

Strange hadron production in Au+Au collisions at $\sqrt{s_{NN}} = 3 - 27$ GeV with UrQMD

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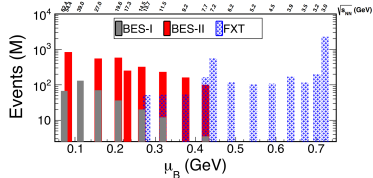
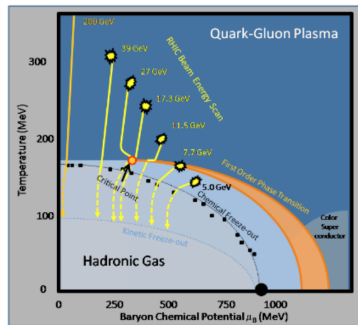
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Overview

- Motivation
- Particle reconstruction at STAR (work in progress)
- Strange hadrons production
- Near threshold kinematics
- Summary

Motivation

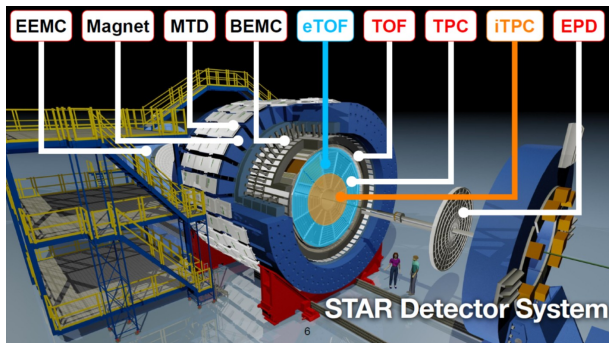
- Strange hadrons are suggested as exceptional probe for identifying phase boundary and onset of deconfinement
- Strange baryons-to-meson and antibaryons-to-baryons ratios may give insights on hadronization mechanism and baryon stopping
- UrQMD model comparison to STAR data



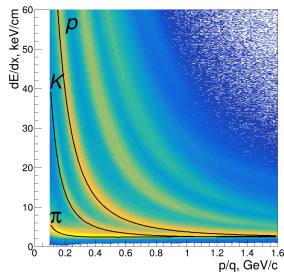
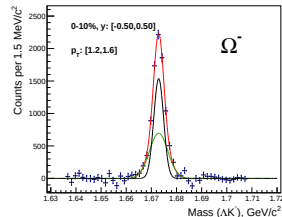
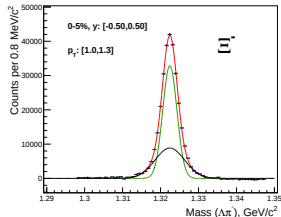
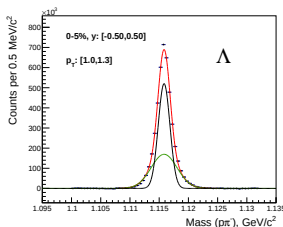
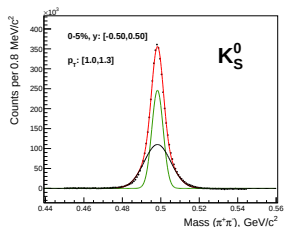
STAR Experiment

TPC (Time Projection Chamber):

- ✚ Tracking and particle identification by dE/dx
- ✚ Magnetic field: 0.5 Tesla
- ✚ 2π azimuthal angle acceptance
- ✚ Pseudorapidity $|\eta| < 1$ acceptance
- ✚ $p_T > 0.1 \text{ GeV}/c^2$



