

# Performance of the Pixel Luminosity Telescope for Luminosity Measurement at CMS during Run2

**NEC2019 - 30 Sept-04 Oct 2019 - Becici**

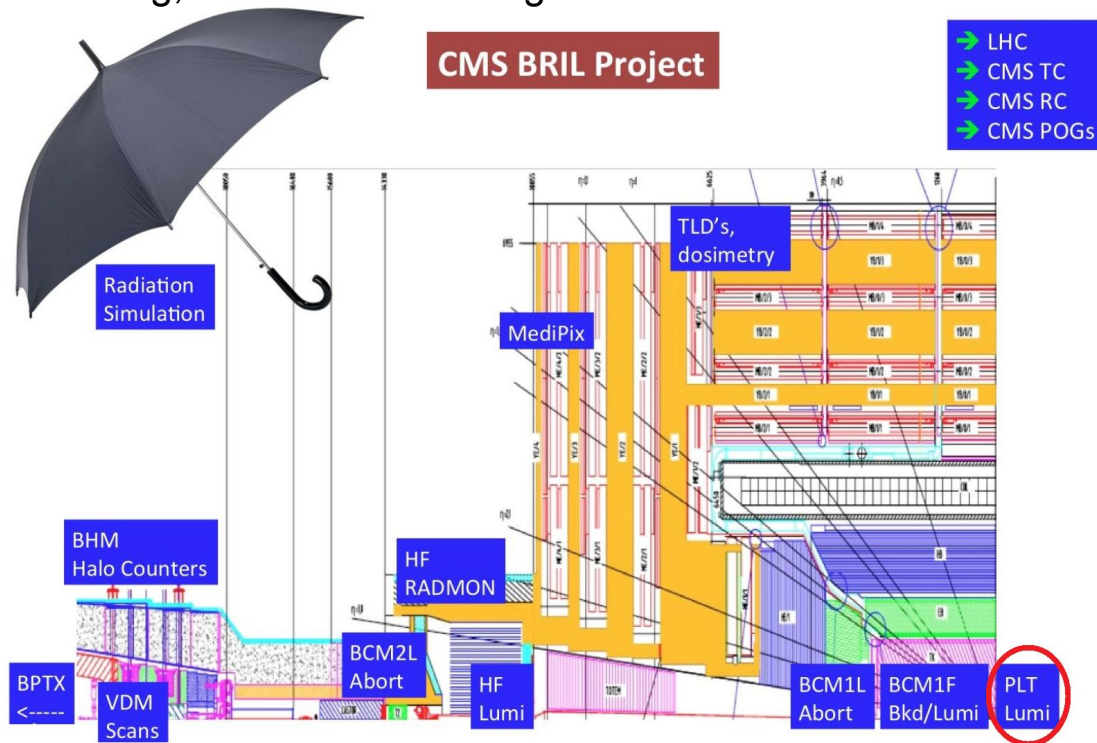
**Francesco Romeo**  
on behalf of the CMS collaboration



VANDERBILT  
UNIVERSITY

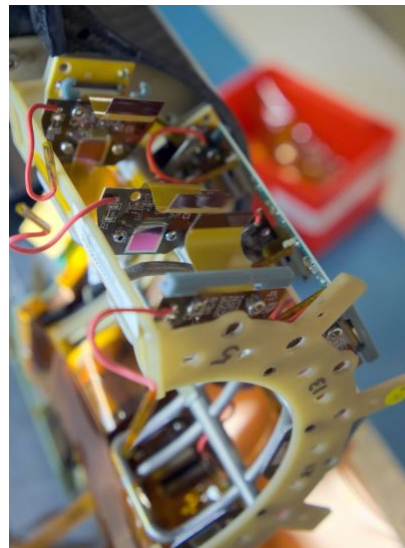
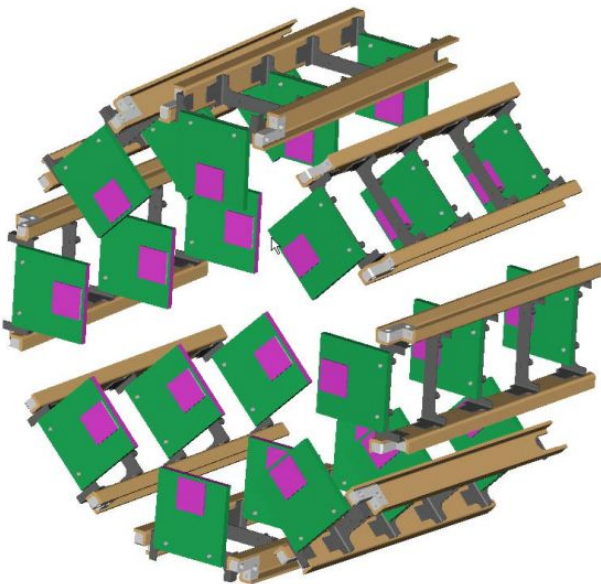
# Motivations and overview

