

Baikal-GVD: Deep-Underwater Neutrino Telescope: Status and Results

60th meeting of the PAC for Nuclear Physics Bair Shaibonov on behalf of the Baikal-GVD collaboration, Dubna, 23.01.2025



Outline

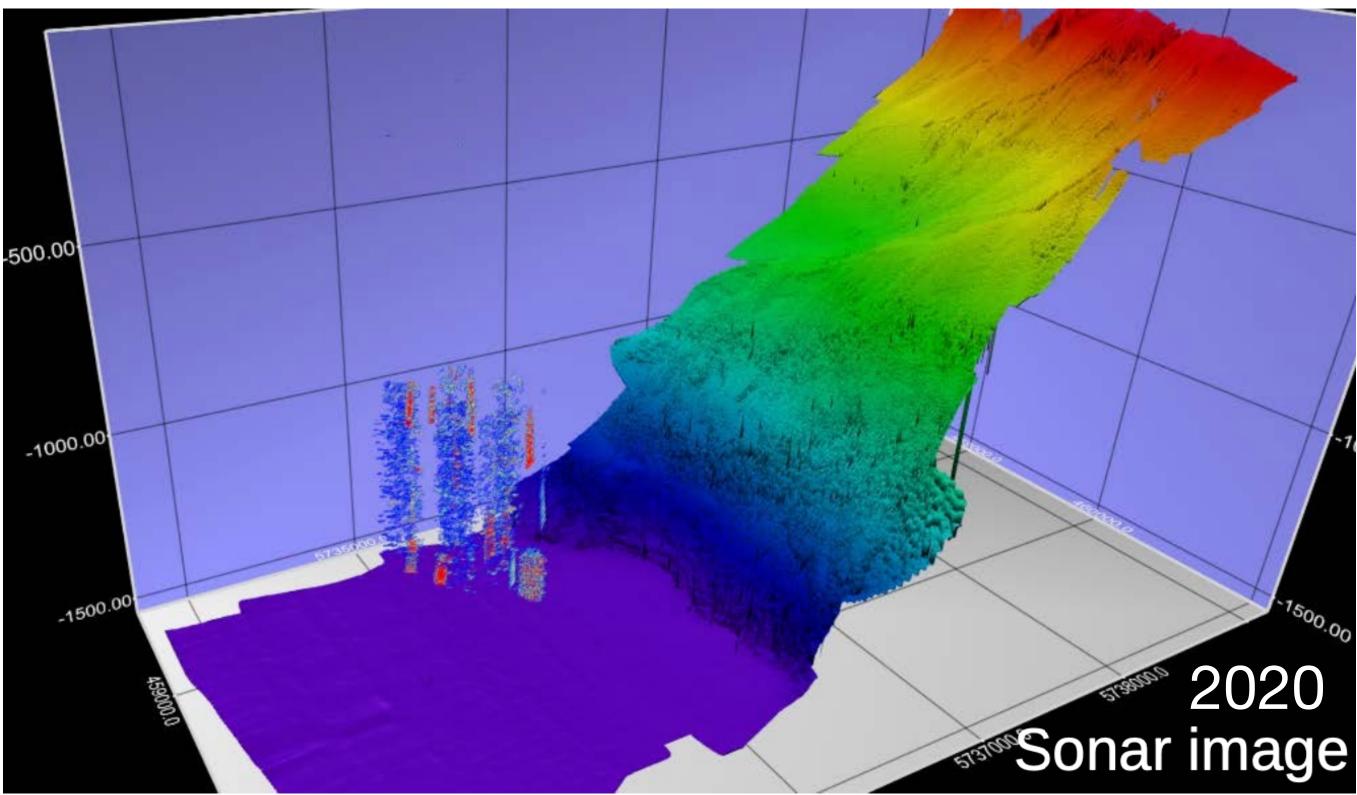
- Baikal-GVD Telescope Description and Status
- Nearest Plans
- Recent Results:
 - Characterisation of diffuse astrophysical flux of high-energy neutrino
 - Search for astrophysical neutrino point sources
 - Search for extended neutrino source: Galactic plane
 - Follow-up activities



Baikal-GVD Site

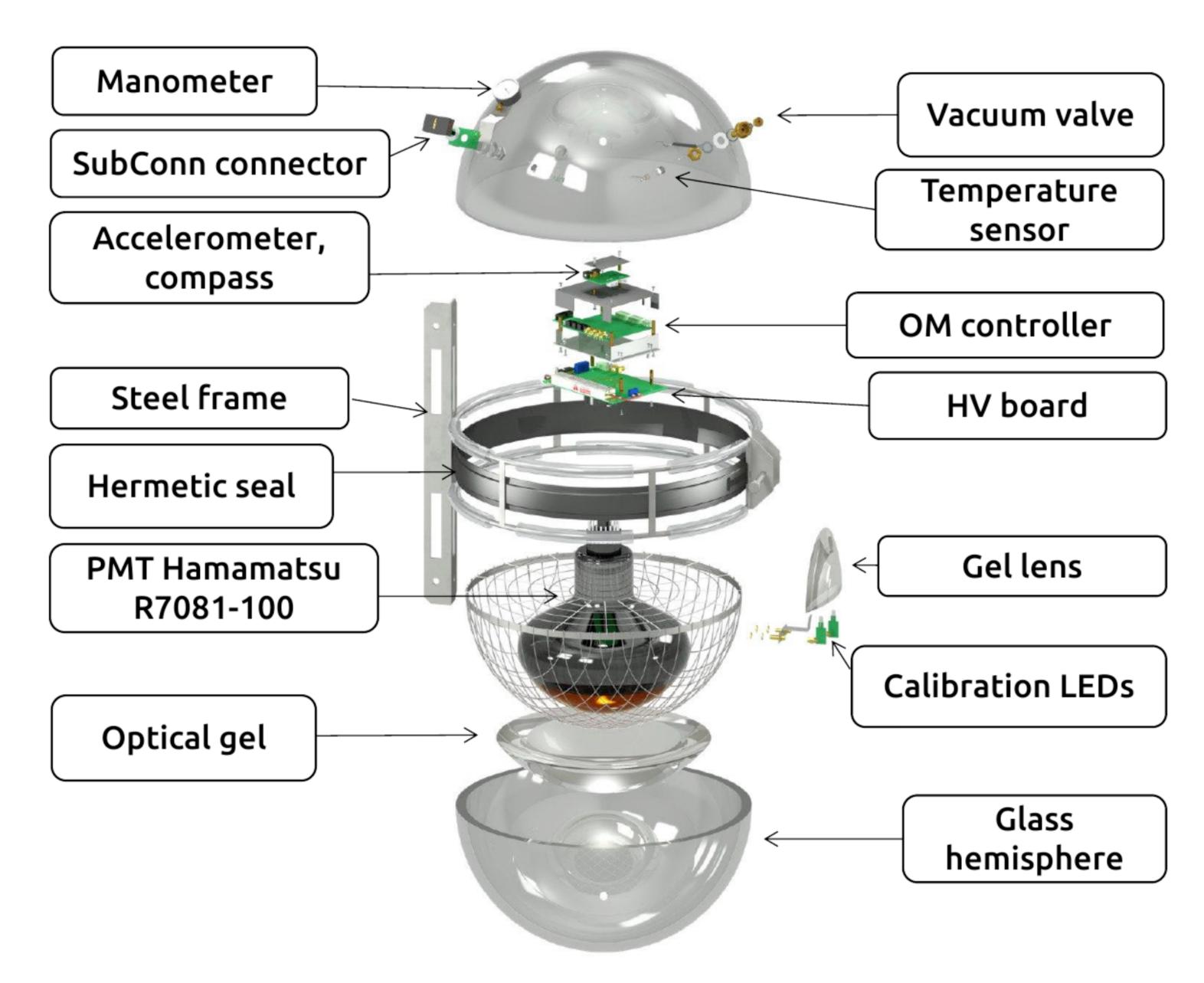


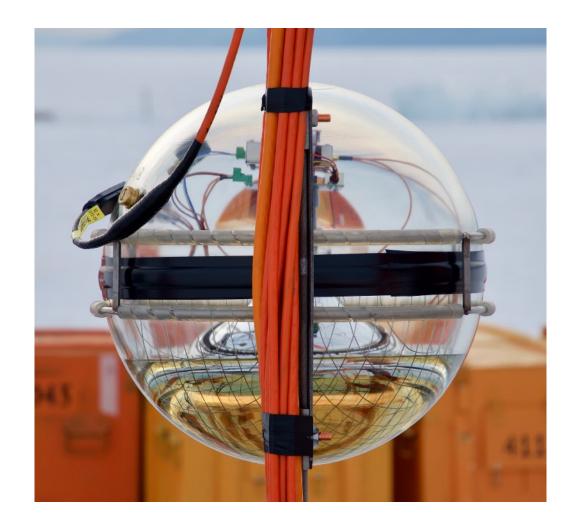
- Southern basin of the lake \bullet
- ~3.6 km offshore lacksquare
- Flat area at depths 1366–1367 m \bullet
- High water transparency: \bullet
 - Absorption length: 22 m \bullet
 - Effective scattering length: 480 m \bullet
- Moderately low optical background: 15–50 kHz
- Deployment from the ice cover of the lake





