

Search for dark matter particles predicted by extended Higgs models

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Introduction: Dark Matter

Arguments for dark matter existence

Astrophysical

- Curved rotation of galaxies, virial theorem "violation", ultra diffuse galaxies and satellites of galaxies, spiral structures of galaxies
- Gravitational lensing, evaluation of potentials and masses of galaxies/clusters of galaxies ("Bullet" cluster etc.)

Cosmological

- Anisotropy of cosmic microwave background, flatness of the Universe, the prevalence of the elements and necessity of DM.
- Forming of the early Universe structure, the growth of the initial inhomogeneities

Dark matter (DM) is

- Not interacting (very weakly interacting) with ordinary matter
- **Electrically neutral**
- Stable in terms of cosmological time (14 bill. years)
- Initiated in the early stage of the Universe (till the change of the regimes, from radiation-dominant epoch to epoch of matter domination)

DM candidates

- Baryonic matter (massive astrophysical compact halo objects - MACHO)
- Non-baryonic matter (sterile neutrinos, weakly interacting massive particles – WIMPs, axions, supersymmetric particles, etc.)

DM in particle accelerators

Dark matter can be produced within particle accelerators if:

- The dark matter mass is low enough
- Its production cross section is large enough
- Dark and ordinary matter interact at least weakly with each other

The LHC is able to probe energies higher than ever with huge luminosities:

- Largest dataset to date to analyze at 13 TeV
- Perfect tool to try and detect DM particles
- Able to study a large range of particle masses and cross-sections
- The two multipurpose detectors (CMS, ATLAS) are mostly able to search for DM particles

However, if producing dark matter particles is theoretically possible, detecting them directly is impossible, as they are not expected to interact with detector.

How to detect DM?

The CMS detector is not able to directly detect eventual dark matter particles

The key variable to detect DM is the missing transverse energy (MET):

• Defined as the imbalance in transverse momentum in the plane perpendicular to the beam direction ${\sf p_T}^{\sf miss}$

$$
p_T^{\text{miss}}=-|\sum \overrightarrow{p_T}|=0
$$

• This quantity is $p_T^{miss} \neq 0$ if something escapes the detector undetected

Most of the DM searches are therefore dependant on p_T^{miss}

However, $p_T^{miss} \neq 0$ does not mean that we discovered new physics, as common processes can have the same effect:

- Neutrino production
- Limited detector resolution

2HDM+a/S Models

The two-Higgs-doublet model (2HDM) is a way to extend Higgs sector

- neutral CP-even scalars *h, H*
- neutral CP-odd pseudoscalar *A*
- charged *H⁺ , H-*

arXiv:1701.07427

2HDM + a (neutral pseudoscalar singlet)

h (bbar) + a $(\chi \chi)$ = bbar + MET

 $Z + a(y\chi) = Z (II) + MET$

Dark Mater with Z and MET at the LHC 27.03.2024 Tepesitory, the contracted models, reduced to the state of the S repository/tree/master/models/Pseudoscalar_2HDM

Model Parameters and Signal Simulation

Generator: **MadGraph5MC@NLO.2.9.2** (PS, frag./hadr. - **Pythia 8**)

Models: 2HDM+s or 2HDM+a + NNPDF 3.1 NNLO

Process: $p p > Z \chi \chi$ (16 diagrams)

Free parametersfor **2HDM + a**:

- masses of heavy higgses, $m_{H+/-} = m_{H} = m_A = [600:2000]$ GeV
- mass of dark matter particle, $my = [1:2000]$ GeV
- mass of light pseudoscalar/scalar states, m_a = m_S = [300:1000] GeV
- the ratio of the vacuum expectation values of the two Higgs doublets, $tan(β) = [0.5:50]$
- the mixing angle of the two CP-odd weak spin-0 eigenstates (a/A), $sin(\Theta) = [0.15:0.7]$
- the mixing angle between the two CP-even weak spin-0 eigenstates (h/H), sin(O O) = 1

Free parameters for $2HDM + S$:

 $H = \cos \theta S_1 - \sin \theta S_2,$

 $S = v_S + \sin \theta S_1 + \cos \theta S_2.$

- Yukawa couplings
- Couplings of DM and mediators (a/S)
- Mass and widths of the w/ new states Interaction constants between two Higgs doublets (different for 2HDM+s and 2HDM+s)

RUN3 Expectations (cross sections)

2HDM+S 2HDM+a

pp → Z χ \bar{x}

In total about 1000 sets of model parameters

Background

Main background sources: $(pp \rightarrow Z + MET \rightarrow H + MET)$

 $WZ \rightarrow IVII$

 $ZZ \rightarrow 4I$

 $V\gamma \rightarrow$ II + missreconstructed γ

VVV: (WWZ, WZZ, and ZZZ) ttW \rightarrow WWbbW, ttZ \rightarrow WWbbZ, and tty \rightarrow WWbby

Experiment CMS at LHC

dark photons in VBF

Higgs boson

 $137 fb^{-1}$

EXO-20-005

Summary of the kinematic selections for the signal region

mmmmm

CMS RUN2 Results

Observed number of events and post-fit background estimates

Expected yields and the product of acceptance and efficiency for several models probed in the analysis

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CMS RUN2 Exclusion Plot for 2HDM+a

CMS RUN2 Exclusion Plot for 2HDM+a

DM-nucleon upper limits on the cross section for simplified DM

CMS,

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SUMMARY

- A search for dark matter particles can be performed using events with a Z boson and large missing transverse momentum
- Recent search has been performed with proton-proton collision data at a center-of-mass energy of 13 TeV, collected by the CMS experiment at the LHC in 2016-2018, corresponding to an integrated luminosity of 137 fb⁻¹
	- ‒ no evidence of physics beyond the standard model is observed
	- ‒ limits are set on dark matter particle production in the context of a two-Higgs-doublet model with an additional pseudoscalar mediator.
- For the preparation of LHC RUN3 data analysis, the cross sections of dark matter production in association with a leptonically decaying Z boson have been calculated for
	- ‒ 2HDM + a model (additional pseudoscalar mediator)
	- ‒ 2HDM + S model (additional scalar mediator)
- These processes were simulated for about 1000 sets of model parameters
- The next steps are full simulation (right now) and RUN3 data analysis (waiting for data of 2023)

THANK YOU FOR YOUR ATTENTION!

DM-nucleon upper limits on the cross section

CMS The upper limits on the cross section as a function of WIMP mass

The 90% upper limits (solid lines) and expected sensitivity (dotted) on the spin-dependent cross section as a function of WIMP mass obtained by seven years of IceCube DeepCore data in this work. The shaded bands show the central 90% expected limits. Also shown are limits from the Super-K, PICO-60, and ANTARES experiments. Credit: IceCube Collaboration

The simplified dark matter

model for a spin-1 (vector or

CMS RUN2 Exclusion Plots:

Search DM in collider experiment

predictions SM

directly in accelerators.

Models BSM

Available models

Наш интерес