

Particle physics: present and prospects

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- Three gauged symmetries SU(3)xSU(2)xU(1)
- Fixe families of quarks and leptons (3x2, 3x1, 1x2, 1x1)
- Brout-Englert-Higgs mechanism of spontaneous EW symmetry breaking -> Higgs boson
- CKM and PMNS mixing of flavours
- CP violation via phase factors
- Confinement of quarks and gluons inside hadrons
- Baryon and lepton number conservation
- CPT invariance -> existence of antimatter

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- Majorana or Dirac nature of neutrino
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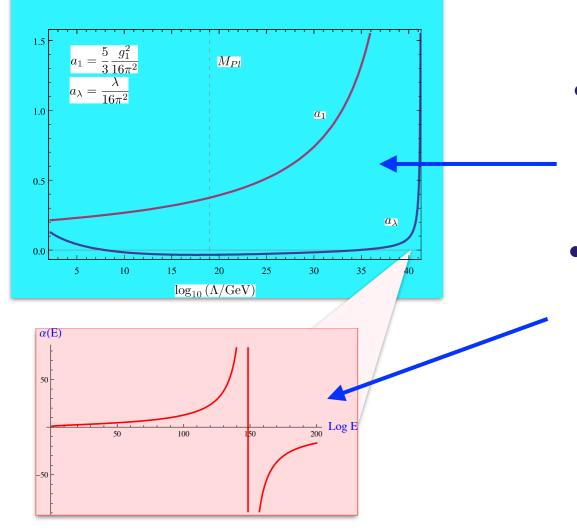
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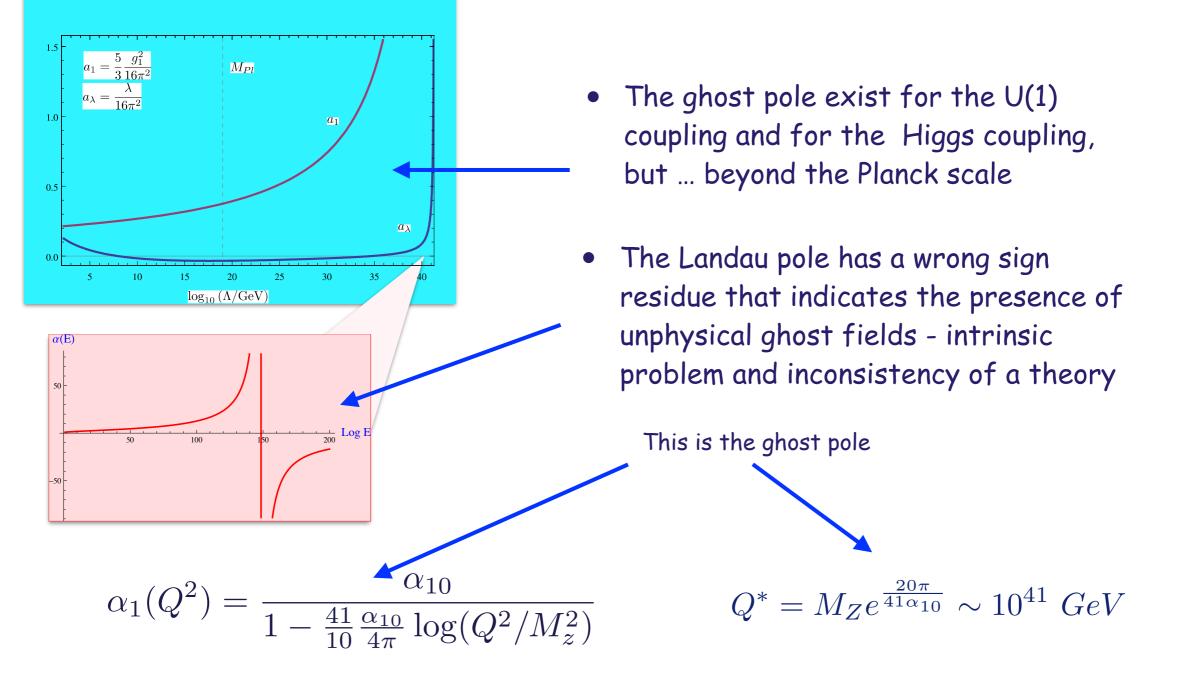
- how confinement actually works ?
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- how CP violation occurs in the Universe?
- how to protect the SM from would be heavy scale physics?

First Strain States and the second se

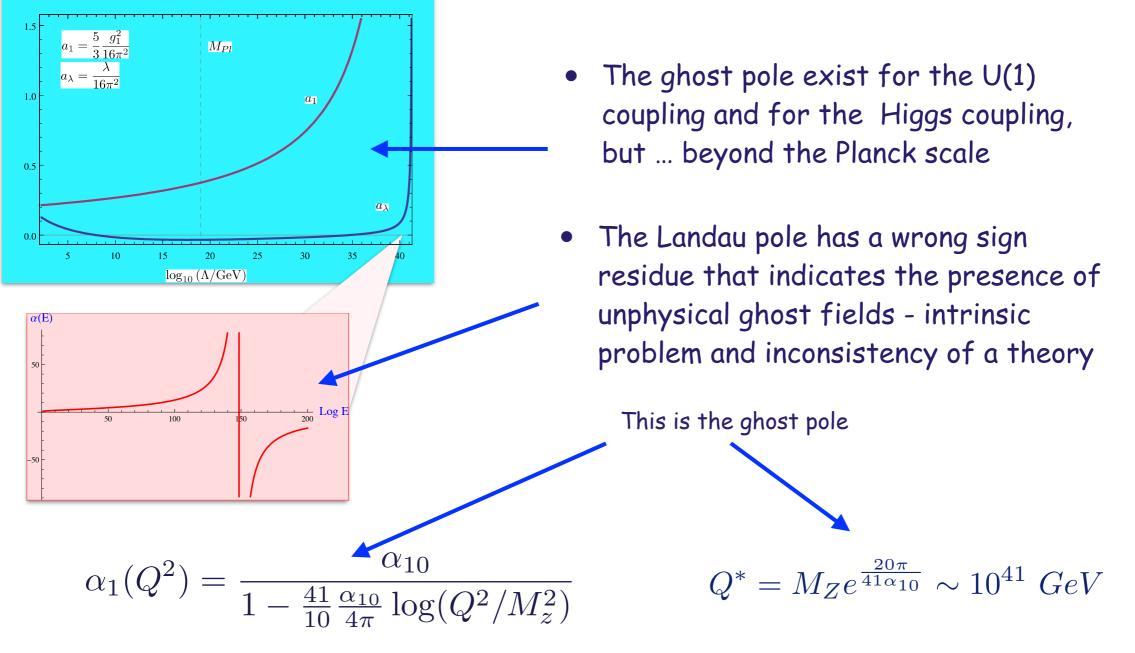


- The ghost pole exist for the U(1) coupling and for the Higgs coupling, but ... beyond the Planck scale
- The Landau pole has a wrong sign residue that indicates the presence of unphysical ghost fields - intrinsic problem and inconsistency of a theory

First State For the State of the Formation of the second state of the second state of the second state of the second state of the state of the second state of the state of th

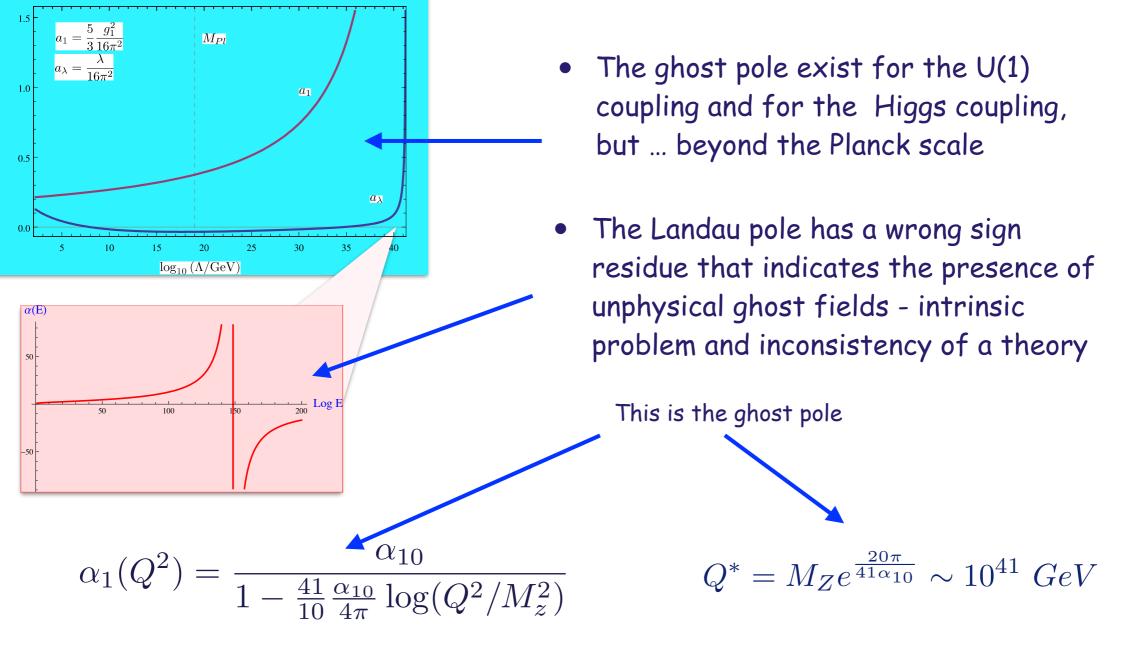


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• The situation may change in GUTs due to new heavy fields @ the GUT scale

First Strain Formation For



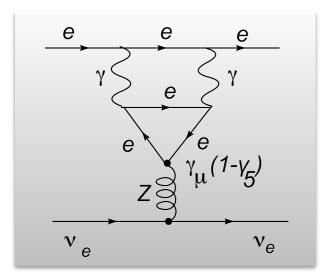
• The situation may change in GUTs due to new heavy fields @ the GUT scale

 requires modification of the ST at VERY high energies

Quantum anomalies may ruin the ST if not cancelled among quarks and leptons

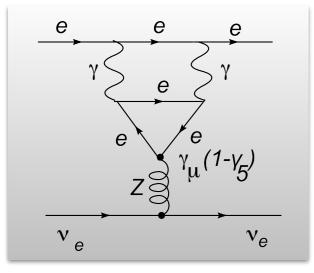
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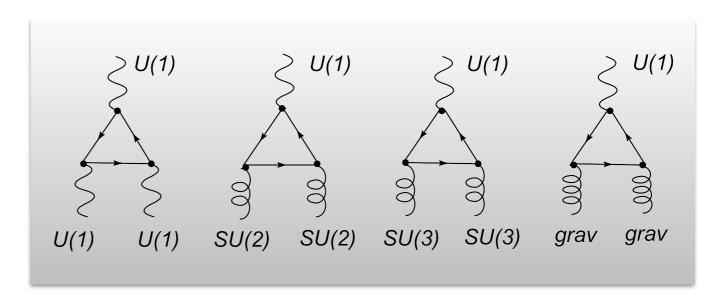
This is the anomalous diagram



Quantum anomalies may ruin the ST if not cancelled among quarks and leptons

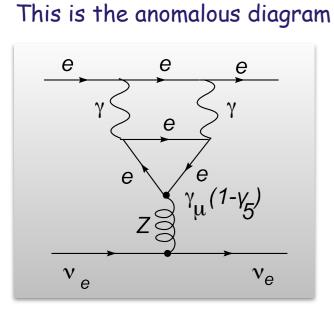


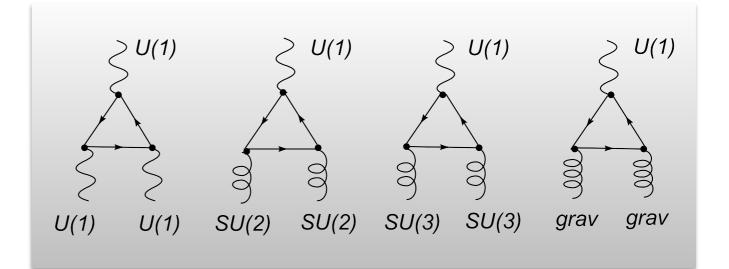




Anomalies in the SM

Quantum anomalies may ruin the ST if not cancelled among quarks and leptons

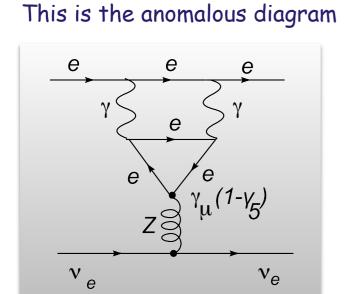


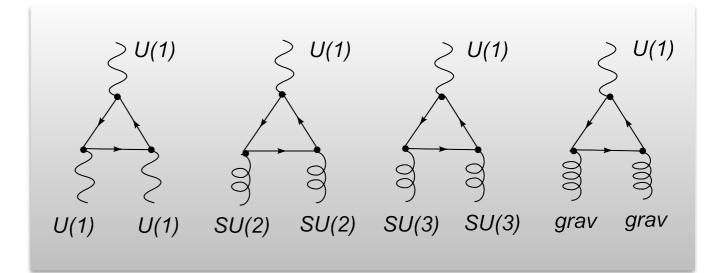


Anomalies in the SM

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Quantum anomalies may ruin the ST if not cancelled among quarks and leptons





Anomalies in the SM

$$TrY^{3} = = 3\left[\left(\frac{1}{3}\right)^{3} + \left(\frac{1}{3}\right)^{3} - \left(\frac{4}{3}\right)^{3} - \left(-\frac{2}{3}\right)^{3}\right] + (-1)^{3} + (-1)^{3} - (-2)^{3} = 0$$

$$\uparrow \uparrow \uparrow$$

$$colour \ u_{L} \ d_{L} \ u_{R} \ d_{R} \ \nu_{L} \ e_{L} \ e_{R}.$$

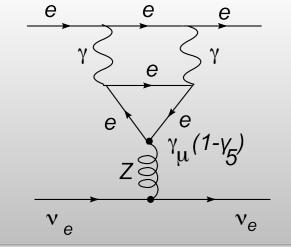
$$TrY_{L} = 3\left(\frac{1}{3} + \frac{1}{3}\right) - 1 - 1 = 0,$$

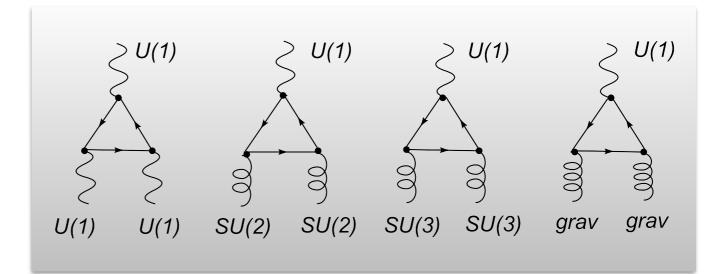
$$TrY_{q} = 3\left(\frac{1}{3} + \frac{1}{3} - \frac{4}{3} - \left(-\frac{2}{3}\right)\right) = 0,$$

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Anomalies in the SM

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- cancellation of anomalies requires quark-lepton symmetry
- this is a hint towards the Grand Unified Theories

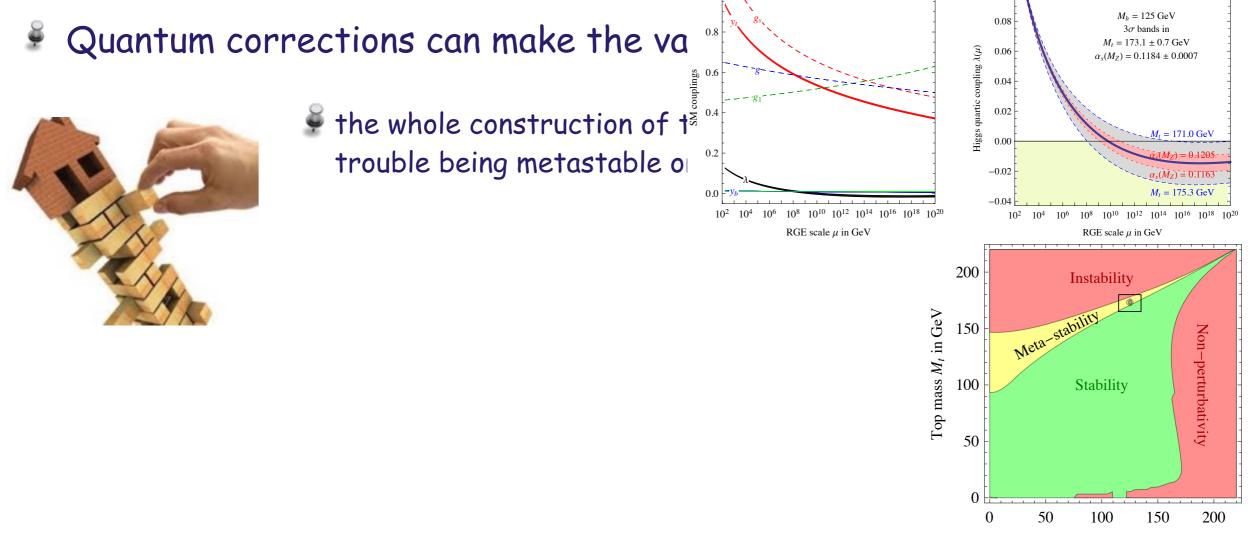
Quantum corrections can make the vacuum unstable



