

# MCORD

## MPD Cosmic Ray Detector for NICA

by Polish consortium NICA-PL



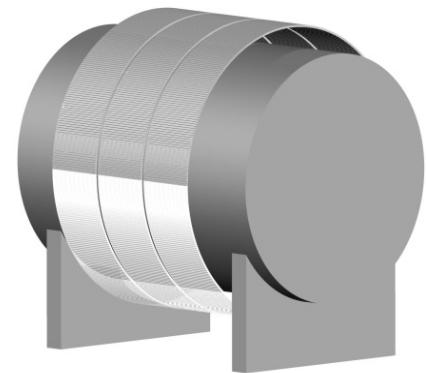
NARODOWE  
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JĄDROWYCH  
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# Outline



1. NICA collider
2. Cosmic Ray Detector – Goals
3. Main tasks and schedule
4. Design, modeling proposition
5. Conclusion



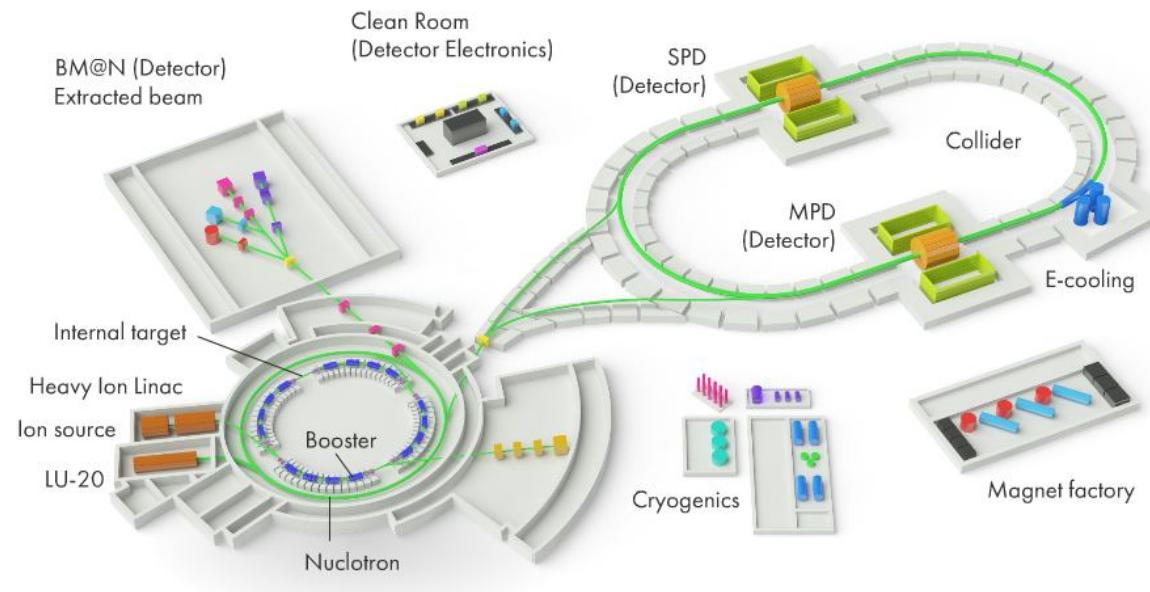
# 1. NICA complex - Glossary



**NICA - Nuclotron Ion Collider fAcility**

**MPD - Multi-Purpose Detector**

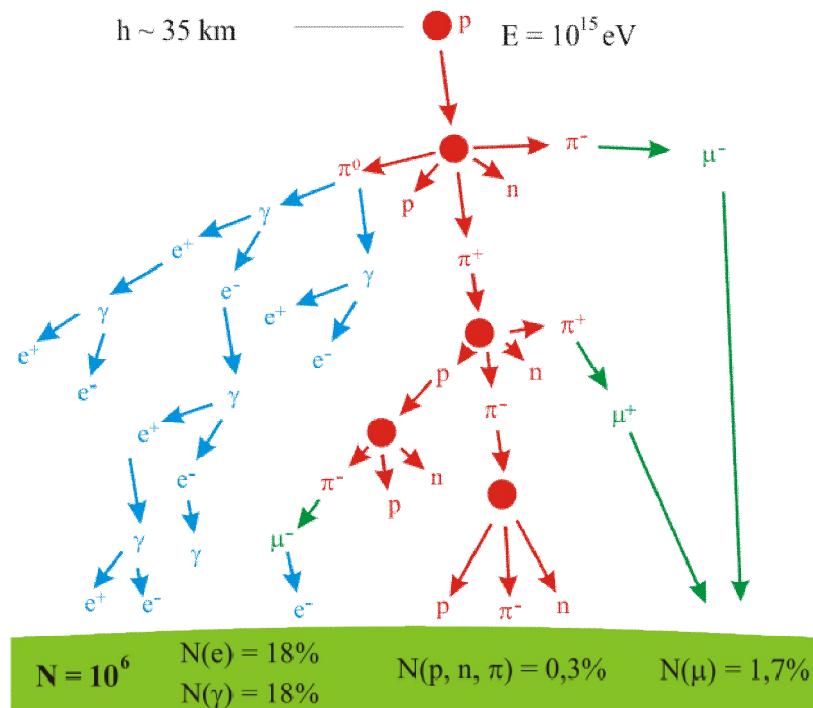
**MCORD - MPD Cosmic Ray Detector**



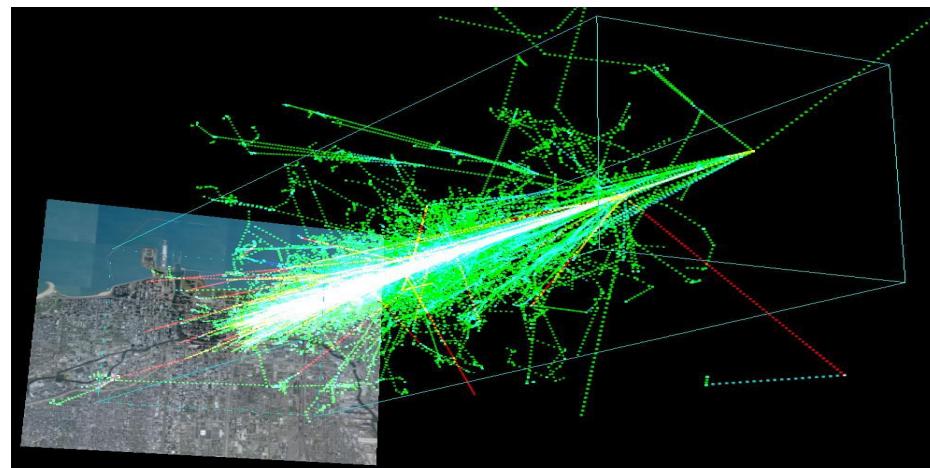
# 1. NICA complex - Cosmic Ray



## PRIMARY PARTICLE



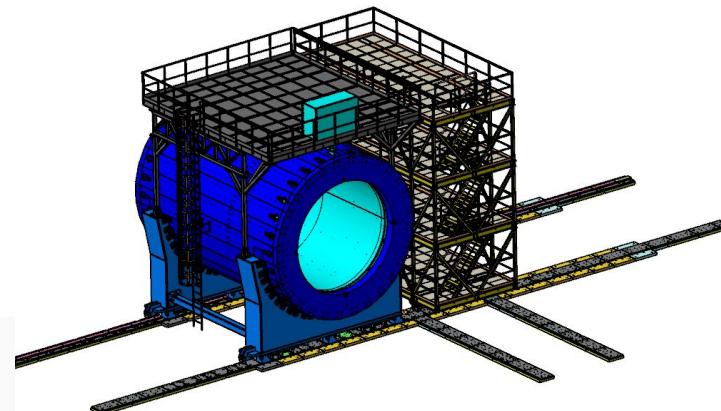
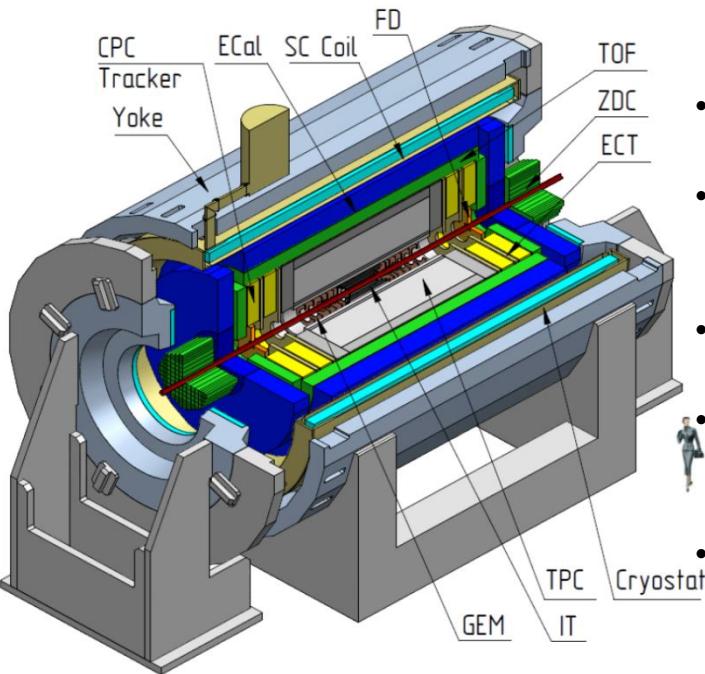
## GROUND LEVEL



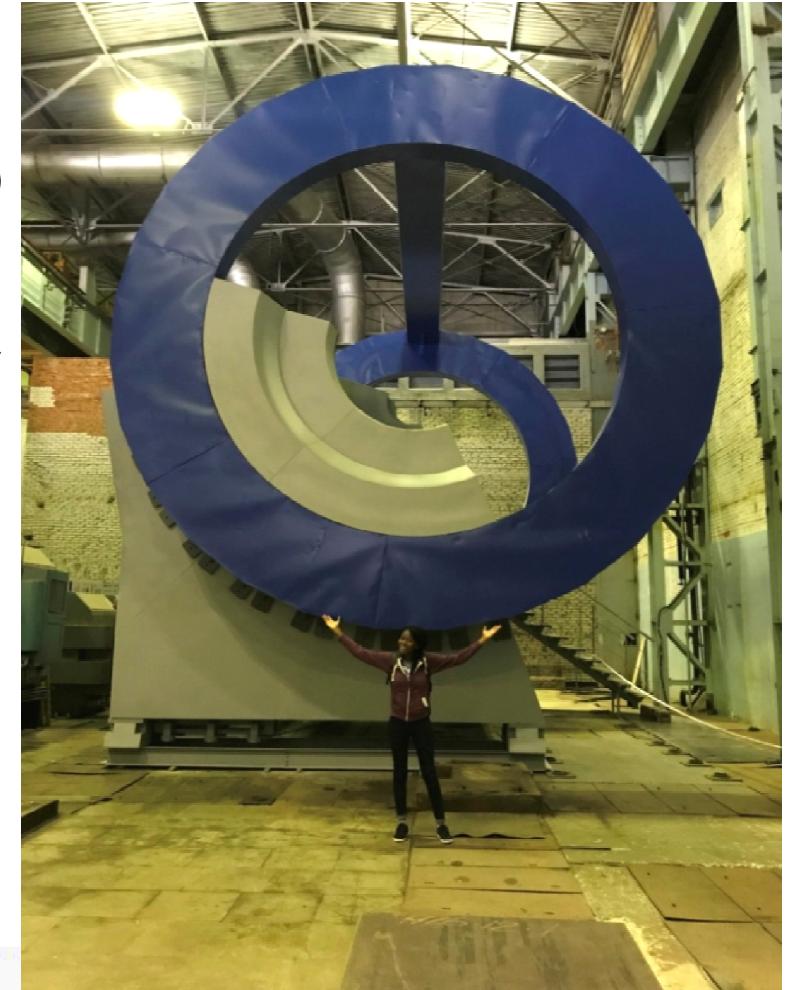
Cosmic ray air shower created by a 1TeV proton hitting the atmosphere 20 km above the Earth. The shower was simulated using the [AIRES](#) package.



