Sergey Morozov, INR RAS, Moscow on behalf of HGND team





13th Collaboration Meeting of the BM@N Experiment at NICA

8-10 October 2024, JINR, Dubna

Outline:

- Physics motivation of measuring the neutrons (reminder)
- Highly Granular Neutron Detector (HGND) construction status
- Performance studies for HGND optimization (update)
- Current status of HGND hardware and electronics

New time-of-flight neutron detector for the BM@N experiment is under development and construction now



### **EOS** for high baryon density matter

The binding energy per nucleon:  $E_A(
ho,\delta) = \left| E_A(
ho,0) \right| + \left| E_{sym}(
ho) \right| \delta^2 + O(\delta^4)$ Isospin asymmetry:  $\delta = (\rho_n - \rho_p)/\rho$ Symmetric matter Symmetry energy



Being extensively studied nowadays using observables (flow, meson yields, etc) to explore incomressibility

$$K_0 = 9\rho^2 \frac{d^2 E_A}{d\rho^2}$$

One of the main sources of uncertainty: discrepancy between experimental data

One of the main parameters to study is the E<sub>sym</sub> slope

$$L = 3\rho \frac{dE_{sym}(\rho)}{d\rho}$$

- No experimental data for beam energies  $E_{kin} > 0.4 \text{ GeV}$
- One needs to establish observables sensitive to *L* and obtain new experimental data

Ch. Fuchs and H.H. Wolter, EPJA 30 (2006) 5

**Observables to study symmetry energy** 

# Using $v_1^{np}$ to study L



One can define free neutron-proton differential directed flow:

$$\nu_1^{np} = \frac{N_n(y)}{N(y)} \langle \nu_1^n(y) \rangle - \frac{N_p(y)}{N(y)} \langle \nu_1^p(y) \rangle$$

 $N_n(y), N_p(y), N(y)$  - total number of neutrons, protons and nucleons respectively  $\langle v_1^n(y) \rangle, \langle v_1^p(y) \rangle$  - flow of neutrons and protons respectively

- $v_1^{np}$  sensitive to both  $K_0$  and L which may lead to ambigous interpretation
  - More observables might be necessary for robust study of L

### **Observables to study symmetry energy**

Rapidity and kinetic energy distributions of n/p ratios show strong dependence on L and weak dependence on  $K_{\theta}$ X.X. Long, G.F. Wei, Phys.Rev.C 109 (2024) 5, 054619 IBUU, Au+Au, 0.4A GeV



- n/p ratio requires less statistics than anisotropic flow measurements

### Positioning of the HGND at the BM@N experiment



1) previous proposed 16 layer

(1 veto + 15 active Scint./absorber) HGND detector configuration in position (1) at 17 deg shows limited rapidity range for neutrons

2) in order to extend neutron rapidity range the new position (2) has been found at 10 deg but the distance is 7m from target now, resulting in lost of acceptance

3) in order to keep the acceptance for neutrons the new system has been checked: two 8 layers (1 veto + 7 active Scint./absorber) detectors



### Highly Granular Neutron time-of-flight Detector (HGND) with SiPM readout



- structure of layer: 3 cm Cu (absorber) + 2.5cm Scint. + 0.5cm (SiPM+FEE)
- size of scintillation detectors (cells): 4x4x2.5 cm<sup>3</sup>, 121 cells in each layer
- light readout: one SiPM with sensitive are 6 x 6 mm<sup>2</sup> per cell (EQR-15), measured time res. ~ 130ps
- total length of one HGND half-detector: ~ 48 cm (~1.5  $\lambda_{in}$ )



- height and angle adjustment
- adjustment of the distance between blocks

Total weight: ~800 kg

See presentation of A. Makhnev for details..



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See presentation of D. Finogeev for details..

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Time [ns]

Threshold

### **Readout scheme**

- 1.Plastic scintillator light flash
- 2.SiPM EQR15 11-6060D-S
- 3.High-speed comparator with differential LVDS output

4.FPGA-based TDC

= Response time + ToT

### Per channel

- Dynamic range: 0.5-7 MIP ٠
- Time resolution: 130 ps ٠
- Amplitude resolution: < 20%• (reconstructed from ToT)

F. Guber, et al., Instrum. Exp. Tech. 66 (2023) 4, 553-557. D. Finogeev, et al., Nucl. Instrum. Meth. A 1059 (2024) 168952. N. Karpushkin, et al., Nucl. Instrum. Meth. A 1068 (2024) 169739.

### SiPM: Beijing NDL EQR15 11-6060D-S

- •Active area  $6 \times 6 \text{ mm}^2$
- •Pixel size15×15 µm<sup>2</sup>
- •Total pixels: 160 000
- •PDE: 45%
- •Gain: 4\*10<sup>5</sup>

Single neutrons detection efficiency for different kinetic energies on the HGND surface

Efficiency = 1 – <u>Nevents without selected hits in HGND</u> Nevents

Hit selection: minimum 2 hits with > 3 MeV ( $\sim$ 1/2 MIP) signal



Comparison of primary neutrons rapidity and pT distributions BiBi@3A GeV on the HGND entrance surface for different positions DCM-QGSM-SMM



Reconstruction of neutron kinetic energy and energy resolution



Experimentally measured time resolution of the HGND scintillation cell ~ 150ps has been applied

BiBi@3A GeV

DCM-QGSM-SMM

ToF vs kinetic energy of different type of particles at the HGND 700cm, 10 deg -4.5 deg



BiBi@3A GeV

DCM-QGSM-SMM

number of registered primary neutrons at the HGND



During **1 month** of the BM@N run  $\sim 1.2*10^{9}$  single primary neutrons with kinetic energy > 300 MeV can be collected with 2 x HGNDs

Upper limit is **1.5** \* **10**<sup>9</sup> neutrons (additional multi-neutron event recognition is required).



