# Anisotropic flow at high-baryon density

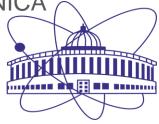
### Arkadiy Taranenko (NRNU MEPhI, JINR)





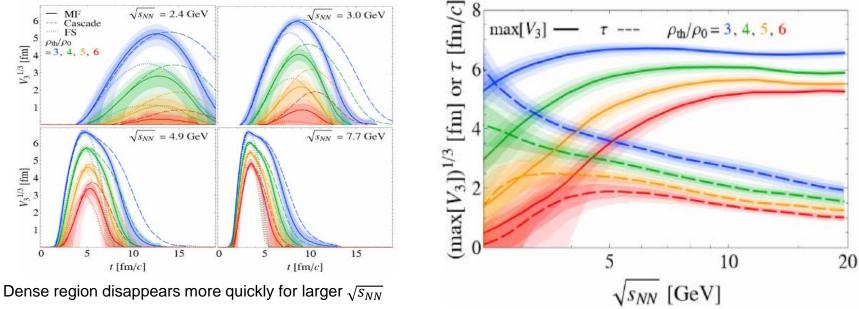
13th Collaboration Meeting of the BM@N Experiment at NICA 8-10 October 2024, JINR

The work has been supported by the Ministry of Science and Higher Education of the Russian Federation, Project "Fundamental and applied research at the NICA megascience experimental complex" № FSWU-2024-0024



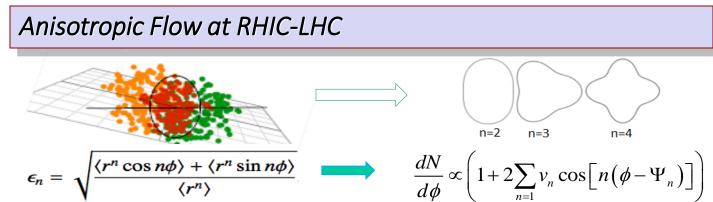
# **Optimal collision energy for realizing high baryon-density matter**

H. Taya, A. Jinno, M. Kitazawa, Y. Nara https://arxiv.org/abs/2409.07685

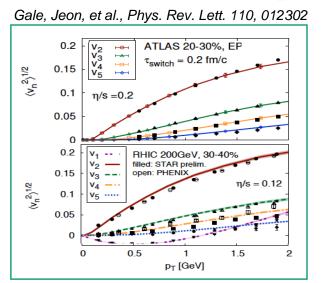


 $\sqrt{s_{NN}}$  dependence of the maximum volume max[V3] (solid) and the lifetime  $\tau$  (dashed)

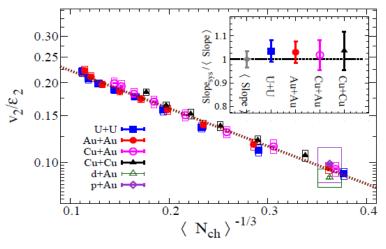
The optimal energy is around  $\sqrt{s_{NN}}$ =3-4GeV, where a baryon density  $\rho/\rho_0$  = 3 nuclear density is realized with a substantially large space-time volume. Higher and lower energies are disfavored due to short lifetime and low density



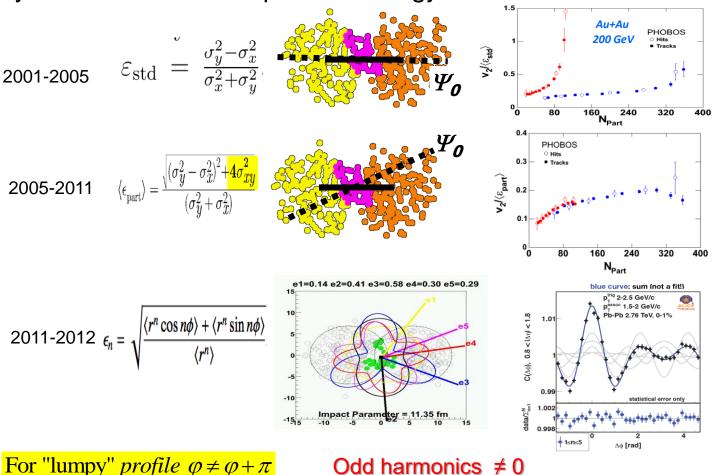
#### Initial eccentricity (and its attendant fluctuations) $\epsilon_n$ drive momentum anisotropy $v_n$ with specific viscous modulation



