

# Energy/System Size/Rapidity Scan - Flow

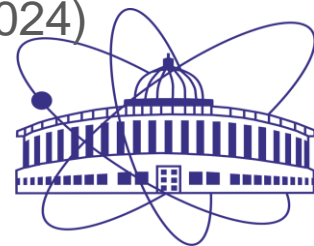
Arkadiy Taranenko  
(NRNU MEPHI, JINR)

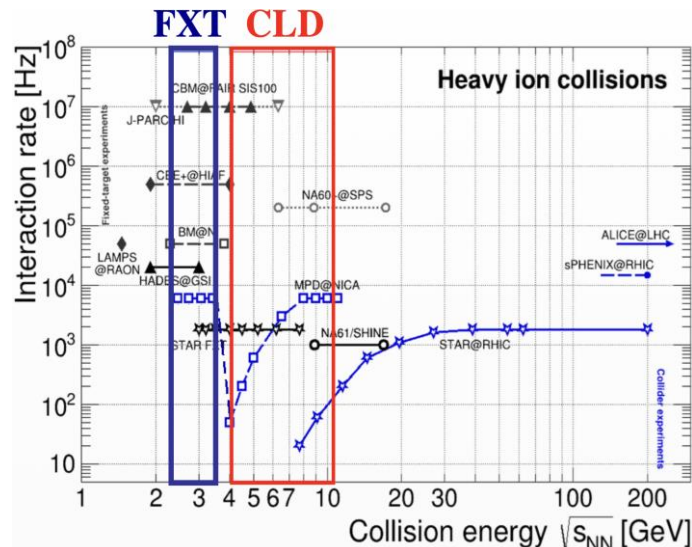
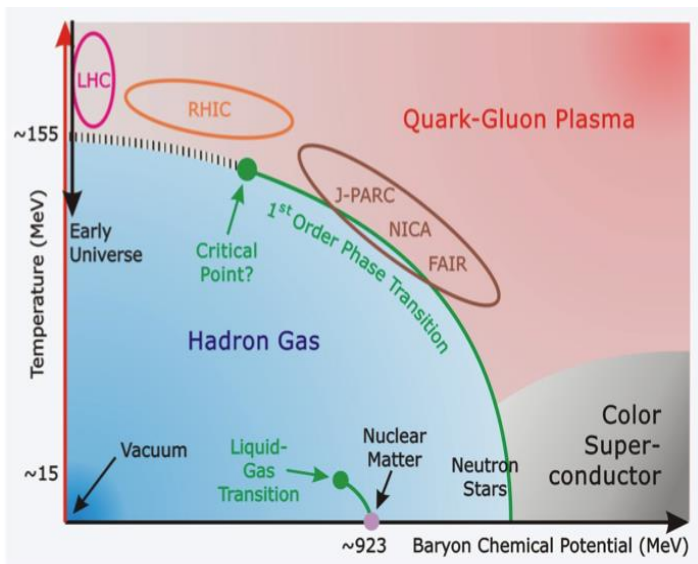


Workshop on physics performance studies at NICA (NICA-2024)  
25-27 November 2024, JINR-MEPHI



The work has been supported by the Ministry of Science and Higher Education of the Russian Federation, Project "Fundamental and applied research at the NICA megascience experimental complex" № FSWU-2024-0024



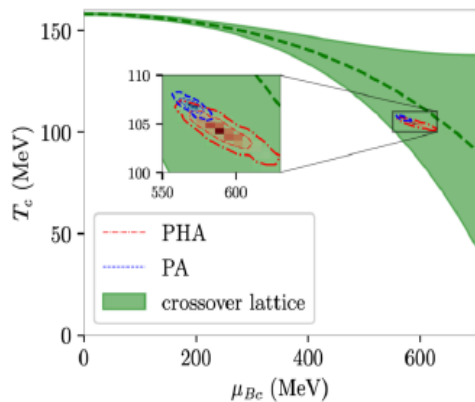


Experiments BM@N ( $\sqrt{s_{NN}} = 2.3-3.3 \text{ GeV}$ ) and MPD ( $\sqrt{s_{NN}} = 4-11 \text{ GeV}$ )

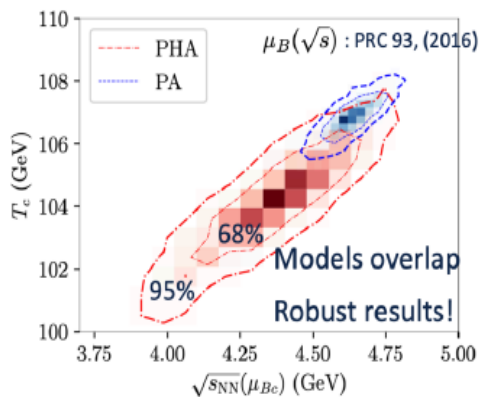
Ion sources: C (A=12), N (A=14), Ar (A=40), Fe (A=56), Kr (A=78-86), Xe (A=124-134), Bi (A=209). 2



# Location of the QCD Critical Point : Theoretical Estimation/Prediction



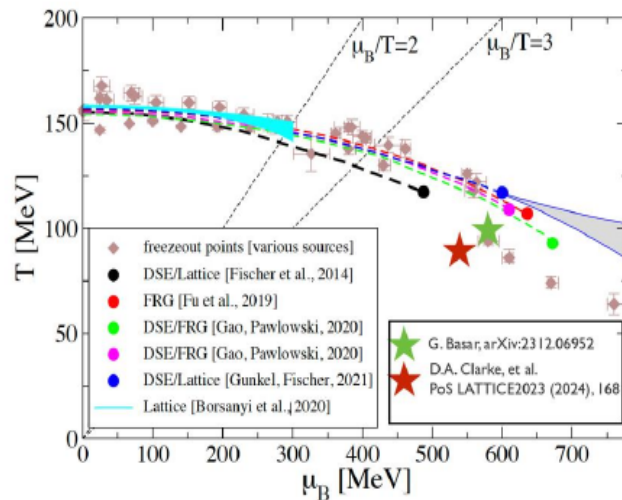
Holography+ Bayesian : Hippert et al., arXiv : 2309.00579



## CPOD2024

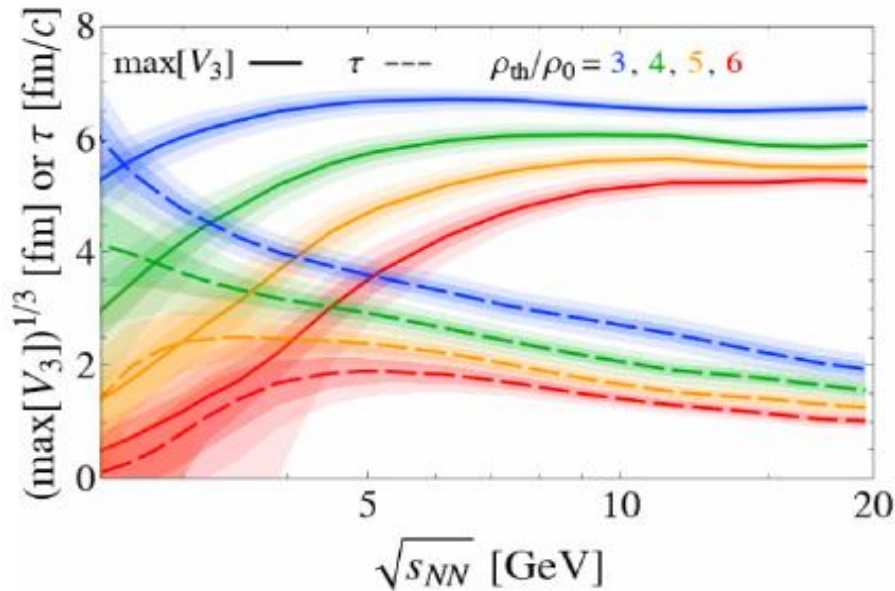
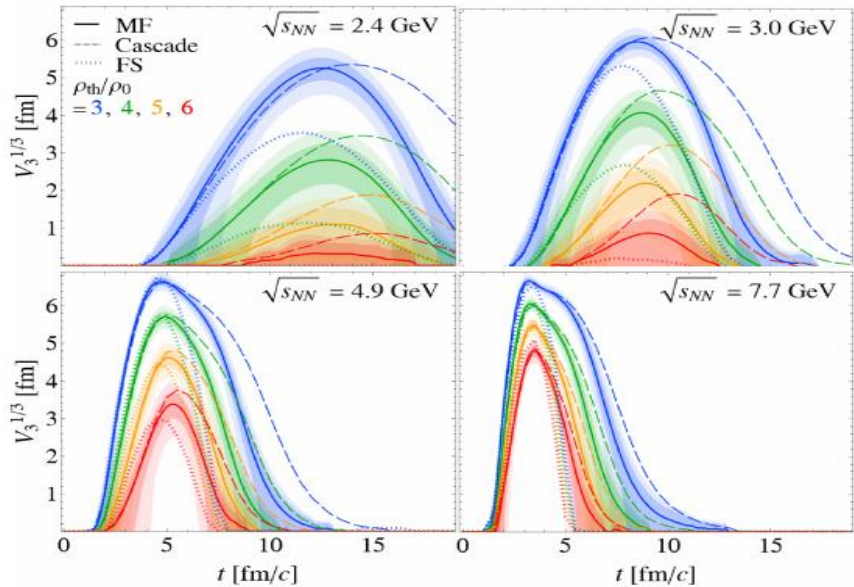
Method	$\mu_c$ (MeV)	$T_c$ (MeV)
Holography + Bayesian	560 - 625	101 - 108
FRG/DSE	495 - 654	108 - 119
Lee-Yang edge singularities	500 - 600	100 - 105
Lattice QCD	$\mu_c/T_c > 3$	F. Karsch et al.
<b>Summary</b>	<b>495 - 654</b>	<b>100 - 119</b>

$(\mu_c, T_c) = (495 - 654, 100 - 119) \text{ MeV} \longrightarrow 3.5 < \sqrt{s_{NN}} < 4.9 \text{ GeV}$



# Optimal collision energy for realizing high baryon-density matter

H. Taya, A. Jinno, M. Kitazawa, Y. Nara, <https://arxiv.org/abs/2409.07685>



Dense region disappears more quickly for larger  $\sqrt{s_{NN}}$

$\sqrt{s_{NN}}$  dependence of the maximum volume  $\max[V_3]$  (solid) and the lifetime  $\tau$  (dashed)

The optimal energy is around  $\sqrt{s_{NN}}=3-5$  GeV, where a baryon density  $\rho/\rho_0 = 3$  nuclear density is realized with a substantially large space-time volume. Higher and lower energies are disfavored due to short lifetime and low density

# EOS for high baryon density matter

The binding energy per nucleon:  $E_A(\rho, \delta) = E_A(\rho, 0) + E_{sym}(\rho)\delta^2 + O(\delta^4)$

Isospin asymmetry:

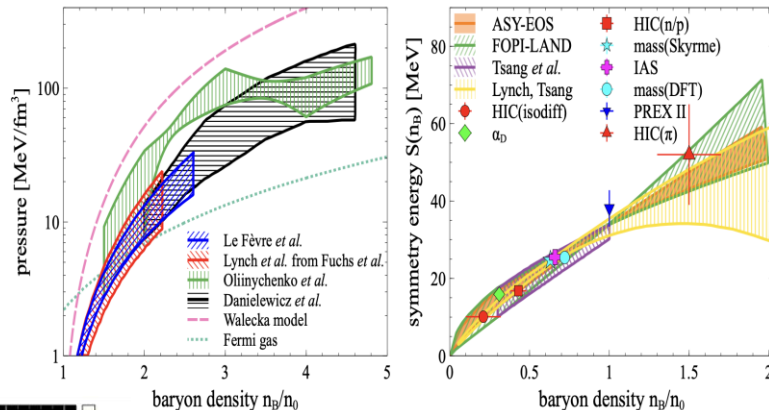
$$\delta = (\rho_n - \rho_p)/\rho$$

Symmetric  
matter

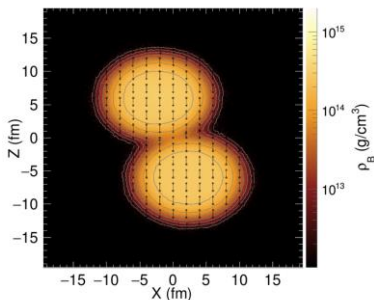
Symmetry energy

## Heavy-Ion Collisions and Merging Neutron Stars

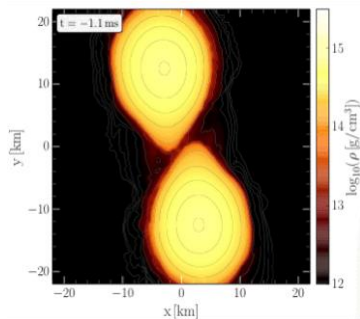
- $T < 70$  MeV,  $\rho \approx 3\rho_0$  for both
- Probe NS merger matter in the laboratory



Au+Au 1.25A GeV



NS mergers

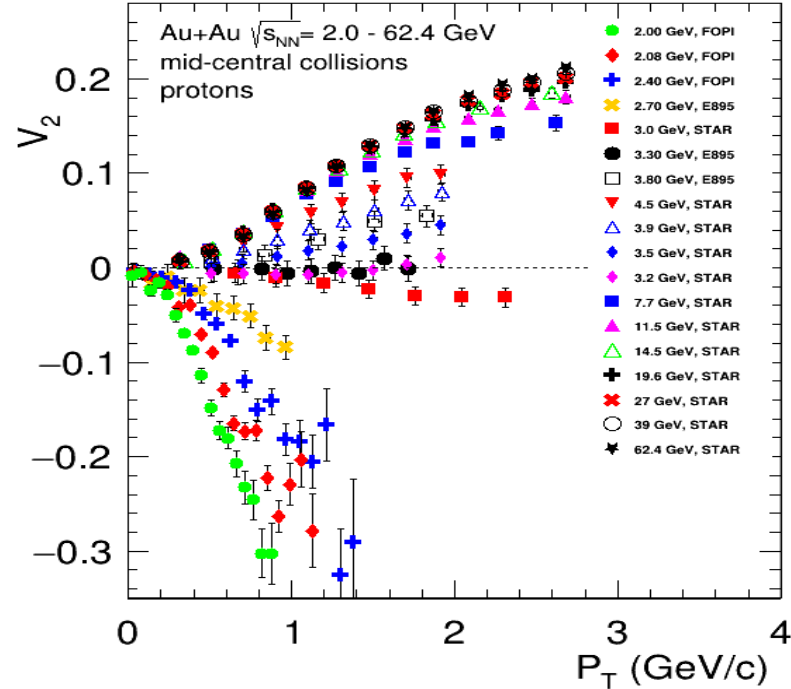
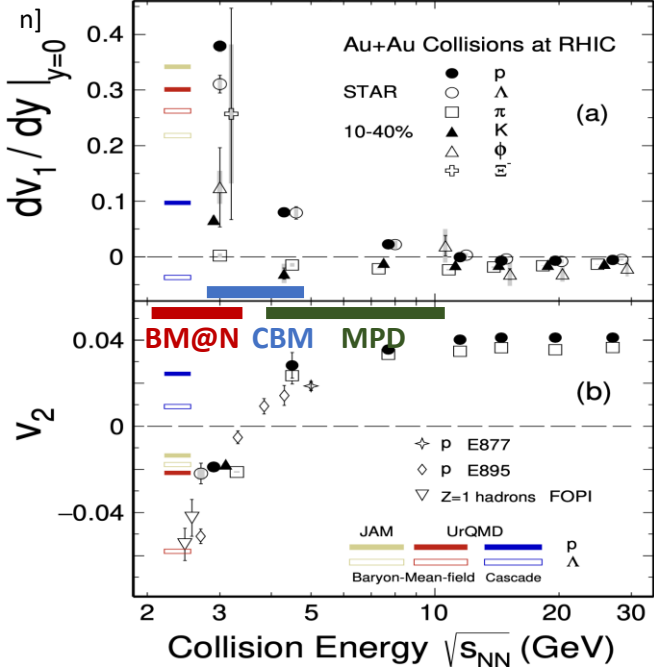