- Precision **measurement of the luminosity** is extremely important
  - Online: real-time feedback on LHC performance and operation
  - Offline: crucial component of every physics analysis (cross-section, coupling, upper limit, ...)
- **BRIL (Beam Radiation, Instrumentation, and Luminosity)** group oversees luminosity measurements, beam condition monitoring, radiation monitoring and simulation



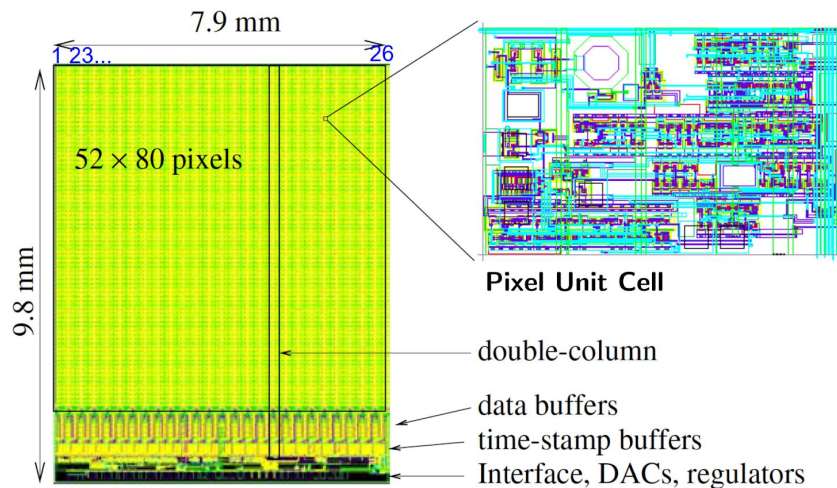
# Pixel Luminosity Telescope (PLT)

- Silicon pixel detector, installed at CMS in 2015 for LHC Run2
- Located 1.75m from the IP (interaction point) at both ends of CMS ( $|\eta| \approx 4.2$ )
- 16 telescopes (8 one side + 8 other side of CMS)
  - Each telescope is 7.5 cm long, 5 cm away from the beam pipe, and contains 3 individual sensors
  - Sensor and PSI46v2 readout chips (ROCs) developed for the CMS phase-0 Pixel detector

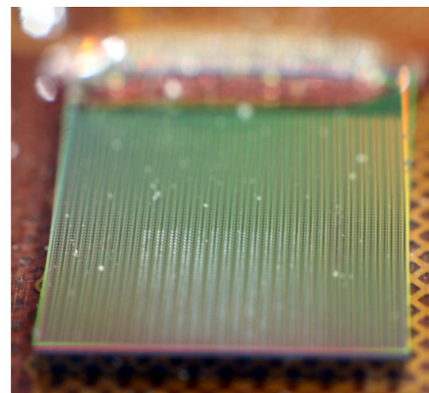


# Sensor and readout chips

- **n-in-n silicon sensor**. Pixel size:  $150 \times 100 \mu\text{m}^2$  and active thickness of  $285 \mu\text{m}$
- **PSI46v2 ROC**
  - records hit position and amount of charge deposited in the silicon sensor
  - time resolution of 25 ns
  - active area is an array of  $52 \times 80$  pixel unit cells (26 double columns of 160 pixels each)



