Directed flow in heavy-ion collisions at high baryon densities

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Exploring Nuclear Phase Diagram



From M. Stephanov, EPJ Web Conf. 314, 00042 (2024)

Hydrodynamics directly addresses Equation of State

However, nonequilibrium prevents direct application of Hydrodynamics

Hydrodynamics versus Kinetics

Hydrodynamics

- ✓ directly addresses Equation of State (EoS)!
- ✓ Phase transition in QGP is accessible through EoS
 However, there are certain problems

Standard hydrodynamics (e.g.,vHLLE)

- Requires local equilibrium, therefore
 Pre-equilibrium (kinetic or parametrised) stage is required
- Is not applicate at very end of collision because of nonequilibrium
 Kinetic afterburner is required
- ✓ At $\sqrt{s_{NN}} \le 6$ GeV the time for hydro turns out to be very short
- ✓ Phase transition into QGP is inaccessible in kinetics
 Exceptions: PHSD, PHQMD, AMPT → only crossover transition

✓ 3FD overcomes these problems

3-fluid dynamics (3FD) model describes initial equilibration

The 3FD approximation simulate the early, nonequilibrium stage of the strongly-interacting matter:

baryon-rich fluids: nucleons of the projectile (p) and the target (t) nuclei.

They are separated in momentum space

 fireball (f) fluid: newly produced particles which dominantly populate the midrapidity region.



momentum along beam

3FD model

Target-like fluid: $\partial_{\mu}J_{t}^{\mu}=0$ $\partial_{\mu}T_{t}^{\mu\nu}=-F_{tp}^{\nu}+F_{ft}^{\nu}$ Leading particles carry bar. chargeexchange/emissionProjectile-like fluid: $\partial_{\mu}J_{p}^{\mu}=0$, $\partial_{\mu}T_{p}^{\mu\nu}=-F_{pt}^{\nu}+F_{fp}^{\nu}$ Fireball fluid: $J_{f}^{\mu}=0$, $\partial_{\mu}T_{f}^{\mu\nu}=F_{pt}^{\nu}+F_{tp}^{\nu}-F_{fp}^{\nu}-F_{ft}^{\nu}$ Baryon-free fluidSource termExchangeThe source term is delayed due to a formation time τ

Total energy-momentum conservation: $\partial_{\mu}(T_{p}^{\mu\nu} + T_{t}^{\mu\nu} + T_{f}^{\mu\nu}) = 0$

Physical Input

- ✓ Equation of State (EoS)
- ✓ Friction
- ✓ Freeze-out energy density ϵ_{frz} = 0.4 GeV/fm³

3FD: YI, Russkikh, Toneev, PRC 73, 044904 (2006)

EoS:

 hadronic EoS (no phase transition)

Mishustin, Russkikh and Satarov,

Sov. J. Nucl. Phys. 54, 260 (1991)

- hadronic+QGP with 1storder Phase Transition (1PT)
- hadronic+QGP with crossover

Khvorostukhin, Skokov, Toneev, Redlich, EPJ C48, 531 (2006)

- Friction in hadronic phase is estimated [Satarov, Yad. Fiz. 52, 412 (1990)].
- Friction in QGP is parametrized [Y.I. PRC 87, 064904 (2013)]

Afterburner (kinetic evolution after hydro)

Afterburner by means of event generator THESEUS

(Three-fluid Hydrodynamics-based Event Simulator Extended by UrQMD final State interactions)

THESEUS-v1, Batyuk, Blaschke, Bleicher, YI, Karpenko, Merts, Nahrgang, Petersen and Rogachevsky, PRC 94, 044917 (2016) **THESEUS-v2** updated, Kozhevnikova, YI, Karpenko, Blaschke and Rogachevsky, PRC 103, 044905 (2021)

THESEUS:

- transformation of 3FD output (fields of T, μ_B and μ_S) into set of particles (particlization)
- Post-freeze-out kinetic evolution by means of UrQMD (afterburner)

