

in collaboration with

I.R. Gabdrakhmanov¹

in collaboration with

N.A Gramotkov^{2,3}, A.V. Kotikov², O.V. Teryaev², D.A. Volkova^{2,4} and
I.A. Zemlyakov^{2,5}

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Introduction:

Analytic Perturbation
Theory

1/L expansion

Bjorken sum rule

Higher twist
resummation

GDH constraint

Conclusion

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Outline:

Правило сумм
Бёркена с
аналитической
связью при малых
значениях Q^2

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Introduction:

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Bjorken sum rule

Higher twist resummation
Low Q^2
GDH constraint
Photoproduction limit ($Q^2 \rightarrow 0$)

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Analytic Perturbation Theory ("Minimal Approach")

- ▶ Q^2 analyticity
- ▶ preserve RG-invariance
- ✓ Landau singularity removed

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Power series is replaced by non-power functional series^{1,2}

$$\sum_k c_k \alpha_s^k \rightarrow \sum_k c_k A_k$$

$$A_n^{(i)}(Q^2) = \frac{1}{\pi} \int_0^\infty \frac{\rho_n^{(i)}(\sigma)}{\sigma + Q^2} d\sigma, \quad \rho_n^{(i)}(\sigma) = \text{Im}[\alpha_{(i)}^n(-\sigma)] \quad (1)$$

Later generalized to the non-integer powers of α_s - Fractional APT^{3 4 5},

¹I. L. Solovtsov and D. V. Shirkov. In: *Theor. Math. Phys.* 120 (1999). [Teor. Mat. Fiz. 120, 482 (1999)], pp. 1220–1244. arXiv: hep-ph/9909305 [hep-ph].

²K. A. Milton, I. L. Solovtsov, and O. P. Solovtsova. In: *Phys. Lett. B* 415 (1997), pp. 104–110. arXiv: hep-ph/9706409.

³A. P. Bakulev, S. V. Mikhailov, and N. G. Stefanis. In: *Phys. Rev. D* 72 (2005), p. 074014. arXiv: hep-ph/0506311.

⁴A. P. Bakulev, S. V. Mikhailov, and N. G. Stefanis. In: *Phys. Rev. D* 75 (2007). [Erratum: *Phys. Rev. D* 77, 079901 (2008)], p. 056005. arXiv: hep-ph/0607040.

⁵A. P. Bakulev, S. V. Mikhailov, and N. G. Stefanis. In: *JHEP* 06 (2010), p. 085. arXiv: 1004.4125 [hep-ph].

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FAPT 1/L expansion ($L = \ln(\frac{Q^2}{\Lambda^2})$)

$$\tilde{a}_{n+1}^{(k)}(Q^2) = \frac{(-1)^n}{n!} \frac{d^n a_s^{(k)}(Q^2)}{(dL)^n}, \quad a_s^{(k)}(Q^2) = \frac{\beta_0 \alpha_s^{(k)}(Q^2)}{4\pi} = \beta_0 \bar{a}_s^{(k)}(Q^2), \quad (2)$$

The series of derivatives $\tilde{a}_n(Q^2)$ can successfully replace the corresponding series of a_s -powers (see, e.g. [\(Kotikov, Zemlyakov:2022\)](#)). Indeed, each derivative reduces the a_s power but is accompanied by an additional β -function \tilde{a}_s^2 . Thus, each application of a derivative yields an additional a_s , and thus it is indeed possible to use a series of derivatives instead of a series of a_s -powers. In the LO, the series of derivatives $\tilde{a}_n = a_s^n$.

Beyond LO, the relationship between \tilde{a}_n and a_s^n was established in [\(Cvetic, Valenzuela: 2006\)](#), [\(Cvetic, Kogerler, Valenzuela: 2010\)](#) and extended to the fractional case, where n is a non-integer ν , in [\(Cvetic, Kotikov: 2012\)](#)

$$\tilde{a}_\nu^{(i+1)}(Q^2) = \tilde{a}_\nu^{(1)}(Q^2) + \sum_{m=1}^{(i)} \frac{\Gamma(\nu+m)}{m!\Gamma(\nu)} \left(\hat{R}_m \tilde{a}_{\nu+m}^{(1)}(Q^2) \right) \quad (3)$$

$$\tilde{A}_{MA,\nu}^{(i+1)}(Q^2) = \tilde{A}_{MA,\nu}^{(1)}(Q^2) + \sum_{m=1}^i \frac{\Gamma(\nu+m)}{m!\Gamma(\nu)} \left(\hat{R}_m \tilde{A}_{MA,\nu+m}^{(1)}(Q^2) \right) \quad (4)$$

