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## Deconvolution in IceCube

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October 4th 2018

## Outline

- Challenges in Deconvolution
- Selected Deconvolution Algorithms
- Some results on atmospheric neutrino spectra
- Dortmund Spectrum Estimation Algorithm

## Deconvolution in a Nutshell

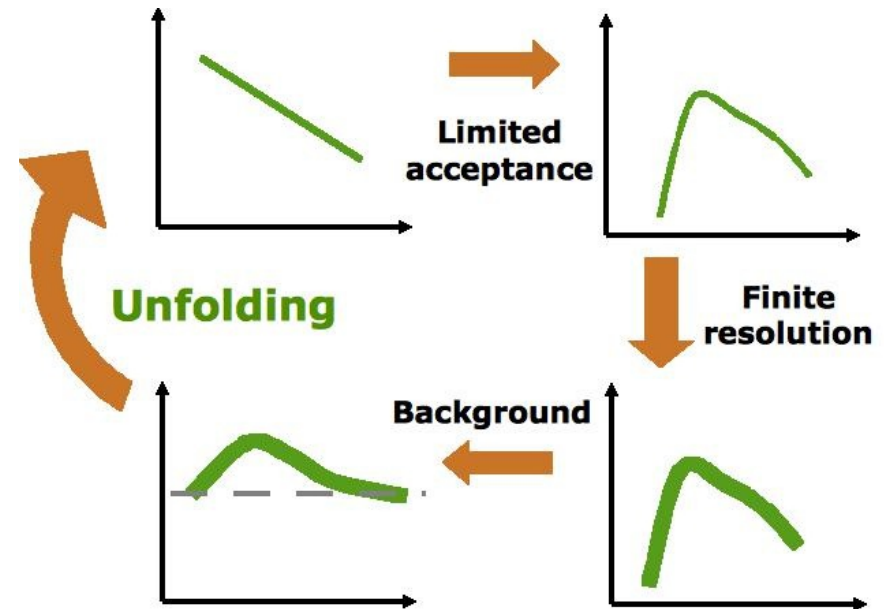
$$\underbrace{\frac{dN_\mu}{dE_\mu}}_{\text{Muon energy spectrum}} = \int_{E_\mu}^{\infty} \underbrace{\left(\frac{dN_\nu}{dE_\nu}\right)}_{\text{Neutrino energy spectrum}} \underbrace{\left(\frac{dP(E_\nu)}{dE_\mu}\right)}_{\text{Physics of neutrino interaction}} dE_\nu$$

Muon energy spectrum

Neutrino energy spectrum

Physics of neutrino interaction

- Production of charged lepton in neutrino interaction is governed by stochastic processes
- Additional smearing, due to several detector effects



Mathematically: Fredholm integral equation of the first kind:

$$g(y) = \int_{E_{min}}^{E_{max}} A(E, y) f(E) dE$$

## Deconvolution in a Nutshell

Physicists generally happy with a discrete result

$$g(y) = \int_{E_{min}}^{E_{max}} A(E, y) f(E) dE$$

Integral equation



Discretizing operation

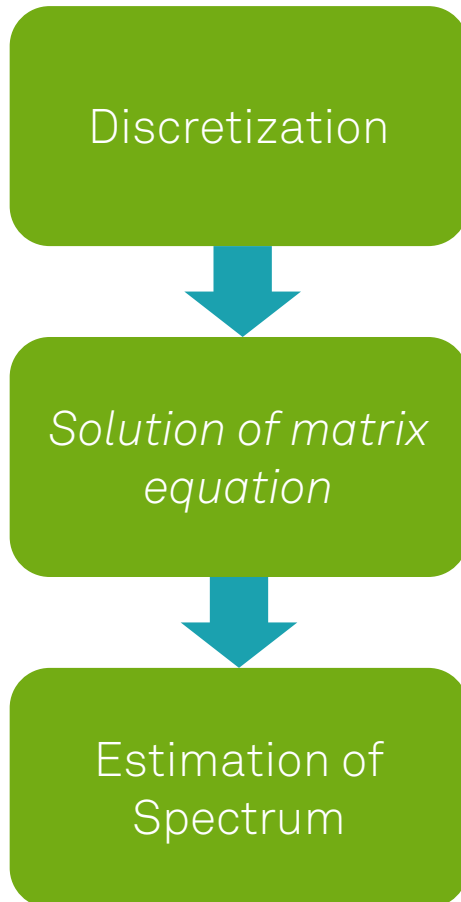
$$\vec{g}(y) = A(E, y) \vec{f}(E) dE$$

Matrix equation

$A(E, y)$  generally not known analytically.

Replaced my probability matrix, obtained from Monte Carlo simulations.

## Deconvolution in a Nutshell



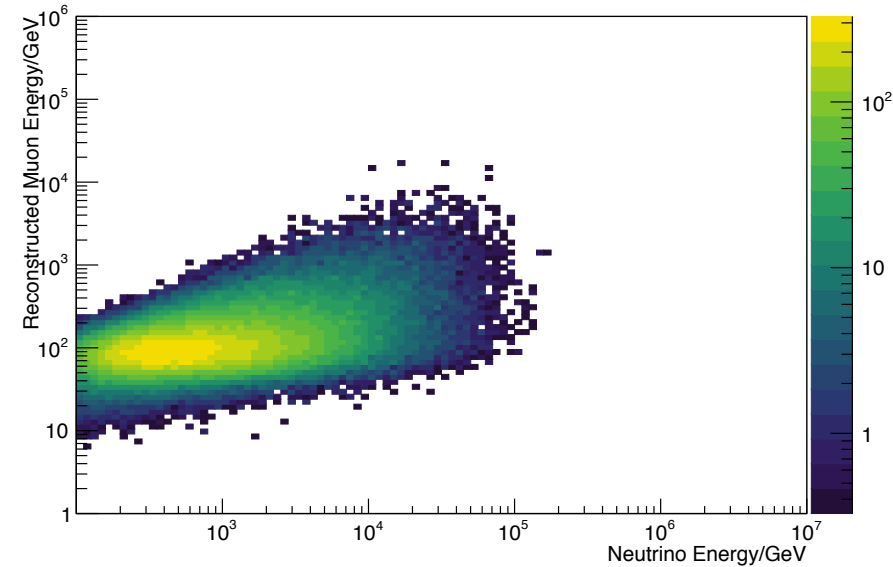
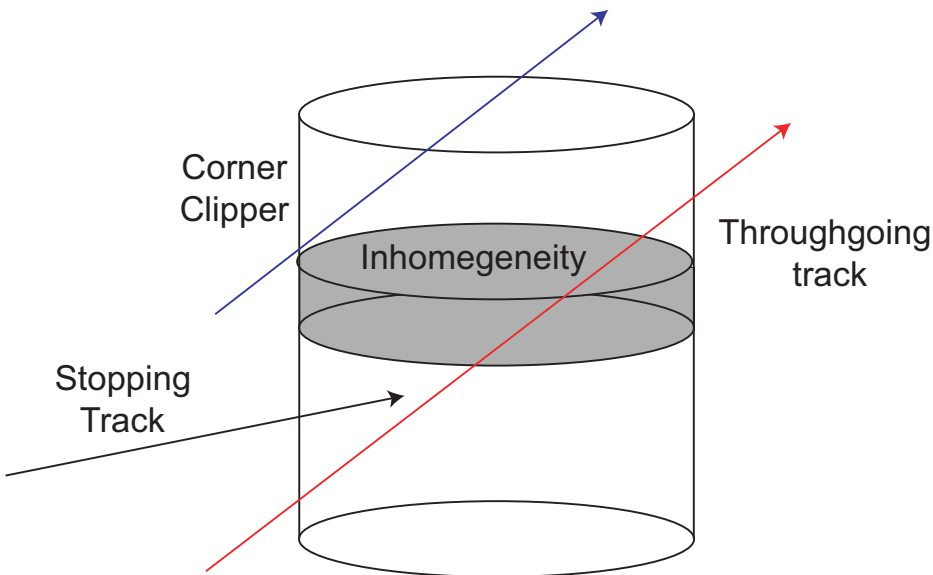
Direct inversion of  $A$  not feasible or leads to oscillating solutions.

