^{-19}$ 

 $10^{-20}$ 

0.0

0.5

1.0

s.  $[peV^{2}$ 

A

40

50

60

1PN

1.5PN

 $\diamond$ 

 $\nabla$ 

1.5

A > 0

A < 0

2PN

 $\bigcirc$ 

Ŧ

70

80

Final mass  $M_f(M_{\odot})$ 

Final spin  $a_j$ 

0.2

0.1

-0.1-0.2 $-0^{-1}$ 

 $\delta \hat{p}_{i}$ 



1.0

0.5

0.0

-0.5

-1.0

 $\Delta a_f/a_f$ 

inspiral

90

GW150914 + GW151226

100

110 12

Abbott et al. PRL 116, 221101 (2016)

#### Astrophysical results

- Bounds on merger rates:  $12 230 \text{ Gpc}^{-3} \text{yr}^{-1}$
- Observed black holes in unexpected mass range
  - Hints on formation mechanisms and environment
  - Estimates of mass distribution law
  - Population estimates (needs more data)
- Hints on formation channels
  - Low-metallicity needed for high-mass BH remnant
  - Rates too high for some formation channels (but could contribute to total)
  - Misaligned spins slightly favored





#### AdV quest to join O2

AdV best BNS range from May 7 (C8) to July 30 (ER12)





#### Detectors sensitivity in O2



#### LIGO: O1 and O2... up to August 2017



#### GW170814: 3-detectors BBH observation

Abbott et al. PRL 119, 141101 (2017)

Virgo

- 31 + 25 solar masses BBH coalescence
- Virgo contribution:
  - Improved sky localization
  - First ever polarization studies



Hanford

SNR

Livingston



#### GW170817: the birth of multi-messenger astronomy

- Signal promptly recognized as BNS candidate
  - Long duration
  - High merger frequency
- Smoking gun: short gamma ray burst coincidence
- Virgo non-detection!
  - Accurate localization by accounting for antenna pattern







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#### An unprecedented coordinated observational campaign



### GW170817: scientific highlights

- Confiremed BNS mergers as sources of sGRB
  - Chace association probability  $> 5.3~\sigma$
- (weak) constraints on NS Equation of State
  - Less compact stars disfavored
- Independent measurement of the Hubble Constant
  - By comparison of GW distance with source redshift
- Measurement of the speed of GW (compared to EM)



0.04

0.03

*p*(H<sub>0</sub> | GW170817) ( 0.0

0.00

50

Mpc)

F

#### GW170817: kilonova emission

- Contribution of heavy elements observed in emission spectrum
  - Support hypothesis of formation by neutron capture (r-process)
- Evidence supports kilonova event
  - Atypical luminosity vs distance





#### To observe or to improve? This is the question...



#### Gravitational wave observatories are here to stay (and be useful)!

- We now have a proven network of gravitational wave *observatories* 
  - "Routine" detection demonstrated
  - A roadmap exist for a steady increase in sensitivity:
    - Detection rates to increase steadily and substantially run-by-run
- Virgo is an essential complement to the LIGO detectors
  - Improved parameter estimation
  - More precise sky localization -> Multi-messenger astronomy
  - More detectors will join in the future
- Extremely *rich scientific output* even with few GW detections
  - Consequences in fundamental physics, astrophysics, astronomy...

## Thanks for listening... to the tune of gravitational waves!