

Modelling of light (hyper)nuclei production in heavy-ion collisions at NICA energies based on generator THESEUS

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61st meeting of the PAC
for Particle Physics



INTRODUCTION

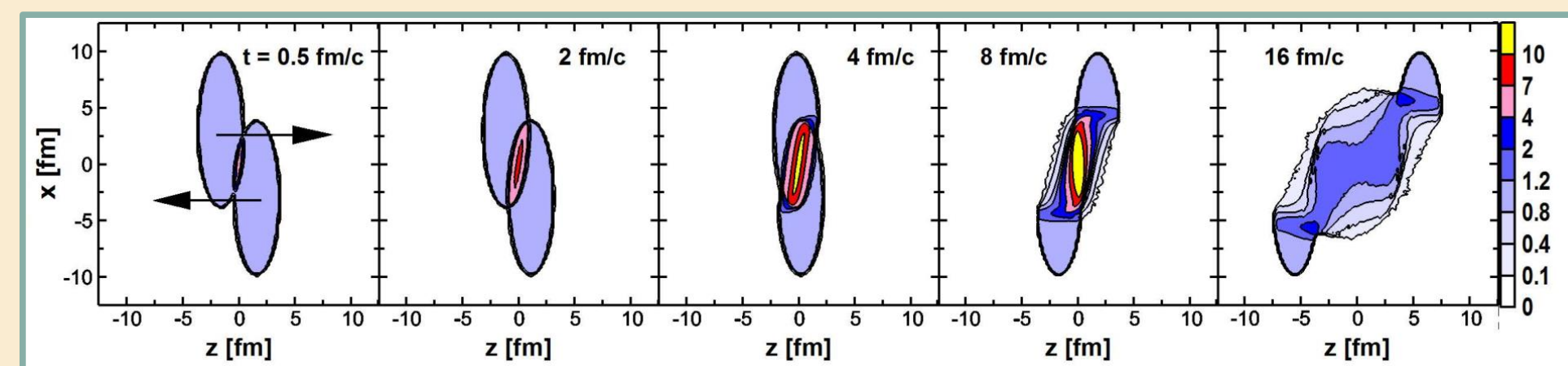
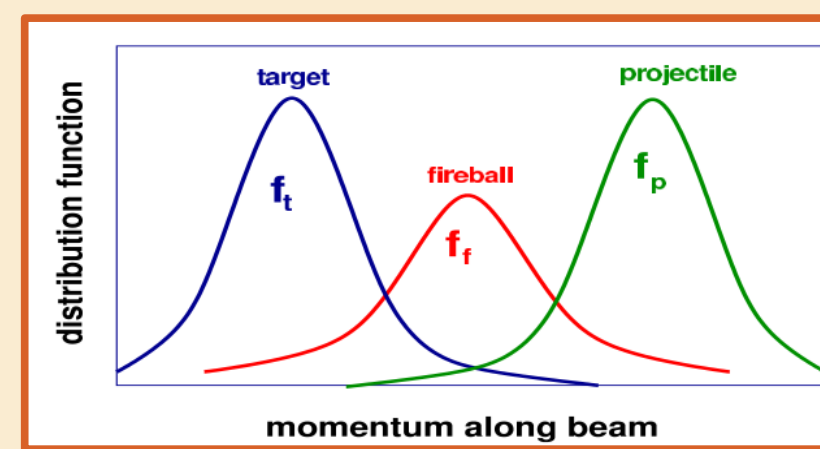
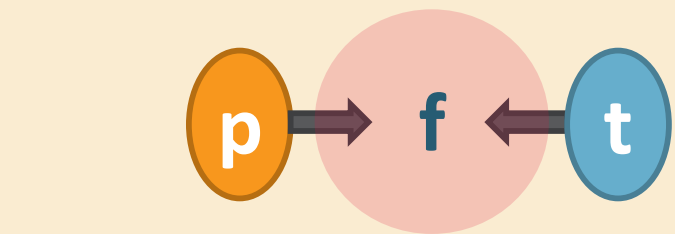
- Light-nuclei production is related to search for critical point in QCD phase diagram.
- Spinodal instability: abundant production of light nuclei due to density fluctuations in spinodal region.
- At moderately relativistic energies light nuclei give significant contribution to the total baryon charge of the system of colliding nuclei ($\approx 30\%$ in Au+Au collisions at the energy $\sqrt{s_{NN}} = 3$ GeV).
- Various 3D dynamical models describing light nuclei exist: with coalescence mechanism, microscopic dynamical approaches (SMASH, PHQMD).
- THESEUS (Three-fluid Hydrodynamics-based Event Simulator Extended by UrQMD (Ultrarelativistic Quantum Molecular Dynamics) final State interactions) generator uses **thermodynamical approach** for modelling of light (hyper)nuclei on the equal basis with hadrons. **No additional parameters are needed.**
- Main areas of research: study the light-(hyper)nuclei production at collision energies of the BES-RHIC, SPS, NICA and FAIR.