Regularisation is required,

- Assumes smoothness of  $f$
- Adds certain amount of bias
- In most cases small second derivative

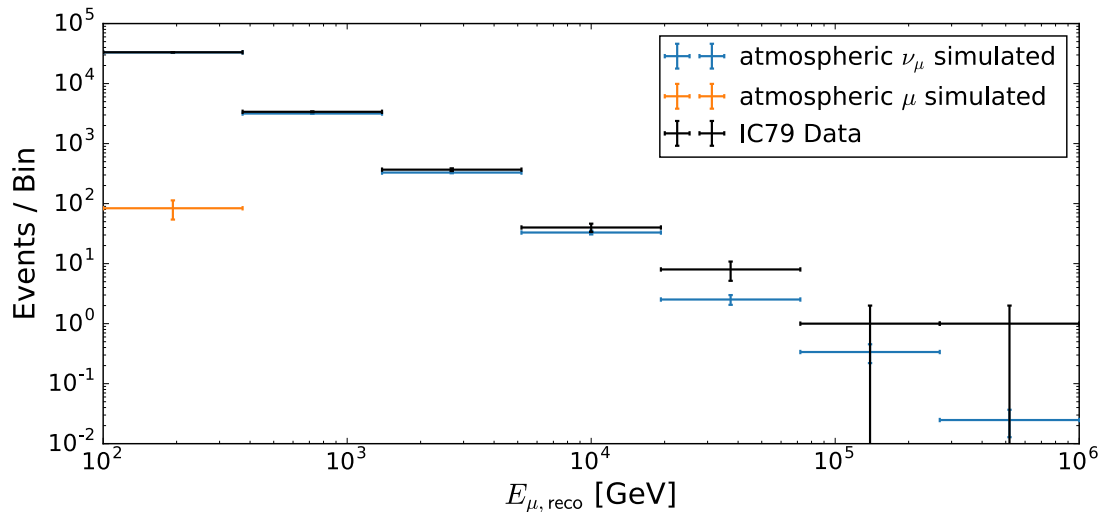
## Additional Challenges

- Amount of smearing
- Small statistics at high energies
- Muon energy only an estimate
- Estimation of systematic uncertainties



Neutrinos with exactly the same energy can create very different signatures in the detector.

## Before Unfolding: Purity Requirements



Purity generally above 99%.

**Even more important:** Make sure there are no muons in bins with small statistics.

## Selected Deconvolution Algorithms

- Forward Folding
- Iterative Bayesian Unfolding D'Agostini, 2010
- Dortmund Spectrum Estimation Algorithm (DSEA), machine learning-based  
Ruhe et al., 2016
- Singular Value Decomposition (SVD)
- TRUEE

Based on RUN-Algorithm,  
uses B-splines.

Minimizes a log-likelihood.

Milke et al., 2013

Obtain inverse matrix  $A^{-1}$  via  
factorization of the form  $A = U \cdot S \cdot V^T$

Challenge: Small eigenvalues in  $S$   
enhance statistically insignificant  
contributions. Spectrum is distorted.

Höcker and Kartvelishvili, 1996

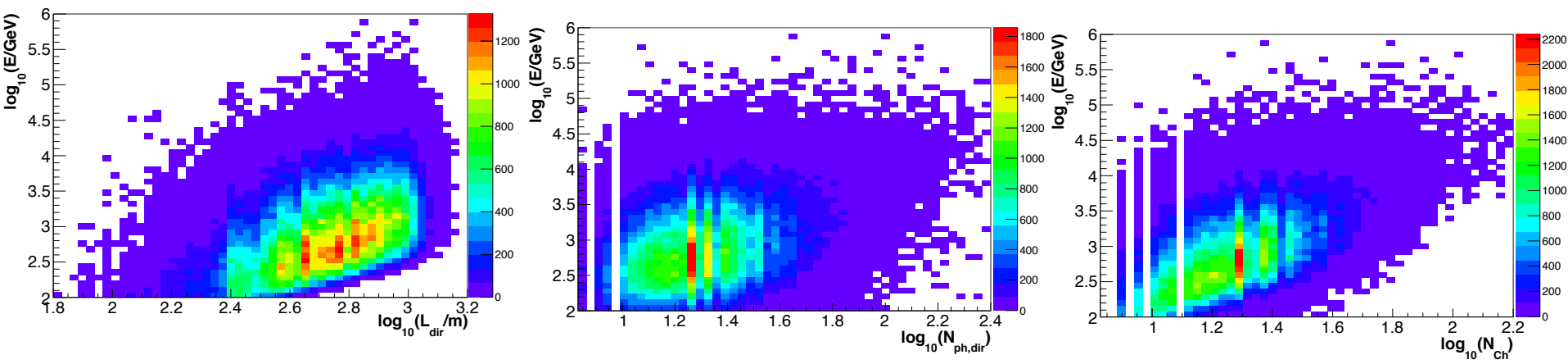


## IC-59 NuMu Unfolding

Variables utilized:

- Length of track
- Number of unscattered photons
- Number responding DOMs

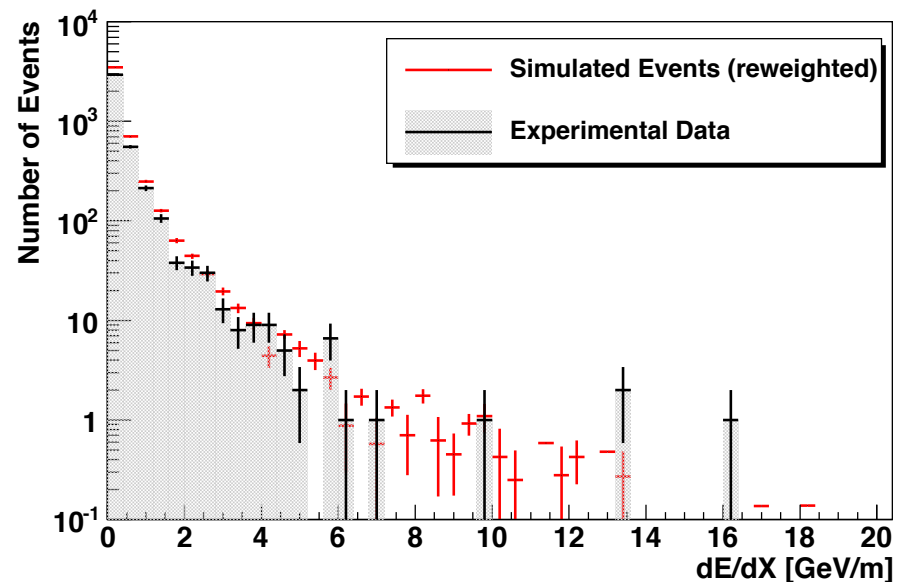
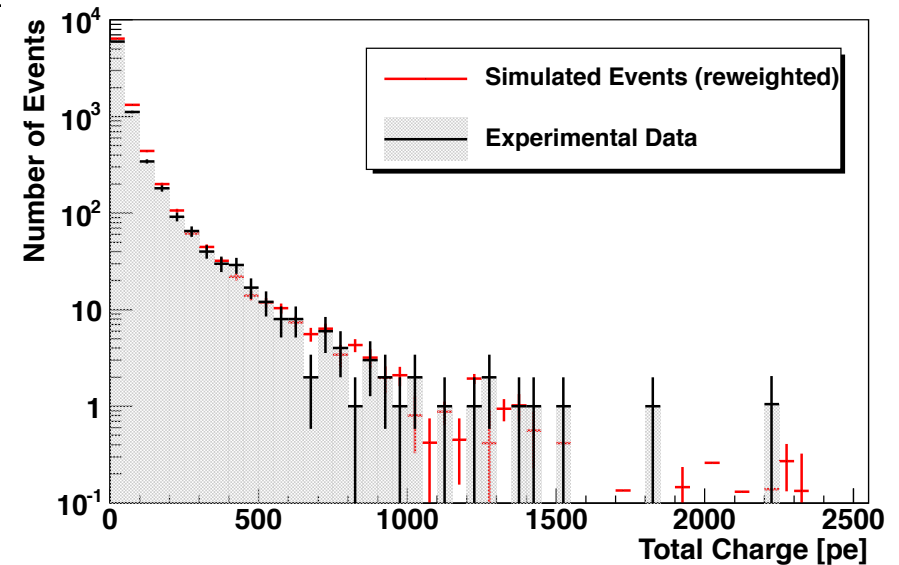
Not the beste energy estimators individually, but the combination of variables with the best unfolding result.



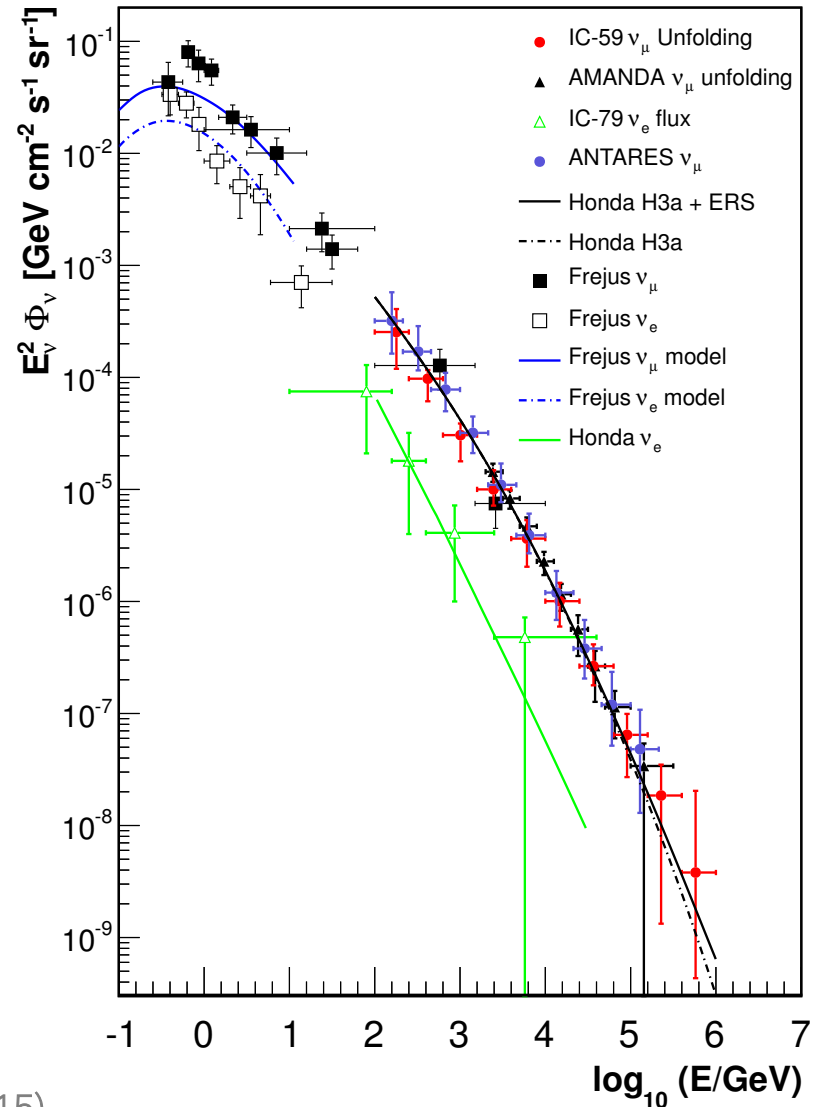
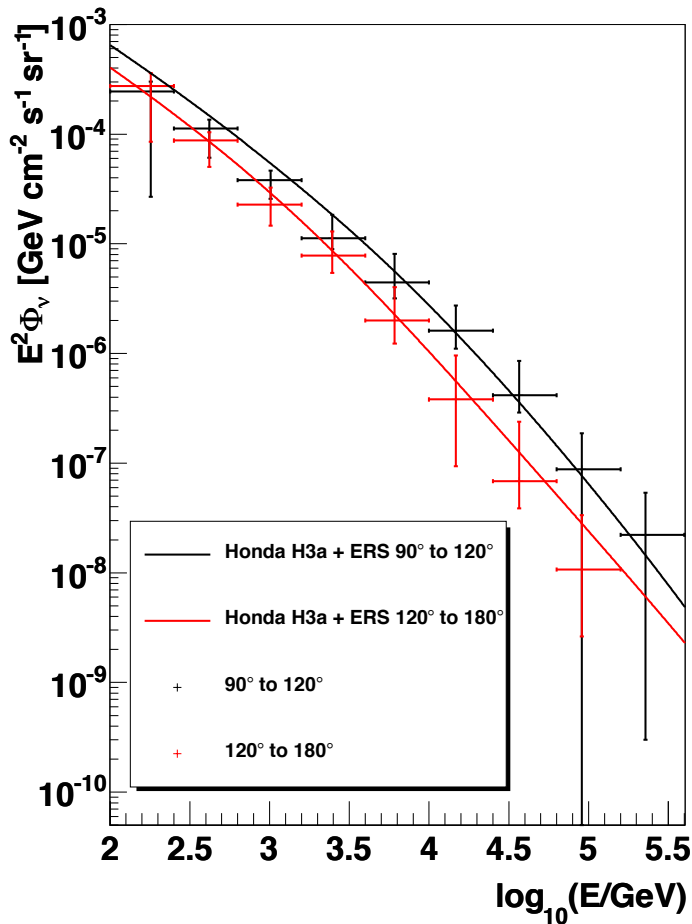
## Sanity Checks in TRUEE

