

Status of neutrino oscillations

Yury Kudenko

Institute for Nuclear Research, Moscow

Baldin ISHEPP XXIV, Dubna, Russia,
September 17-22, 2018



OUTLINE

□ Neutrino oscillations

- 3-neutrino scheme
- running accelerator and reactor experiments
- future projects

□ Light sterile neutrinos

- neutrino anomalies
- new experimental tests



ν oscillations and mixing

Standard Model: neutrinos are *massless* particles

3 families

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix}$$

U parameterization:

three mixing angles θ_{12} θ_{23} θ_{13}
 CP violating phase δ_{CP}

atmospheric

link between
atmospheric and solar

solar

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

SuperK, K2K,
MINOS, T2K, NovA(?)

T2K
MINOS

Daya Bay, RENO
Double Chooz

Solar experiments, SuperK
KamLAND

$$\theta_{23} \sim 45^\circ$$

$$\theta_{13} \approx 9^\circ$$

$$\theta_{12} \approx 34^\circ$$

$$|\Delta m_{32}^2| \cong |\Delta m_{31}^2| =$$

$$|\Delta m_{atm}^2| \approx 2.4 \times 10^{-3} \text{ eV}^2$$

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$

$$\Delta m_{12}^2 + \Delta m_{23}^2 + \Delta m_{31}^2 = 0$$

$$\Delta m_{21}^2 = \Delta m_{sol}^2 \approx 7.5 \times 10^{-5} \text{ eV}^2$$

two independent Δm^2



Main goals of oscillation experiments

- CP violation in lepton sector

Strength of CP violation in neutrino oscillations

$$J_{CP} = \text{Im}(U_{e1} U_{\mu 2} U_{e2}^* U_{\mu 1}^*) = \text{Im}(U_{e2} U_{\mu 3} U_{e3}^* U_{\mu 2}^*) \\ = \cos\theta_{12} \sin\theta_{12} \cos^2\theta_{13} \sin\theta_{13} \cos\theta_{23} \sin\theta_{23} \sin\delta_{CP}$$

all mixing angles $\neq 0 \rightarrow$
 $\rightarrow J_{CP} \neq 0$ if $\delta_{CP} \neq 0$

First indication from T2K: $\delta_{CP} = -\pi/2$

neutrinos

$$V_{MNS} \sim \begin{pmatrix} 0.8 & 0.5 & 0.2 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

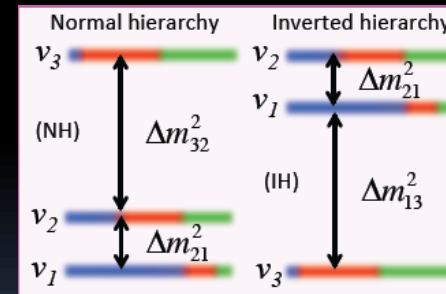
quarks

$$V_{CKM} \sim \begin{pmatrix} 1 & 0.2 & 0.001 \\ 0.2 & 1 & 0.01 \\ 0.001 & 0.01 & 1 \end{pmatrix}$$

Quark sector $J_{CP} \approx 3 \times 10^{-5}$

Lepton sector $J_{CP} \sim 0.02 \times \sin\delta_{CP}$

- Neutrino mass hierarchy



- θ_{23} – maximal? If not, what octant ($\theta_{23} > \pi/4$ or $\theta_{23} < \pi/4$)?

Neutrino cross sections

- Sterile neutrinos



Experimental methods

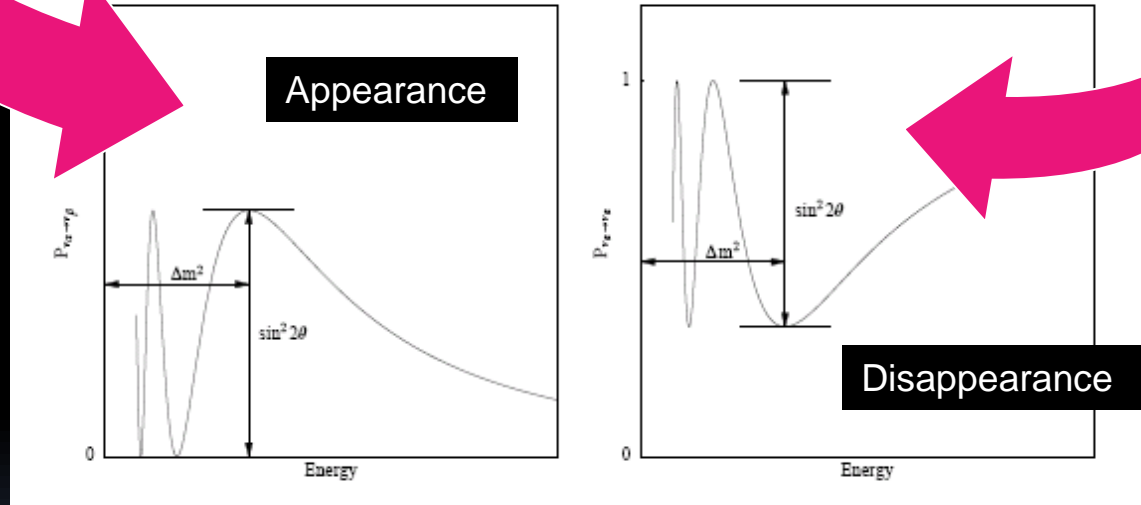
$$P(\nu_\alpha \rightarrow \nu_\beta) = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2 \Phi_{ij} \mp 2 \sum_{i>j} \text{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin 2\Phi_{ij}$$

$$P_{\nu_\alpha \rightarrow \nu_\beta} = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E} \right),$$

$$P_{\nu_\alpha \rightarrow \nu_\alpha} = 1 - \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E} \right)$$

Appearance

Disappearance



Search for CP violation in neutrino oscillations

Matter effect

$$A_{CP} = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} \cong \frac{\Delta m_{12}^2 L}{4E_\nu} \cdot \frac{\sin 2\theta_{12}}{\sin \theta_{13}} \cdot \sin \delta$$

Mass Hierarchy

Current experiments



about 500 members
59 institutions
from 11 countries

LONG-BASELINE NEUTRINO OSCILLATION EXPERIMENT



Super-K

Toyama
Kamioka Mine



JPARC

Tokai

Tokyo

Tokyo/Narita Airport

JAPAN



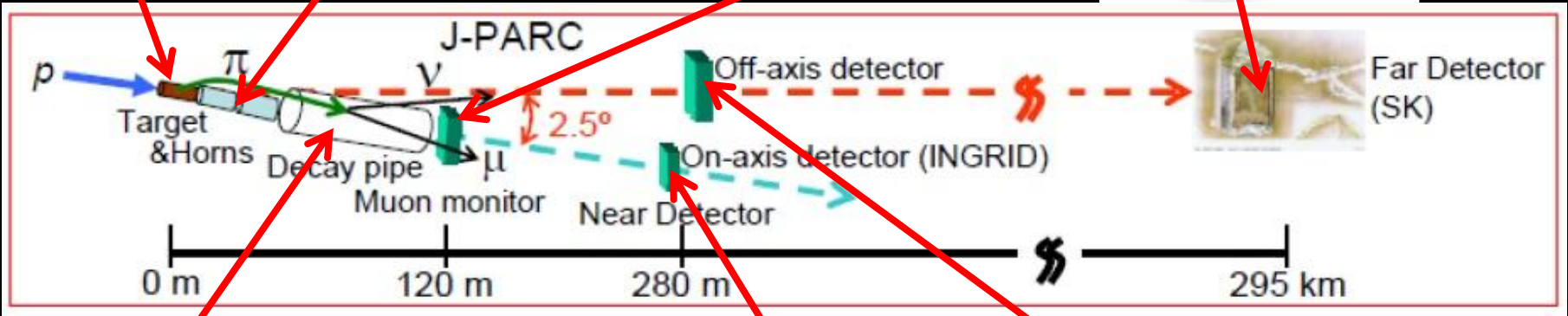
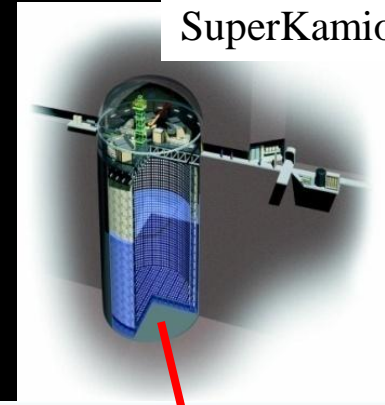
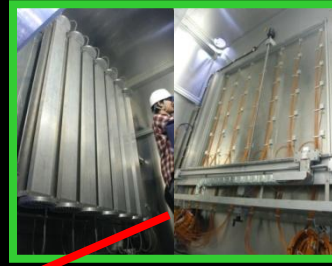
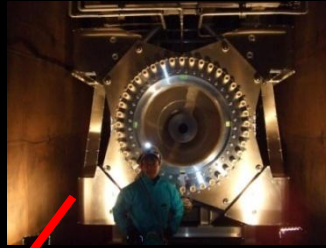
Y.Kudenko



T2K experiment

Collect data since 2010

Far neutrino detector
SuperKamiokande

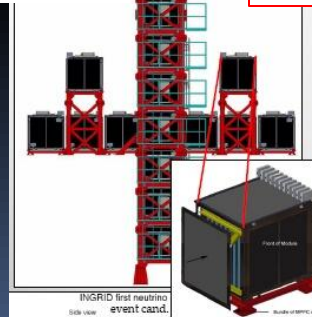
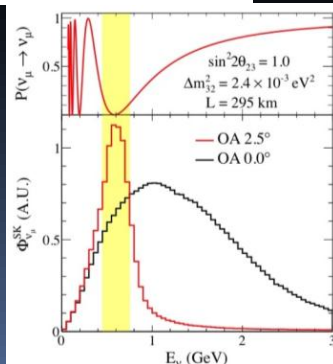


Off-axis neutrino beam

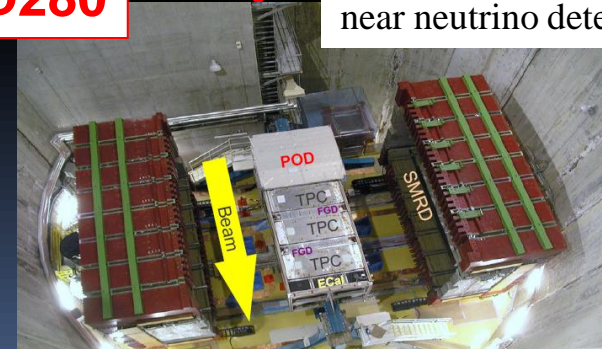
Neutrino monitor INGRID

ND280

Off-axis near neutrino detector



INGRID first neutrino event candid.





T2K data

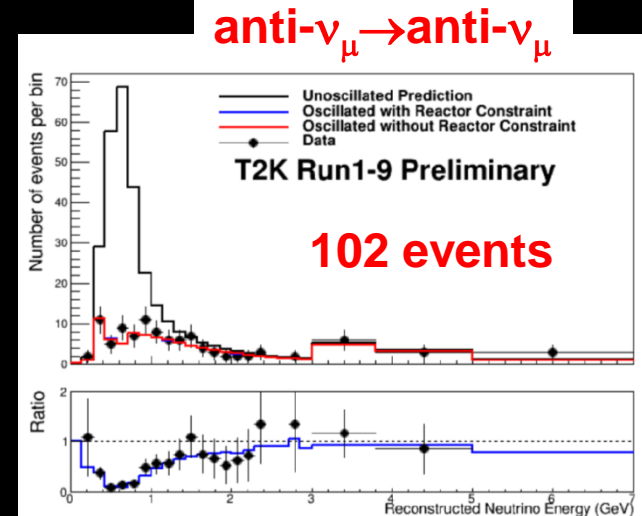
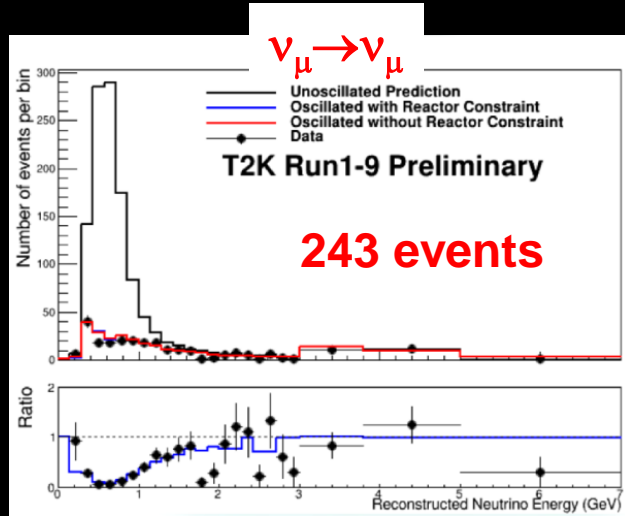
P.Litchfield, ICHEP2018

Neutrino mode

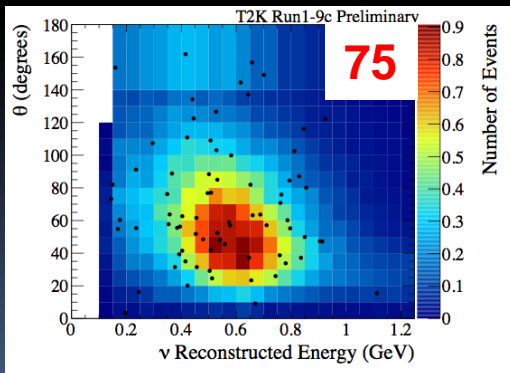
1.49×10^{21} POT

Antineutrino mode

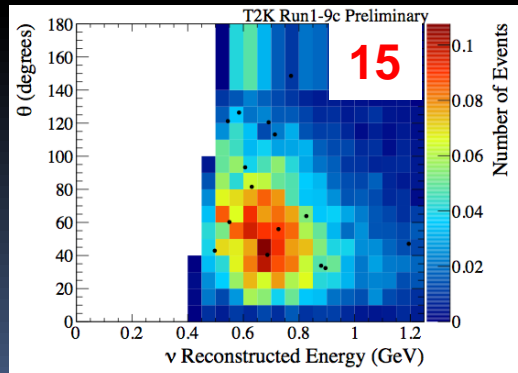
1.12×10^{21} POT



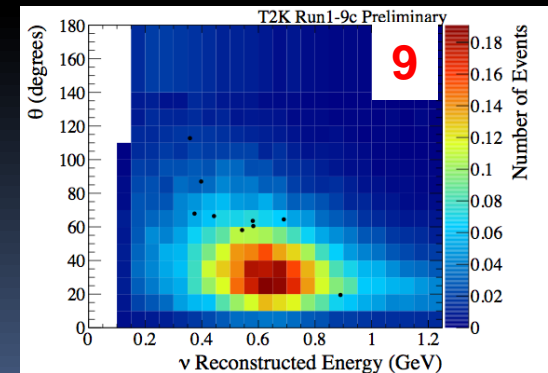
$\nu_{\mu} \rightarrow \nu_e$



$\nu_{\mu} \rightarrow \nu_e + 1\pi$



$\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$





T2K data and expectation

Event rate			
Beam mode	Not Oscillated	Oscillated (maximal mixing)	Observed
neutrino	1211.4	268.2	243
antineutrino	314.3	95.3	102

Systematic error		
Beam mode	w/o ND280	ND280 constrained
neutrino	14.5% \longleftrightarrow	4.9%
antineutrino	12.2% \longleftrightarrow	4.3%

Sample	Expectation, $\sin^2 \theta_{23} = 0.528, \delta =$				Observed
	$-\pi/2$	0	π	$+\pi/2$	
FHC 1R- μ	268.5	268.2	268.9	268.9	243
RHC 1R- μ	95.5	95.3	95.8	95.5	102
<i>Sum of 1R-μ</i>	364.0	363.5	364.7	364.5	345
FHC 1R- e	73.8	61.6	62.2	50.0	75
FHC 1R- e + d.e.	6.9	6.0	5.8	4.9	15
RHC 1R- e	11.8	13.4	13.2	14.9	9

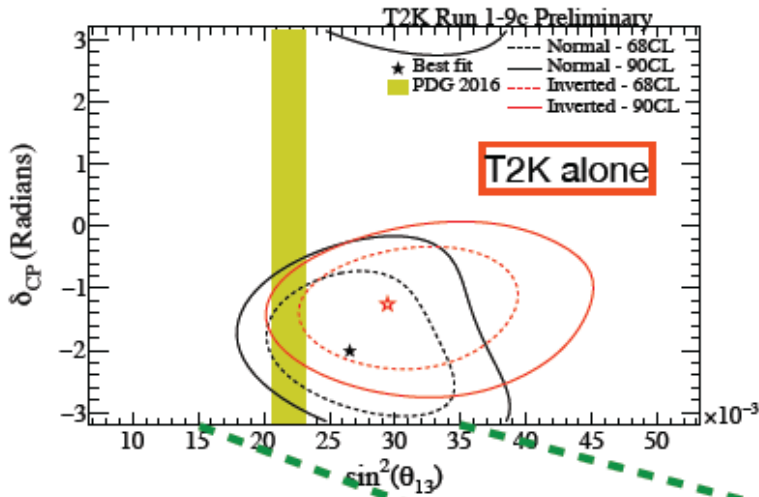
disappearance

appearance

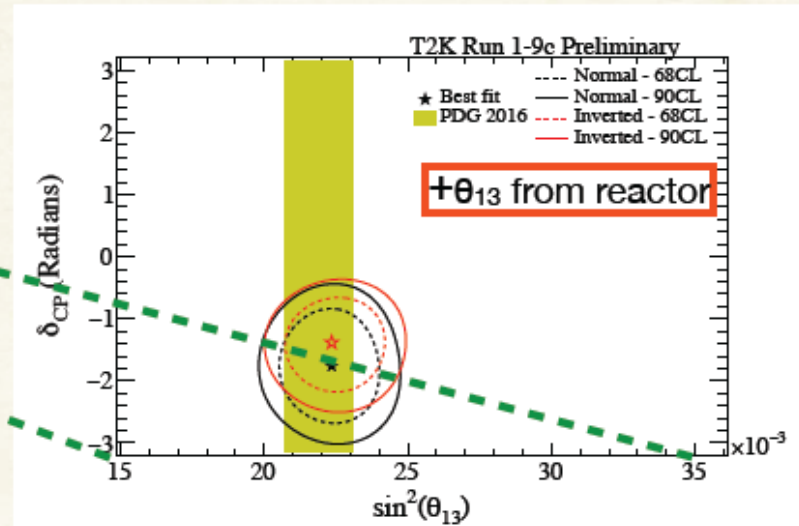


T2K result

T2K $\nu_e / \text{anti-}\nu_e$



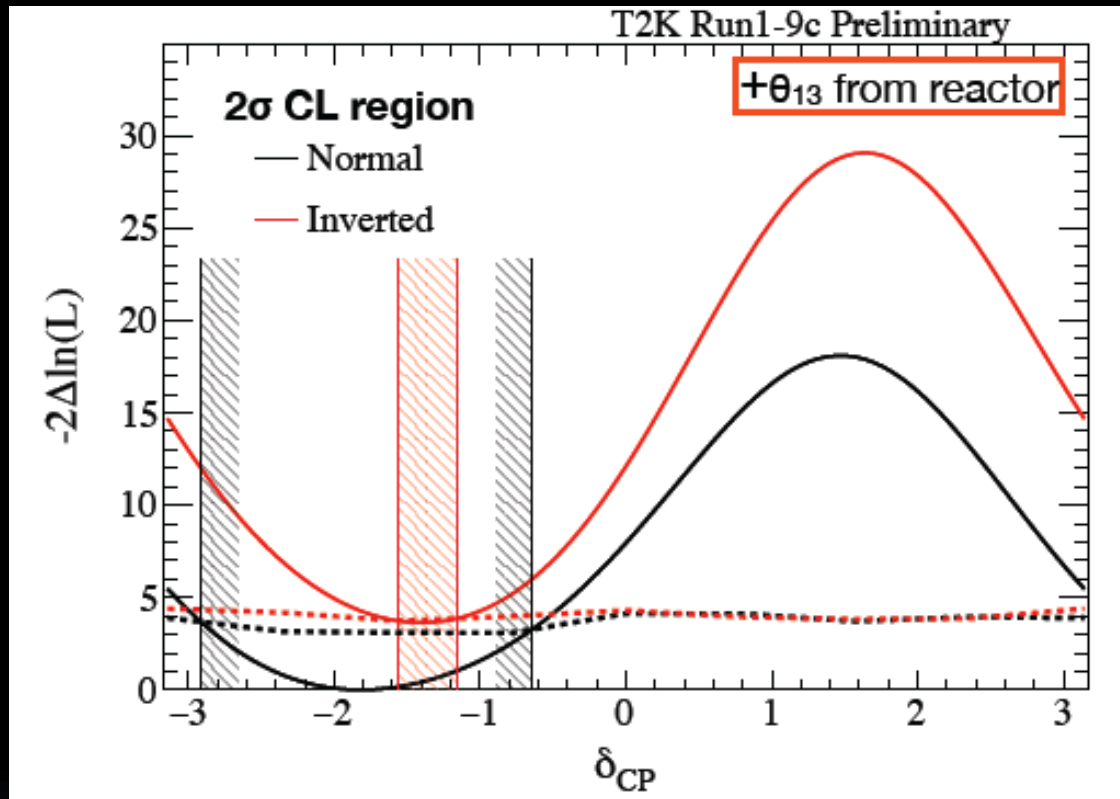
- **Constraint on δ_{CP} with T2K data alone**
- **Tighter constraint with θ_{13} value from reactor**