Optical Module - Basic Element of the Telescope



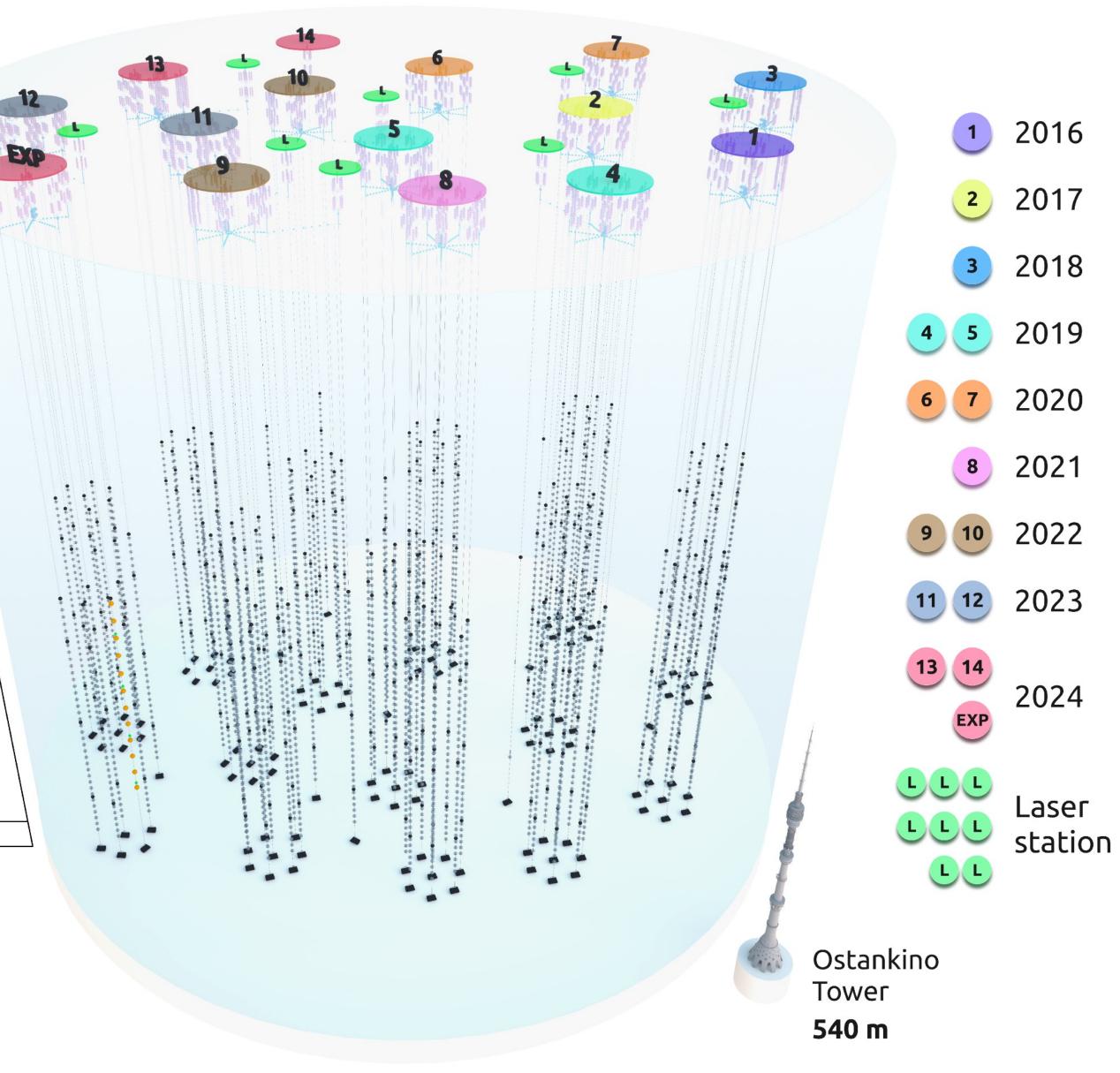


17 inches sphere (42 cm)

10 inch Hamamatsu PMT R7081-100

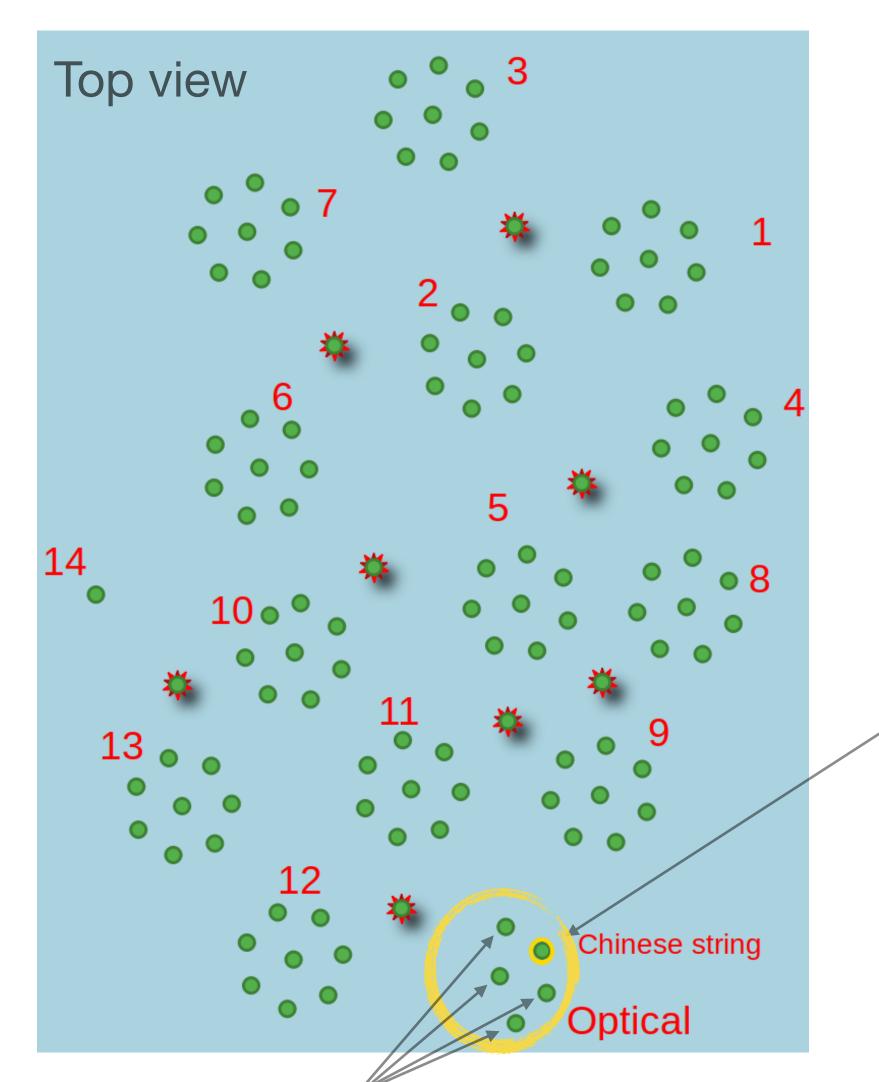


	Baikal-GVD State	JS
	April 2024	
•	4104 Optical modules on 114 strings (13 clusters)	
•	0 8 strings form a cluster - independent array of optical modules	m
•	36 optical modules per string	
•	60 m between strings in a cluster, 250-300 m between clusters	
•	More than 0.6 km ³ of water volume	
•	8 laser stations/inter-cluster strings	750 m —
•	More than 400 acoustic modules for positioning	525 m 36 OM
•	LED beacons and powerful laser sources for calibration	1275 m – 1366 m –
•	4 experimental strings with the fibre-optic DAQ for testing of new equipment	
•	Prototype string for the next-generation telescope (12 new OMs)	