Simulations are reweighted according to obtained spectrum and compared to experimental data.

Especially useful for energy dependent quantities not used in the unfolding.



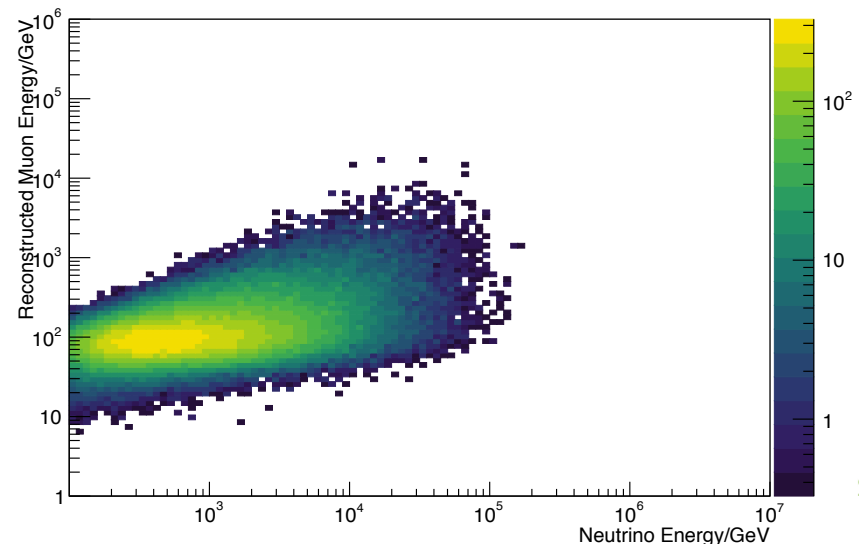
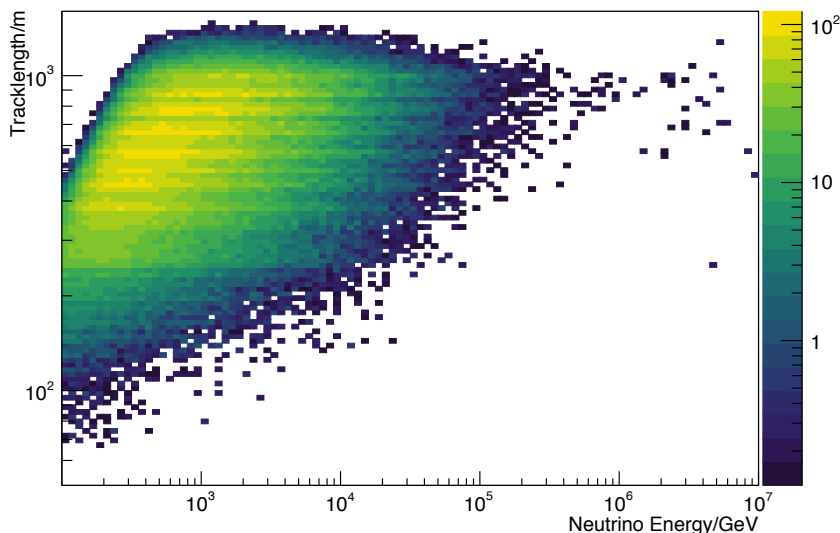
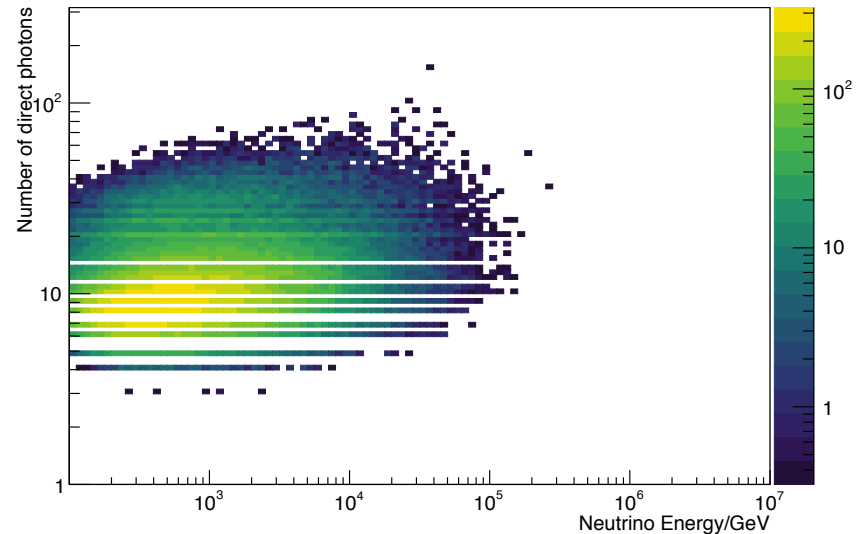
## IC-59 NuMu Unfolding