T2K $\nu_e / \text{anti-}\nu_e$ + reactor θ_{13}



T2K result



Best fit
 $\delta_{cp} = -1.6$ rad
for NH

CP-conservation hypothesis ($\sin\delta_{CP} = 0, \pi$) excluded at 2σ level

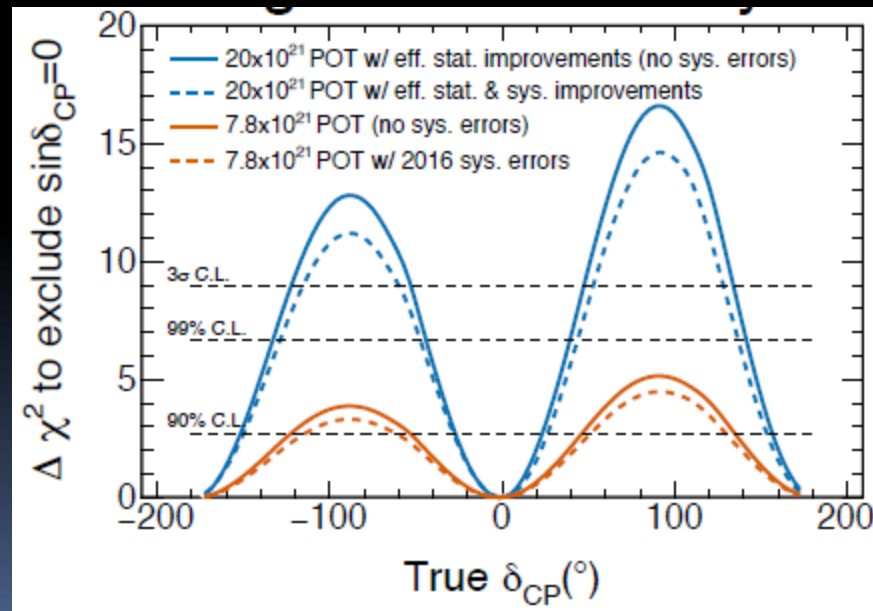
- **First hint for CP violation in the lepton sector**
- **T2K data favour $\delta_{CP} \sim -\pi/2$ and normal hierarchy**



Future plans

T2K expected to accumulate **7.8×10^{21} POT** around 2021
(now **3×10^{21} POT**)

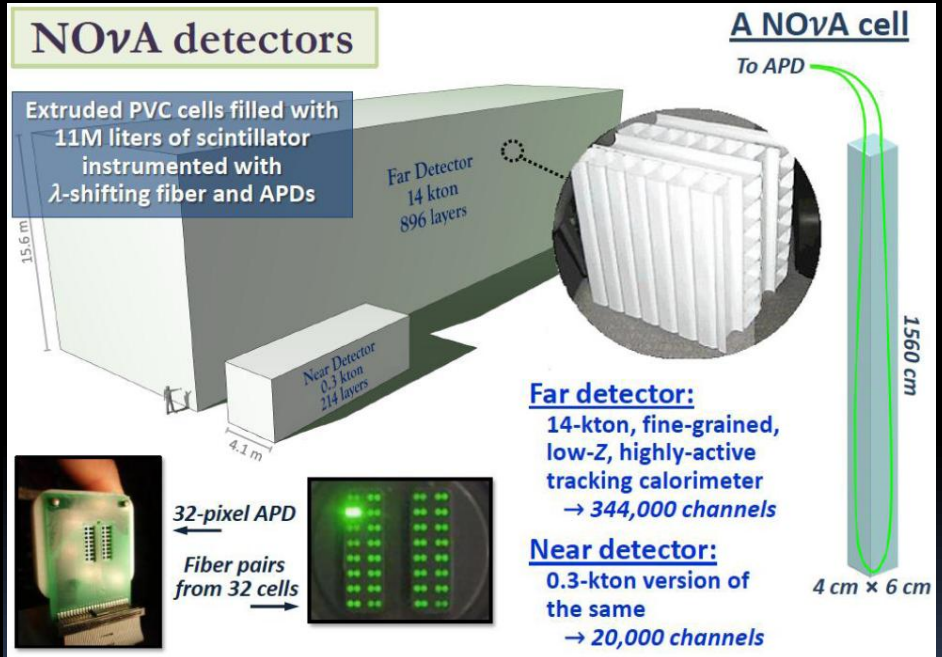
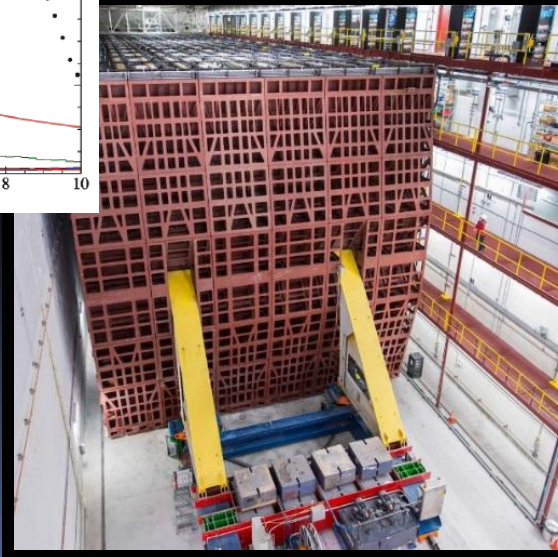
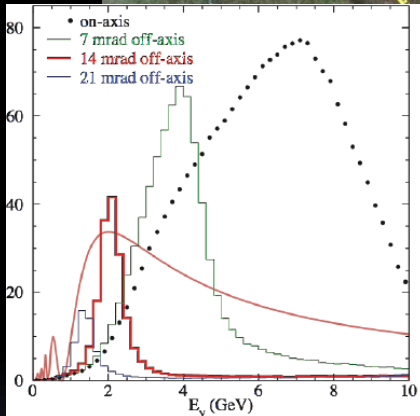
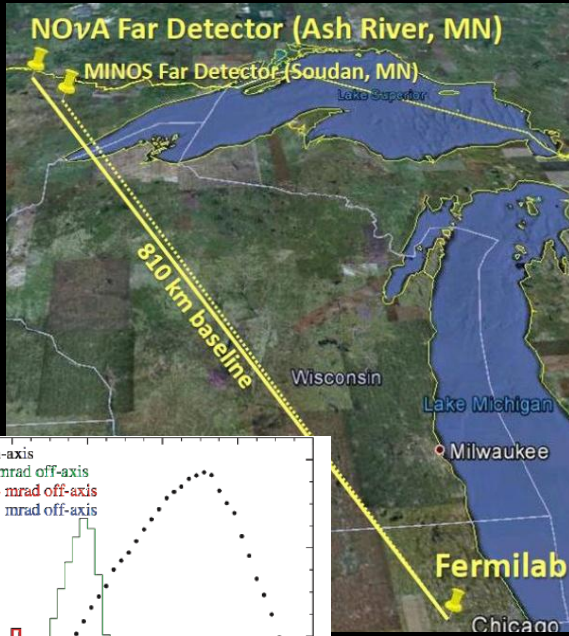
- Upgrade of near detectors to improve systematic uncertainties **18% (2011) \rightarrow 9% (2014) \rightarrow 5% (2018) \rightarrow goal \leq 4% (2021)**
- Plan to increase the beam intensity up to 1 MW in 2021
- Beam power up to 1.3 MW in \sim 2026
- T2K-II: proposed extension up to 2026 for **20×10^{21} POT**
 3σ sensitivity to CP violation for $\delta_{CP} \sim -\pi/2$





NOvA

Neutrino beam from FNAL to Ash River
Baseline 810 km
Neutrino beam 14 mrad off-axis
Far detector : 14 kt fine-grained calorimeter
65% active mass
Near Detector: 0.3 kt fine-grained calorimeter



Taking data since Summer 2014
 Study of $\nu_\mu \rightarrow \nu_\mu$ and $\nu_\mu \rightarrow \nu_e$ oscillations



NOvA: $\nu_{\mu} \rightarrow \nu_{\mu}$

J.Bian ICHEP2018

Neutrino beam: 8.85×10^{20} POT

Antineutrino beam: 6.9×10^{20} POT

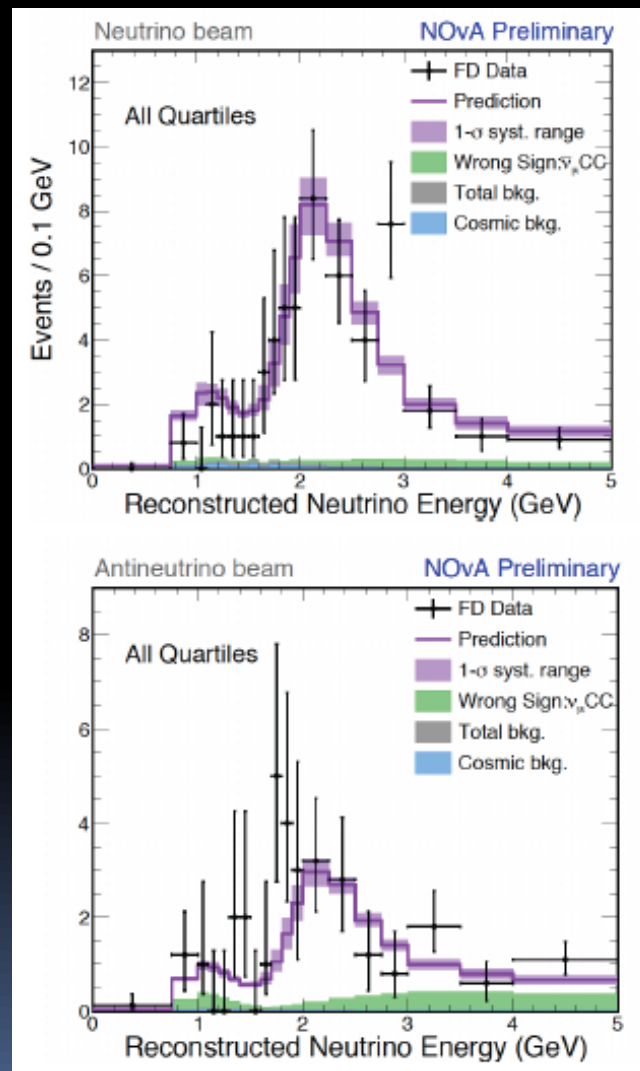
Far detector

Neutrino beam:

- Observe 113 events
- Expect $730 +38/-49$ (syst.) w/o oscillations

Antineutrino beam:

- Observe 65 events
- Expect $266 +12/-14$ (syst.) w/o oscillations



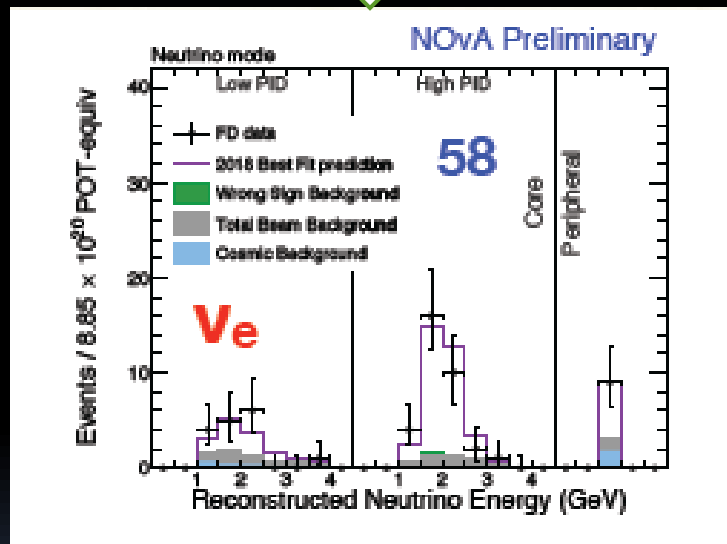


NOvA: ν_e /anti- ν_e

ν_e

58 events observed

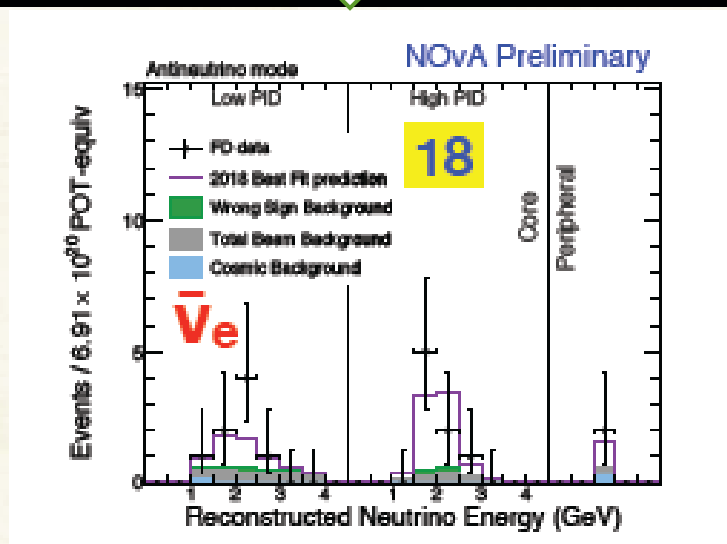
15 background events expected



anti- ν_e

18 events observed

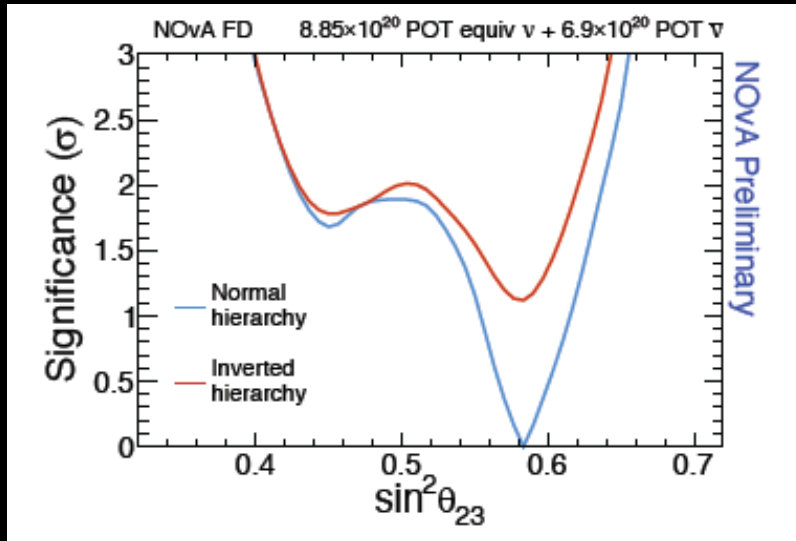
5.3 background events expected



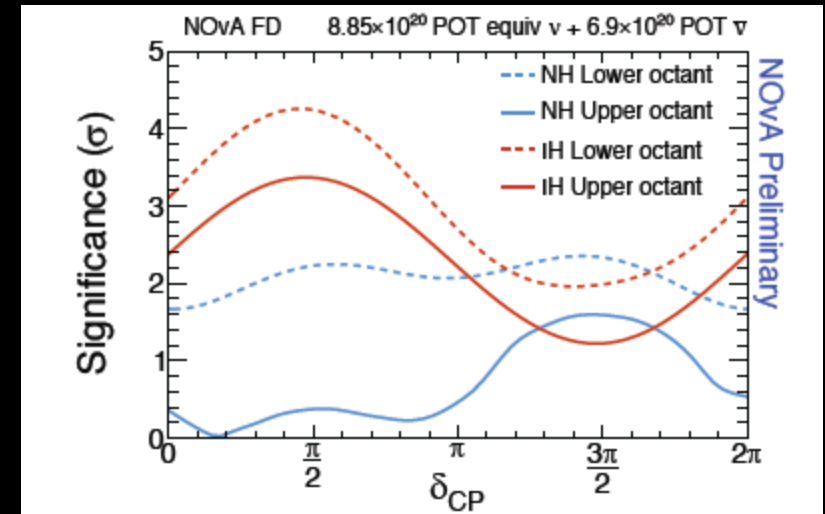
$\bar{\nu}_e$ appearance $> 4\sigma$



NOvA results



**NOvA prefers
Normal Hierarchy at 1.8σ**

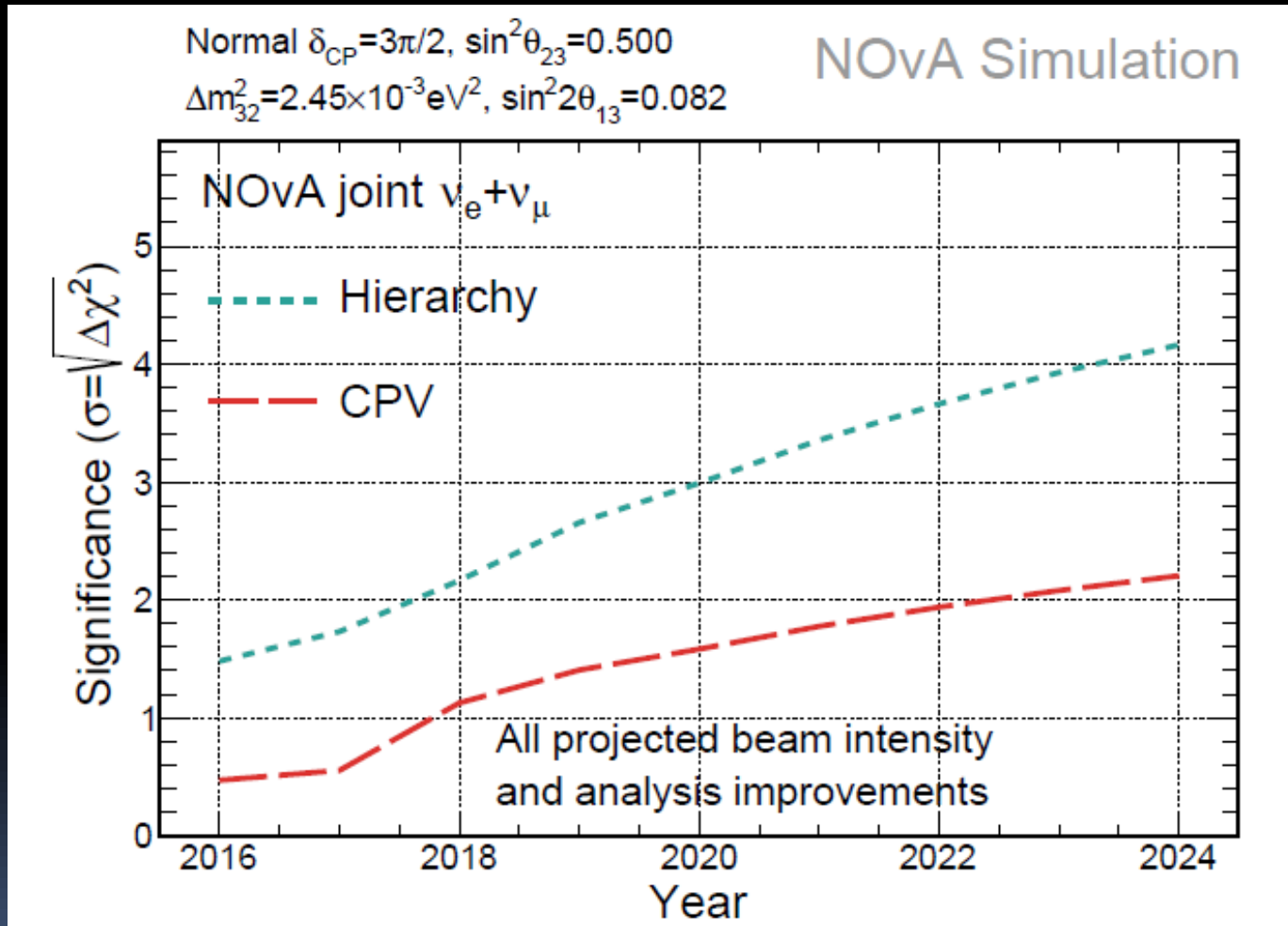


Best fit:

- Normal Hierarchy
- $\delta_{CP} = 0.17\pi$ but consistent with all δ_{CP} values at $<1.6\sigma$



Prospects for NOvA



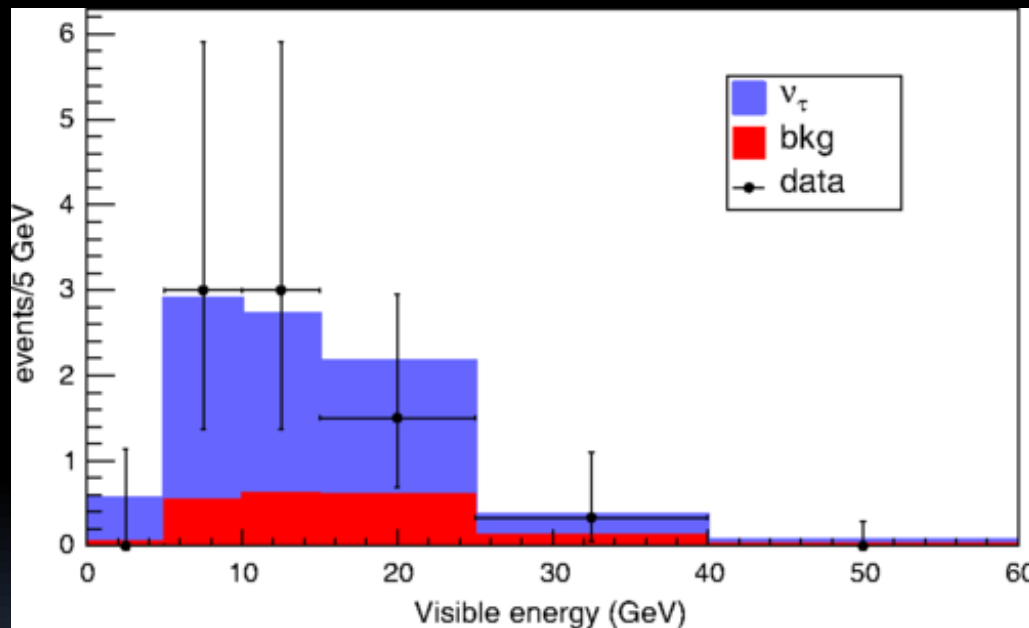


OPERA: final result

PRL 120 (2018) 211801

$\nu_\mu \rightarrow \nu_\tau$ appearance

10 ν_τ events observed for 18×10^{19} POT
Expected 6.4 events for $\Delta m^2_{23} = 2.5 \times 10^{-3} \text{ eV}^2$, $\sin^2 2\theta_{23} = 1.0$
Expected background 2.0 ± 0.4 events



Significance of ν_τ appearance 6.1σ

OPERA: $\Delta m^2_{23} = (2.7 + 0.7 - 0.6) \times 10^{-3} \text{ eV}^2$, assuming $\sin^2 2\theta_{23} = 1.0$



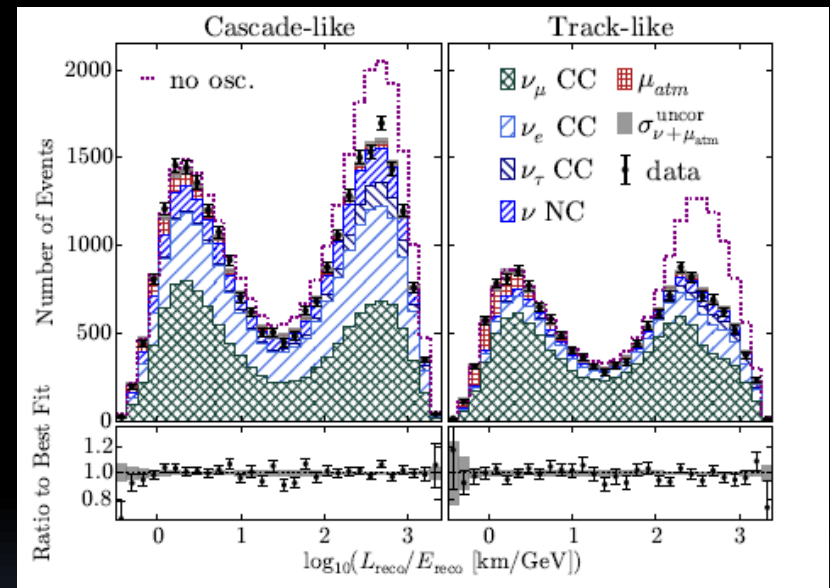
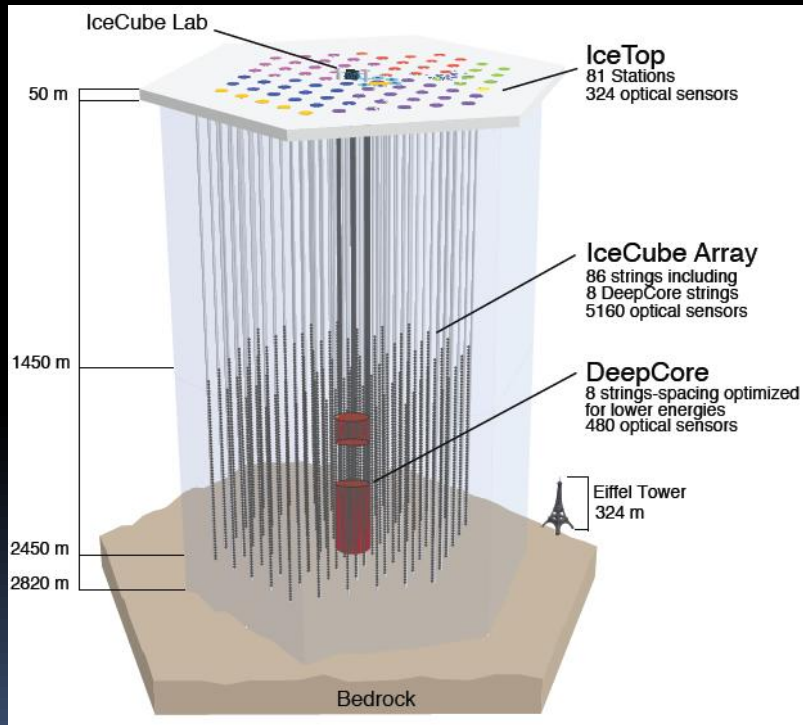
IceCube

Neutrinos have the first maximum of disappearance at about 25 GeV

Energy threshold of Deep Core = 5 GeV

PRL 120 (2018) 071801

Data taking for 3 years

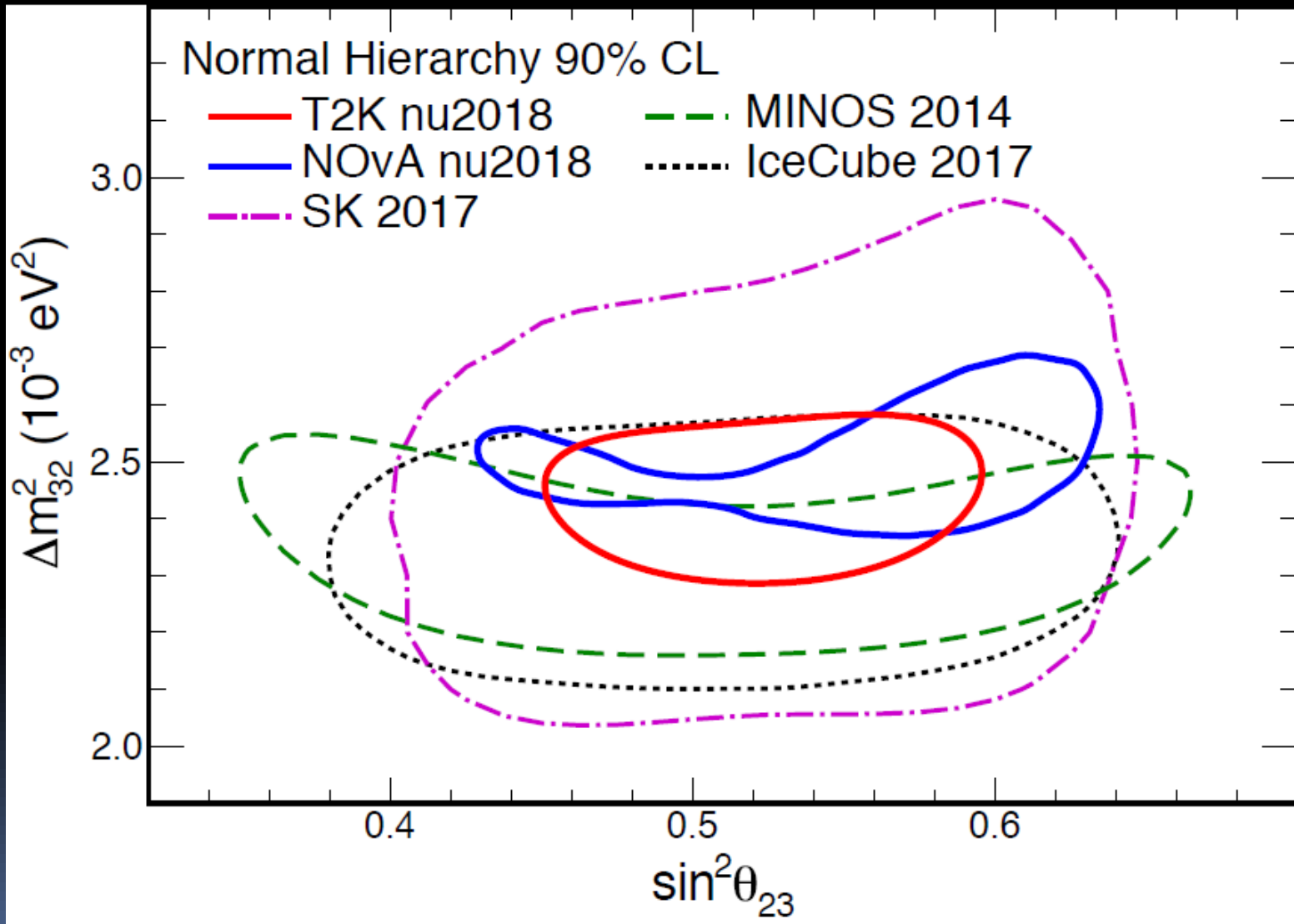


$$\Delta m_{32}^2 = (2.31 +011-0.13) \times 10^{-3} \text{ eV}^2 \quad \sin^2 \theta_{23} = 0.51 +0.07 -0.09 \text{ for NH}$$



Oscillation parameters: $\Delta m^2_{32} - \sin^2\theta_{23}$

M.Yokoyama ICHEP2018





Reactor experiments

Measurement of θ_{13}

Daya Bay, China



17.4 GW

RENO, Korea



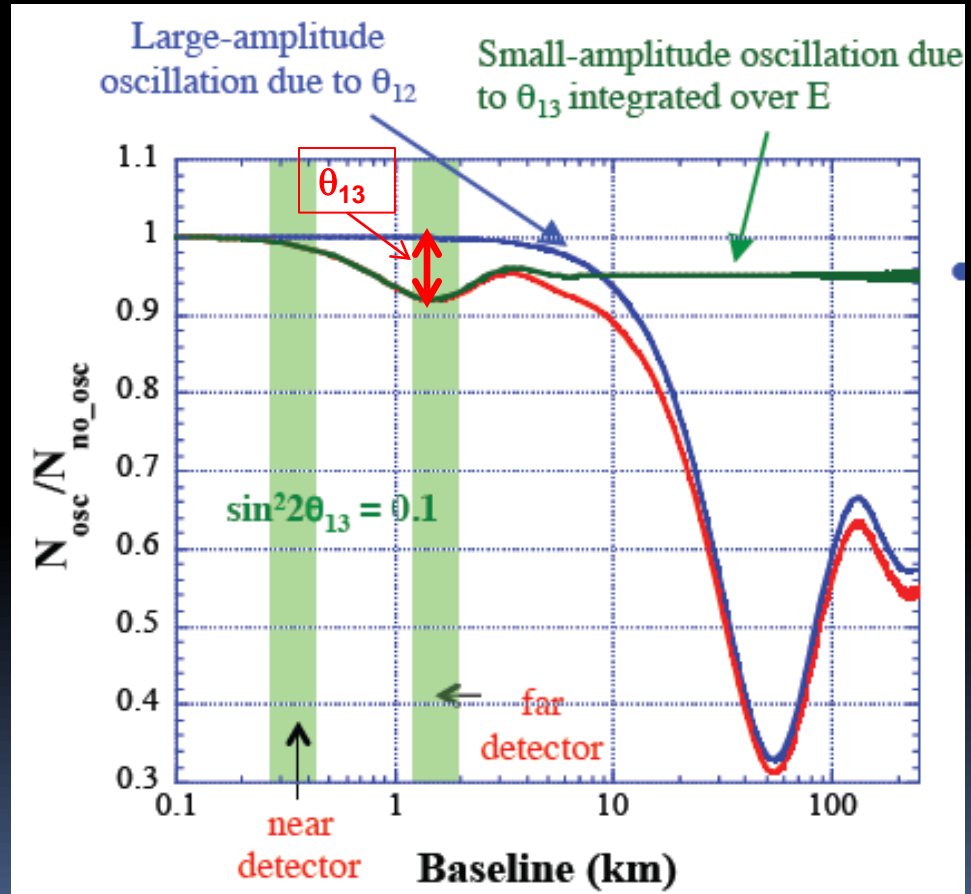
16 GW

Double Chooz, France



8.5 GW

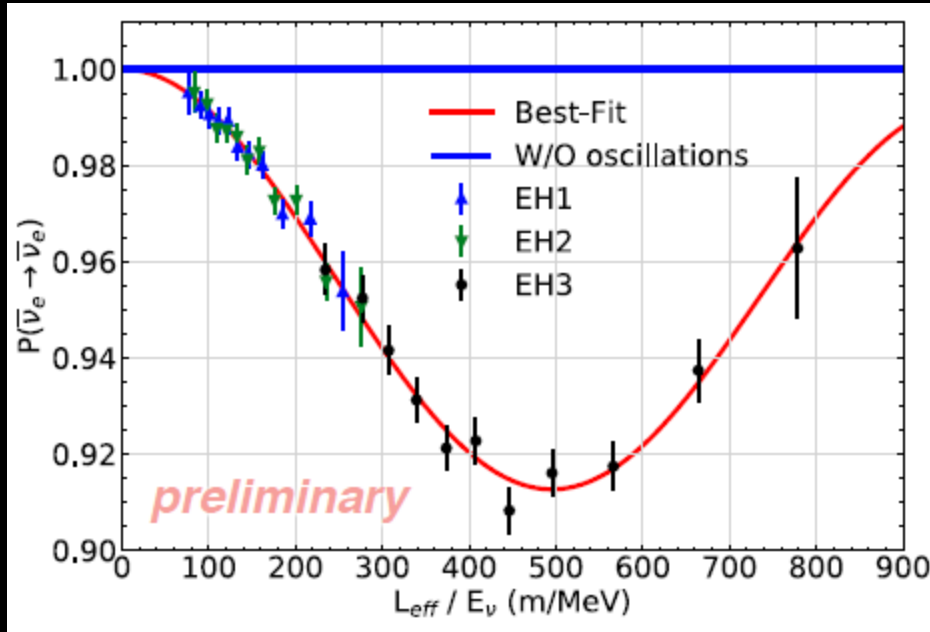
$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{ee}^2 L}{4E} \right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta m_{21}^2 L}{4E} \right)$$



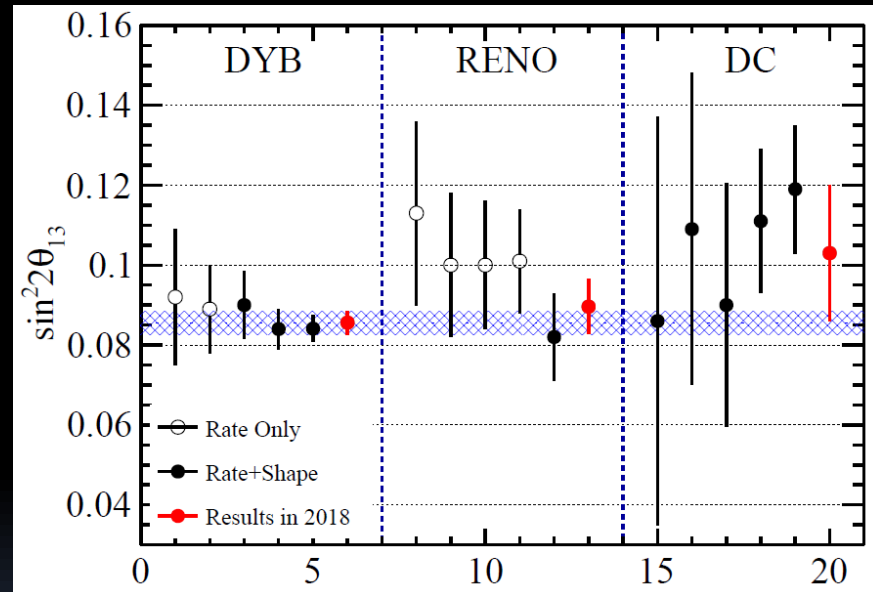


Oscillation results

Daya Bay



Liang Zhan, ICHEP2018



$$\sin^2 2\theta_{13} = 0.0856 \pm 0.0029$$

$$|\Delta m_{ee}^2| = (2.52 \pm 0.07) \times 10^{-3} \text{ eV}^2$$

Future LBL Projects

- Reactor experiment JUNO
- Accelerator LBL experiment DUNE
- HyperKamiokande and T2HK



