Solving Dyson-Schwinger and Bethe-Salpeter equations at zero and finite temperatures

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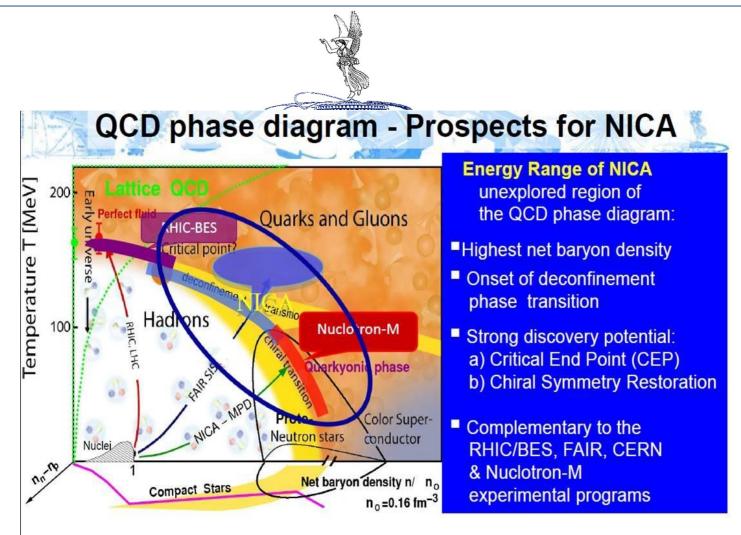
(nucl-th 1807.10075, PhysPNL 15,2018, PRC 89 (2014), PRC 91 (2015), J. Mod. Phys. 7 (2016), Few Body Syst.49, Few Body Syst.42...)

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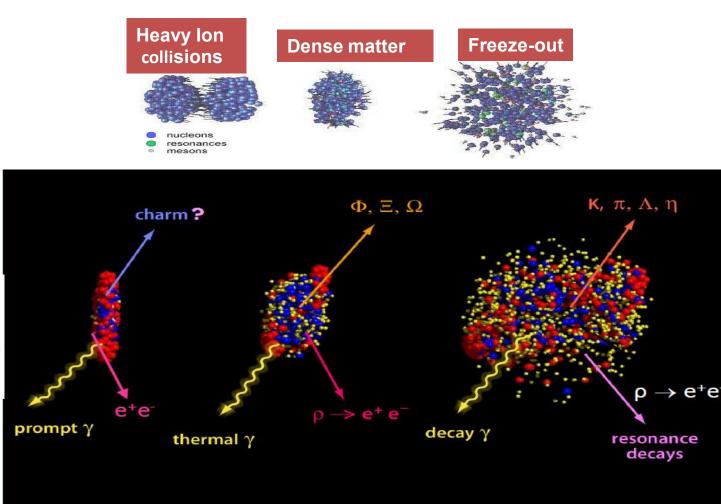
MOTIVATION

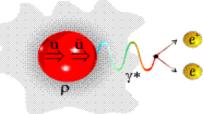
QGP signals (GSI (10-40 GeV/N), RHIC ($\sqrt{S_{NN}}$ > 200 GeV), NICA ($\sqrt{S_{NN}}$ > 4-11 GeV/N)...



MOTIVATION

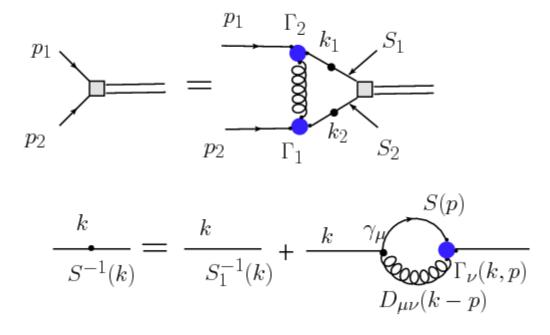
> QGP signals (GSI, RHIC, NICA ..)