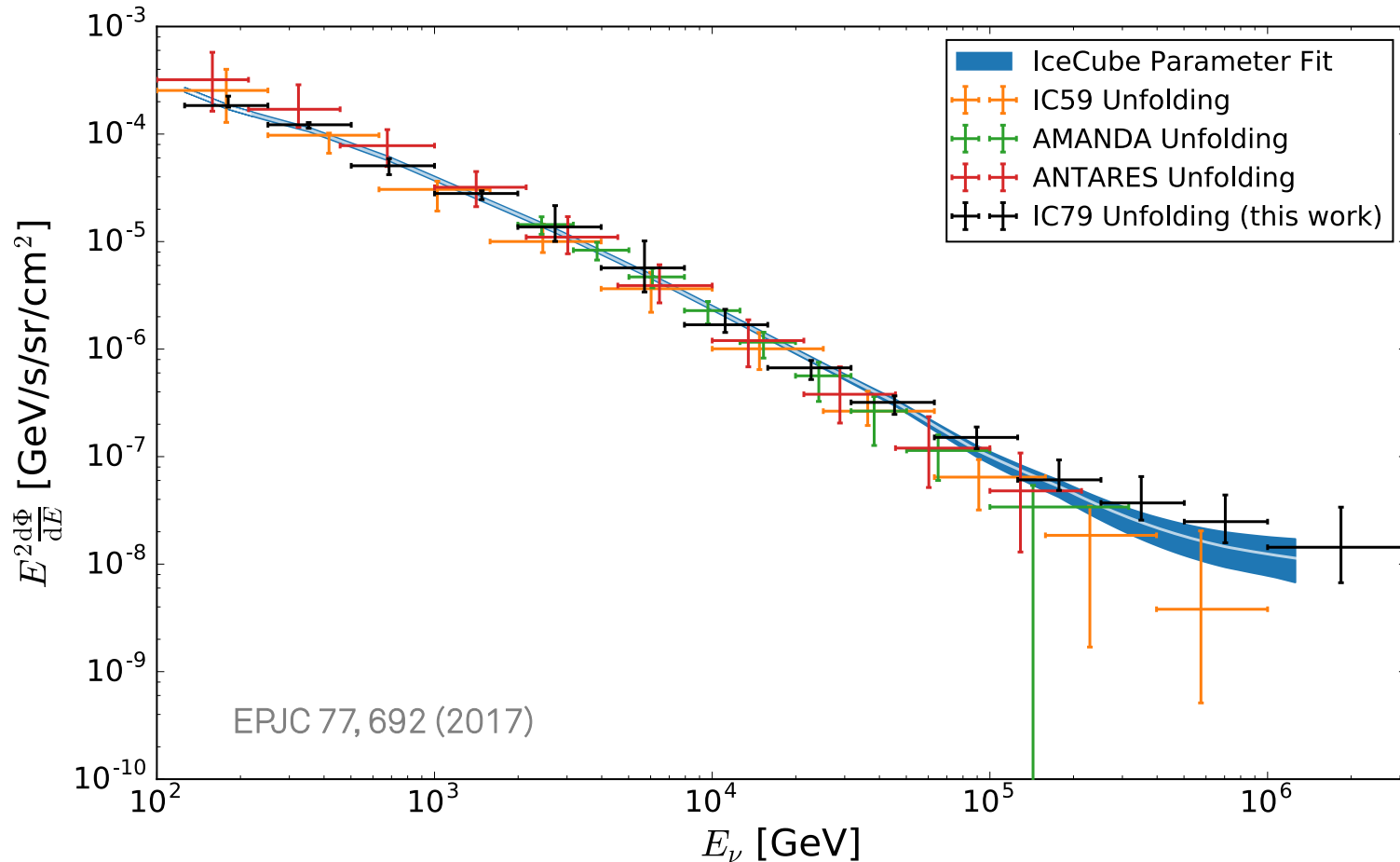
## IC-79 NuMu Unfolding (TRUEE)

Three input parameters (Tracklength, Number of unscattered photons, Reconstructed Muon energy).

