

# KK reduction of Horndeski and the speed of gravity.

based on papers with M.Valencia-Villegas and A.Shtennikova

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I) Horndeski theory and generalizations

II) Kaluza-Klein reduction

III) Profit

IV) Final remarks

## Horndeski theory

$$S = \int d^4x \sqrt{-g} (\mathcal{L}_2 + \mathcal{L}_3 + \mathcal{L}_4 + \mathcal{L}_5),$$

$$\mathcal{L}_2 = F(\pi, X),$$

$$\mathcal{L}_3 = K(\pi, X) \square \pi,$$

$$\mathcal{L}_4 = -G_4(\pi, X) R + 2G_{4X}(\pi, X) \left[ (\square \pi)^2 - \pi_{;\mu\nu} \pi^{;\mu\nu} \right],$$

$$\mathcal{L}_5 = G_5(\pi, X) G^{\mu\nu} \pi_{;\mu\nu} + \frac{1}{3} G_{5X} \left[ (\square \pi)^3 - 3 \square \pi \pi_{;\mu\nu} \pi^{;\mu\nu} + 2 \pi_{;\mu\nu} \pi^{;\mu\rho} \pi_{;\rho}{}^\nu \right],$$

where  $\pi$  is the Galileon field,  $X = g^{\mu\nu} \pi_{;\mu} \pi_{;\nu}$ ,  $\pi_{;\mu} = \partial_\mu \pi$ ,  $\pi_{;\mu\nu} = \nabla_\nu \nabla_\mu \pi$ ,  
 $\square \pi = g^{\mu\nu} \nabla_\nu \nabla_\mu \pi$ ,  $G_{4X} = \partial G_4 / \partial X$

## Motivation

- 1 Avoid the quantum gravity. Being able to construct everywhere-regular weak-gravity solutions.
- 2 Have sufficiently much freedom to modify gravity and scalar dynamics in different ways
- 3 Theory of a very general form under several assumptions

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general covariance, locality and 1 additional degree of freedom.

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(1) requires NEC violation.



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Galileons  $\rightarrow$  Covariant Galileons  $\rightarrow$  Generalized Galileons

(Nicolis et al 0811.2197)

(Deffayet et al 0901.1314)

(Deffayet et al 1103.3260)

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Horndeski

(Horndeski 1974)

$$\mathcal{L} = c_1\phi + c_2X - c_3X\Box\phi + c_4X [(\Box\phi)^2 - \partial_\mu\partial_\nu\phi\partial^\mu\partial^\nu\phi] \\ - \frac{c_5}{3}X [(\Box\phi)^3 - 3\Box\phi\partial_\mu\partial_\nu\phi\partial^\mu\partial^\nu\phi + 2\partial_\mu\partial_\nu\phi\partial^\nu\partial^\lambda\phi\partial_\lambda\partial^\mu\phi]$$

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$$\mathcal{L} = c_1\phi + c_2X - c_3X\Box\phi + \frac{c_4}{2}X^2R + c_4X [(\Box\phi)^2 - \phi^{\mu\nu}\phi_{\mu\nu}] \\ + c_5X^2G^{\mu\nu}\phi_{\mu\nu} - \frac{c_5}{3}X [(\Box\phi)^3 - 3\Box\phi\phi^{\mu\nu}\phi_{\mu\nu} + 2\phi_{\mu\nu}\phi^{\nu\lambda}\phi_{\lambda}^{\mu}]$$

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(Horndeski 1974)

$$\mathcal{L} = G_2(\phi, X) - G_3(\phi, X)\square\phi + G_4(\phi, X)R + G_{4X} [(\square\phi)^2 - \phi^{\mu\nu}\phi_{\mu\nu}] \\ + G_5(\phi, X)G^{\mu\nu}\phi_{\mu\nu} - \frac{G_{5X}}{6} [(\square\phi)^3 - 3\square\phi\phi^{\mu\nu}\phi_{\mu\nu} + 2\phi_{\mu\nu}\phi^{\nu\lambda}\phi_{\lambda}^{\mu}]$$

## Horndeski theory

$$S = \int d^4x \sqrt{-g} (\mathcal{L}_2 + \mathcal{L}_3 + \mathcal{L}_4 + \mathcal{L}_5),$$

$$\mathcal{L}_2 = F(\pi, X),$$

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## General lagrangian with 2 tensor and 1 scalar DOF