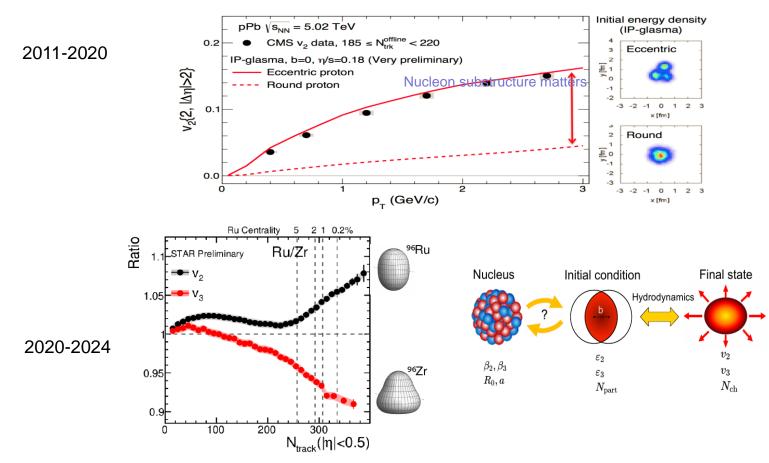
STAR, Phys. Rev. Lett. 122 (2019) 172301



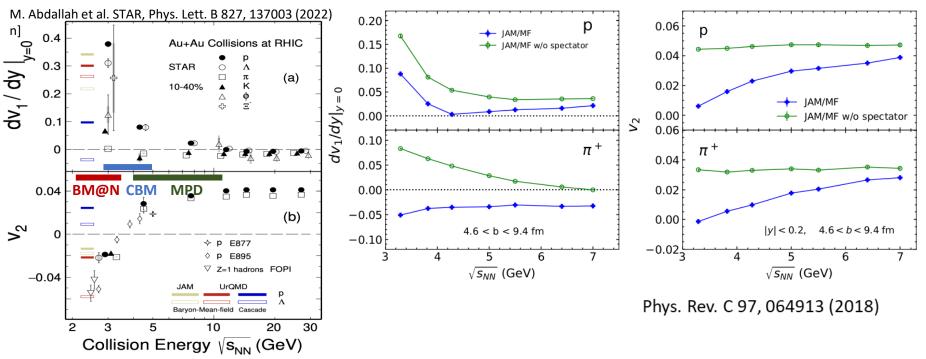
### System size scan at top RHIC energy



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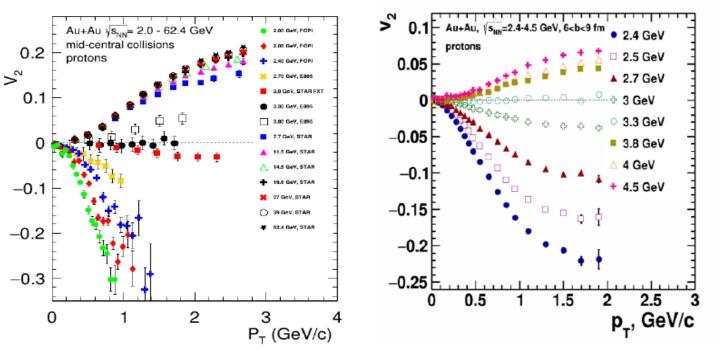
### Anisotropic flow in heavy-ion collisions at high baryon density



#### Anisotropic flow at FAIR/NICA energies is a delicate balance between:

- I. The ability of pressure developed early in the reaction zone ( $t_{exp} = R/c_s$ ,  $c_s = c\sqrt{dp/d\varepsilon}$ ) and
- II. The passage time for removal of the shadowing by spectators ( $t_{pass} = 2R/\gamma_{CM}\beta_{CM}$ )