Floor plan of PSI46v2 readout chip



- **ROC read out mode**
  - **fast-or**: info on double columns with pixels above threshold in a given 25 ns bunch crossing
  - **full pixel**: dedicated trigger at rate of  $\sim 3$  kHz for additional studies

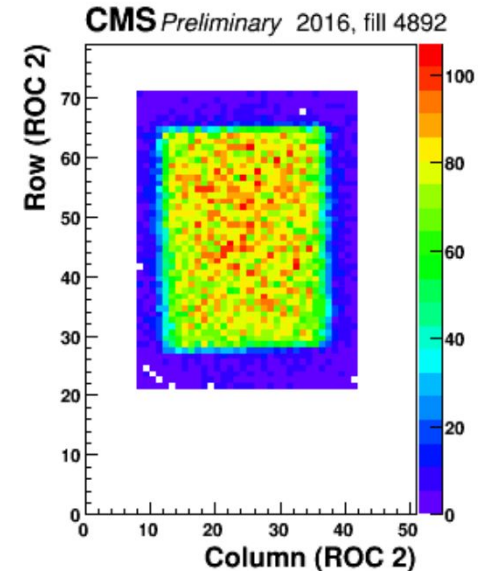
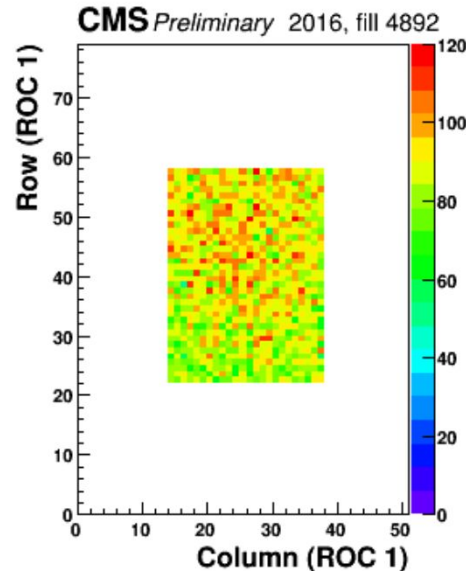
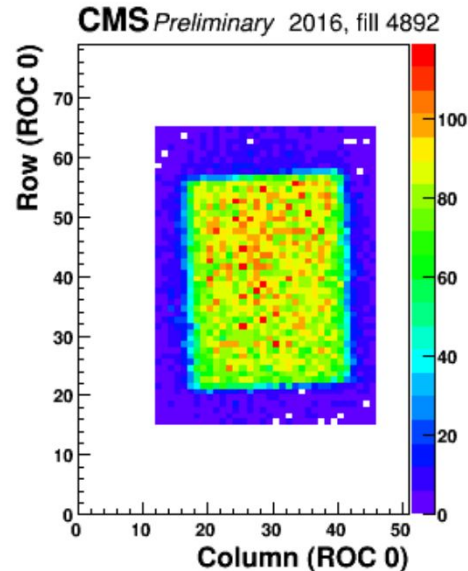
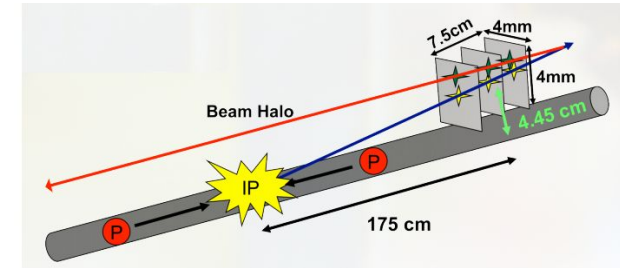
# Luminosity measurement

The PLT provides **bunch-by-bunch luminosity** with statistical precision of  $\sim 1\%$  every 1.5 s