General relativity, 1-field inflations, non-minimal coupling

K-essence/k-inflation

kinetic gravity braiding/G-inflation

f(R)-gravity, Gauss-Bonnet term, f(G-B)

## No Ostrogradski ghost

second order equations of motion in Horndeski, despite second derivatives is the Lagrangian

## Can break NEC without linear instabilities

$$\pi = \pi_0 + \chi, \quad g_{\mu\nu} = \tilde{g}_{\mu\nu} + h_{\mu\nu}$$

$$L_{\zeta}^{(2)} = \frac{1}{2} U \dot{\zeta}^2 - \frac{1}{2} V (\partial_i \zeta)^2 - \frac{1}{2} W \zeta^2$$

$$U \omega^2 = V \mathbf{p}^2 + W,$$

- stability requirement:  $U > 0$ ,  $V > 0$ ,  $W \geq 0$ .

## beyond Horndeski

$$S = \int d^4x \sqrt{-g} (\mathcal{L}_2 + \mathcal{L}_3 + \mathcal{L}_4 + \mathcal{L}_5 + \mathcal{L}_{\mathcal{BH}}),$$

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$$\mathcal{L}_4 = -G_4(\pi, X) R + 2G_{4X}(\pi, X) \left[ (\square \pi)^2 - \pi_{;\mu\nu} \pi^{;\mu\nu} \right],$$

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$$\begin{aligned} \mathcal{L}_{\mathcal{BH}} = & F_4(\pi, X) \epsilon^{\mu\nu\rho}{}_\sigma \epsilon^{\mu'\nu'\rho'\sigma} \pi_{;\mu} \pi_{;\mu'} \pi_{;\nu\nu'} \pi_{;\rho\rho'} + \\ & + F_5(\pi, X) \epsilon^{\mu\nu\rho\sigma} \epsilon^{\mu'\nu'\rho'\sigma'} \pi_{;\mu} \pi_{;\mu'} \pi_{;\nu\nu'} \pi_{;\rho\rho'} \pi_{;\sigma\sigma'} \end{aligned}$$

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$$F_4 G_{5X} X = -3F_5 \left[ G_4 - 2XG_{4X} + \frac{1}{2} G_{5\pi} X \right],$$

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$$H \rightarrow \text{BH} \quad g_{\mu\nu} \rightarrow g_{\mu\nu} + \Gamma(\pi, X) \partial_\mu \pi \partial_\nu \pi.$$



## DHOST theory

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$$A_2 = -A_1,$$

$$A_4 = \frac{1}{8(F_2 - XA_1)^2} \left[ -16XA_1^3 + 4(3F_2 + 16XF_{2X})A_1^2 \right. \\ \left. - (16X^2F_{2X} - 12XF_2)A_3A_1 - X^2F_2A_3^2 \right. \\ \left. - 16F_{2X}(3F_2 + 4XF_{2X})A_1 + 8F_2(XF_{2X} - F_2)A_3 + 48F_2F_{2X}^2 \right],$$

$$A_5 = \frac{(4F_{2X} - 2A_1 + XA_3)(-2A_1^2 - 3XA_1A_3 + 4F_{2X}A_1 + 4F_2A_3)}{8(F_2 - XA_1)^2}.$$

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$$L_6^{(3)} = (\pi_{\rho\sigma})^2(\pi_{\mu\nu}\pi^\mu\pi^\nu), \quad L_7^{(3)} = \pi^{\mu\nu}\pi_{\nu\rho}\pi^{\rho\sigma}\pi_\mu\pi_\sigma,$$

$$L_8^{(3)} = (\pi^{\mu\nu}\pi_\mu)^2(\pi^{\rho\sigma}\pi_\rho\pi_\sigma), \quad L_9^{(3)} = \square\pi(\pi^{\rho\sigma}\pi_\rho\pi_\sigma)^2,$$

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+ Relations between  $F_3$  and  $B_j$

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+ Relations

$$H \rightarrow \text{DHOST} \quad g_{\mu\nu} \rightarrow \Omega^2(\pi, X) g_{\mu\nu} + \Gamma(\pi, X) \partial_\mu \pi \partial_\nu \pi.$$