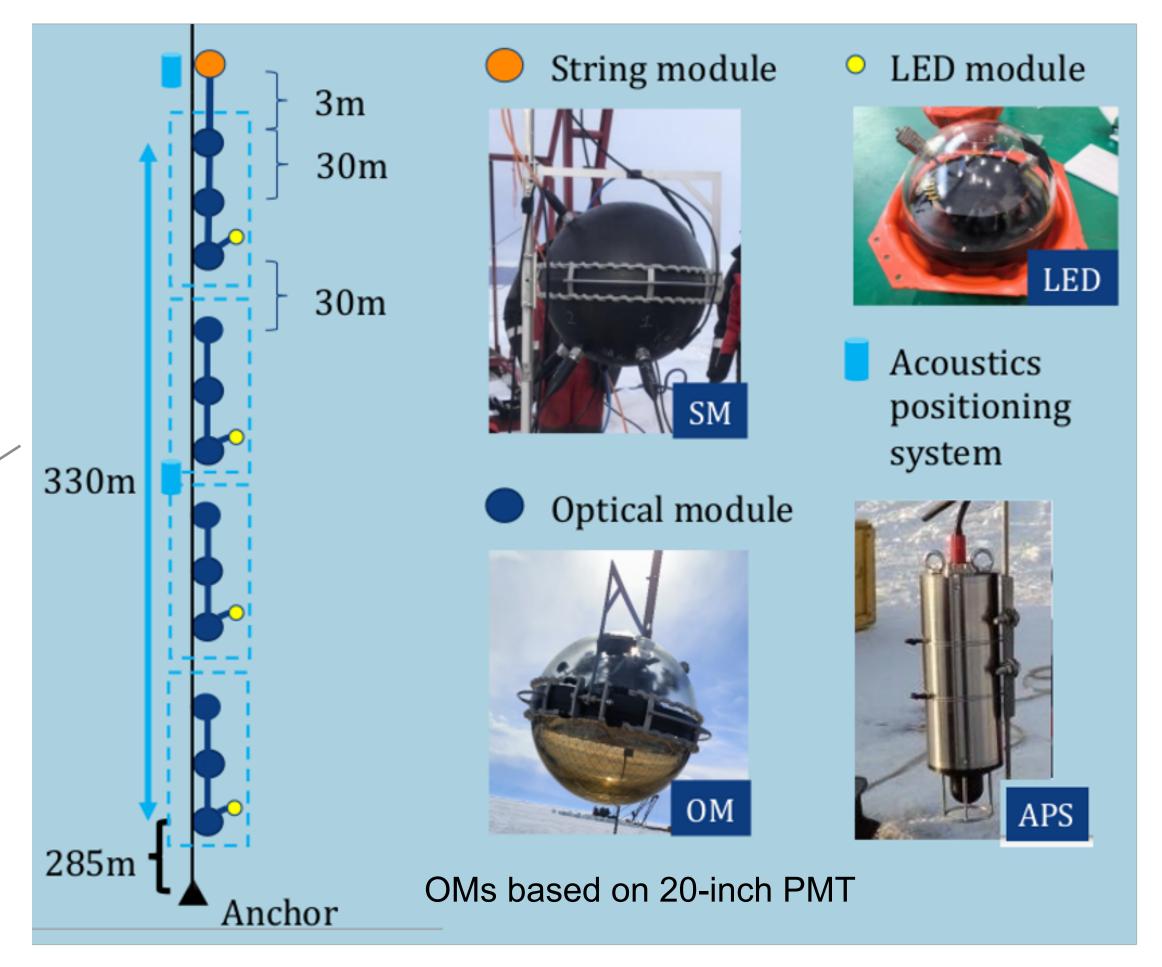


Technological prototype strings (2024)



Four "experimental" strings with new fibre-optic technology for data transmission

Next-generation prototype string deployed in 2024 (IHEP (Beijing) & Baikal-GVD joint effort)



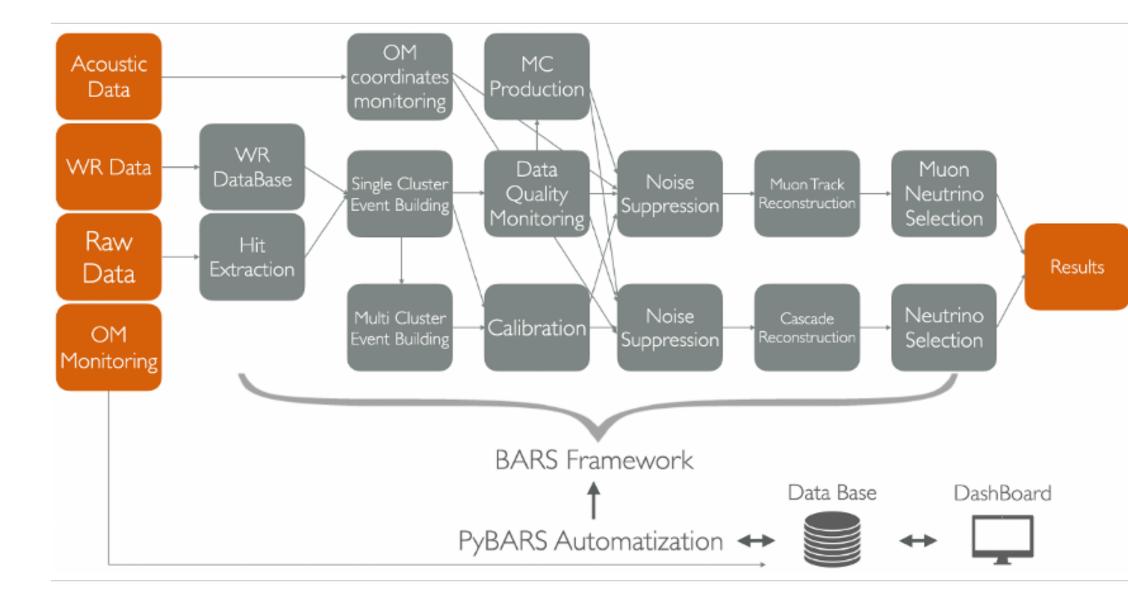
Next generation neutrino telescope project [PoS(ICRC2023)1080]











Data handling

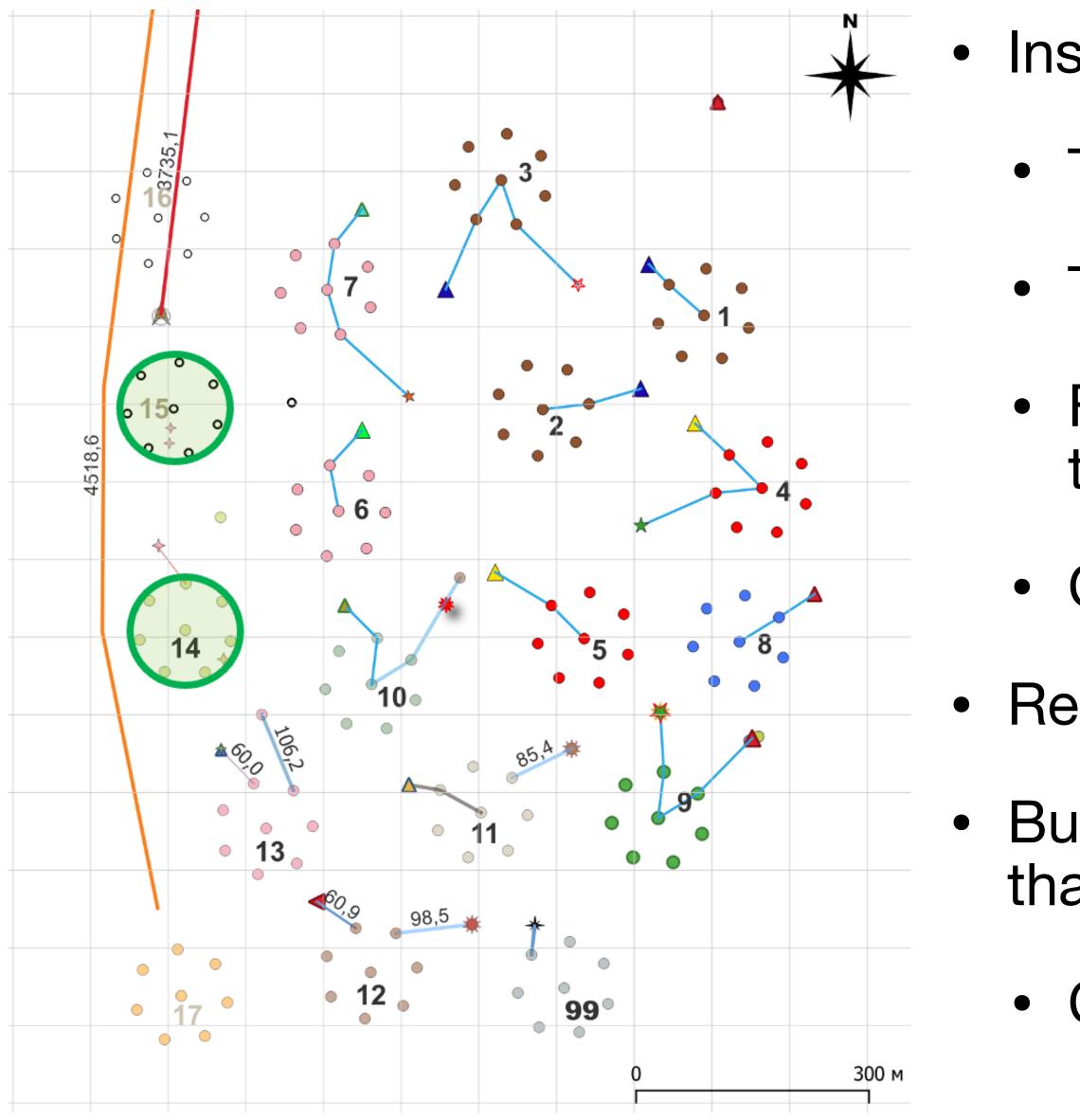
Raw data are transferred from the Shore Center to JINR:

- Shore center
 → Baikalsk: 300 Mbit/s radiochannel
- Baikalsk → JINR: Internet
- Compressed data volume ~10-40 GB per day per cluster
- Full-scale reconstruction at JINR
- Delay due to shore JINR data transfer < 1 min
- Full processing delay is about 10-15 min





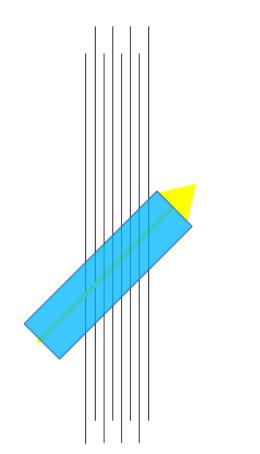
Next Expedition Plans (2025)



- Installation of new equipments:
- Two new clusters: 14th and 15th
- Two bottom cable lines
- Full-scale string for the next-generation telescope
- Cluster Center for cluster 17
- Repairing some parts
- But: autumn and winter in Siberia were warmer than usual
 - Challenging ice conditions