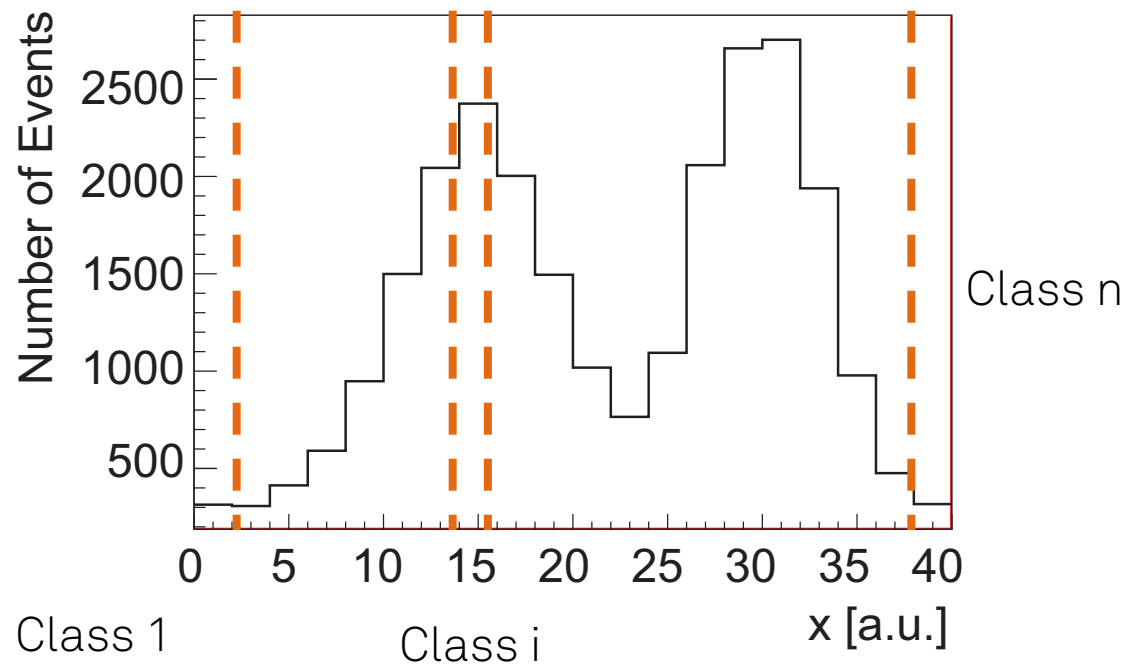
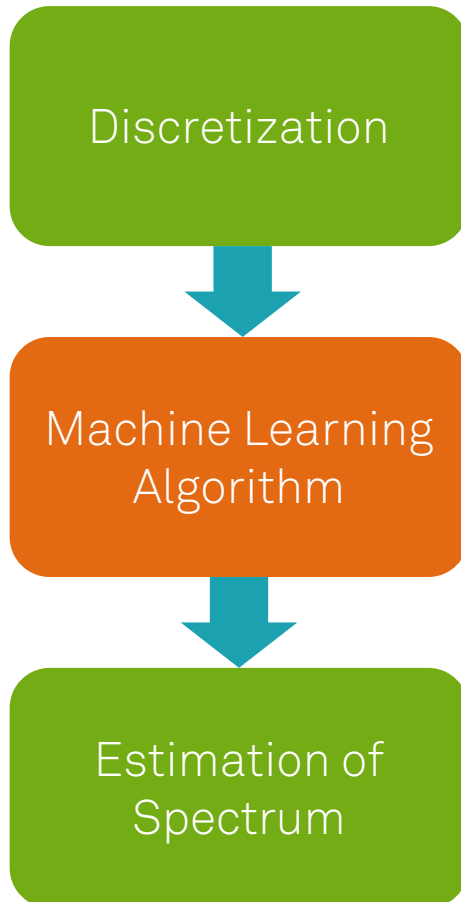
Better estimators available, but above combination gave the best unfolding result.



## IC-79 NuMu Unfolding (TRUEE)

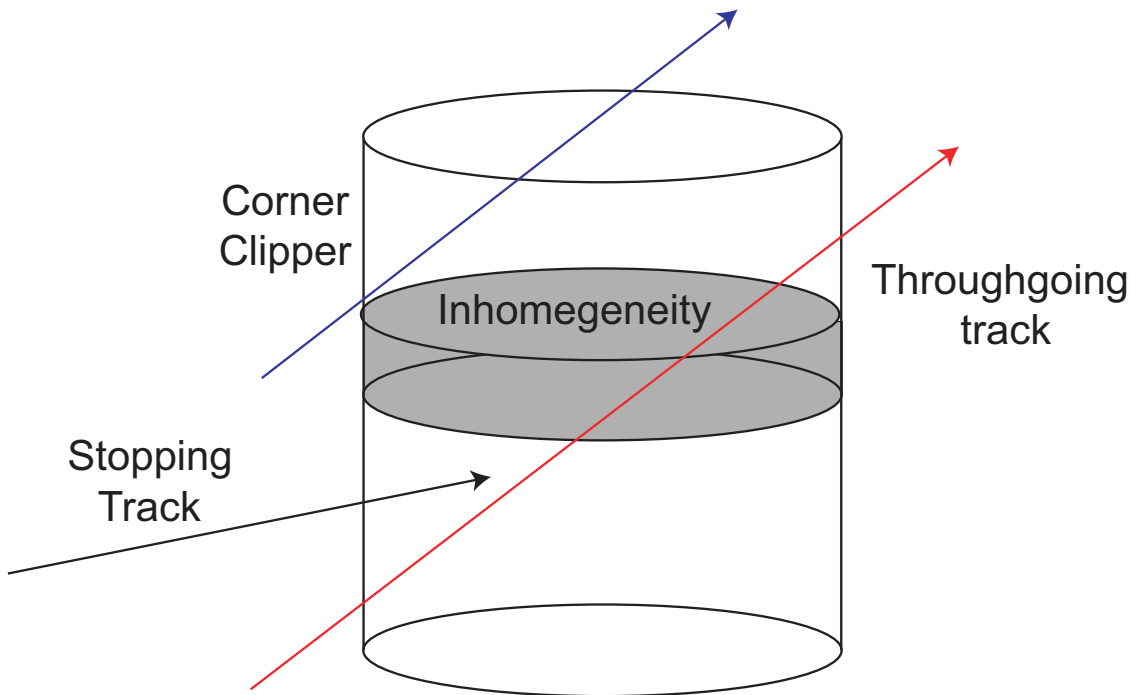


## Dortmund Spectrum Estimation Algorithm (DSEA)



Inverse problem is transferred into multinominal classification problem.

## Motivation for DSEA

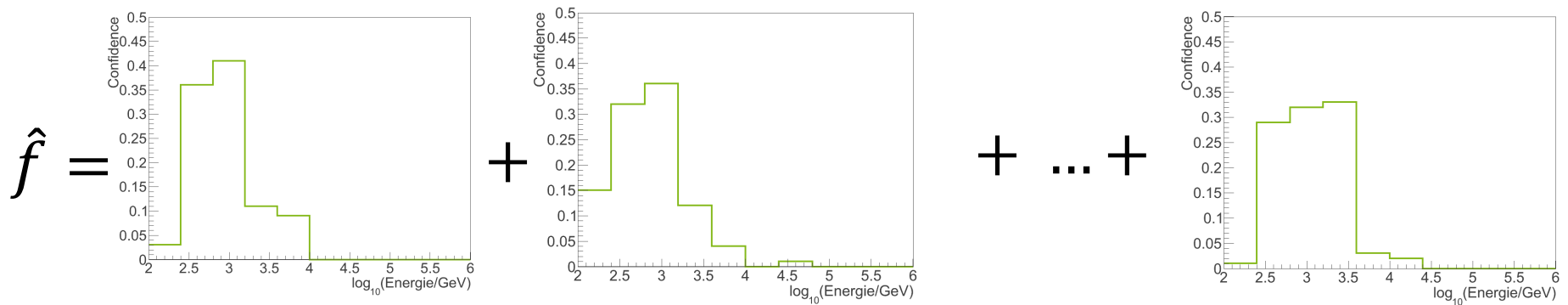


Geometric information is hard to include.