Quantum corrections can make the vacuum unstable

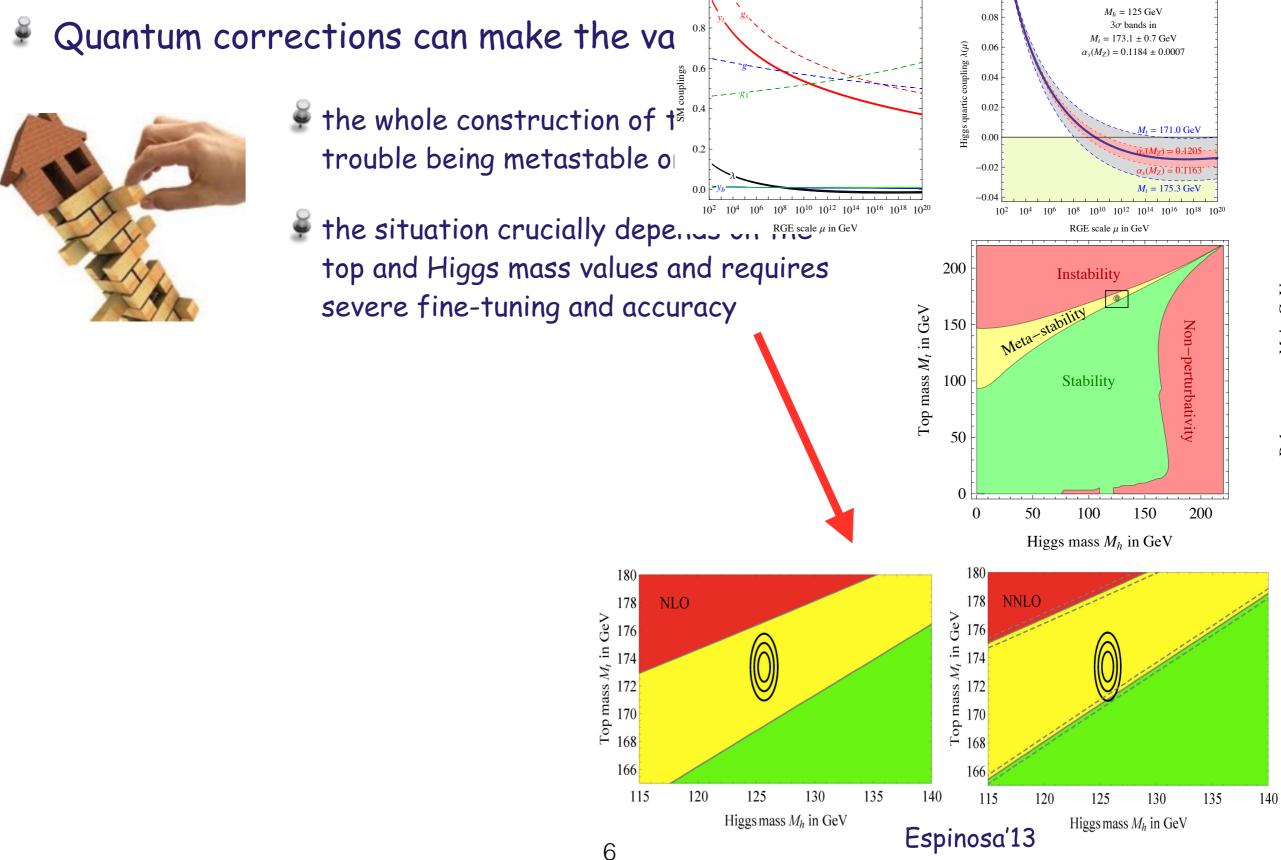


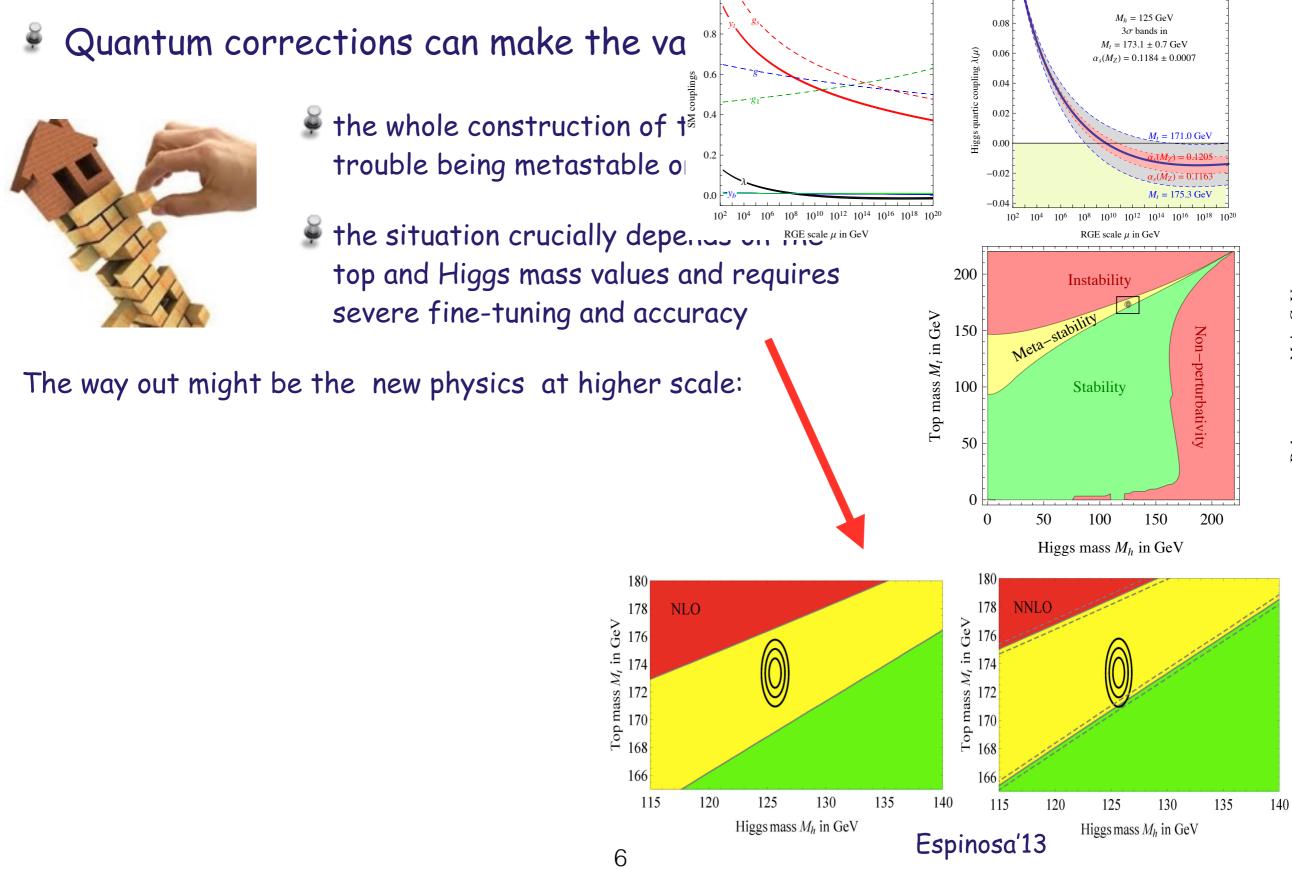
the whole construction of the SM may be in trouble being metastable or even unstable

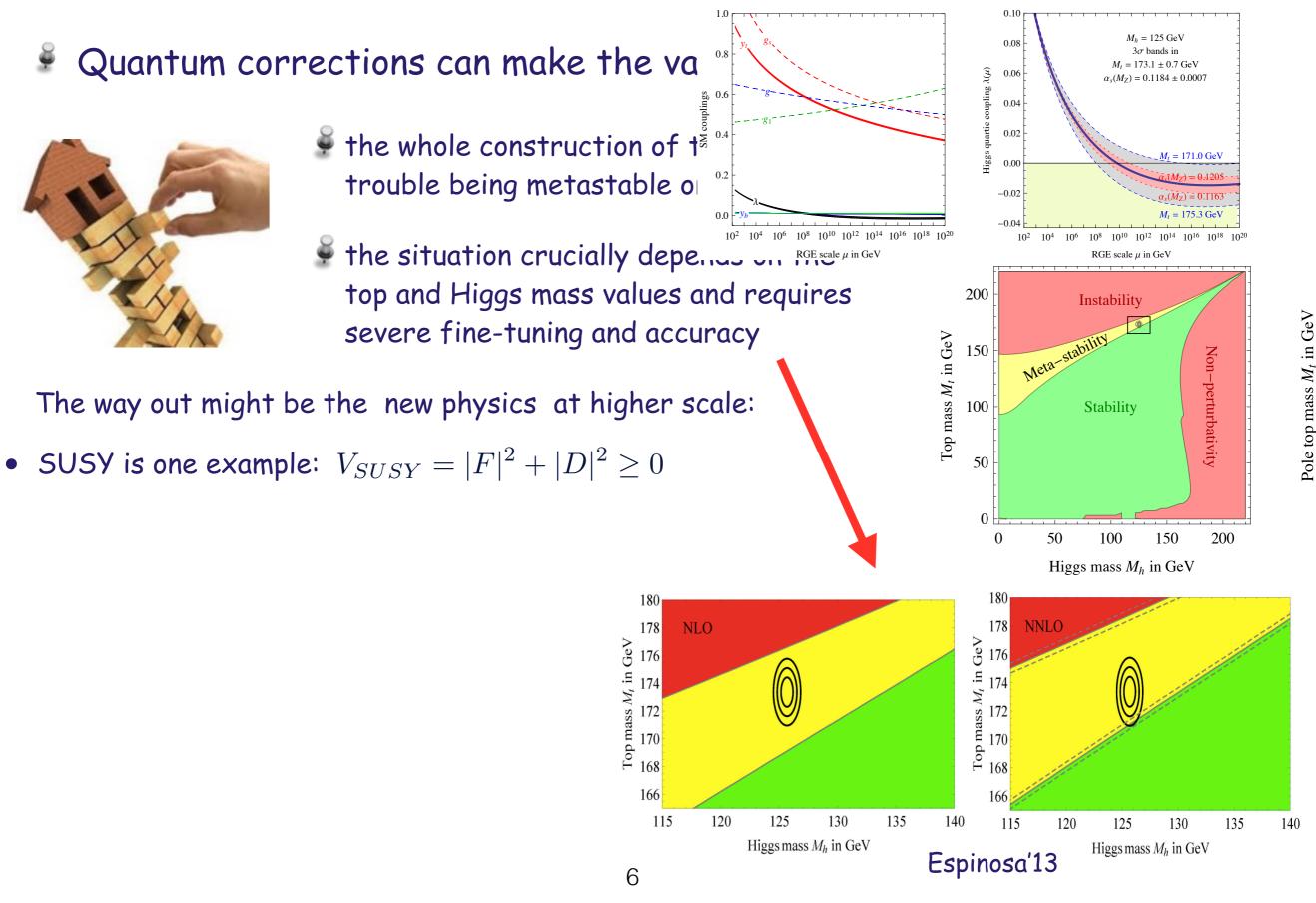


Higgs mass M_h in GeV

Pole top mass M_t in GeV





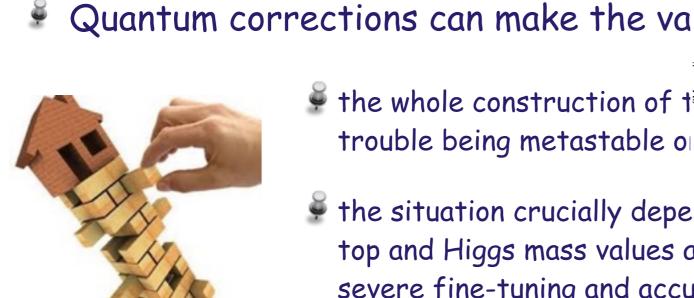


 $M_h = 125 \text{ GeV}$ 0.08 Quantum corrections can make the va 3σ bands in $M_t = 173.1 \pm 0.7 \text{ GeV}$ coupling $\lambda(\mu)$ 0.06 $\alpha_s(M_Z) = 0.1184 \pm 0.0007$ 0.04 quartic 0.02 the whole construction of t[§] Higgs $M_t = 171.0 \, \text{GeV}$ 0.00 trouble being metastable o -0.02 10^8 10^{10} 10^{12} 10^{14} 10^{16} 10^{18} 10^{20} $10^8 \ 10^{10} \ 10^{12} \ 10^{14} \ 10^{16} \ 10^{18}$ 10^{2} 10^{6} the situation crucially dependent. RGE scale μ in GeV RGE scale μ in GeV top and Higgs mass values and requires 200 Instability mass M_t in GeV severe fine-tuning and accuracy Top mass M_t in GeV Meta-stability 150 Non-perturbativity The way out might be the new physics at higher scale: 100 Stability Pole top SUSY is one example: $V_{SUSY} = |F|^2 + |D|^2 \ge 0$ 50 Extended Higgs sector is another 150 50 100 200 0 example: Higgs mass M_h in GeV Several Higgs fields with several Higgs-like 180 180 couplings push the smallest coupling up NLO 178 178 **NNLO** 290 176 Fop mass M_t in GeV (might have also several minima) 176 U 174 U 174 U 172 U 170 U 168 174 172 170 168 166 166 120 125 130 135 115 140 130 120 125 135 115 140 Higgs mass M_h in GeV

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Higgs mass M_h in GeV

Espinosa'13



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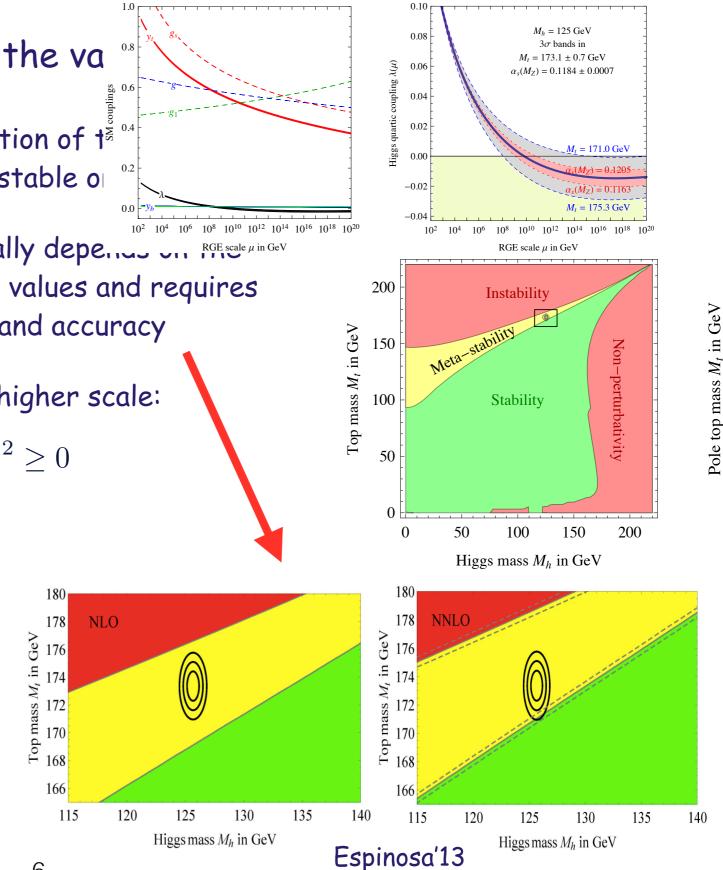
The way out might be the new physics at higher scale:

- SUSY is one example: $V_{SUSY} = |F|^2 + |D|^2 \ge 0$
- Extended Higgs sector is another example:

Several Higgs fields with several Higgs-like couplings push the smallest coupling up (might have also several minima)

GUT's provide the third example:

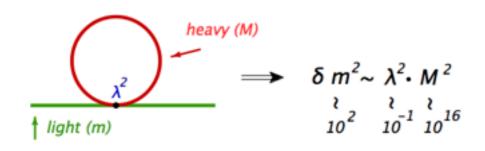
In a unified theory the Higgs coupling might be attracted by the gauge coupling and stabilize the potential



New physics at high scale may destroy the EW scale of the ST

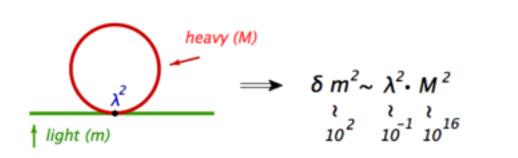
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Quantum corrections to the Higgs potential due to New physics



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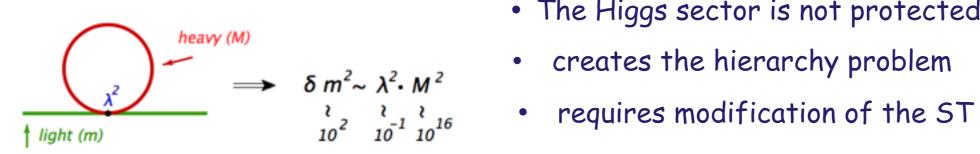
- The Higgs sector is not protected by any symmetry
- creates the hierarchy problem

$$\frac{m_H}{m_{GUT}} \sim 10^{-14}$$

requires modification of the ST

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Quantum corrections to the Higgs potential due to New physics