$qar{q}$ Bound States: BS & Dyson-Schwinger Equations

Mesons as $q\bar{q}$ bound systems: $\pi^+(\sim 0.140, u\bar{d}) \ \rho^+(\sim 0.77, u\bar{d}) K^+(\sim 0.494, u\bar{s}), K^{*+}(\sim 0.891, u\bar{s}), \eta_c(\sim 2.98, c\bar{c}), D^+(1.869, c\bar{d}) \dots$



$$S^{-1}(p) = S_0^{-1}(p) + \frac{4}{3} \int \frac{d^4k}{(2\pi)^4} \left[g^2 \mathcal{D}_{\mu\nu}(p-k) \right] \gamma_{\mu} S(k) \gamma_{\nu} ,$$

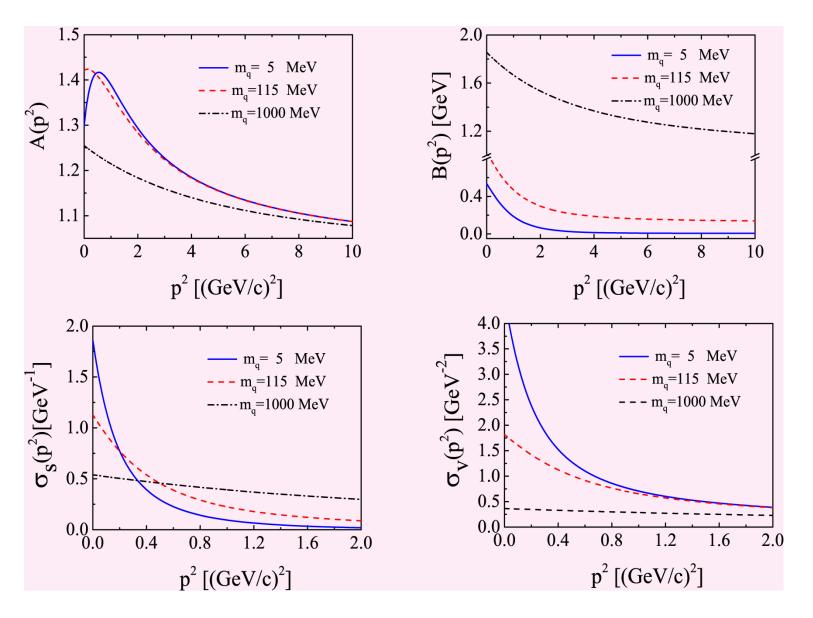
$$S(p) = \frac{Z(p^2)}{i\gamma \cdot p + M(p^2)} = \frac{1}{i\gamma \cdot pA(p^2) + B(p^2)} = -i\gamma \cdot p\sigma_V(p^2) + \sigma_S(p^2)$$

Rainbow approximation ($\Gamma^a_{\nu}(l,p) = \frac{\lambda^a}{2}\gamma_{\nu}$) + effective model for gluon propagator:

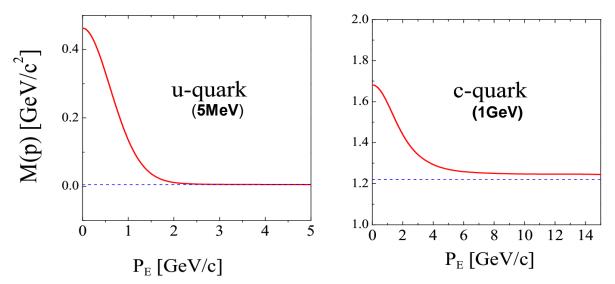
$$\frac{D(q^2)}{q^2} = \frac{4\pi^2}{\omega^6} C q^2 \mathrm{e}^{-q^2/\omega^2} + \left(\frac{8\pi^2 \gamma_m}{\ln\left[\tau + \left(1 + \frac{q^2}{\Lambda_{QCD}^2}\right)^2\right]} F(q^2) \rightarrow \left\langle \bar{q}q \rangle = -\mathrm{N_c} \int \frac{\mathrm{d}^4 \mathbf{p}}{(2\pi)^4} \mathrm{Tr}\left[\mathbf{S}(\mathbf{p})\right] = -\mathrm{N_c} \int \frac{\mathrm{d}^4 \mathbf{p}}{(2\pi)^4} \sigma_s(\mathbf{p}) \mathrm{d}\mathbf{p} \mathrm{d}\mathbf{p}$$

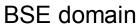
$$\begin{aligned} A(p) &= 1 + 2C \int dk \frac{k^4}{p} \frac{A(k)}{k^2 A^2(k) + B^2(k)} e^{-(p-k)^2/\omega^2} \left\{ \frac{[p^2 + k^2 + 2\omega^2]}{kp} I_2^{(s)}(z) - 2I_1^{(s)}(z) \right\}, \\ B(p) &= m_q + 2C \int dk k^3 \frac{B(k)}{k^2 A^2(k) + B^2(k)} e^{-(p-k)^2/\omega^2} \left\{ \frac{[p^2 + k^2]}{kp} I_1^{(s)}(z) - 2I_2^{(s)}(z) \right\}, \end{aligned}$$