1 Model 3FD and generator THESEUS

The 3FD [1] simulates the early, nonequilibrium stage of the collision and describes it until the freeze-out.

- baryon-rich fluids:** nucleons of the projectile (p) and the target (t) nuclei;
- fireball (f) fluid:** newly produced particles which dominantly populate the midrapidity region during the evolution process.

Problems: no afterburner, application.



Hadronic EoS
only hadronic states (hadron gas), no phase transition (Mishustin, Russkikh, Satarov)

1PT EoS
hadronic states+QGP with 1st-order phase transition (Khvorostukhin, Skokov, Toneev, Redlich)

Crossover EoS
hadronic states + QGP with a smooth transition (crossover) between phases (Khvorostukhin, Skokov, Toneev, Redlich)

THESEUS [2] is based on the 3FD model, performs the procedure of particlization. The kinetic stage (or afterburning) is modeled by means of UrQMD, which describes hadronic rescatterings.

THESEUS = 3FD + Monte-Carlo hadron sampling + afterburner via UrQMD.

THESEUS uses 3FD output (T, μ_B, μ_S) to generate particles ($x, y, z, p_x, p_y, p_z, E, \dots$);

It presents the output in terms of a **set of observed particles.**

It is suitable for application of experimental acceptance!

1. Initial state & hydrodynamic evolution :
3FD uses hydrodynamical equations with friction terms modelling of the interaction between fluids