Event Topologies

Single-cluster tracks



- Low energy threshold
- Optimal sensitivity to nearly vertical tracks
- 90% of recorded track events

Single-cluster cascades

- High energy threshold
- Good energy resolution
- Relatively rare events

Main results for today

Multi-cluster tracks

- Moderately low energy threshold
- Optimal sensitivity to inclined tracks
- Best angular resolution

NC, $\nu \nu$ CC

 $\nu_{\mu} CC$

Multi-cluster cascades

- Very high energy threshold
- Excellent energy resolution
- Very rare events



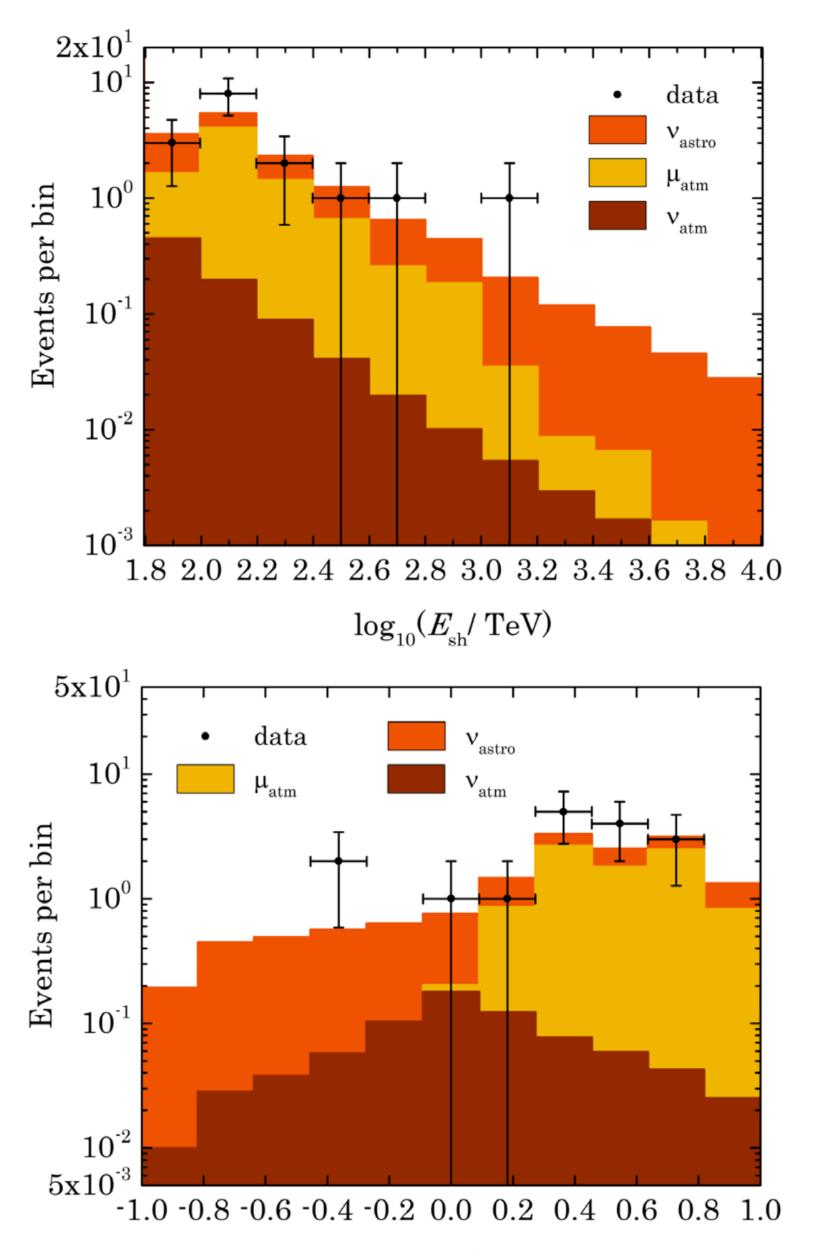


Astrophysical Diffuse Neutrino Flux: All-sky

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 \bullet

lacksquare



Excess over the atmospheric background: 2.22σ

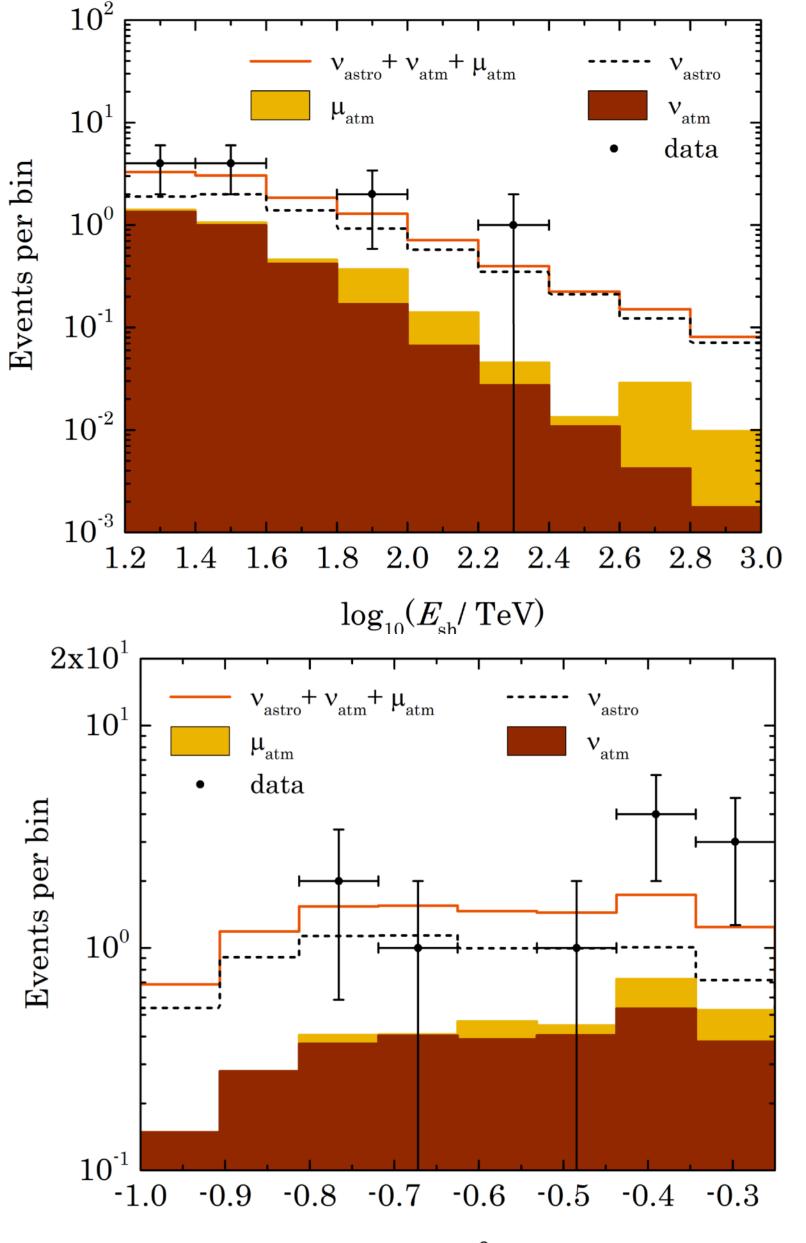
- Data analysed April 2018 March 2022
- 14328 cascades reconstructed with E > 10 TeV, $N_{hit} > 11$
- Cascade energy E >70 TeV and $N_{hit} > 19$

	Events
Atm. muons MC	7.4
Atm. neutrino MC	0.8
Astro neutrino MC best fit	5.8
Data	16

Phys. Rev. D 107, 042005, February 2023



Astrophysical Diffuse Neutrino Flux: Upward-Going Events



 $\cos \theta$

- Data analysed April 2018 March 2022
- Less background from below:
 - Improving purity and lower energies
- Cascade energy >15 TeV

	Events
Atm. muons MC	0.5
Atm. neutrino MC	2.7
Astro neutrino MC best fit	6.3
Data	11