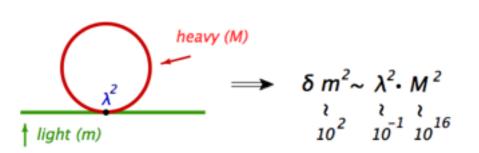
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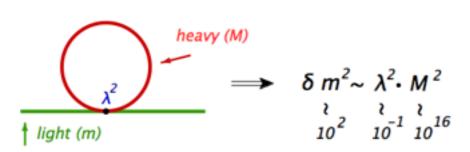
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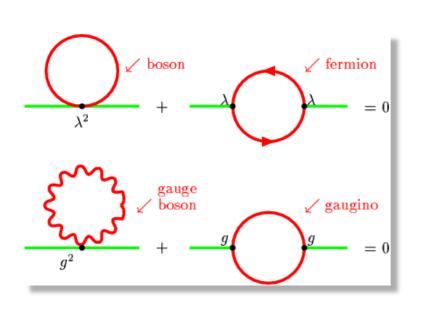
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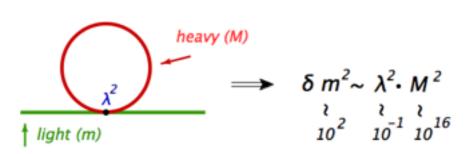
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• cancellation with superpartners up to $\Delta m^2 \sim 1 \ TeV$

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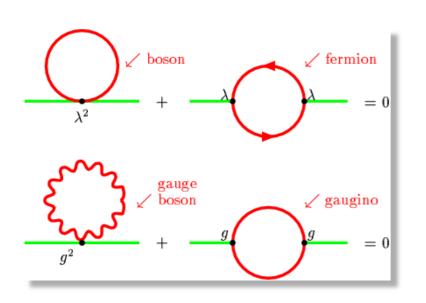
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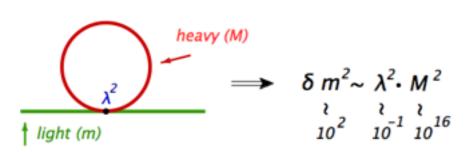


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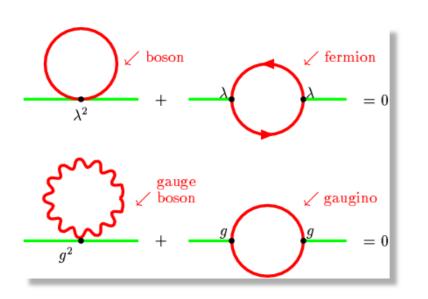


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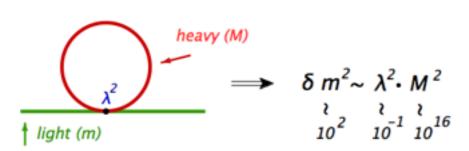
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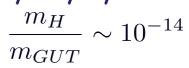
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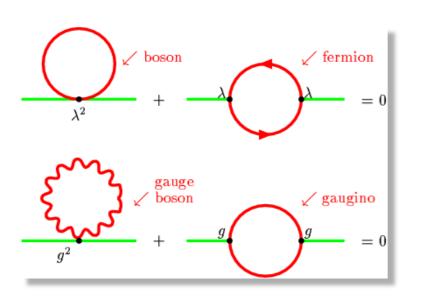


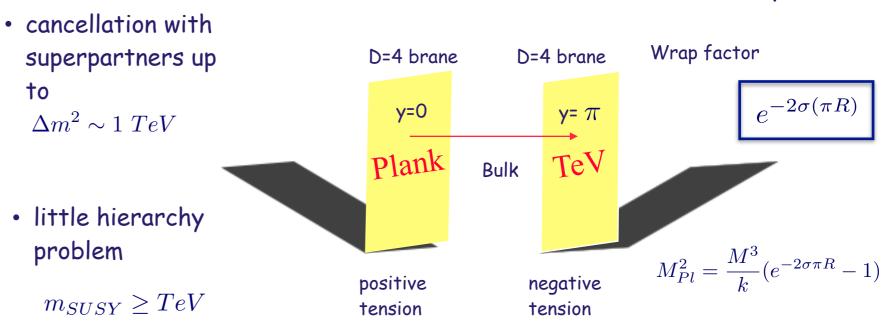
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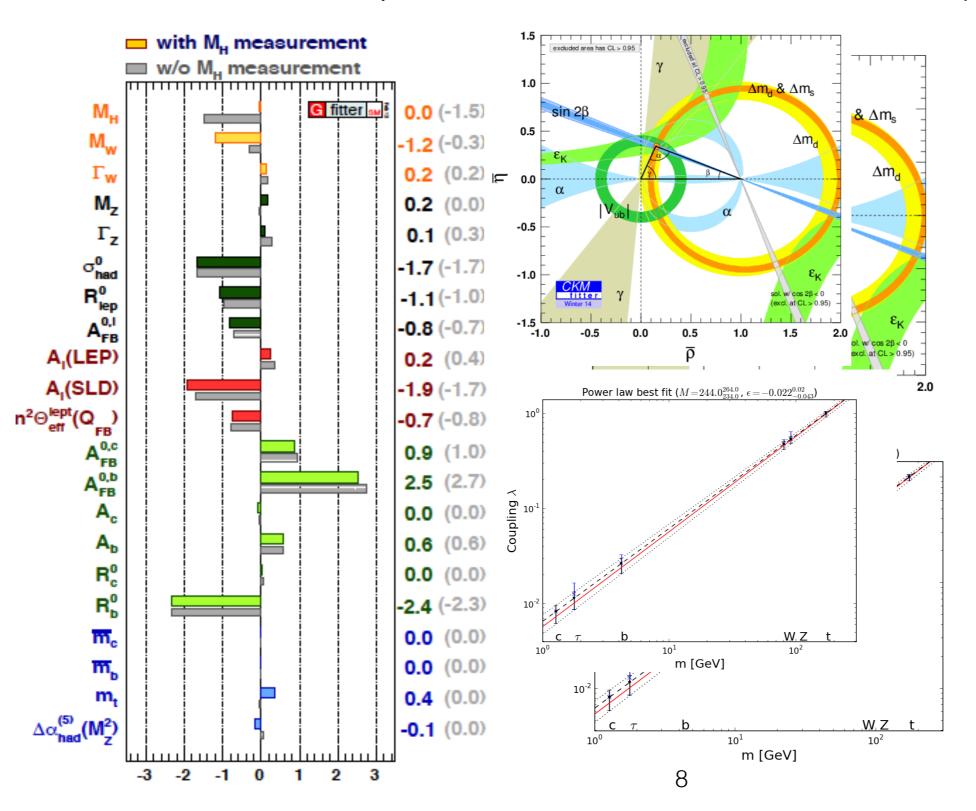




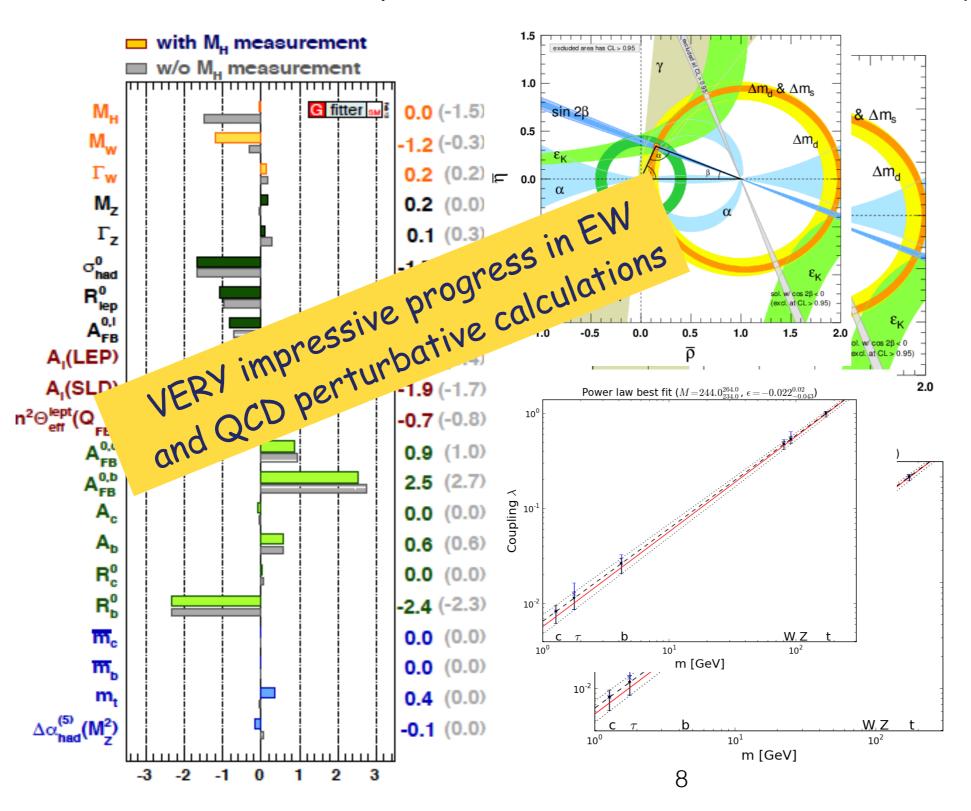


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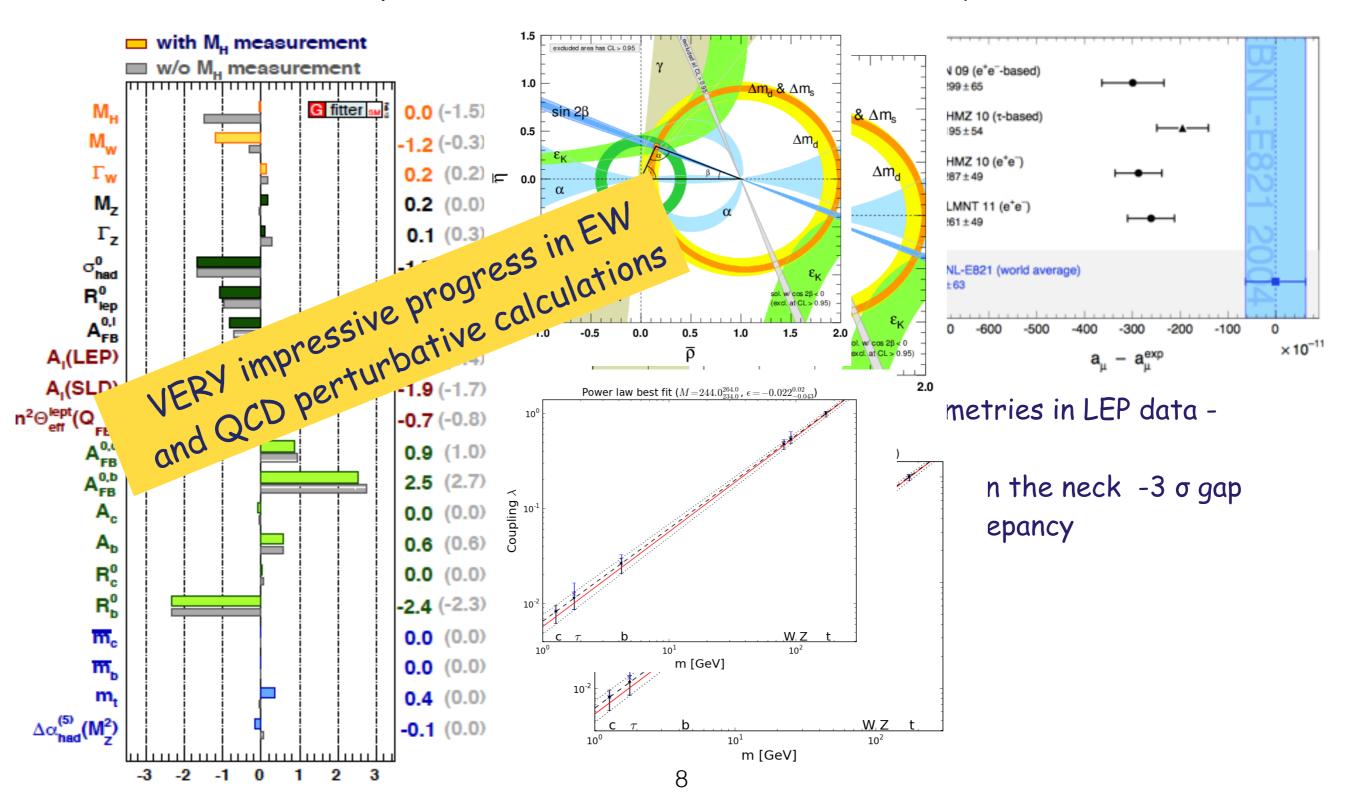
EW observables pool



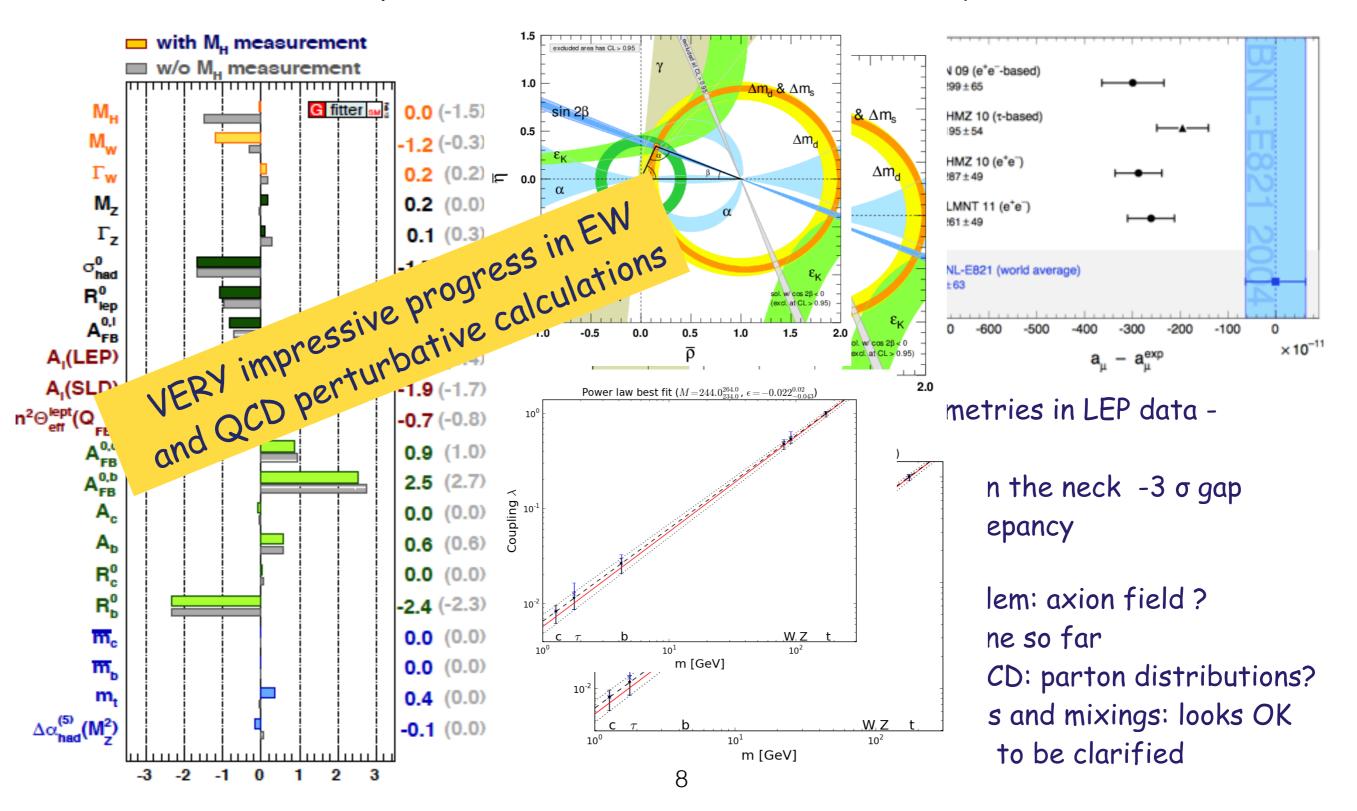
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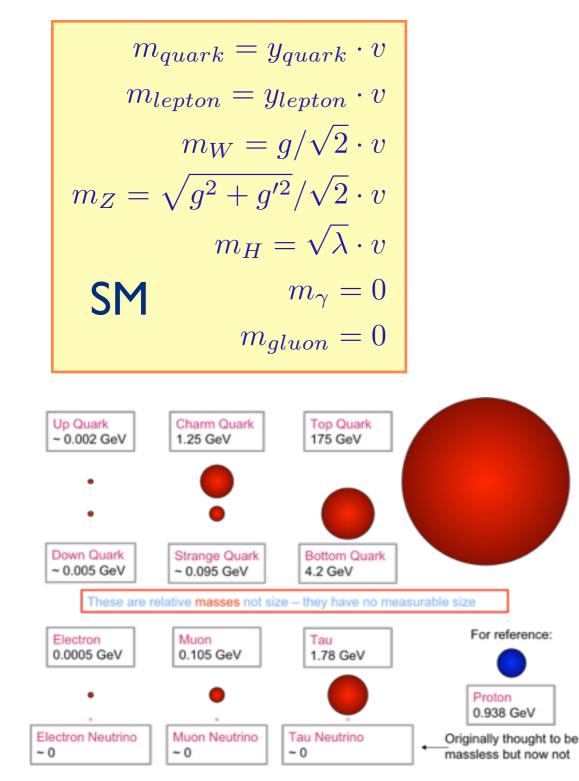
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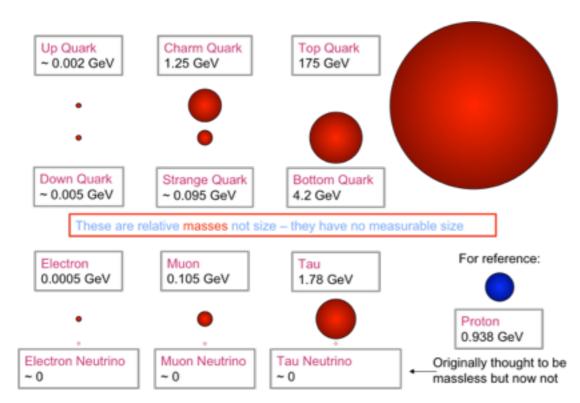


$$egin{aligned} m_{quark} &= y_{quark} \cdot v \ m_{lepton} &= y_{lepton} \cdot v \ m_W &= g/\sqrt{2} \cdot v \ m_Z &= \sqrt{g^2 + g'^2}/\sqrt{2} \cdot v \ m_H &= \sqrt{\lambda} \cdot v \ \mathbf{SM} & m_\gamma &= 0 \ m_{gluon} &= 0 \end{aligned}$$



• Mass spectrum?