Information on individual events is lost.

Don't know which event contributed to which energy bin and how much...

## DSEA in greater detail



### Iterate:

1. Discretize
2. Train Model
3. Apply Model
4. Reconstruct spectrum
5. Update weights according to unfolding result

Choice of learning algorithm largely arbitrary (and probably somewhat problem dependent).

Some overlap with IBU in case Naive Bayes is used as a learner.



## Variable Step Width in DSEA+

Step Width

$$p_k = f_k - f_{k-1}$$

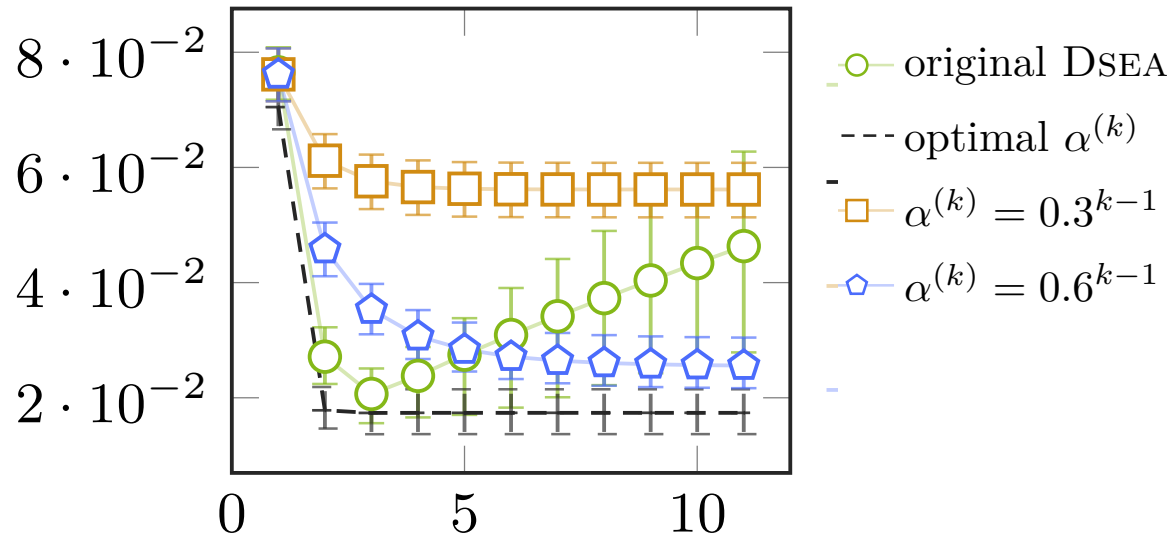
$\delta_H$

Next estimate then becomes

$$f_k = f_{k-1} + \alpha_k p_k$$

Find optimal  $\alpha$  via:

$$\alpha = \arg \min_{\alpha \geq 0} l(f_{k-1} + \alpha p_k)$$

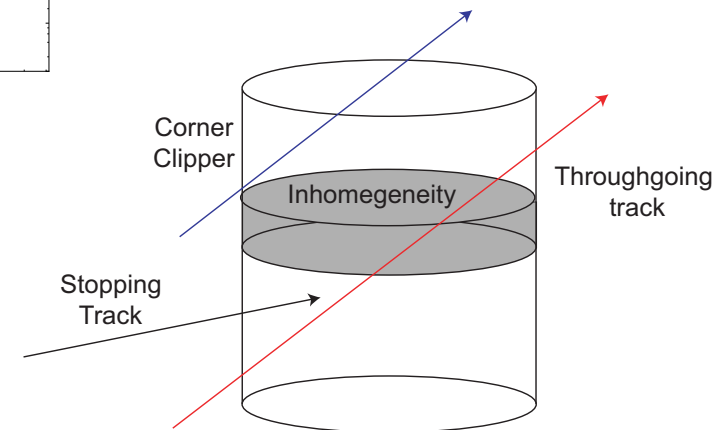
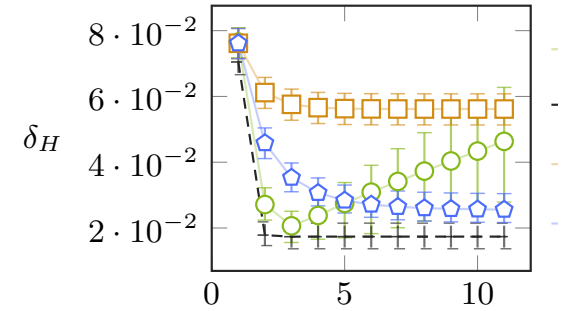
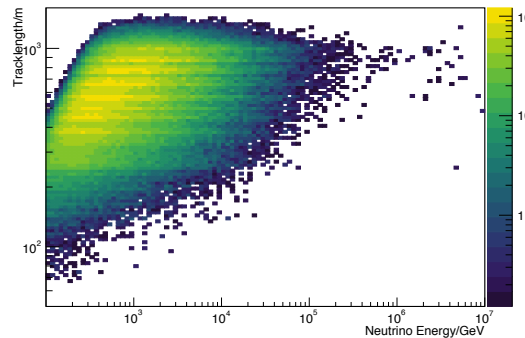
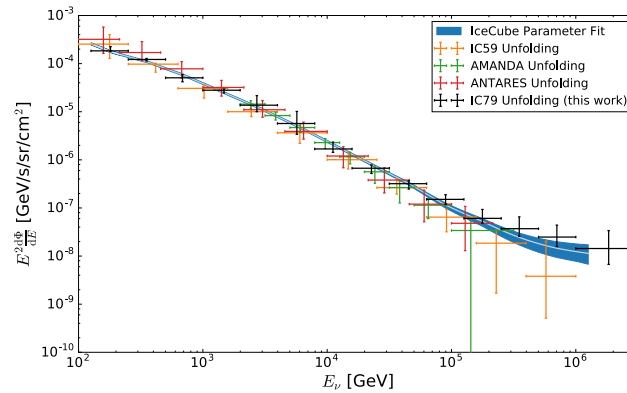
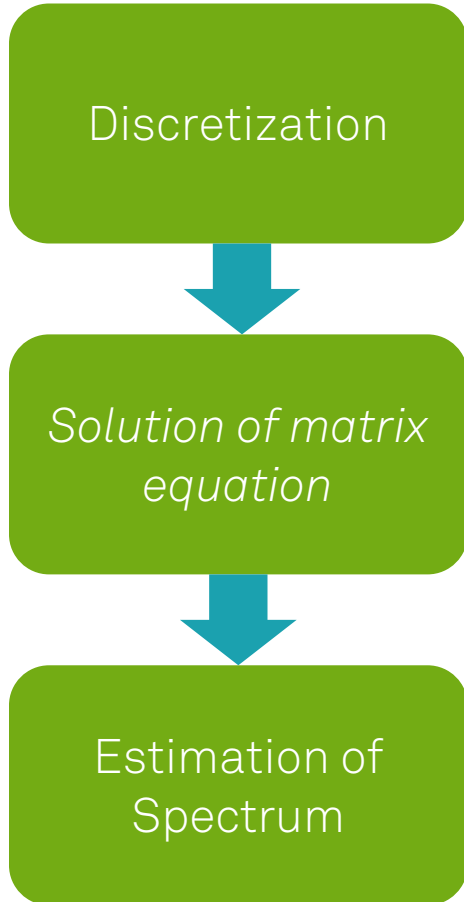