A. Sorensen *et al.*, *Prog.Part.Nucl.Phys.* 134 (2024) 104080

**New data is needed to further constrain transport models with hadronic d.o.f.**

# Anisotropic flow in heavy-ion collisions at high baryon density

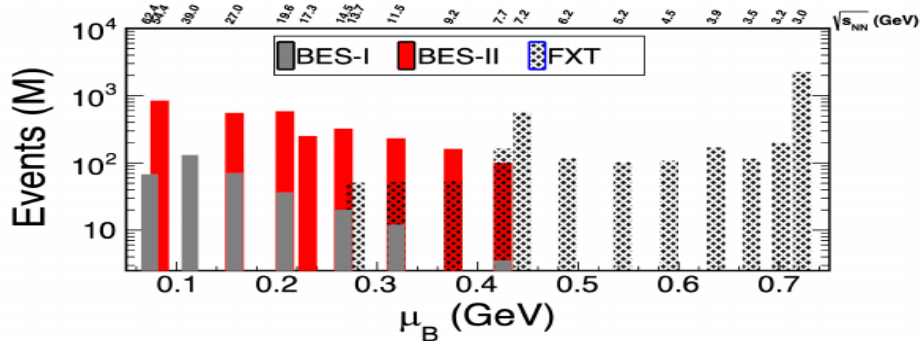
M. Abdallah et al. STAR, Phys. Lett. B 827, 137003 (2022)



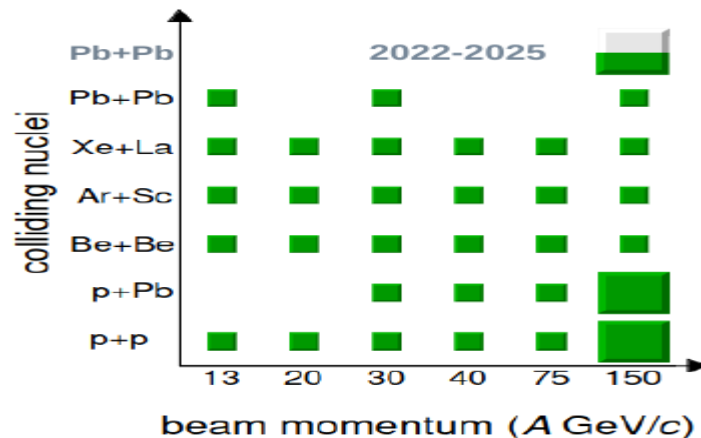
**Anisotropic flow at NICA energies : strong energy dependence**

# Beam Energy / System size Scan programs

STAR at RHIC:  $3 < \sqrt{s_{NN}} < 200$  GeV ( $750 < \mu_B < 25$  MeV)



NA61SHINE at SPS:  $5.1 < \sqrt{s_{NN}} < 17$  (27) GeV



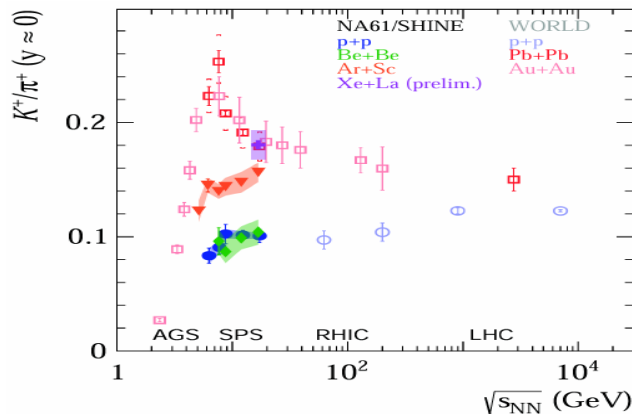
Au+Au Collisions at RHIC

Collider Runs

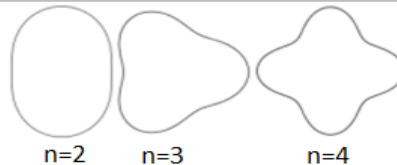
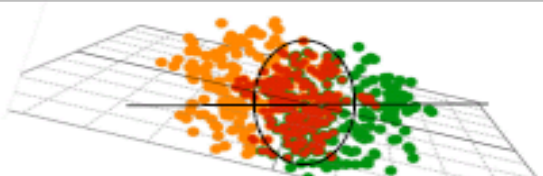
Fixed-Target Runs

Collider Runs					Fixed-Target Runs						
	$\sqrt{s_{NN}}$ (GeV)	#Events	$\mu_B$	$\gamma_{beam}$	run		$\sqrt{s_{NN}}$ (GeV)	#Events	$\mu_B$	$\gamma_{beam}$	run
1	200	380 M	25 MeV	5.3	Run-10, 19	1	13.7 (100)	50 M	280 MeV	-2.69	Run-21
2	62.4	46 M	75 MeV		Run-10	2	11.5 (70)	50 M	320 MeV	-2.51	Run-21
3	54.4	1200 M	85 MeV		Run-17	3	9.2 (44.5)	50 M	370 MeV	-2.28	Run-21
4	39	86 M	112 MeV		Run-10	4	7.7 (31.2)	260 M	420 MeV	-2.1	Run-18, 19, 20
5	27	585 M	156 MeV	3.36	Run-11, 18	5	7.2 (26.5)	470 M	440 MeV	-2.02	Run-18, 20
6	19.6	595 M	206 MeV	3.1	Run-11, 19	6	6.2 (19.5)	120 M	490 MeV	1.87	Run-20
7	17.3	256 M	230 MeV		Run-21	7	5.2 (13.5)	100 M	540 MeV	-1.68	Run-20
8	14.6	340 M	262 MeV		Run-14, 19	8	4.5 (9.8)	110 M	590 MeV	-1.52	Run-20
9	11.5	157 M	316 MeV		Run-10, 20	9	3.9 (7.3)	120 M	633 MeV	-1.37	Run-20
10	9.2	160 M	372 MeV		Run-10, 20	10	3.5 (5.75)	120 M	670 MeV	-1.2	Run-20
11	7.7	104 M	420 MeV		Run-21	11	3.2 (4.59)	200 M	699 MeV	-1.13	Run-19
						12	3.0 (3.85)	2000 M	750 MeV	-1.05	Run-18, 21

beam momentum ( $A$  GeV/c)



# Anisotropic Flow at RHIC-LHC



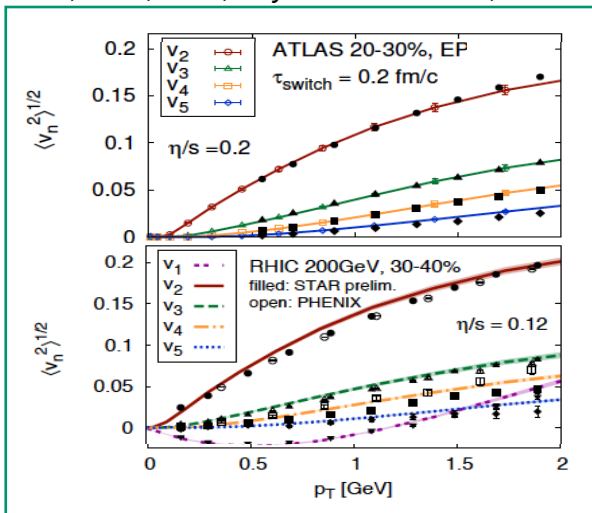
$$\epsilon_n = \sqrt{\frac{\langle r^n \cos n\phi \rangle + \langle r^n \sin n\phi \rangle}{\langle r^n \rangle}}$$



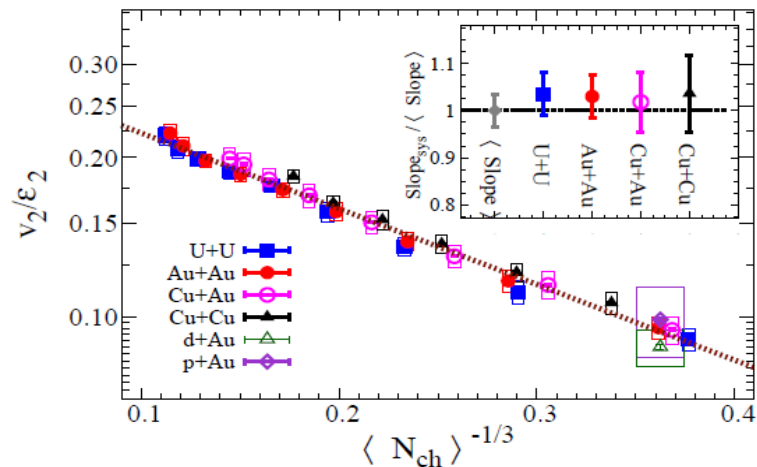
$$\frac{dN}{d\phi} \propto \left( 1 + 2 \sum_{n=1} v_n \cos[n(\phi - \Psi_n)] \right)$$

**Initial eccentricity (and its attendant fluctuations)  $\epsilon_n$  drive momentum anisotropy  $v_n$  with specific viscous modulation**

Gale, Jeon, et al., Phys. Rev. Lett. 110, 012302



STAR, Phys. Rev. Lett. 122 (2019) 172301





# Anisotropic Flow at RHIC/LHC is acoustic

PRC 84, 034908 (2011)  
P. Staig and E. Shuryak.

- $v_n$  measurements are sensitive to system shape ( $\epsilon_n$ ), system size ( $RT$ ) and transport coefficients  $\left(\frac{\eta}{s}, \frac{\zeta}{s}, \dots\right)$ .

arXiv:1305.3341  
Roy A. Lacey, et al.

- Acoustic ansatz

✓ Sound attenuation in the viscous matter reduces the magnitude of  $v_n$ .

- Anisotropic flow attenuation,

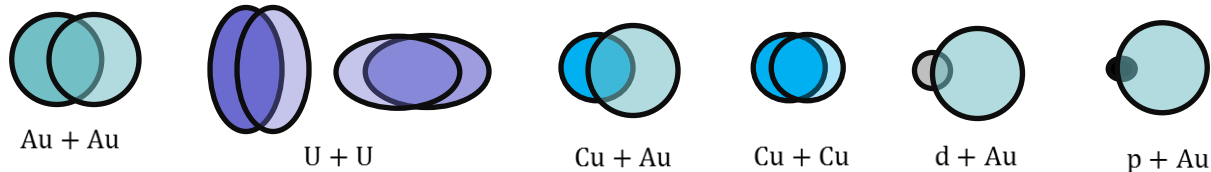
$$\frac{v_n}{\epsilon_n} \propto e^{-\beta n^2}, \quad \beta \propto \frac{\eta}{s} \frac{1}{RT}$$

arXiv:1601.06001  
Roy A. Lacey, et al.

- From macroscopic entropy considerations  $S \sim (RT)^3 \propto \frac{dN}{d\eta}$

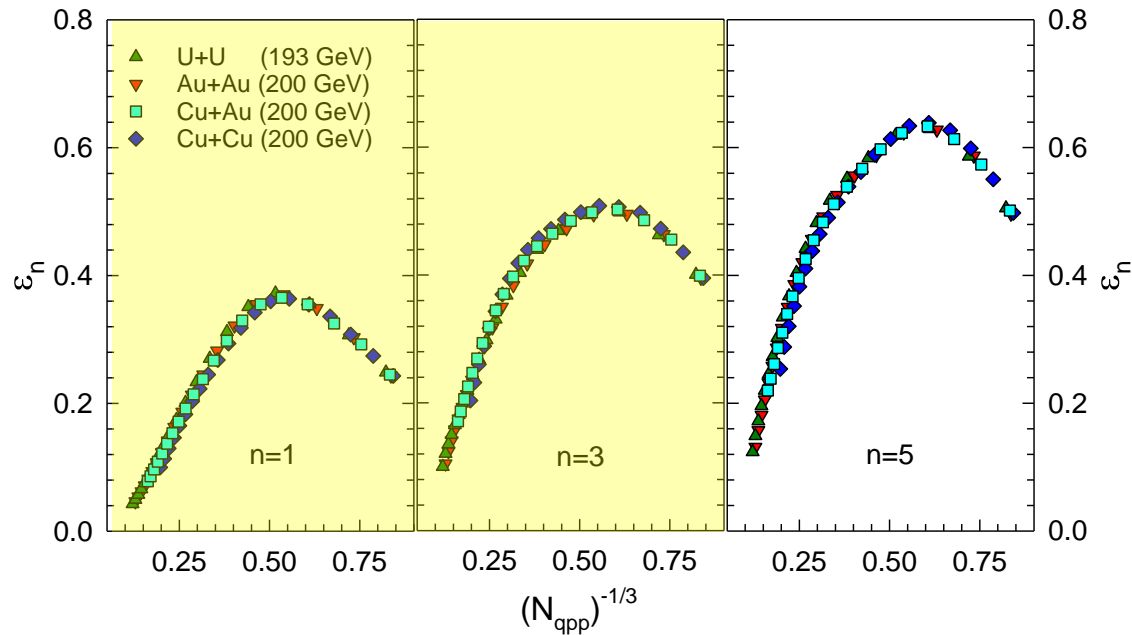
PRC 88, 044915 (2013)  
E. Shuryak and I. Zahed

$$\ln\left(\frac{v_n}{\epsilon_n}\right) \propto A \frac{\eta}{s} \left(\frac{dN}{d\eta}\right)^{-\frac{1}{3}}$$