Reactor experiment JUNO

China



66 institutions
> 400 collaborators

Main target:
Measurement of
neutrino mass hierarchy

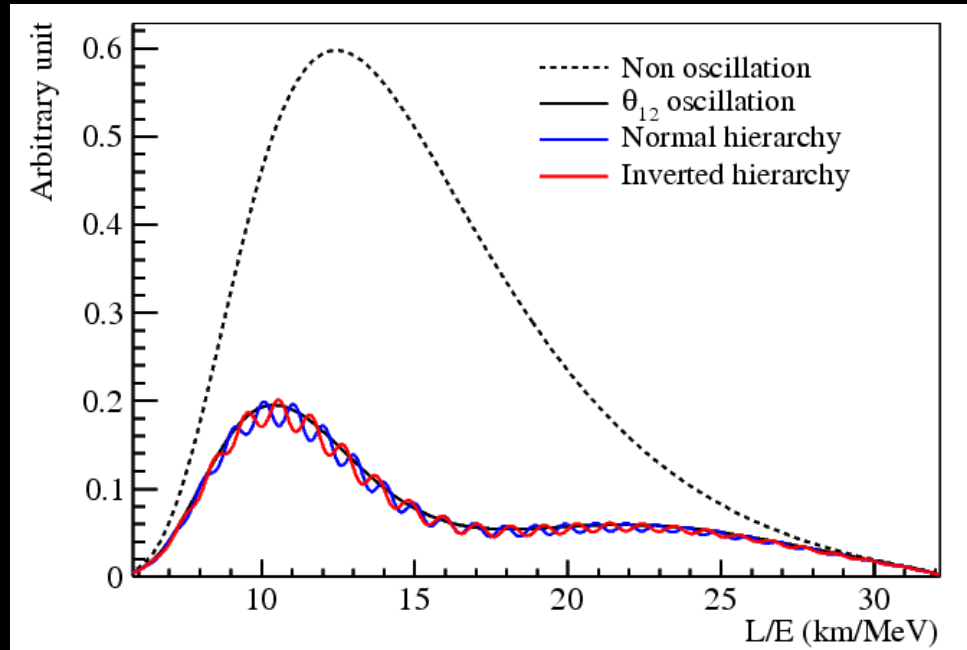
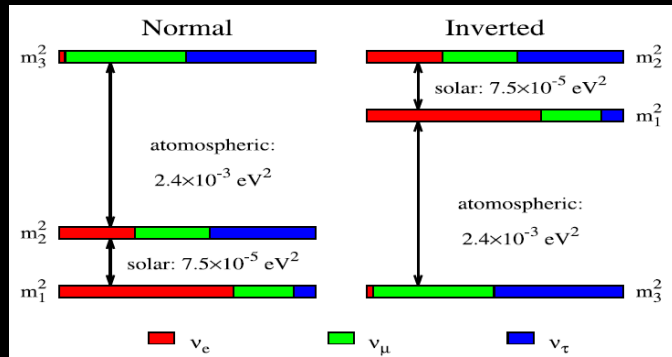
- 700 m deep underground
- 36 GW reactor power
- 53 km baseline -> **oscillation maximum θ_{12}**
- 20 kton LS detector
- **3%** energy resolution at 1MeV
- **<1%** energy scale uncertainty

!?



JUNO goals

Main goal: determination of neutrino mass hierarchy



FRD 88, 013008(2013)	Hierarchy discrimination power	With info on $\Delta m_{\mu\mu}^2$ from LBL expts
Statistics only	4 σ	5 σ
Realistic case	3 σ	4 σ

Oscillation Parameter	Current accuracy (global 1 σ) **	Dominant experiment(s)	JUNO Potentiality
Δm_{21}^2	2.3%	KamLAND	0.59%
$\Delta m^2 = m_3^2 - \frac{1}{2}(m_1^2 + m_2^2) $	1.6%	MINOS, T2K	0.44%
$\sin^2(\theta_{12})$	~4-6%	SNO	0.67%

Supernova neutrino
+ Geoneutrinos
Solar neutrinos



Detector JUNO

Requirements:

- PMT coverage 75% of total surface
- QE ~ 35%
- Sci. att. length >20 m

Calibration

Top Tracker

Central detector

Acrylic sphere
20kt Liquid Scin
~17000 20'' PMT
~36000 3'' PMT

Water Cherenkov

~2000 20'' PMT

h=44 m

3'' PMT

d=43.5 m

20'' PMT



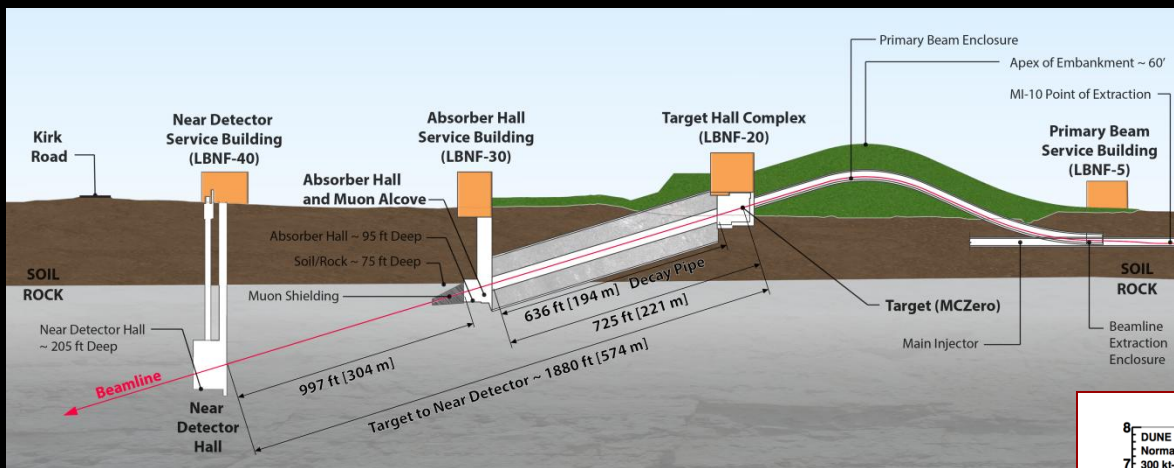
LBNF/DUNE Project

Flagship FNAL project

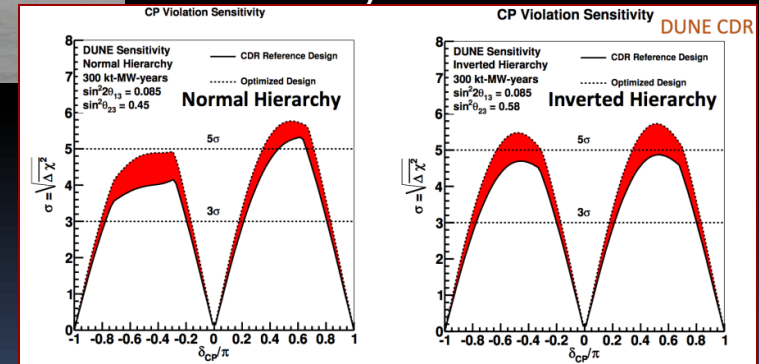
- Main goals:**
- discovery of CP violation in leptonic sector
 - neutrino mass hierarchy at $>5\sigma$ level
 - neutrino astronomy
 - proton decay search

30 countries
 161 institutions
 ~1000 collaborators

$E_p = 60-120$ GeV
 Beam power 1.2 \rightarrow 2.4 MW
 On axis neutrino beam
 $E_\nu \sim 1-6$ GeV
 L=1300 km from FNAL to SURF, S.Dakota

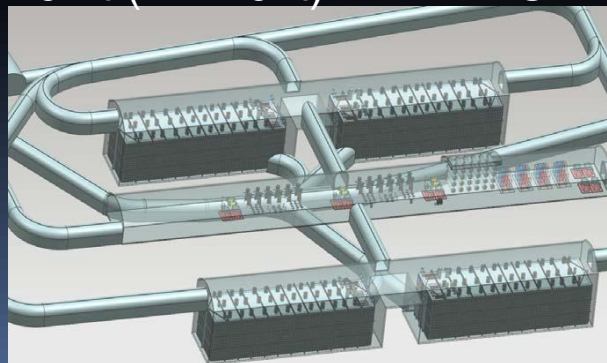


Sensitivity to CP violation



Far detector 40 kt (4 x 10kt) LAr TPC

Single
and
Dual
phase
detectors

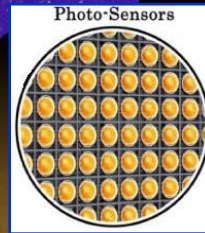
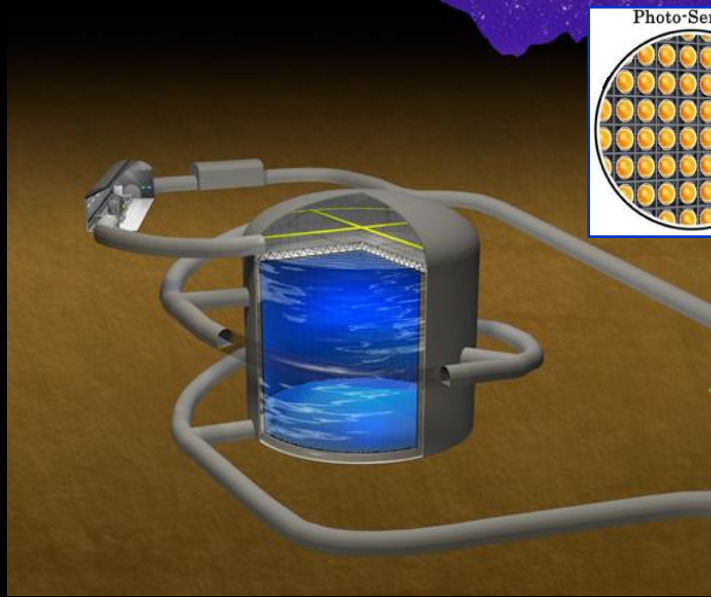


- 2021 – installation of 1st far detector
- 2024 – 2 modules operational
- 2026 – deliver neutrino beam



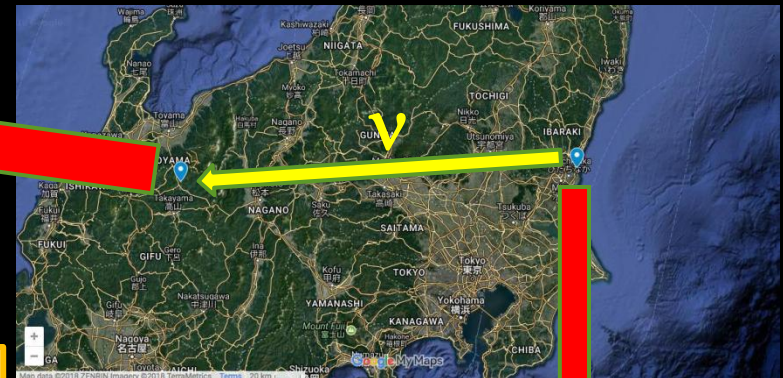
Hyper-Kamiokande project

Hyper-K water



Main goals:

- Search for CP violation
- Proton decay
- Neutrino astrophysics



J-PARC

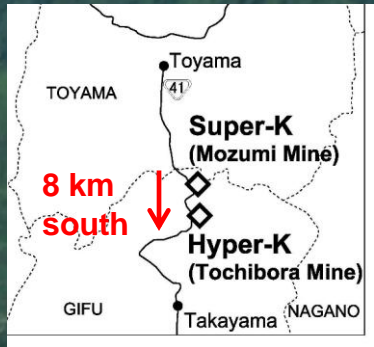


Water tank 60 m(H) x 74m(D)
Total volume 260 kt
Fiducial volume 190 kt ~10xSuper-K
40000 50 cm ID PMTs
PMT coverage 40%
6700 20 cm OD PMT's
Photon sensitivity ~2 times better than Super-K
Construction of 2nd tank in Korea
(1-3 deg off axis, 2nd oscill. maximum) is under study

Mt. Ikeno-yama

SK

1000 m

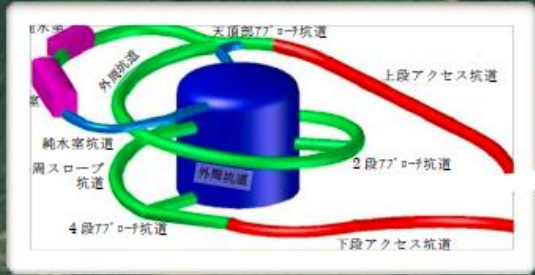


Maruyama



Excavated rock disposal site

Mt. Nijyugo-yama



650 m

HK



Tunnel Entrance

Route 41

Funatsu Bridge

Kamioka Town

Wasabo

Google



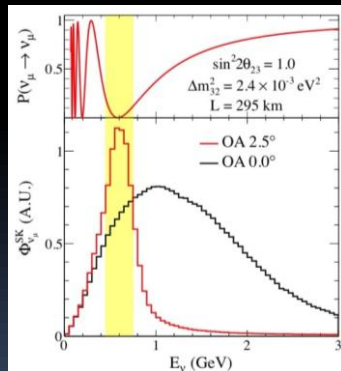
T2HK (Tokai-to-Hyper-Kamiokande)

Hyper-K

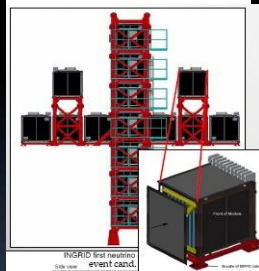
J-PARC



Off-axis neutrino beam



Neutrino monitor INGRID



Off-axis near neutrino detector



Decay tunnel



Horn



Target



Neutrino beam elements

Near neutrino detector at 280 m from target

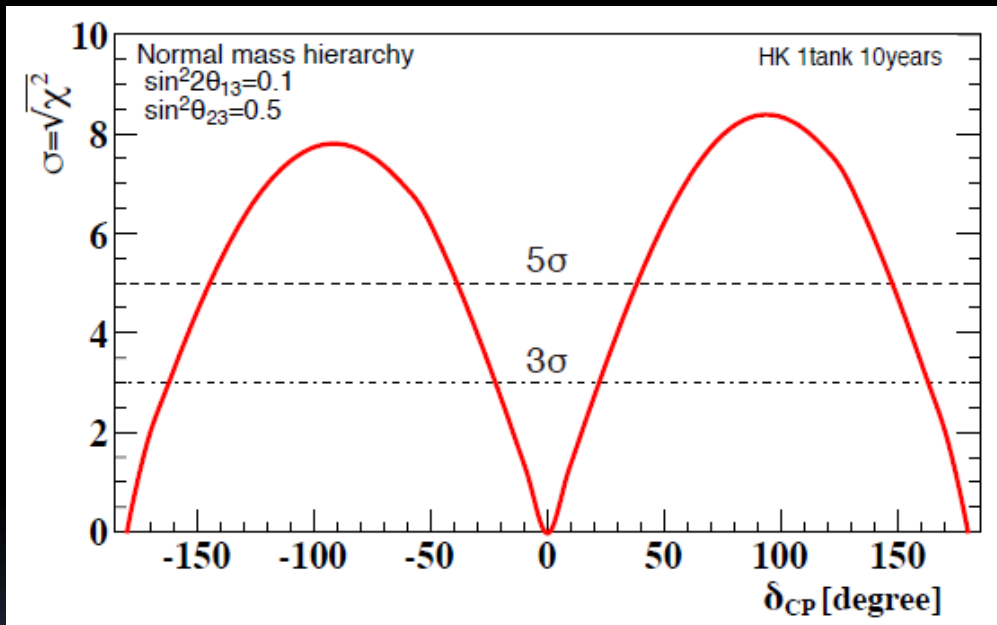
J-PARC neutrino beam

2.5° off-axis, peak energy 600 MeV (oscillation maximum), current beam power 485 kW



Sensitivity to CP

Integrated beam power $1.3 \text{ MW} \times 10^8 \text{ s}$
→ 2.7×10^{22} POT with 30 GeV proton beam
 $\sin^2 2\theta_{13} = 0.1$



Hyper-K: uncertainties of expected number events

$\nu_{\mu} \rightarrow \nu_e$	3.2%
$\nu_{\mu} \rightarrow \nu_{\mu}$	3.6%
<hr/>	
$\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$	3.9%
$\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{\mu}$	3.6%

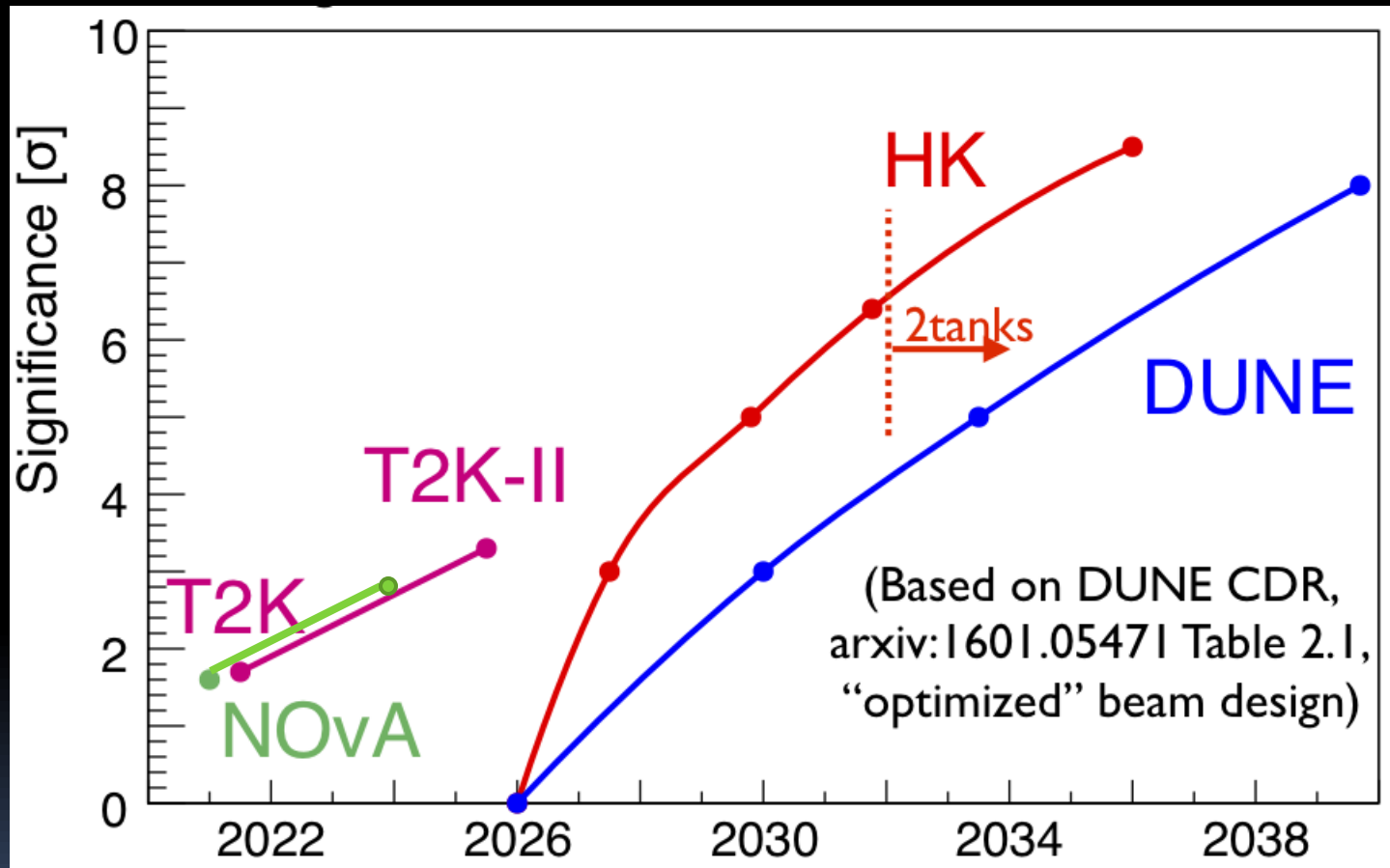
T2K
systematic
uncertainties
5-6 %

**Exclusion of $\delta=0$ at 8σ (for $\delta = -\pi/2$)
 5σ (3σ) significance for 57 (80)% of possible δ values**



Expected sensitivity to CP

Significance for $\delta_{CP} = -\pi/2$





Start of Hyper-Kamiokande

- Seed funding is allocated within 2019 fiscal year in Japan
- Construction will start in April 2020



The University of Tokyo
Hongo, Bunkyo-ku, Tokyo 113-8654, Japan

September 12th, 2018

Concerning the Start of Hyper-Kamiokande

Seed funding towards the construction of the next-generation water Cherenkov detector Hyper-Kamiokande has been allocated by the Ministry of Education, Culture, Sports, Science and Technology (MEXT) within its budget request for the 2019 fiscal year. Seed fundings in the past projects usually lead to full funding in the following year, as it was the case for the Super-Kamiokande project.