# 1. NICA complex - MPD



- FD Forward det.
- Superconductor solenoid (SC Coil)
- Inner Tracker (IT)
- straw-tube tracker (ECT)
- Time-projection chamber (TPC)
- Time-of-flight system (TOF)
- Electromagnetic calorimeter (EMC - ECal)
- Zero degree calorimeter (ZDC).

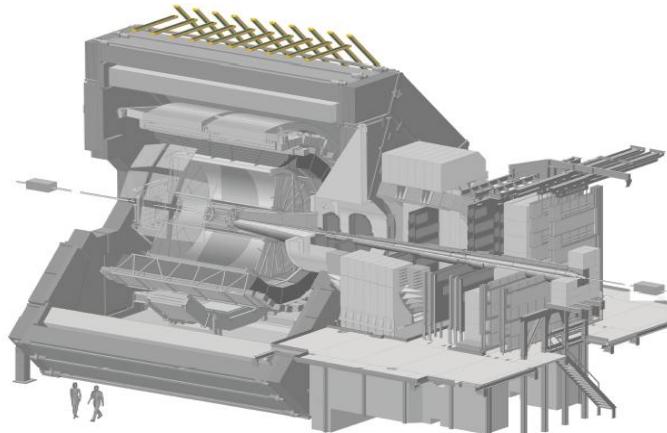


[nica.jinr.ru/video/general\\_compressed.mp4](http://nica.jinr.ru/video/general_compressed.mp4)

## 2. Cosmic Ray Detector – Goals



Examples from other experiments



ALEPH Exp.

140 m under. (thr. 70 GeV) (1997-99y)



Available online at [www.sciencedirect.com](http://www.sciencedirect.com)

SCIENCE @ DIRECT<sup>E</sup>

Astroparticle Physics 19 (2003) 513–523

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Astroparticle  
Physics  

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[www.elsevier.com/locate/astropart](http://www.elsevier.com/locate/astropart)

Cosmic multi-muon events observed in the underground  
CERN-LEP tunnel with the ALEPH experiment

V. Avati <sup>a,\*</sup>, L. Dick <sup>a,1</sup>, K. Eggert <sup>a</sup>, J. Ström <sup>a,2</sup>, H. Wachsmuth <sup>a,3</sup>,  
S. Schmeling <sup>b</sup>, T. Ziegler <sup>b</sup>, A. Brühl <sup>c</sup>, C. Grupen <sup>c</sup>

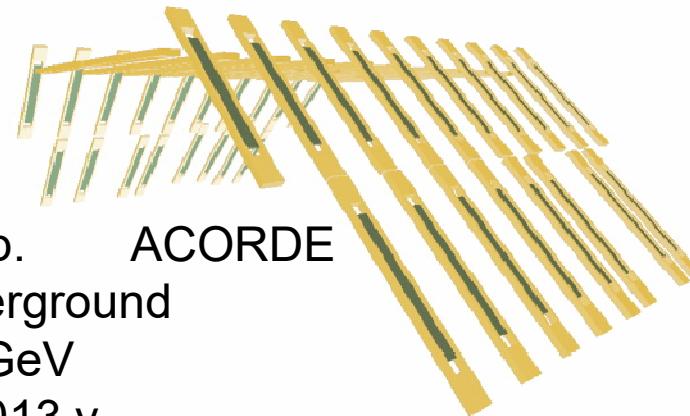
<sup>a</sup> European Laboratory for Particle Physics (CERN), Geneva, Switzerland

<sup>b</sup> Institut für Physik, Johannes Gutenberg-Universität Mainz, Germany

<sup>c</sup> University of Siegen, Siegen, Germany

Received 26 July 2002; received in revised form 27 October 2002; accepted 26 November 2002

ALICE Exp.  
55 m underground  
thr. 16 GeV  
2010-2013 y



DELPHI Exp.

100 m under. (thr. 52 GeV) (99-2000y)



Available online at [www.sciencedirect.com](http://www.sciencedirect.com)

ScienceDirect

Astroparticle Physics 28 (2007) 273–286

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Physics  

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[www.elsevier.com/locate/astropart](http://www.elsevier.com/locate/astropart)

Study of multi-muon bundles in cosmic ray showers detected  
with the DELPHI detector at LEP

DELPHI Collaboration  
J. Abdallah <sup>z</sup>, P. Abreu <sup>w</sup>, W. Adam <sup>bc</sup>, P. Adzic <sup>1</sup>, T. Albrecht <sup>r</sup>, R. Alemany-Fernandez <sup>i</sup>,  
T. Allmendinger <sup>r</sup>, P.P. Allport <sup>x</sup>, U. Amaldi <sup>ad</sup>, N. Amapane <sup>av</sup>, S. Amato <sup>az</sup>, E. Anashkin <sup>ak</sup>,  
A. Andreazza <sup>ac</sup>, S. Andringa <sup>w</sup>, N. Anjos <sup>w</sup>, P. Antilogus <sup>z</sup>, W.-D. Apel <sup>r</sup>, Y. Arnoud <sup>o</sup>,  
S. Ask <sup>aa</sup>, B. Asman <sup>au</sup>, A. Augustinus <sup>i</sup>, P. Baillon <sup>i</sup>, A. Ballestrero <sup>aw</sup>, P. Bambade <sup>u</sup>,  
D. Basso <sup>ab</sup>, D. Bazzocchi <sup>g</sup>, G. Belotti <sup>be</sup>, A. Benlliure <sup>an</sup>, M. Bernabeu <sup>in</sup>, M. Bertaina <sup>z</sup>,



M.Bielewicz, 23.I.2019 DAC NICA Dubna

## 2. Cosmic Ray Detector – Goals



- a) Trigger (for testing or calibration)
  - testing before completion of MPD (testing of TOF, ECAL modules and TPC)
  - calibration before experimental session
- b) Veto (normal mode - track and time window recognition)  
Mainly for TPC and eCAL

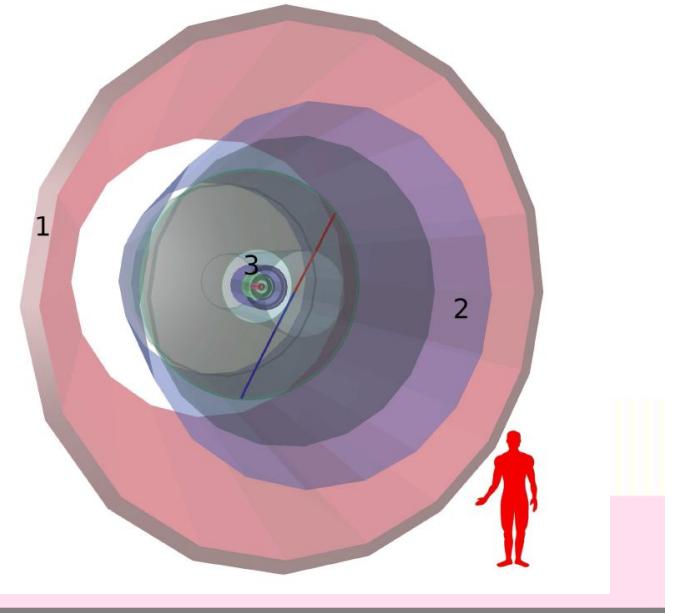
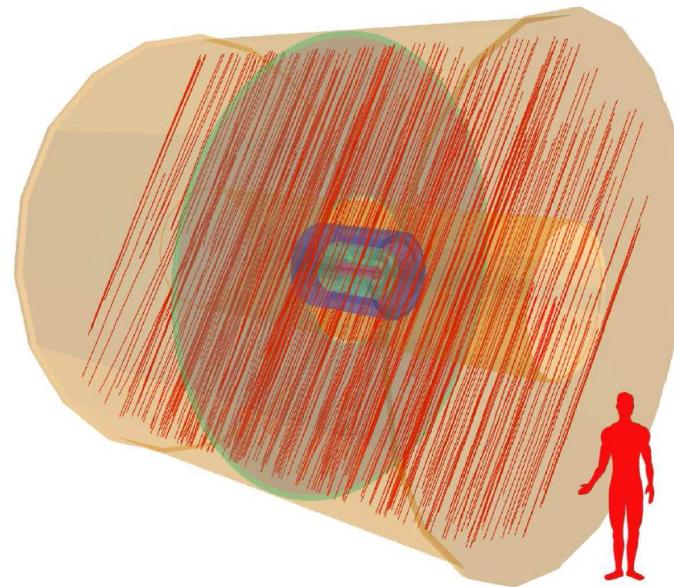
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Additionally

- c) Astrophysics (muon shower and bundles)
  - unique for horizontal events

Working in cooperation with TPC

DECOR exp. 2002-2003y (near horizontal observation  
(60-90 deg. angular range) - 1-10 PeV primary  
particle)





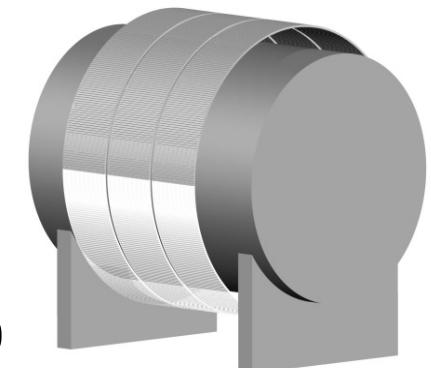
### 3. Main tasks and schedule

#### I. Conceptual Design (9-12 months)

(Preliminary Technical and Electronic Design, Market Research, Literature Studies, Cost Estimates, TPC and TOF requirements)

#### II. Module Optimization (9-15 months )

(Scintillator, Power Supply, Front-End and Analog electronics characterization; Detector response, Cosmic-ray and MPD spectra simulation, Veto response, Integration )



#### III. Demonstrator constr. (2-4 ready to use modules) (10-20

months) (Demo. detailed design, Procurement of Modules, Lab. Tests, Installation in MPD

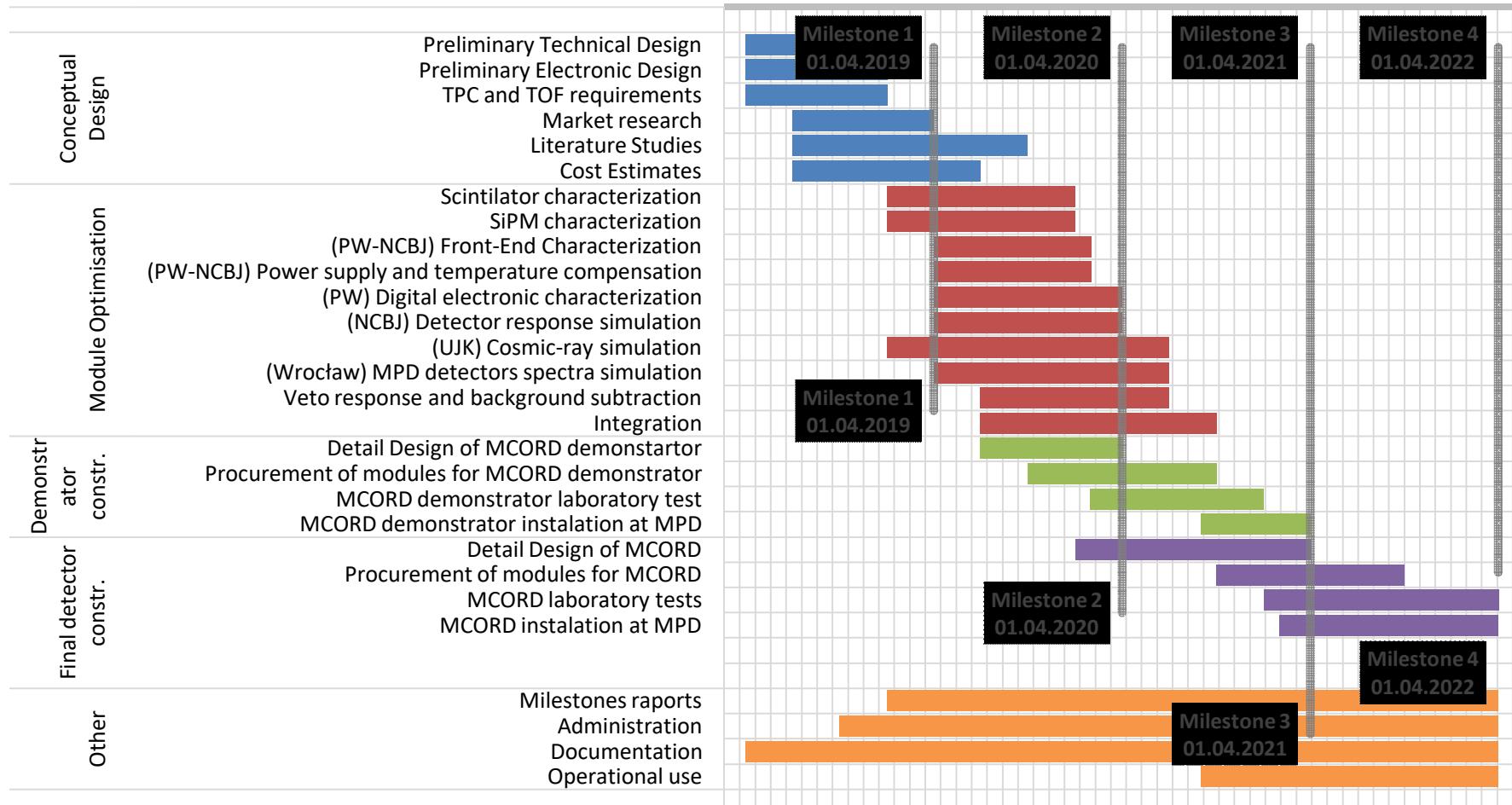
#### IV. Final Detector constr. (procurement of all modules , test installation and integration –12 months)