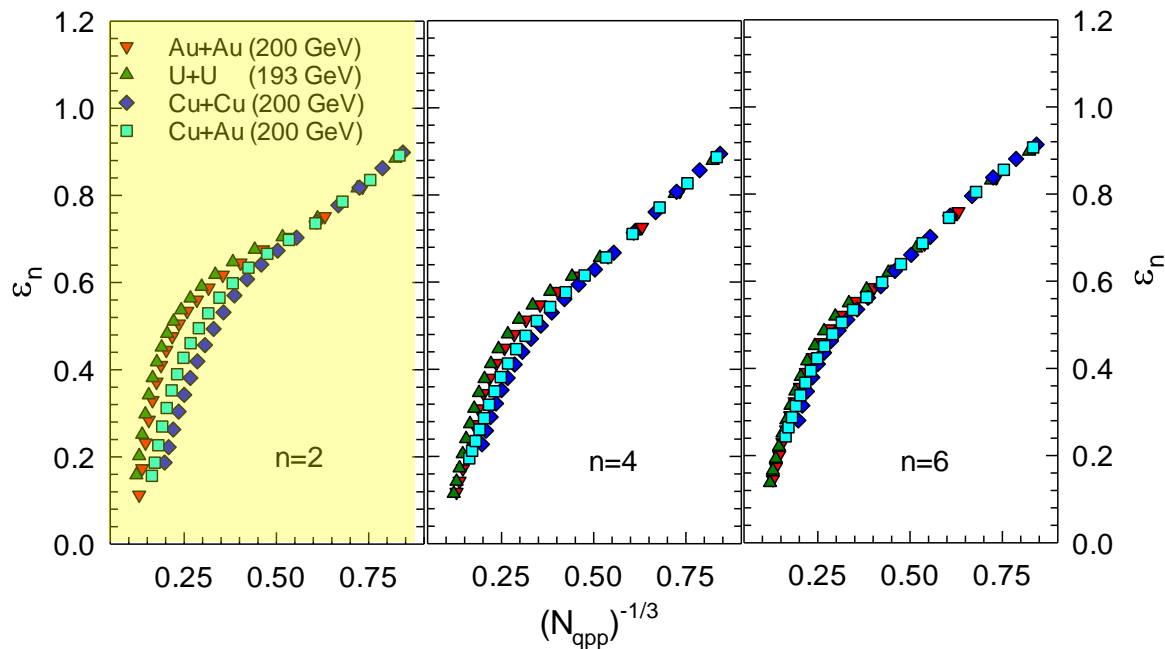
Scaling expected For *similar*  $\frac{\eta}{s}$  and  $\frac{dN}{d\eta}$

## Acoustic Flow – Expected Shape Response

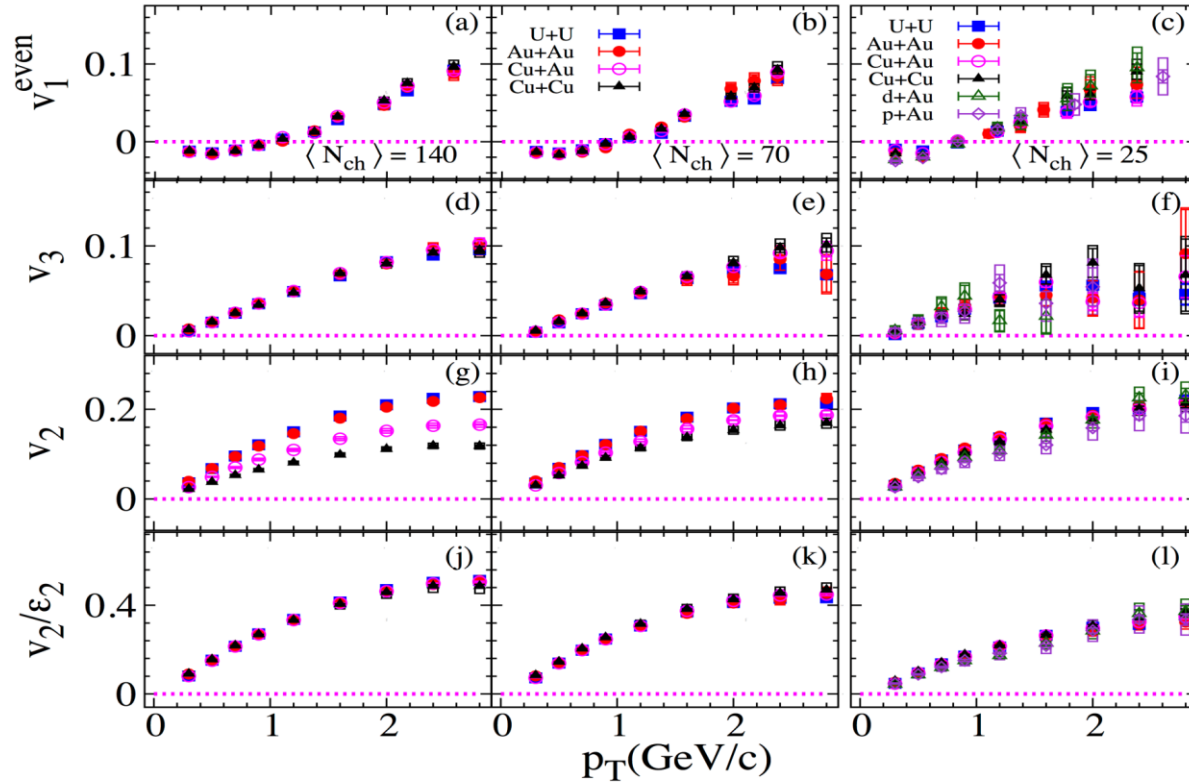


- **Odd eccentricity moments are fluctuations driven**
  - ✓ Little, if any, system dependence for A+A(B) collisions for similar geometric size

## Acoustic Flow – Expected Shape Response



- **Even eccentricity moments are shape driven**
  - ✓ **Sizeable system dependence for A+A(B) collisions in central & mid-central collisions**
  - ✓ **System independence in peripheral collisions**

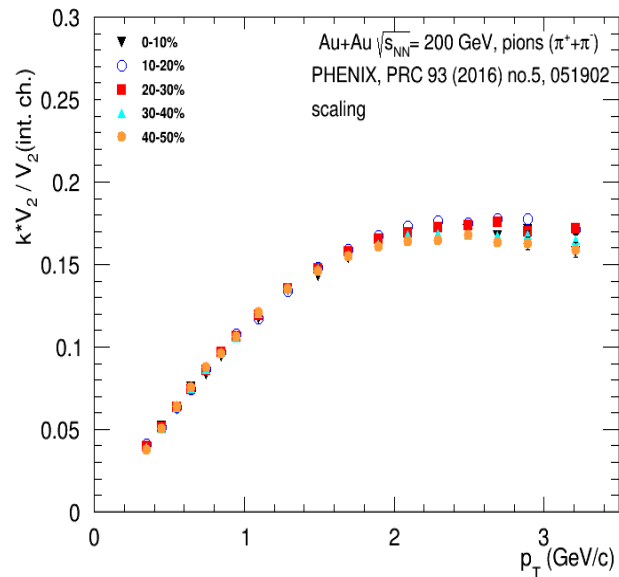
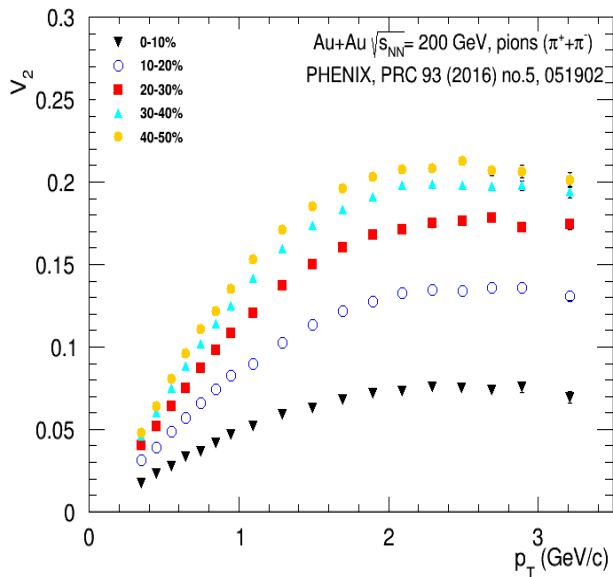


For a fixed  $\langle N_{ch} \rangle$ , the  $v_2/\epsilon_2$  are similar for different systems

Even harmonics are system dependent. 1

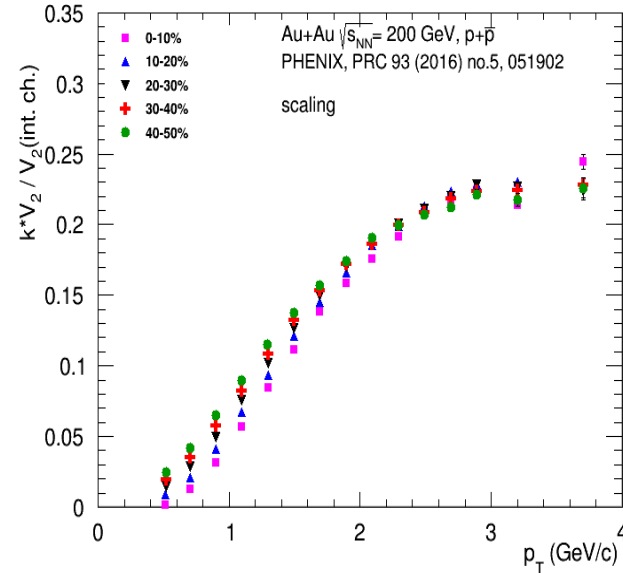
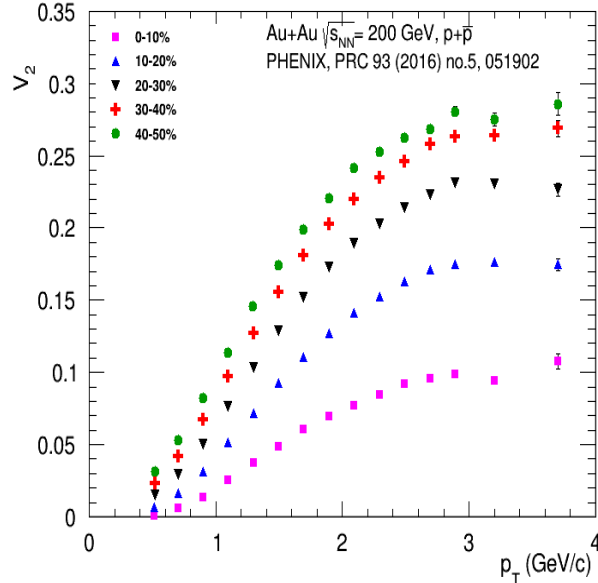
Odd harmonics are system independent.

## Scaling with integral flow of charged hadrons



13  $V_2(PID, p_T, centrality, \sqrt{s_{NN}}) = V_2(h, centrality, \sqrt{s_{NN}}) * V_2(PID, p_T) ???$

## Scaling with integral flow of charged hadrons

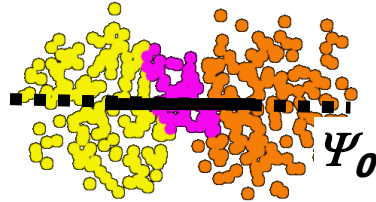


**for protons the strong radial flow “blueshifts” the entire flow signal to higher  $p_T$  :  $p_T \sim p_T^{th} + mc\beta$**

# System size scan at top RHIC energy

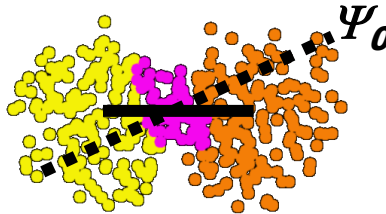
2001-2005

$$\epsilon_{\text{std}} = \frac{\sigma_y^2 - \sigma_x^2}{\sigma_x^2 + \sigma_y^2}$$



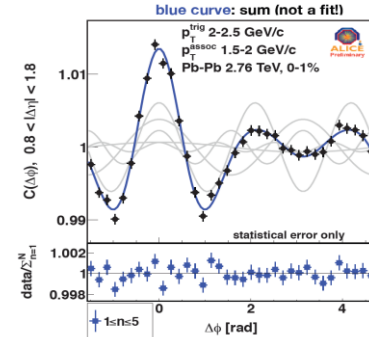
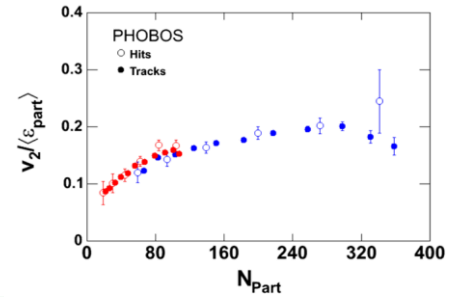
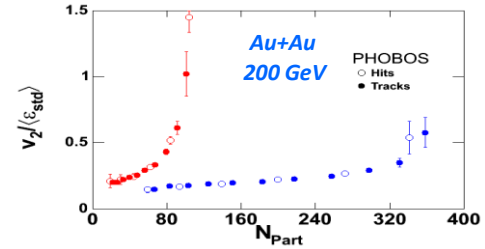
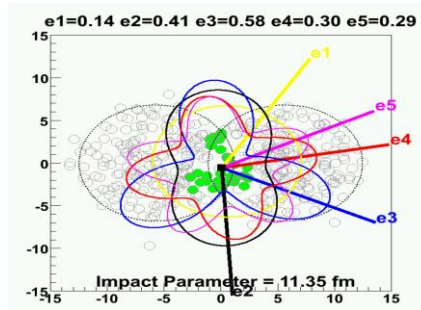
2005-2011

$$\langle \epsilon_{\text{part}} \rangle = \frac{\sqrt{(\sigma_y^2 - \sigma_x^2)^2 + 4\sigma_{xy}^2}}{(\sigma_y^2 + \sigma_x^2)}$$



2011-2012

$$\epsilon_n = \frac{\sqrt{\langle r^n \cos n\phi \rangle + \langle r^n \sin n\phi \rangle}}{\langle r^n \rangle}$$