### Anisotropic flow in Au+Au collisions at Nuclotron-NICA energies

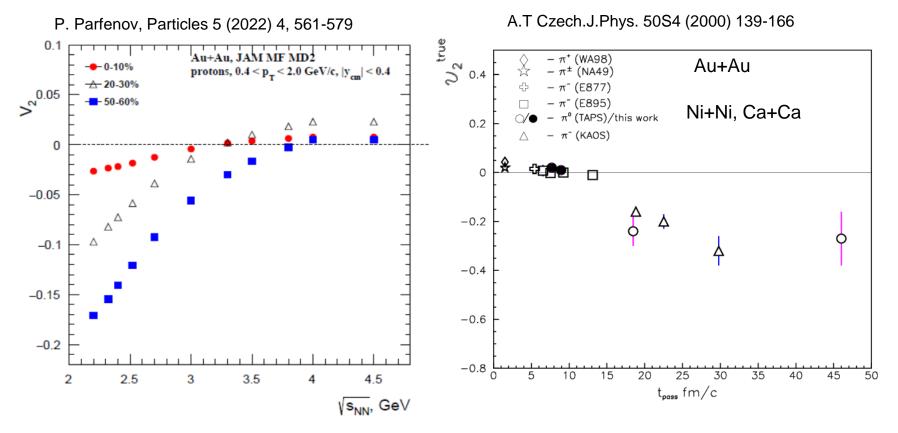


M. Mamaev, Particles 6 (2023) 2, 622-637

Anisotropic flow at FAIR/NICA energies is a delicate balance between:

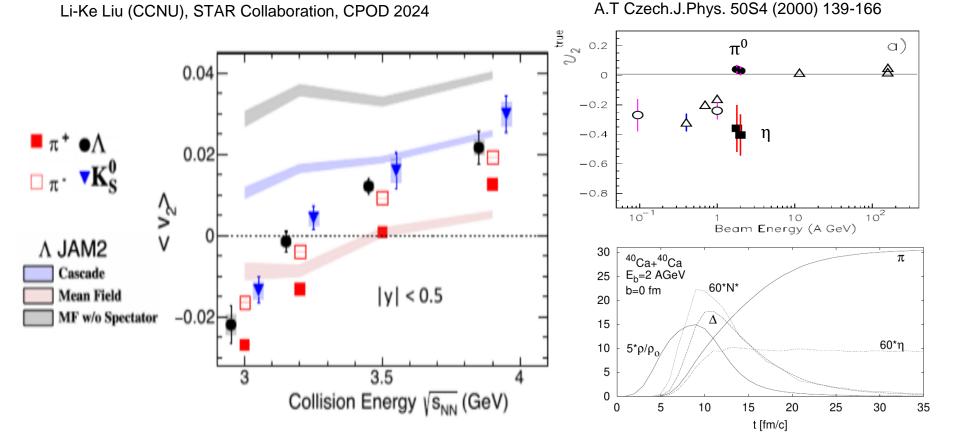
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### Elliptic flow: transition from out-of-plane to in-plane: geometry

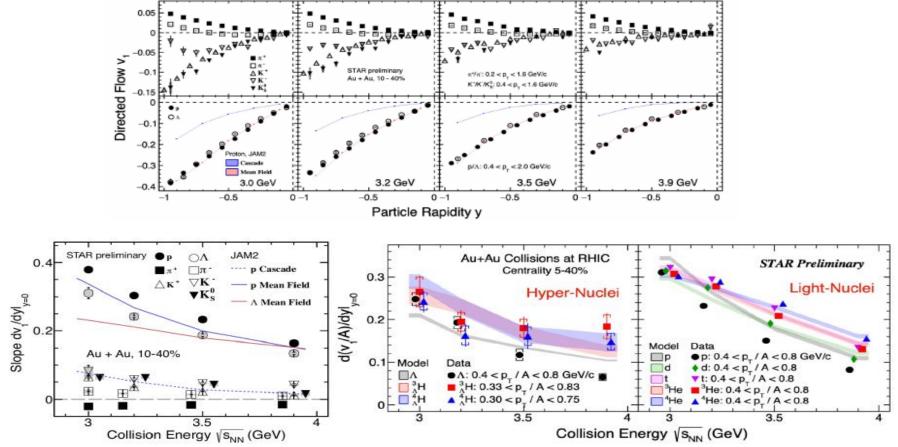


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# Elliptic flow: transition from out-of-plane to in-plane: PID