The University of Tokyo pledges to ensure construction of the Hyper-Kamiokande detector commences as scheduled in April 2020. The University of Tokyo has made this decision in recognition of both the project's importance and value both nationally and internationally.

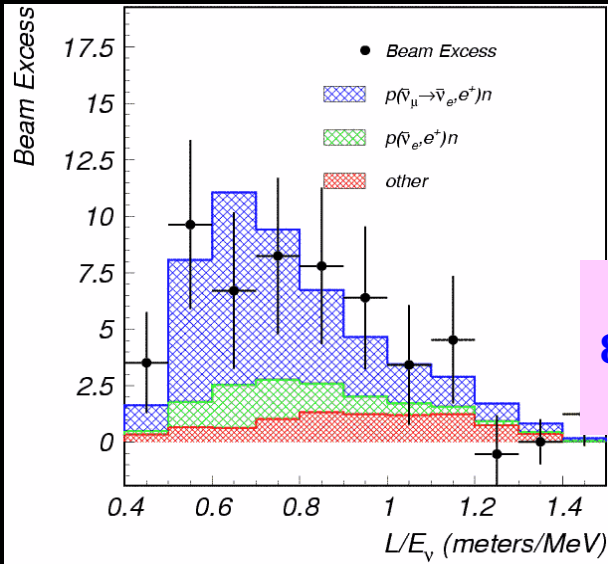
The neutrino research that led to Nobel prizes for Special University Professor Emeritus Koshiba and Distinguished University Professor Kajita has entered a new era. The international community has demonstrated the need for Hyper-Kamiokande. The considerable expertise and achievements of the University of Tokyo and Japan, and unique and invaluable contributions from national and international collaborators will ensure the project will make significant contributions to the intellectual progress of the world.

Makoto Gonokami
President, The University of Tokyo

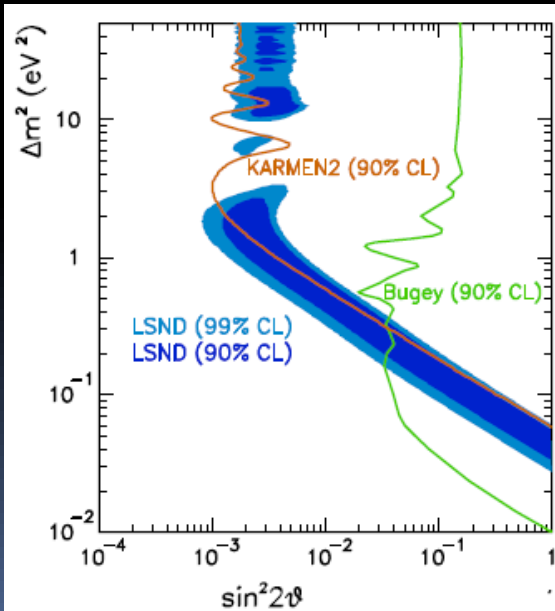
Light sterile neutrinos



LSND

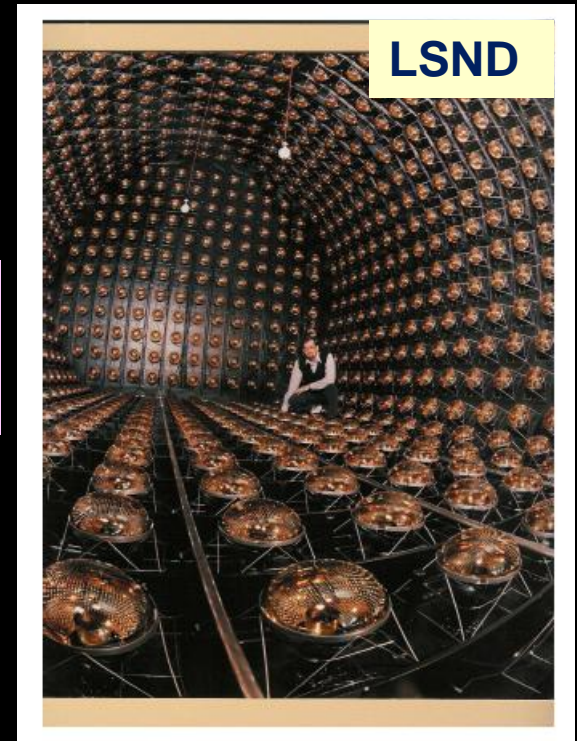


$\text{anti-}\nu_{\mu} \rightarrow \text{anti-}\nu_e$
 $87.9 \pm 22.4 \pm 6.0$ events
 Excess 3.8σ



$$\begin{aligned}
 P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e) &= \sin^2(2\theta) \sin^2\left(\frac{1.27 L \Delta m^2}{E}\right) \\
 &= 0.245 \pm 0.067 \pm 0.045 \%
 \end{aligned}$$

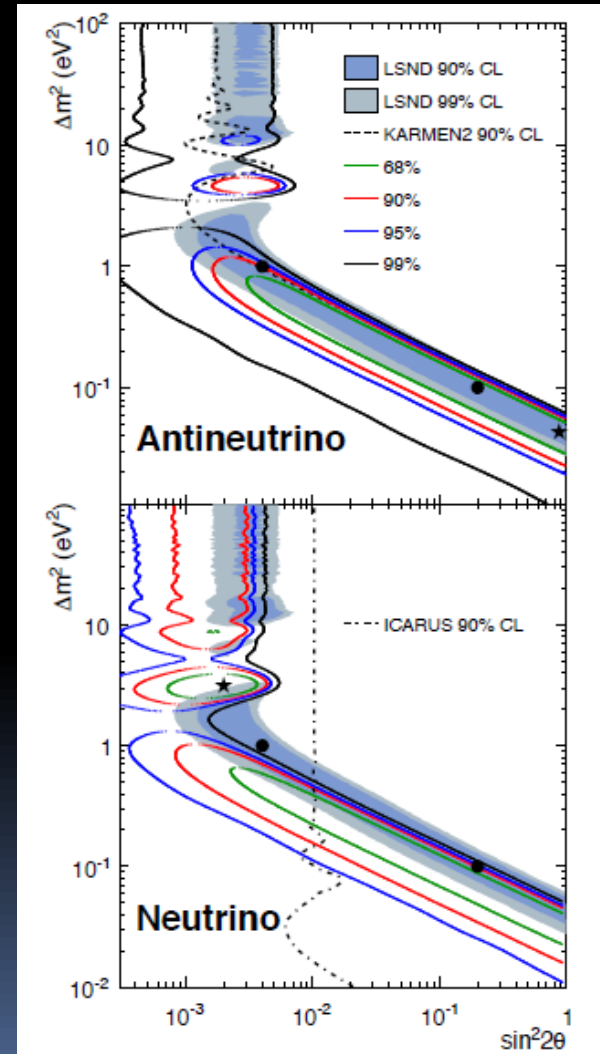
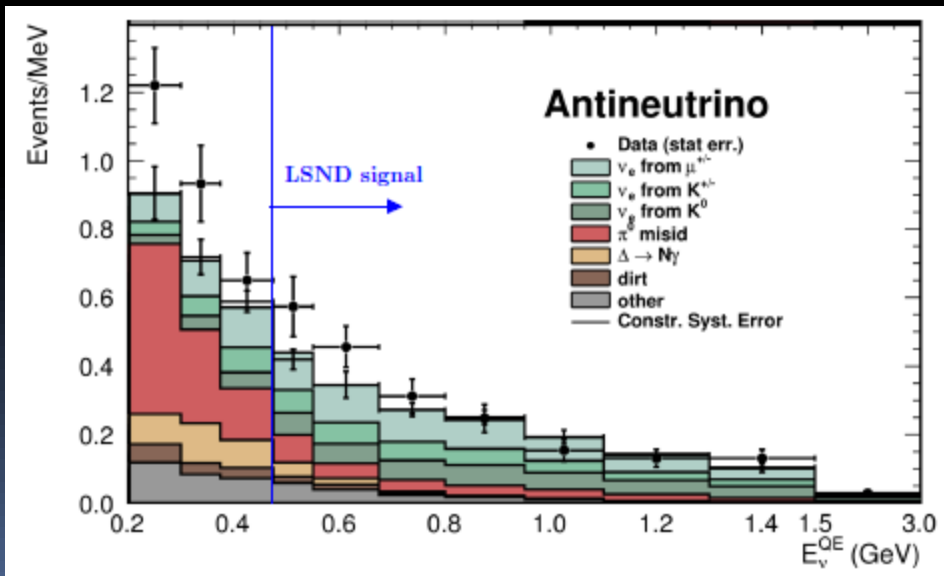
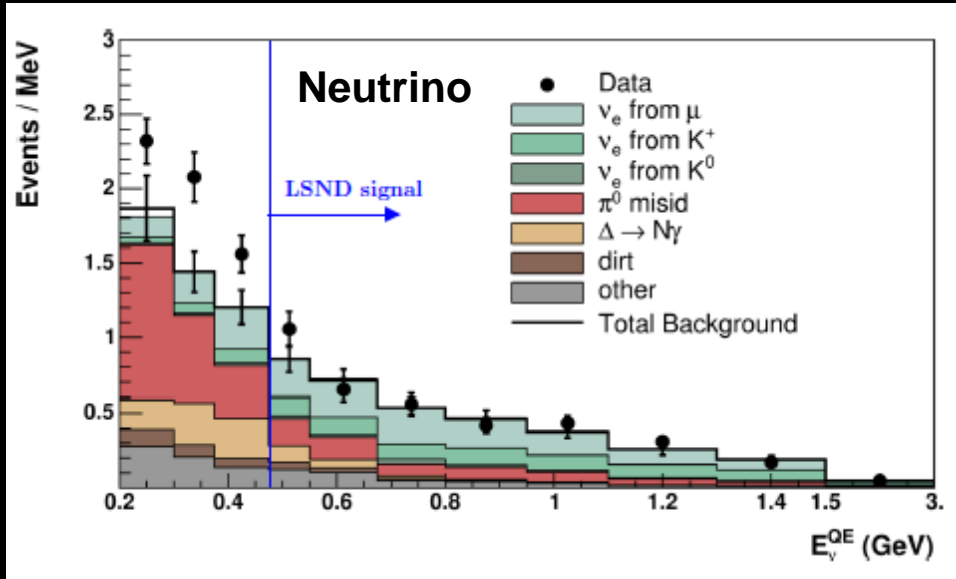
$$0.2 \leq \Delta m^2 \leq 1 \text{ eV}^2 \quad 2 \times 10^{-3} \leq \sin^2 2\theta \leq 4 \times 10^{-2}$$





MiniBooNe

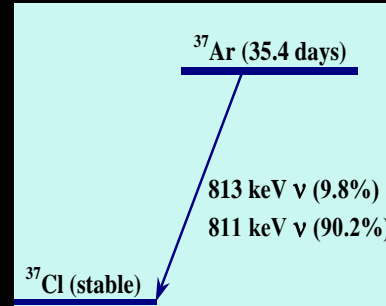
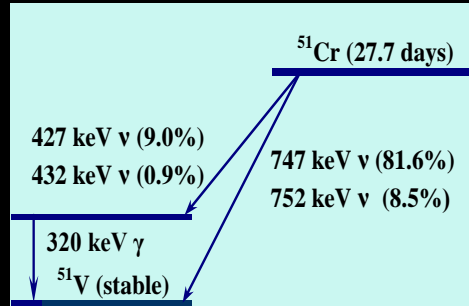
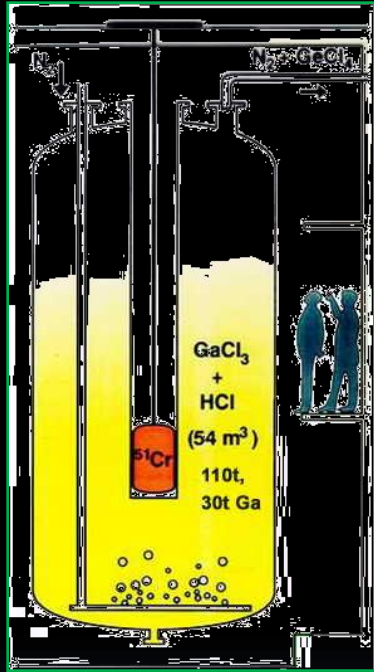
$\nu_{\mu} \rightarrow \nu_e$
 $\text{anti-}\nu_{\mu} \rightarrow \text{anti-}\nu_e$
 $L \approx 540 \text{ m}$ $E_{\nu} = 0.2\text{-}3 \text{ GeV}$



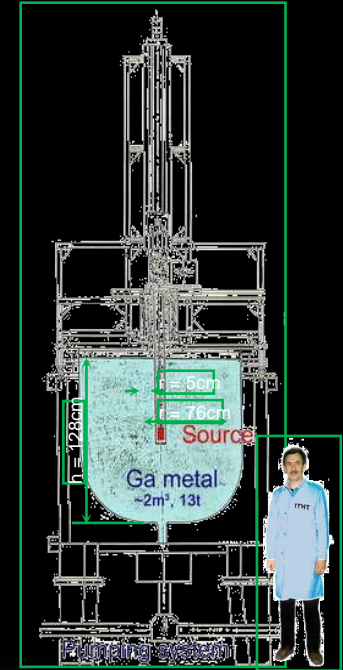


Gallium anomaly

GALLEX

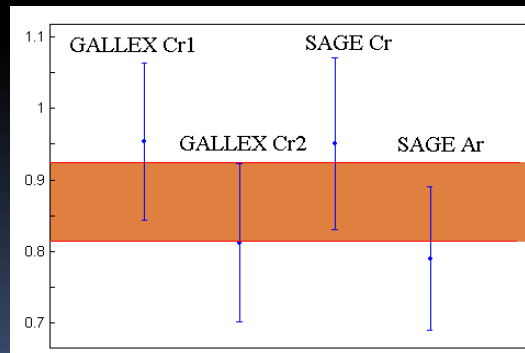


SAGE



Detection process: $\nu_e + {}^{71}\text{Ga} \rightarrow {}^{71}\text{Ge} + e^-$

	GALLEX m(Ga)=30 t		SAGE m(Ga)=13 t	
Source	${}^{51}\text{Cr}$ -1	${}^{51}\text{Cr}$ -2	${}^{51}\text{Cr}$	${}^{37}\text{Ar}$
Intensity (Mci)	1.714	1.868	0.517	0.409
$R = (p_{\text{exp}}/p_{\text{theory}})$	0.95 ± 0.11	0.81 ± 0.11	0.95 ± 0.12	0.79 ± 0.10
R_{comb}	0.88 ± 0.08		0.86 ± 0.08	

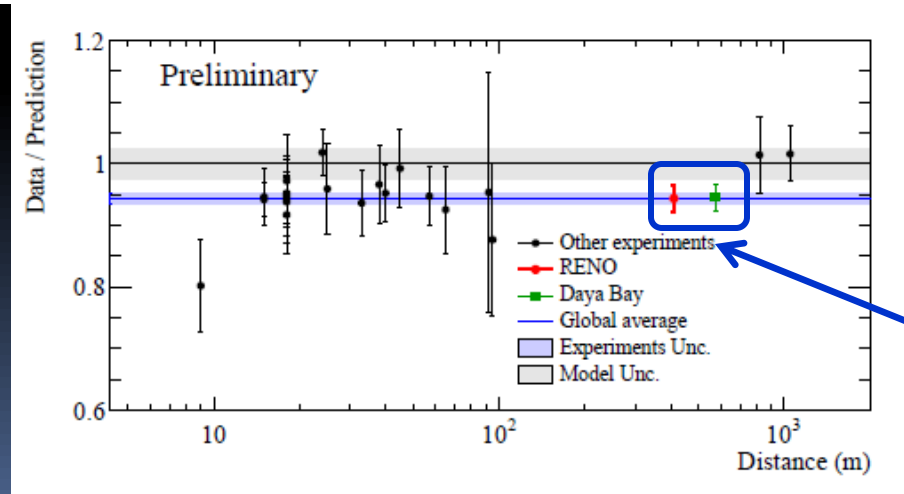
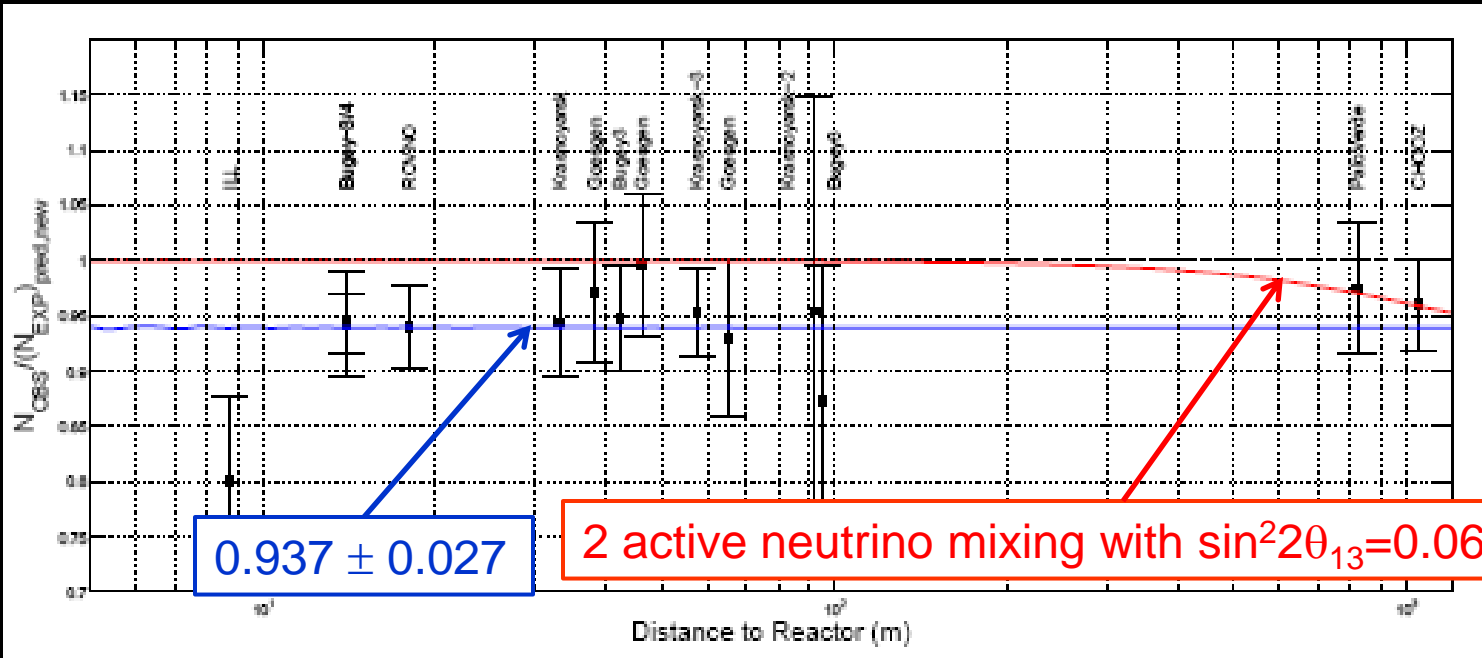