# 3. Main tasks and schedule



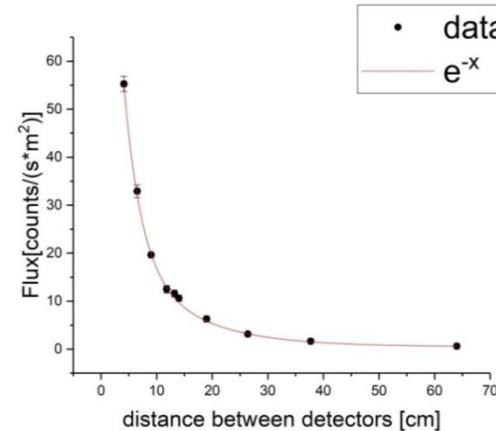
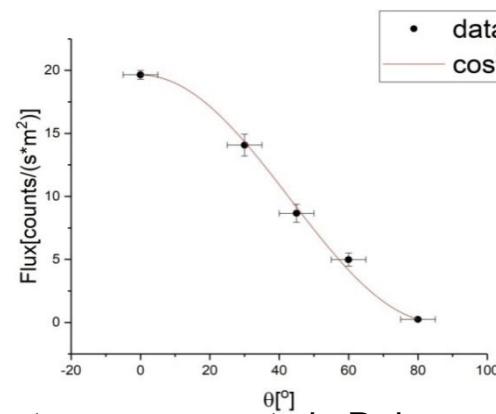
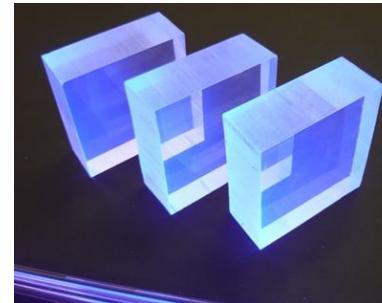
FM\AMJuJAS\CNDJ:FM\AMJuJAS\CNDJaIM\AMJuJAS\CNDJ:FM\AMJuJAS\CNDJ:FM Apr 29, 2022



# Conceptual design – tasks



1. Measurement of cosmic radiation - azimuthal and horizontal distribution
2. Angular influence of the building (concrete walls and ceiling)
3. Optimization of scintillator shape and light readout method
4. Design of single module and scaling up to full size instrument
5. Simulations (MCNPX, GEANT4 and CORSIKA, Showersim)



- First measurements in Dubna
- Azimuthal and horizontal distribution
  - Angular influence of the building
  - With Pb filter



# Conceptual design – laboratory tests



## Laboratory tests at NCBJ Swierk - Plans for 2019

- Laboratory tests of plastic tiles with and without wavelength shifting fibers
- Readout by PMTs for reference and with silicon PMs for optimization of single detector performance.
- We will study light yield and coincidence resolving time for different configurations of scintillator and photodetectors.
- The tests will be supported with grant from Polish contribution to JINR.

**Cooperation with scintillators manufacturers and providers:**

- Uniplast (Russia)
- Amcrys-H (Ukraine)
- Saint-Gobain (FR)
- Scionix (Holland)
- Nuvia (Czech)

**Cooperation with SiPM vendors:**

- Hamamatsu (Japan)
- AdvanSiD (Italy),
- Ketek (Germany),
- SensL (Ireland)

**Wide range of sizes:**  
 $1 \times 1 \text{ mm}^2$  to  $51 \times 51 \text{ mm}^2$



Uniplast



NUVIA



Hamamatsu



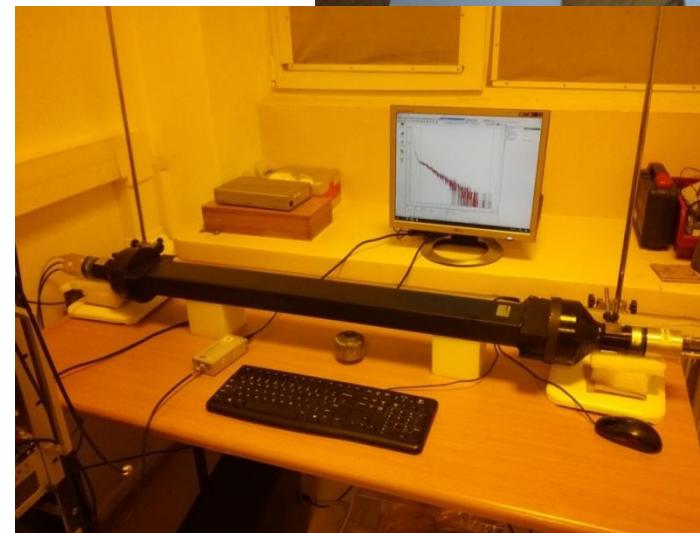
# Conceptual design – laboratory tests



## Laboratory tests at NCBJ Swierk

Available equipment:

- long tiles (~100-150 cm) from NUVIA (Czech Rep.) and UNIPLAST (Russia) with and without Wavelength Shifting (WLS) fibers
- 5" dia PMTs (XP45D2 and ETL9390)
- medium and small SiPMs (6x6 and 1x1 mm) from Hamamatsu (in future 25x25 mm)
- first measurements of light output and light attenuation along 100x10x5 cm plastic tile
- double-side 5 inches dia PMTs readout
- Co-60 gamma-rays energy calibration

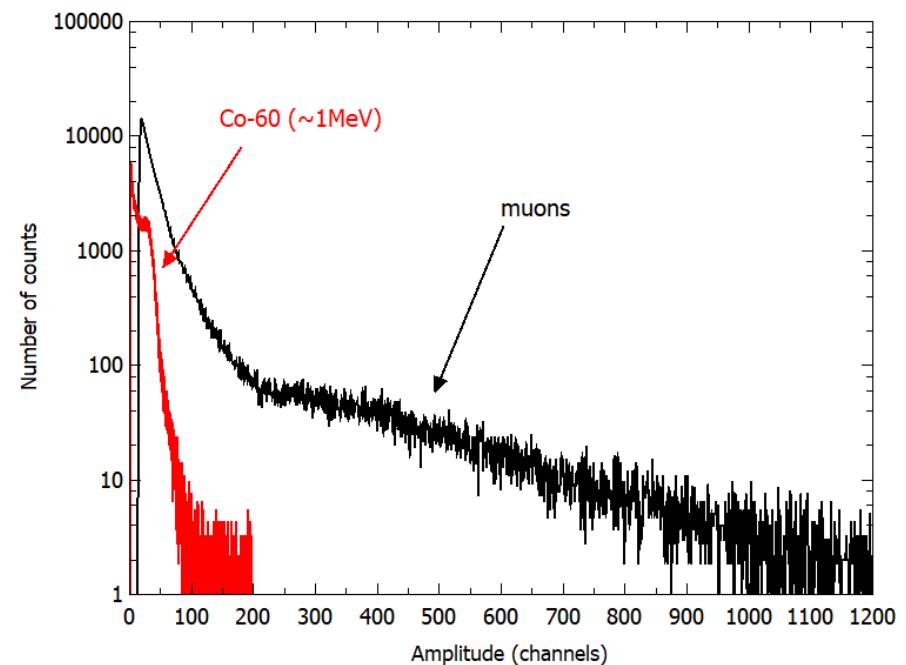
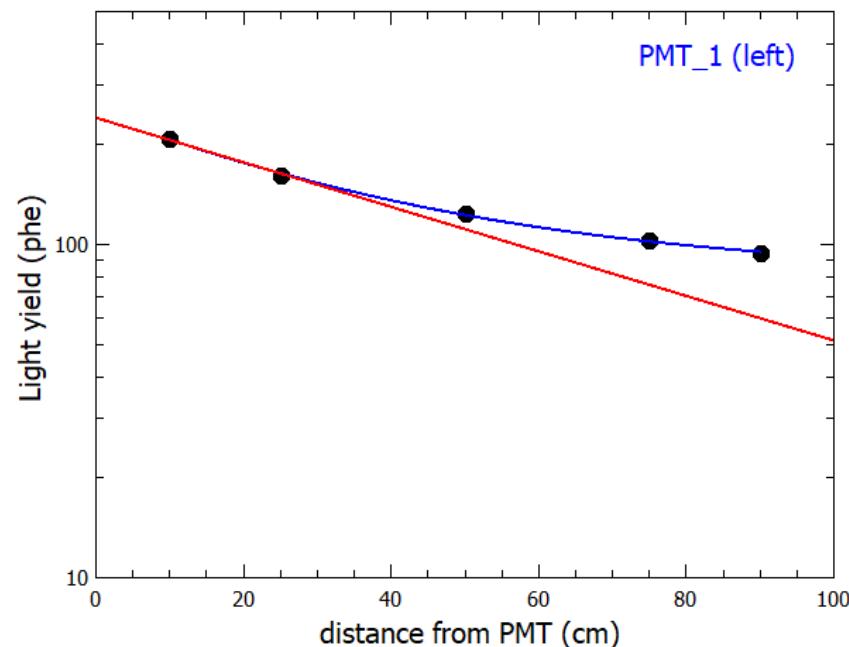


# Conceptual design – laboratory tests



## Laboratory tests at NCBJ Swierk

- attenuation length in Nuvia plastic (100×10×5 cm)  
 $\lambda > 65$  cm
- muon energy spectra along entire slab recorded (RT = 1000s)