Afterburner is important at moderately relativistic energies



⁽prepared by M. Kozhevnirjva)

3DF+THESEUS-v2 can predict almost all except for few things:

- $\checkmark\,$ Fluctuations, because there are no fluctuations in the initial state
- Collisions of light nuclei and very peripheral collisions because hydro is inapplicable to few-particles systems
- ✓ Collisions at $\sqrt{s_{NN}} \ge 40$ GeV because of numerical problems

Equation of State

Pressure vs baryon density



Phase transition → EoS softening (in dense baryon matter)

Hadronic phase: Friction was estimated by Satarov [Sov.J.Nucl.Phys. 52, 264 (1990)] QGP phase: Phenomenological friction fitted to reproduce baryon stopping [YI, PRC 87 (2013), 064904]



Dynamical trajectories of matter in central region

Temperature vs baryon density

✓ Turning points at 3 GeV does not reach the QGP.

✓ 4.5-GeV trajectories fall well into crossover QGP region and even enter the the 1PT mixed phase.

Speed of sound vs baryon density

Evolution starts from instants (indicated by \bigstar) when matter is equilibrated.





 $c_s = (dP/d\epsilon)$ along trajectory

softest-point region is probed at 4.5 GeV within 1PT scenario.

How are these features manifested in directed flow?

Directed flow (v₁)

 $E\frac{d^3N}{d^3p} = \frac{1}{2\pi} \frac{d^2N}{p_T dp_T dy} (1 + \sum_{n=1}^{\infty} 2v_n \cos(n(\phi - \Psi_{\rm RP})))$ $\varphi = \text{azimuthal angle}$

 v_1 = directed flow v_2 = elliptic flow

- Proton directed flow at √sNN < 7 GeV is well reproduced in terms of (P_x) but not in terms of v₁.
- v₁ is very sensitive to onset of deconfinement transition and softest point
- Onset of deconfinement transition happens in the NICA energy range.
- ✓ Crossover EoS is preferable.



Directed flow at 3 GeV



- ✓ v_1 in midrapidity does not depend on EoS → no QGP transition
- ✓ *p* and \land v₁ are well reproduced with all EoS's → **no QGP tr.**

✓ Afterburner is important for π , Λ and K^{-1} :

shadowing of participants by spectators

Description of this shadowing still is not perfect

✓ K⁺ is not equilibrated or requires early freeze-out

v₁ of light nuclei and hypernuclei at 3 GeV



3FD-THESEUS: Kozhevnikova and YI Light nuclei: PRC109 (2024), 014913 Light hypernuclei: PRC 109 (2024), 034901 Hadrons: PRC 110 (2024), 014907

Data: STAR, PRC 110 (2024) 5, 054911; PLB 827, 136941 (2022); PRL 130, 212301 (2023); PRC 103, 034908 (2021); PLB 827, 137003 (2022)

Directed flow at 4.5 GeV



- \checkmark v₁ does depend on EoS \rightarrow QGP transition
- \checkmark p, π and K⁰ are well reproduced with crossover EoS. Afterburner is important, especially for π , because of shadowing of participants by spectators
- ✓ Different particles probe EoS in different parts and at different stages of the colliding system

Proton v₁ evolution between 2.7 and 4.5 GeV

✓ Note that $v_1 = \langle p_x / p_T \rangle$

✓ v₁ does not depend on EoS at 2.7 GeV
 → no QGP transition

✓ Crossover EoS gradually becomes preferable at $\sqrt{s_{NN}}$ > 3 GeV

 Hadronic and 1PT scenarios give almost identical v₁



YI and Soldatov, PRC

Consensus: QGP onset happens between $\sqrt{s_{NN}} = 3$ and 4.5 GeV

Model	Potential interaction	QGP phase	Transition to QGP	type
UrQMD	optional	No	Above $4n_0^*$	QMD
QGSM	no	No	No	Kin. Eq.
PHSD	Mean-field (optional)	yes	E > 0.5 GeV/fm ³	Kin. Eq.
PHQMD	2-body potential	yes	E > 0.5 GeV/fm ³	QMD
AMPT	Mean-field (optional) Various versions	yes	lower than $\sim 2.5n_0$ 35n_0 4.35.7n_0	Kin. Eq.
SMASH	Mean-field	no	3-4n ₀ *	Kin. Eq.
JAM	Mean-field (optional)	no	Above $3-4n_0^*$	Kin. Eq.
3FD (THESEUS)	Non-gas EoS	yes	Above 45n ₀	hydro