- Neutron energy in events with only one primary neutron can be correctly reconstructed by determining fastest time in the HGND cells.
- Neutron energy reconstruction for events with more than 1 neutron in events requires development of more sophisticated methods of energy reconstruction.

### Status and steps of the HGND construction









Status of the HGND development Conclusions:

- 2-arms HGND detector system is proposed (7m from the target position and 10 deg angle)
- the performance study of the 2-arms HGND is done based on BiBi 3A GeV simulations
- neutron resonstruction algorythms (cluster, ML) is under development (A. Shabanov, V. Bocharnikov)
- construction of the HGND is ongoing (see presentation of A. Makhnev)
- FPGA based tdc electronics is under development now (see presentation of D. Finogeev)

Outlook:

- the HGND is planned to be ready next year (2025)

Members of HGND development group:

 INR RAS: D. Finogeev, M. Golubeva, F. Guber, A. Izvestnyy, N. Karpuskin, A. Makhnev, S. Morozov, P. Parfenov, I. Pshenichnov, S. Savenkov, D. Serebryakov, A. Shabanov, A. Svetlichnyi, V.Volkov, A. Zubankov JINR: S. Afanasiev, D. Sakulin, E. Sukhov, V. Ustivov,
 Kurchatov Institute NRC: P. Alexeev, A. Martemianov, A. Stavinsky, G. Taer, N. Zhigareva HSE: V. Bocharnikov, F. Ratnikov

### Thank you for your attention!

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### Backup

### Measurements of time resolution of scintillation detectors (scint + SiPM)

F.Guber et.al., Instruments and Experimental Techniques, 2023, Vol. 66, No. 4, pp. 553–557 (JINR + Hamamatsu, SensL photodetectors) F.Guber et.al., arXiv:2309.03614v1 [hep-ex] 7 Sep 2023 (JINR, EJ230 scint. + EQR photodetector)

**Photodetector:** 

### EQR15 11-6060D-S

(sensitive area -  $6x6 \text{ mm}^2$ , 15mkm pixel pitch,  $160\ 000$  pixels, PDE - 45%, gain -  $4x10^5$ )

### Scintillator:

- 1) JINR produced (40x40x25mm<sup>3</sup>), 1.5% paraterphenyl and 0.01% POPOP) with light time decay of  $3.9 \pm 0.7$  ns
- 2) EJ230 with light time decay of  $2.8 \pm 0.5$  ns

**FEE:** LMH6629MF preamp (20 dB gain, bandwidth of 600 MHz at a 3 dB level, and noise of  $<2.2 \text{ nV}/\sqrt{\text{Hz}}$ ) + rapid discriminator (ADCMP553) with a fixed threshold.

Readout: CAEN DT5742

Test results on e-beam at LPI

JINR scintillators will be used for the HGN detector because they are available and significantly cheaper than EJ230.

## Symmetry energy in high-density region



- Nuclotron-NICA density region:  $2 \lesssim n_B/n_0 \lesssim 8$
- Symmetry energy  $E_{sym}$  has strong density dependence and can be described with its slope *L*:  $L = 3\rho \frac{dE_{sym}(\rho)}{dE_{sym}(\rho)}$

What observables can we use to extract information about *L*?

# Using dN/dy(n,p), $dN/dE_{kin}(n,p)$ to study L



X.X. Long, G.F. Wei, Phys.Rev.C 109 (2024) 5, 054619

Rapidity and kinetic energy distributions of n/p ratios can be used to study *L* 

- n/p ratios show strong dependence on L and significantly weaker dependence on K<sub>0</sub>
- n/p ratios require less statistics than anisotropic flow measurements

Neutron measurements are required in order to extract robust information about symmetry energy

BiBi@3A GeV

DCM-QGSM-SMM

### Estimation of primary neutrons count rate at BiBi@3 AGeV run

Beam rate - 10<sup>6</sup> per spill, Duty factor of the beam - 50% Efficiency of accelerator operation - 70%

Target interaction length - 2%, Mean primary neutron yield: 0.17 (single) - 0.23 (all) neutron / interaction Mean efficiency of the HGND detector - 50%

During 1 month of the BM@N run ~  $1.2*10^{\circ}$  single primary neutrons with kinetic energy > 300 MeV can be collected with 2 x HGNDs

Upper limit is **1.5** \* **10**<sup>9</sup> neutrons (additional multi-neutron event recognition is required).

#### nDet:

**780 cm,** 44x44 cm, 121 mods, 4x4 cm, 100000 ev., **w/o magnetic field, 45.8 cm** PLA 0.2cm + **Veto** 2.5 cm + PCB 0.2cm + PLA 0.2cm + **7 layers (Cu** 3 cm + PLA 0.2cm + **Sc** 2.5 cm + PCB 0.2cm + PLA 0.2cm) **Time cut in nDet:** time < **55** nsec (in simulations) Vac. in cave, **BOX generator, neutrons huge spot, multiplicity=1** 

Neutrons kinetic energy [GeV]:

### Neutrons reconstructed kinetic energy (Cu absorbers)

*Kinetic energy is reconstructed with hit with min Time With Veto* 

### **Corrected module geometry**

Edep > 3 MeV in cells >1 hit in nDet Time resol 150 ps



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