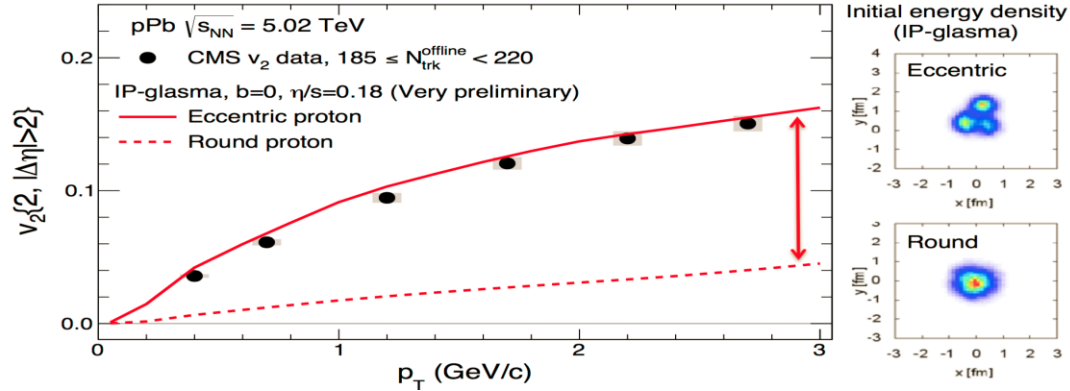
For "lumpy" profile  $\varphi \neq \varphi + \pi$

Odd harmonics  $\neq 0$

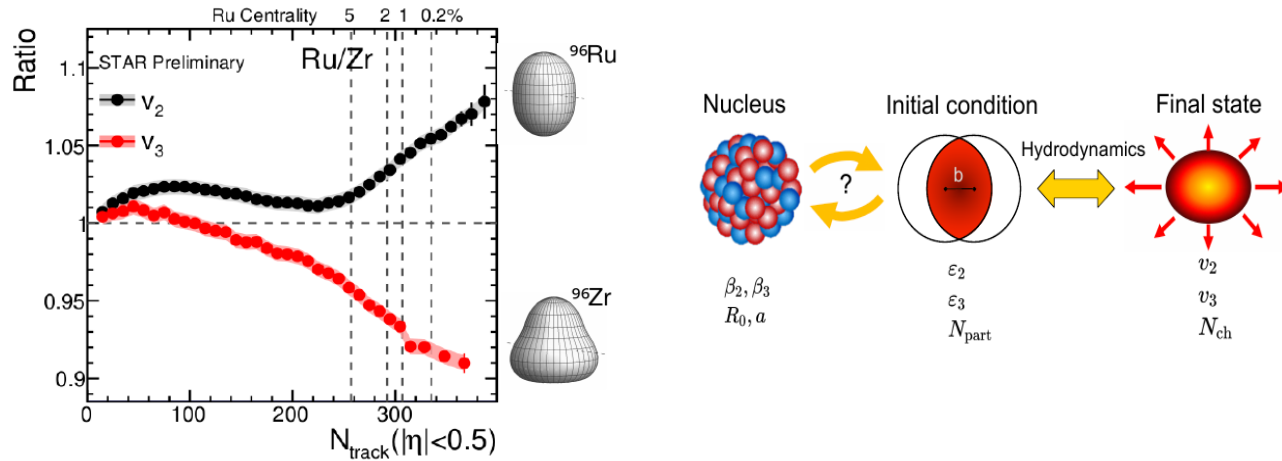
# System size scan at top RHIC energy

Nucleon substructure matters

2011-2020



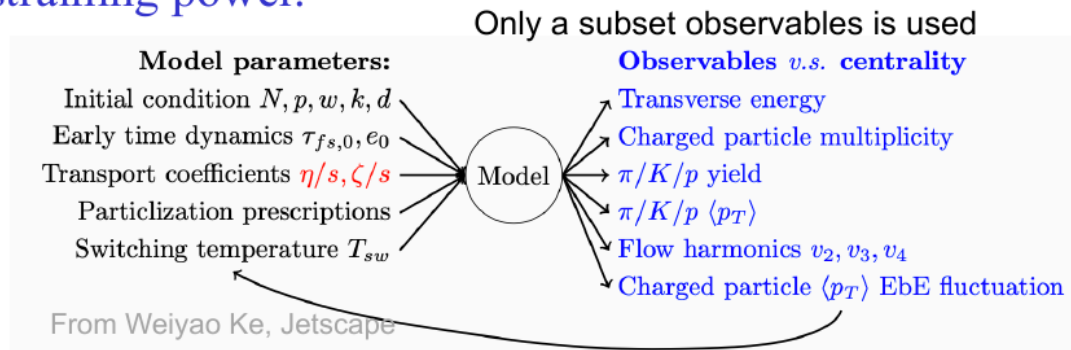
2020-2024



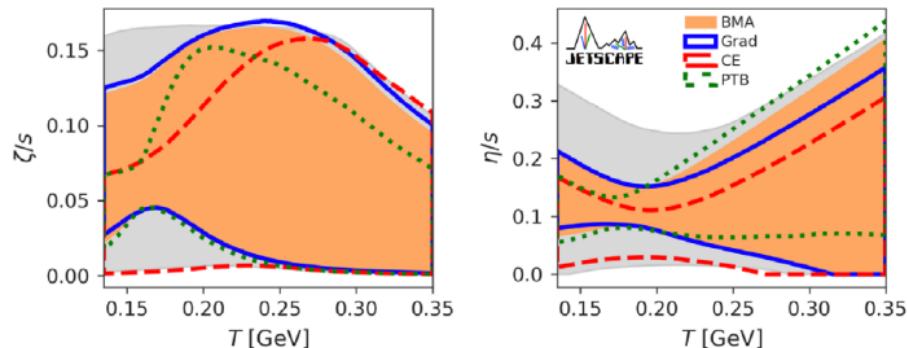


# State-of-the-art modeling of HI collisions

- Data-model comparison via Bayesian inference to optimize constraining power.



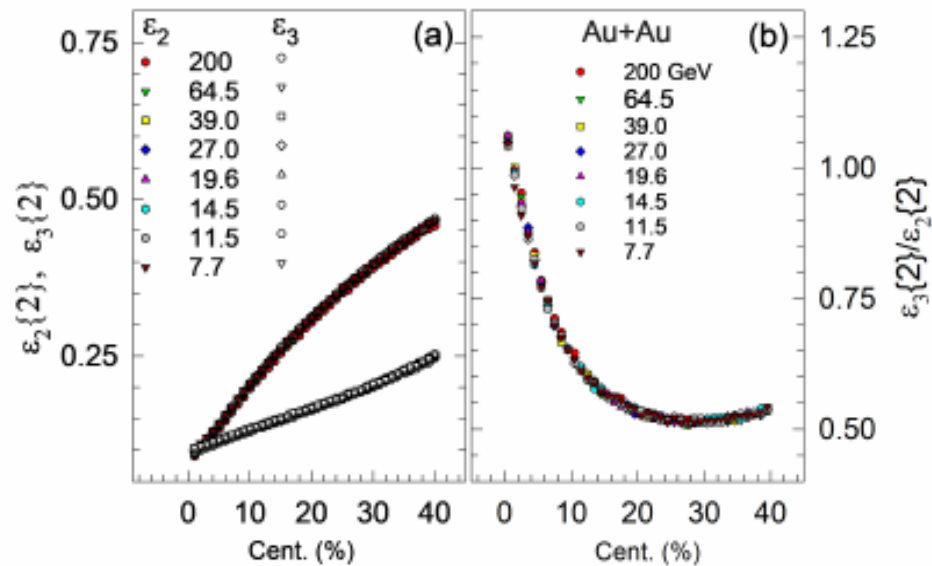
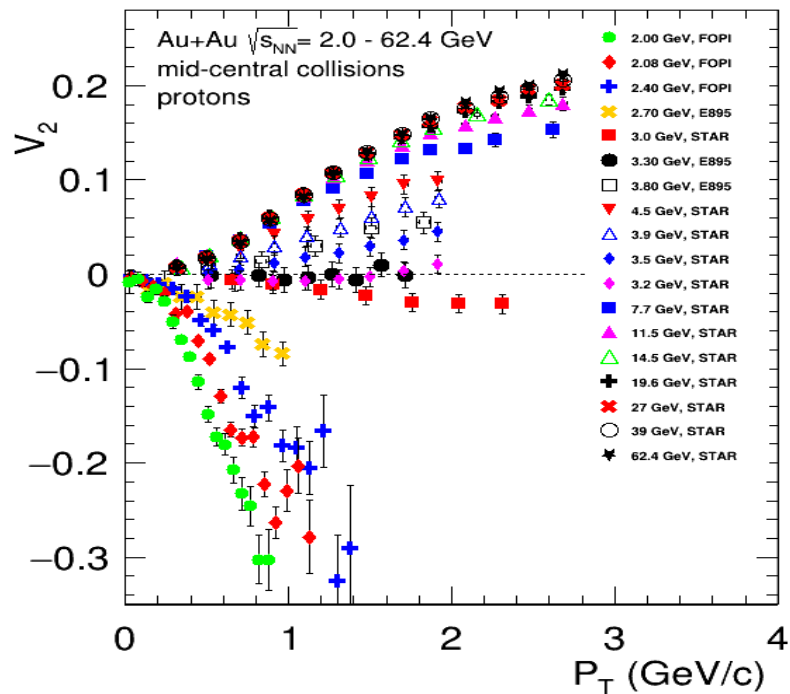
- Detailed temperature dependence of viscosity!



Jetscape PRL.126.242301  
Trjactum PRL.126.202301

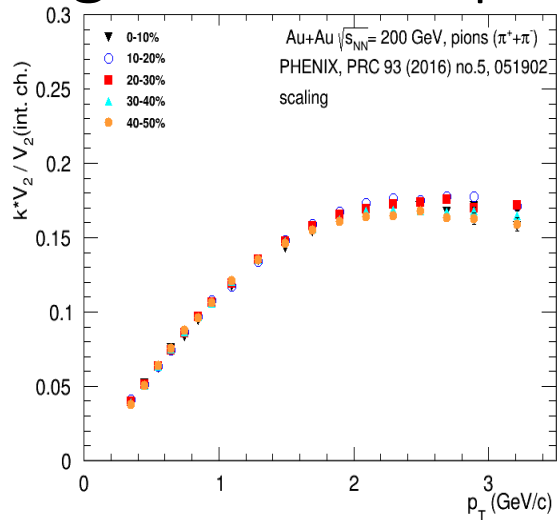
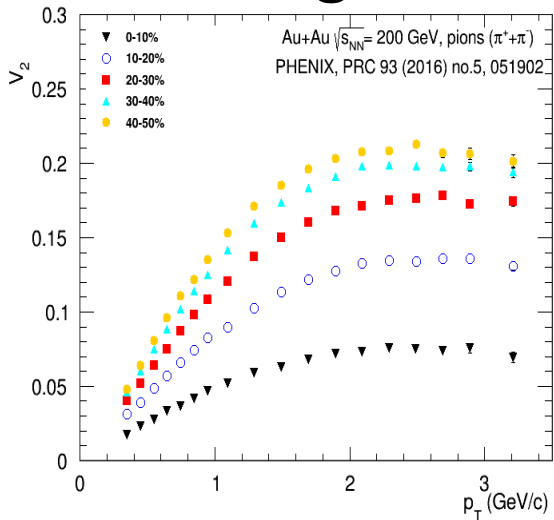
Major uncertainty: initial condition and pre-hydro phase

# Anisotropic flow in Au+Au collisions: Beam Energy dependence



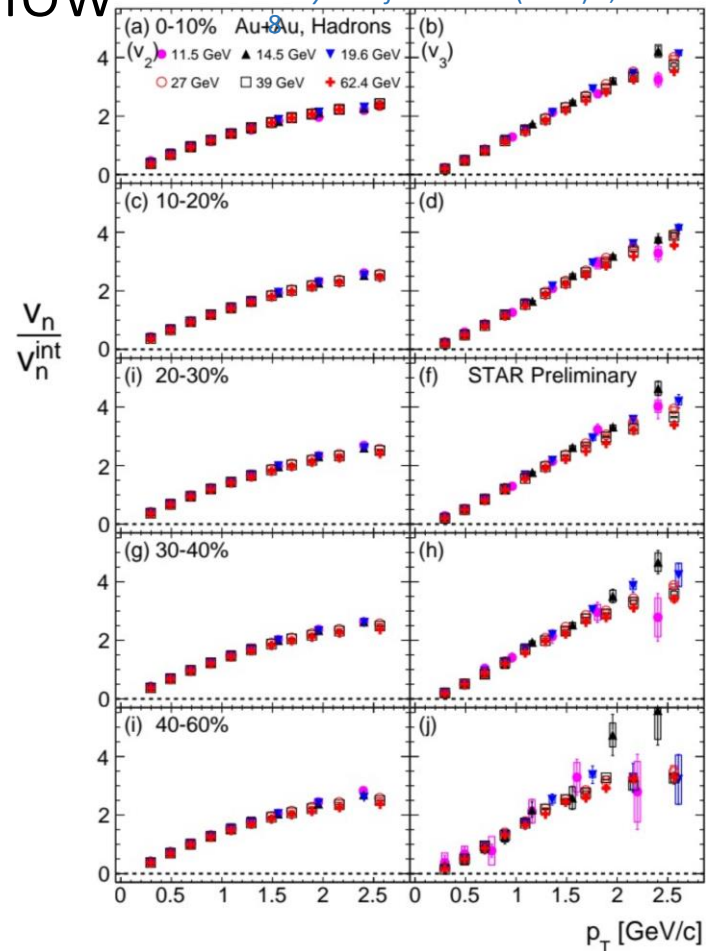
courtesy of R.A. Lacey

# Scaling with integral anisotropic flow

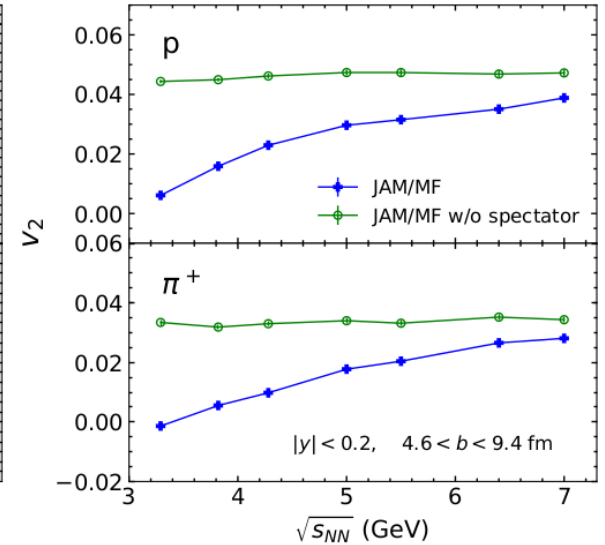
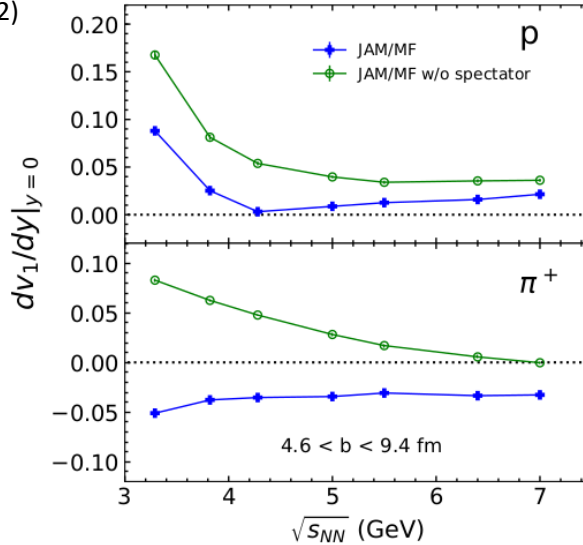
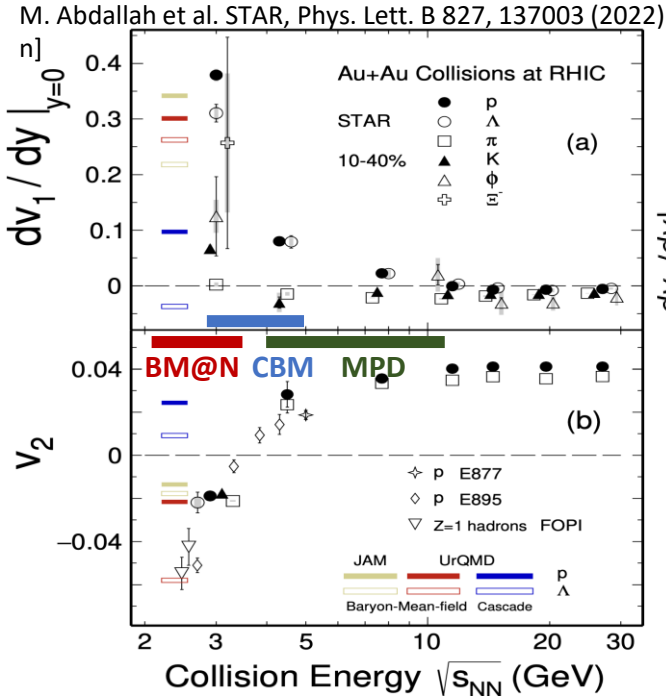


$$v_n(int.) \equiv |v_n^{int}| = |\langle v_n(p_T, y, \text{centrality, PID}) \rangle_{p_T, y}|$$