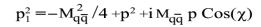
RESULTS

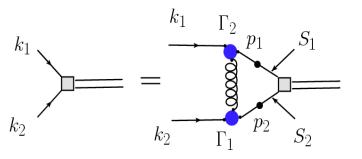


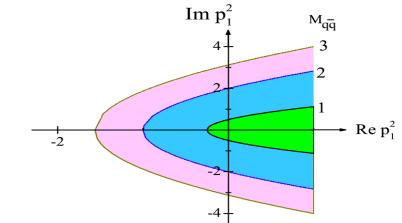


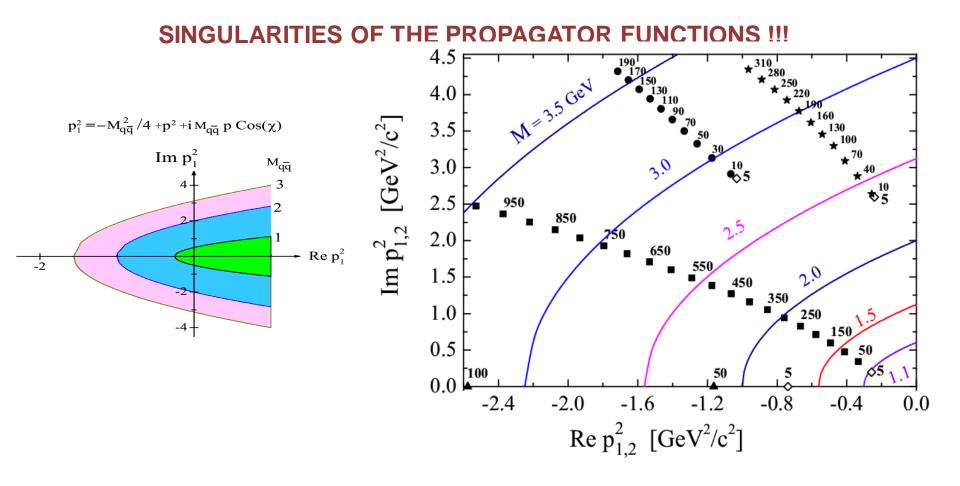




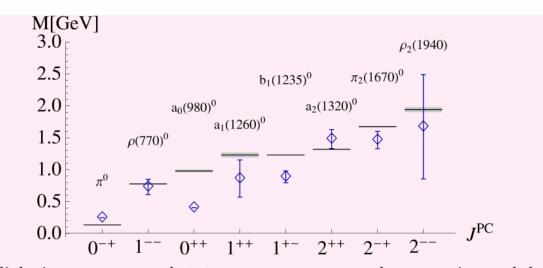




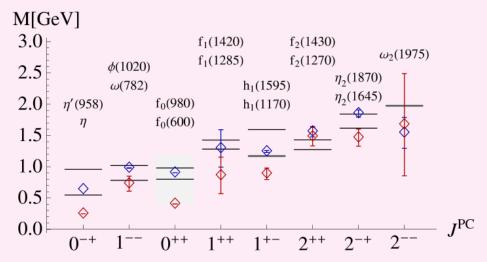




$$\sigma_{s,v}(\tilde{k}^2) = \tilde{\sigma}_{s,v}(\tilde{k}^2) + \sum_i \frac{\operatorname{res}[\sigma_{s,v}(\tilde{k}_{0i}^2)]}{\tilde{k}^2 - k_{0i}^2}, \qquad \tilde{\sigma}_{s,v}(\tilde{k}^2) = \frac{1}{2\pi i} \oint_{\gamma} \frac{\tilde{\sigma}_{s,v}(\xi)}{\xi - \tilde{k}^2} d\xi = \frac{1}{2\pi i} \oint_{\gamma} \frac{\sigma_{s,v}(\xi)}{\xi - \tilde{k}^2} d\xi.$$



The light-isovector ground-state spectrum compared to experimental data.