$$R = p_{\text{exp}}/p_{\text{theory}} = 0.87 \pm 0.05$$



Reactor anomaly

anti- $\nu_e \rightarrow \text{anti-}\nu_e$



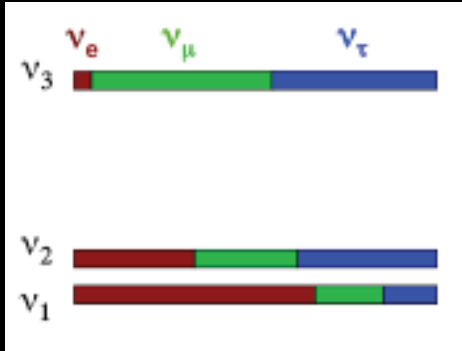
S.Gariazzo et al
arXiv:1703.08860

**Daya Bay
and
RENO**

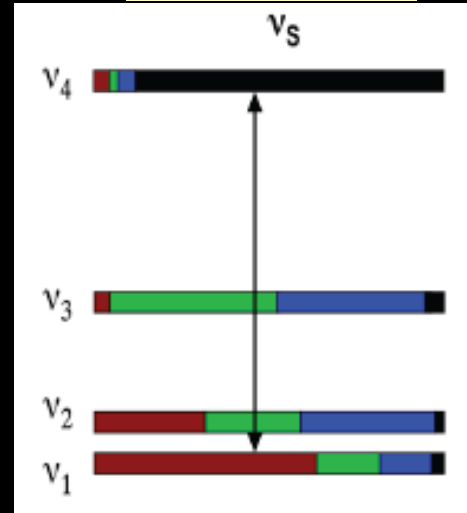


Sterile neutrino?

3ν



3ν + 1s



$$\Delta m_{14}^2 \sim 1 \text{ eV}^2$$

?



PNMS matrix

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_s \end{bmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix} \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{bmatrix}$$

$$\left. \begin{aligned} |U_{e4}|^2 &= \sin^2 \theta_{14} \\ |U_{\mu4}|^2 &= \sin^2 \theta_{24} \cdot \cos^2 \theta_{14} \\ |U_{\tau4}|^2 &= \sin^2 \theta_{34} \cdot \cos^2 \theta_{24} \cdot \cos^2 \theta_{14} \end{aligned} \right\}$$

$$P_{\nu_e \rightarrow \nu_e} \simeq 1 - 2|U_{e4}|^2(1 - |U_{e4}|^2)$$

$$P_{\nu_\mu \rightarrow \nu_\mu} \simeq 1 - 2|U_{\mu4}|^2(1 - |U_{\mu4}|^2)$$

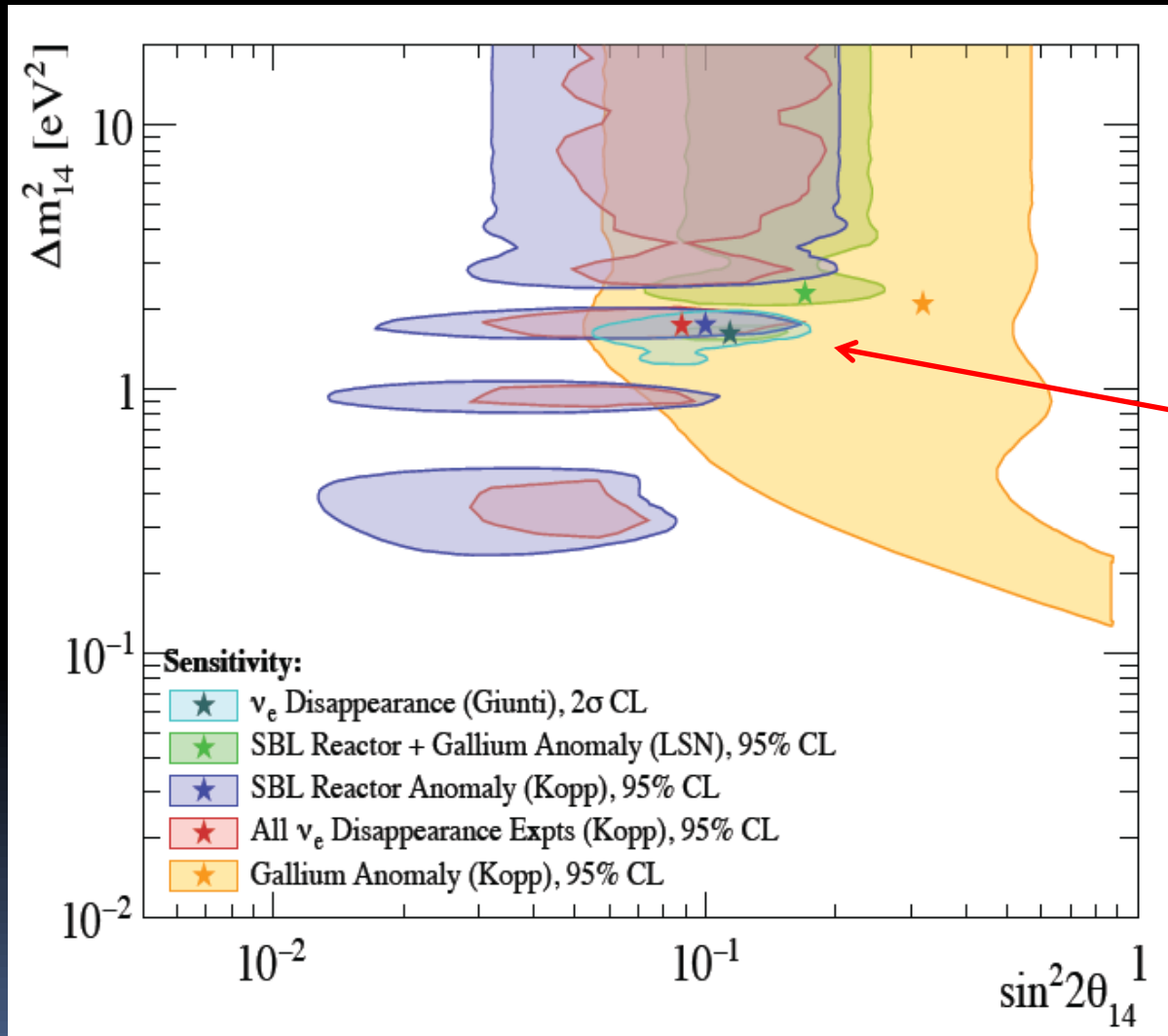
$$P_{\nu_\mu \rightarrow \nu_e} \simeq 2|U_{e4}|^2|U_{\mu4}|^2$$



ν_e and anti- ν_e disappearance

arXiv:1512.02202

Global fit of reactor and Gallium data

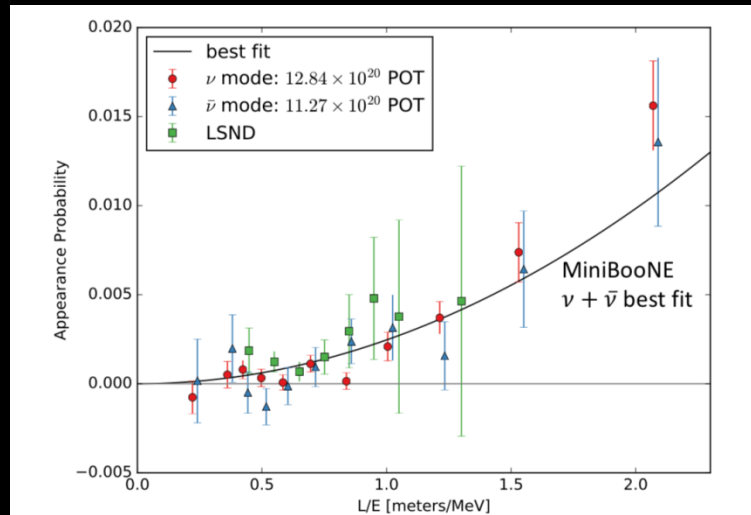


$\Delta m^2 \approx 2 \text{ eV}^2$
 $\sin^2 2\theta \sim 0.1$



New MiniBooNE result

arXiv:1805.12028



MiniBooNE doubled
ν data since 2012

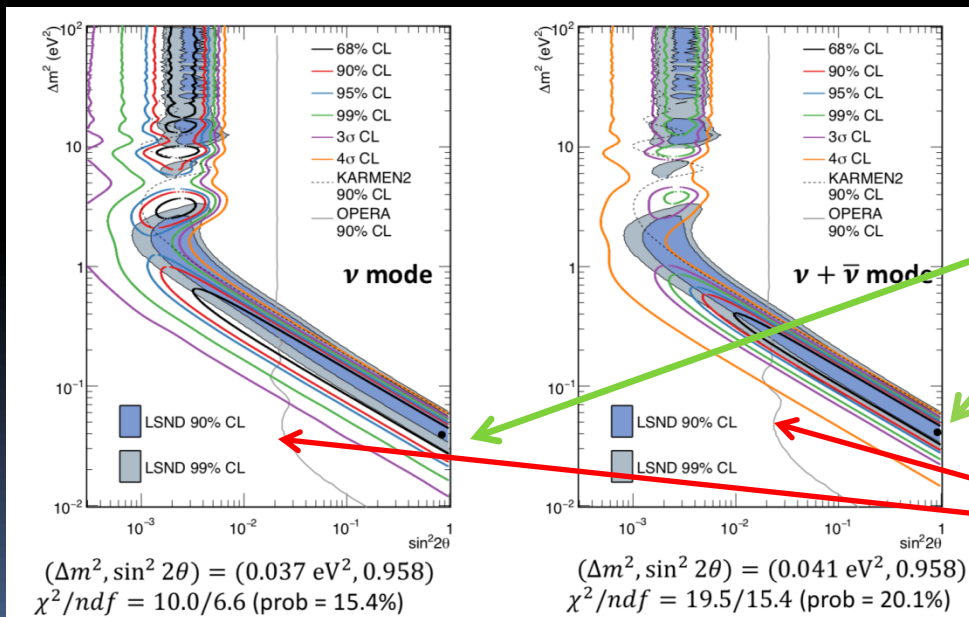
ν	12.84×10^{20} POT
anti-ν	11.27×10^{20} POT



**neutrino +antineutrino
total excess
 460.5 ± 95.8 events (4.8σ)**

**Best fit point:
 $\Delta m^2_{41} = 0.041 \text{ eV}^2$
 $\sin^2 2\theta = 0.958$**

**Excluded by
OPERA
and ICARUS**



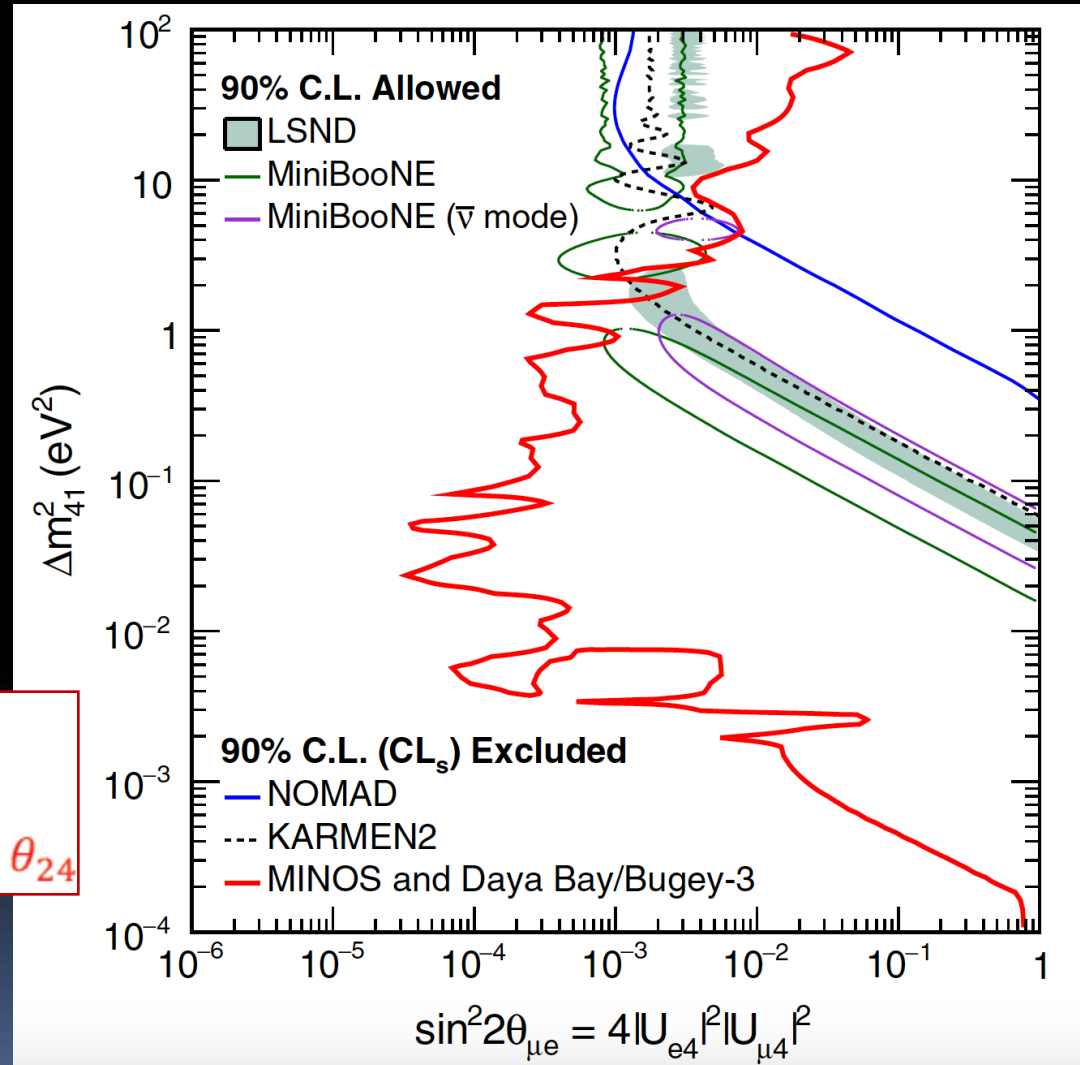


Sterile ν 's: Daya Bay + MINOS+ Bugey-3

PRL117 (2016) 151801

- Daya Bay data
 - Constrains Δm_{41}^2 (mainly 10^{-4} to 10^{-1} eV²) and $\sin^2 2\theta_{14}$
- Bugey-3 data
 - constrains Δm_{41}^2 (mainly 10^{-1} to 10 eV²) and $\sin^2 2\theta_{14}$
- MINOS data
 - Constrains Δm_{41}^2 (mainly 10^{-3} to 10^2 eV²) and $\sin^2 \theta_{24}$

- Combined all three
 - Constrains Δm_{41}^2 and $\sin^2 2\theta_{\mu e} = \sin^2 2\theta_{14} \cdot \sin^2 \theta_{24}$





Sterile ν 's: IceCube

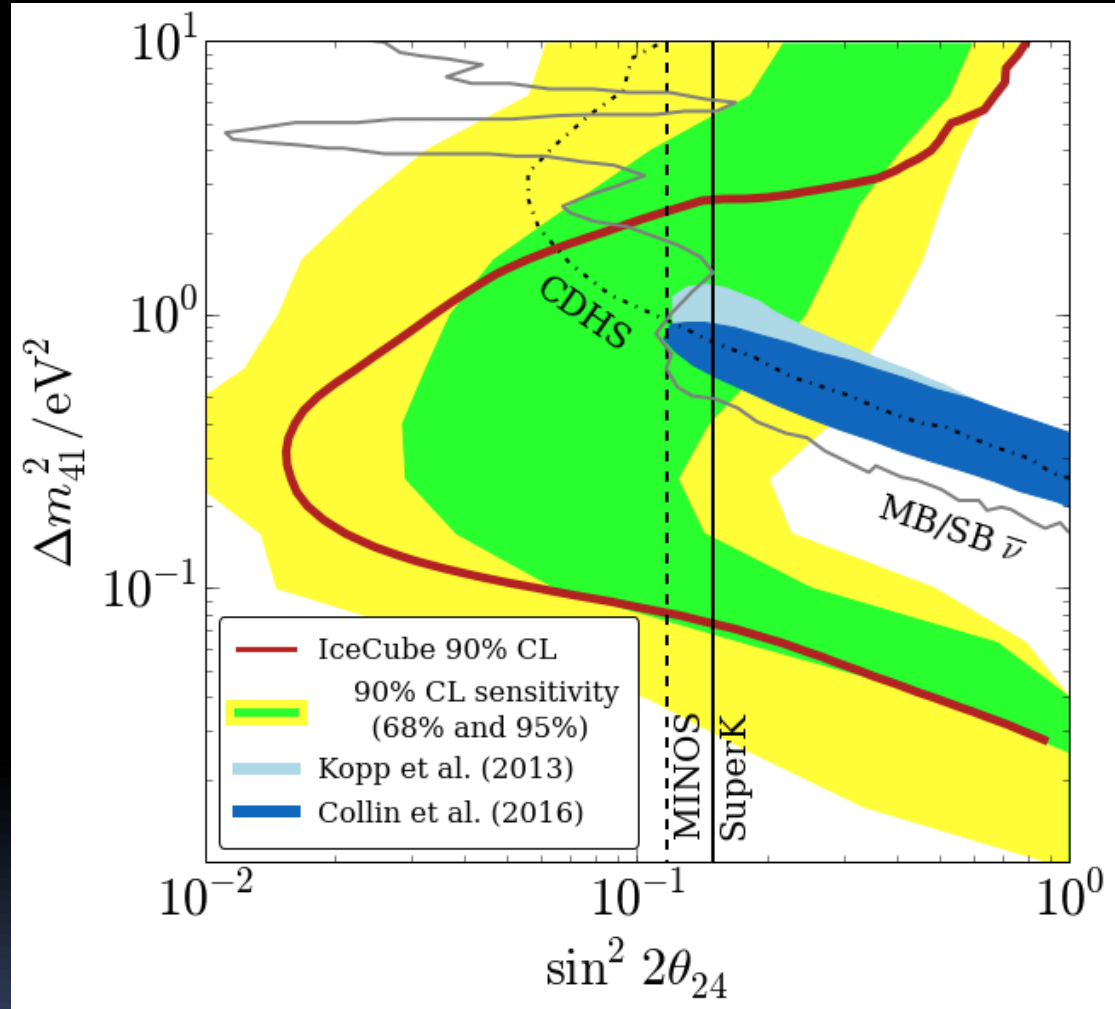
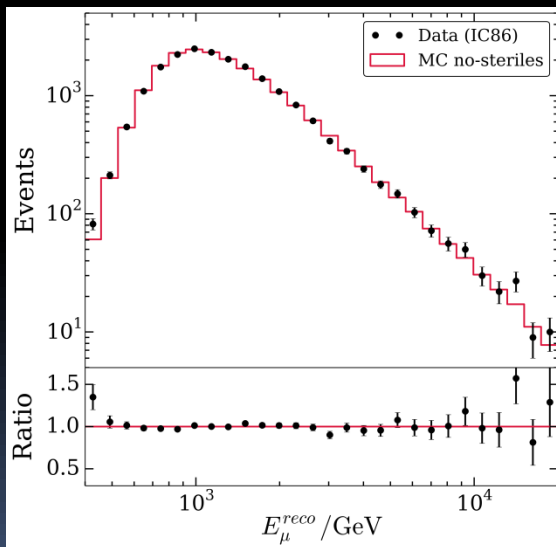
PRL 117 (2016) 071801

