

# Event by event strangeness fluctuation in MPD-NICA experiment

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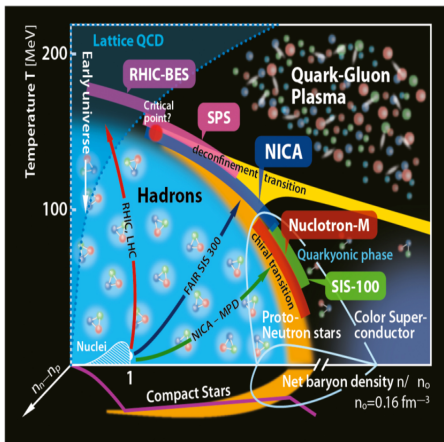


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## Motivation: QCD Phase diagram

- Some QCD calculations indicate that there exists a critical point in the phase transition from hadron gas to quark-gluon plasma.
- The Critical Point can be explored experimentally by studying Quantum Number Fluctuations. This work focuses on analyzing strangeness.



## Cumulants

Let  $\Delta N = N - \bar{N}$  be the net multiplicity of a particle, then the standard deviation is  $\delta N = \Delta N - \langle \Delta N \rangle$ , and the first order cumulants are defined as:

$$C_1 = \langle \Delta N \rangle, \quad C_2 = \langle (\delta N)^2 \rangle, \quad C_3 = \langle (\delta N)^3 \rangle, \quad (1)$$

$$C_4 = \langle (\delta N)^4 \rangle - 3\langle (\delta N)^2 \rangle^2.$$

The cumulants are related with the statistical moments as:

$$M = C_1, \quad \sigma^2 = C_2, \quad S = \frac{C_3}{(C_2)^{3/2}}, \quad \kappa = \frac{C_4}{(C_2)^2} \quad (2)$$

## Cumulants from thermodynamic point of view

Modeling the particle system with a partition function for a grand canonical ensemble  $Z(z, V, T)$ , it can be shown that

$$C_n = \frac{\partial^n \log Z}{\partial (\mu/T)^n}. \quad (3)$$

Using the definition of generalized susceptibilities:

$$\chi_q^{(n)} = \frac{1}{VT^3} \frac{\partial^n \log Z}{\partial (\mu_q/T)^n} \Rightarrow C_n = VT^3 \chi_q^{(n)}. \quad (4)$$

where  $\chi_q^{(n)}$  can be theoretically calculated. Additionally, the cumulants are also related to the correlation length  $\xi$  as:

$$C_2 \sim \xi^2, \quad C_3 \sim \xi^{4,5}, \quad C_4 \sim \xi^7. \quad (5)$$

In the ideal thermodynamic limit,  $\xi$  diverges at the critical point.

# Data Analysis

## Data sample

The following events were generated using UrQMD.

Collision Type	$\sqrt{S_{NN}}$	Events	Analysis
Au+Au	4.5 GeV	190,000	Monte Carlo
Au+Au	7.7 GeV	190,000	Monte Carlo
Au+Au	9.2 GeV	190,000	Monte Carlo
Au+Au	11.5 GeV	190,000	Monte Carlo
Bi+Bi (Request 25)	9.2 GeV	300,000	Reconstructed

## MC: Kaon selection for cumulants calculation

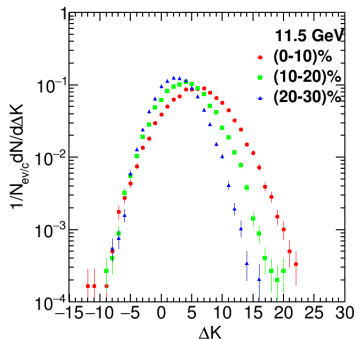
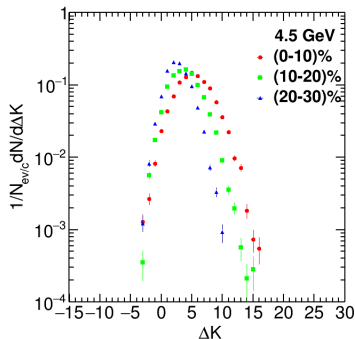
Cuts over kaons:

- $0,2 \leq p_T \leq 1,2 \text{ GeV}/c.$
- $|y| \leq 0,5.$

The cuts used are the same as those used by the STAR collaboration.