Particle reconstruction at STAR



- ▶ Particle identification with $n\sigma$ on dE/dx
- ▶ Cuts on η , p_T , nHitsFit, nHitsdEdx are applied on π , K , p
- ▶ Topological cuts are applied on strange particles and daughter particles

$$K_S^0 \rightarrow \pi^+ + \pi^- \quad (69.20\%)$$

$$\Xi^- (\Xi^+) \rightarrow \Lambda (\bar{\Lambda}) + \pi^- (\pi^+) \quad (99.887\%)$$

$$\Lambda (\bar{\Lambda}) \rightarrow p (\bar{p}) + \pi^- (\pi^+) \quad (63.9\%)$$

$$\Omega^- (\Omega^+) \rightarrow \Lambda (\bar{\Lambda}) + K^- (K^+) \quad (67.8\%)$$

UrQMD

UrQMD is transport model that solves a relativistic Boltzmann-type transport equation by tracking hadrons and resonances, incorporating string excitation/fragmentation at high energies.

Initialization via Gaussian wave packets:

$$\varphi_j(x_j, t) = \left(\frac{2\alpha}{\pi}\right)^{3/4} \exp \left[-\alpha(x_j - r_j(t))^2 + \frac{i}{h} p_j(t) \cdot x_j \right]$$

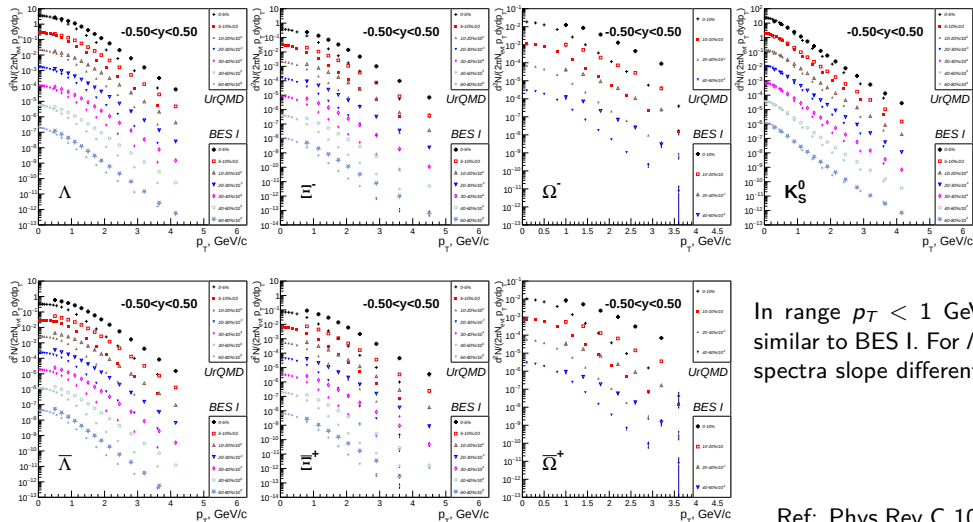
Collision condition via impact parameter and total cross section:

$$d_{\text{trans}} \leq \sqrt{\frac{\sigma_{\text{tot}}}{\pi}}, \quad \sigma_{\text{tot}} = \sigma(\sqrt{s}, \text{particle types})$$

In this work UrQMD 3.4 (CASCADE mode, EoS = 0) was used.