- Scaling works at top RHIC and BES energy range
- Similar trend for pions, kaons and protons



# Anisotropic flow in heavy-ion collisions at high baryon density



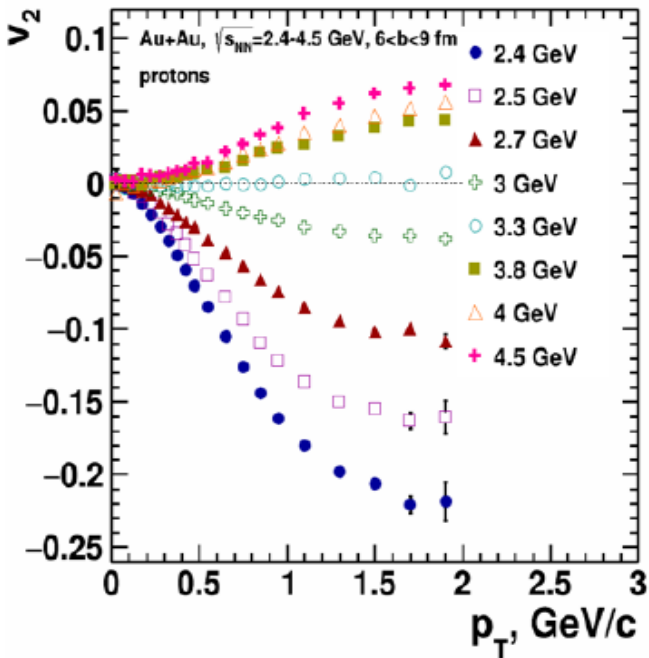
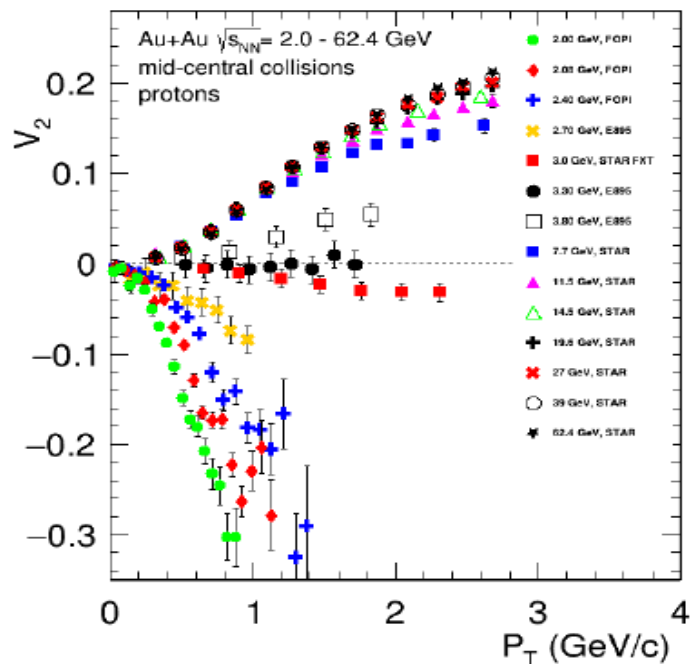
Phys. Rev. C 97, 064913 (2018)

**Anisotropic flow at FAIR/NICA energies is a delicate balance between:**

- I. The ability of pressure developed early in the reaction zone ( $t_{exp} = R/c_s$ ,  $c_s = c\sqrt{dp/d\varepsilon}$ ) and
- II. The passage time for removal of the shadowing by spectators ( $t_{pass} = 2R/\gamma_{CM}\beta_{CM}$ )

# Anisotropic flow in Au+Au collisions at Nuclotron-NICA energies

Particles 6 (2023) 2, 622-637

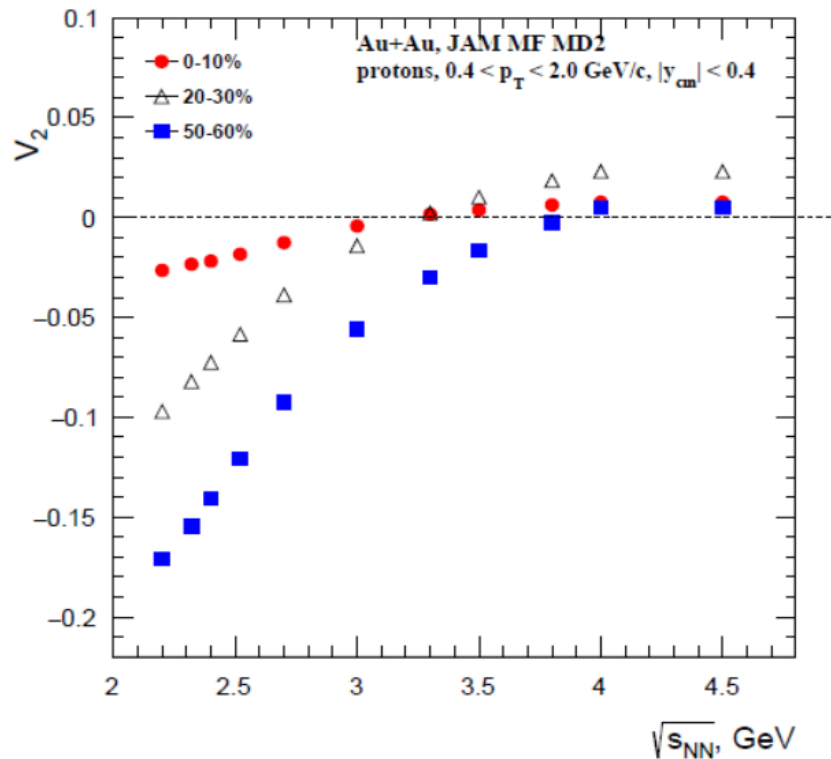


**Anisotropic flow at FAIR/NICA energies is a delicate balance between:**

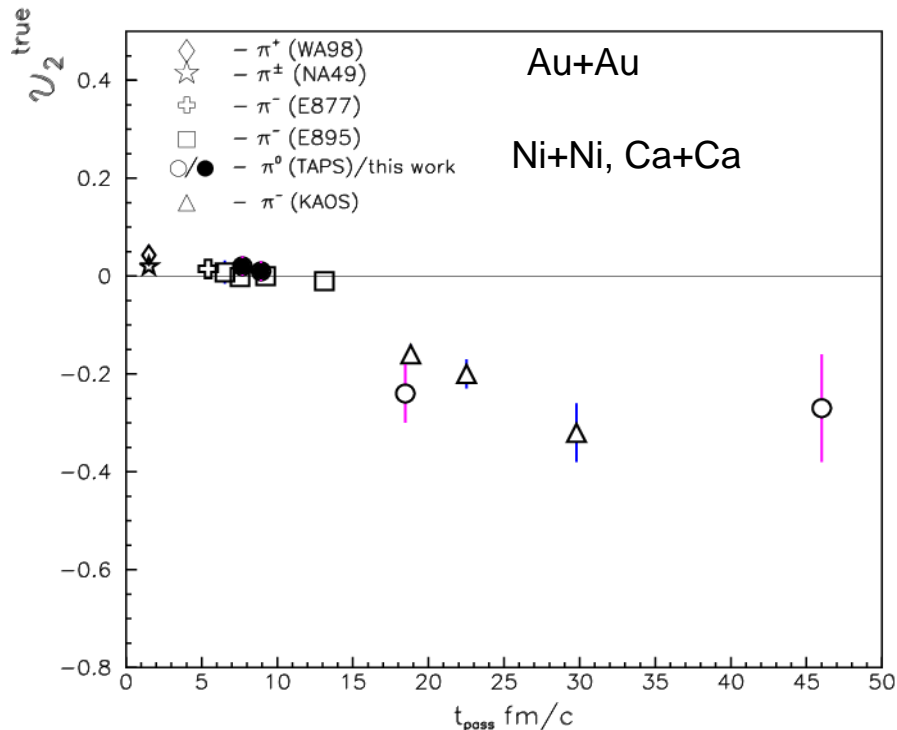
- I. The ability of pressure developed early in the reaction zone ( $t_{exp} = R/c_s$ ,  $c_s = c\sqrt{dp/d\varepsilon}$ ) and
- II. The passage time for removal of the shadowing by spectators ( $t_{pass} = 2R/\gamma_{CM}\beta_{CM}$ )

# Elliptic flow: transition from out-of-plane to in-plane: geometry

P. Parfenov, Particles 5 (2022) 4, 561-579

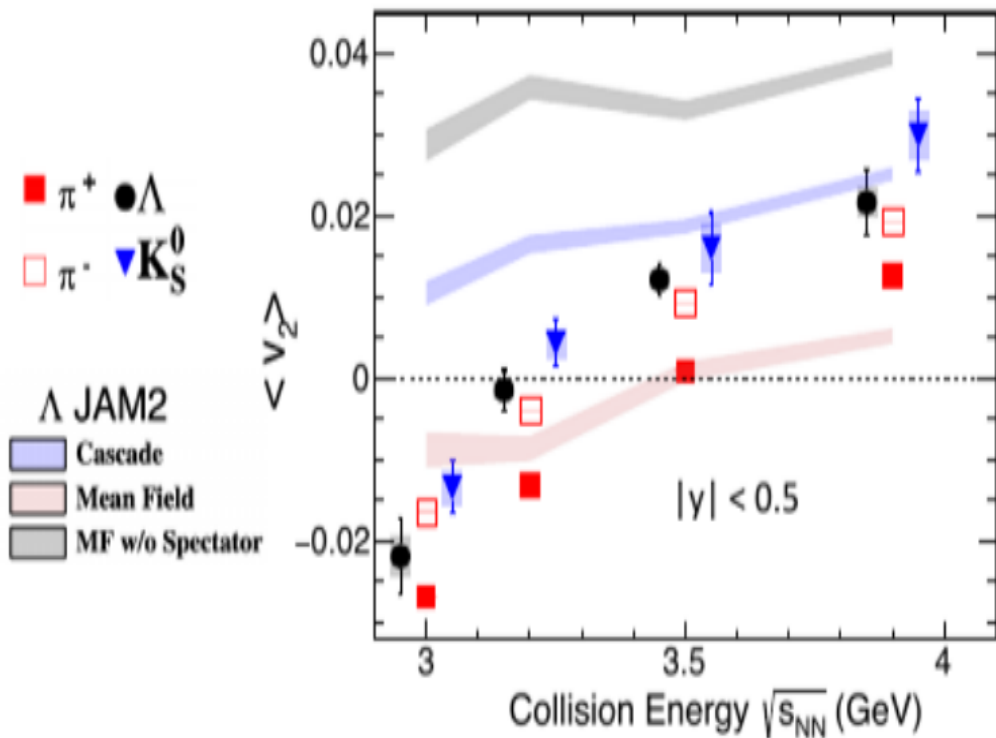


A.T Czech.J.Phys. 50S4 (2000) 139-166

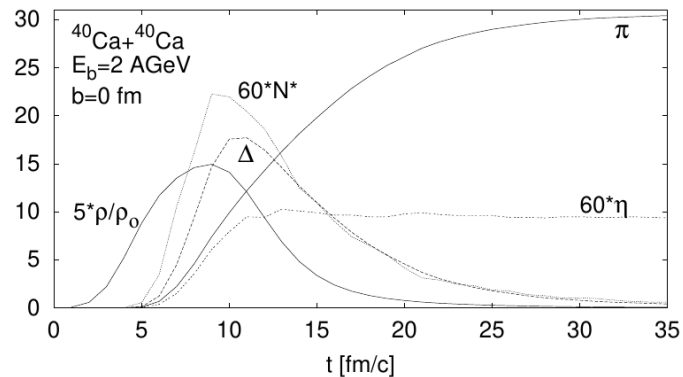
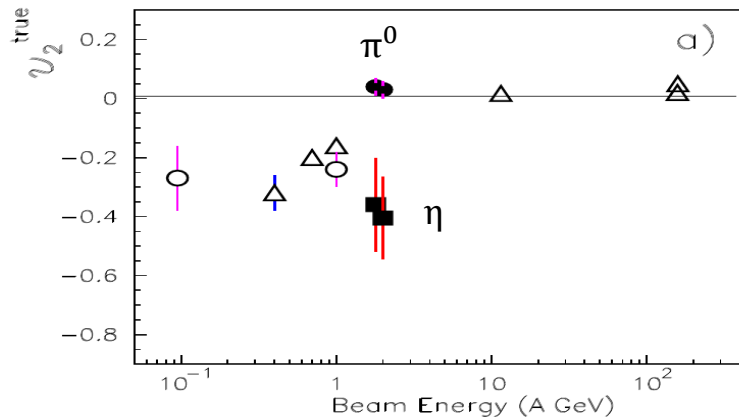