The light-isoscalar ground-state spectrum compared to experimental data (LPK& SMD PRC **89** (2014), PRC **91** (2015), Few Body Syst. **49** (2011), arXiv: 1012.5372; M. Blank and A. Krassnigg, PRD **84** (2011)

	experiment calculated	
	(estimates)	(† fitted)
$m^{u=d}_{\mu=1 { m GeV}}$	5 - 10 MeV	$5.5 { m MeV}$
$m_{\mu=1 { m GeV}}^s$	100 - 300 MeV	$125 { m MeV}$
- $\langle \bar{q}q \rangle^0_\mu$	$(0.236 \text{ GeV})^3$	$(0.241)^3$
m_{π}	$0.1385~{ m GeV}$	0.138
f_{π}	$0.131~{ m GeV}$	0.131
m_K	$0.496~{ m GeV}$	0.497
f_K	$0.160~{ m GeV}$	0.155
$m_{ ho}$	$0.770~{ m GeV}$	0.742
$f_{ ho}$	$0.216~{ m GeV}$	0.207
$m_{K^{\star}}$	$0.892~{ m GeV}$	0.936
$f_{K^{\star}}$	$0.225~{ m GeV}$	0.241
m_{ϕ}	$1.020~{ m GeV}$	1.072
f_{ϕ}	$0.236~{ m GeV}$	0.259

Overview of the results of the model for the meson masses and decay constant, adapted from P. Maris and P. C. Tandy, PRC **60**, 055214 (1999).

INTERMEDIATE SUMMARY

The Dyson–Schwinger-Bethe-Salpeter approach + rainbow approximation with only two-three effective parameters, describe fairly well the vacuum (T=0) properties (masses, decay constants...) of the scalar, pseudoscalar, vector etc., mesons and allow for Poincarè covariant studies of reactions with mesons

Finite Temperatures

The statistical density matrix: Ansamble average of an operator : $\langle \hat{A} \rangle = \frac{Tr[\hat{A}]}{Tr[\hat{A}]}$

$$\hat{\rho} = \exp\left[-\beta\left(\hat{H} - \mu_i \hat{N}_i\right)\right]$$
$$\langle \hat{A} \rangle = \frac{Tr[\hat{A}\hat{\rho}]}{Tr[\hat{\rho}]} \equiv \frac{Tr[\hat{A}\hat{\rho}]}{Z}$$

$$\mathbf{Z} = \mathbf{Tr}\left[\exp\left\{-\beta\left(\hat{\mathbf{H}} - \boldsymbol{\mu}_{\mathbf{i}}\hat{\mathbf{N}}_{\mathbf{i}}\right)\right\}\right] = \int \mathbf{d}\phi_{\mathbf{a}}\left\langle\phi_{\mathbf{a}}\left|\exp\left\{-\beta\left(\hat{\mathbf{H}} - \boldsymbol{\mu}_{\mathbf{i}}\hat{\mathbf{N}}_{\mathbf{i}}\right)\right\}\right|\phi_{\mathbf{a}}\right\rangle$$

Transition amplitude from a state $|\phi_a\rangle$ to a state $|\phi_b\rangle$ after a time t is $\langle \phi_{\mathbf{b}} | e^{-i\hat{\mathbf{H}}\mathbf{t}} | \phi_{\mathbf{a}} \rangle$, where $|\phi_a\rangle$ is an eigenstate of the field operator $\hat{\phi}(\mathbf{x})$, i.e. $\hat{\phi}(\mathbf{x}) | \phi_{\mathbf{a}} \rangle = \phi(\mathbf{x}) | \phi_{\mathbf{a}} \rangle$

Imaginary time (Matsubara) formalism $\tau = it$ and periodic condition= after a time $\tau = \beta$ the system comes back to its initial state: $\phi(\mathbf{x}, 0) = \pm \phi(\mathbf{x}, \beta)$

$$Z = \int [d\pi] \int d[\boldsymbol{\phi}] \exp\left\{\int_0^\beta d\tau \int d^3x \left(i\pi \frac{\partial \boldsymbol{\phi}}{\partial \tau} - H(\boldsymbol{\phi}, \pi) + \mu_i N_i(\boldsymbol{\phi}, \pi)\right)\right\}$$

Main difference

Furie transform (O(4)symmetry lost!!)

2.