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Down Quark Strange Quark Bottom Quark ~ 0.005 GeV ~ 0.095 GeV 4.2 GeV These are relative masses not size - they have no measurable size For reference: Electron Muon Tau 0.105 GeV 0.0005 GeV 1.78 GeV ٠ Proton 0.938 GeV Tau Neutrino Electron Neutrino Muon Neutrino Originally thought to be ~ 0 ~ 0 ~ 0 massless but now not

- Mixing Matrices?
- Quark-Lepton Symmetry

$$m_{quark} = y_{quark} \cdot v$$

$$m_{lepton} = y_{lepton} \cdot v$$

$$m_W = g/\sqrt{2} \cdot v$$

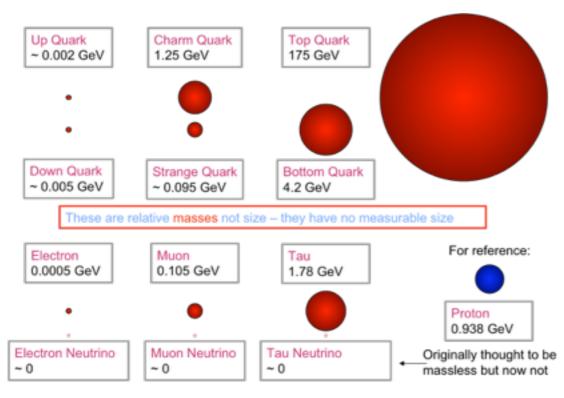
$$m_Z = \sqrt{g^2 + g'^2}/\sqrt{2} \cdot v$$

$$m_H = \sqrt{\lambda} \cdot v$$

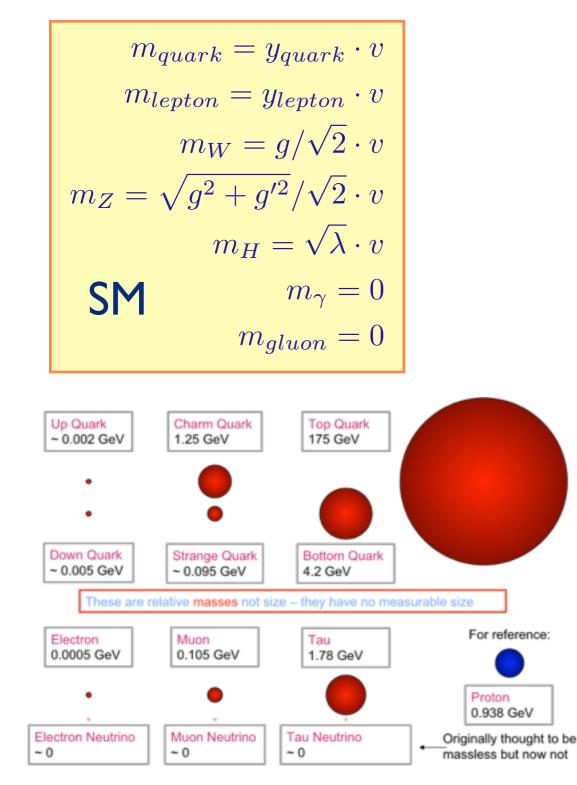
$$M_H = \sqrt{\lambda} \cdot v$$

$$m_{\gamma} = 0$$

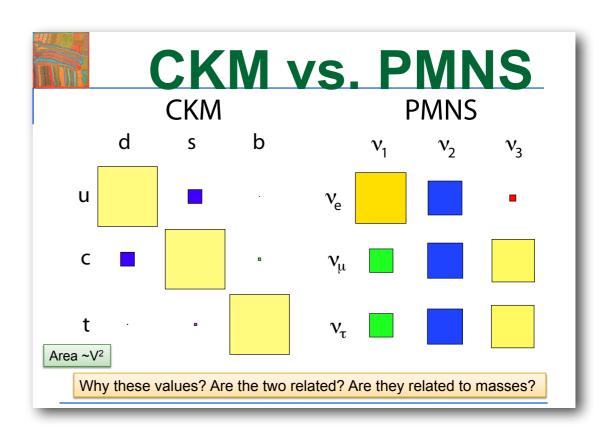
$$m_{gluon} = 0$$



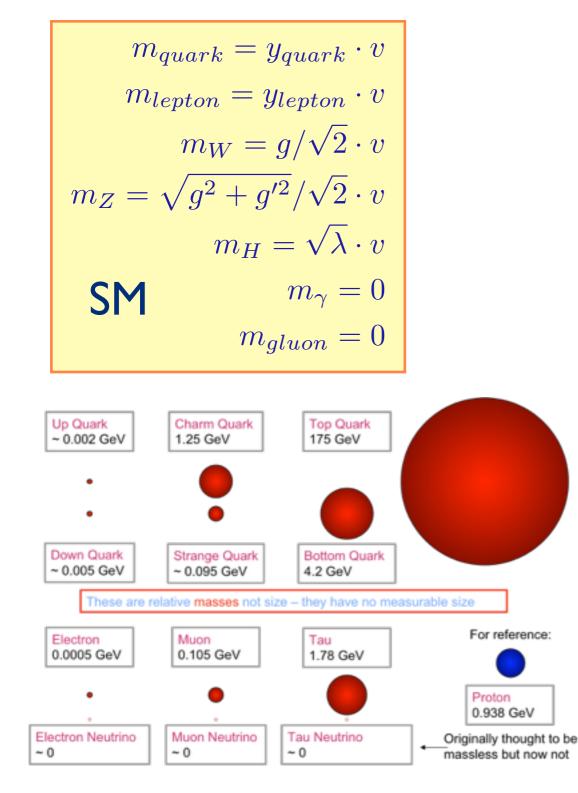
- Mixing Matrices?
- Quark-Lepton Symmetry
- Strong difference in parameters



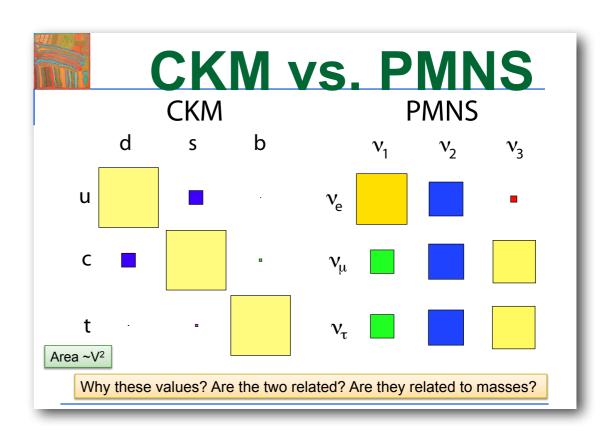
- Mixing Matrices?
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- Strong difference in parameters



• Mass spectrum?

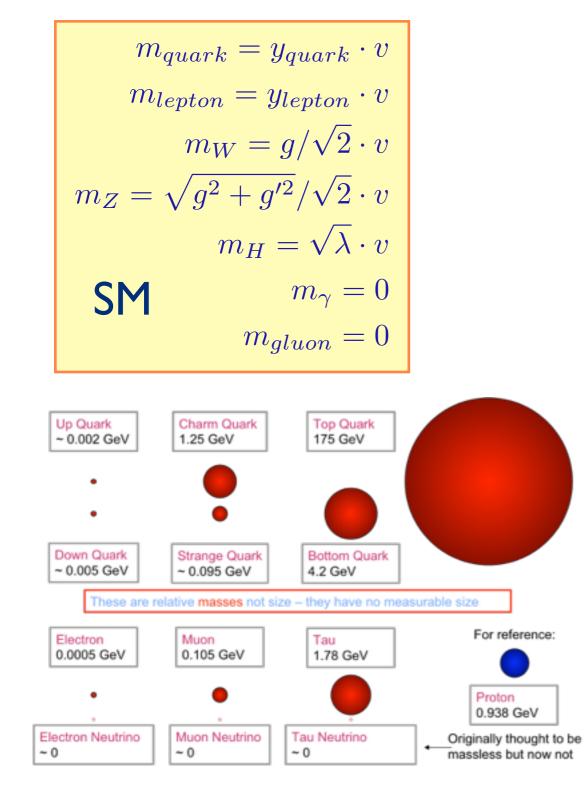


- Mixing Matrices?
- Quark-Lepton Symmetry
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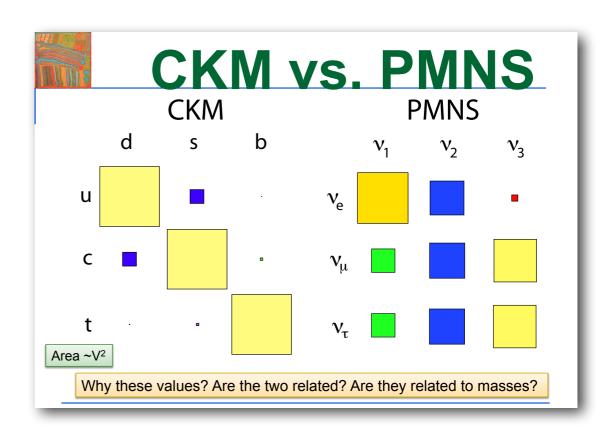


• What re the CKM and PMNS phases?

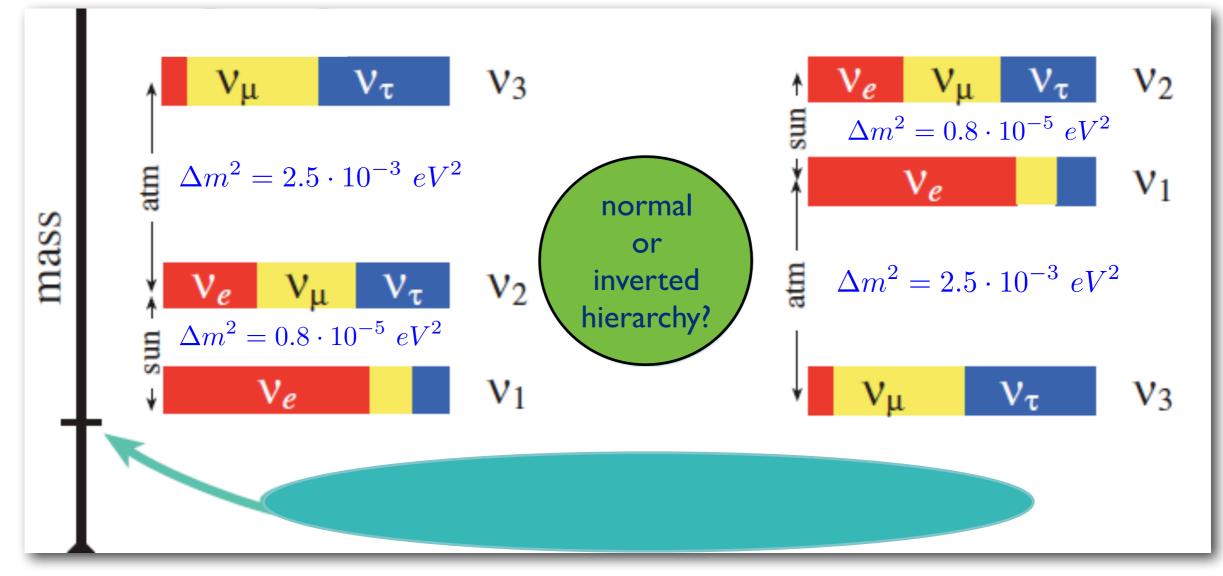
• Mass spectrum?

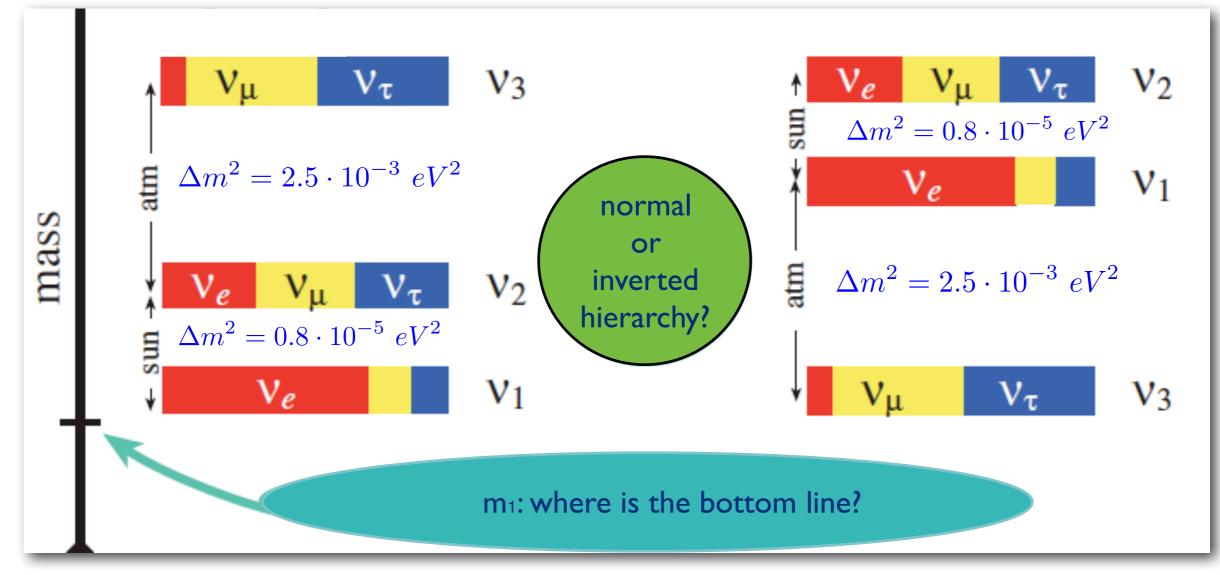


- Mixing Matrices?
- Quark-Lepton Symmetry
- Strong difference in parameters

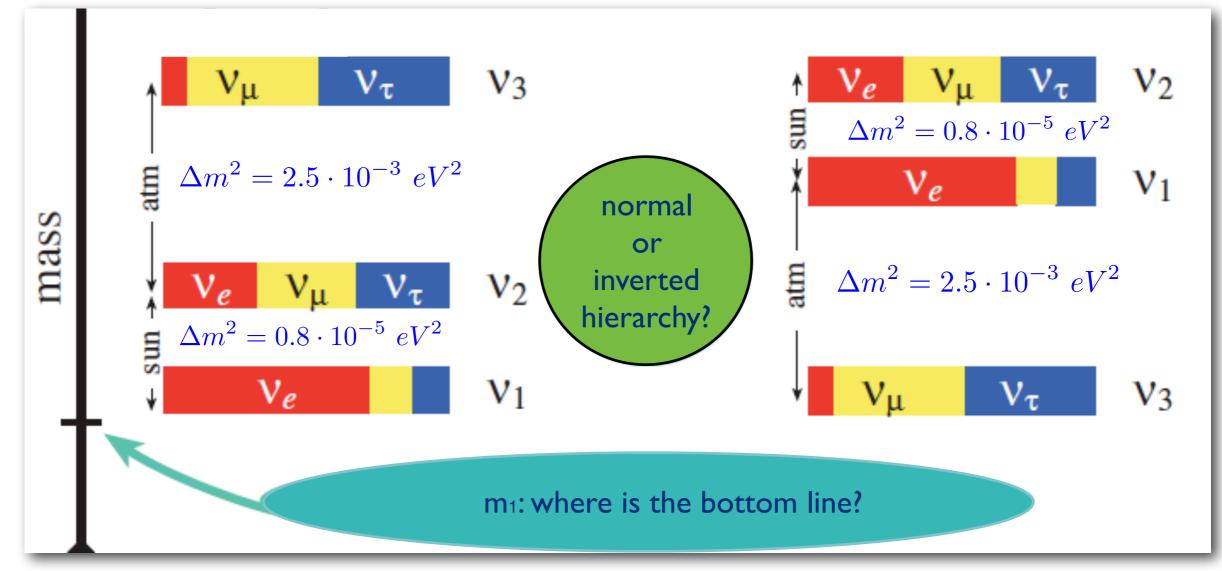


What re the CKM and PMNS phases?
Where lies the source of CP violation: in qurk or lepton sector?