Excess over the atmospheric background: 3.05σ

Phys. Rev. D 107, 042005, February 2023



Single Power-Law Model of Astrophysical Flux

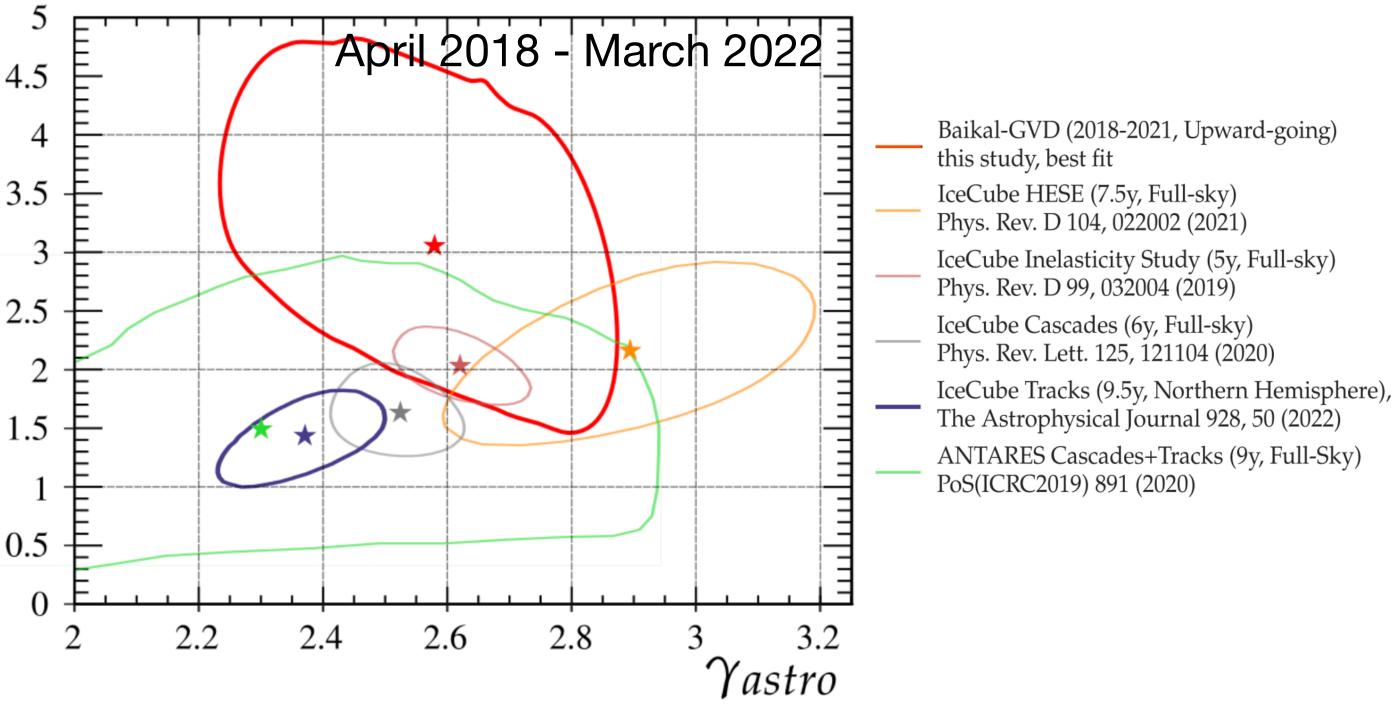
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The best fit parameters for the single power law model:

$$\Phi_{astro}^{\nu+\bar{\nu}} = 3 \times 10^{-18} \phi_{astro} \left(\frac{E_{\nu}}{E_0}\right)^{-\gamma_{astro}} \text{GeV}^{-1} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1} \qquad 3.4$$

$$\gamma_{astro} = 2.58^{+0.27}_{-0.33} \qquad 1.4$$

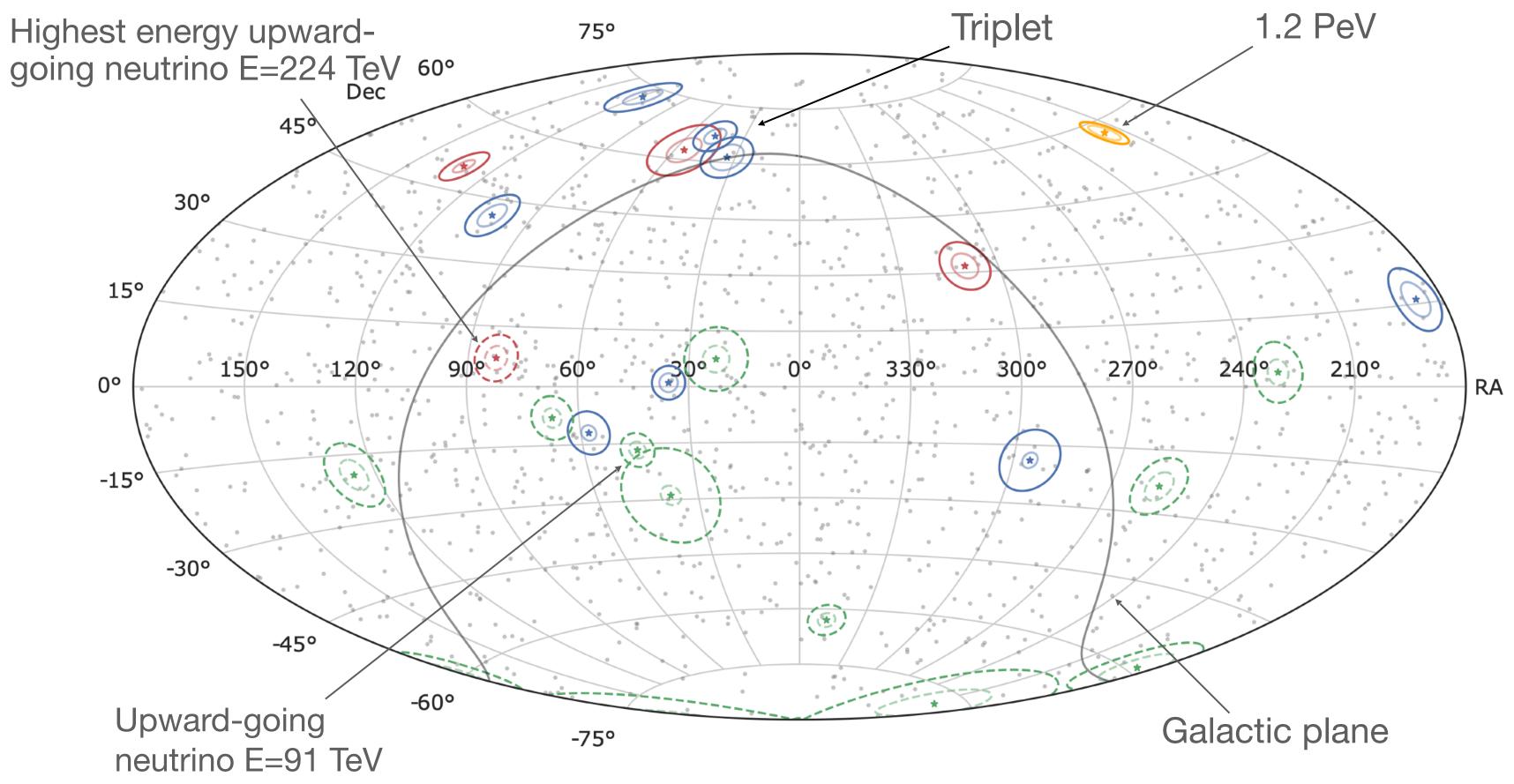
$$\phi_{astro} = 3.04^{+1.52}_{-1.27} \qquad 0.4$$





High-Energy Cascade Sky Map

Opens a possibility to use the cascade channel for searches for neutrino point sources



Best fit positions and 90% angular uncertainty regions

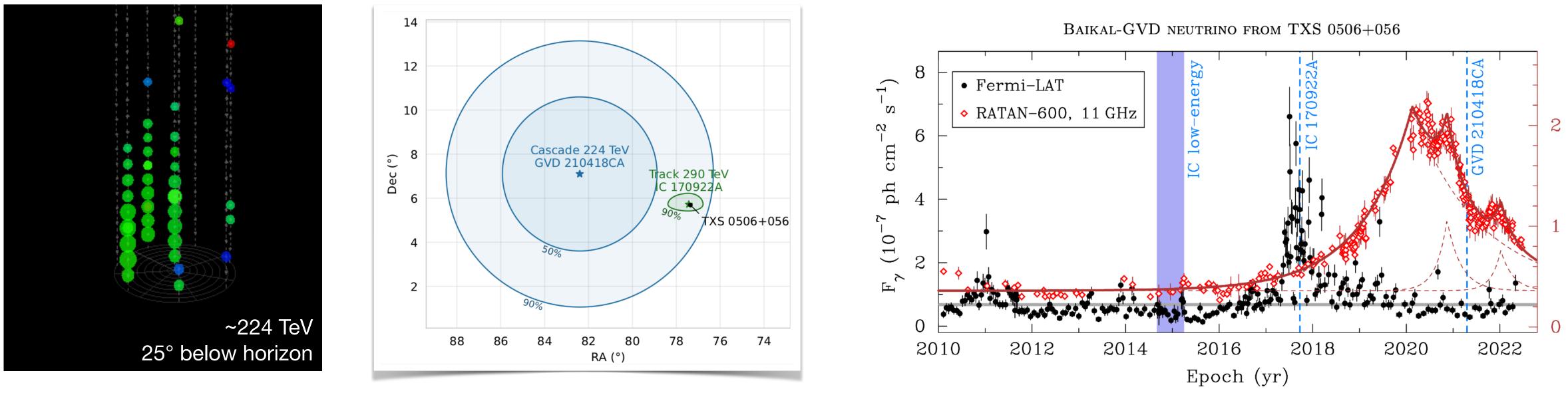
Monthly Notices of the Royal Astronomical Society, Volume 526, Issue 1, November 2023, Pages 942-951

- About half of the events are background from atmospheric muons and neutrinos
- Grey dots: radio-bright blazars (3.6 σ correlation with IceCube tracks with E > 200 TeV)
- No significant correlation between Baikal-GVD cascades with E > 100 TeV and radio-bright blazars was found





Most energetic upgoing cascade event **Best candidate for neutrino events of astrophysical origin**



Closest sources (in 6 degrees):