$E_\nu = 320 \text{ GeV} - 20 \text{ TeV}$

sterile neutrinos produce distortions of $\nu_\mu + \text{anti-}\nu_\mu$ flux (energy and angle) in the range

$$0.01 \leq \Delta m^2 \leq 10 \text{ eV}^2$$

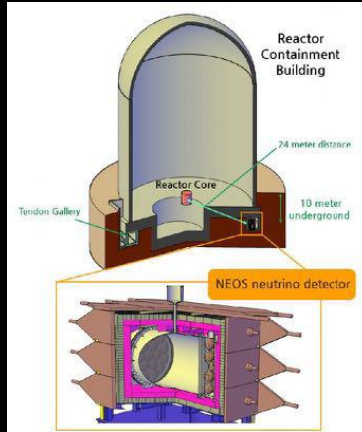
- 1 year of data
- statistics limited



Result compatible with **no-sterile hypothesis**

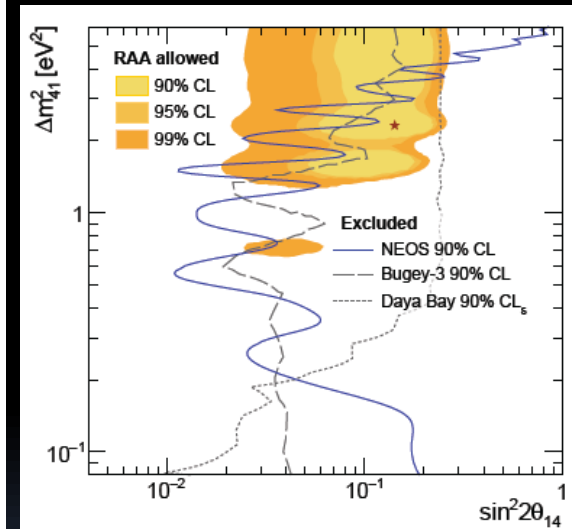
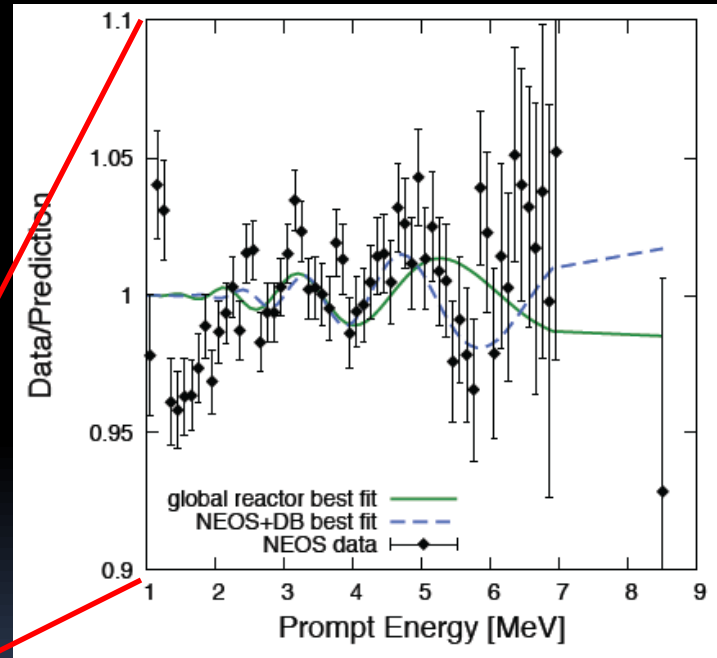
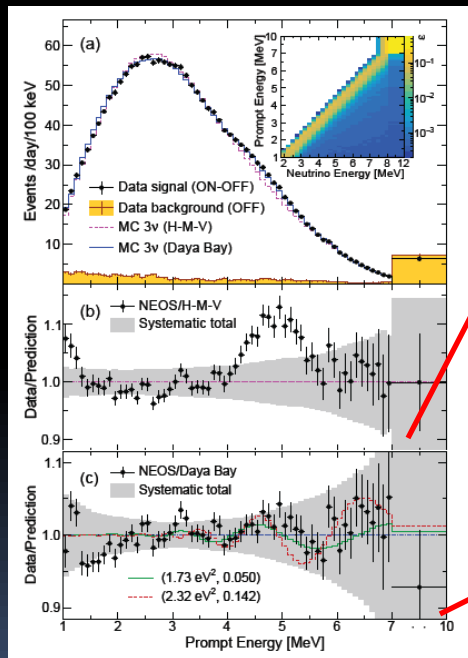


NEOS: reactor anti- ν disappearance



Korea, Reactor 2.8 GW
Core: \varnothing 3.1 m h=3.8 m
Detector 1t LS + Gd, 24 m from reactor core
S/N \sim 22

PRL 118 (2017) 121802



No evidence for sterile neutrino with mass \sim 1 eV



Neutrino-4

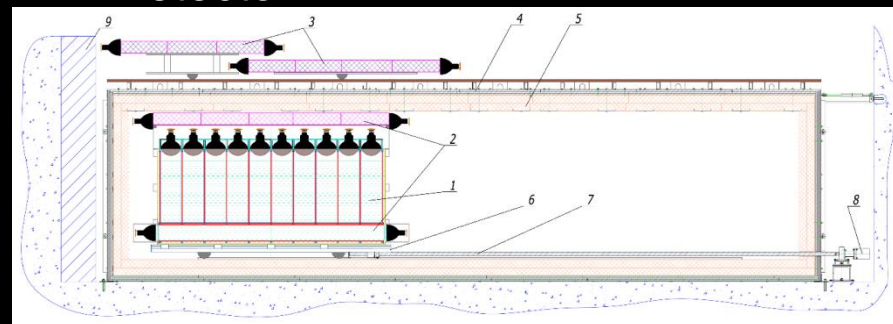
arXiv:1708.00421

SM-3 reactor, Dimitrovgrad

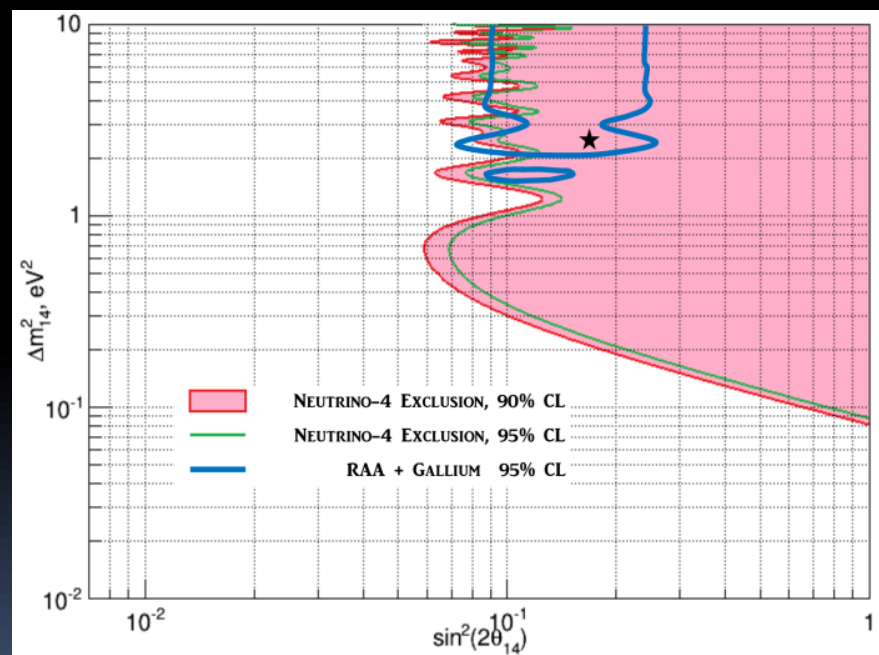
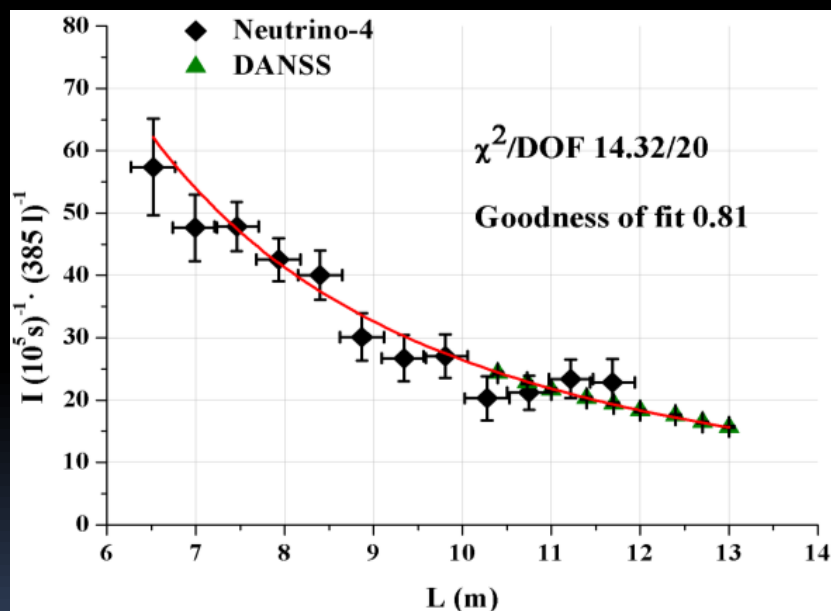
Detector on Earth surface

Signal/background = 0.6

Detector



~ 6.5 – 11.5 m from active zone



Best fit point of Reactor and Gallium anomalies excluded at 95% CL



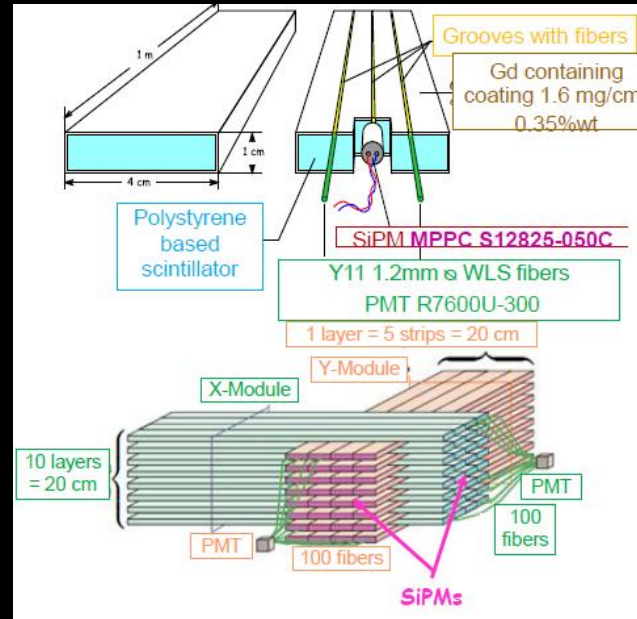
DANSS experiment

arXiv:1804.0404

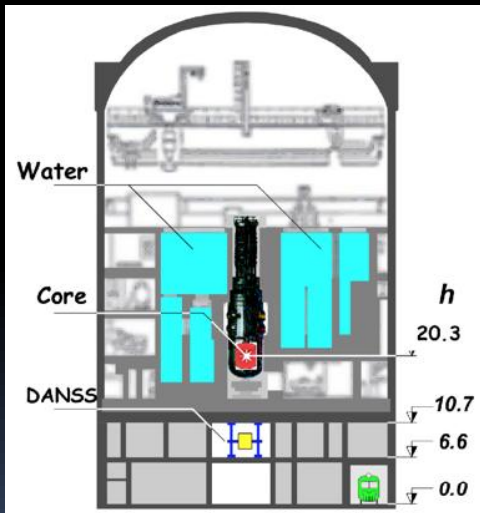
3.1 GW_{th} Kalinin Power Plant



Detector configuration

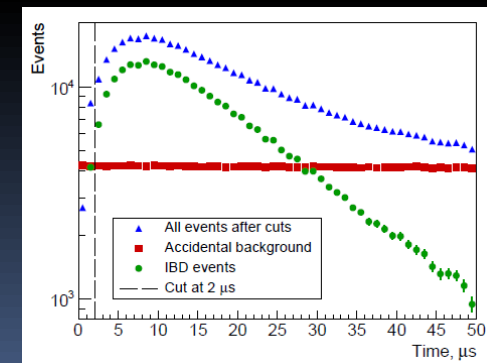
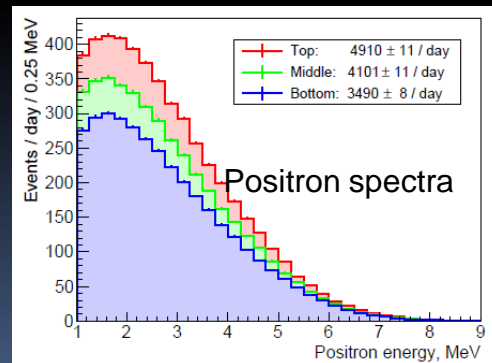


- 2500 scintillator strips with Gd containing coating for neutron capture
- Light collection with 3 WLS fibers
- Central fiber read out with individual SiPM
- Side fibers from 50 strips make a bunch of 100 on a PMT cathode = Module



Distance from the reactor core 10.7-12.7 m

5000 anti- ν_e detected per day with $< 3\%$ cosmic background

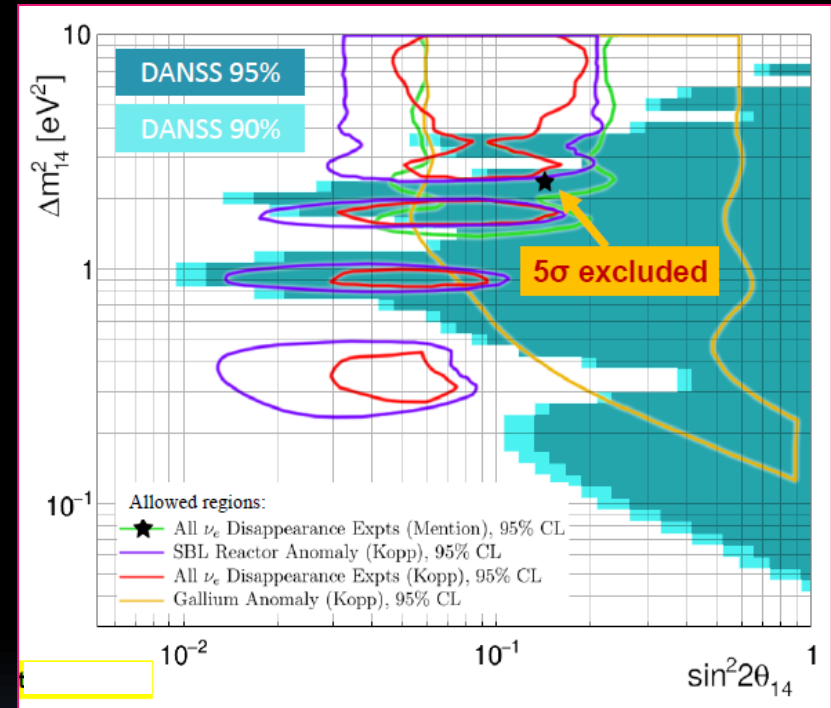
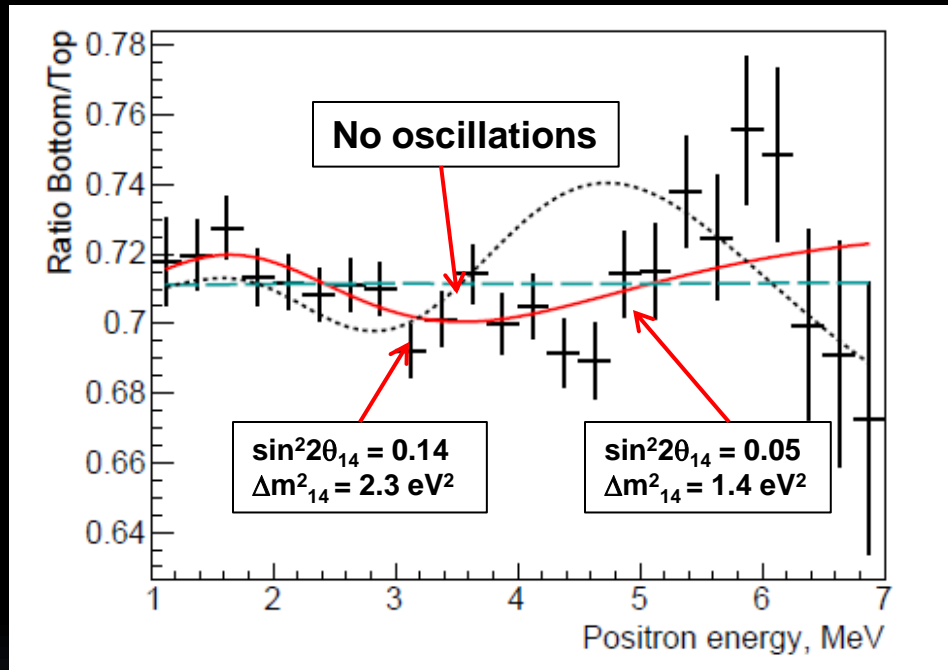




DANSS result

arXiv:1804.0404

Result is based on 663×10^3 events

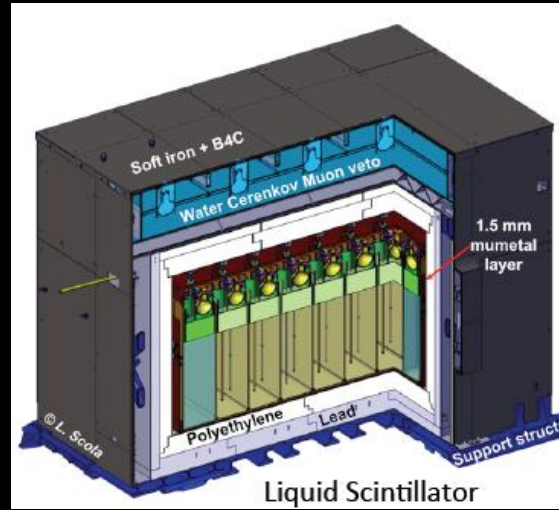


The Reactor Anomaly best fit point ($\sin^2 2\theta_{14} = 0.14$ $\Delta m^2_{14} = 2.3 \text{ eV}^2$) is excluded at $> 5\sigma$ CL



STEREO experiment

ILL, Grenoble, France,
58.3 MW_{th}
compact core Ø40x80 cm

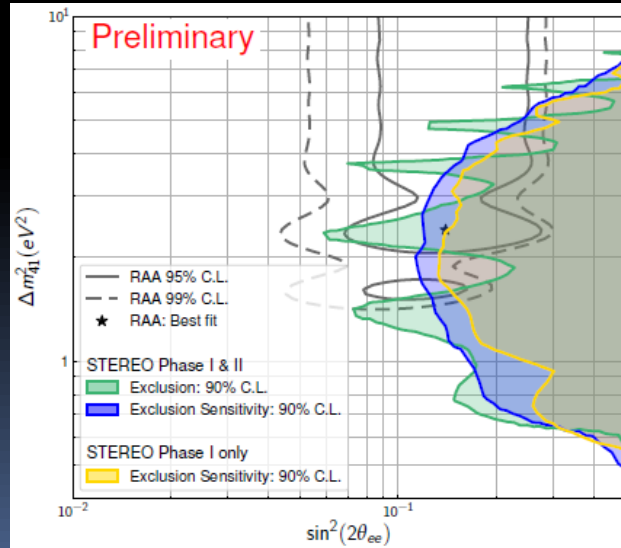


L. Bernard, ICHEP2018

- Detector 6 identical cells
- Gd loaded (0.2% in mass)
- $V_{\text{tot}} = 2.2 \times 0.9 \times 0.9 \text{ m}^3$
- 15 mwe overburden
- Neutrino rate 396/day
- First result based on 65.8 days of data taking

The first results using ratios of cells compatible with no oscillation, rejects the best fit of the Reactor Antineutrino Anomaly at 98.8% C.L.