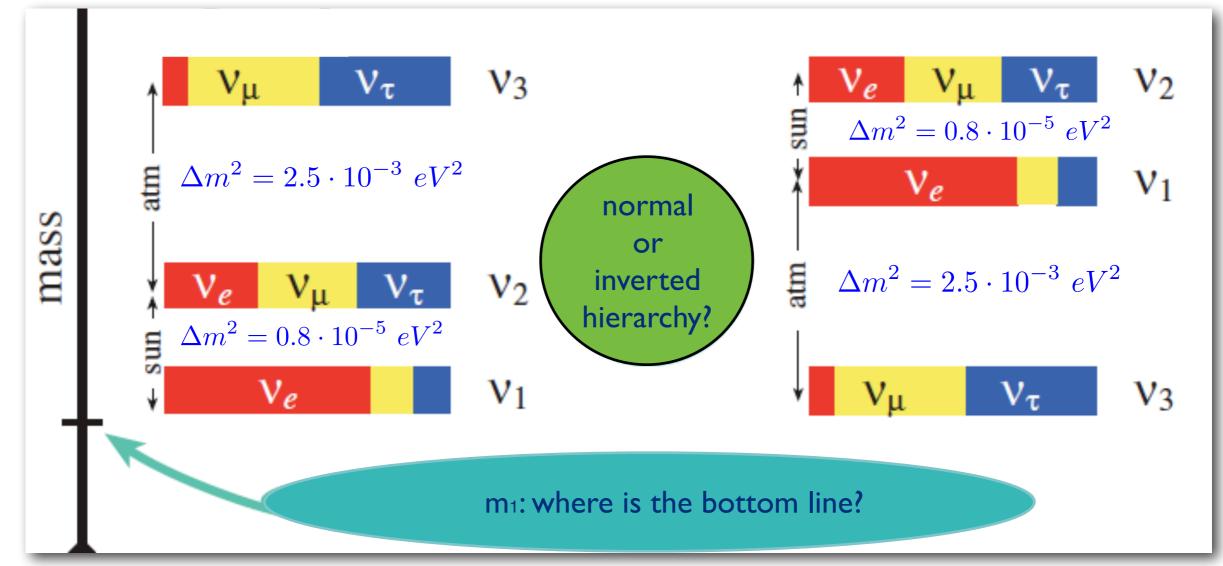
Neutrino masses



$$\sum m_{\nu} < 0.23 \ eV$$

cosmology: the CMB spectrum Planck

Neutrino masses

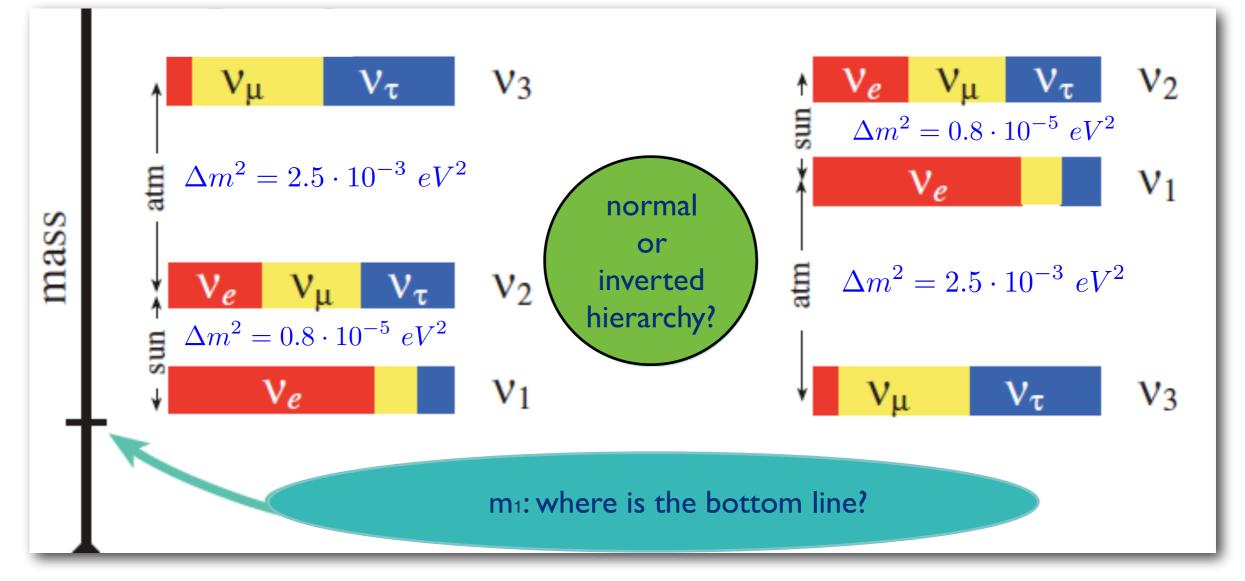


$$\sum m_{\nu} < 0.23 \ eV$$

cosmology: the CMB spectrum Planck

$$m_{\nu_e} < 2 \ eV$$

β-decay Troitsk-Mainz

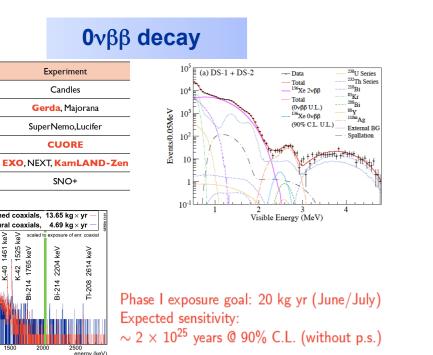


$$\nu_D = \begin{pmatrix} \nu_L \\ \nu_R \end{pmatrix} \quad \nu_{M_1} = \begin{pmatrix} \xi_1 \\ \xi_1^* \end{pmatrix}, \quad \nu_{M_2} = \begin{pmatrix} \xi_2 \\ \xi_2^* \end{pmatrix}$$

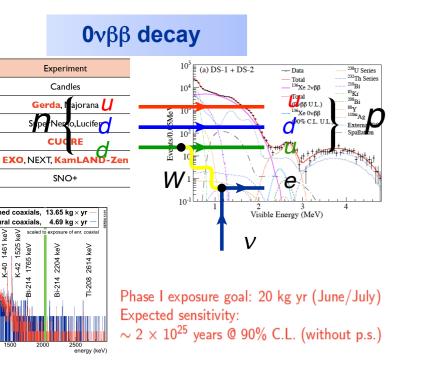
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$$\boxed{\begin{array}{c} \nu_{D} \neq \nu_{D}^{*} \\ m_{\nu_{L}} = m_{\nu_{R}} \end{array}}$$

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$$? \qquad \boxed{\nu_{M} = \nu_{M}^{*}}$$
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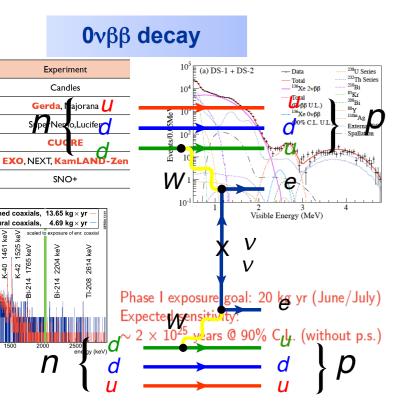
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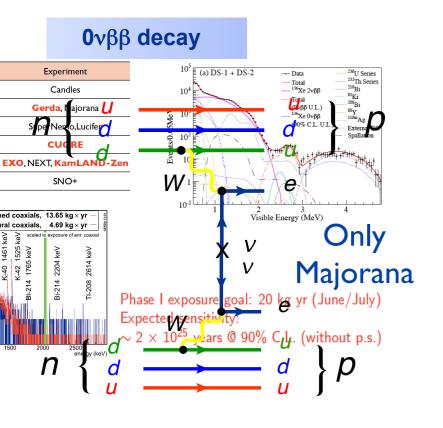
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Dirac or Majorana?

Experiment Candles Gerda, lajorana U

o,Lucifer

Si per Nei

ned coaxials, 13.65 kg × yr 4.69 kg × yr

n

EXO, NEXT, KamLAND-Z SNO+

$$\nu_{D} = \begin{pmatrix} \nu_{L} \\ \nu_{R} \end{pmatrix} \quad \nu_{M_{1}} = \begin{pmatrix} \xi_{1} \\ \xi_{1}^{*} \end{pmatrix}, \quad \nu_{M_{2}} = \begin{pmatrix} \xi_{2} \\ \xi_{2}^{*} \end{pmatrix}$$

$$\nu_{D} \neq \nu_{D}^{*}$$

$$m_{\nu_{L}} = m_{\nu_{R}}$$

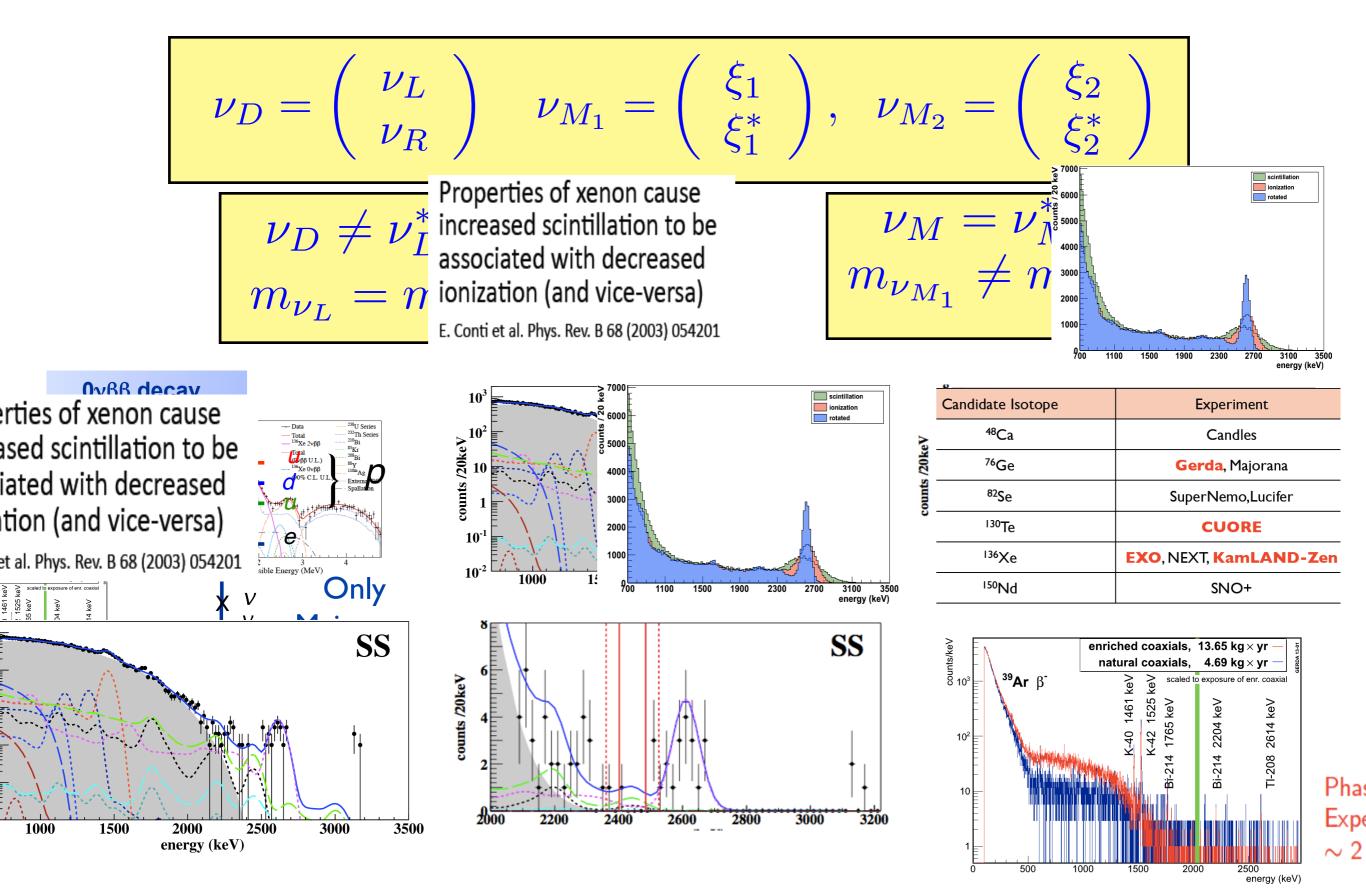
$$(P) \qquad P = \nu_{M}^{*}$$

$$m_{\nu_{M_{1}}} \neq m_{\nu_{M_{2}}}$$

Energy, keV

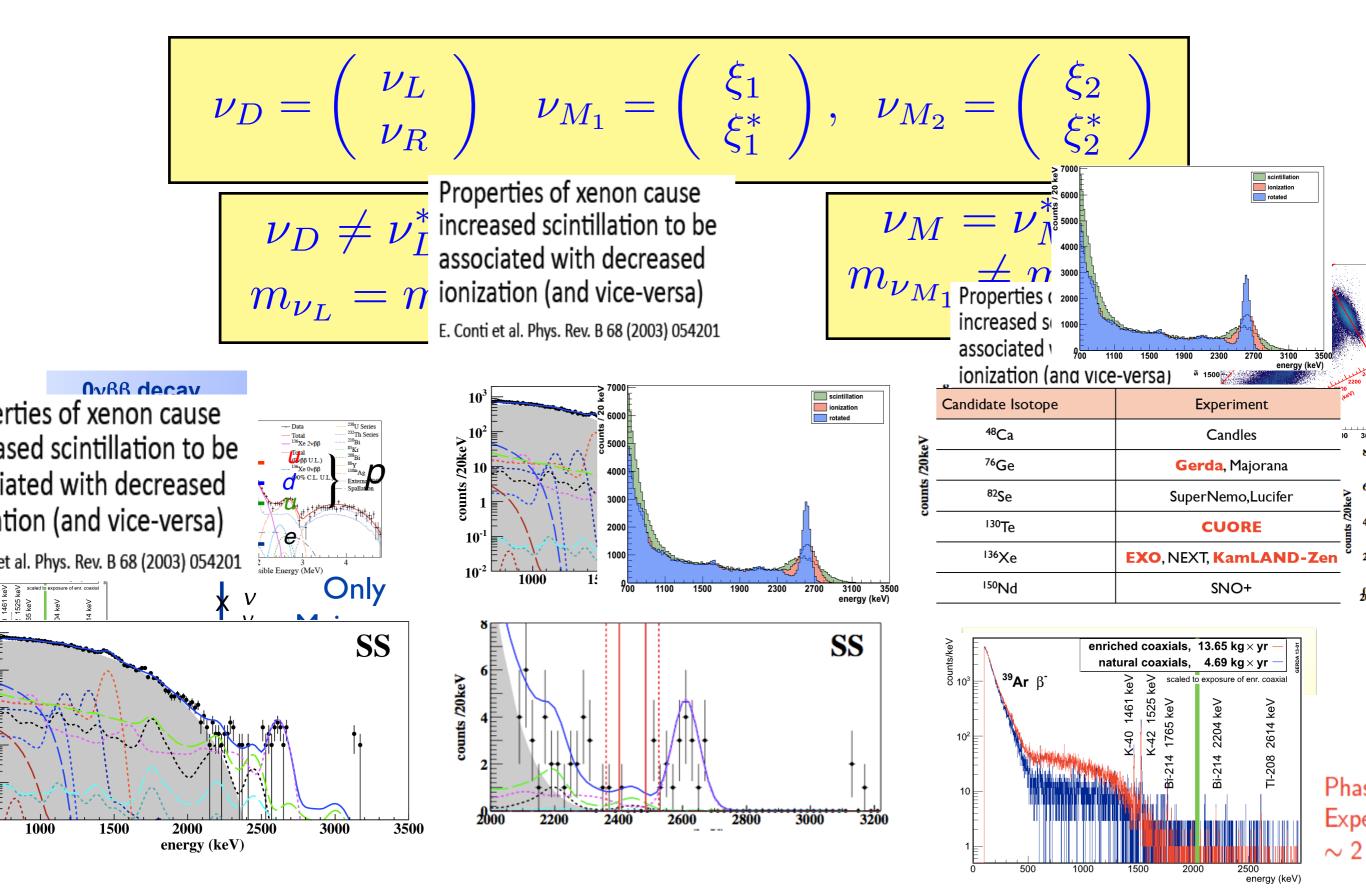
Dirac or Majorana?