# Conceptual design – simulations



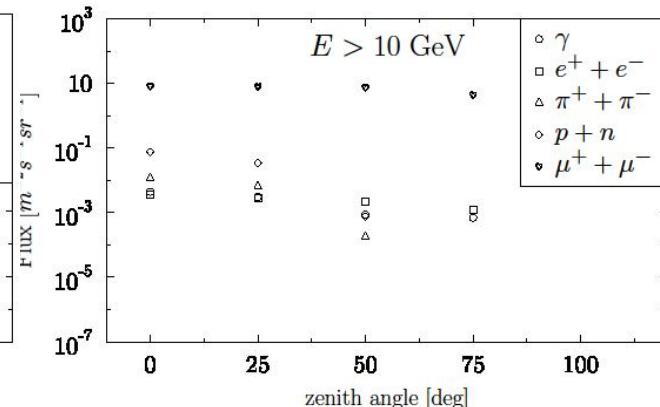
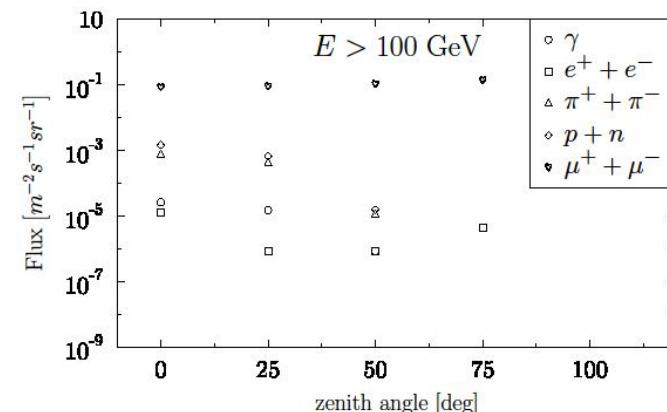
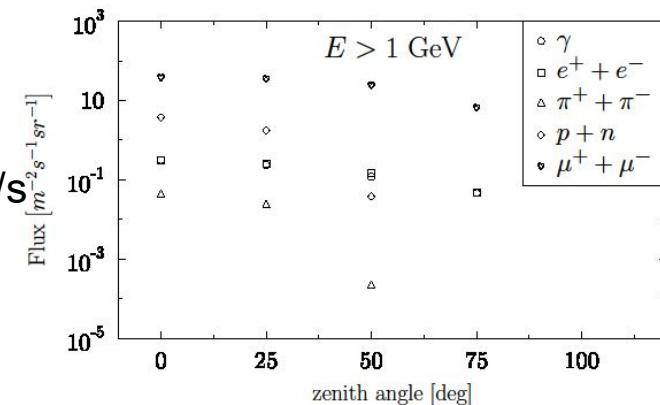
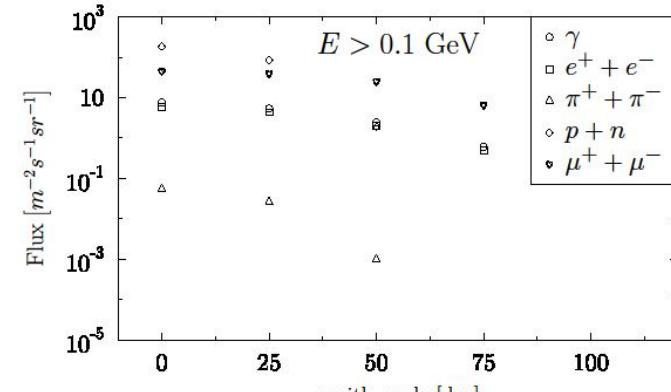
## CORSIKA simulation of Cosmic Showers

Angular distribution of atmospheric cosmic shower particles

E [GeV]	0	25	50	75
0.1- 1	5.912	3.829	0.612	0
1-10	29.931	26.935	16.637	2.224
10-100	7.976	7.884	7.291	4.186
100~1000	0.087	0.092	0.108	0.142

Muon Flux from all angle for E>1GeV is about 100-150 count/m<sup>2</sup>/s

It gives (6.72m x 4.7m = 31.58m<sup>2</sup>) about 3000-4500 count/s



# Conceptual design – simulations



## MCNP calculations for MCORD muon detector

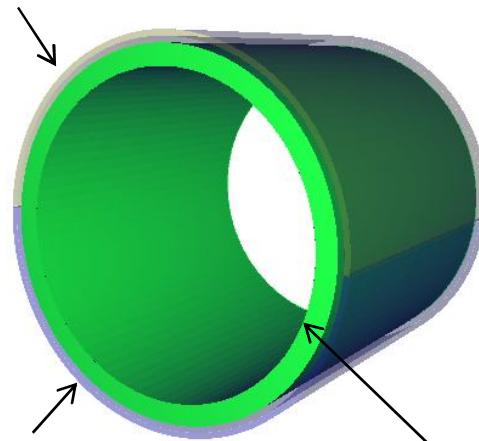
(MCNP 6.11, MCNPX 2.7.0. number of iteration 1E9)

Develop the simple model of MCORD detector

divided into top and bottom semicylinder (monolithic) + yoke

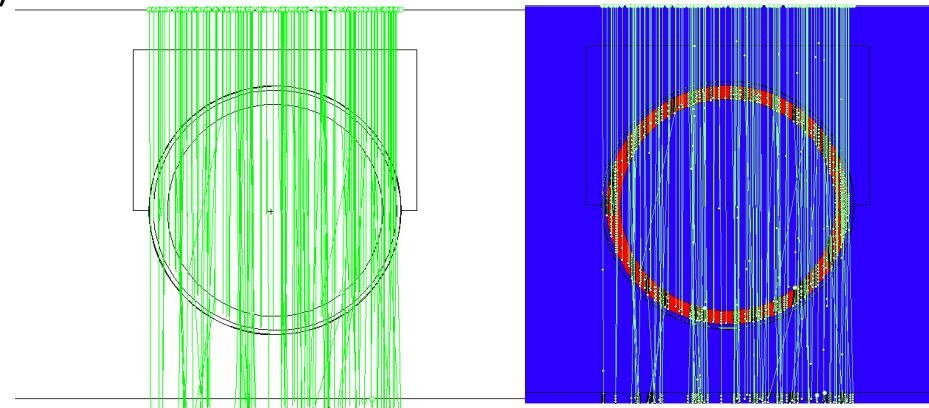
- Two half-cylinders of plastic (2.5 cm and 1 cm thick)
- Implemented energy distribution as a function of muon incident angle:
  - Angles: 0 and 45 degrees
  - Energy: 0.1 – 1000 GeV
- Implemented surface emission
- Estimated bottom-to-top muon coincidence ratio

## Top detectors



Bottom detectors

Yoke



# Conceptual design – simulations



Pure facts:

- Coincidence rate: when **both top and bottom detectors detect muon**
- The coincidence rate for two types of detector and for two muon incident degrees:

Detector thickness (cm)	Angle (deg)	Coincidence rate (%)
1	0	98
1	45	98
2.5	0	96
2.5	45	98

- Coincidence rate lower for thicker detector. Worse coincidence rate for thicker detector and incident muon angle of 0 degree (lower average muon energy) – probably due to muon scattering/glazing on the edge.
- Other muon emission angles are under calculations.

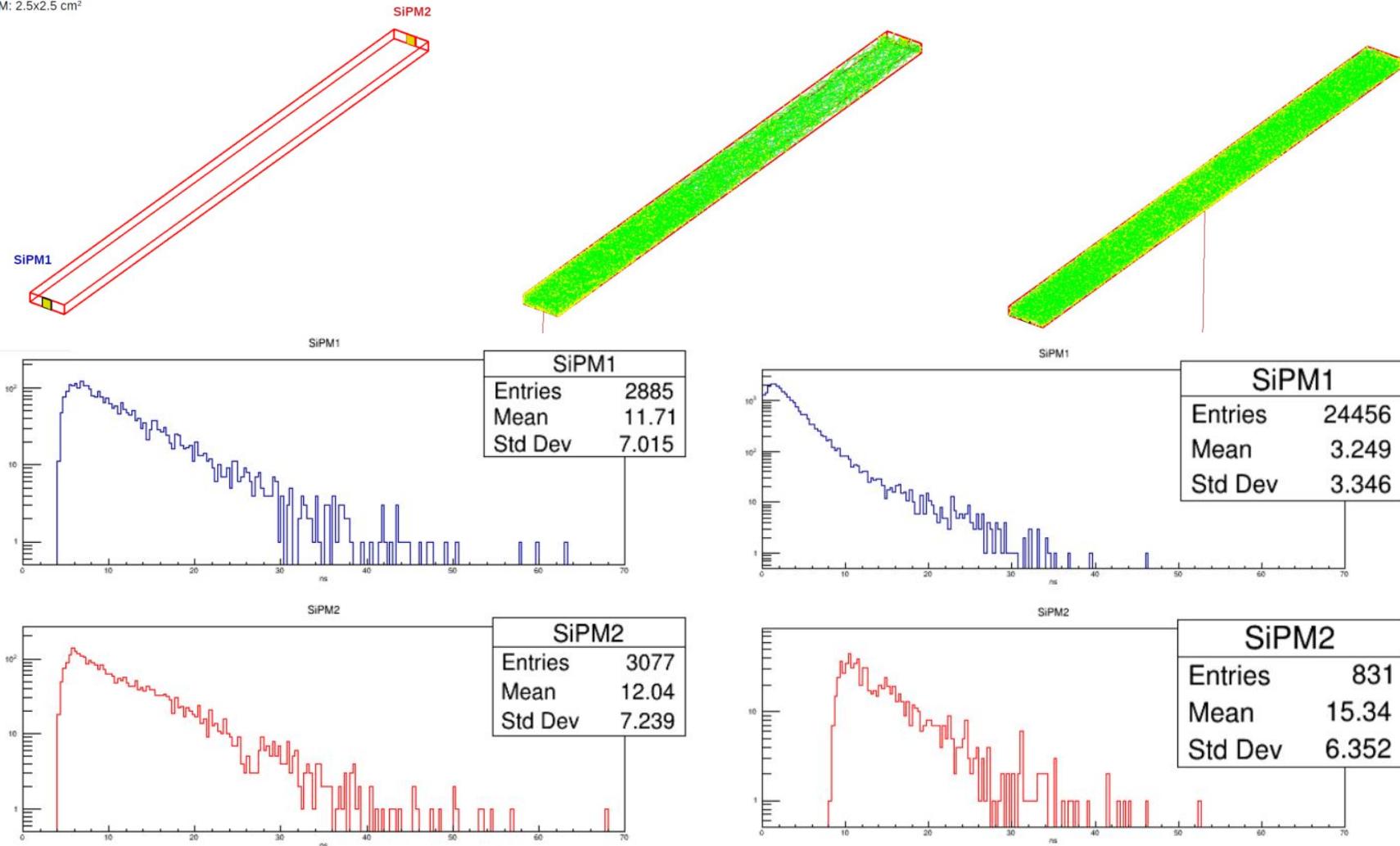


# Conceptual design – simulations



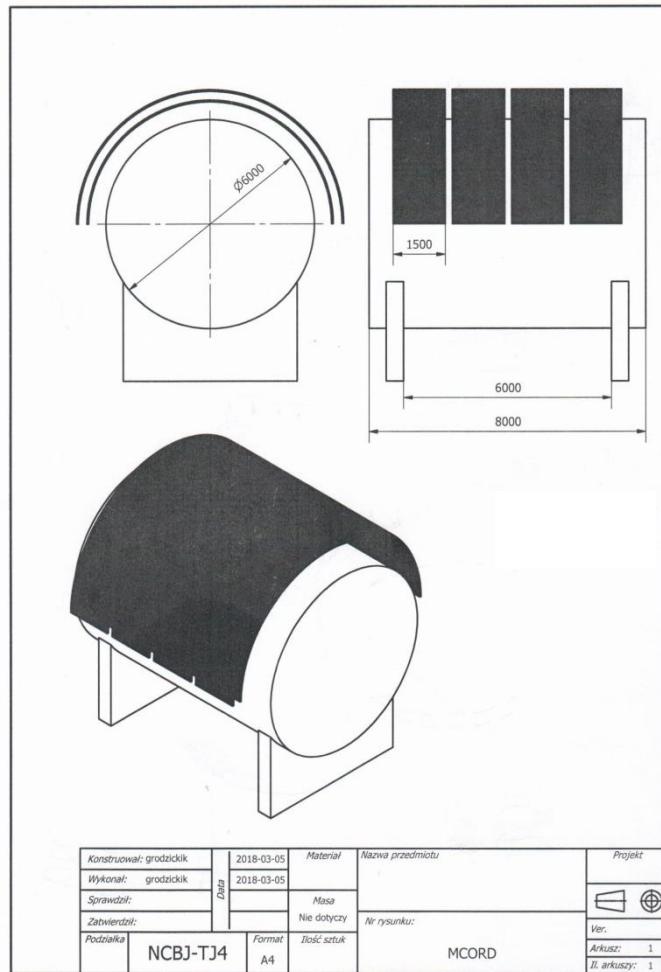
GEANT4 simulation of photon production and distributions in scintillators

