# Quark Matter under Rotation 



## Kenji Fukushima

The University of Tokyo

- Theory of Hadronic Matter under Extreme Conditions -


## Rotation in Collisions



## $L$ remains longer than $B$



## $L$ is ubiquitous in the nature Deformed Nuclei Neutron Stars Electron Vortices ...

## Rotation in Collisions

## $L$ is an intrinsic property of matter



## Rotation in Collisions

## Global Polarization of $\Lambda$



$$
\begin{aligned}
& P_{\text {Vortical }}=\frac{1}{2}\left(P_{\Lambda}+P_{\bar{\Lambda}}\right) \\
& P_{\text {Magnetic }}=\frac{1}{2}\left(P_{\Lambda}-P_{\bar{\Lambda}}\right)
\end{aligned}
$$

## Becattini, Csernai, Wang, ...

September 16,2019@ JINR, Dubna

## Theoretical Formulation



## Fluid

## Rotating QFT



Coordinate Transformation Finite Size (causality)

## Theoretical Formulation



## Fluid


$\nabla \times u$

## Rotating QFT



## Coordinate Transformation Finite Size (causality)

## Theoretical Formulation

$$
\left[i \gamma^{\mu}\left(\partial_{\mu}+\Gamma_{\mu}\right)-m\right] \psi=0
$$

$$
g_{\mu \nu}=\left(\begin{array}{cccc}
1-\left(x^{2}+y^{2}\right) \Omega^{2} & y \Omega & -x \Omega & 0 \\
y \Omega & -1 & 0 & 0 \\
-x \Omega & 0 & -1 & 0 \\
0 & 0 & 0 & -1
\end{array}\right)
$$

Solve this in a finite cylinder (radius R)
Not only the affine connection but gamma's changed
Vierbeins are needed !

## Theoretical Formulation

Theoretical treatment for deformed nuclei
Cranking Hamiltonian $\quad H_{\text {rot }}=H-\omega J_{z}$
Chemical Potential Like


Angular Momentum

## Theoretical Formulation

Theoretical treatment for deformed nuclei
Cranking Hamiltonian $\quad H_{\text {rot }}=H-\omega J_{z}$
Chemical Potential Like


Angular Momentum
$\mu$

## Rotation ~ Density

Chen-KF-Huang-Mameda, PRD (2015)

## Inverse Magnetic Catalysis driven by rotation


More interestingly,
Rotation $+\boldsymbol{B}=($ Genuine) Density

$$
n=-\left.\frac{\partial \Omega}{\partial \mu}\right|_{\mu=0}=\frac{e B \omega}{4 \pi^{2}}
$$

interpreted as anomaly
Hattori-Yin, PRL (2016)
Can be given another interpretation from the Floquet theory

## Rotation ~ Density

Jiang-Liao, PRL (2017)



Completely analogous to chemical potential... BUT!

## Rotation ~ Density

## Is it really possible to change the QCD vacuum just by rotation ???

The answer is negative:
Ebihara-Fukushima-Mameda, PLB (2017)
Causality System size should be finite $\sim \boldsymbol{R}$
$\omega R<1 \quad$ Energy dispersion should be gapped $\sim \boldsymbol{J} / \boldsymbol{R}$
Induced chemical potential $\sim \omega J$
Gap is always bigger than the chemical potential

## Rotation ~ Density

## Is it really possible to change the QCD vacuum just by rotation ???

The answer is negative:
Ebihara-Fukushima-Mameda, PLB (2017)
If one wants to see nontrivial effects of rotation, it should be coupled with...
$\mu_{\text {Gauge CVE }}^{\mu}$
$B$
Chiral Pumping Effect
T
Gravity CVE

# There are still lots of interesting challenges in physics and theoretical computations! 

But, these are mostly technical issues, and there seems to be no conceptual problem.

Let's move on to a more subtle thing now...

## Switch the gear into...




## Two Choices

Hydrodynamics with Local Vorticity Vectors
Derivative expansion? (vorticities are second order)
Discrimination of $L$ and $S$ ?