Ref: arXiv:nucl-th/9803035v2

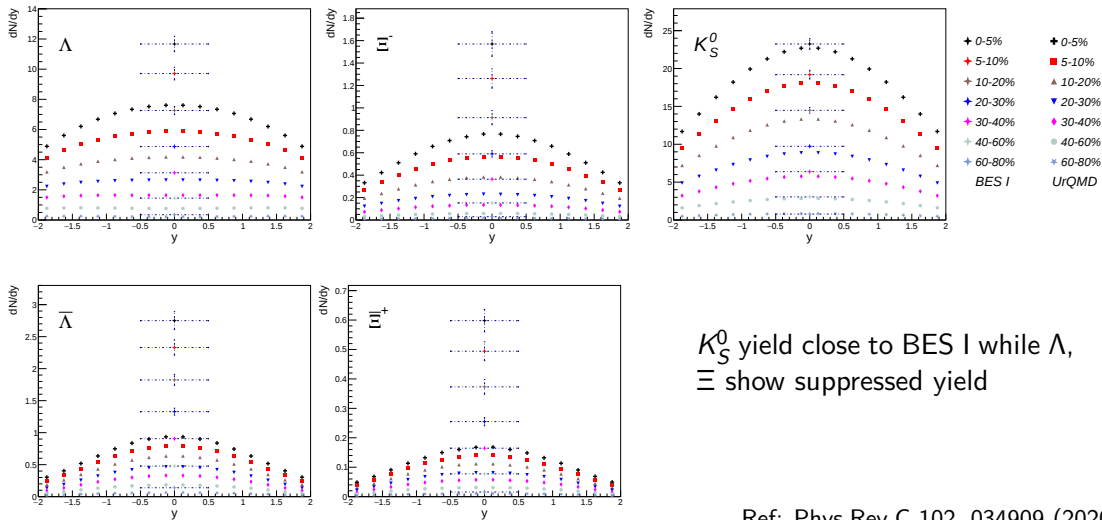
p_T Spectra for 27 GeV



In range $p_T < 1$ GeV/c K_S^0 spectra similar to BES I. For Λ , Ξ , Ω UrQMD spectra slope different to BES I

Ref: Phys.Rev.C 102, 034909 (2020)

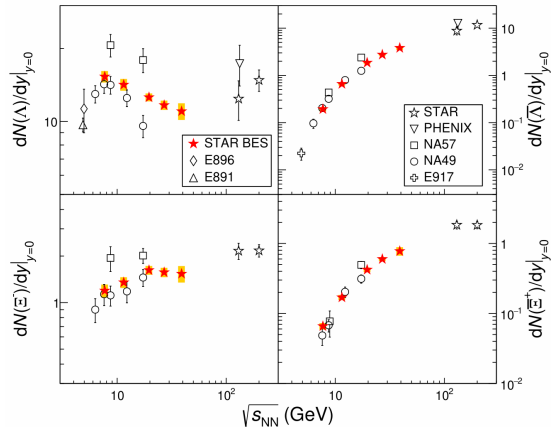
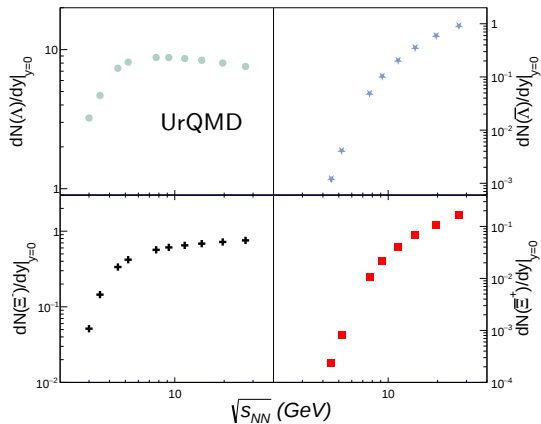
Particle Yield at 27 GeV



K_S^0 yield close to BES I while Λ , Ξ show suppressed yield

Ref: Phys.Rev.C 102, 034909 (2020)

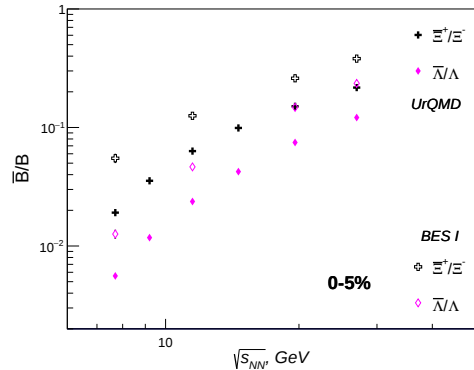
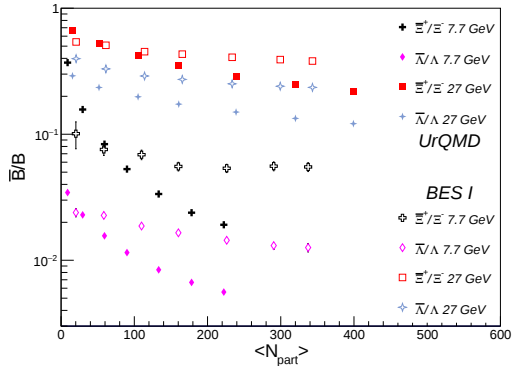
Energy dependence of the yield