scintillator: BC404, 150x10x2.5 cm<sup>3</sup>  
SiPM: 2.5x2.5 cm<sup>2</sup>

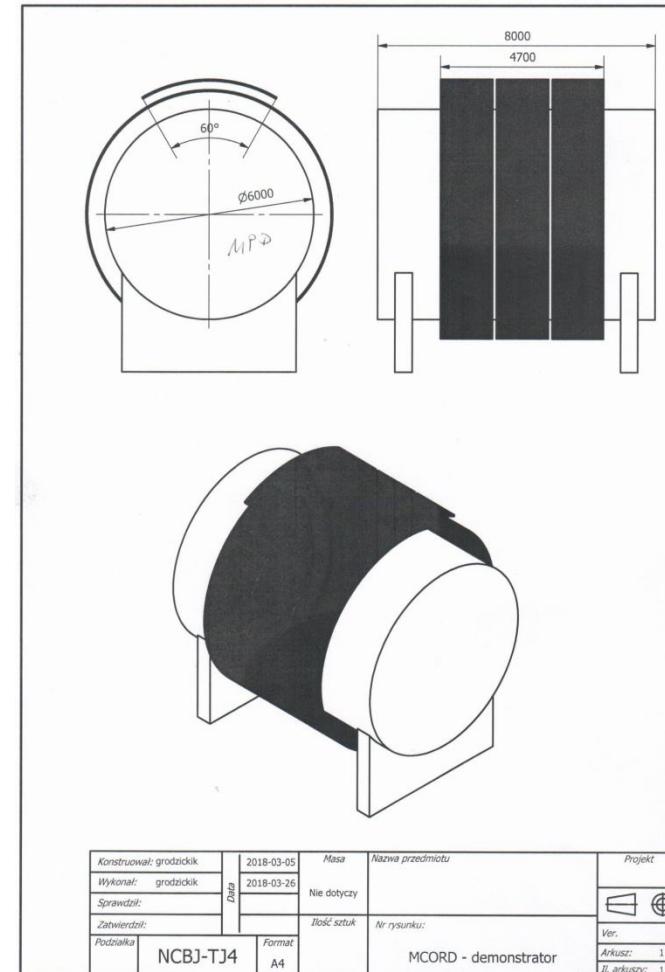


M.Bielewicz, 23.I.2019 DAC NICA  
Dubna

# 4. Design, modeling variants



Two surfaces on half circumference

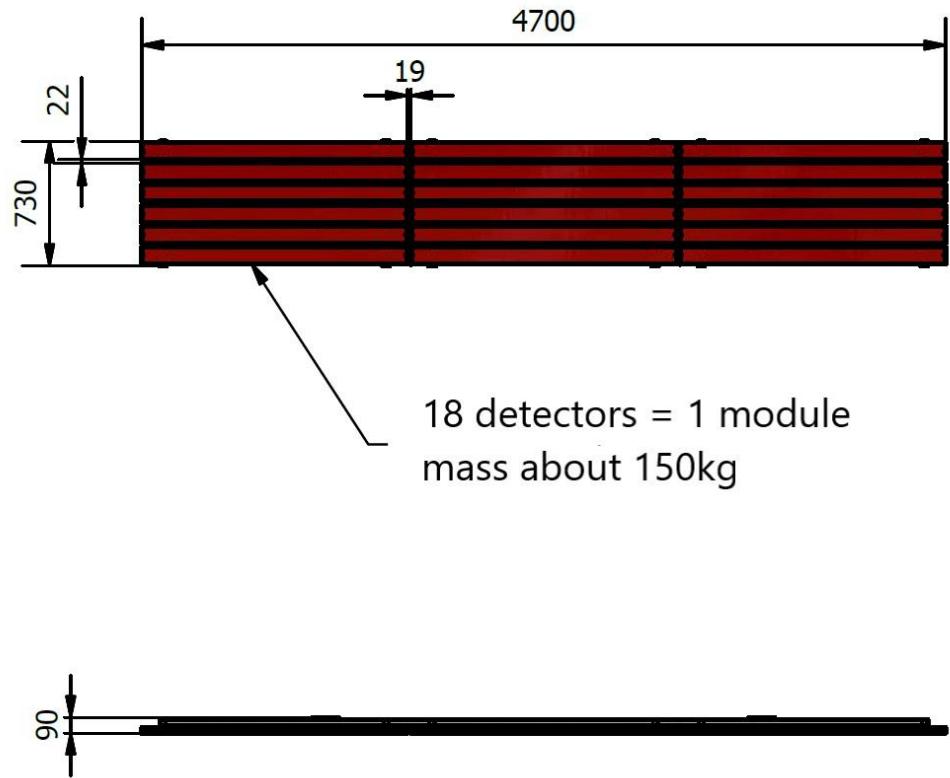


One surface on full circumference

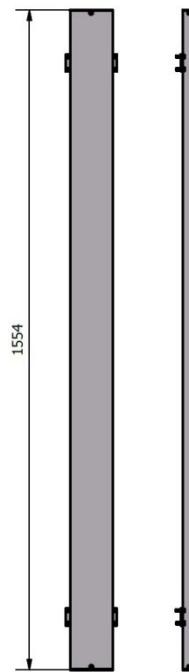
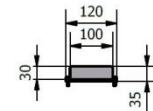




## 4. Design, modeling variants

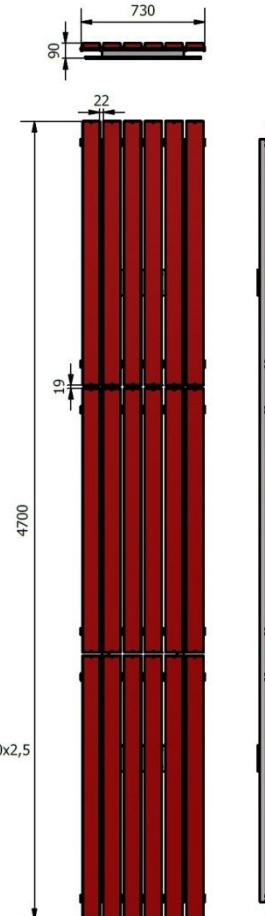


## Scintillators and modules

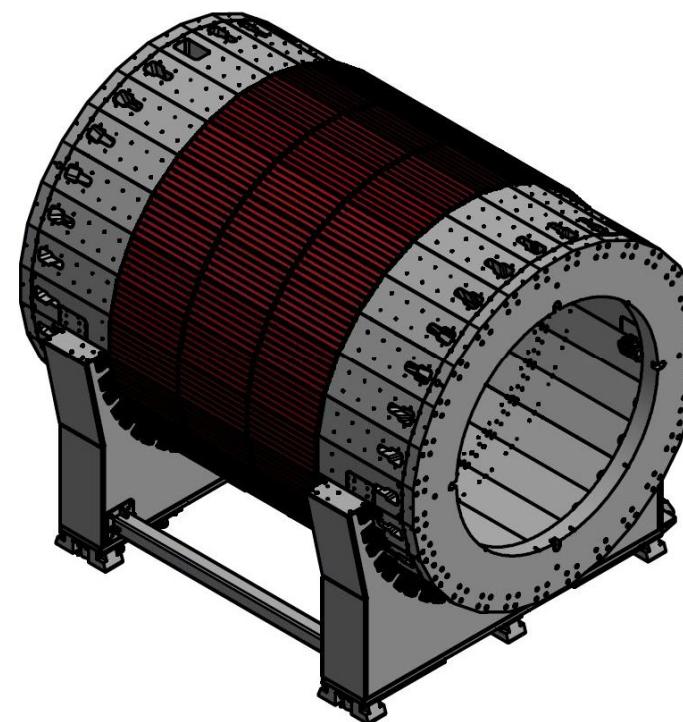
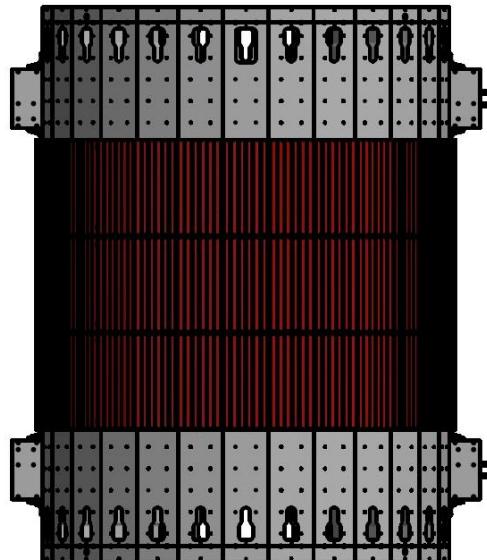


**DETECTOR**  
Dimensions of scintillator - 95x25x1500  
Dimensions of detector - 100x30x1554  
Material of casing - Aluminium alloy  
Scintillator placed in rectangle profile 100x30x2,5  
Weight of detector - 6,5kg

Module of detectors  
Number of detectors: 18  
Dimensions of module: 730x90x4700  
Weight of module: 150kg  
Detectors mounted to steel frame.  
Steel frame built with square profiles.  
Frame mounted to MPD by screws.



## 4. Design, modeling variant



**MCORD at  
MPD scheme  
Basic variant**

One surface on full circumference

### Modules Mounting on MPD surface

Detector mounted to steel frame

Steel frame build with square profiles:

40x40 [mm]

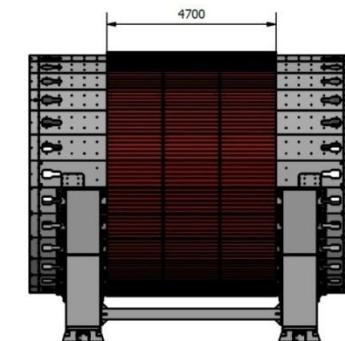
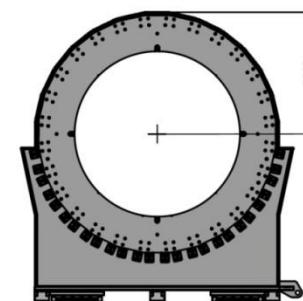
Number of Modules:

28

Frame mounted to MPD by screws

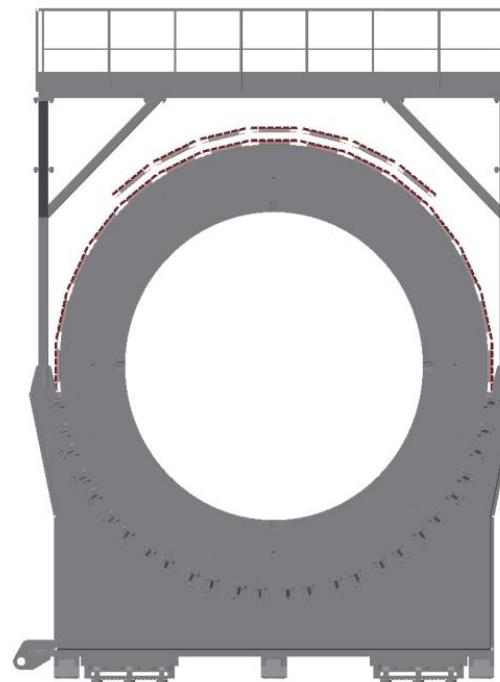
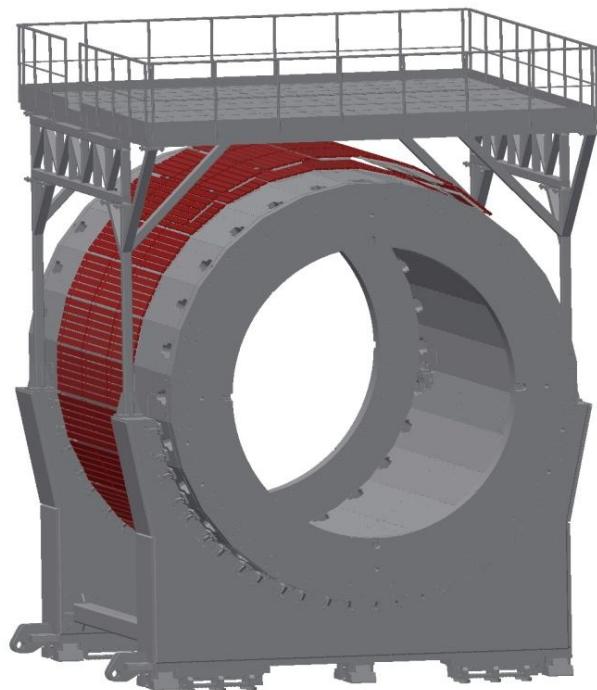
Weight of all modules:

4200 kg





## 4. Design, modeling variants



### MCORD at MPD scheme

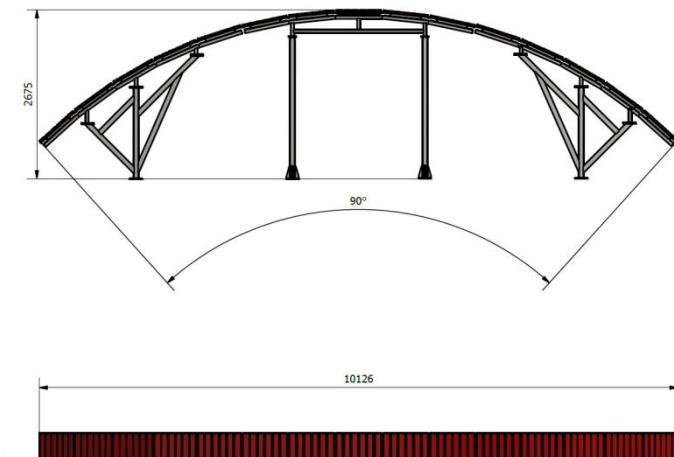
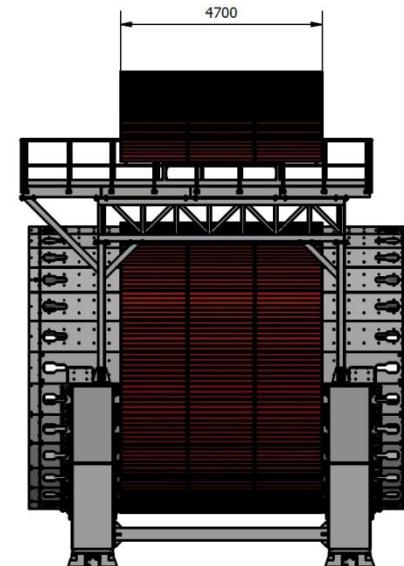
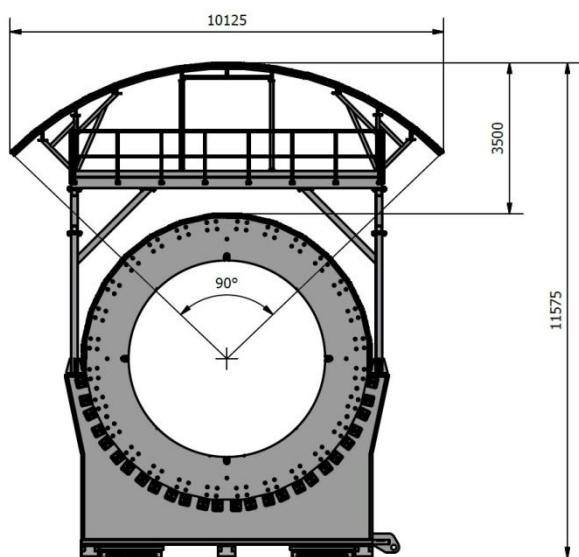
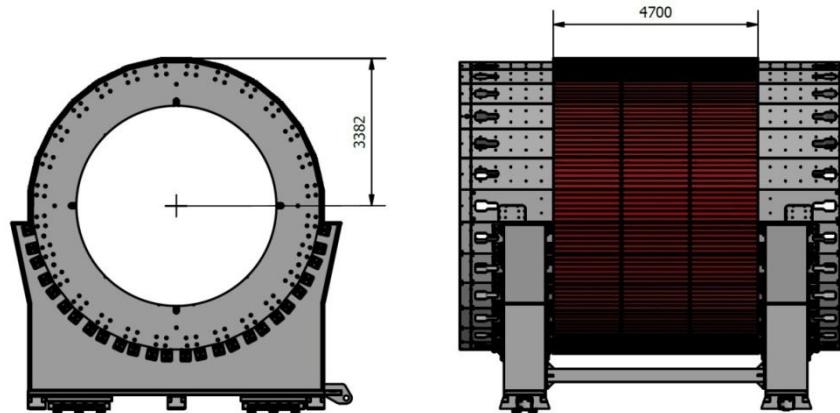
One surface on full circumference + additional surface on the top ver.1





## 4. Design, modeling variants

MCORD at MPD scheme



One surface on full circumference + additional surface on the top ver.2



# 4. Design, modeling variants – FIBER?

## OPTION 1

Scintillator:

length:

width:

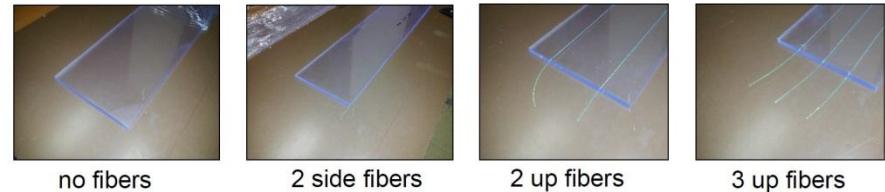
thickness:

**PVT or PS plastic (no fibers)**

**150 cm**

**10 cm**

**2.5 cm**



no fibers

2 side fibers

2 up fibers

3 up fibers

## Light Yield estimation

Muon  $dE/dx = 2 \text{ MeV} / (\text{g}\cdot\text{cm}^2)$  - light production by one particle (muon)

Muon Energy Deposition	Scinti. Light Yield (LY)	Muon LY
~5 MeV	$\times \sim 8000 \text{ ph/MeV}$	~40000 ph

Light collection eff.: ~20% (?)

**Attenuation length ( $\lambda$ ): ~75 cm**

Light atten.:  $\text{LY}/[2\exp(x/\lambda)]$

$\text{LY}_{\text{mid}}: \text{LY} \times 0.2 / (2e) \approx 7000 \text{ ph}$

$\text{LY}_{\text{end}}: \text{LY} \times 0.2 / (2e^2) \approx 2500 \text{ ph}$

Photodetector: **SiPM** size: **25×25 mm<sup>2</sup>** (**1/4 scintillators end area**)  
 Dark Count Rate (DCR): ~50 Mcps (@ 0.5 phe)  
 DCR @ 5 phe (threshold level): ~10 kcps

Light Yield (LY)	Photodetector QE: ~25%	Photoelectron Yield (PHE)
	Relative area: 25%	
~7000 ph	$\times 25\% \times 25\%$	<b>~500 phe</b>
~2500 ph	$\times 25\% \times 25\%$	<b>~150 phe</b>





## 4. Design, modeling variants – FIBER?

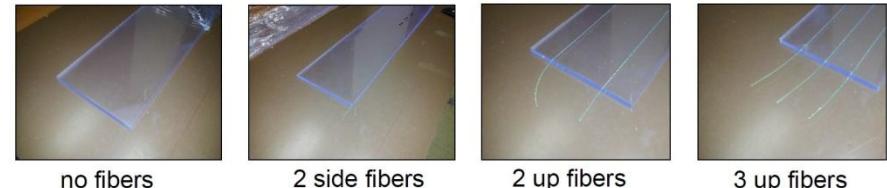
### OPTION 2

scintillator: **PVT or PS plastic + WLS (fibers)** (fi:1 mm)

length: **150 cm**

width: **10 cm**

thickness: **1 cm (less thickness)**



### Light Yield estimation

Muon  $dE/dx = 2 \text{ MeV} / (\text{g}\cdot\text{cm}^2)$  - light production by one particle (muon)

Muon Energy Deposit	Scinti. Light Yield (LY)	Muon LY
$\sim 2 \text{ MeV}$	$\times \sim 8000 \text{ ph/MeV}$	$\sim 16000 \text{ ph}$

Light collection eff.: ~10% (?)

**Atten. length ( $\lambda$ ): ~350 cm** Light atten.:  $LY/[2\exp(x/\lambda)]$

**$LY_{mid}$ :  $LY \times 0.1 / [2\exp(75/350)] \approx 600 \text{ ph}$**

**$LY_{end}$ :  $LY \times 0.1 / [2\exp(150/350)] \approx 500 \text{ ph}$**

photodetector:

**SiPM**

size:  **$1 \times 1 \text{ mm}^2$  (100% fiber end area)**

dark count rate (DCR): ~0.5 Mcps (@ 0.5 phe)

**DCR @ 3 phe (less threshold level) : ~10 kcps**

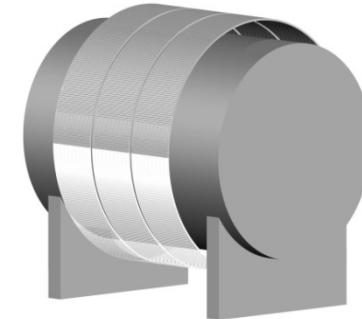
Light Yield (LY)	Photodetector QE: ~25%	Photoelectron Yield (PHE)
	Relative area: 100%	
$\sim 600 \text{ ph}$	$\times 25\% \times 100\%$	<b><math>\sim 150 \text{ phe}</math></b>
$\sim 500 \text{ ph}$	$\times 25\% \times 100\%$	<b><math>\sim 120 \text{ phe}</math></b>





# 4. Design, modeling variants – FIBER?

<b>Size estimation</b>	<b>Detector MCORD</b>
Diameter:	7 m
Length:	4.5 m
Circumference:	22 m
No. of scintillators:	<b>660 pcs</b>
No. of SiPM:	<b>1320 pcs</b>
No. of mTCA:	<b>4 crates</b>



## False trigger rate (FTR) estimation

SiPM-SiPM coinc\_gate: 20 ns (two ends of one scintillator)  
scintillator trigger\_gate: 100 ns (two scintillators on MCORD cylinder)  
cosmic muon rate: <150 cps/m<sup>2</sup>

## **DCR (dark count rate)**

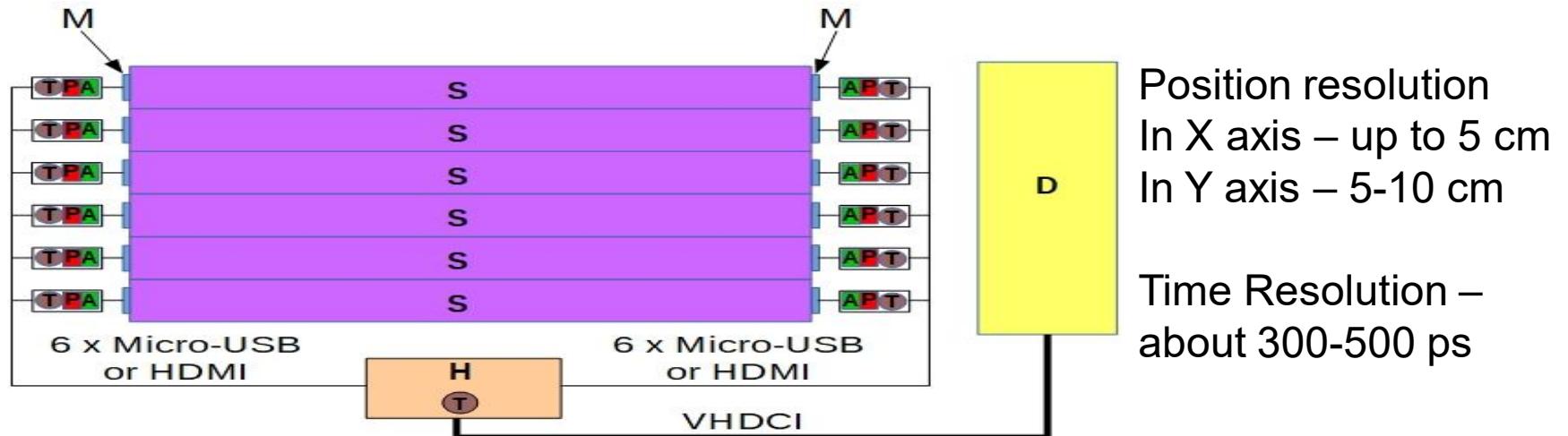
PVT or PS plastic	(No fiber)	PVT or PS plastic + WLS	
DCR (@ 5 p.e.)	~10 kcps	DCR (@ 3 p.e.):	~10 kcps
noise-noise FTR:	<0.1 cps (8,67xE-2)	noise-noise FTR:	<0.1 cps
noise-cosmic FTR:	<0.1 cps (7,81xE-2)	noise-cosmic FTR:	<0.1 cps



# 4. Design, modeling variants



## THE MUON DETECTOR SCHEME OF ANALOG SIGNAL PATH



Legend: S (violet) – plastic scintillator, M (blue) – SiPM, P (red) – power supply with temperature compensation circuit, T (brown) – temperature sensor, A (green) – amplifier, D (yellow) – MicroTCA system with ADC boards, H (orange) – Passive Signal Hub & Power Splitter.