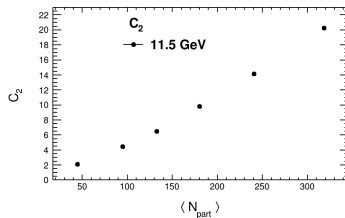
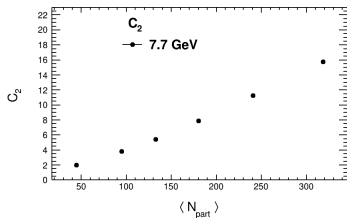
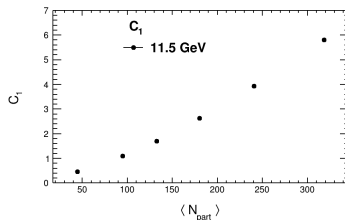
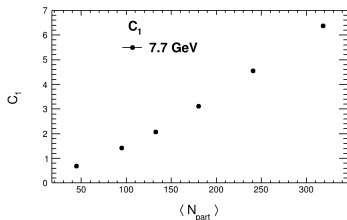


# MC: Kaon net number at different energy



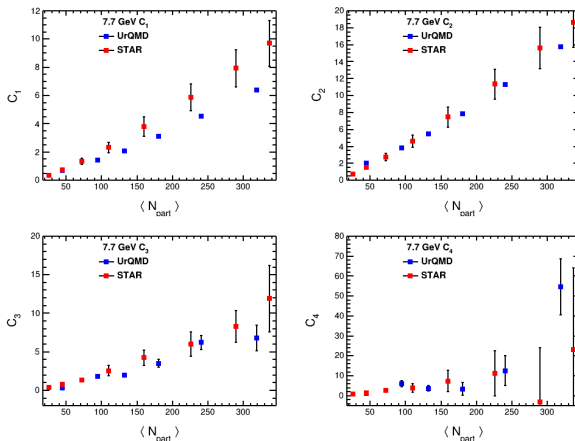
$\Delta K$  Distributions. Distributions from most central collisions are wider. As centrality decreases, the peak shifts towards zero. The distributions are not symmetrical.

# MC: Net kaon cumulants as a function of participants for Au+Au



The volume dependency is observed.

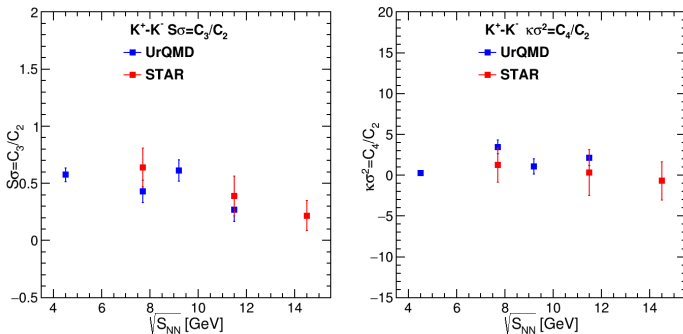
# MC: Comparison with STAR data. $\Delta K, \sqrt{s_{NN}} = 7,7$ GeV



The first cumulant does not match the experimental data; however, the agreement with the other cumulants is very good.

## MC: Ratio of Cumulants and its comparison with STAR data

Comparison with STAR data for central Au+Au collisions at different energies.



A prediction is shown for energies of 4.5 GeV and 9.7 GeV; however, it is necessary to test different models.

# Reconstructed Data

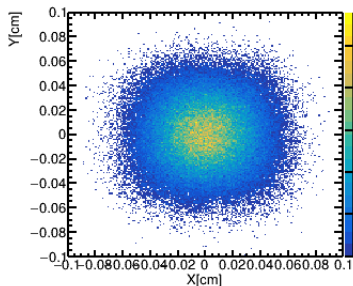
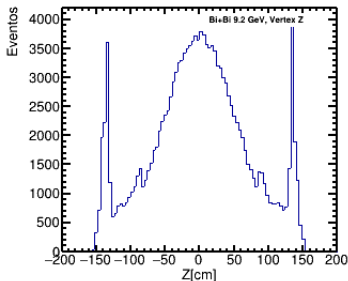
## Sample Data

UrQMD generated Bi+Bi events (Request 25).