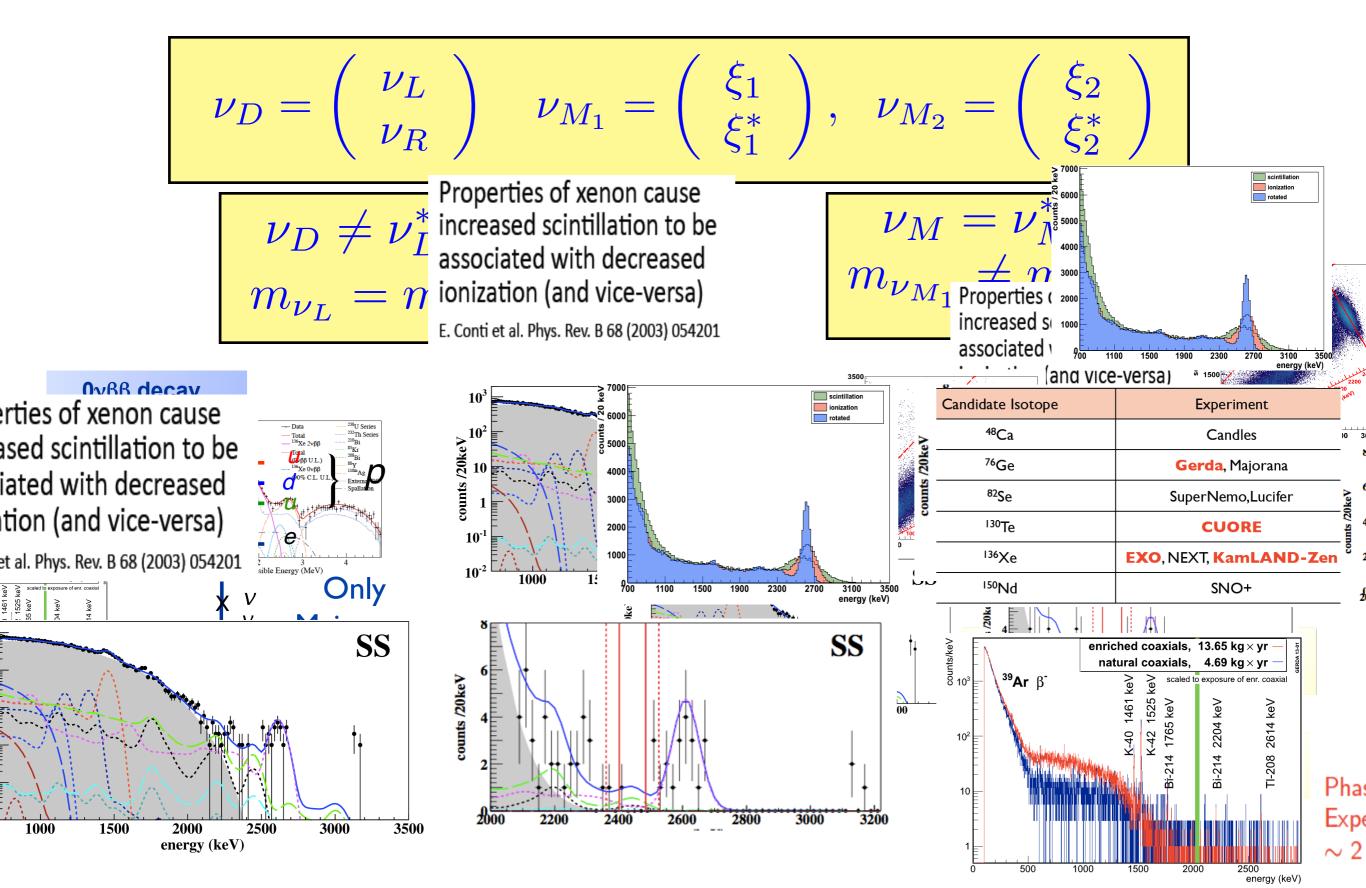
1000

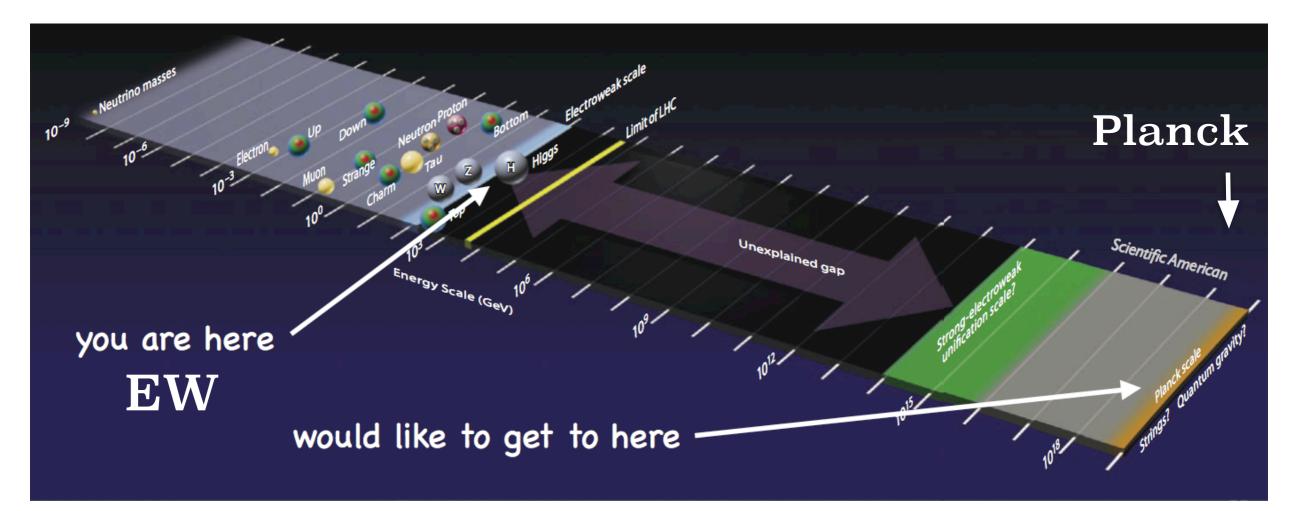


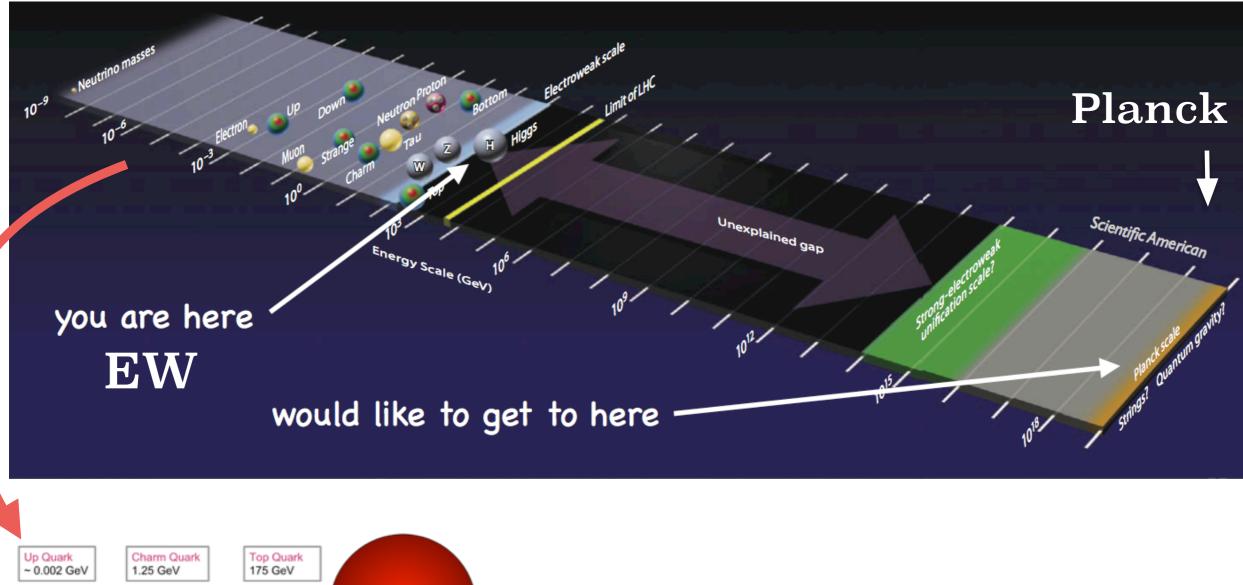
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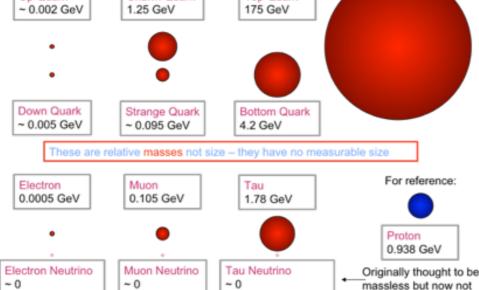
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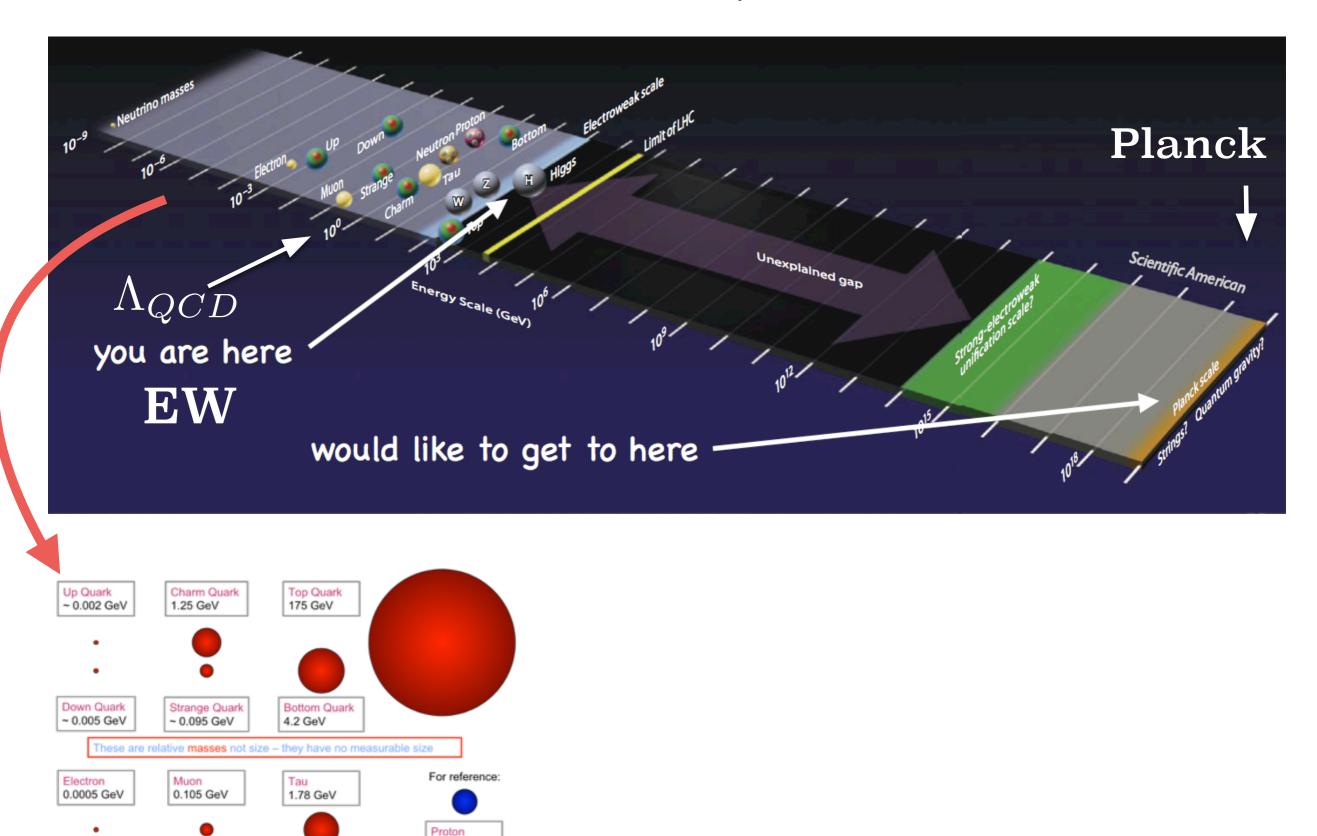