Connector type examples:



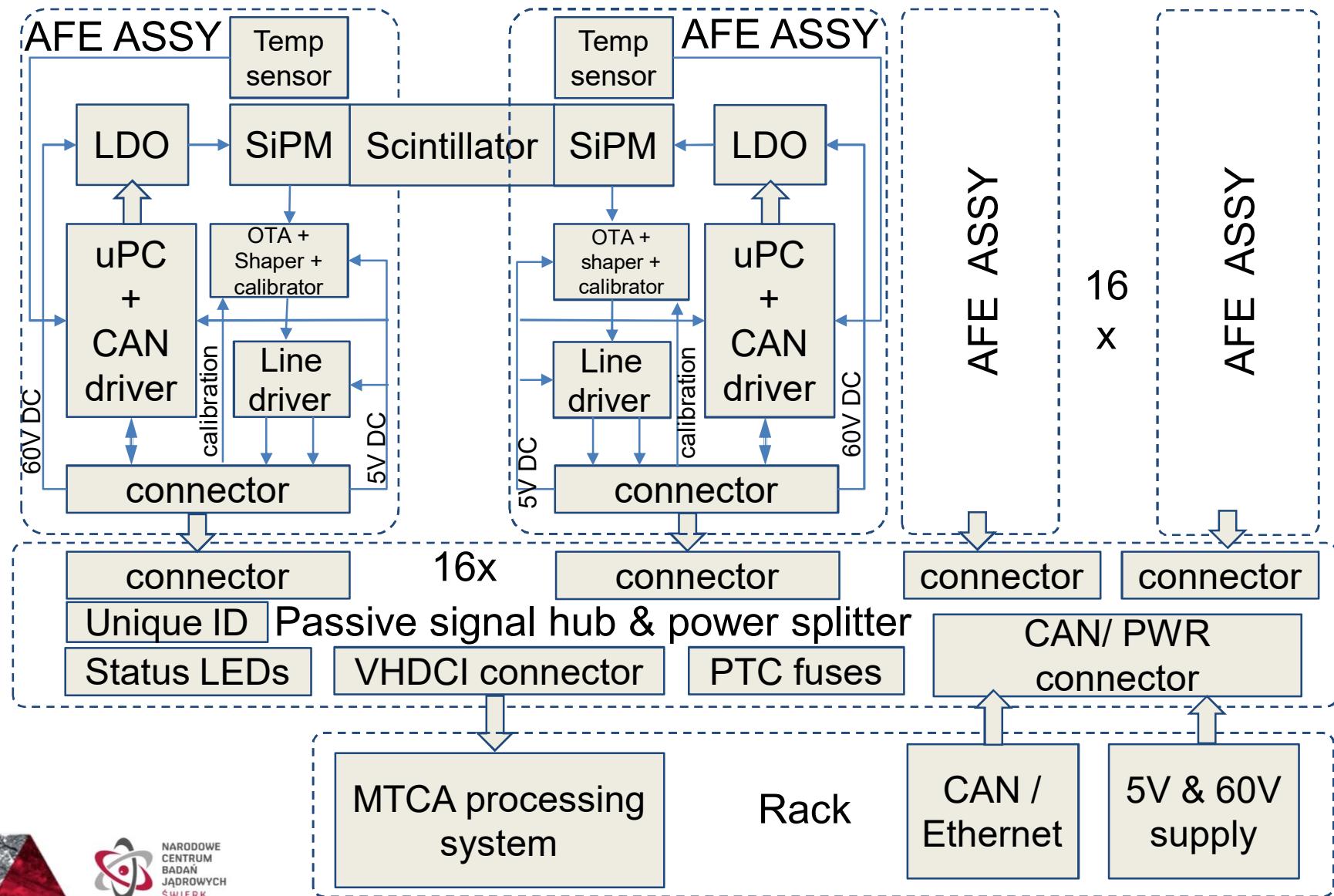
1. Rugged Micro-USB
2. C3 HDMI Rugged
3. HDMI Industrial
4. HVCDI



# 4. Design, modeling variants



## THE MUON DETECTOR SCHEME OF Analog Front END

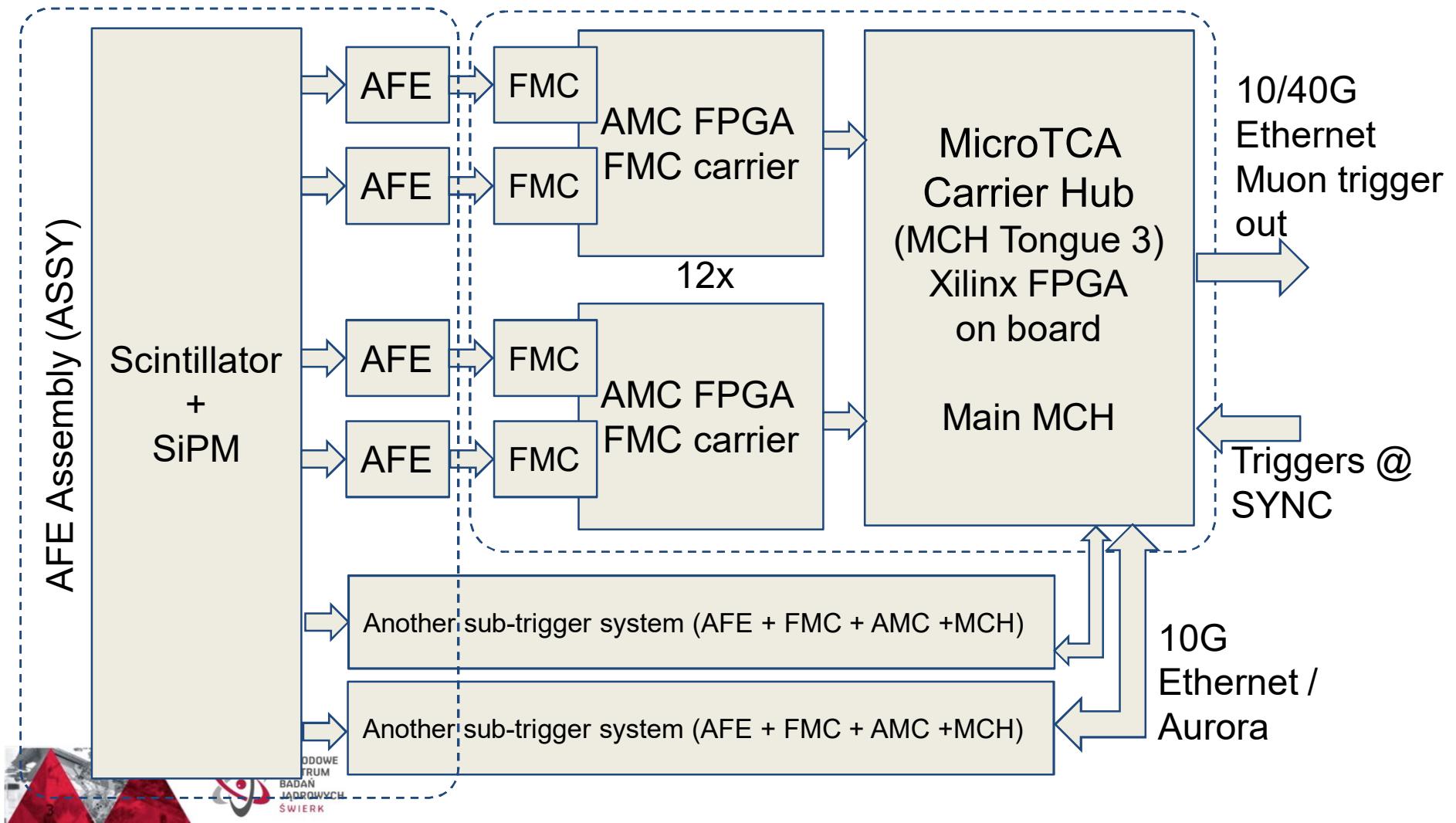


# 4. Design, modeling variants



## THE MUON DETECTOR SCHEME OF DIGITAL SIGNAL PATH

uTCA based modular muon trigger (signal flow only)



## 4. Design, modeling variants



# Analog Front End configuration

- Dedicated AFE Assembly per SiPM
- Embedded uPC + temperature sensor + LDO for SiPM set point adjust
- CAN network connectivity with unique ID chip as CAN address
- Unique ID in every hub for VHDCI cabling checking and identification
- Hardware ID for every AFE ASSY
- Low cost LDO instead of expensive switching power supply. No inductors required and lowers EMI.
- SiPM voltage, AFE current monitoring, latchup detection & protection for AFE
- Low cost shielded VHDCI cables – COTS components available as 1-10m length and custom versions
- Local passive hub with PTC fuses for 5V and 60V rails, distribution of power, CAN and signals from 16 AFE ASSY to single VHDCI cable
- Status LEDs on AFE ASSY and hub for quick fault identification
- Central power supply – custom built 2U rack box with COTS resonant 5V SMPS, 60V flyback SMPS, IEC outlets and fuses.
- CAN to Ethernet converter – standard COTS component.



## 4. Design, modeling variants



# MicroTCA (MTCA) configuration



- Standard MTCA crate (14U)  
(cable fit 1,5cm 24 channels +8)  
(additional cable for 5V and 40V power)
- Crate number depends on channel count and sampling speed

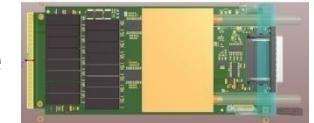
At 250MS/s: 192 channels / crate

At 125MS/s: 384 channels / crate  
(16 cables)

At 80MS/s: 576 channels / crate

At 50MS/s: 768 channels / crate

Analog Front-End module



FPGA mezzanine card (FMC)



AMC FMC carrier board



MTCA Carrier Hub



**For several MTCA modules one main MCH concentrates data from slave MCHs to generate the final muon trigger**



## 4. Design, modeling variants



# Data processing

- Latency estimation for L1 trigger (event without parameters)
  - ✓ AFE cabling 8ns/m, with 10m cabling latency is 80ns
  - ✓ ADC + SERDES latency: 400ns

Latency estimation for L2 trigger (event with parameters)

- ✓ MGT latency: 500ns
- ✓ Algorithm latency : 2-5us
- ✓ Formatter and transmitter latency: 1us

Estimated total latency: 3.5 – 7.5us

Latency estimation for L3 trigger (between MTCA systems)

- ✓ MGT latency: 500ns
- ✓ Fiber latency: 500ns + 8ns/m
- ✓ Algorithm latency : 2-5us
- ✓ Formatter and transmitter latency: 1us

Estimated total latency: 10 – 15us



# 5. MCORD Detector



## SCINTILLATORS

Number of scintillators:	660 pcs
Dimensions of scintillators:	95x25x1500 [mm]
Dimensions of detector:	100x30x1554 [mm]
Scintillators are placed in the rectangle profile	10x30x2.5 [mm]
Weight of detector:	6.5 kg
Material of scintillators casing:	Aluminum alloy

## MODULES

Number of detector in one module:	18
Number of Modules:	28
Dimensions of module:	730x90x4700 [mm]
Weight of one module:	150 kg

## SiPM/MMPC

Number of SiPMs (Channels)	1320
Number of SiPMs (with two fibers)	2640

## RESOLUTION

Position resolution: In X axis – up to 5 cm, In Y axis – 5-10 cm

Time Resolution – about 300-500 ps

Number of events (particles):	about 100-150 per sec per m <sup>2</sup>
Calculated Coincidence factor:	about 98%



## 5. MCORD Team

30 people

Including

18 scientists (professors and doctors)

12 engineers

from

2 universities and 3 science institutes

With knowledge and experience in digital and analog electronics, scintillators and photodetectors, nuclear and astrophysics, data acquisition and analysis, simulation and experiments.