3.

$$\int \frac{d^4x}{(2\pi)^4} \exp(ipx)\phi(p) \longrightarrow T \sum_{n=-\infty}^{\infty} \int \frac{d^3x}{(2\pi)^3} \exp(i \mathbf{px} + \omega_n \tau)\phi(\mathbf{p}, \omega_n)$$

 $\omega_n = \pi (2n+1)T$ (fermions); $\omega_n = 2n\pi T$ (bosons) The inverse quark propagator is now parametrized as

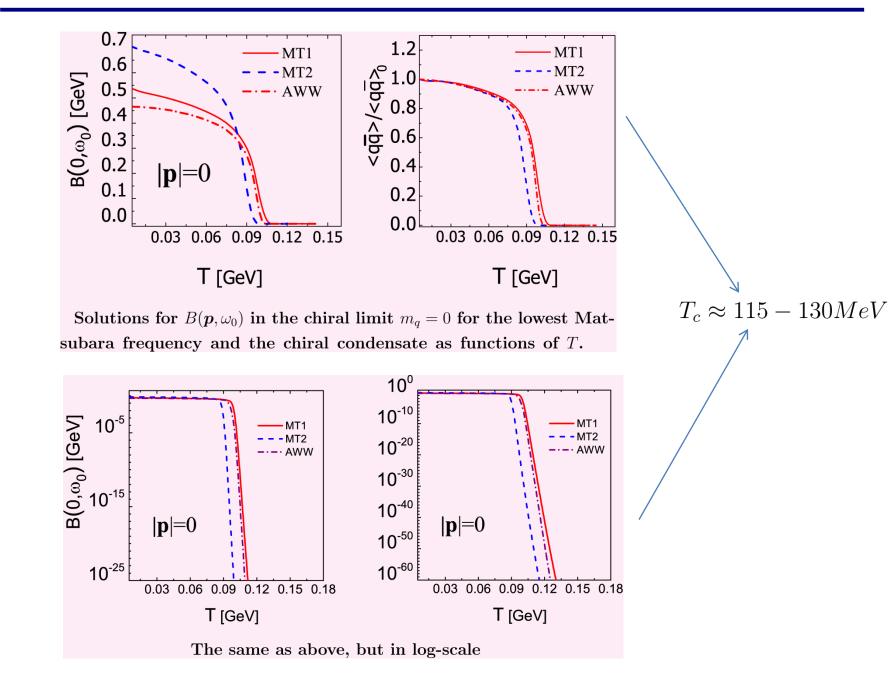
Debye mass (leading order)

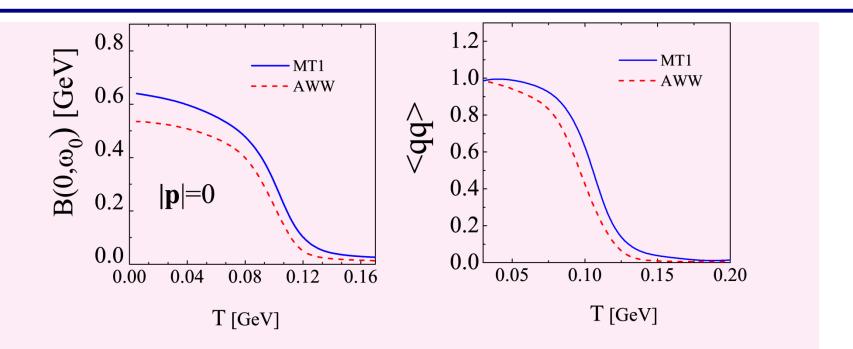
$$m_g^2 = \alpha_s \frac{\pi}{3} \left[2N_c + N_f \right] T^2, \qquad \alpha_s(E) \equiv \frac{g^2(E)}{4\pi} = 2 \frac{12\pi}{11N_c - 2N_f}$$
$$D^T(\mathbf{q}, \Omega_{mn}(T), 0) = D_{IR}(\mathbf{q}^2 + \Omega_{mn}(T)^2) + D_{UV}(\mathbf{q}^2 + \Omega_{mn}(T)^2),$$
$$D^L(\mathbf{q}, \Omega_{mn}(T), m_g(T)) = D_{IR}(\mathbf{q}^2 + \Omega_{mn}(T)^2 + m_g^2(T)) + D_{UV}(\mathbf{q}^2 + \Omega_{mn}(T)^2)$$