1. **Detect “triple coincidences”**:  
events where each plane registers at least one fast-or hit
2. **Use “Zero-counting” method**:  
count fraction of events where no triple coincidence is found and then use  $L \sim -\ln \langle p_0 \rangle$   
avoid systematic issues from events with two or more tracks
3. **Calibrate with “Van der Meer” scan**:  
calibrate absolute luminosity
4. Luminosity **corrections**:  
“**accidentals**” events where we have a triple coincidence but not from a real track from the IP  
(either combination of random hits, or track from other source)  
“**linearity/slope**”: detector response

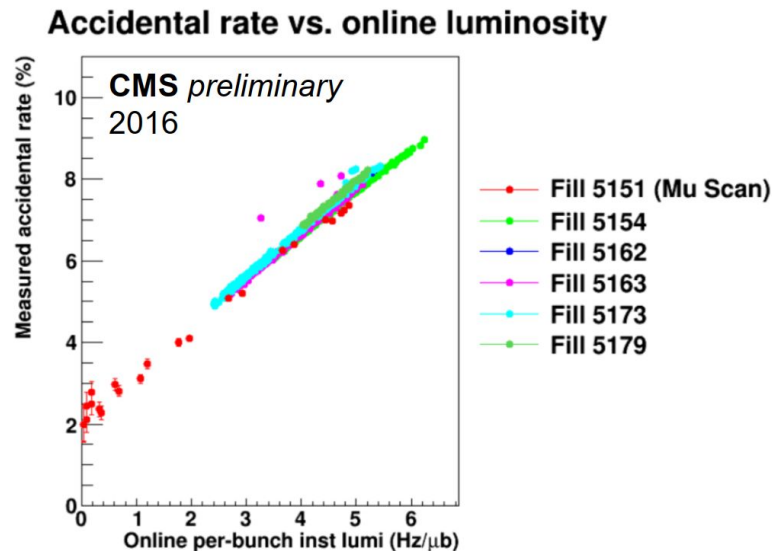
# Track, occupancy, and accidentals

- Using full pixel readout, look at events in a single telescope, and select only hits which can be reconstructed as a single track
- The **active area** of the sensor planes is
  - $3.6 \times 3.6$  ( $3.9 \times 3.8$ ) mm<sup>2</sup> on the center (outer planes)
  - > to **reduce** the contribution from **accidentals**
  - **displaced** (1st, 3d) or **centered** (2nd)
  - > **correct** for **alignment** effects



# Corrections of accidentals

- “**Accidentals**” are triple coincidences observed not caused by a track originating from the IP:
  - tracks from particles elsewhere (e.g., beam halo, or activated material surrounding PLT)
  - random coincidence of hits from different tracks (for example, two tracks that do not individually pass through all three planes of the PLT could combine to make a triple coincidence)
- **Algorithm:**
  - look at distribution of tracks in reconstructed slopes
  - measure the mean slope and fit residual
  - reject any track more than 5 standard deviations away from the mean



# Van der Meer scan

The luminosity is calibrated with  $\sigma_{\text{vis}}$  measured with the Van der Meer scan

- Assume beams with a Gaussian profile  $L = N_1 N_2 v_{\text{orb}} N_b / 2\pi \Sigma_x \Sigma_y$   