### STAR preliminary results from BES – II program – directed flow



#### Vn of protons in Au+Au collisions at 2.4 GeV - HADES

#### **Determination of EOS**

New level of precision - multi differential Additional information from higher orders

#### Models:

JAM 1.9 NS3 (hard EOS, mom.-indep.) JAM 1.9 MD1 (hard EOS, mom.-dep.) JAM 1.9 MD4 (soft EOS, mom.dep.) UrQMD 3.4 (hard EOS, mom.-indep.) GiBUU Skyrme 12 (soft EOS)

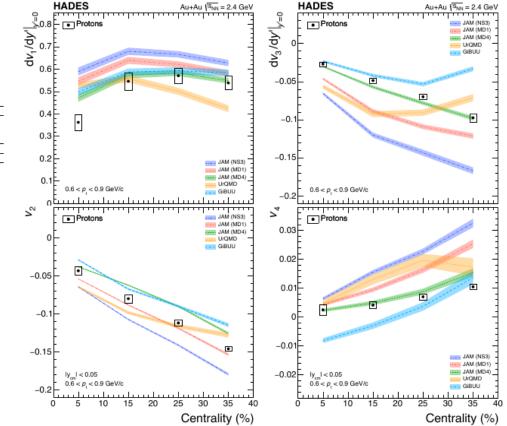
EOS	K (MeV)	$m^*/m$	mom-dep.
NS1	380	0.83	no
MD1	380	0.65	yes
MD4	210	0.83	yes
Hard	380		no
Skyrme 12	240	0.75	no
	NS1 MD1 MD4 Hard	NS1 380   MD1 380   MD4 210   Hard 380	NS1 380 0.83   MD1 380 0.65   MD4 210 0.83   Hard 380 0.65

#### Conclusions

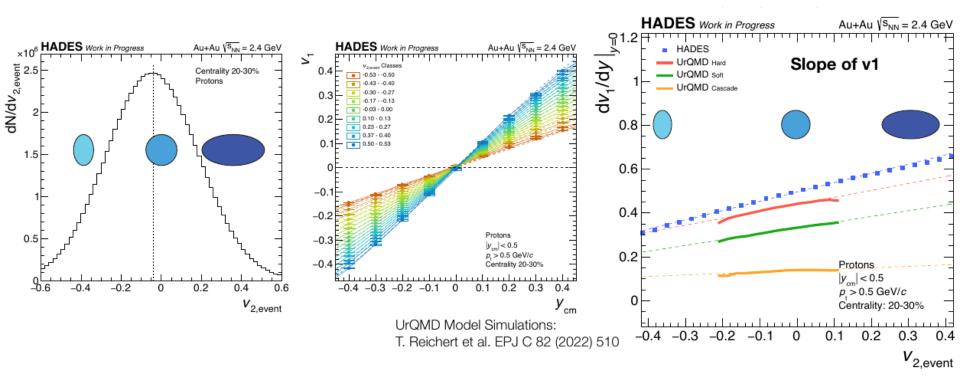
Overall trend reasonably described, but no model works everywhere

Several systematic deviations can be linked to different implementation in transport codes

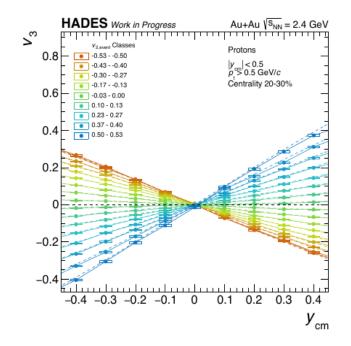
Mechanism of light nuclei production is essential for the description of the data