 $+m_{o}$

$$\begin{split} A(\boldsymbol{p}^{2},\omega_{n}^{2}) &= 1 + \frac{4}{3}T\sum_{m=-\infty}^{\infty}\int \frac{d\boldsymbol{k}}{(2\pi)^{3}} \left\{ 2\frac{\boldsymbol{q}\boldsymbol{k}}{\boldsymbol{q}^{2}} \left(1 - \frac{\boldsymbol{p}\boldsymbol{k}}{\boldsymbol{p}^{2}}\right) \boldsymbol{\sigma}_{A}(\boldsymbol{k},\omega_{m}) D^{T}(\boldsymbol{q},\Omega_{nm}) + \left[\frac{\boldsymbol{p}\boldsymbol{k}}{\boldsymbol{p}^{2}}\boldsymbol{\sigma}_{A}(\boldsymbol{k},\omega_{m}) + 2\frac{\Omega_{mn}}{q^{2}} \left(1 - \frac{\boldsymbol{p}\boldsymbol{k}}{\boldsymbol{p}^{2}}\right) \omega_{m}\boldsymbol{\sigma}_{C}(\boldsymbol{k},\omega_{m}) - \\ -2\frac{\Omega_{mn}^{2}}{q^{2}}\frac{\boldsymbol{q}\boldsymbol{k}}{\boldsymbol{q}^{2}} \left(1 - \frac{\boldsymbol{p}\boldsymbol{k}}{\boldsymbol{p}^{2}}\right) \boldsymbol{\sigma}_{A}(\boldsymbol{k},\omega_{m}) \right] D^{L}(\boldsymbol{q},\Omega_{nm},m_{g}) \right\}, \\ B(\boldsymbol{p}^{2},\omega_{n}^{2}) &= m_{q} + \frac{4}{3}T\sum_{m=-\infty}^{\infty}\int \frac{d\boldsymbol{k}}{(2\pi)^{3}} \left[D^{L}(\boldsymbol{q},\Omega_{nm},m_{g}) + 2D^{T}(\boldsymbol{q},\Omega_{nm},0) \right] \boldsymbol{\sigma}_{B}(\boldsymbol{k},\omega_{m}), \\ C(\boldsymbol{p}^{2},\omega_{n}^{2}) &= 1 + \frac{4}{3}T\sum_{m=-\infty}^{\infty}\int \frac{d\boldsymbol{k}}{(2\pi)^{3}} \left\{ 2\frac{\omega_{m}}{\omega_{n}}\boldsymbol{\sigma}_{C}(\boldsymbol{k},\omega_{m}) D^{T}(\boldsymbol{q},\Omega_{nm},0) + \\ \left[- \left(1 - 2\frac{\Omega_{mn}^{2}}{\boldsymbol{q}^{2}}\right) \frac{\omega_{m}}{\omega_{n}}\boldsymbol{\sigma}_{C}(\boldsymbol{k},\omega_{m}) + 2\frac{\boldsymbol{q}\boldsymbol{k}}{\boldsymbol{q}^{2}}\frac{\Omega_{nm}}{\omega_{n}}\boldsymbol{\sigma}_{A} \right] D^{L}(\boldsymbol{q},\Omega_{nm},m_{g}) \right\}, \end{split}$$

$$\sigma_F(\boldsymbol{k},\omega_m) = \frac{F(\boldsymbol{k},\omega_m)}{\boldsymbol{k}^2 A^2(\boldsymbol{k},\omega_m) + \omega_m^2 C^2(\boldsymbol{k},\omega_m) + B^2(\boldsymbol{k},\omega_m)}$$
(1)



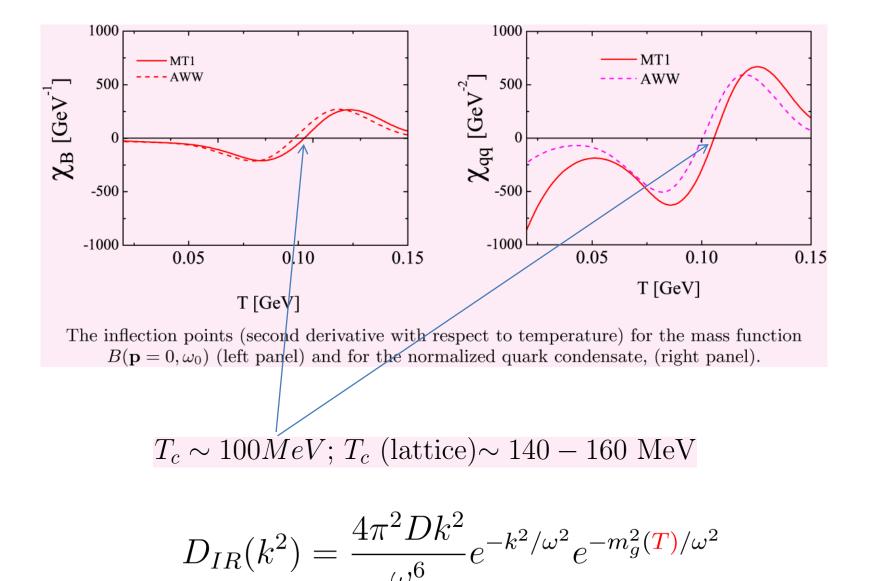


The solutions $B(\mathbf{p} = 0, \omega_0)$ (left panel) and the quark condensate (right panel) for the light quark $m_l = 5$ MeV for the lowest Matsubara frequency.