2. Particlization:
THESEUS allows to move from the description in terms of liquid to a kinetic description.

3. Afterburner:
UrQMD (optionally) describes hadronic rescattering (or afterburning) processes.

4. Detector response:
Experimental cuts can be easily applied

HYDRODYNAMICAL MODELLING OF NUCLEAR COLLISIONS

2 RESULTS

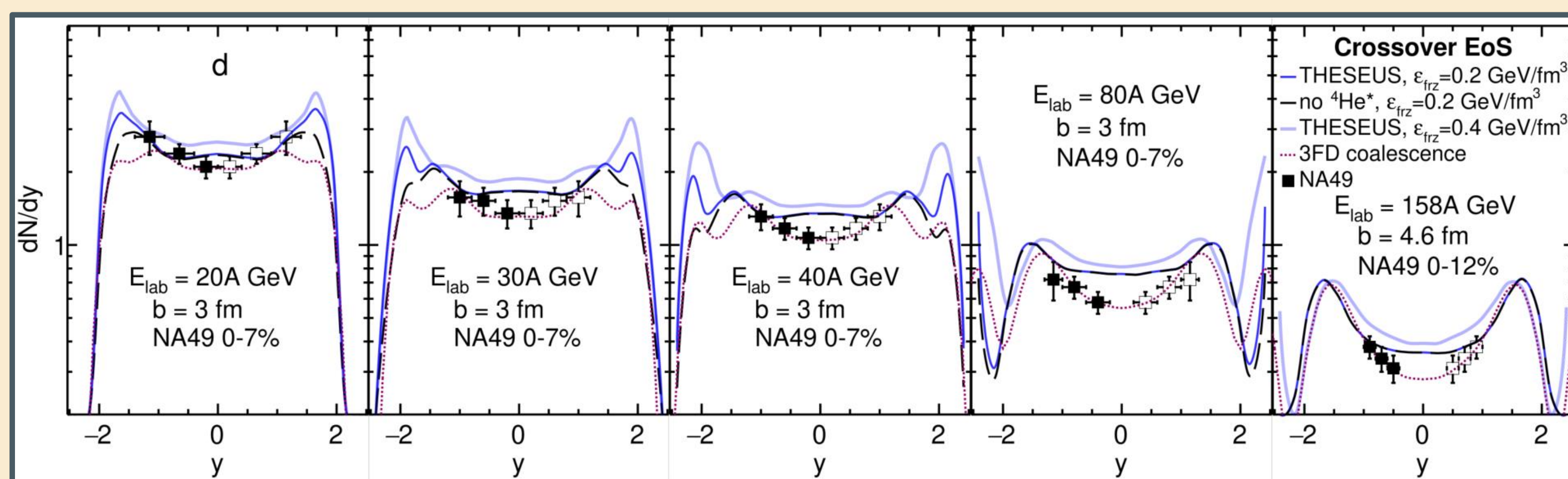


Fig.1: THESEUS and 3FD results [3] on rapidity distributions of deuterons in comparison with NA49 [4] data calculated for three different EoS and with late freeze-out ($\epsilon_{frz} = 0.2$ GeV/fm³). For comparison, the result with standard freeze-out ($\epsilon_{frz} = 0.4$ GeV/fm³) for crossover EoS is shown.