Phase-I + Phase-II
combined results
(66+47) days reactor-ON
 $(396 \pm 4) \bar{\nu}_e \text{ day}^{-1}$

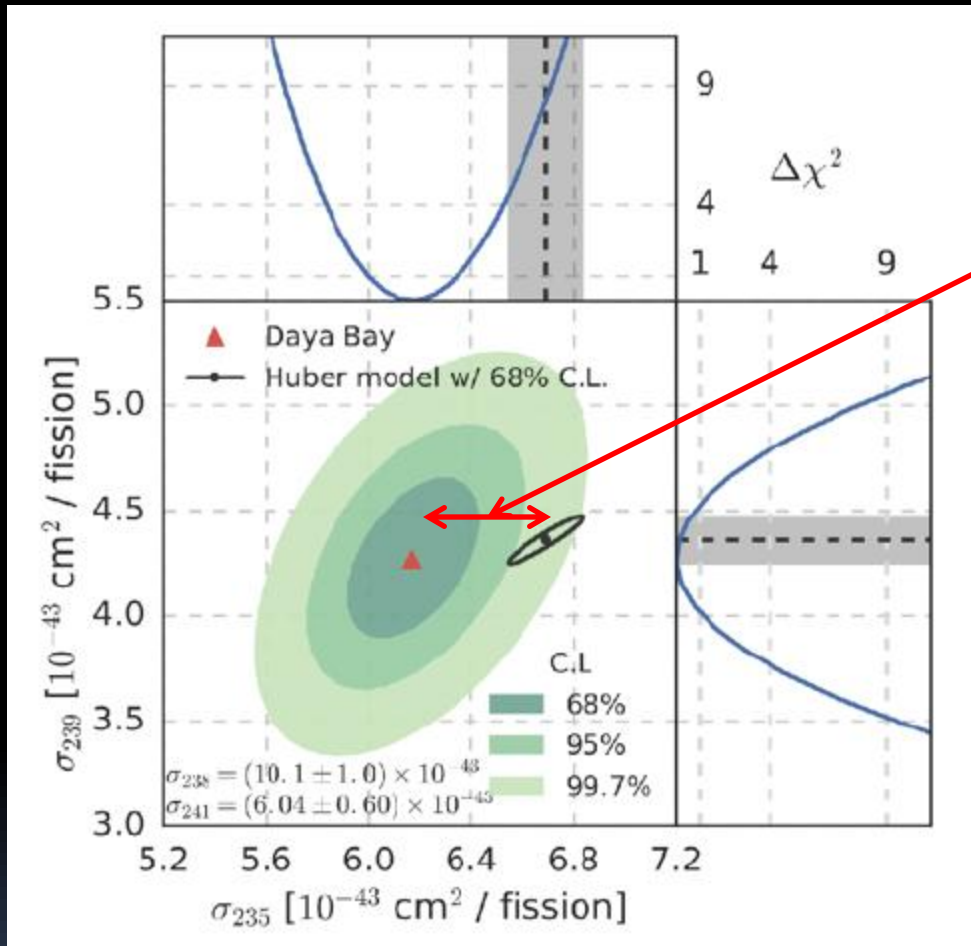


Preliminary result is compatible with no oscillation. The best fit of the Reactor Anomaly is rejected at 98%.



Daya Bay: anti-neutrino flux

PRL 118 (2017) 251801



This discrepancy gives an overestimation of predicted antineutrino flux by 7.8%.

U-235 is a possible source of the Reactor Anomaly?

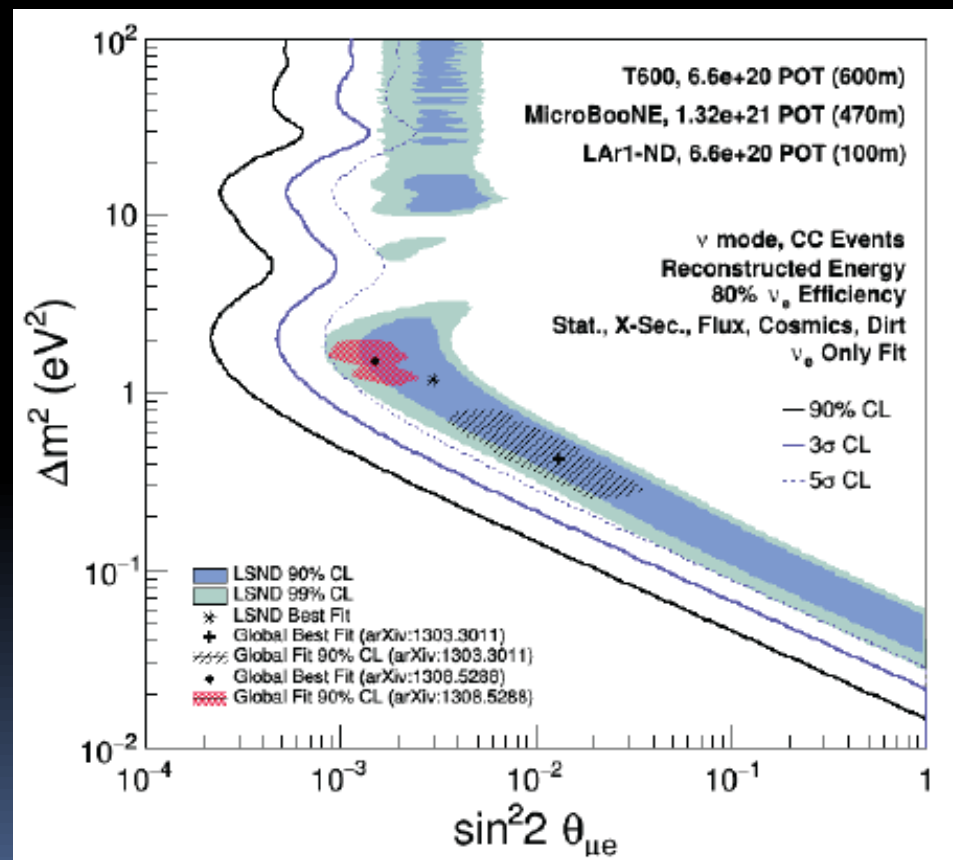
Short baseline experiments at U-enriched reactors are needed



FNAL: Short Baseline Neutrino program

arXiv:1503.01520

Detector	Distance from BNB Target	LAr Total Mass	LAr Active Mass
LAr1-ND	110 m	220 t	112 t
MicroBooNE	470 m	170 t	89 t
ICARUS-T600	600 m	760 t	476 t





Hunting for light sterile neutrinos

Accelerator

MiniBooNe

MINOS+

SBN at FNAL

Reactor *(running or under construction)*

Daya Bay

DANSS

Neutrino-4

NEOS

STEREO

Solid

PROSPECT

NuLat

Neutrino sources

BEST

Atmospheric neutrinos

SuperKamiokande

IceCube



Conclusion

Current LBL experiments T2K + NOvA

main goals: CP violation (3σ), Mass Hierarchy, θ_{23}

T2K: first hint of CP violation in lepton sector

Next generation experiments: discovery/measurement of CP violation, determination of Mass Hierarchy

JUNO (MH) *under construction*

DUNE (CP, MH) *approved*

HyperK and T2HK (CP) *approved, seed funding*

Light sterile neutrinos:

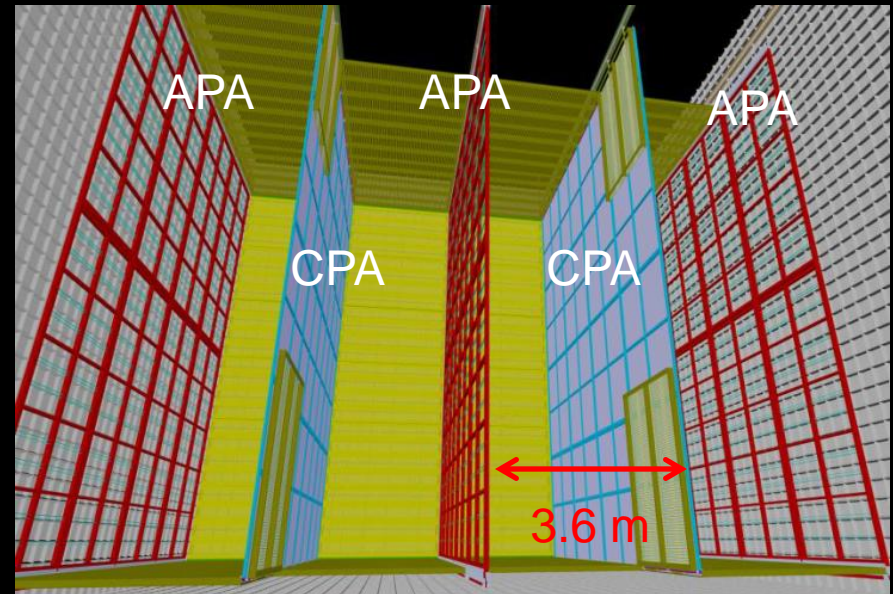
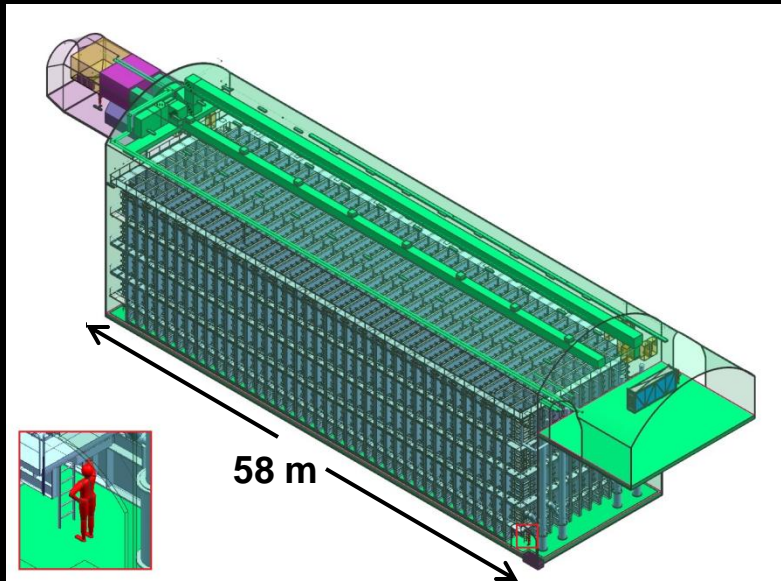
- no positive signal from running experiments
- crucial tests are coming

Thank you for attention!

Backup slides

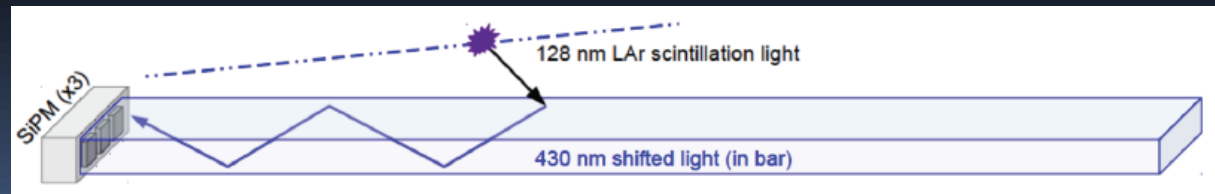


Single-phase LAr TPC



1st 10 kt module of DUNE - single-phase TPC
6m x 2.3 m anode and cathode planes 3.6 m spacing
Photon detectors – light guides + SiPMs embedded in APAs

J.Insler, talk at LLWI2017





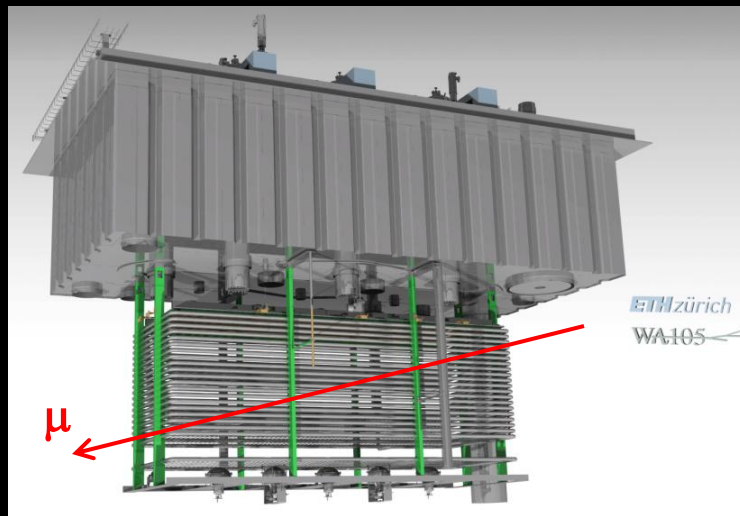
LAr detectors at CERN Neutrino Platform

NP02: WA105

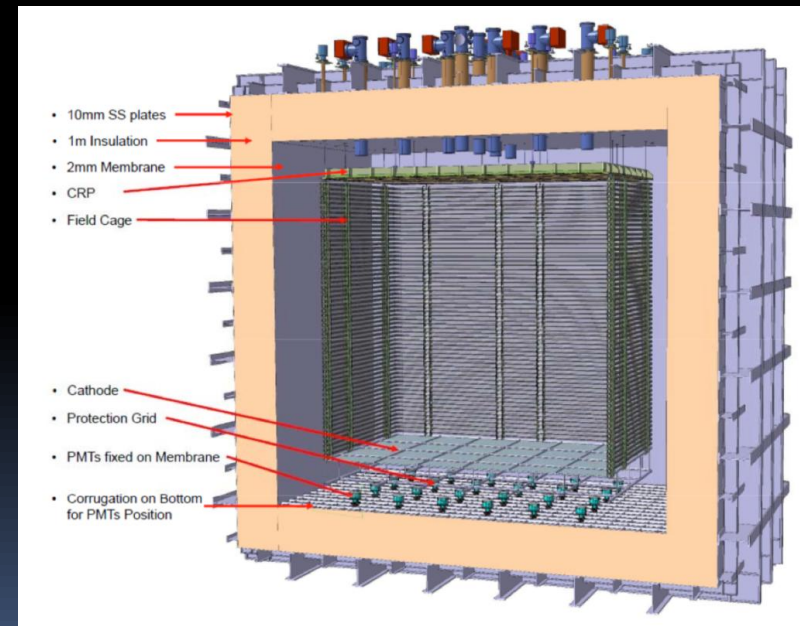
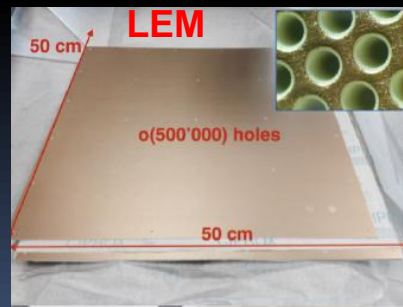
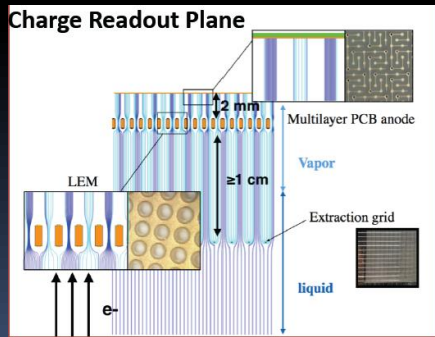
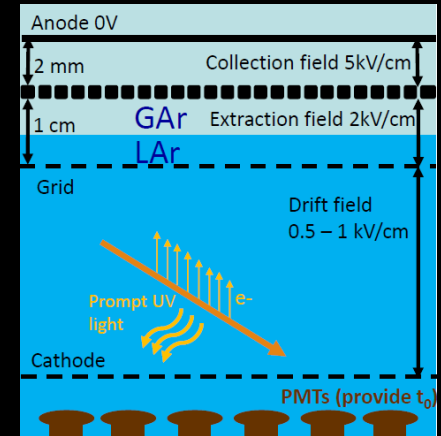
(DP demonstrator + ProtoDUNE DP)

S.Murthy, talk at TPC-2016

Demonstrator: 3x1x1 m³ – 5 tons



ProtoDUNE DP:
6x6x6 m³
300 tons active mass



Cosmic data taking gas begun

Measurements with test beam in 2018



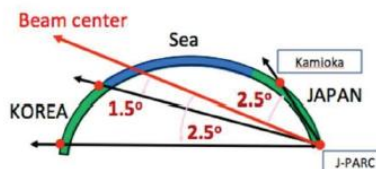
T2HKK

Second tank in Korea

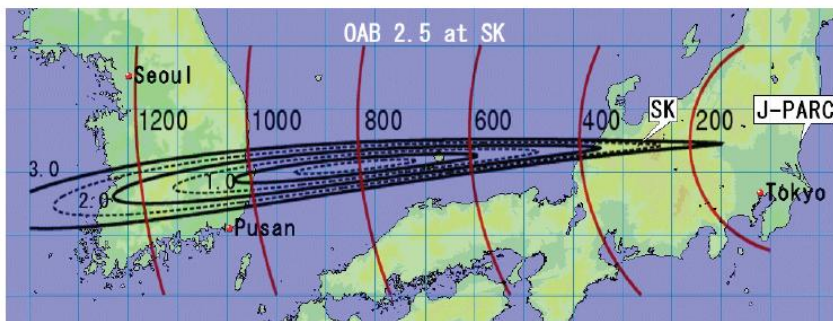
arXiv:1611.06118

Build second tank in Korea to enhance mass hierarchy and δ_{CP} sensitivities

- 1000 – 1200 km baseline
- 1.3° – 3.0° off axis beam direction



Neutrino and antineutrino spectra in T2HKK cover 1st and 2nd oscillation maximum



- A_{CP} ~3 times larger in 2nd maximum
 - Sensitivity to MH