12



0.938 GeV

Originally thought to be

massless but now not

Muon Neutrino

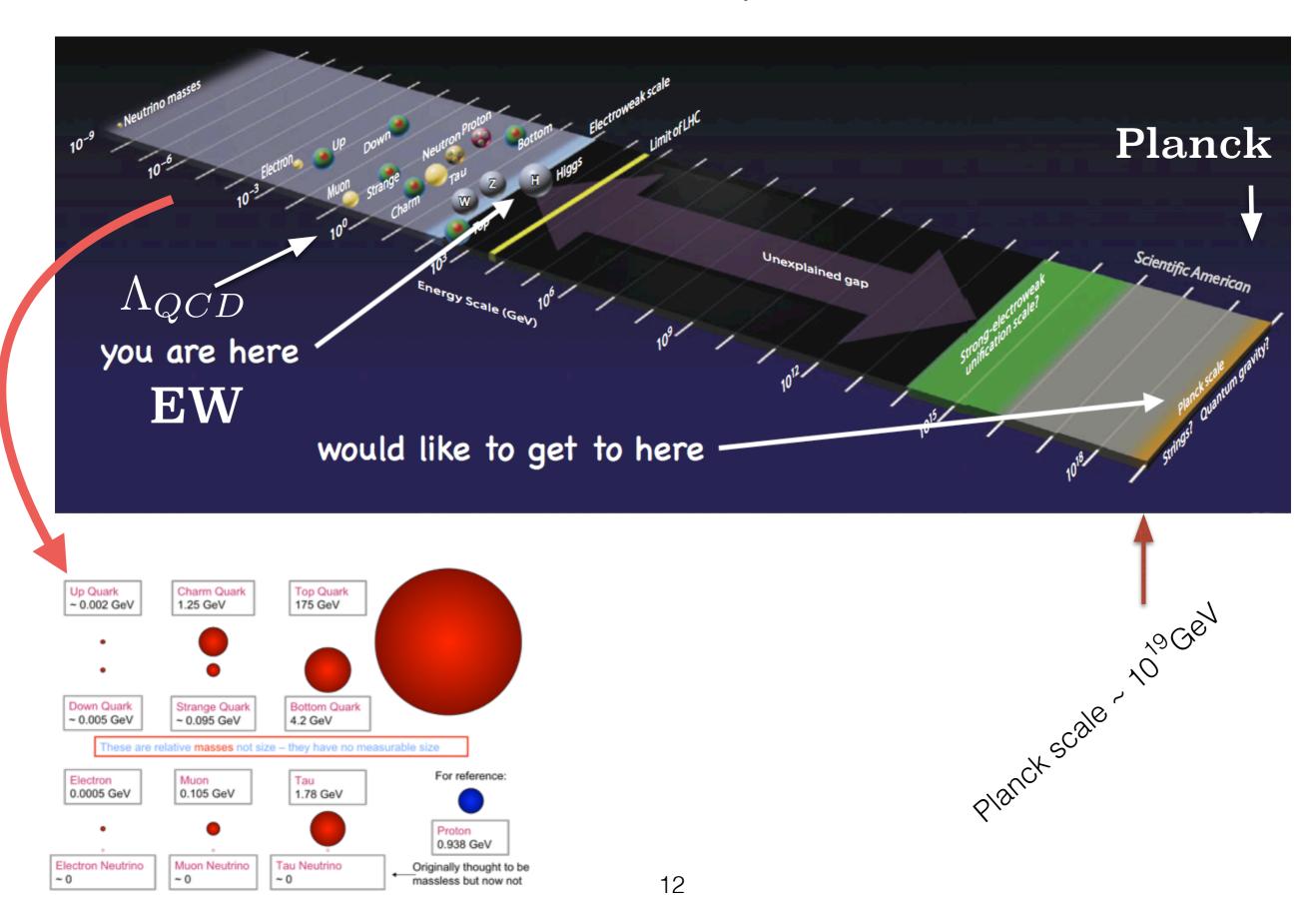
~ 0

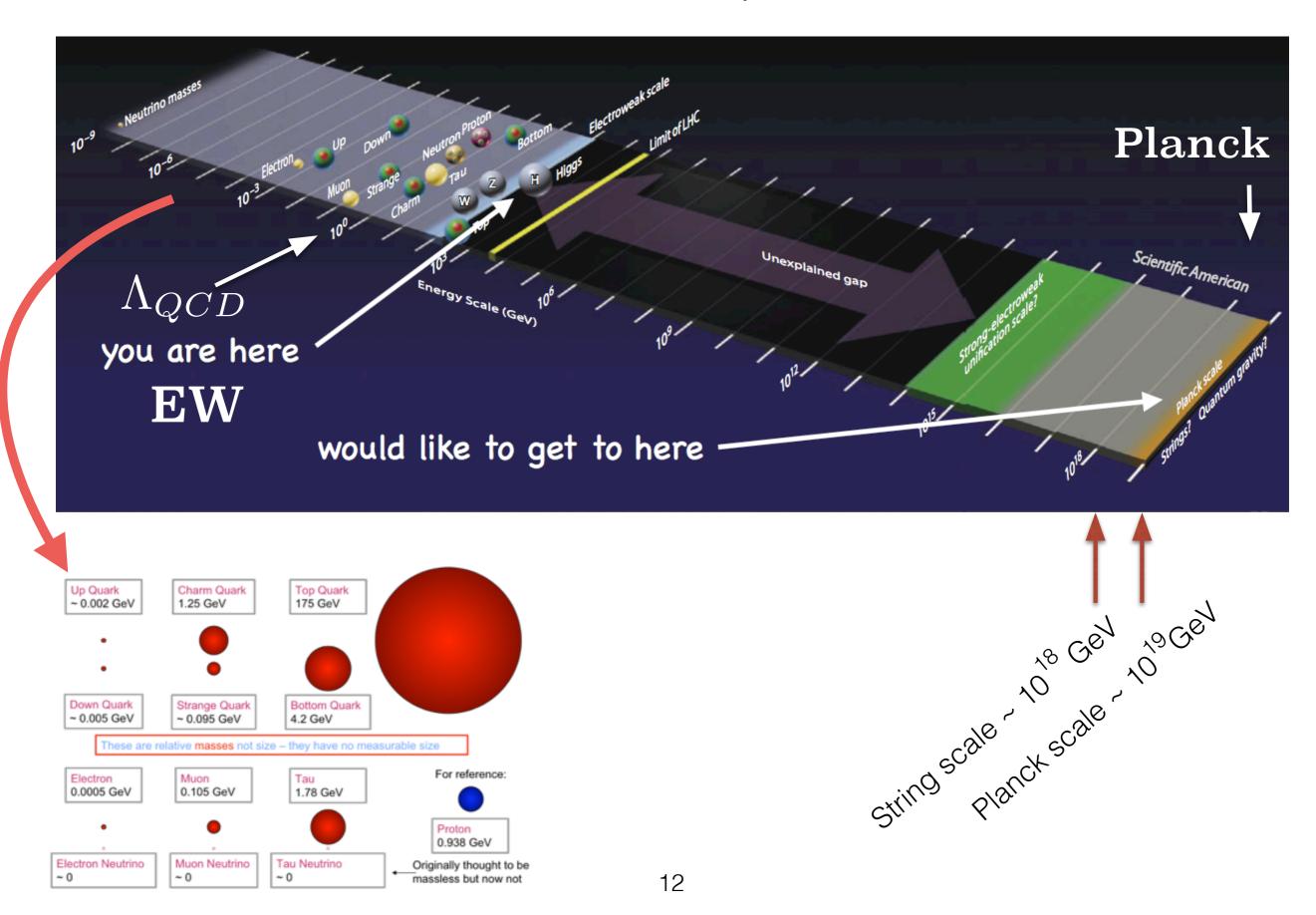
Tau Neutrino

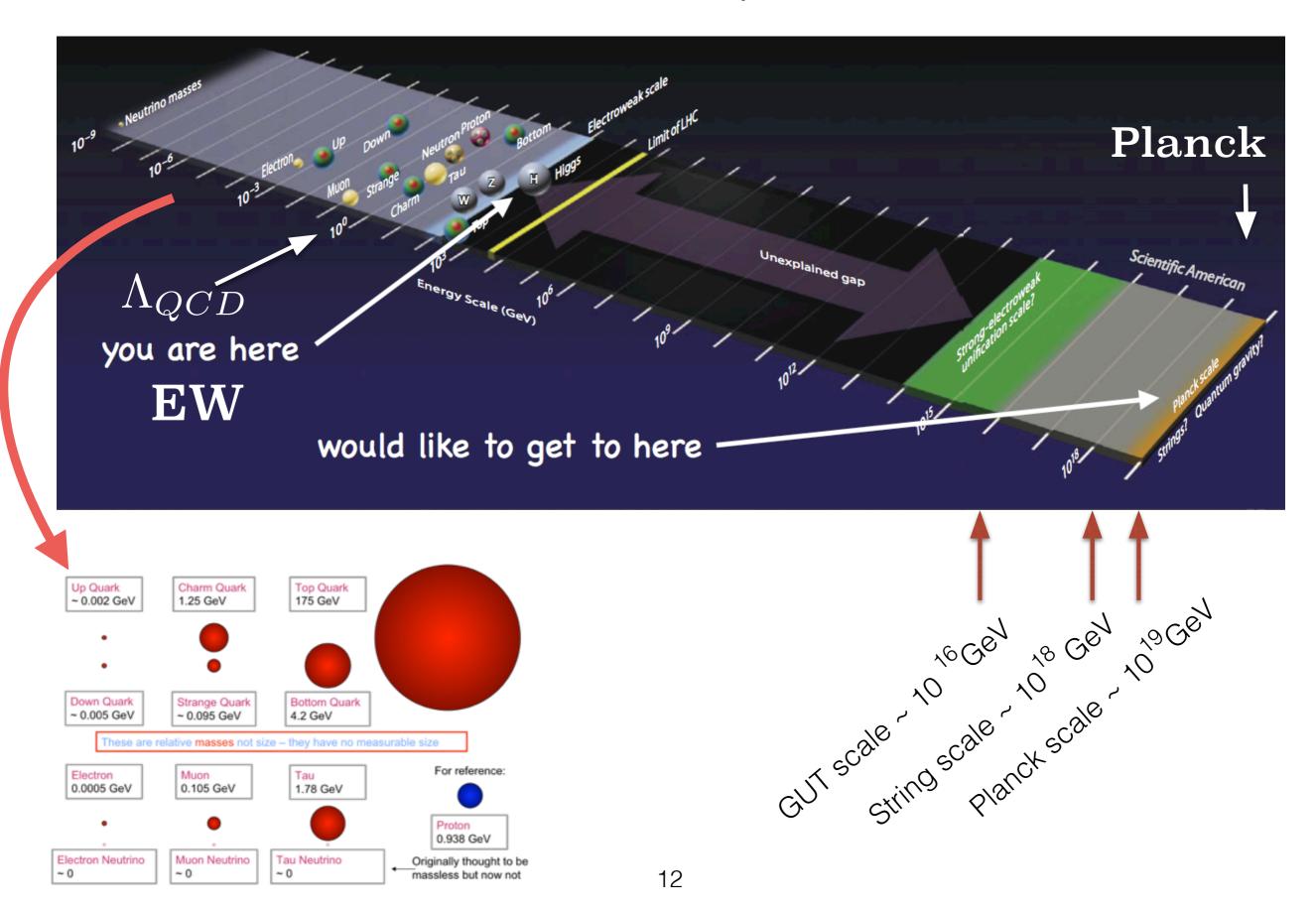
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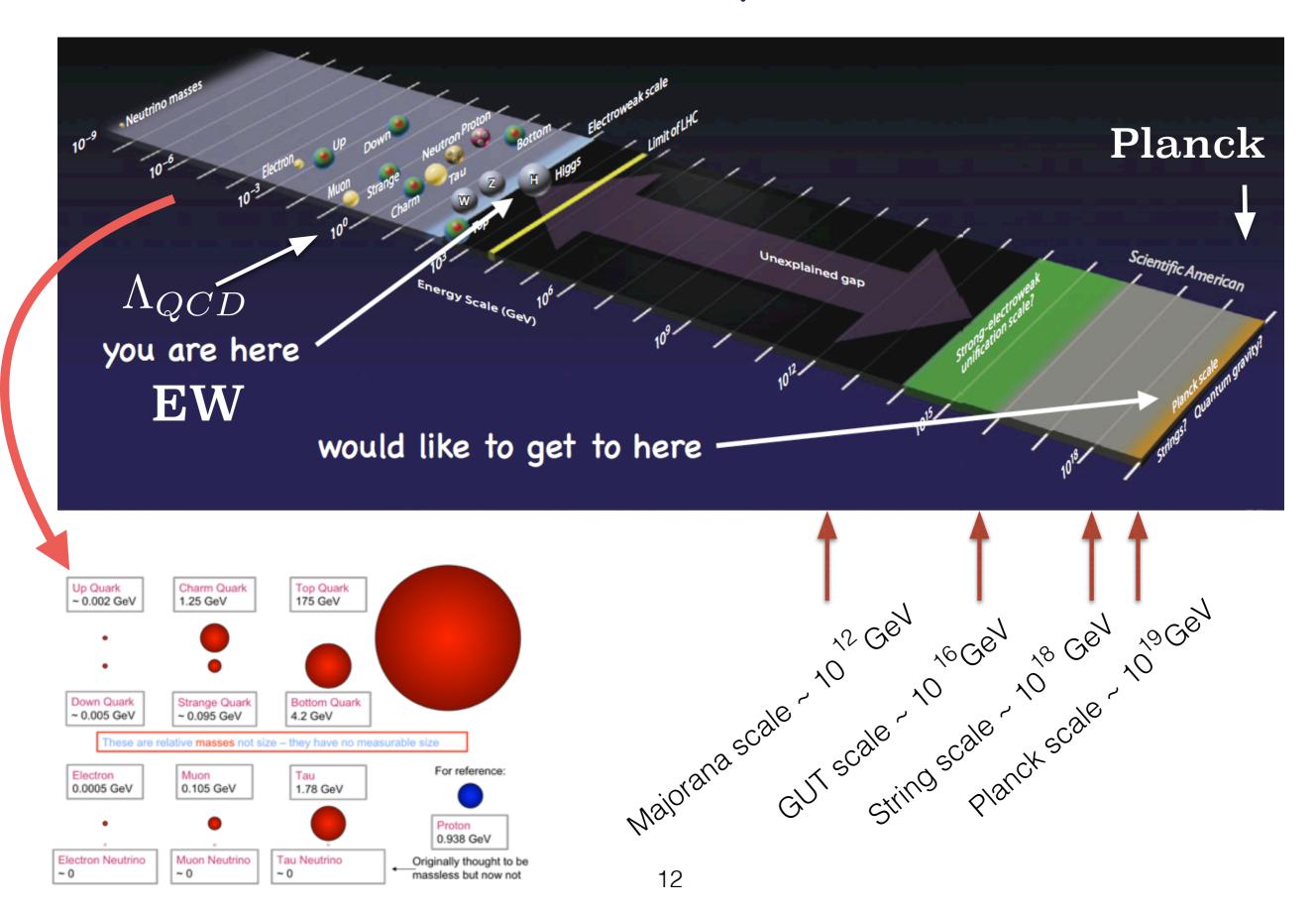
Electron Neutrino