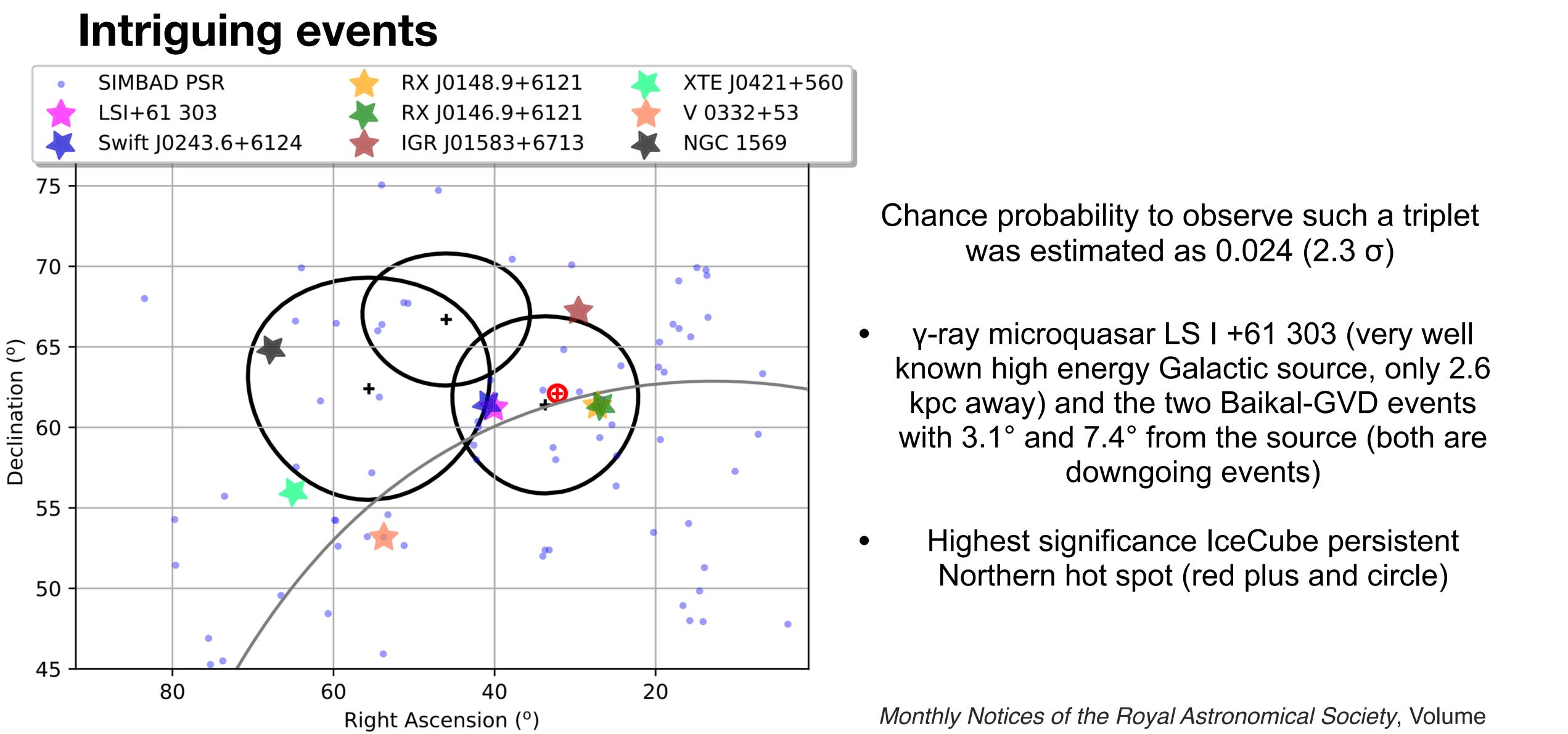
- This event is probably of astrophysical origin (signalness = 97%).
- Chance probability of coincidence p=0.0074 (2.7 σ)

TXS 0506+056 Blazar (BL Lac) at z= 0.34 (5.7 Gly) is IceCube neutrino source observed at 3.7 σ

Monthly Notices of the Royal Astronomical Society, Volume 527, Issue 3, January 2024, Pages 8784–8792



Event Triplet near Galactic Plane Intriguing events

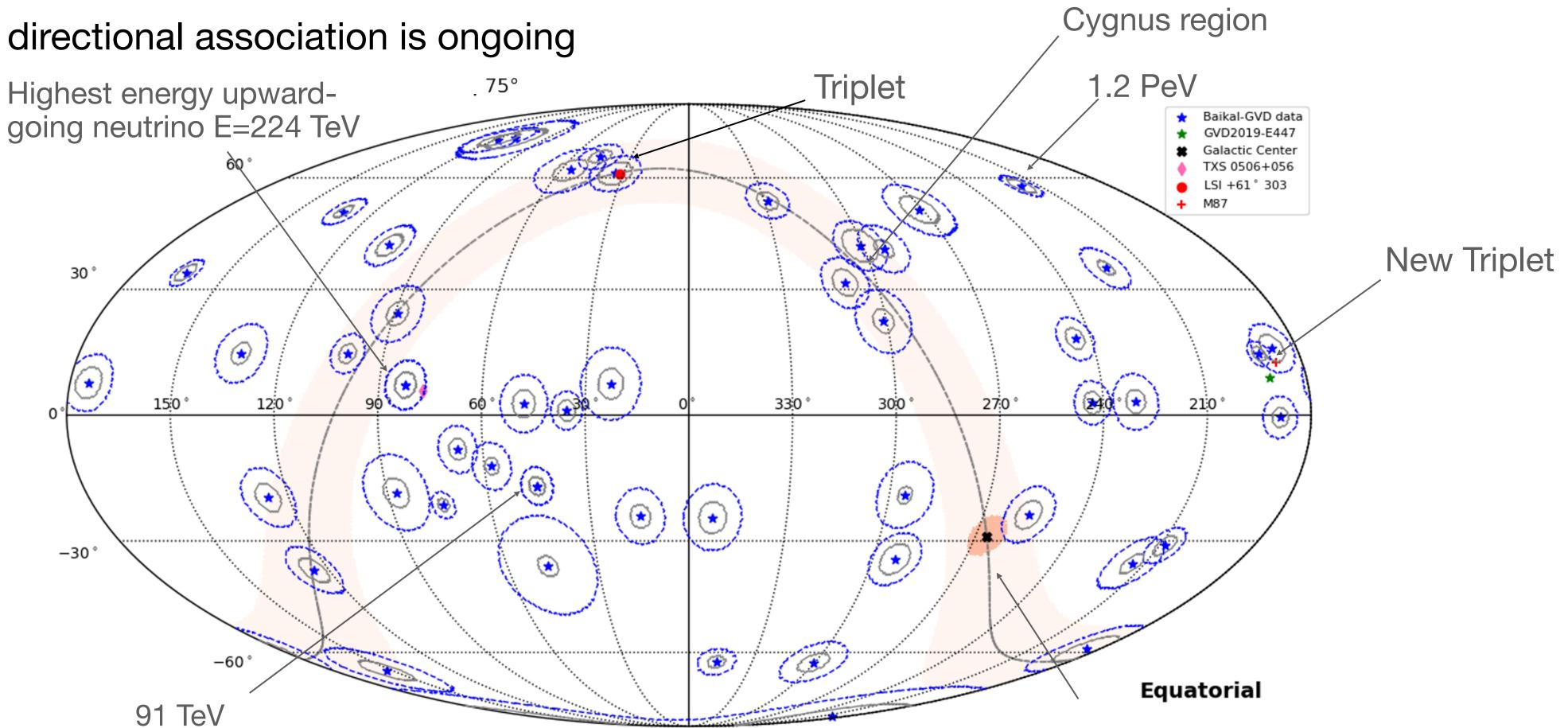


526, Issue 1, November 2023, Pages 942–951

New High-Energy Cascade Sky Map

Data from April 2022 to March 2024 double the statistics:

- Excess over the atmospheric background is 5.54σ .
- Search for directional association is ongoing