# Elliptic flow: transition from out-of-plane to in-plane: PID

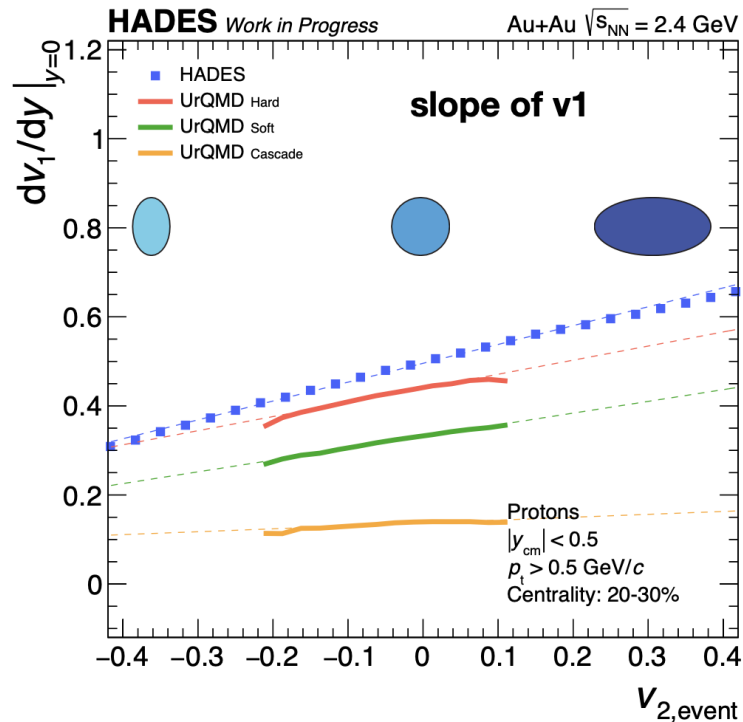
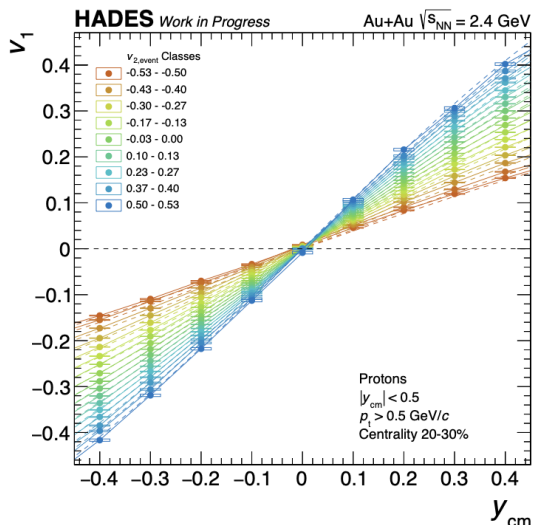
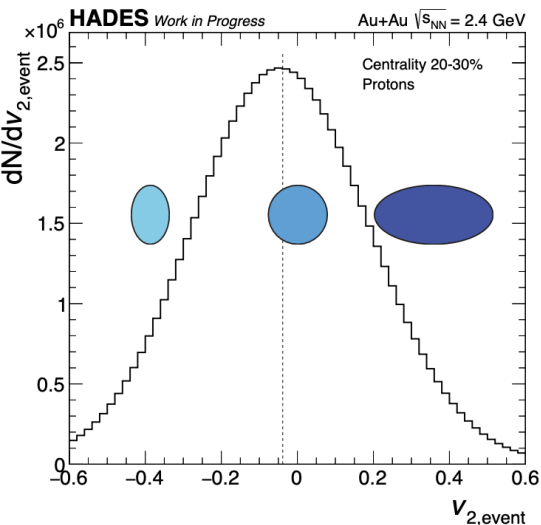
Li-Ke Liu (CCNU), STAR Collaboration, CPOD 2024



A.T Czech.J.Phys. 50S4 (2000) 139-166

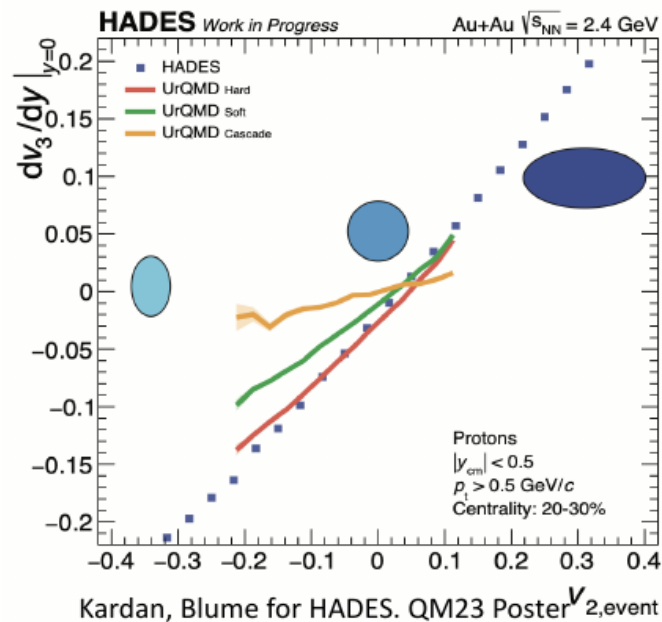
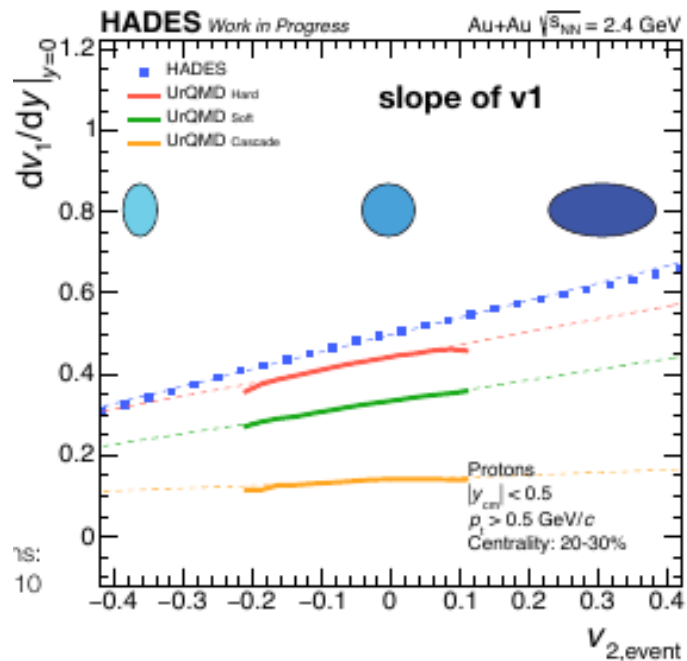


# Event-wise flow correlations



- Events can be characterized based on the event-wise magnitude of the elliptic flow  $v_{2,event}$
- UrQMD can not describe  $dv_1/dy|_{y=0}$  of protons as a function of  $v_{2,event}$
- Strong sensitivity to the EOS

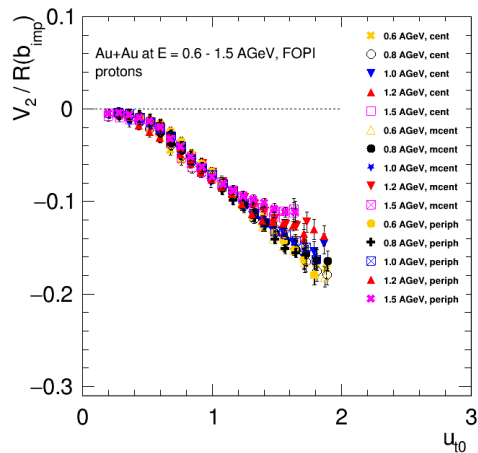
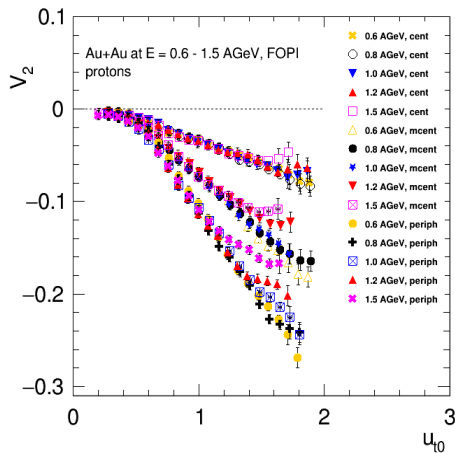
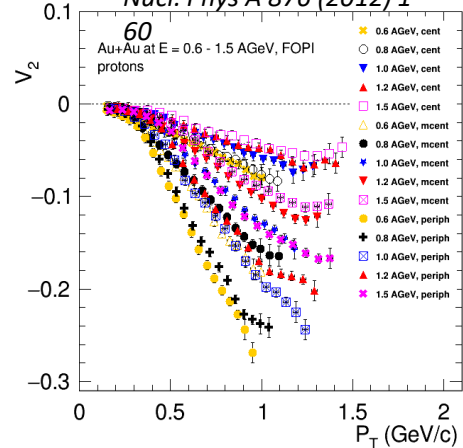




**New HADES results on flow correlations**

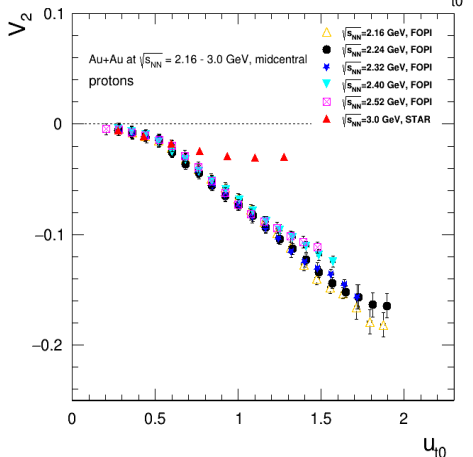
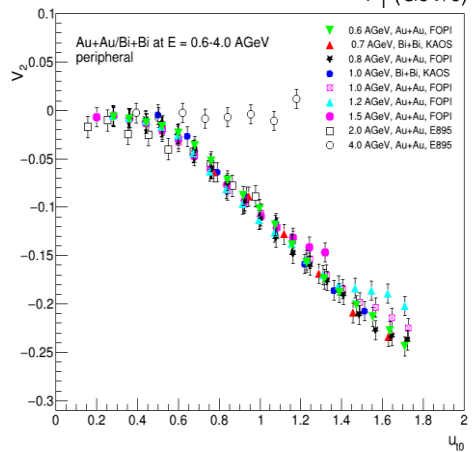
# Scaling relations at SIS – scaling with passage time

Nucl. Phys A 876 (2012) 1-



$$u_{t0} = \frac{p_T}{m_0 \beta_{CMYCM}} \equiv \frac{p_T t_{pass}}{2Rm_0}$$

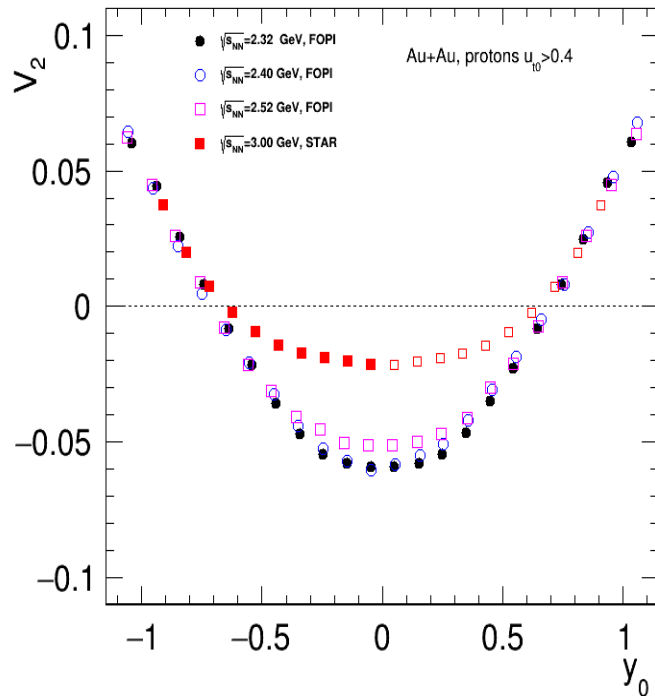
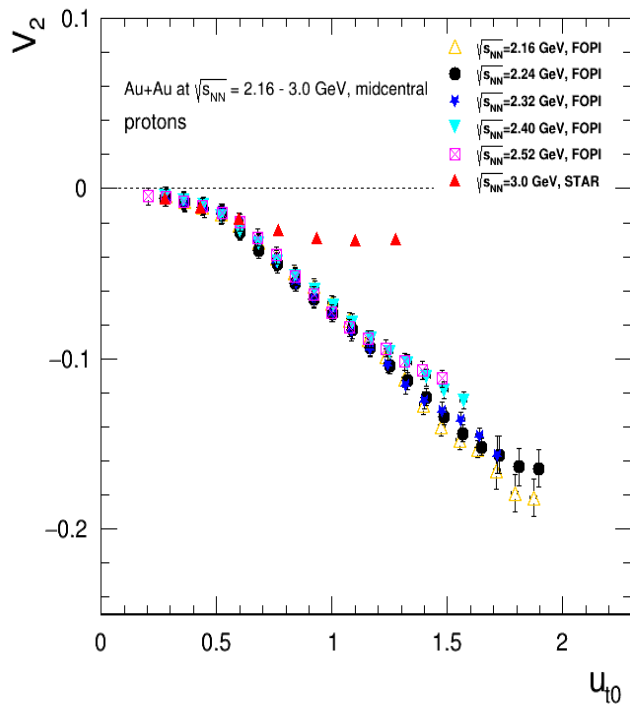
$$t_{pass} = \frac{2R}{\beta_{CMYCM}}$$



- The rather good scaling observed suggests that  $c_s$  does not change significantly over beam energy range  $E_{kin} = 0.4 - 2$  AGeV ( $\sqrt{s_{NN}} = 2 - 2.7$  GeV)
- Scaling breaks at  $E_{kin} = 2.9$  AGeV ( $\sqrt{s_{NN}} = 3$  GeV)

# $u_{t0}$ scaling: FOPI/STAR data

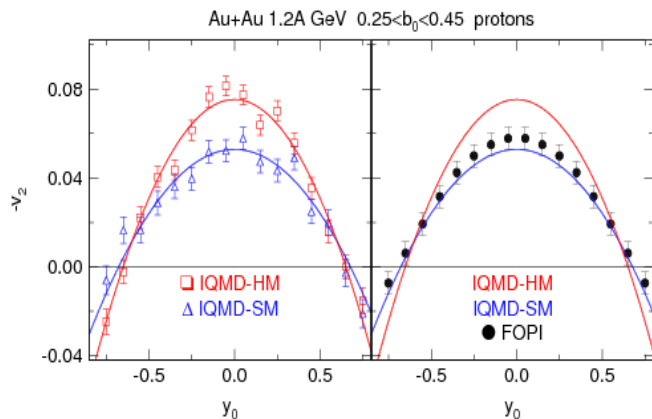
**STAR published results for protons : Scaling breaks at  $\sqrt{s_{NN}}=3\text{GeV}$  – but holds at forward rapidity?**