Kinetic Equations with Local Vorticity Vectors

$$
\begin{aligned}
& \varepsilon_{\mathrm{rot}}=p-\underline{\boldsymbol{\omega} \cdot(\boldsymbol{x} \times \boldsymbol{p}+\hbar \lambda \hat{\boldsymbol{p}})} \\
& =\omega \cdot \boldsymbol{\omega} \\
& f(\varepsilon) \rightarrow f\left(\varepsilon_{\mathrm{rot}}\right) \quad \text { Corrections in the Kinetic Eqs.? }
\end{aligned}
$$

## Barnett Effect

## 

## Gyromagnetic Effect



September 16, 2019 @ JINR, Dubna

## Barnett Effect

 Gyromagnetic Effect


## Barnett Effect

## 

 Spin Alignment in response to Rotation"Gyroscopic" Motion



## Barnett Effect

## Quickest Derivation

$$
\omega \cdot J=\mu \cdot B
$$

Magnetization

$$
\boldsymbol{M}=\underset{\text { magnetic susceptibility }}{\chi_{B} \boldsymbol{B}}
$$

Magnetic moment $\boldsymbol{\mu}=\underset{\text { gyromagnetic ratio }}{\gamma \boldsymbol{J}}$


Roughly speaking, the Barnett effect is a transport from the orbital to the spin angular momentum.

To make this phenomenon well-defined, the orbital and the spin components must be well separated.

## HOW?

## Decomposition of $L$ and $S$

Angular Momentum
$=$ Noether Current from Rotational Symmetry

$$
J^{\lambda \mu \nu}=\int_{\bar{\psi} i \hbar\left(\gamma^{\lambda} x^{\mu} \partial^{\nu}-\gamma^{\lambda} x^{\nu} \partial^{\mu}\right) \psi}^{\sum_{\frac{1}{2}} \bar{\psi} i \hbar \gamma^{\lambda} \gamma^{\mu} \gamma^{\nu} \psi}
$$

Neither $L$ nor $S$ conserved separately

$$
\partial_{\lambda} L^{\lambda \mu \nu}=-\partial_{\lambda} S^{\lambda \mu \nu}=\bar{\psi} i \hbar\left(\gamma^{\mu} \partial^{\nu}-\gamma^{\nu} \partial^{\mu}\right) \psi
$$

## Decomposition of $L$ and $S$

## Different decomposition

$$
\begin{aligned}
& \tilde{L}^{\lambda \mu \nu}=\frac{1}{2} L^{\lambda \mu \nu}+\frac{1}{2} \bar{\psi} i \hbar\left[\left(x^{\mu} \gamma^{\nu}-x^{\nu} \gamma^{\mu}\right) \partial^{\lambda}\right] \psi \\
& \tilde{S}^{\lambda \mu \nu}=J^{\lambda \mu \nu}-\tilde{L}^{\lambda \mu \nu}
\end{aligned}
$$

$$
\partial_{\lambda} \tilde{L}^{\lambda \mu \nu}=\partial_{\lambda} \tilde{S}^{\lambda \mu \nu}=0
$$

Separately
conserved?
Belinfante angular momentum (Only the orbital part remains, and the spin part turns out to be trivial...)

## Decomposition of $L$ and $S$

We believe the former decomposition makes sense:

1) Reduced to ordinary $L$ and $S$ in non-rela limit
2) $S$ is related to the axial current

$$
S^{0 i j}=\epsilon^{i j k} \frac{\hbar}{2} \bar{\psi} \gamma^{k} \gamma_{5} \psi=\epsilon^{i j k} \frac{j_{5}^{k}}{2}
$$

Corresponding Spin Operator:

$$
\boldsymbol{S} \rightarrow \hbar \lambda\left(\hat{\boldsymbol{p}}-\hbar \lambda \frac{\hat{\boldsymbol{p}}}{2 p} \times \boldsymbol{\nabla}\right)
$$