- $\sqrt{s_{NN}} = 9,2 \text{ GeV}$ .
- Reconstructed Data

## Event Selection

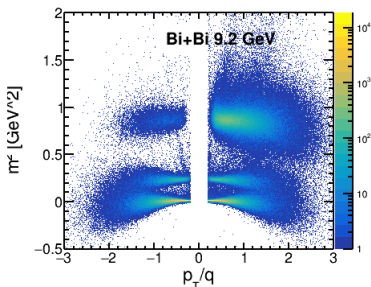
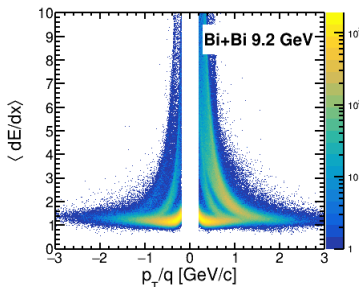
Vertex cut  $|z| \leq 80$  cm.



Collision	$\sqrt{S_{NN}}$	Events	Events after cuts
Bi+Bi	9.2 GeV	300,000	151,662

## Particle Identification

TPC and TOF information to identify  $K$  using the PID wagon.  
 Only primary tracks were selected, and the following cuts were applied:  $0,4 \leq p_T \leq 0,8 \text{ GeV}/c$ ,  $|y| \leq 0,5$ ,  $\langle n\text{Hits} \rangle > 20$ .



$$p_T \leq 0,6$$

- $\sigma_{TPC} \leq 2$

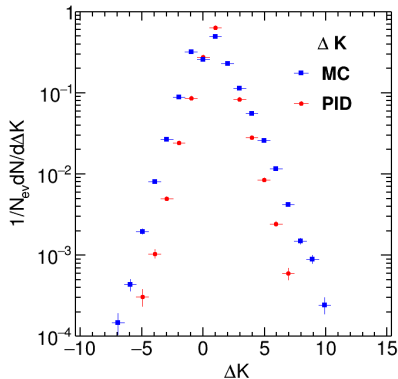
$$p_T > 0,6 (\sigma_{\phi}, \sigma_z \leq 2)$$

- $\sigma_{TPC} \leq 2$

- $\sigma_{TOF} \leq 2$

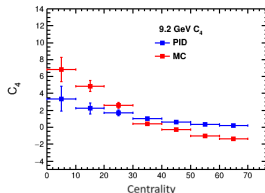
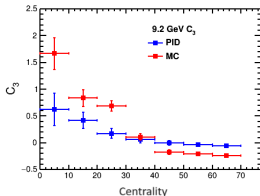
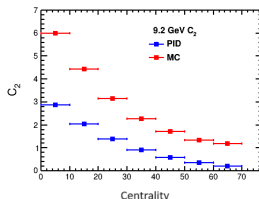
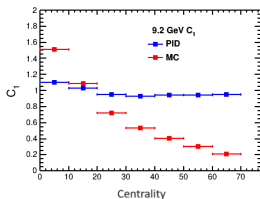


# Net Kaon Distribution



Net kaon distribution from Bi+Bi collisions at 9.2 GeV. Monte Carlo and PID uncorrected distribution.

## Calculation of cumulants (uncorrected)



Statistical cumulants (uncorrected) compared with Monte Carlo. It is not possible to perform the physical analysis without correction.

## Cumulants Corrections

To perform the correction, we assume that the difference between the real distribution  $P$  and the measured distribution  $p$  can be modeled as a binomial distribution, so defining the factorial moments of  $p$  and  $P$  as

$$f_{ik} = \left\langle \frac{n_1!}{(n_1 - i)!} \frac{n_2!}{(n_2 - k)!} \right\rangle, \quad F_{ik} = \left\langle \frac{N_1!}{(N_1 - i)!} \frac{N_2!}{(N_2 - k)!} \right\rangle \quad (6)$$

we can get the relation

$$F_{ik} = \frac{1}{p_+^i p_-^k} f_{ik}. \quad (7)$$

$p_+$  and  $p_-$  the acceptance of the identification. With this relation, is possible to obtain the real value of the cumulants.