$N_1, N_2$   
beam intensities

$N_b$   
# of colliding bunches

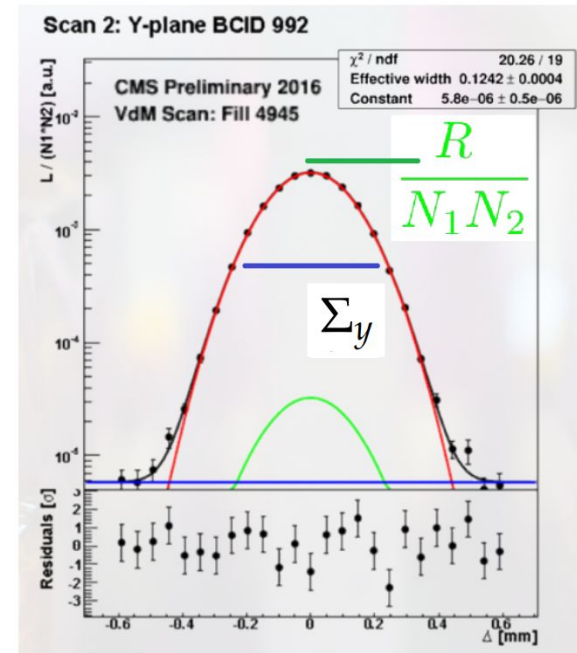
$v_{\text{orb}}$   
orbit frequency

$\Sigma_x, \Sigma_y$   
effective beam sizes
- Gradually vary beam separation in X and Y directions
- The resulting luminosity curve is fit with a double (2016, 2017) or single (2018) Gaussian, and the background either modelled with a constant term (2016), or subtracted (2017, 2018)

$$\sigma_{\text{vis}} = R / N_1 N_2 2\pi \Sigma_x \Sigma_y / v_{\text{orb}}$$

$N_1 N_2$  provided by LHC  
from fit extract

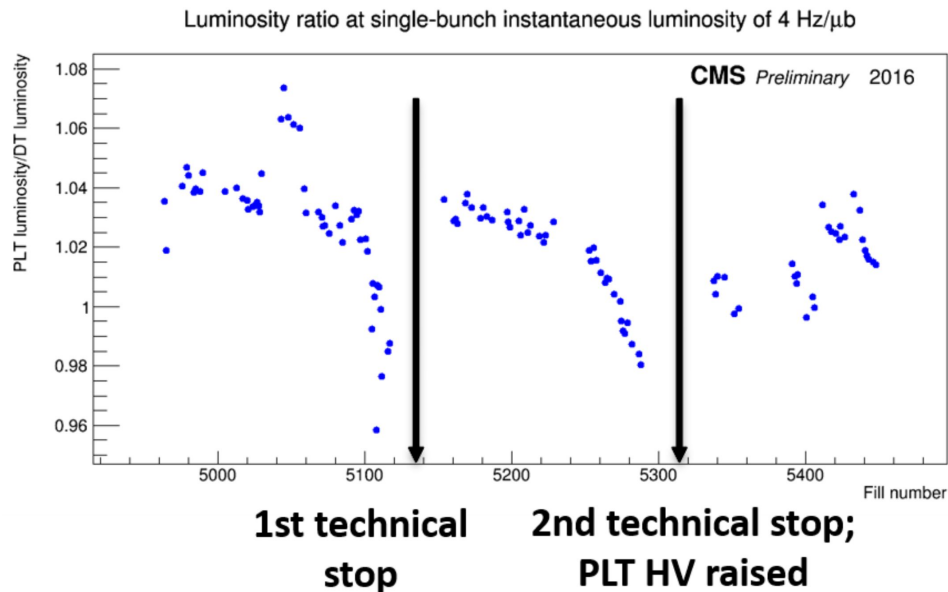
- peak:  $R / N_1 N_2$
- profile width:  $\Sigma_x$  and  $\Sigma_y$



# Relative rate monitoring, and detector performance

- PLT rates are periodically compared to other CMS luminosity detectors
- Relative change in measured rates between detectors can be indicative of issues, e.g. radiation damage from accumulated dose

Ratio of PLT luminosity to luminosity from muon rate in the drift tubes (DT) per fill at a given single-bunch instantaneous luminosity (SBIL)



- Drop in ratio resulted from increased radiation damage and mitigated by
  - increase in HVs (2016-2017)
  - adjust the thresholds used in the PLT readout chips to determine hit pixels (2018)