Chen-Son-Stephanov, PRL (2015)
Torque from gyromagnetic effect

## Relativistic Barnett Effect

Spin Expectation Value
Energy in a rotating fluid $\varepsilon_{\mathrm{rot}}=p-\boldsymbol{\omega} \cdot(\boldsymbol{x} \times \boldsymbol{p}+\hbar \lambda \hat{\boldsymbol{p}})$

$$
\langle\boldsymbol{S}\rangle=\int_{\boldsymbol{p}} \lambda \hbar\left(\hat{\boldsymbol{p}}-\lambda \hbar \frac{\hat{\boldsymbol{p}}}{2 p} \times \nabla\right) f\left(\varepsilon_{\mathrm{rot}}\right)
$$

Vilenkin (1978)

$$
=-\hbar \lambda(\boldsymbol{\omega} \times \boldsymbol{x}) \int_{\boldsymbol{p}} \frac{p}{3} f^{\prime}(p)-\hbar^{2} \boldsymbol{\omega} \int_{\boldsymbol{p}} f^{\prime}(p)
$$

$\begin{aligned} &\langle\boldsymbol{S}\rangle \\ & \perp \text { "Transverse" Barnett Effect } \begin{aligned} \text { Chiral Vortical Effe } \\ \sim\end{aligned} \\ & \sim \text { Barnett Effect }\end{aligned}$

## Relativistic Barnett Effect

Spin Expectation Value

$$
\begin{gathered}
\langle\boldsymbol{S}\rangle_{\perp}=-\hbar \sum_{R, L} \lambda(\boldsymbol{\omega} \times \boldsymbol{x}) \int_{\boldsymbol{p}} \frac{p}{3} f_{\lambda}^{\prime}(p) \\
=\frac{\hbar}{2}(\boldsymbol{\omega} \times \boldsymbol{x}) \int_{\boldsymbol{p}}\left[f_{R}(p)-f_{L}(p)\right]=\frac{\hbar}{2}(\boldsymbol{\omega} \times \boldsymbol{x}) n_{5} \\
\boldsymbol{j}_{5}=n_{5} \boldsymbol{v} \\
\text { Transverse Barnett appears } \\
\text { for massless and chirally } \\
\text { imbalanced fermions }
\end{gathered}
$$

## Relativistic Barnett Effect

Magnetic moment

$$
\langle\boldsymbol{\mu}\rangle=\left\langle\boldsymbol{\mu}_{L}\right\rangle+\left\langle\boldsymbol{\mu}_{S}\right\rangle
$$

$$
=\left\langle\boldsymbol{\mu}_{L}\right\rangle_{\mathrm{mech}}-\hbar \lambda \frac{q_{e}}{6}(\boldsymbol{\omega} \times \boldsymbol{x}) \int_{\boldsymbol{p}} f^{\prime}(p)
$$


up to $O\left(\hbar^{1}\right)$


Eddy magnetic moment
Fukushima-Pu-Qiu, PRA (2018)

## Eddy Magnetization



# Possible Evidence for Free Precession of a Strongly Magnetized Neutron Star in the Magnetar 4U 0142+61 

K. Makishima, ${ }^{1,2,3}$ T. Enoto,,${ }^{4,5}$ J. S. Hiraga, ${ }^{2}$ T. Nakano, ${ }^{1}$ K. Nakazawa, ${ }^{1}$ S. Sakurai, ${ }^{1}$ M. Sasano, ${ }^{1}$ and H. Murakami ${ }^{1}$<br>${ }^{1}$ Department of Physics, Graduate School of Science, the University of Tokyo, 7-3-1 Hongo, Bunkyo-ku, Tokyo 113-0033, Japan<br>${ }^{2}$ Research Center for the Early Universe, Graduate School of Science, the University of Tokyo, 7-3-1 Hongo, Bunkyo-ku, Tokyo 113-0033, Japan<br>${ }^{3}$ MAXI team, RIKEN, 2-1 Hirosawa, Wako, Saitama 351-0198, Japan<br>${ }^{4}$ High Energy Astrophysics Laboratory, RIKEN Nishina Center, 2-1 Hirosawa, Wako, Saitama 351-0198, Japan<br>${ }^{5}$ NASA Goddard Space Flight Center, Astrophysics Science Division, Code 662, Greenbelt, MD 20771, USA<br>(Dated: February 19, 2018)