Order parameters? One can use the inflection point of $B(0,\omega_0)$ and of $\langle q\bar{q} \rangle$, i.e. the maximum of the corresponding derivative with respect to the temperature

$$\chi_B(T) = \frac{d^2 B(0, \omega_0)}{dT^2}; \quad \chi_{qq}(T) = \frac{d^2 \langle q\bar{q} \rangle}{dT^2}.$$

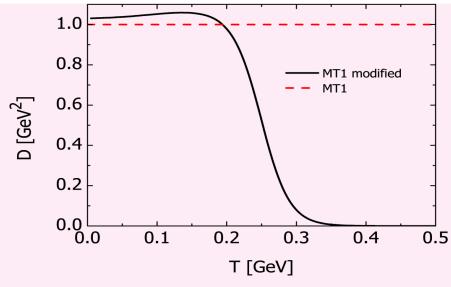
Then the (pseudo) critical temperature T_c is fixed by the condition that $\chi_B(T)|_{T=T_c} = 0$ and/or $\chi_{qq}(T)|_{T=T_c} = 0$.



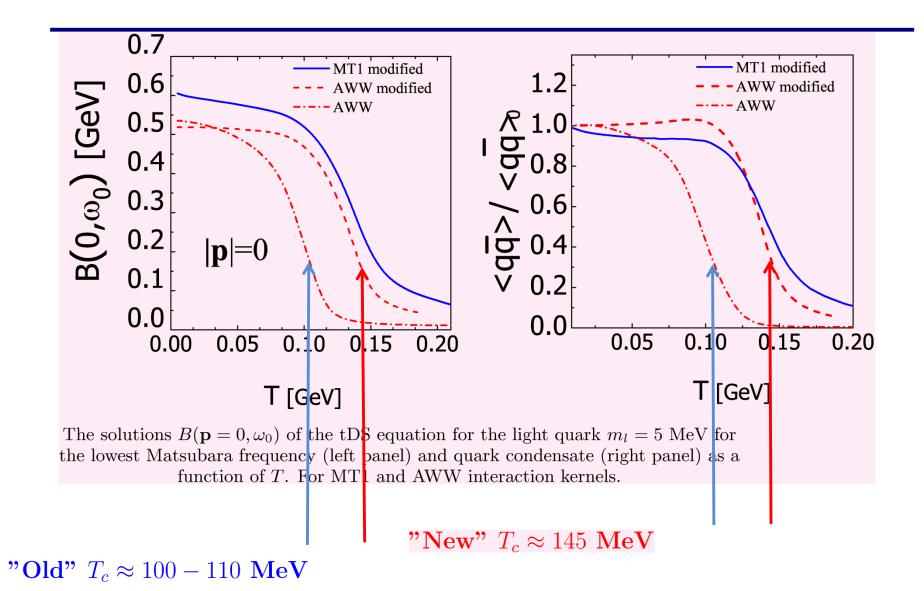
(i) in the IR term the Debye mass is omitted,
(ii) the parameter D receives a T-dependence,
(iii) the UV term, being inspired by perturbative QCD calculations, remains unchanged,

$$D_{IR}(k^2) = \frac{4\pi^2 D(\mathbf{T})k^2}{\omega^6} e^{-k^2/\omega^2}, \qquad D_{UV}(k^2 + \mathbf{m_g}^2) = \frac{8\pi^2 \gamma_m F(k^2 + \mathbf{m_g}^2)}{\ln[\tau + (1 + \frac{k^2 + \mathbf{m_g}^2}{\Lambda_{QCD}^2})^2]}.$$

$$D(T) = D \left[a \left\{ 1 + \tanh\left(-\frac{T - T_p}{\beta}\right) \right\} + b \left\{ 1 - \tanh\left(-\frac{T - T_p}{\beta'}\right) \right\} \exp[-\alpha^2 T^2] \right],$$



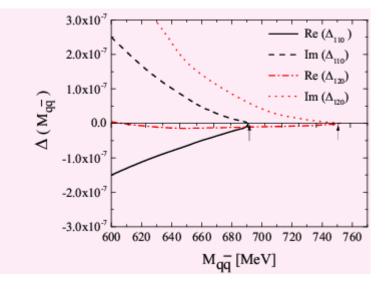
An illustration of a possible dependence of the IR term D of the MT1 model on the temperature T.