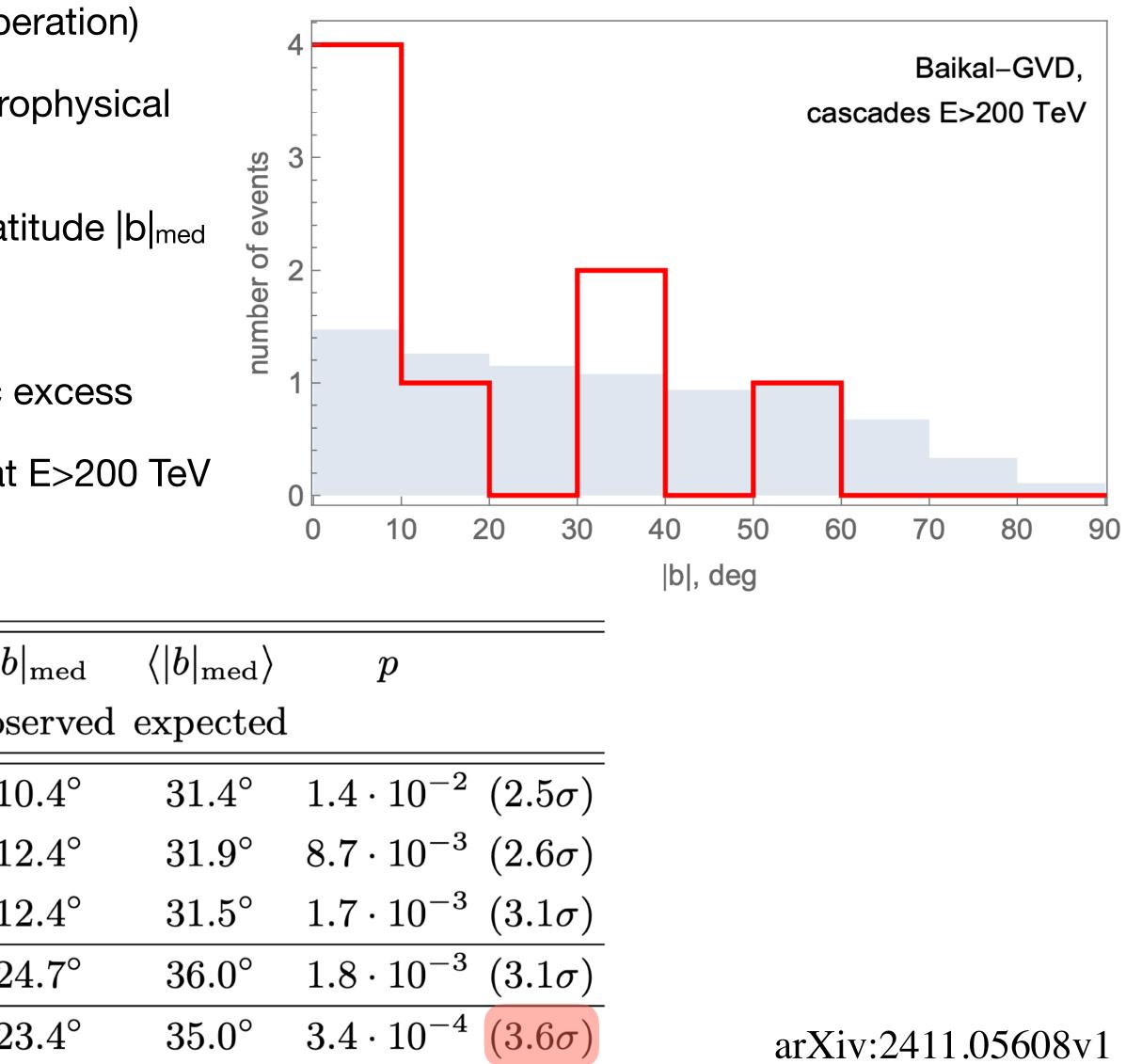
- Best fit positions and 90% angular uncertainty regions
- About half of the events are background from atmospheric muons and neutrinos



Galactic Neutrinos with the Highest Energies

- High-energy cascades April 2018- March 2024 (6 years of operation) \bullet
- Test the Galactic excess at E>200 TeV (8 events, 64% of astrophysical) \bullet origin)
- Simplest model-independent test using median of galactic latitude |b|_{med}
- Galactic component is visible with a significance of 2.5σ
- IceCube cascades and tracks also demonstrate the Galactic excess \bullet
- Fraction of Galactic events reaches several tens of percent at E>200 TeV \bullet disagreeing many theoretical predictions

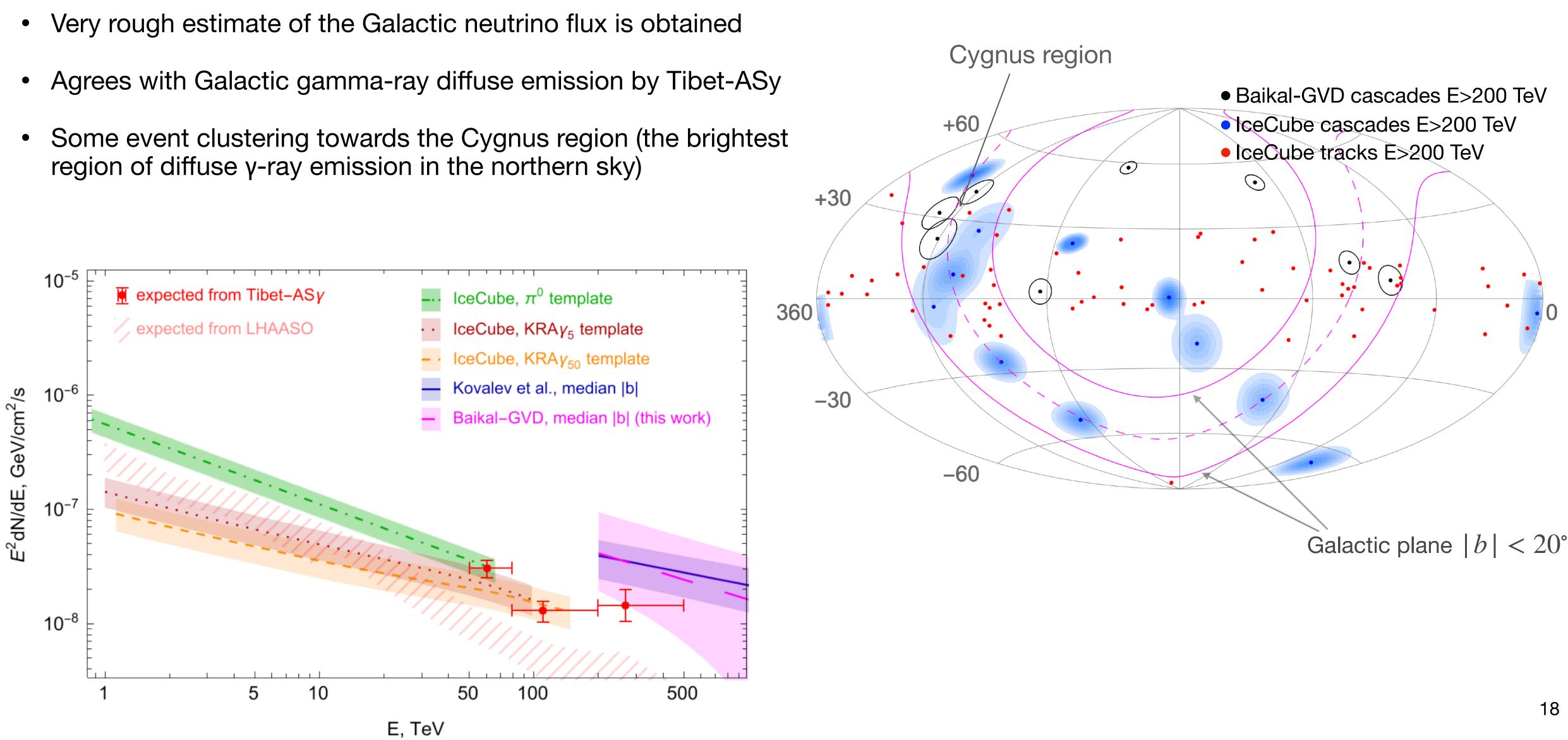
Sample	t
	obs
Baikal-GVD cascades	1
IceCube cascades	1
combined cascades	1
IceCube tracks	2
all cascades+tracks	2





Galactic Neutrinos with the Highest Energies

- region of diffuse γ -ray emission in the northern sky)





Baikal-GVD Follow-up of IceCube-211208A / PKS 0735+17

- Fast processing system for transient sources has been working since 2021
- Dec 8, 2021 20:02: IceCube "Astrotrack Bronze" net event in the vicinity of the bright blazar PKS 0735+1
- Active state of PKS 0735+17 reported in optical (MASTER), HE gamma-rays (Fermi LAT), X-rays (Swi and radio
- Baikal-GVD found a downward-going (30° above ho ulletcascade-like event 4 hours after the IceCube alert a in 5.3° from it and 4.7° from PKS 0735+17
 - E ≈ 43 TeV
 - PSF 50% (68%) containment radius = 5.5 deg (8.
 - Pre-trial p-value = 0.0044 (2.85 σ) [24 hr, 5.5 deg \bullet
 - Trial factor ~ 40 (total number of IceCube alerts) analysed)

Nutrino	
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orizon) and	TXS 0626+180 PKS 0735+17 VSS 1073948174304 SDSS 1075936.13+1321740 Gb6 10812 ttk50 VVSS 1080204+100639 NVSS 1080204+100639
8.1 deg)	0.0086 0.026 0.06 0.13 0.27 0.54 1.1 2.2 4.3
, cone]	

Astronomy telegram ATeL 15112 was sent https://www.astronomerstelegram.org/?read=15112

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PKS0735+178 poter
Search for neutrino counterpart to the b PKS0735+178 poter associated with Ice 211208A and Baikal 211208A with the KI
211208A with the KI neutrino detectors.
NIR followup of the PKS 0735+178
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observation of a Ge candidate event at t
a gamma-ray flare o
a gamma-ray flare of blazar PKS 0735+17 possible source of
IceCube and Baikal energy neutrinos
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0735+178
Optical view of neut emitter candidate P
+178 Re-brightening of th
object PKS 0735+17 observed by Swift
Fermi-LAT observat flaring activity from
27 and PKS 0735+1
NuSTAR observatio blazar PKS 0735+17
Baikal-GVD observa high-energy neutrin
candidate event fro blazar PKS 0735+17
of the IceCube-2112 neutrino alert from
direction
Swift monitoring of Lac object PKS 073 during a bright state
SRG/eROSITA obse PKS 0735+17
Search for counterp IceCube-211208A w
ANTARES
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600 programs find a radio flare in PKS07
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Event Topologies

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- 90% of recorded track events

Results are coming

Single-cluster cascades

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- Good energy resolution
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Multi-cluster tracks

- Moderately low energy threshold
- Optimal sensitivity to inclined tracks
- Best angular resolution