# Rapidity dependence of $v_2$ and EOS

**HM** – stiff momentum dependent with  $K=376$  MeV

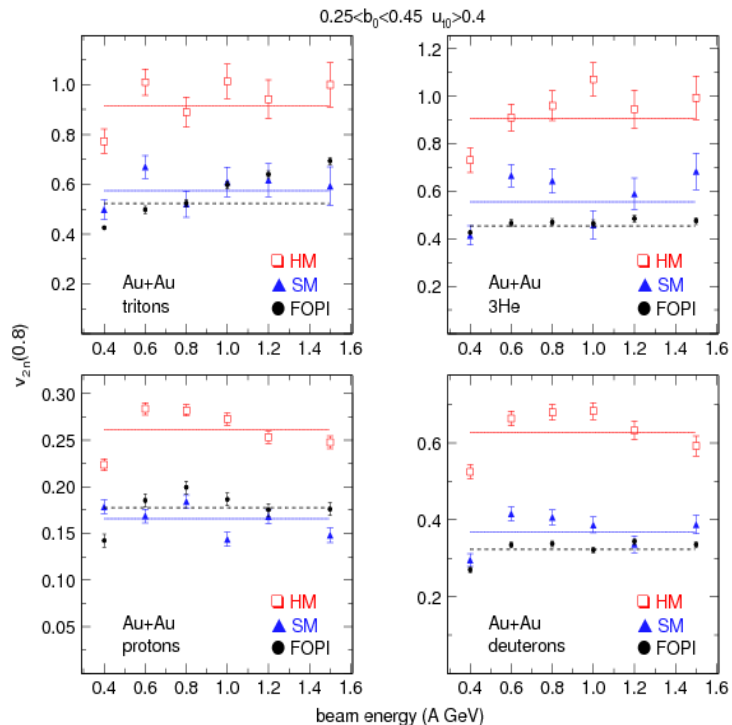
**SM** – soft momentum dependent with  $K=200$  MeV



$$V_{2n} = |V_{20}| + |V_{22}|$$

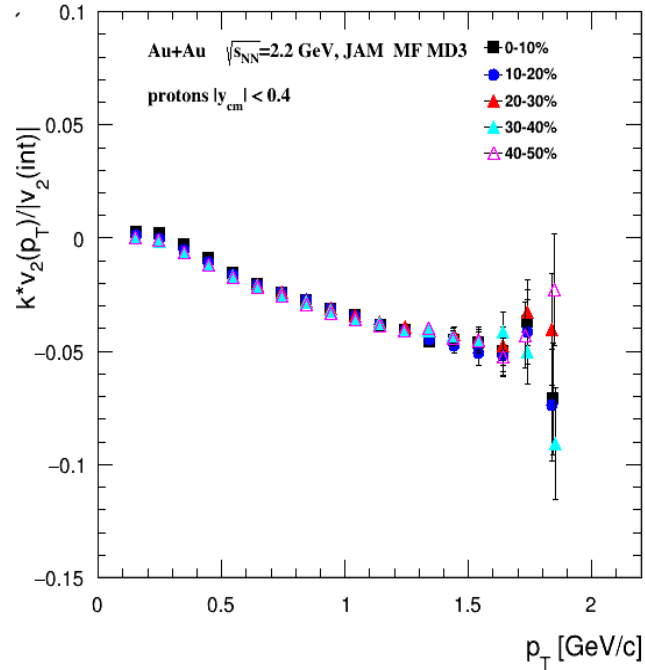
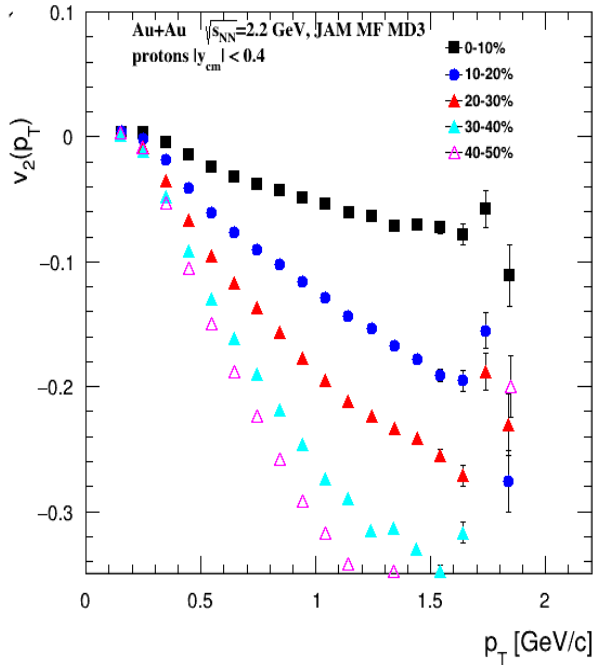
$$\text{Fit: } V_2(y_0) = V_{20} + V_{22} \cdot Y_0^2$$

FOPI data : Nucl. Phys. A 876 (2012) 1  
IQMD : Nucl Phys. A 945 (2016)



**Large rapidity coverage is important for flow measurements: MPD forward upgrade**

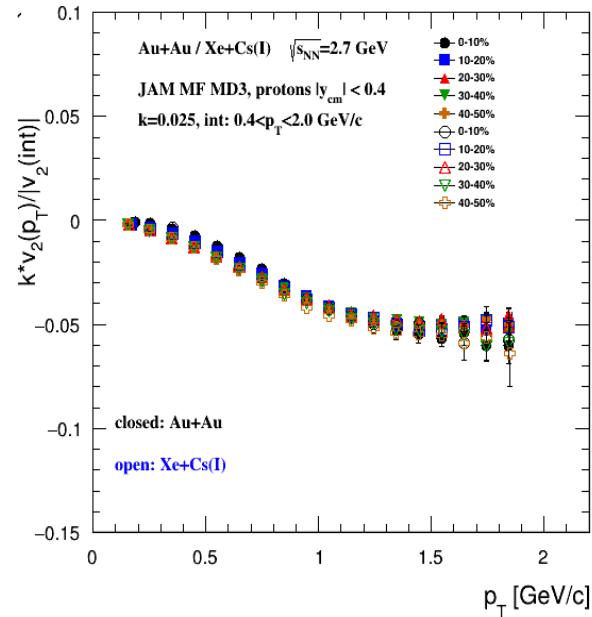
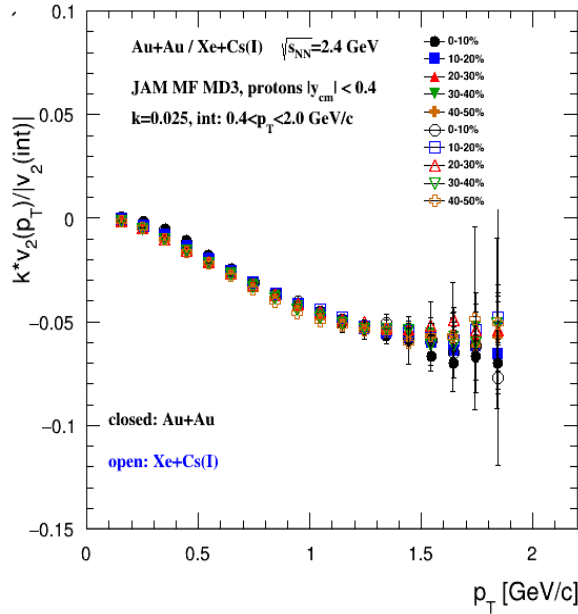
# Scaling with integral flow of charged hadrons. Will it work at $\sqrt{s_{NN}}=2.2$ GeV? (JAM mean field MD3)



In all plots  $k=0.025$  and  $v_2(int)$  for  $0.4 < p_T < 2.0$  GeV/c

29  $V_2(PID, p_T, centrality, \sqrt{s_{NN}}) = V_2(h, centrality, \sqrt{s_{NN}}) * V_2(PID, p_T) ???$

# Scaling with integral flow of charged hadrons. Will it work at $\sqrt{s_{NN}}=2.4$ GeV for different colliding systems? (JAM mean field MD3)

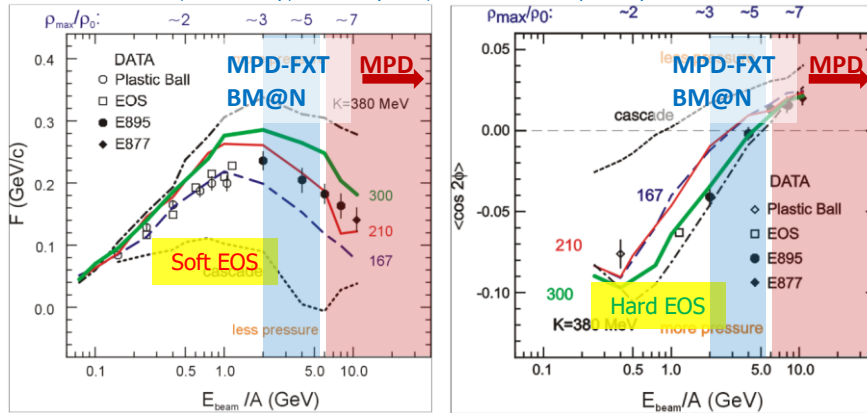


**Scaling works for Au+Au and Xe+Cs(I) – general feature of the flow?**

30  $V_2(PID, p_T, centrality, \sqrt{s_{NN}}) = V_2(h, centrality, \sqrt{s_{NN}}) * V_2(PID, p_T) ???$

# Sensitivity of the collective flow to the EOS

P. Danielewicz, R. Lacey, W.G. Lynch, Science 298 (2002) 1592



$$F = \left. \frac{d\langle p_x/A \rangle}{d(y/y_{cm})} \right|_{y/y_{cm}=1}$$

$$v_2 \equiv \langle \cos(2(\phi - \Psi_{RP})) \rangle$$

Mean field usually can be defined using Skyrme potential with:

$$U(n_B) = A \left( \frac{n_B}{n_0} \right) + B \left( \frac{n_B}{n_0} \right)^\tau$$

Discrepancy in the interpretation:

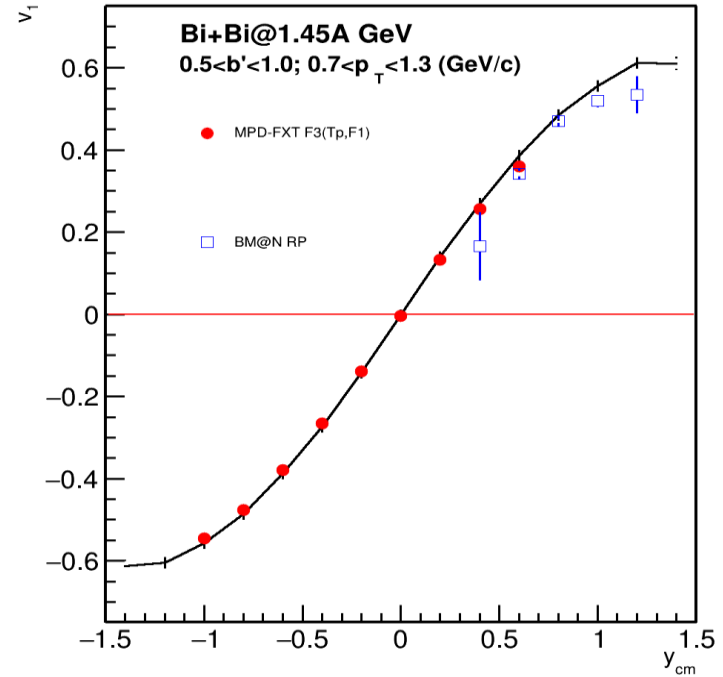
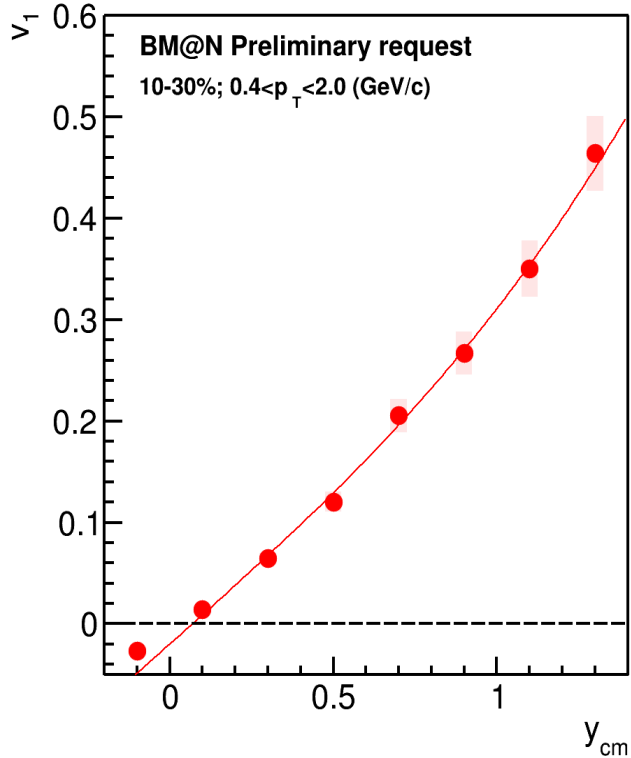
- $v_1$  suggests soft EoS
- $v_2$  suggests hard EoS