Bjorken polarized sum rule

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Bjorken sum rule is defined as the 1st Mellin moment of the g_1 by the Bjorken variable of the difference for proton and neutron for the fixed Q^2 :

$$\Gamma_1^{p-n}(Q^2) = \int_0^1 [g_1^p(x, Q^2) - g_1^n(x, Q^2)] dx. \quad (5)$$

BSR in the OPE form (twist-2 + twist-4):

$$\Gamma_1^{p-n}(Q^2) = \frac{|g_{A/V}|}{6} \left[1 - D_{Bj}(Q^2) \right] + \sum_{i=2}^{\infty} \frac{\mu_{2i}}{Q^{2i-2}}, \quad (6)$$

where $|g_{A/V}| = 1.2762 \pm 0.0005$

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Higher twist resummation

Non-perturbative HT resummation via spectral function^a:

$$\sum_{n=1}^{\infty} \mu_{2n+2}^{p-n} \left(\frac{M_{HT}^2}{Q^2} \right)^n = \int_{-\infty}^{\infty} \frac{f(x) M^2}{Q^2 - x M^2} dx, \quad (7)$$

^aO. Teryaev. In: *Nucl. Phys. Proc. Suppl.* 245 (2013), pp. 195–198. arXiv: 1309.1985 [hep-ph].

δ -like like spectral density^a

$$\Delta_{HT}(Q^2) = \mu_4 \frac{M_{HT}^2}{Q^2 + M_{HT}^2} \quad (8)$$

^aV. L. Khandramai, O. V. Teryaev, and I. R. Gabdrakhmanov. In: *J. Phys. Conf. Ser.* 678.1 (2016), p. 012018.

Both using Massive APT^a with $\ln(Q^2) \rightarrow \ln(Q^2 + M_{gl}^2)$ modified coupling $A(Q^2, M_{gl}^2)$

Breit-Wigner like spectral density^a

$$\Delta_{HT}(Q^2) = \mu_4 \frac{M_{HT}^2 (M_{HT}^2 + Q^2)}{(M_{HT}^2 + Q^2)^2 + M_{HT}^2 \sigma^2} \quad (9)$$

^aI. R. Gabdrakhmanov, O. V. Teryaev, and V. L. Khandramai. In: *J. Phys. Conf. Ser.* 938.1 (2017), p. 012046.

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1/L FAPT expansion

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$$D_{BS}^{(k)}(Q^2) = \frac{4}{\beta_0} \left(\tilde{a}_1^{(k)} + \sum_{m=2}^k \tilde{d}_{m-1} \tilde{a}_m^{(k)} \right), \quad (10)$$

$$D_{MA,BS}^{(k)}(Q^2) = \frac{4}{\beta_0} \left(\tilde{A}_{MA}^{(k)} + \sum_{m=2}^k \tilde{d}_{m-1} \tilde{A}_{MA,v=m}^{(k)} \right), \quad (11)$$

^aA. V. Kotikov and I. A. Zemlyakov. In: *J. Phys. G* 50.1 (2023), p. 015001. arXiv: 2203.09307 [hep-ph].

$1/L$ expansion with the massive twist-4 HT representations⁶

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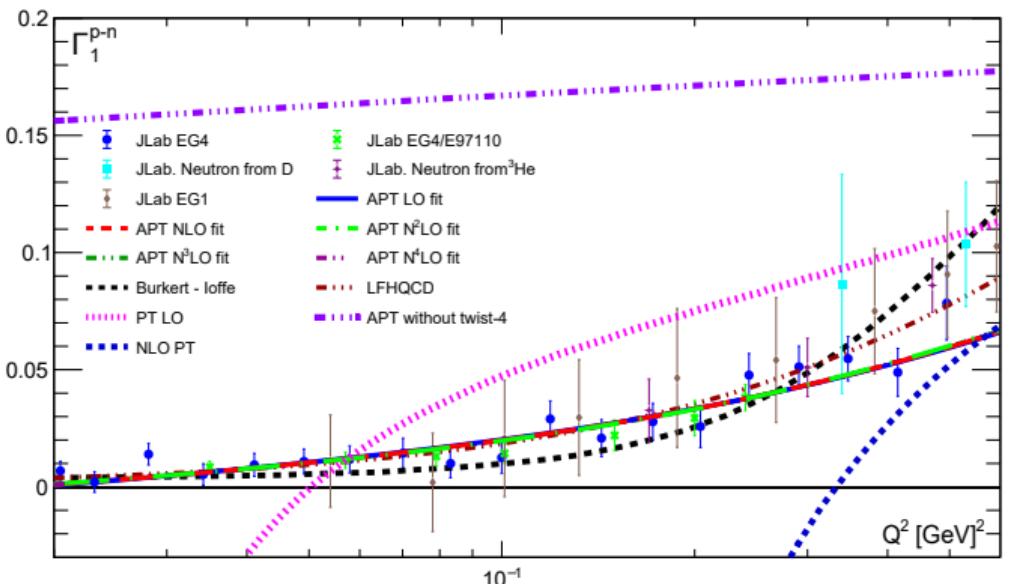


Figure: The results for $\Gamma_1^{p-n}(Q^2)$ in the first four orders of APT with $\sigma = \sigma_p$.