Magnetars are a special type of neutron stars, considered to have extreme dipole magnetic fields reaching $\sim 10^{11} \mathrm{~T}$. The magnetar $4 \mathrm{U} 0142+61$, one of prototypes of this class, was studied in broadband X-rays ( $0.5-70 \mathrm{keV}$ ) with the Suzaku observatory. In hard X-rays ( $15-40 \mathrm{keV}$ ), its 8.69 sec pulsations suffered slow phase modulations by $\pm 0.7 \mathrm{sec}$, with a period of $\sim 15$ hours. When this effect is interpreted as free precession of the neutron star, the object is inferred to deviate from spherical symmetry by $\sim 1.6 \times 10^{-4}$ in its moments of inertia. This deformation, when ascribed to magnetic pressure, suggests a strong toroidal magnetic field, $\sim 10^{12} \mathrm{~T}$, residing inside the object. This provides one of the first observational approaches towards toroidal magnetic fields of magnetars.

PRL112, 171102 (2014)

September 16, 2019 @ JINR, Dubna

## Eddy Magnetization

Precession has been measured


Spherically non-symmetric moment of inertia inferred

Deformation is assumed to be sustained by (toroidal) field energy

$$
B \sim 10^{16} \text { gauss }
$$

Stronger than the surface $\boldsymbol{B}$ of magnetar !
How can it be created? $\leftarrow$ Chiral Barnett Effect?

## Subtlety in Hydro

## Can we compute the same quantities in hydro?

## YES, BUT NO!

$$
\begin{aligned}
& T_{\text {hydro }}^{\mu \nu}=(E+P) u^{\mu} u^{\nu}-P g^{\mu \nu}+\hbar n_{5}\left(u^{\mu} \omega^{\nu}+u^{\nu} \omega^{\mu}\right) \\
& L_{\text {hydro }}^{i j}=x^{i} T_{\text {hydro }}^{0 j}-x^{j} T_{\text {hydro }}^{0 i}
\end{aligned}
$$

$$
\boldsymbol{L}_{\mathrm{hydro}}=(E+P)(\boldsymbol{x} \times \boldsymbol{u})-\hbar n_{5}(\boldsymbol{\omega} \times \boldsymbol{x})
$$

Is this consistent with kinetic theory results?

## Subtlety in Hydro

$$
\begin{aligned}
& \boldsymbol{L}_{\mathrm{hydro}}=(E+P)(\boldsymbol{x} \times \boldsymbol{u})-\hbar n_{5}(\boldsymbol{\omega} \times \boldsymbol{x}) \\
& \boldsymbol{x} \times(\boldsymbol{\omega} \times \boldsymbol{x}) \frac{4}{3} \int_{\boldsymbol{p}} p\left(f_{R}+f_{L}\right) \\
& =\langle\boldsymbol{L}\rangle_{\text {mech }} \quad \text { OK } \\
& =2\langle\boldsymbol{S}\rangle_{\perp} \\
& \text { Owice bigger !? }
\end{aligned}
$$

The difference comes from the energy momentum tensor. In hydro the energy momentum tensor is a symmetrized one. Belinfante form? Should be pseudo-gauge invariant...

Becattini-Florkowski-Speranza, PLB (2019)

## Summary of the Talk

## Rotation ~ Density

$\square$ Phase Diagram
$\square$ Finite-size System / Inhomogeneous Condensates
Rotation ~ Magnetic Field
$\square$ Barnett Effect expected
$\square$ Chiral Vortical Effect is nothing but the Barnett Effect.
$\square$ Decomposition of $L$ and $S$ still assumed...
$\square$ Ideas testable in optics and electron vortex systems


Works along these lines ongoing