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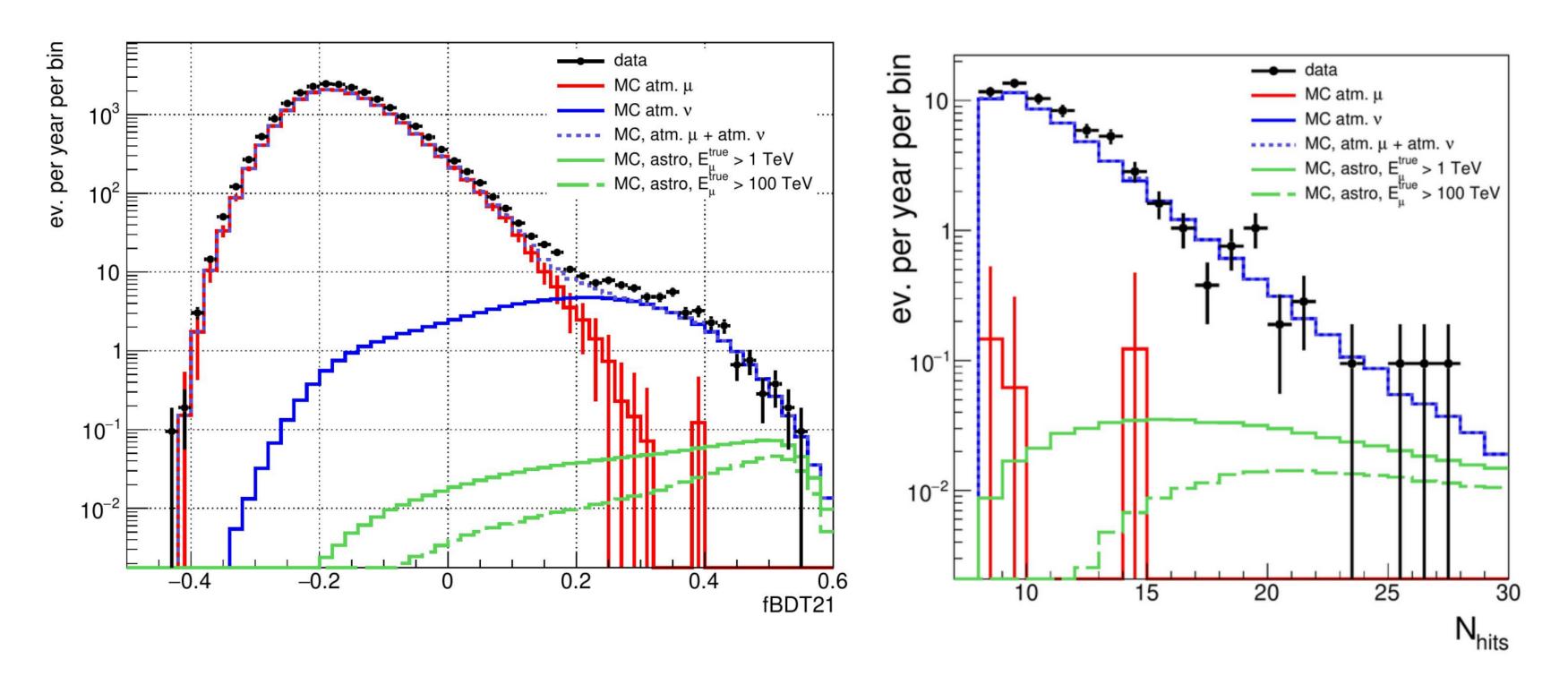
Multi-cluster cascades

- Very high energy threshold
- Excellent energy resolution
- Very rare events





Muon-Track Analysis



- Direction resolution: 0.3-1.0 degrees
- Energy resolution: factor of 3 or 2
- Good agreement with MC expectation

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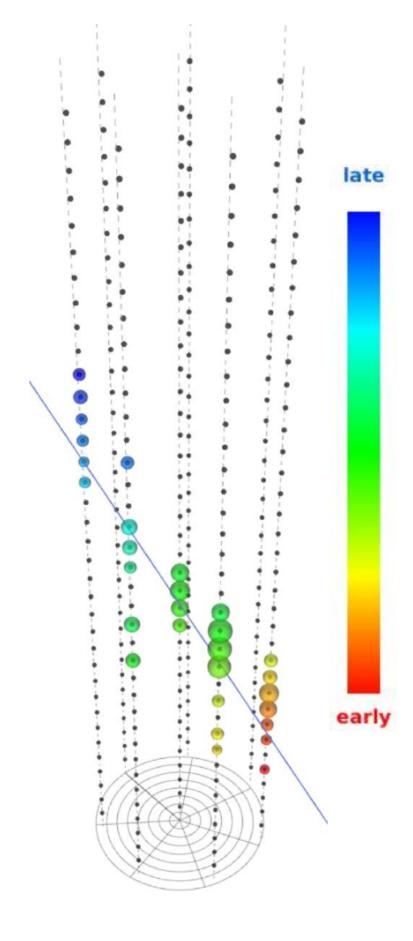
2613 neutrino candidates (2020-2023):

33 events E > 60 TeV

20 events E > 100 TeV

Work in progress...

Neutrino candidate example

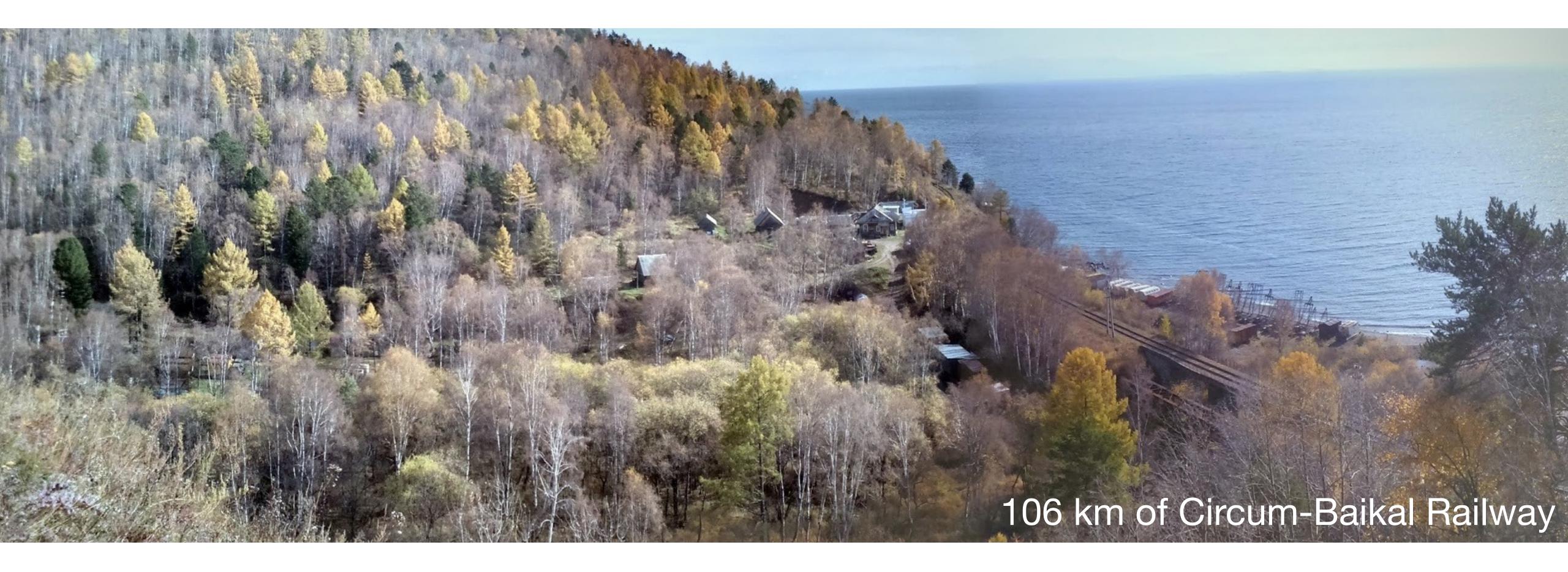


E = 100 TeV



Conclusion

- Baikal-GVD is the largest neutrino telescope in the Northern hemisphere:
 - Volume approaching 0.6 km3 for high-energy cascades
 - Angular resolution better than 1° for tracks
 - Field of view complementary to IceCube
- Nearest plans:
 - Installation of two new clusters + full-scale string for the next-generation project (if the ice conditions) are favorable to us)
- Partially installed telescope produces astrophysical results:
 - Diffuse neutrino flux is confirmed with $> 5\sigma$ significance \bullet
 - Hints of Galactic and extragalactic neutrino sources are accumulating
- The completion of work on the creation of 1 km³ Baikal-GVD detector with ~6000 OM is planned in 2027/2028



Thank you for attention!



Baikal-GVD Collaboration









- Joint Institute for Nuclear Research, Russia
- Institute for Nuclear Research of the Russian Academy of Sciences, Russia \bullet
- Comenius University, Slovakia
- Czech Technical University in Prague, Czech Republic
- Irkutsk State University, Russia
- Skobeltsyn Research Institute of Nuclear Physics, Russia
- Institute of Nuclear Physics ME RK, Kazakhstan
- AO 'LATENA' (Joint Stock Company), Russia
- St. Petersburg State Marine Technical University, Russia

~ 65 physicists and engineers