**New measurements using new data and modern analysis techniques will address this discrepancy**

**More detailed model study should be done to address  $n_B$ -dependence of incompressibility  $K_0$**

**Additional measurements are essential to clarify the previous measurements**

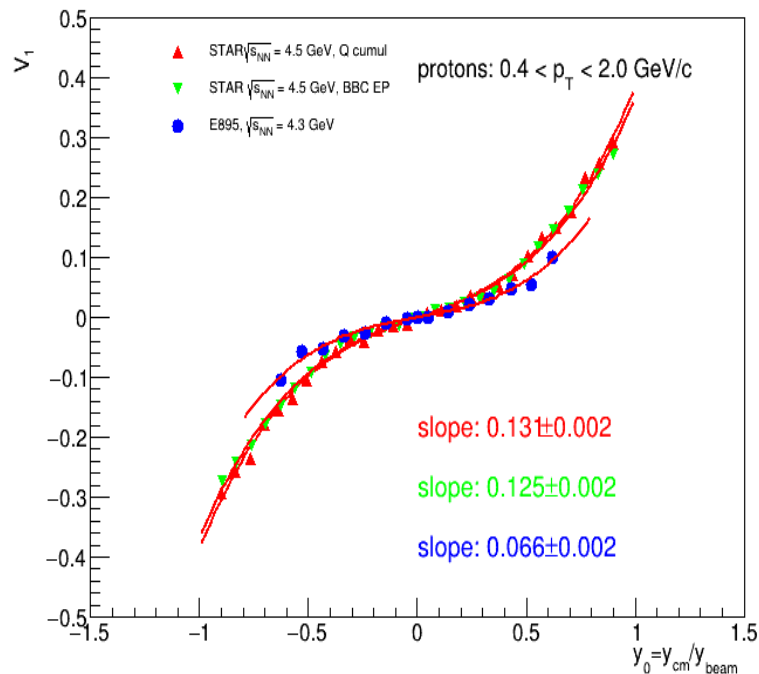
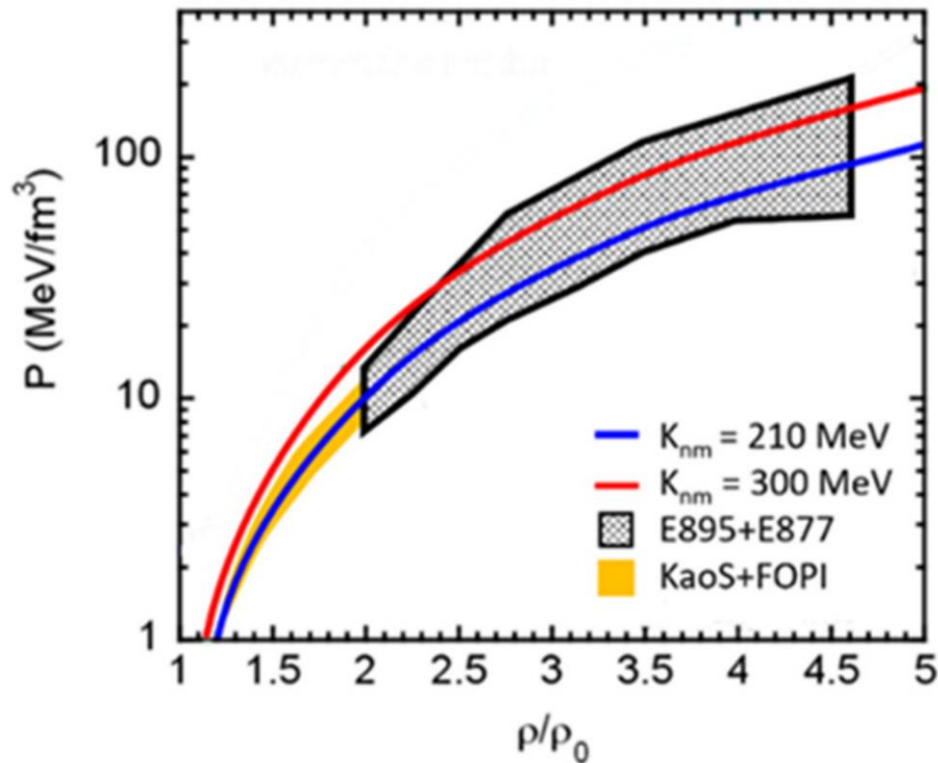
# Directed flow of protons: BM@N – MD FXT



Please see Mikhail Mamaev talk at the workshop



# Directed flow / Elliptic flow of protons: EOS



# Summary and outlook

- Measurements of anisotropic flow, flow fluctuations, correlations between flow of different harmonics are sensitive to many details of the initial conditions and the system evolution. It may provides access to the transport properties of the medium: EOS, sound speed ( $c_s$ ), viscosity, etc.
- $v_n$  at energies 2.5-11 GeV (SIS, STAR BESII, NICA, FAIR) shows strong energy dependence: possible transition between hadronic and partonic matter.
- System size scan is very important in order to understand the effect of spectators on the experimental observables
  
- Feasibility study for anisotropic flow in MPD/MPD FXT/ BM@N:
  
- Programs for flow analysis are available for MPD/BM@N collaborations – first preliminary flow results from BM@N will be published soon.

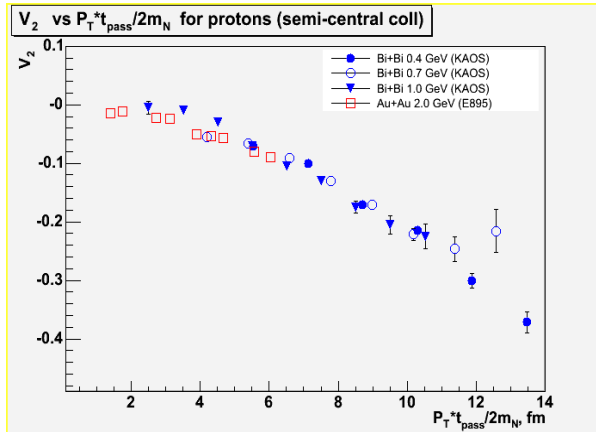
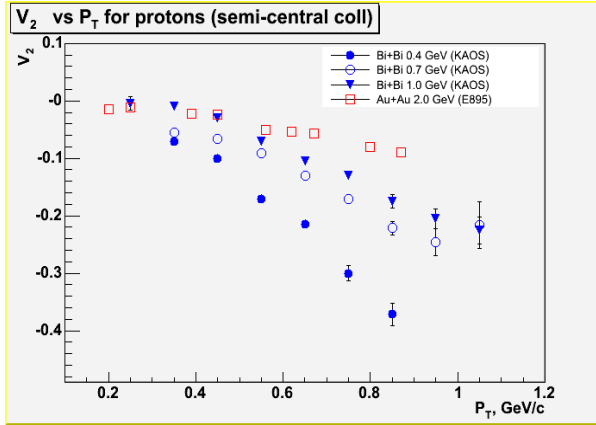
Backup

MPD

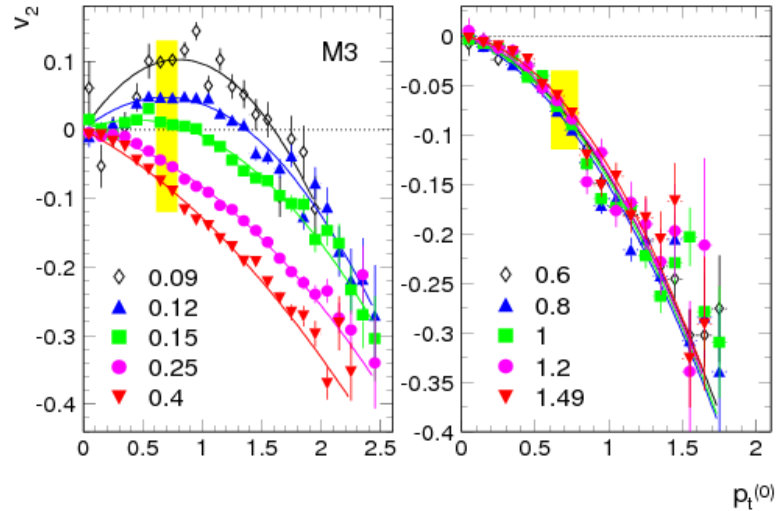
BM@N

# $v_2$ Flow at SIS-AGS: scaling relations

(KAOS – *Z. Phys. A*355 (1996);  
(E895) - *PRL* 83 (1999) 1295



**FOPI:  $v_2$  of protons from  
 $Elab=0.09$  to  $1.49$  GeV**  
*Phys.Lett. B*612 (2005) 173-180



**The rather good scaling observed suggest that  $c_s$  does not change significantly over beam energy range 0.4 – 2.0 AGeV. .**

# Vn of protons in Au+Au collisions at 2.4 GeV - HADES

## Determination of EOS

New level of precision - multi differential  
Additional information from higher orders

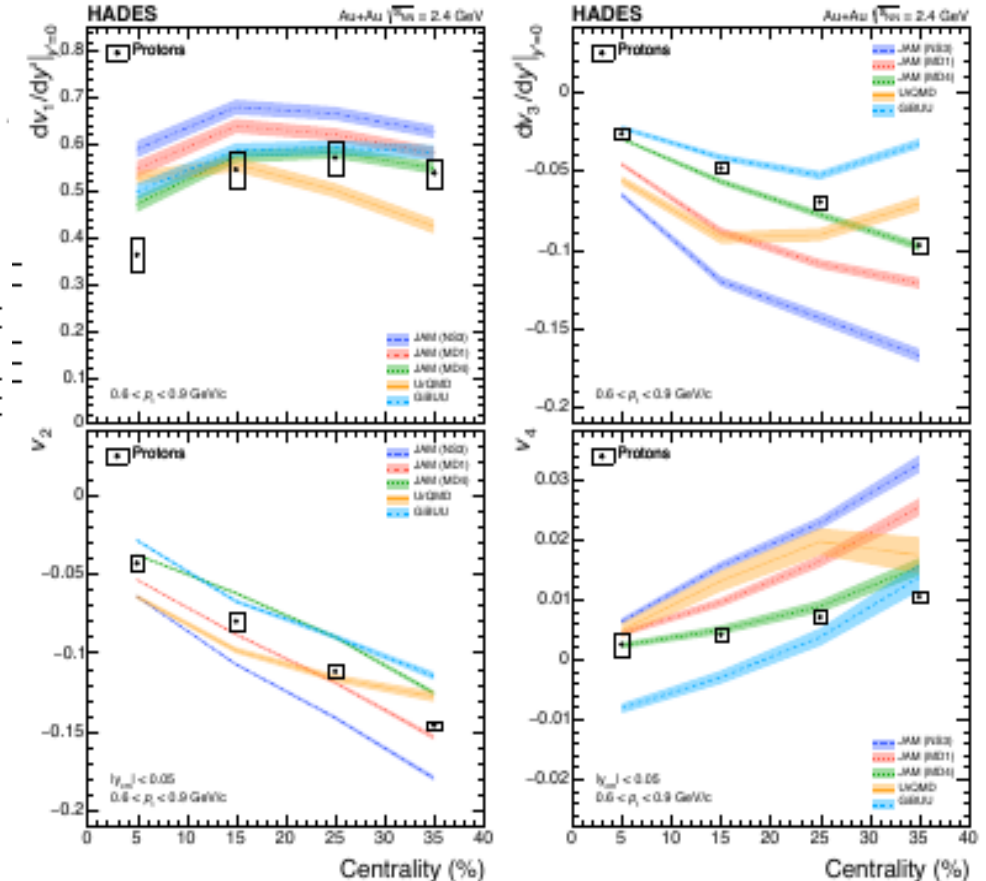
## Models:

JAM 1.9 NS3 (hard EOS, mom.-indep.)  
JAM 1.9 MD1 (hard EOS, mom.-dep.)  
JAM 1.9 MD4 (soft EOS, mom.dep.)  
UrQMD 3.4 (hard EOS, mom.-indep.)  
GIBUU Skyrme 12 (soft EOS)

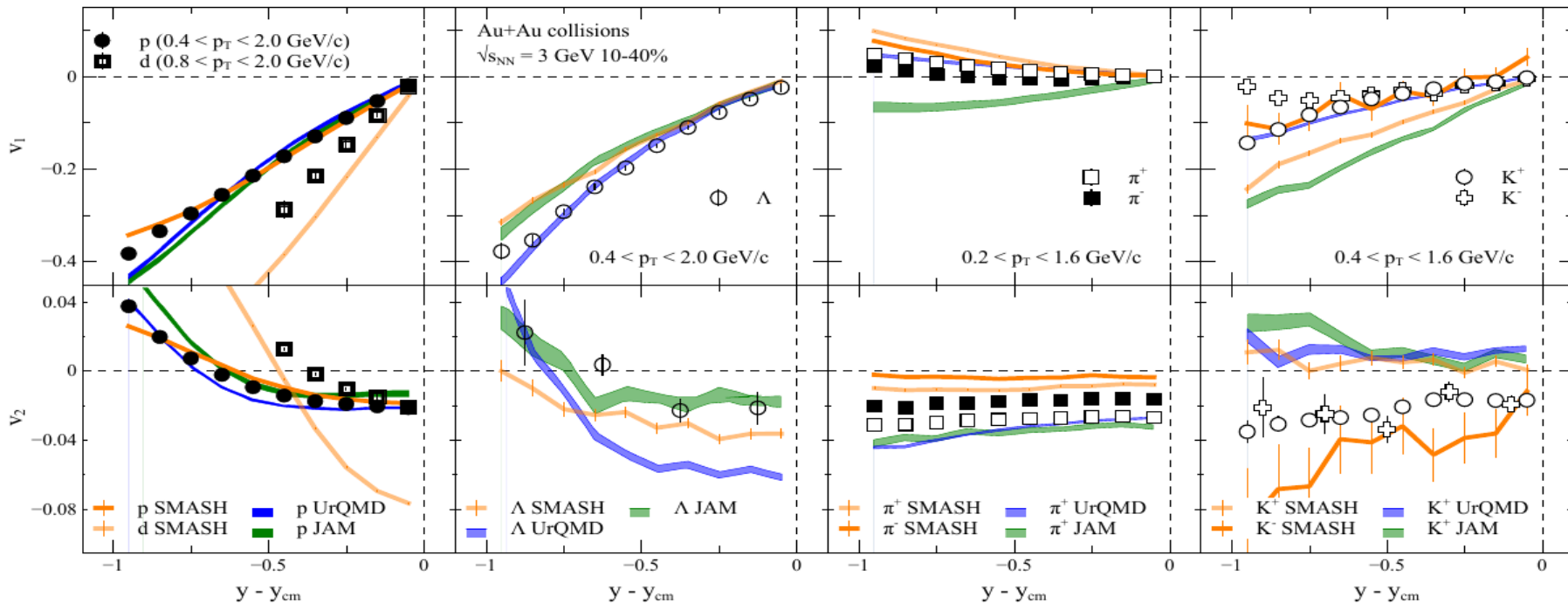
Model	EOS	$K$ (MeV)	$m^*/m$	mom.-dep.
JAM 1.9 NS3	NS3	380	0.83	no
JAM 1.9 MD1	MD1	380	0.65	yes
JAM 1.9 MD4	MD4	210	0.83	yes
UrQMD 3.4	Hard	380	no	no
GIBUU Skyrme 12	Skyrm+ J2	240	0.75	no

## Conclusions

Overall trend reasonably described, but no model works everywhere



# Describing proton flow is not enough



Strange baryons are not well described  
 — the results may depend on:

- nucleon-hyperon and hyperon-hyperon interactions
- in-medium modifications of interactions

Pions and kaons NOT described!  
 Not very surprising: UrQMD, JAM, and SMASH  
 don't have mean-fields for mesons