# Corrections due to detector performance

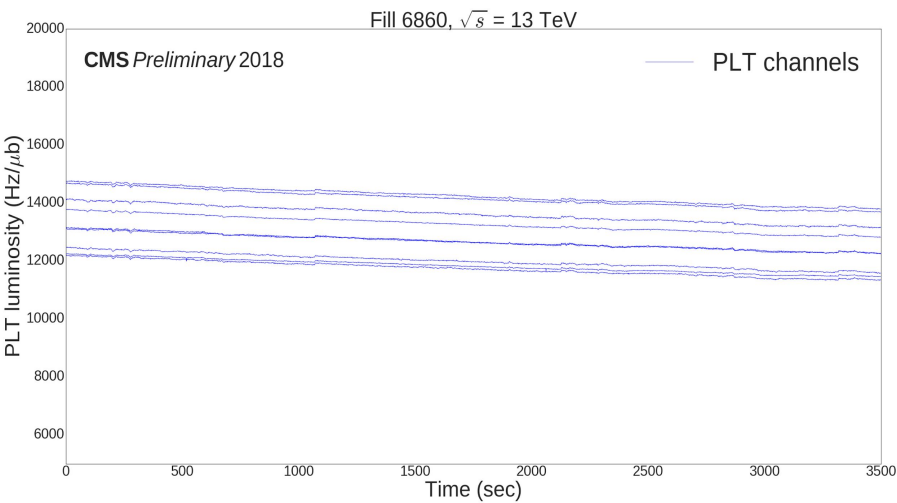
Ideal luminometer would have a perfectly linear response as a function of the instantaneous luminosity with a single calibration constant relating the two. In reality, need to take into account corrections for:

- linearity (Slope)
- efficiency (Eff)

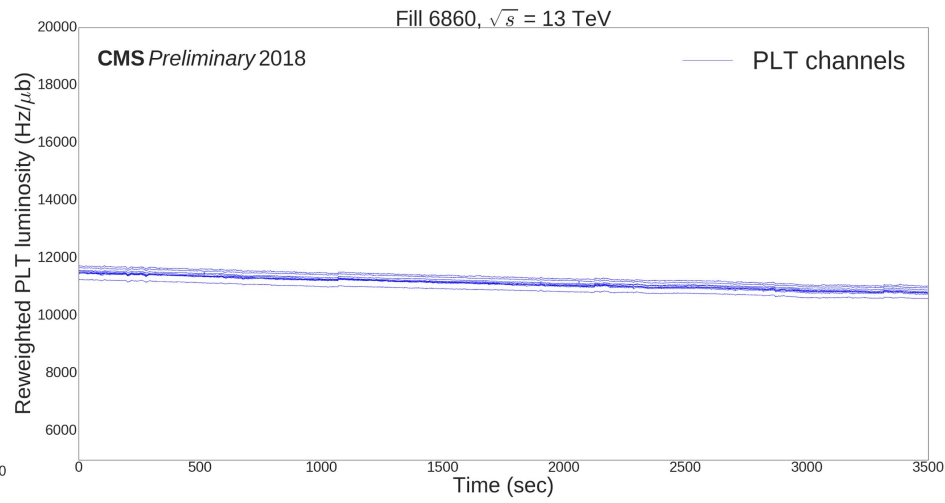
Use **emittance scans (ES)** (mini Van der Meer scans at the beginning and end of a fill)

- **linearity**:  $\sigma_{\text{ES}}^{\text{ES}}$  vs SBIL as individual bunches within a single fill can exhibit substantially different SBIL
- **efficiency**:  $\sigma_{\text{ES}}^{\text{vis}} / \sigma_{\text{vdM}}^{\text{vis}}$
- Linearity and efficiency corrections are applied channel by channel, and weights are measured separately for leading and trailing bunches:  $\text{SBIL}_{\text{corr}} = (\text{SBIL}_{\text{raw}} / \text{Eff}) - (\text{SBIL}_{\text{raw}} / \text{Eff})^2 \cdot \text{Slope} / 100$

No correction applied



Correction applied (much better agreement across channel)



# Conclusions

- Luminosity measurement very important at CMS  
Luminosity studies under BRIL group, which includes several dedicated detectors
- Pixel Luminosity Telescope luminometer has been operating successfully during the LHC Run2
- PLT provides online and offline luminosity measurement with high precision
- With the current Long Shutdown 2, PLT is under upgrades to substitute ROCs, sensors, and other components close to their expected end-of-life, mainly due to radiation damage
- Look forward to excellent data-taking within LHC Run 3 at CMS

# **Back up slides**

# Readout scheme for a single PLT quarter