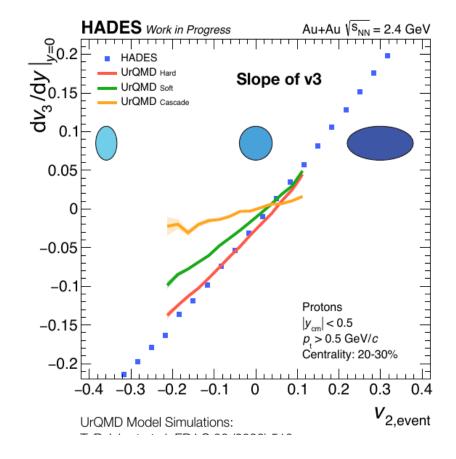
#### New HADES results on flow correlations



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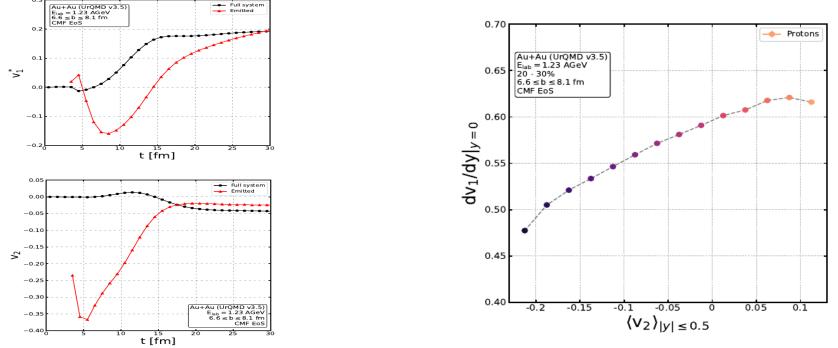


Slope of the Triangular Flow v<sub>3</sub> A strong sensitivity to the EoS is seen



### Decoding the flow evolution using hadron $V_1$ - $V_2$ correlations and dileptons

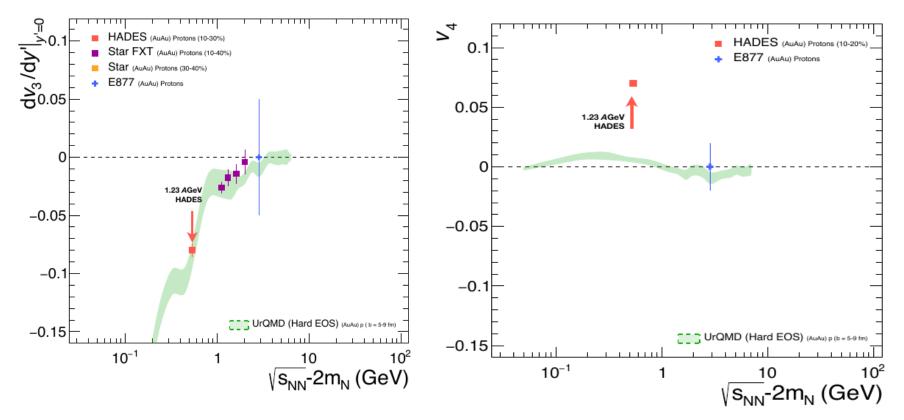
Tom Reichert et al., Phys.Lett.B 841 (2023) 137947



The elliptic flow is initially positive (v2 > 0) due to the early pressure gradient. This positive v2 transfers its momentum to the spectators, which creates the directed flow v1.

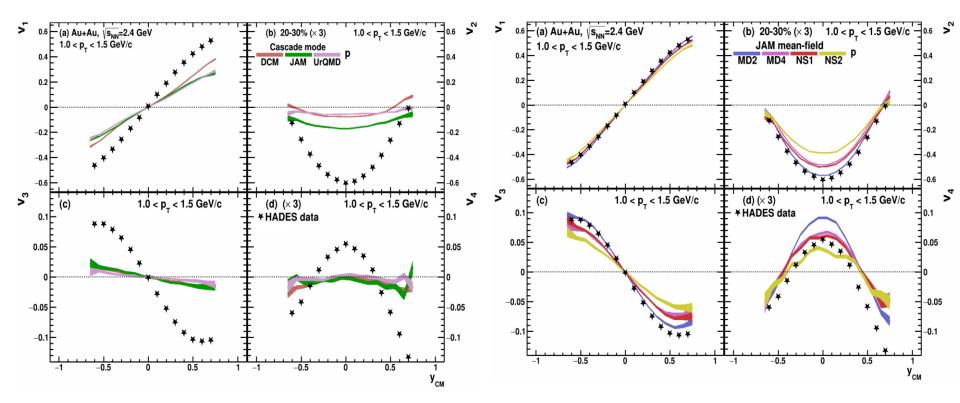
In turn, the spectator shadowing of the in-plane expansion results in a negative v2 for the observable final state hadrons. propose a measurement of v1–v2 flow to pin down this evolution pattern.