# Horndeski theory in 5d

$$S = \int d^5x \sqrt{g} \mathcal{L}_\pi,$$

$$\mathcal{L}_\pi = \mathcal{L}_2 + \mathcal{L}_3 + \mathcal{L}_4 + \mathcal{L}_5 + \mathcal{L}_6$$

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$$\mathcal{L}_5 = G_5(\pi, X) G^{MN} \pi_{;MN} + \frac{1}{3} G_{5X}(\pi, X) \left[ (\square \pi)^3 - 3 \square \pi \pi_{;MN} \pi^{;MN} + 2 \pi_{;MN} \pi^{;MP} \pi_{;P}{}^N \right],$$

$$\mathcal{L}_6 = \frac{3}{4} G_6(\pi, X) \left( R^2 - 4R^{AB} R_{AB} + R^{ABCD} R_{ABCD} \right)$$

$$+ 3 G_{6X}(\pi, X) *$$

$$\left( -R \left( (\square \pi)^2 - \pi^{;AB} \pi_{;AB} \right) + 4R^{AB} \left( \square \pi \pi_{;AB} - \pi_{;A}{}^C \pi_{;CB} \right) - 2R^{ABCD} \pi_{;AC} \pi_{;BD} \right)$$

$$+ G_{6XX}(\pi, X) *$$

$$\left( (\square \pi)^4 - 6 \pi^{;AB} \pi_{;AB} (\square \pi)^2 + 8 \square \pi \pi^{;AB} \pi_{;B}{}^C \pi_{;CA} + 3 \left( \pi^{;AB} \pi_{;AB} \right)^2 - 6 \pi^{;AB} \pi_{;B}{}^C \pi_{;C}{}^D \pi_{;DA} \right)$$

## KK reduction

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Let us perform **KK** reduction for **H**, **BH** and **DHOST** theories

## KK compactification of Horndeski theory and generalizations

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\* Metric + scalar $_{\pi}$   $\longrightarrow$  Metric + vector + scalar $_{\pi}$  + scalar $_{\phi}$   
[U(1) gauge]

## KK compactification of Horndeski theory and generalizations

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$$\mathbf{H}(g_{MN} + \pi) \longrightarrow \mathbf{H}(g_{\mu\nu} + \pi) + \dots$$

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$\dots =$  Modified Maxwell theory + dilaton interactions



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$$g_{AB} = \begin{pmatrix} g_{\mu\nu} - \phi^2 A_\mu A_\nu & \phi^2 A_\mu \\ \phi^2 A_\nu & -\phi^2 \end{pmatrix}$$

$$\mathcal{L}_{H\pi}^{5d} \rightarrow \mathcal{L}_H^{KK} = \mathcal{L}_{H\pi} + \mathcal{L}_A + \mathcal{L}_\phi,$$

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$$\mathcal{L}_A = \mathcal{L}_{4A} + \mathcal{L}_{5A} + \mathcal{L}_{6A},$$

$$\mathcal{L}_{K\phi} = \frac{1}{\phi} K \phi^{;\alpha} \pi_{;\alpha},$$

$$\mathcal{L}_{4\phi} = \frac{2}{\phi} G_4 (\square\phi) + \frac{4}{\phi} G_{4X} (\square\pi) \phi^{;\alpha} \pi_{;\alpha},$$

$$\begin{aligned} \mathcal{L}_{5\phi} = & \frac{1}{\phi} G_5 ((\square\phi) (\square\pi) - \phi^{;\alpha\beta} \pi_{;\alpha\beta}) - \frac{1}{2\phi} G_5 R \phi^{;\alpha} \pi_{;\alpha} \\ & + \frac{1}{\phi} G_{5X} \phi^{;\alpha} \pi_{;\alpha} \left( (\square\pi)^2 - \pi_{;\alpha\beta} \pi^{;\alpha\beta} \right), \end{aligned}$$