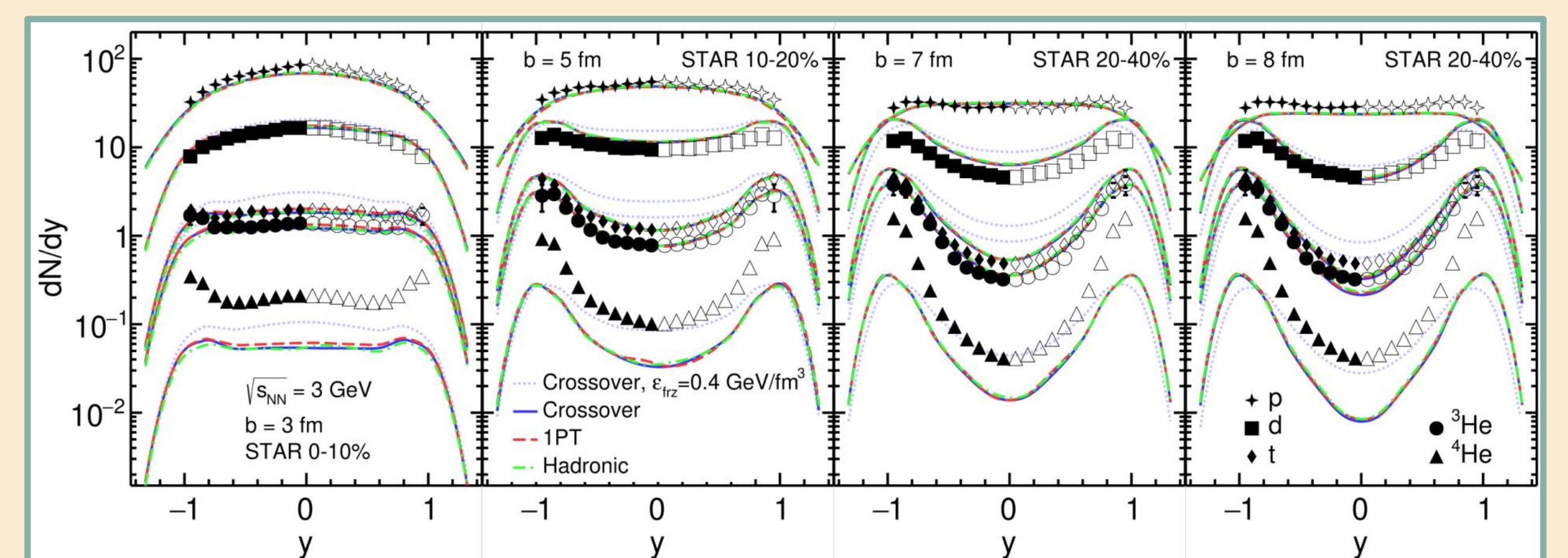


Fig.2: THESEUS results [5] on rapidity distributions of protons and light nuclei (deuterons, tritons, ³He, ⁴He) at $\sqrt{s_{NN}} = 3$ GeV in comparison with STAR data [6] calculated for three different EoS and with late freeze-out ($\epsilon_{frz} = 0.2$ GeV/fm³). The result with standard freeze-out ($\epsilon_{frz} = 0.4$ GeV/fm³) for crossover EoS is shown.