### Collision energy dependence of v3 and v4



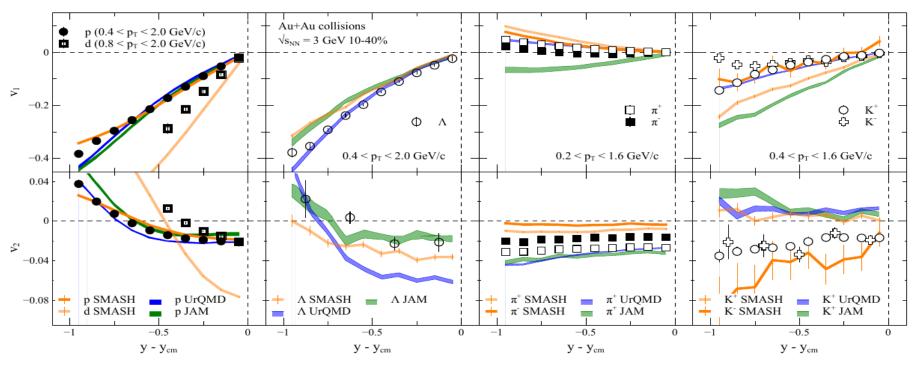
#### Flow of protons - models vs data

P. Parfenov, Particles 5 (2022) 4, 561-579



# **Describing proton flow is not enough**

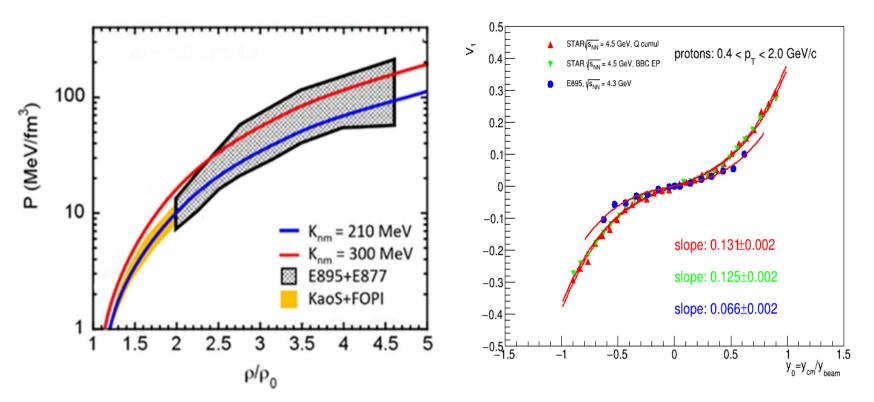
Prog.Part.Nucl.Phys. 134 (2024) 104080



Strange baryons are not well described

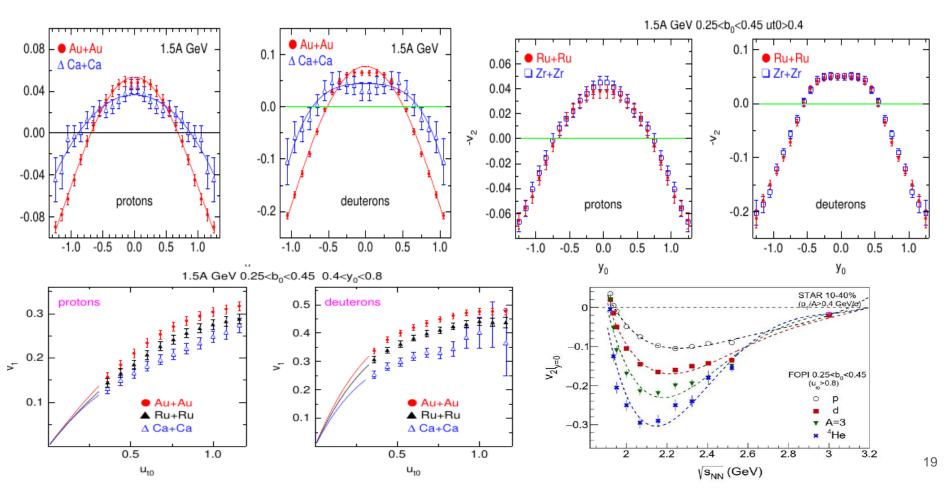
- the results may depend on:
- nucleon-hyperon and hyperon-hyperon interactions
- · in-medium modifications of interactions