Get the software:

<https://sfb876.tu-dortmund.de/deconvolution/index.html>

For details see: [https://sfb876.tu-dortmund.de/PublicPublicationFiles/bunse\\_2018a.pdf](https://sfb876.tu-dortmund.de/PublicPublicationFiles/bunse_2018a.pdf)

## Summary and Outlook

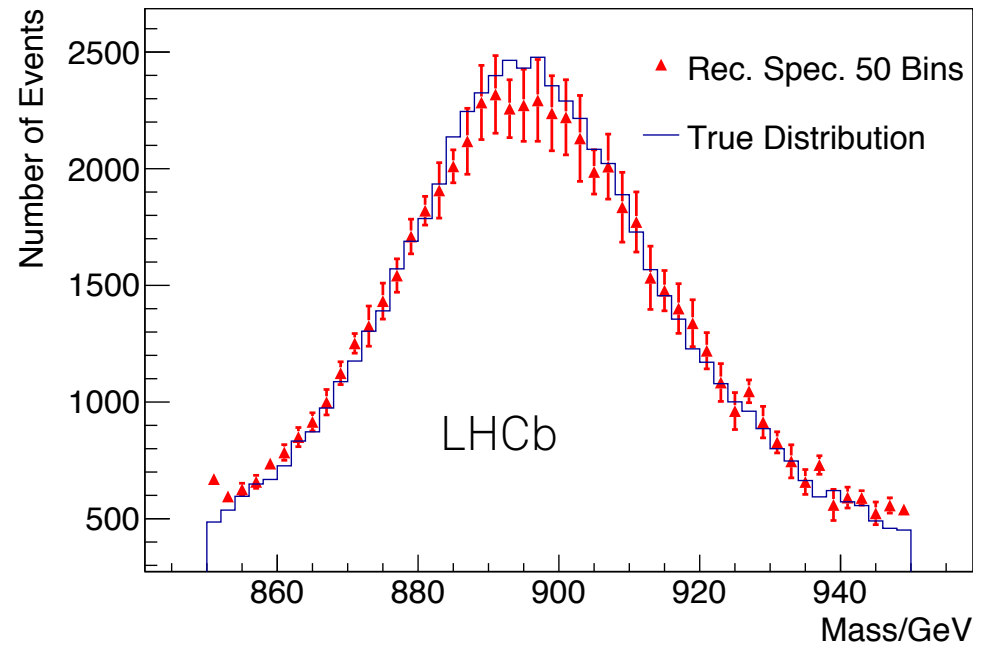
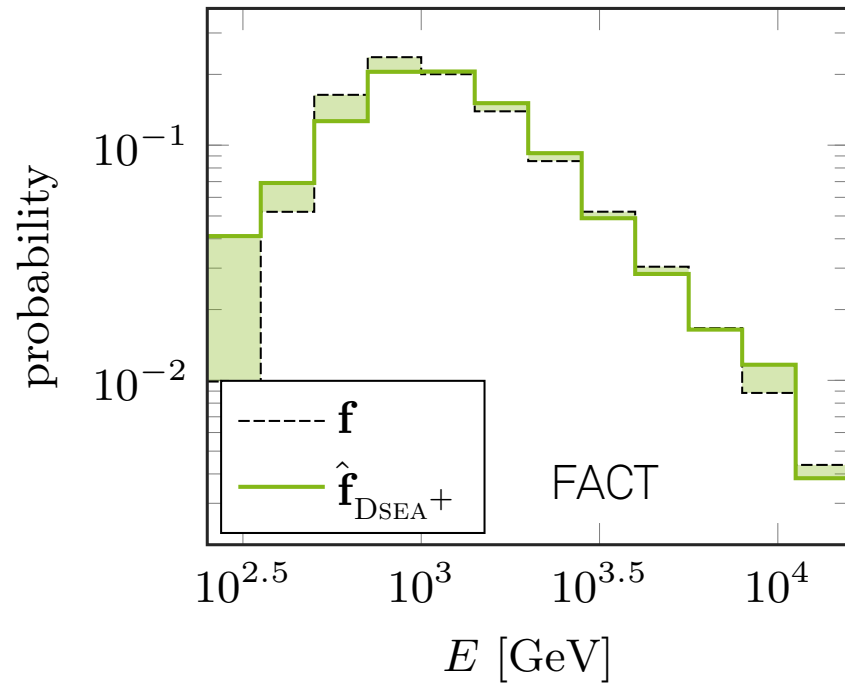


$$g(y) = \int_{E_{min}}^{E_{max}} A(E, y) f(E) dE$$

## Hellinger Distance

$$H^2(f, \hat{f}) = \frac{1}{2} \sum_{j=1}^m \sqrt{f_j} - \sqrt{\hat{f}_j}$$

## Some Preliminary Results



## Strategies for choosin the step-size

strategy	parameter	step size
constant factor	$\alpha > 0$	$\alpha^{(k)} = \alpha$
multiplicative decay (slow)	$0 < \eta < 1$	$\alpha^{(k)} = k^{(\eta-1)}$
exponential decay (fast)	$0 < \eta < 1$	$\alpha^{(k)} = \eta^{(k-1)}$

**Table 3.1:** Some simple strategies determine the step size  $\alpha^{(k)}$ .  $\eta$  is referred to as the *decay rate*. Each parameter controls the speed of the convergence in DSEA<sup>+</sup>.

## IC-79 NuMu Unfolding