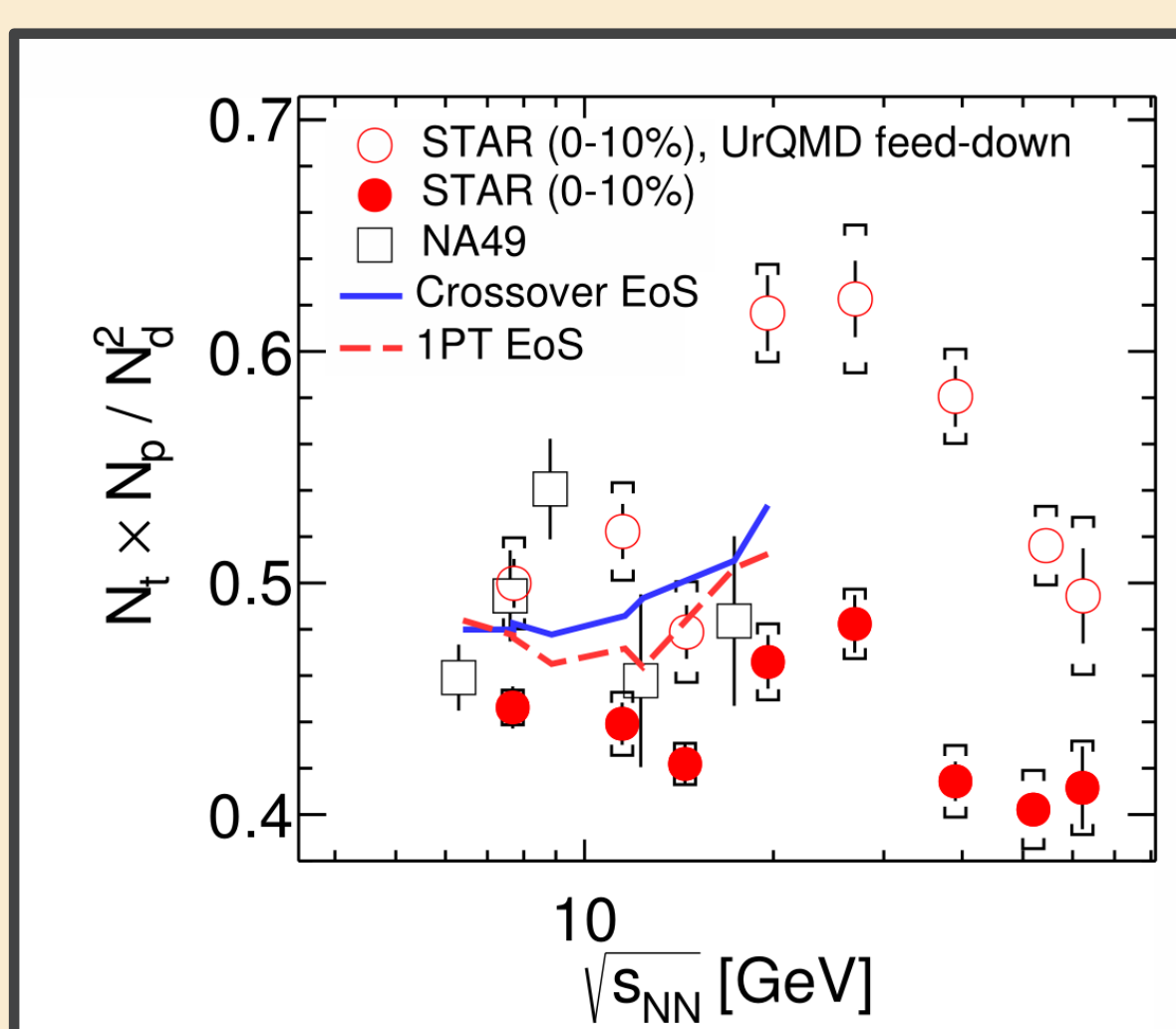


Fig.3: Energy dependence of the midrapidity light-nuclei-yield ratio in central Au+Au and Pb+Pb collisions [3]. NA49 data [4] and STAR data (preliminary [7] and final [8]) are shown.

Fig.4: Particle ratios [9] in dependence of rapidity for t/p, ³He/p, ³H/ Λ , ⁴He/ Λ in comparison with STAR data (Yuanjing Ji, talk at QM 2023). Different EoS are used. For p and Λ the standard freeze-out ($\epsilon_{frz} = 0.4$ GeV/fm³) with UrQMD afterburner is used. For t and ³He the late freeze-out $\epsilon_{frz} = 0.2$ GeV/fm³ is used. For ⁴He and ⁴H the standard one is used.