* "Softenning of EoS" is required

Summary

- \checkmark v₁ is one of most sensitive observable to the QGP transition
- ✓ v_1 data indicate onset of QGP transition at $\sqrt{s_{NN}} \ge 4$ GeV
- ✓ Onset of QGP transition happens at $n \ge 4$ --5n₀ and T ≈ 150 MeV
- v₁ of various particles provide information on EoS in different parts and at different stages of the colliding system
- information on EoS is not always directly accessible because of strong influence of afterburner stage or insufficient equilibration.

✓ Crossover EoS gives the best overall description of the data

- ✓ This EoS crossover may mimic a weak 1PT, as the critical point is likely located in the upper part of NICA energy range [STAR, 2504.00817]
- The used 1PT EoS corresponds to strong 1PT



Thank you

for your attention!

Backup



v₁ of medium modified Kaons



$$E(\mathbf{p}) = \left[m_K^2 + \mathbf{p}^2 - \frac{\Sigma_{KN}}{f_K^2}\rho + \left(\frac{3}{8}\frac{n}{f_K^2}\right)^2\right]^{1/2} \pm \frac{3}{8}\frac{n}{f_K^2}$$

$$n = \sum_{B} \langle \bar{B} \gamma^{0} B \rangle,$$

$$\rho = \sum_{B} \langle \bar{B} B \rangle$$

upper(lower) sign refers to K(anti-K). n and p are proper baryon and scalar baryon densities, respectively.

 v_1 of K and anti-K are changed, while v_1 of K⁰_s remains almost unchanged

Initial Equilibration

vHLLE: **UrQMD (till the equilibration)** \rightarrow **hydro** \rightarrow **UrQMD (afterburner)** This is a typical scheme for any 1-fluid hydro model



- ✓ Time of initial equilibration is long at lower NICA energies.
- ✓ Therefore, the time for hydro turns out to be short
- ✓ Therefore, vHLLE applicability is $\sqrt{s_{NN}} \ge 6$ GeV
- ✓ Therefore, starts already in the QGP phase because of Consensus: QGP is present at $√s_{NN} = 4.5$ GeV and

no QGP at $\sqrt{s_{NN}} = 3$ GeV in Au+Au collisions

Kinetics: advantages and disadvantages

Advantages:

Kinetic and QMD models

- ✓ directly address nonequilibrium
- ✓ can treat non-gas systems (mean fields, 2-body potentials)
- ✓ naturally treat fluctuations (CEP search)

Disadvantages:

- Huge amount of input data (often experimentally unknown)
- ➢ In practice, kinetics → only binary collisions/interactions Approximation of binary collisions is bad in dense system!
- Phase transition into QGP is inaccessible in kinetics, as a rule

Exceptions: PHSD, PHQMD, AMPT - crossover transition

Hydrodynamics versus Kinetics

Hydrodynamics

✓ takes into account any multi-particle interactions

✓ directly addresses Equation of State (EoS)!

✓ Phase transition in QGP is accessible through EoS However, there are certain problems

Standard hydrodynamics (vHLLE)

 Requires local equilibrium, therefore it cannot be applied from the very beginning of the collision
 Pre-hydro (kinetic or parametrised) stage is required
 3FD overcomes this problem (see below)

 Is not applicate at the very end of the collision because of nonequlibrium

Kinetic afterburner is required



illustration:

Time evolution of energy density in reaction plane