Pions and kaons NOT described! Not very surprising: UrQMD, JAM, and SMASH don't have mean-fields for mesons The main source of existing systematic errors in  $v_n$  measurements

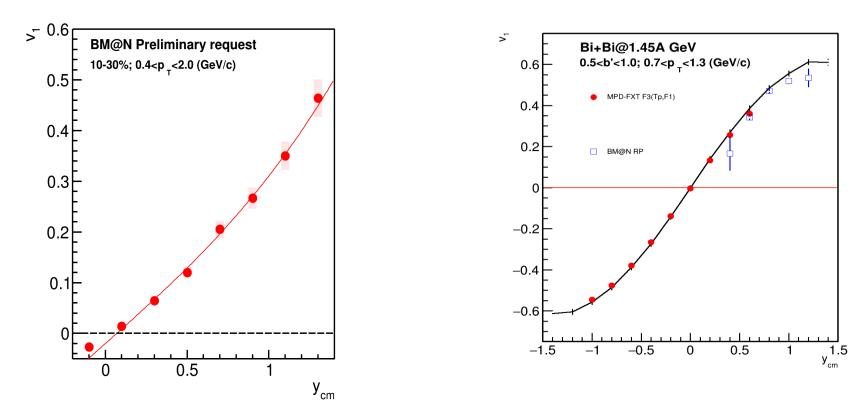


The main source of existing systematic errors in  $v_n$  measurements is the difference between results from different experiments (for example, FOPI and HADES, E895 and STAR) <sup>18</sup>

# FOPI Ion-Ion collisions at 1.5 AGeV



# Directed flow of protons: BM@N – MD FXT

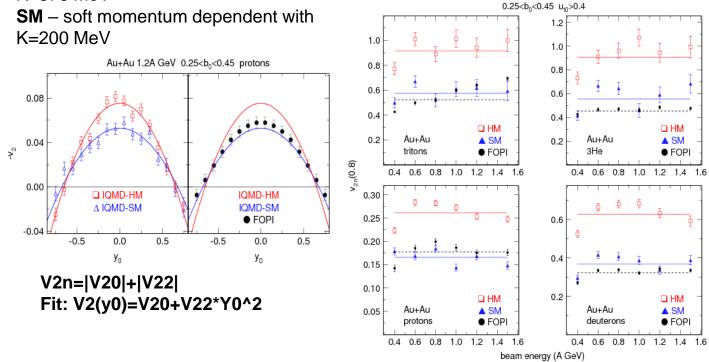


#### Rapidity dependence of v2 and EOS

**HM** – stiff momentum dependent with K=376 MeV

**SM** – soft momentum dependent with K=200 MeV

FOPI data : Nucl. Phys. A 876 (2012) 1 IQMD : Nucl Phys. A 945 (2016)



#### Large rapidity coverage is important for flow measurements

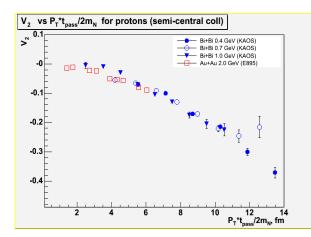
# Summary and outlook

- Measurements of anisotropic flow, flow fluctuations, correlations between flow of different harmonics are sensitive to many details of the initial conditions and the system evolution. It may provides access to the transport properties of the medium: EOS, sound speed (cs), viscosity, etc.
- v<sub>n</sub> at energies 2.5-11 GeV (SIS, STAR BESII, NICA, FAIR) shows strong energy dependence: possible transition between hadronic and partonic matter.
- System size scan is very important in order to understand the effect of spectators on the experimental observables