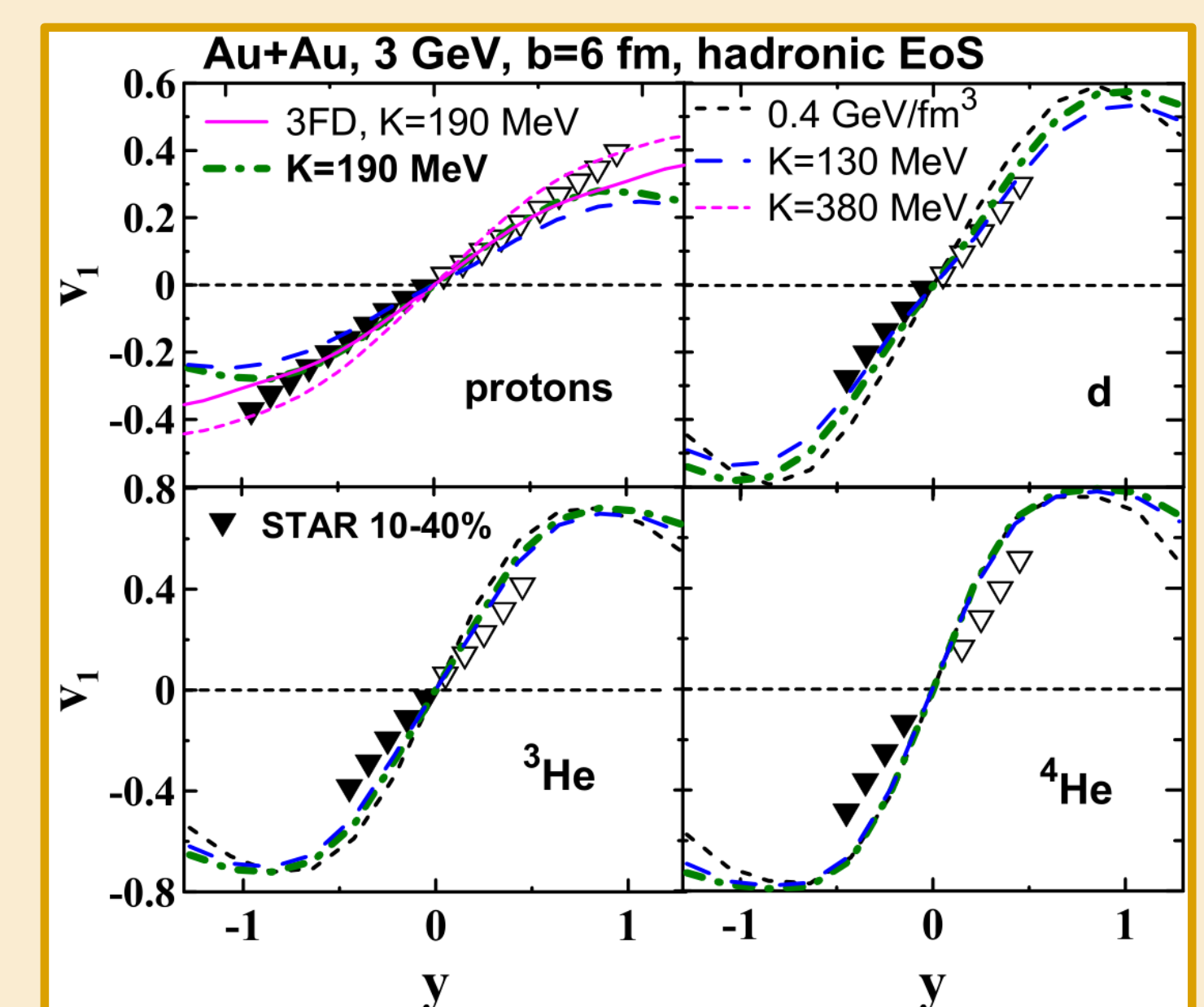
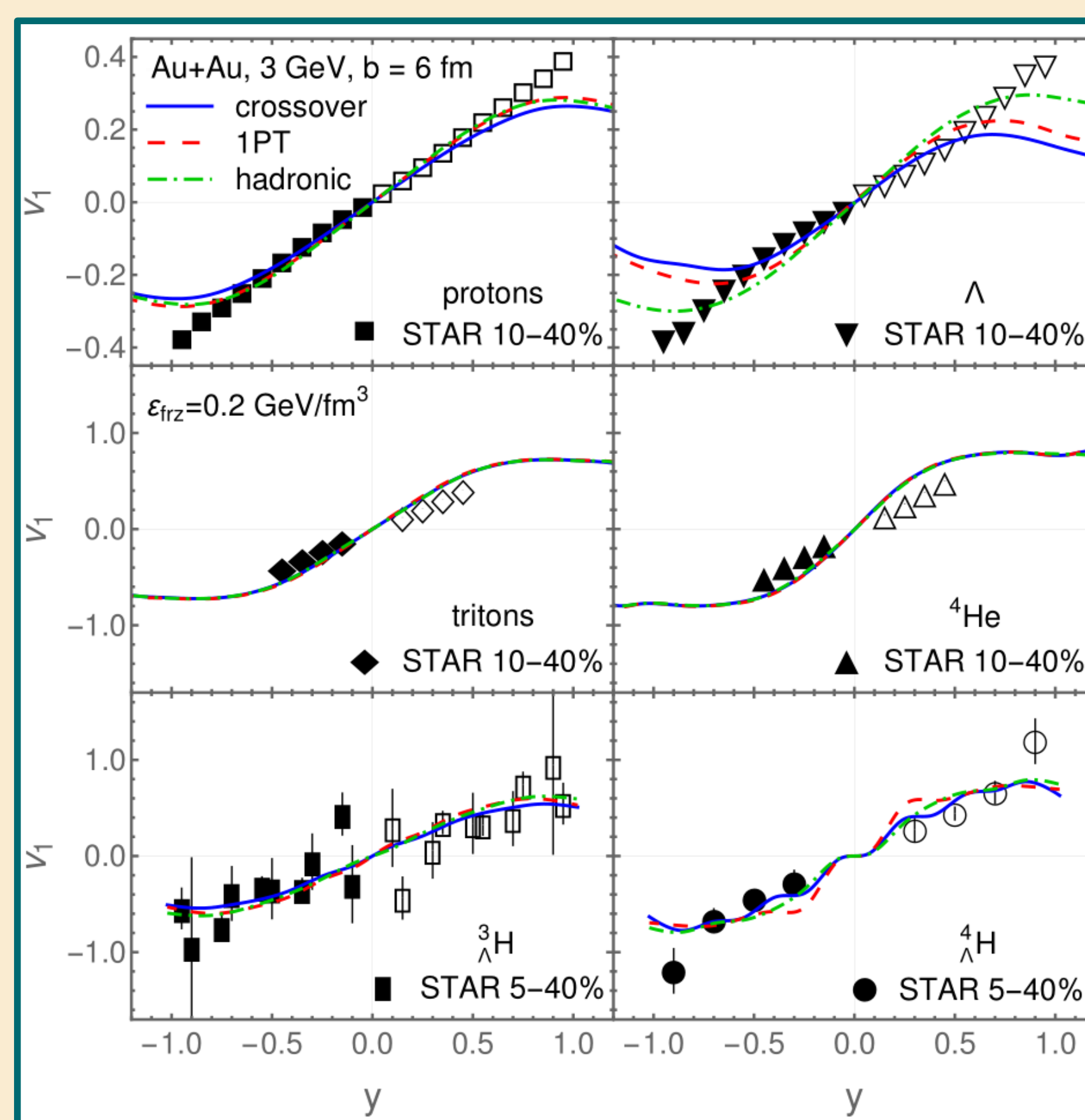