⁶I. R. Gabdrakhmanov, N. A. Gramotkov, A. V. Kotikov, D. A. Volkova, and I. A. Zemlyakov. In: *JETP Letters* 118 (2023). arXiv: 2307.16225 [hep-ph].

Low Q^2 behavior

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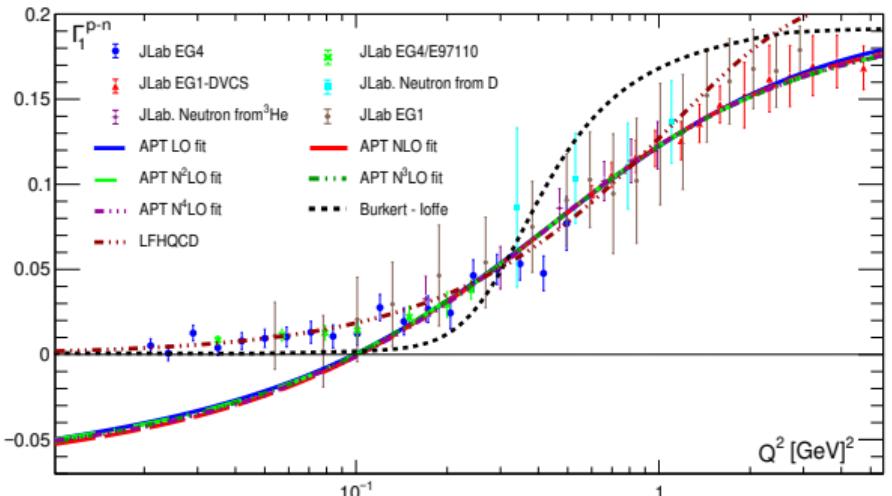


Figure: The results for $\Gamma_1^{p-n}(Q^2)$ in the first four orders of APT for $Q^2 \leq 0.6 \text{ GeV}^2$.

$\Gamma_{MA,1}^{p-n}(Q^2)$ takes negative unphysical values for low $Q^2 < 0.02 \text{ GeV}^2$

The Gerasimov-Drell-Hearn sum rule

Правило сумм Бёркена с аналитической связью при малых значениях Q^2

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The GDH^{7,8} sum rule for the nucleon connects its anomalous magnetic moment and the difference of the photoabsorption cross sections for parallel - σ_P and antiparallel σ_A spins of the photon and the nucleon

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$$\frac{2\pi e^2 \mu^2}{M^2} = \int_0^\infty \frac{d\nu}{\nu} (\sigma_P(\nu) - \sigma_A(\nu)), \quad (12)$$

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⁷S. B. Gerasimov. In: Sov. J. Nucl. Phys. 2 (1966). [Yad. Fiz. 2, 598 (1965)], pp. 430-433.

Photoproduction limit

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Following^a and analogically to^b we take into account Gerasimov-Drell-Hearn sum rule and get the equation

$$\frac{d}{dQ^2} \Gamma_1^{p-n}(Q^2 = 0) = G, \quad G = \frac{-(\mu_p - 1)^2 + \mu_n^2}{8M^2} = 0.0631, \quad (13)$$

also finiteness of cross section for real photons implies:

$$\Gamma_1^{p-n}(Q^2 = 0) = 0, \quad (14)$$

^aJ. Soffer and O. Teryaev. In: *Phys. Rev. Lett.* 70 (1993), pp. 3373-3375.

^bKhandramai, Teryaev, and Gabdrakhmanov, "Infrared modified QCD couplings and Bjorken sum rule"; Gabdrakhmanov, Teryaev, and Khandramai, "Infrared models for the Bjorken sum rule in the APT approach".