$$\begin{aligned} \mathcal{L}_{6\phi} = & \frac{6}{\phi} G_6 G^{\alpha\beta} \phi_{;\alpha\beta} + \frac{12}{\phi} G_{6X} G^{\alpha\beta} \pi_{;\alpha\beta} \phi^{;\gamma} \pi_{;\gamma} + \frac{12}{\phi} G_{6X} \phi^{;\alpha\beta} \pi_{;\beta\gamma} \pi^{;\gamma\alpha} \\ & + \frac{6}{\phi} G_{6X} \left( (\square\phi) (\square\pi)^2 - (\square\phi) \pi_{;\alpha\beta} \pi^{;\alpha\beta} - 2 (\square\pi) \phi^{;\alpha\beta} \pi_{;\alpha\beta} \right) \\ & + \frac{4}{\phi} G_{6XX} \phi^{;\alpha} \pi_{;\alpha} \left( (\square\pi)^3 - (\square\pi) \pi_{;\alpha\beta} \pi^{;\alpha\beta} + 2 \pi_{;\alpha\beta} \pi^{;\beta\gamma} \pi^{;\gamma\alpha} \right). \end{aligned}$$

$$\mathcal{L}_{4A} = -\frac{\phi^2}{4} G_4 F_{\alpha\beta} F^{\alpha\beta} + \phi^2 G_{4X} F_{\alpha\gamma} F^{\alpha\beta} \pi^{;\beta} \pi^{;\gamma}$$

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$$\begin{aligned}
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& - \frac{9}{2\phi} F_{\alpha\beta} F_{\gamma}{}^{\alpha} \phi^{;\gamma\beta} - \frac{9\phi^2}{32} F_{\alpha\beta} F_{\gamma}{}^{\alpha} F_{\delta}{}^{\beta} F^{\gamma\delta} - \frac{9}{\phi^2} F_{\alpha\beta} F_{\gamma}{}^{\alpha} \phi^{;\beta} \phi^{;\gamma} - \frac{9}{\phi} F_{\alpha}{}^{\beta} \nabla^{\gamma} F_{\beta\gamma} \phi^{;\alpha} \\
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& \left. - \frac{9}{2\phi^2} F_{\alpha\beta} F^{\alpha\beta} \phi_{;\gamma} \phi^{;\gamma} - \frac{15}{16} \nabla^{\alpha} F_{\beta\gamma} \nabla_{\alpha} F^{\beta\gamma} + \frac{3}{8} \nabla^{\alpha} F_{\beta\gamma} \nabla^{\beta} F_{\alpha}{}^{\gamma} \right) \\
& + G_{6X} \left( -\frac{3}{4} F_{\alpha\beta} F^{\alpha\beta} (\square\pi)^2 - \frac{9}{2\phi} F_{\alpha\beta} F^{\alpha\beta} (\square\pi) \phi^{;\gamma} \pi_{;\gamma} + \frac{3}{2} R F_{\alpha\beta} F_{\gamma}{}^{\alpha} \pi^{;\beta} \pi^{;\gamma} \right. \\
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- Now, one can forget about 5 dimension and KK procedure.  
It can be considered a trick to obtain the desired Lagrangian  $\mathcal{L}_A$
- Alternatively one can find the desired combinations among all general types of terms  
Might be more general, but much harder.

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& + \frac{3}{4} F_{\alpha\beta} F^{\alpha\beta} \pi_{;\gamma\delta} \pi^{;\gamma\delta} + \frac{9\phi^2}{8} F_{\alpha\beta} F_{\gamma\delta} F_f{}^{\alpha} F^{\gamma\delta} \pi^{;\beta} \pi^{;f} - 6F_{\alpha\beta} F_{\gamma}{}^{\alpha} (\square\pi) \pi^{;\gamma\beta} \\
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Profit 2) Phenomenologically favored by **GW170817**

Modifications of Maxwell theory that are obtained from **KK** are self-tuned in a way, so gravitons and photons propagate at the same speed for wide class of Generalized Galileon theories.

$$c_g^2 = c^2$$

This is not very surprising, since both modes come from 5-dimensional metric.

## Modern Universe cosmology

- For Horndeski theory (and beyond Horndeski and DHOST theories)

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- For trivial Maxwell electrodynamics ( $c = 1$ ) it means  $c_{\mathcal{T}} = 1$  too.
- For **KK** modified Maxwell  $c^2 = c_{\mathcal{T}}^2 \neq 1$

- scalar-tensor theories have two dynamical sectors

$$S = \int dt d^3x a^3 \left[ \frac{\mathcal{G}_T}{8} (\dot{h}_{ik}^T)^2 - \frac{\mathcal{F}_T}{8a^2} (\partial_i h_{kl}^T)^2 + \mathcal{G}_S \dot{\zeta}^2 - \mathcal{F}_S \frac{(\partial_i \zeta)^2}{a^2} \right]$$