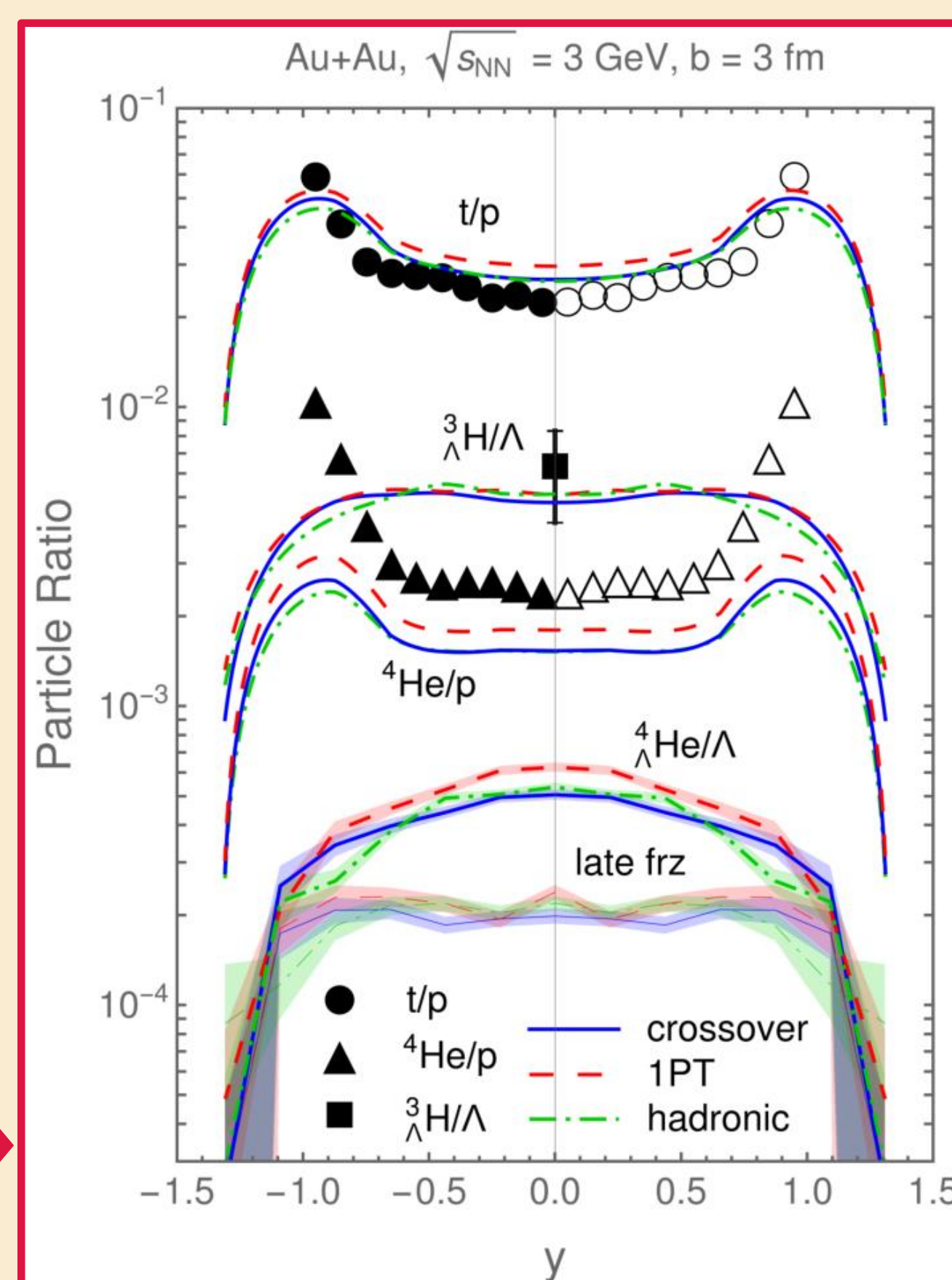