$$|y_{cm}| < 0.5$$

Ref: Phys.Rev.C 102, 034909 (2020)

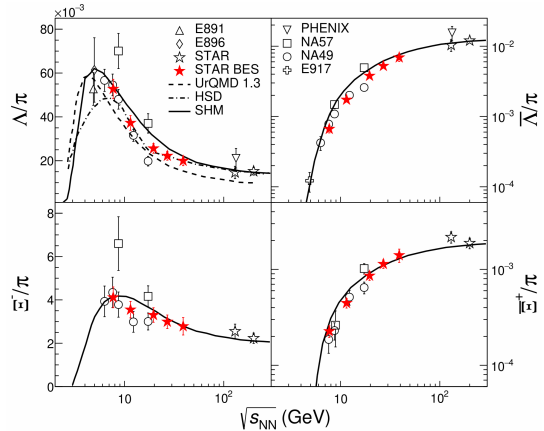
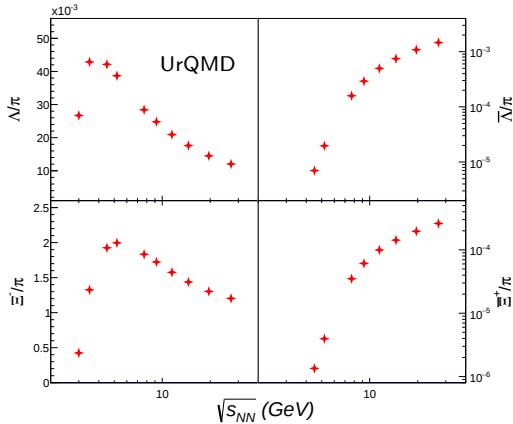
Antibaryon-to-baryon ratio



UrQMD ratios decrease from peripheral to central collisions. Compared to BES I, in central collisions UrQMD underestimate $\bar{B}/B$ ratio and overestimate in peripheral collisions.

Ref: Phys.Rev.C 102, 034909 (2020)

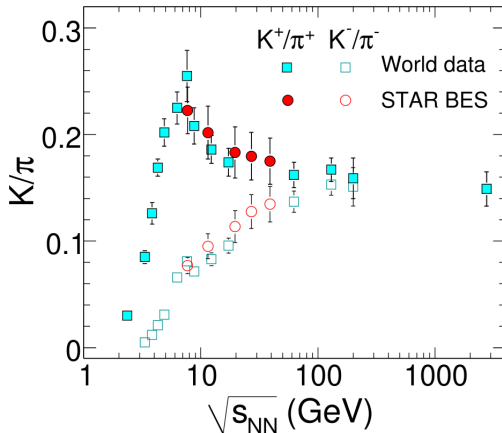
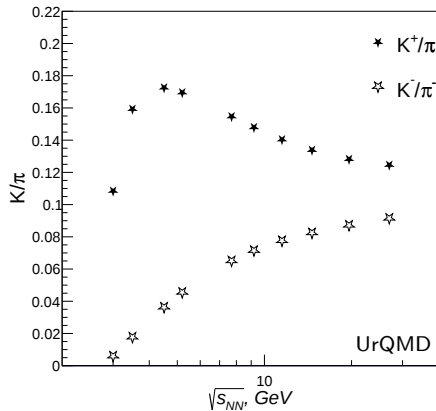
Strangeness production I