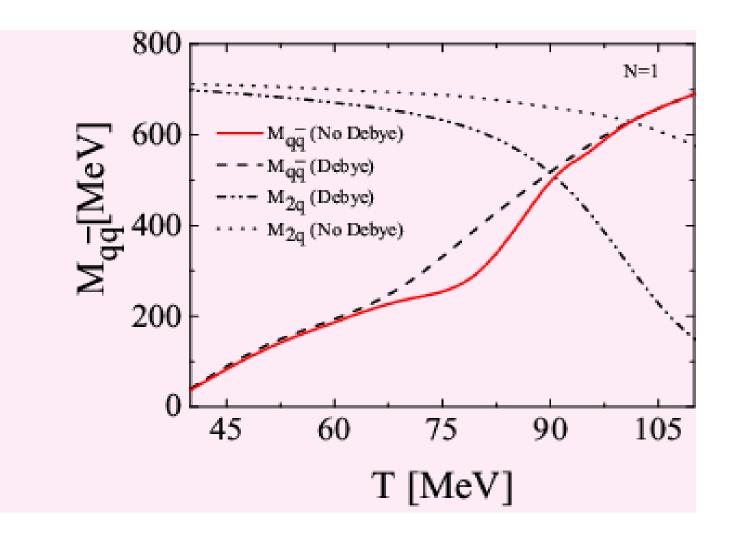
Lattice $T_c = (154 \pm 9)$ MeV (H.-T. Ding et al, Int. J. Mod. Phys. E 24 (2015) 1530007.) Bethe-Salpeter equation for pseudoscalar particles (non-zero T) $\tilde{\Gamma}(P_N, p_n) = \frac{4}{3}T \sum_m \int \frac{d^3q}{(2\pi)^3} \gamma_\mu S^{(+)}(1) \tilde{\Gamma}(P_N, q_m) \tilde{S}^{(-)}(2) \gamma_\nu D_{\mu\nu}(\kappa_{mn}),$

$$\kappa_{mn} = (\boldsymbol{p} - \boldsymbol{q}, \omega_n - \omega_m)$$

$$\tilde{\Gamma}(P_N, p_n) = (g_1(\boldsymbol{p}, \omega_n) \frac{1}{2} \hat{I} + g_2(\boldsymbol{p}, \omega_n) \frac{1}{2} \gamma_4 + ig_3(\boldsymbol{p}, \omega_n) \frac{1}{2} \frac{\vec{\gamma} \boldsymbol{p}}{|\boldsymbol{p}|} + ig_4(\boldsymbol{p}, \omega_n) \frac{1}{2} \frac{\vec{\gamma} \boldsymbol{P}}{|\boldsymbol{P}|}) \gamma_5$$



Solving of the tBS equation



S.M. Dorkin



> The considered model, based on the Dyson-Schwinger-Bethe-Salpeter equations with only two-three effective parameters, describes fairly well the vacuum (T=0) properties (masses, electroweak decay constants...) of the scalar, pseudoscalar, vector etc. mesons, and allows for a Poincarè covariant study of processes with mesons

> Within such effective models one can investigate the analytical structure of the quark propagators related to such fundamental characteristics of QCD as confinement and dynamical chiral symmetry breaking phenomena encoded in the chiral condensate

> A direct generalization of the model for **finite** temperatures demonstrates that it still provides qualitatively descriptions of critical phenomena in hot matter, however qiantitavely the critical temperatures Tc relevant to possible signals of **QGP** are underestimate in comparison with lattice calculation results. This is a clear indication that the interaction kernel must receive an additional dependence on temperature.

> We propose a T-dependence of the interaction kernel, which suppresses the IR part at high temperatures and which provides (pseudo-) critical temperatures close to those from lattice calculation. The solutions of the BS-equation at large T (T>100MeV) give dissociation instability against fragmentation into state of two quasi free quarks.



SOLUTION BSE