Fig.5: Directed flow of protons and light nuclei [5] as function of rapidity in semicentral ($b = 6$ fm) Au+Au collisions. Hadronic EoS with different stiffness K is used. Incompressibility: $K = 9n_0^2 \frac{d^2 \epsilon(n, T=0)}{dn^2}$, where $\epsilon(n, T=0)$ is the energy density of the nuclear matter at $T = 0$, n_0 is the normal nuclear density.

Fig.6: Directed flow of protons, light nuclei and hypernuclei as function of rapidity in semicentral ($b = 6$ fm) Au+Au collisions.

3 DISCUSSION

- Afterburner of almost all light (hyper)nuclei is modelled by the late freeze-out ($\epsilon_{frz} = 0.2$ GeV/fm³), while ⁴He, by standard one ($\epsilon_{frz} = 0.4$ GeV/fm³).
- Reasonable reproduction of data on bulk observables of the light nuclei in the energy range of $\sqrt{s_{NN}} = 3 - 19.6$ GeV.
- Ratio $N(t) \times N(p)/N^2(d)$ is considered. Accurate subtraction of weak-decays feed-down from proton yield is important.
- Good description of difference in the form of proton and light-nuclei distributions at $\sqrt{s_{NN}} = 3$ GeV and its dependence on the centrality. At $\sqrt{s_{NN}} = 3$ GeV the choice of EoS does not affect on the rapidity distributions: no phase transition takes place.
- The directed flow: good agreement of p and Λ especially in midrapidity region, no dependence on EoS for (hyper)nuclei, for hypernuclei results are in agreement with experiment within statistical uncertainties. Standard stiffness with $K = 190$ MeV is the most suitable.

4 CONCLUSION

- THESEUS gives generally good agreement of bulk observables and directed flow with experimental data. THESEUS simulations give reasonable results even for hypernuclei.
- For treatment of almost all studied (hyper)nuclei we need late freeze-out, so afterburner plays significant role in their production. An exception is ⁴He which requires standard freeze-out.
- Dominance of hadronic phase in evolution of the system at the collision energy of $\sqrt{s_{NN}} = 3$ GeV.
- THESEUS can be used for predictions for future experiments.

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DATE :20/01/2025