While the general trend of the dependencies is consistent, the absolute values of the ratios are significantly underestimated in the UrQMD model

Ref: Phys. Rev. C 102, 034909 (2020)

Strangeness production II



Horn-like effect can not be explained by hadron-hadron interactions

Ref: Phys. Rev. C 96, 044904 (2017)

Near threshold kinematics I

$$p + p \rightarrow S(K_S^0, \Lambda, \Xi, \Omega) + X$$

Since $E_{thresh}^* = M_S + M_X$:

$$\left. \begin{array}{l} p + p \rightarrow p + K_S^0 + \Sigma^+ \\ p + p \rightarrow p + \Xi^- + 2K^+ \end{array} \right\} \Rightarrow \begin{cases} E_{thresh, K_S^0}^* \approx 2.63 \text{ GeV} \\ E_{thresh, \Xi^-}^* \approx 3.25 \text{ GeV} \end{cases}$$

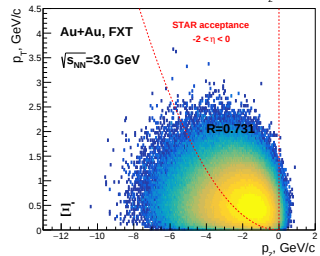
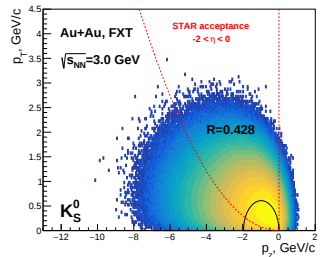
In center-mass system:

$$p_{cm} = \frac{(s - m_S^2 - m_X^2)^2}{2\sqrt{s}}, \quad \beta_{cm} = \frac{p_p}{E_p + m_p}, \quad \gamma_{cm} = \frac{s - m_p^2 + m_p^2}{2m_p\sqrt{s}}$$

Since $\left(\frac{p_T^*}{p^*}\right)^2 + \left(\frac{p_z^*}{p^*}\right)^2 = 1$, using Lorentz boost one obtain:

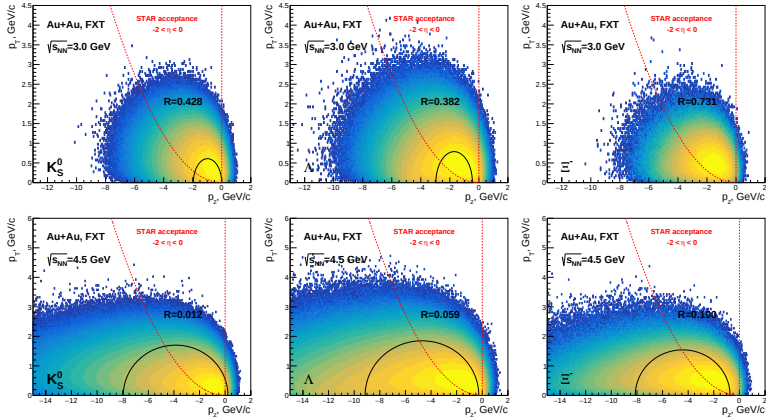
$$\left(\frac{p_T^*}{p^*}\right)^2 + \left(\frac{p_z - \gamma_{cm}\beta_{cm}E_S^*}{\gamma_{cm}p^*}\right)^2 = 1$$

R - ratio of particle inside cumulative region inside STAR acceptance to all particles



Ref: Strokovsky E.A., ISBN: 978-5-317-04607-1

Near threshold kinematics II



A large number of strange particles are produced in the cumulative region. At energies below $\approx 4-5$ GeV it became possible to measure collective effects of multi-nucleon correlations.

Summary

- UrQMD underestimates the yield of strange particles (except K_S^0)
- Particle ratios shows the same trends as STAR data
- At low energies, a significant fraction of particles is produced in the cumulative region, which gives sufficient statistics for measuring cumulative effects.