**Our group is a member of Polish consortium NICA-PL**



NCBJ  
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## 5. Conclusions

1. Cosmic Ray Detector is necessary for good calibration of TPC, TOF and ECAL, MPD detectors before completion of MPD .
2. Cosmic ray detector is helpful for better calibration of TPC TOF, before each experimental session.
3. Additionally MCORD can be used for astrophysics observations similar to past collider experiments. In our case, especially for investigations of near horizontal muon bundles (research of main trivial mechanism of multi-muon event generation (EAS muons)).
4. Our team has a realistic plan and is capable of building this detector.
5. Projected cosmic ray detector will be designed to have required time resolution and position accuracy



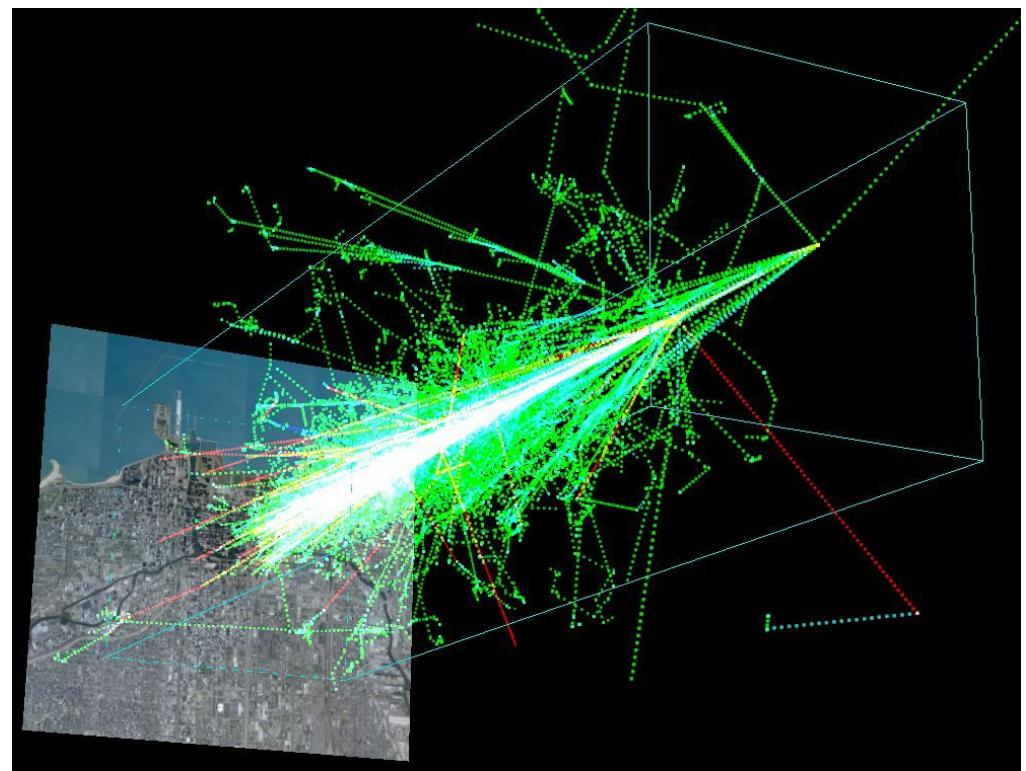
**Our group is a member of the Polish consortium NICA-PL**





# Polish consortium NICA-PL

## Thank You for Attention

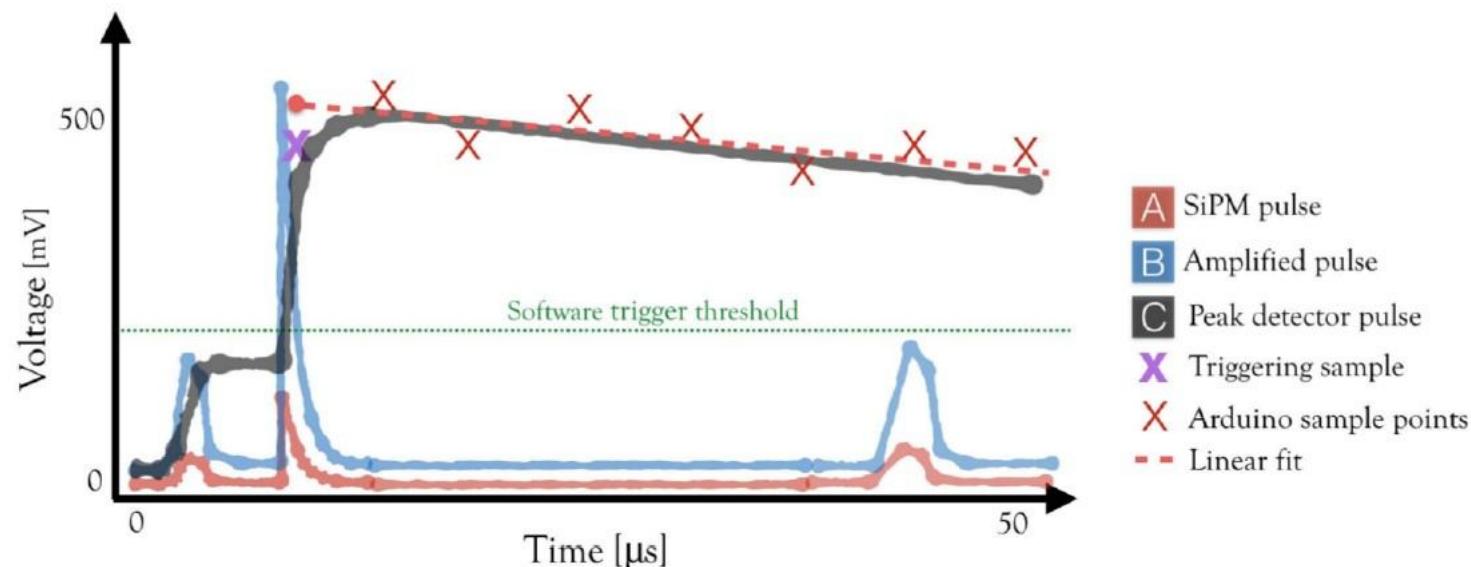


## 4. Design, modeling variants

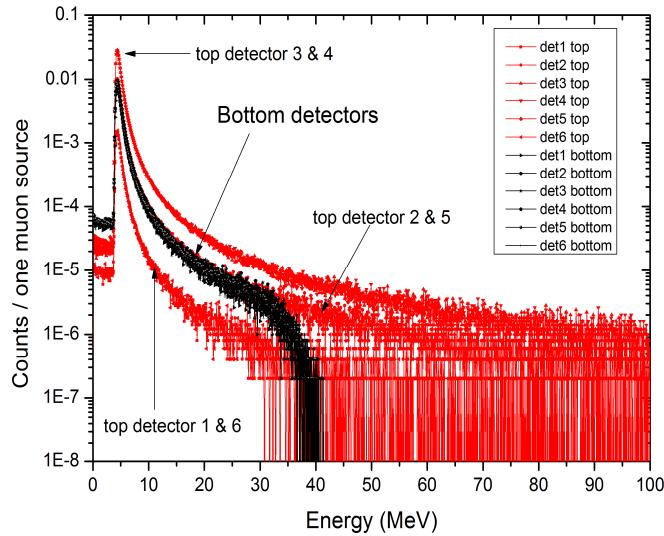
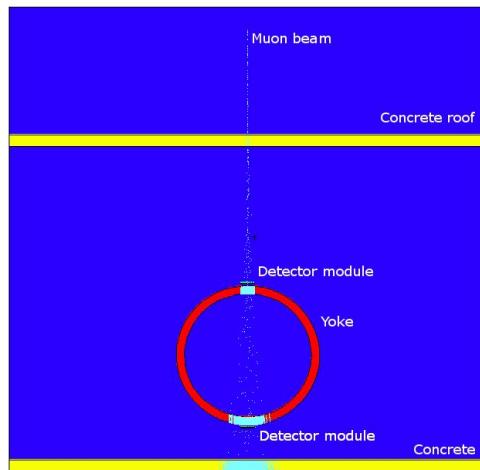


# Data processing

- FPGA gateware:
  - ✓ Time synchronisation: White Rabbit-based sub-ns
  - ✓ Pulse position estimator (advanced DSP processing)
  - ✓ Local triggering
  - ✓ Data quality monitoring
  - ✓ Data acquisition
  - ✓ Data formatter and transmitter



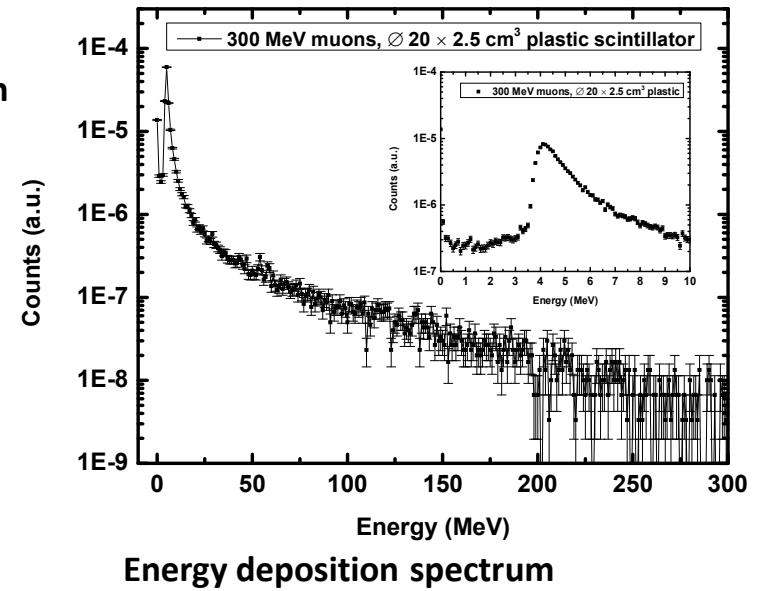
# Conceptual design – simulations



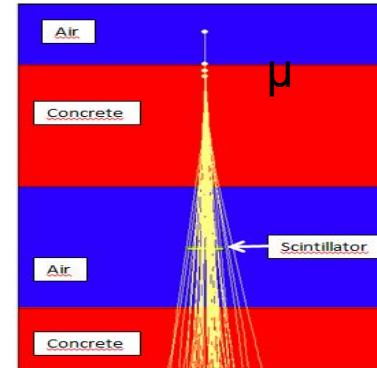
Energy deposition spectra of 2 GeV monoenergetic muons

## The first MCNPX simulation of muon energy deposition in plastic scintillator for NICA Energy spectrum reconstruction

- Cylindrical plastic scintillator  
 $\varnothing 20 \times 2.5 \text{ cm}^3$  – the same used for experiment at NCBJ
- The maximum energy deposition centroid is around 4 MeV - reasonable result, assuming 2 MeV/cm  $dE/dx$  energy loss for muons in organic scintillator. Light generation due to ionization in plastic scintillator will be implemented in the near future.



Energy deposition spectrum



Geometry used (40000 tracks)



## 5. MCORD – estimated cost



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## 4. Design, modeling variants

MicroTCA and other in the Slow Control Racks



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