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For MA coupling:

$$A^{(k)}(Q^2 = 0) = \tilde{A}_{m=1}^{(k)}(Q^2 = 0) = 1, \quad \tilde{A}_{m>1}^{(k)}(Q^2 = 0) = 0, \quad (15)$$

By definition $Q^2 \frac{d}{dQ^2} \tilde{A}_n(Q^2) \sim \tilde{A}_{n+1}(Q^2)$, so

$$\frac{d}{dQ^2} \tilde{A}_n(Q^2 \rightarrow 0) \rightarrow \infty. \quad (16)$$

thus every term in $D_{MA,BS}(Q^2)$ becomes to be divergent at $Q^2 \rightarrow 0$.

So we assume the relation between twist-2 and twist-4 terms, that leads to the appearance of a new contribution.

Taking into account GDH sum rule and (15) possible modification to improve low Q^2 behavior is:

$$\Gamma_{MA,1}^{p-n}(Q^2) = \frac{g_A}{6} \left(1 - D_{MA,BS}(Q^2) \cdot \frac{Q^2}{Q^2 + M^2} \right) + \frac{\hat{\mu}_{MA,4} M^2}{Q^2 + M^2} + \frac{\hat{\mu}_{MA,6} M^4}{(Q^2 + M^2)^2}, \quad (17)$$

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(18)

Thus our modification (17) leads to:

$$\left\{ \begin{array}{l} \hat{\mu}_{MA,6} = -GM^2 + \frac{5g_{A/V}}{54} \approx -GM^2 + 0.1182, \\ \hat{\mu}_{MA,4} = -\frac{g_{A/V}}{6} - \hat{\mu}_{MA,6} = GM^2 - \frac{7g_{A/V}}{27} \approx GM^2 - 0.3309, \end{array} \right.$$

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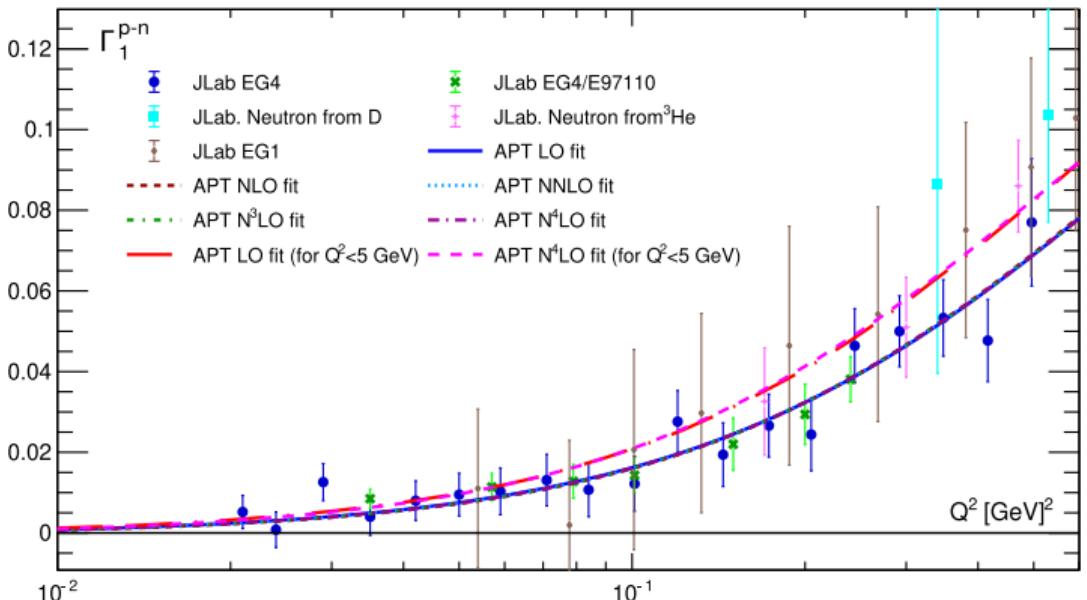


Figure: The results for $\Gamma_1^{p-n}(Q^2)$ (17) in the first four orders of APT for $Q^2 < 0.6 \text{ GeV}^2$ ^a

^aI. R. Gabdrakhmanov, N. A. Gramotkov, A. V. Kotikov, O. V. Teryaev, D. A. Volkova, and I. A. Zemlyakov. In: (2024). arXiv: 2404.01873 [hep-ph].

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- ▶ As previous studies showed: while conventional PT fails to describe BSR data, APT with the massive twist-4 HT term gives a good agreement with the data
- ▶ However the fits extended to very low Q^2 values produce negative results
- ▶ We successfully solved the issue modifying of the OPE and found a good agreement in the whole experimental region with the help of the GDH sum rule
- ▶ The distinctive feature of the solution is continuation of the BSR down to the photoproduction limit with a finite derivative $\frac{d}{dQ^2} \Gamma_{MA,1}^{p-n}$