## Cumulants Corrections

Using the previous relationship and by the definition of statistical cumulants, the following equalities are obtained:

$$C_1 = F_{10} - F_{01},$$

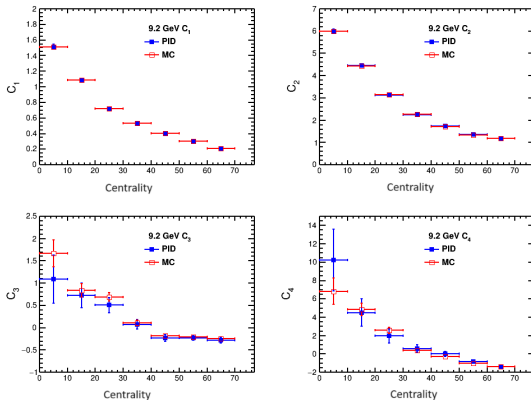
$$C_2 = N - C_1^2 + F_{02} - 2F_{11} + F_{20},$$

$$C_3 = C_1 + 2C_1^3 - F_{03} - 3F_{02} + 3F_{12} + 3F_{20} - 3F_{21} + F_{30} \\ - 3C_1(N + F_{02} - 2F_{11} + F_{20}),$$

$$C_4 = N - 6C_1^4 + F_{04} + 6F_{03} + 7F_{02} - 2F_{11} - 6F_{12} - 4F_{13} \\ + 7F_{20} - 6F_{21} + 6F_{22} + 6F_{30} - 4F_{31} + F_{40} \\ + 12C_1^2(N + F_{02} - 2F_{11} + F_{20}) - 3(N + F_{02} - 2F_{11} + F_{20})^2 \\ - 4C_1(C_1 - F_{03} - 3F_{02} + 3F_{12} + 3F_{20} - 3F_{21} + F_{30}).$$

$C_n$  are the cumulants of the real distribution.

# Calculation of cumulants (corrected)



Statistical cumulants with correction. Higher order cumulants have more discrepancy in central collisions.

## Summary and perspectives


The study of strangeness number fluctuations and the calculation of the first 4 cumulants were presented both at the MC simulation and at the reconstruction level in the MPD experiment.

- UrQMD produce a reasonable description of the strangeness.
- The cumulants ( $C_1$ ,  $C_2$ ,  $C_3$ , and  $C_4$ ) for strangeness were calculated and corrected using factorial moments. Results from cumulants at the most central collision indicate that I need to improve the identification of kaons, and increase the statistic sample.

I need to prove different models and perform the analysis using the net number of protons.

Thank you for your attention <sup>1</sup>.

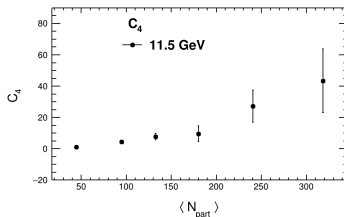
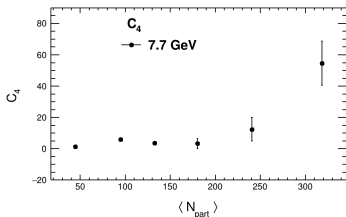
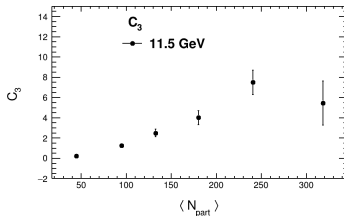
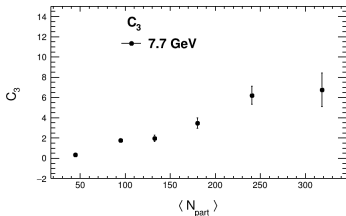
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<sup>1</sup>Special thanks to Eleazar Cuautle and Vadim Kolesnikov 

# Back up



# MC: Net kaon cumulants as a function of participants for Au+Au



The third and fourth cumulants indicate that the distributions  $\Delta K$  are not Gaussian.