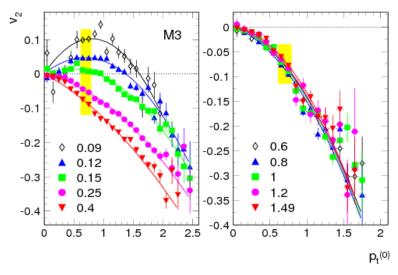


#### v<sub>2</sub> Flow at SIS-AGS: scaling relations

(KAOS – Z. Phys. A355 (1996); (E895) - PRL 83 (1999) 1295 V<sub>2</sub> vs P<sub>T</sub> for protons (semi-central coll) 0.1 > Bi+Bi 0.4 GeV (KAOS) Bi+Bi 0.7 GeV (KAOS) Bi 1.0 GeV (KAOS) Au 2.0 GeV (E895 -0.1 -0.2 -0.3 -0.4 0.2 0.4 0.6 0.8 1.2 P<sub>+</sub>, GeV/c

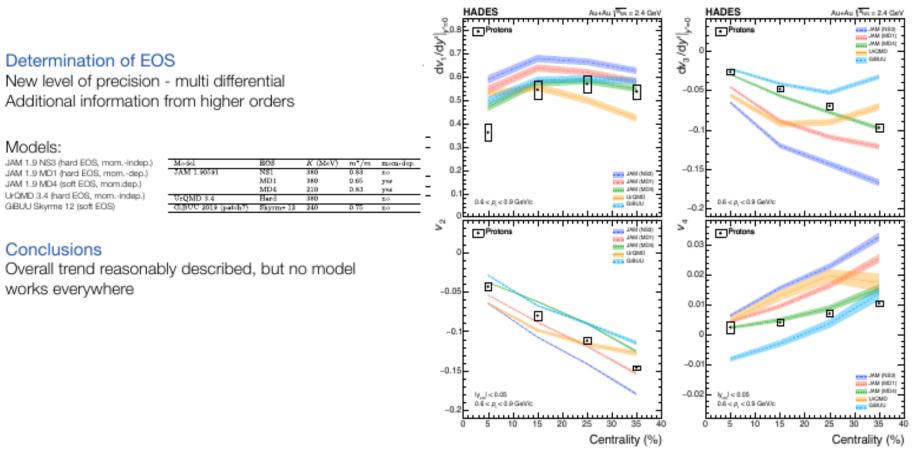


#### *FOPI:* v<sub>2</sub> of protons from *Elab=0.09 to 1.49 GeV* Phys.Lett. B612 (2005) 173-180



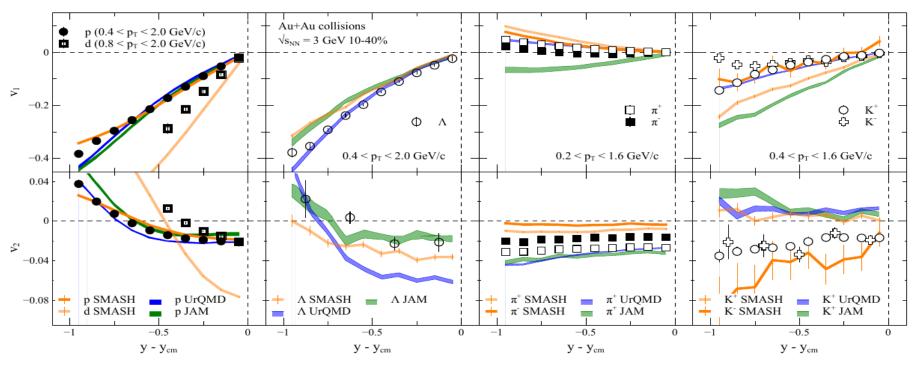
The rather good scaling observed suggest that  $c_s$  does not change significantly over beam energy range 0.4 – 2.0 AGeV.

# Vn of protons in Au+Au collisions at 2.4 GeV - HADES



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