- We do not care about scalar sector now

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- Instead we consider additional U(1) vector field

$$S = \int dt d^3x a^3 \left[ \frac{\mathcal{G}_T}{8} \left( \dot{h}_{ik}^T \right)^2 - \frac{\mathcal{F}_T}{8a^2} \left( \partial_i h_{kl}^T \right)^2 + \mathcal{G}_V \dot{A}_i^2 - \mathcal{F}_V \frac{(\partial_j A_i)^2}{a^2} \right]$$

The speeds of sound for tensor and vector modes are, respectively,

$$c_T^2 = \frac{\mathcal{F}_T}{\mathcal{G}_T}, \quad c^2 = c_V^2 = \frac{\mathcal{F}_V}{\mathcal{G}_V}$$

- Horndeski theory:

$$\begin{aligned}\mathcal{G}_{\mathcal{T}} &= 2G_4 - 4G_{4X}X + G_{5\pi}X - 2HG_{5X}X\dot{\pi}, \\ \mathcal{F}_{\mathcal{T}} &= 2G_4 - G_{5\pi}X - 2G_{5X}X\ddot{\pi}.\end{aligned}$$

- beyond Horndeski theory:

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- DHOST theory:

$$\begin{aligned}\mathcal{G}_{\mathcal{T}} &= 2f_2 + 2\ddot{\pi}Xf_{3,X} - Xf_{3,\pi} - 2Xa_1 \\ &\quad + 2X(3\dot{\pi}H + \ddot{\pi})b_2 + 6\dot{\pi}XHb_3 + 2\ddot{\pi}X^2b_6, \\ \mathcal{F}_{\mathcal{T}} &= 2f_2 - 2\ddot{\pi}Xf_{3,X} + Xf_{3,\pi},\end{aligned}$$

# conventional Maxwell $c = 1$

- Horndeski:
  - $G_4 = G_4(\pi)$
  - $G_5 = \text{const}$
- Beyond Horndeski
  - $F_4 = \frac{2G_4 X}{X}$
  - $G_5 = \text{const}$
- DHOST
  - $a_1 = 0$
  - $f_3 = 0, \quad b_i = 0$



## Modified Maxwell

- Horndeski:

$$G_4 = G_4(\pi, X)$$

$$G_5 = G_5(\pi)$$

- Beyond Horndeski

$$G_4 = G_4(\pi, X)$$

$$F_4 = F_4(\pi, X)$$

$$G_5 = G_5(\pi)$$

- DHOST

$$f_2 = f_2(\pi, X)$$

$$a_1 = a_1(\pi, X)$$

$$a_3 = a_3(\pi, X)$$

## Final Remarks

- Some subclasses of luminal Horndeski with modified Maxwell were known by disformal trick

$$\text{BH with } F_4 = \frac{2G_4 X}{X} \xrightarrow{\text{disformal transformation}} \text{H + modified EM}$$

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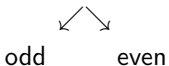
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**! Dark Energy can be made with beyond Horndeski theory**

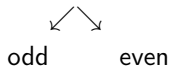
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2 tensor modes



2 vector modes

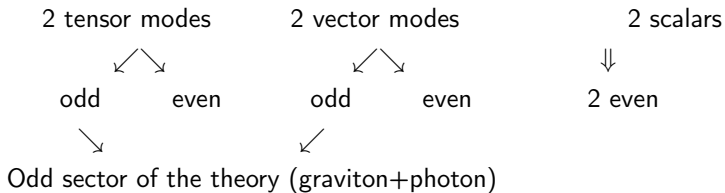


2 scalars

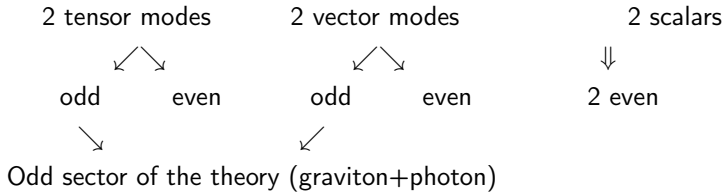




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- Vainshtein mechanism works for modified Maxwell similarly to modified gravity

## Conclusion and outlook

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= ...only second derivatives