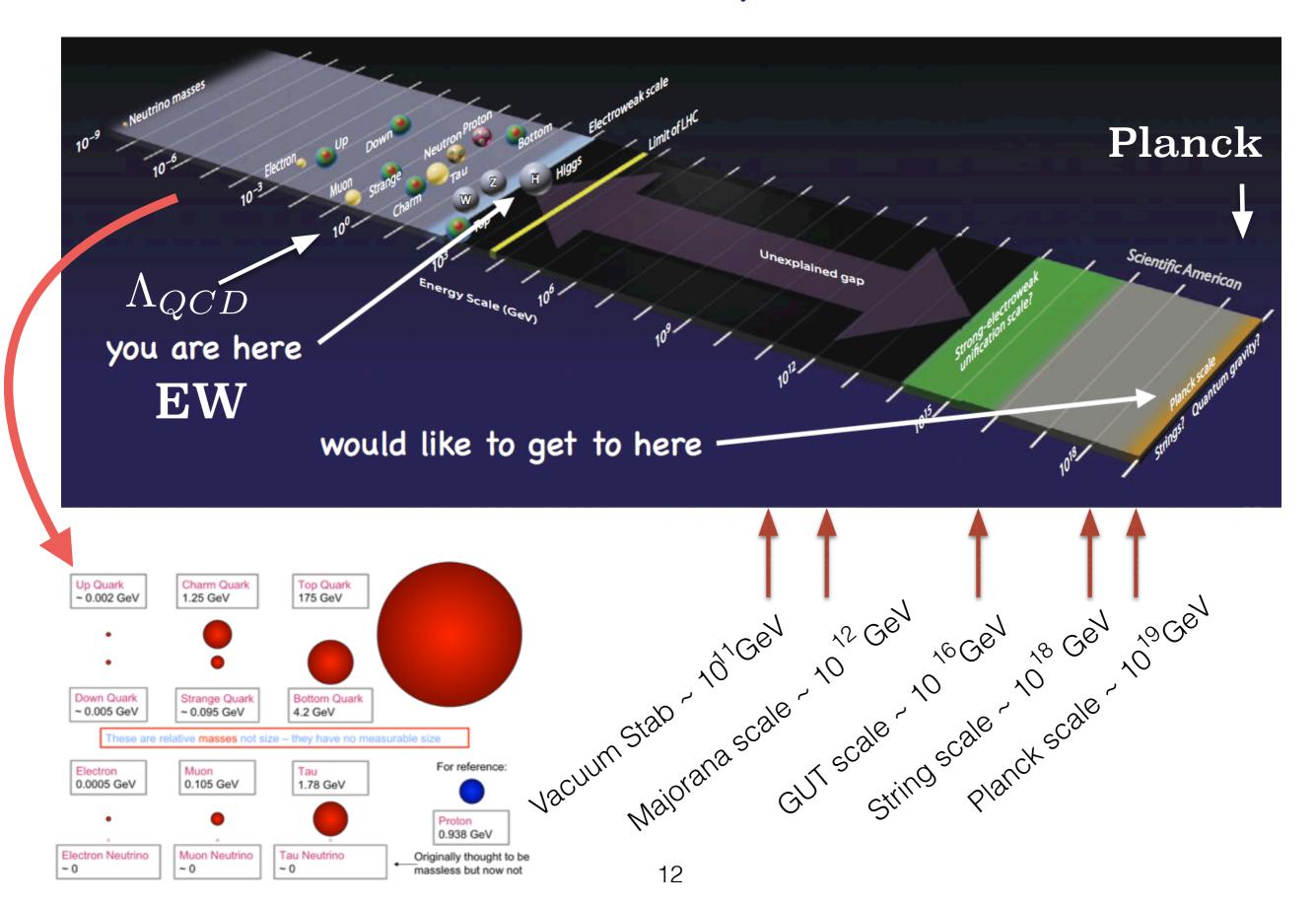
~ 0

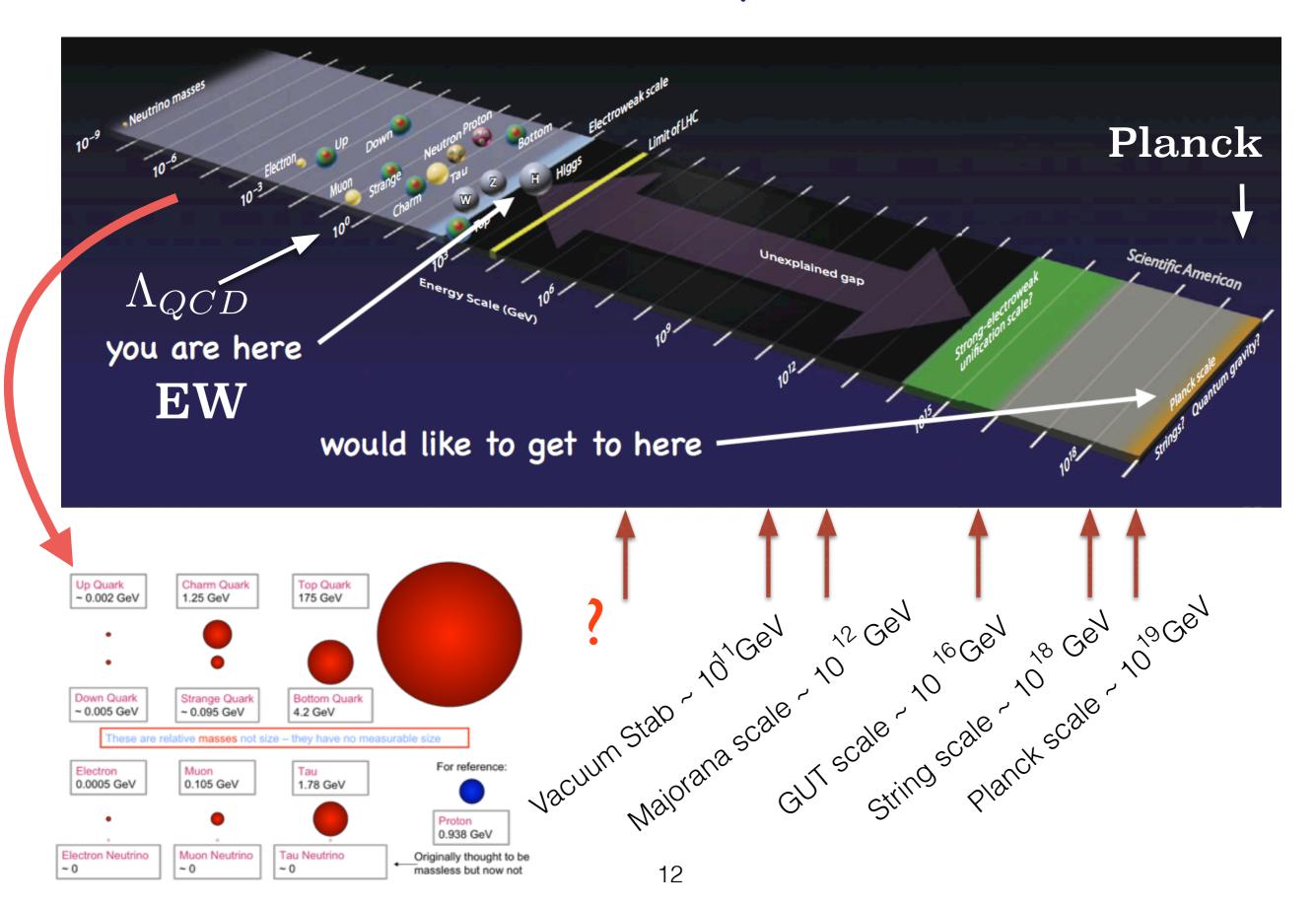


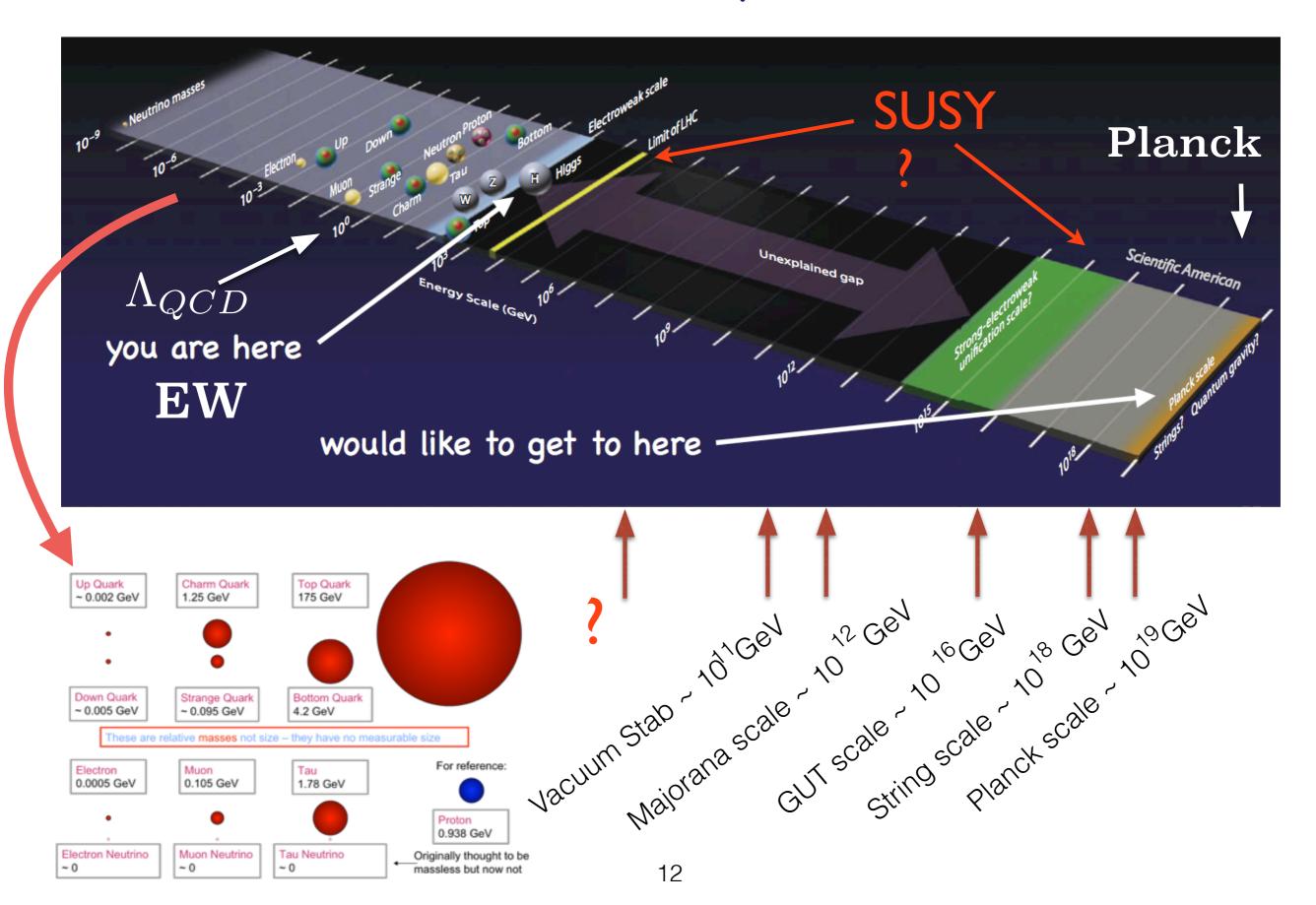


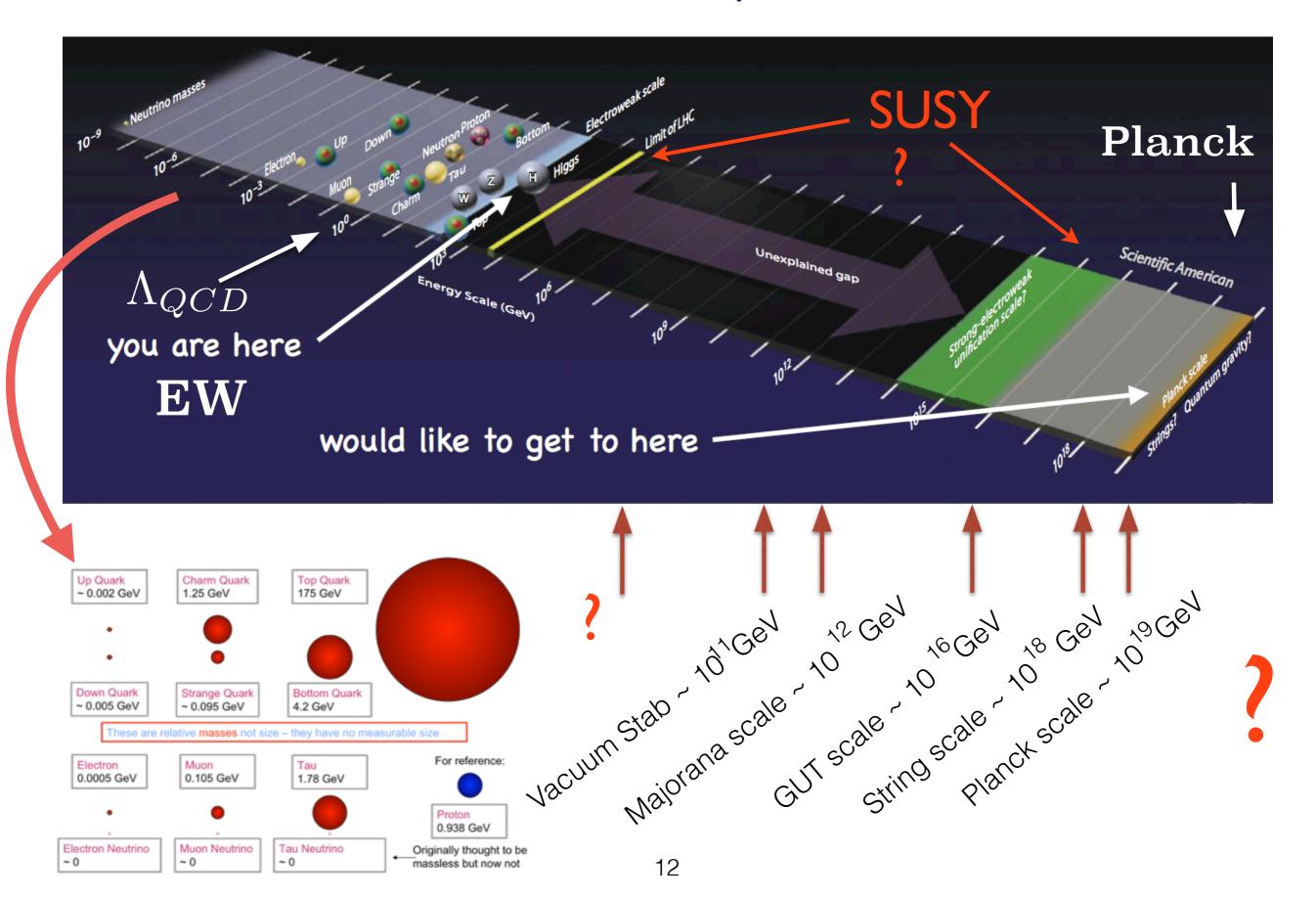




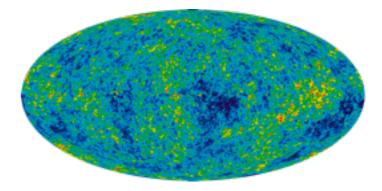






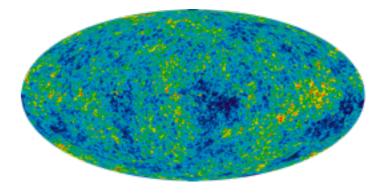


Astrophysics & Cosmology challenge



Astrophysics & Cosmology challenge

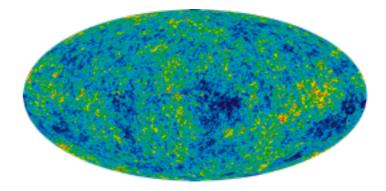
• Baryon asymmetry of the Universe



Astrophysics & Cosmology challenge

• Baryon asymmetry of the Universe

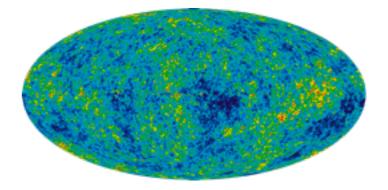
 $\frac{N(B) - N(\bar{B})}{N_{\gamma}} \sim (6.19 \pm 0.14) \times 10^{-10}$



Astrophysics & Cosmology challenge

• Baryon asymmetry of the Universe

$$\frac{N(B) - N(\bar{B})}{N_{\gamma}} \sim (6.19 \pm 0.14) \times 10^{-10}$$



- still not explained
- requires larger GP than in the SM

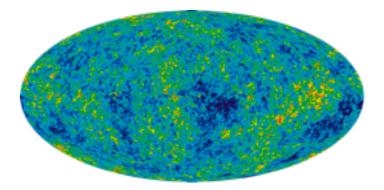
Astrophysics & Cosmology challenge

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• Relic abundance of the Dark Matter

 $OM = 4.9\%, \ DM = 26.8\%, \ DE = 68.3\%$



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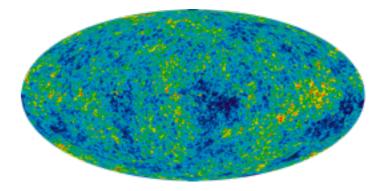
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Astrophysics & Cosmology challenge

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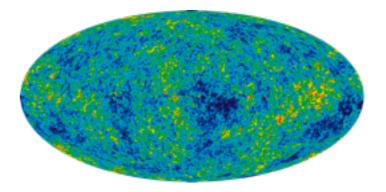
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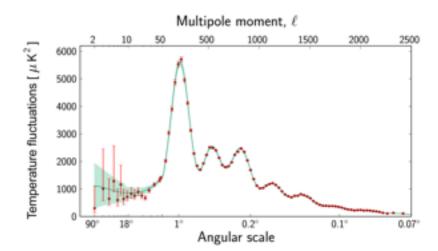
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• Number of neutrinos

 $N_{eff} = 3.52 \pm 0.47 \quad 95\% \ CL$ Planck + WP + highL + BAO + HST



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Astrophysics & Cosmology challenge

• Baryon asymmetry of the Universe

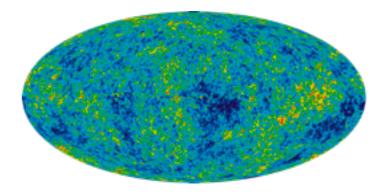
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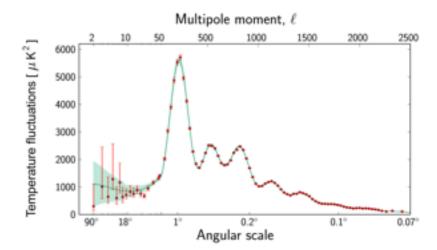
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• Well suits the SM q <-> |



Astrophysics & Cosmology challenge

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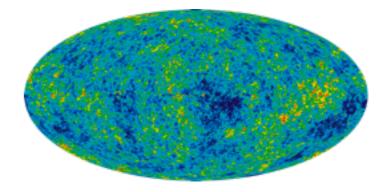
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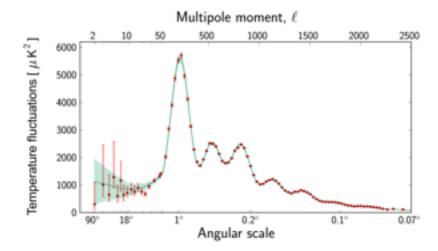
• Masses of neutrinos

 $\Sigma m_{\nu} \ [eV] < 1.11(0.22)$ Planck + WP + lensing + HST



- still not explained
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Astrophysics & Cosmology challenge

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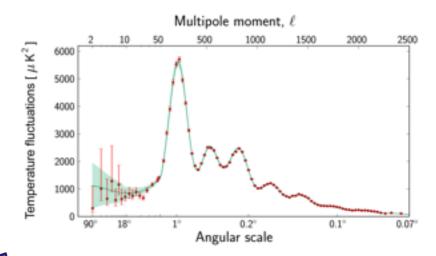
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Masses of neutrinos

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- requires larger GP than in the SM
- Understanding is beyond the SM

- Well suits the SM
 q <-> |
- Probably a hint towards new physics

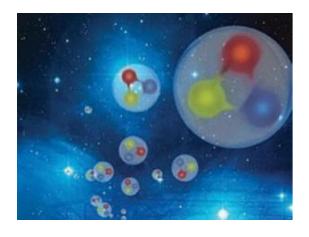


Challenging problem in particle physics well inside the SM





Challenging problem in particle physics well inside the SM





Challenging problem in particle physics well inside the SM

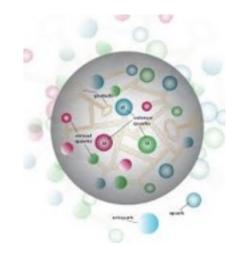
- How confinement actually works?
- Why colourless states?
- Which bound states exist in Nature?

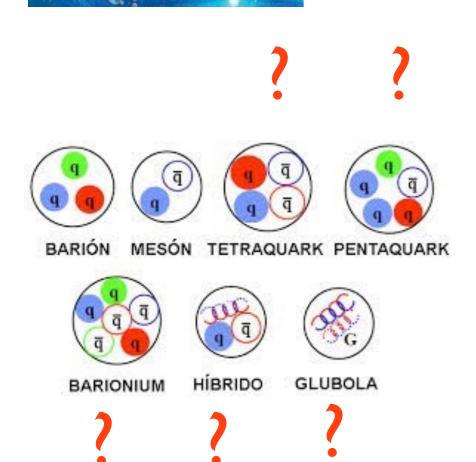




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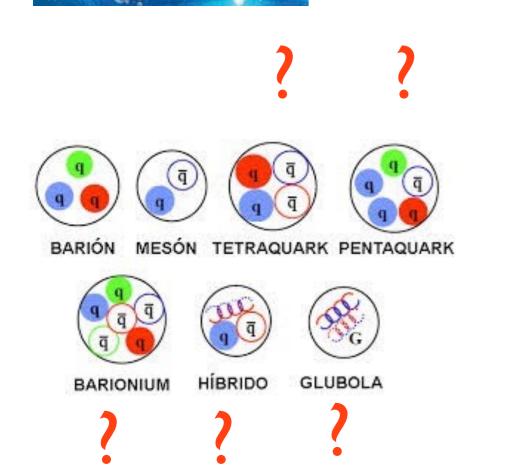


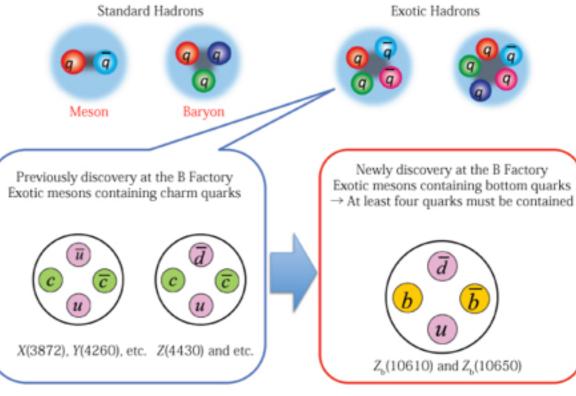


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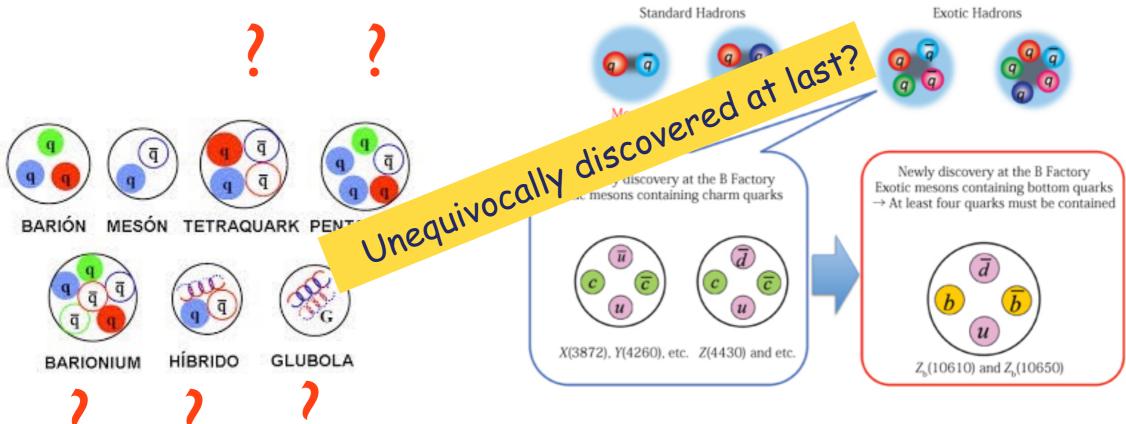




Challenging problem in particle physics well inside the SM

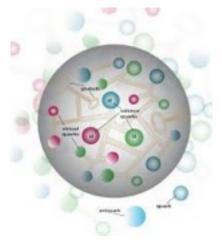
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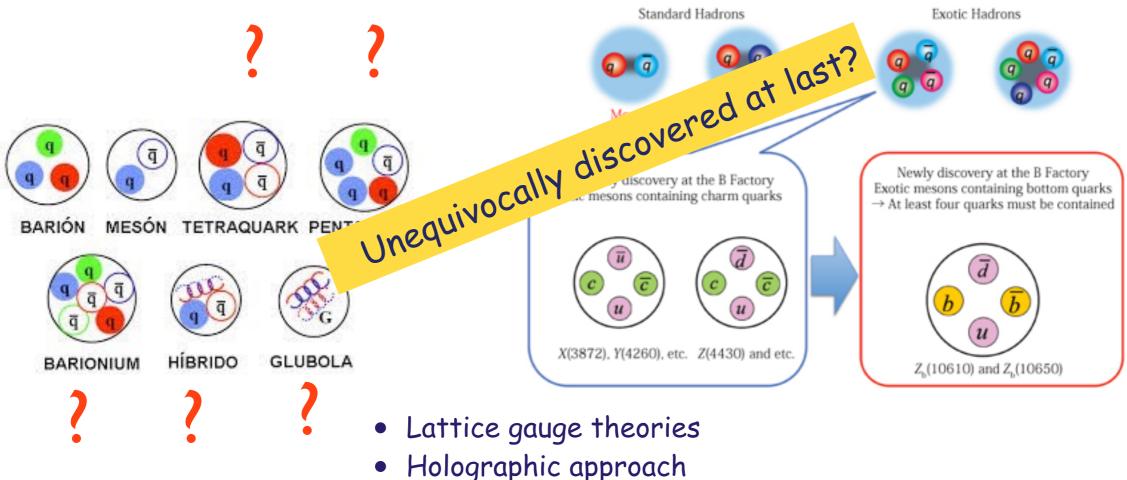




Challenging problem in particle physics well inside the SM

- How confinement actually works?
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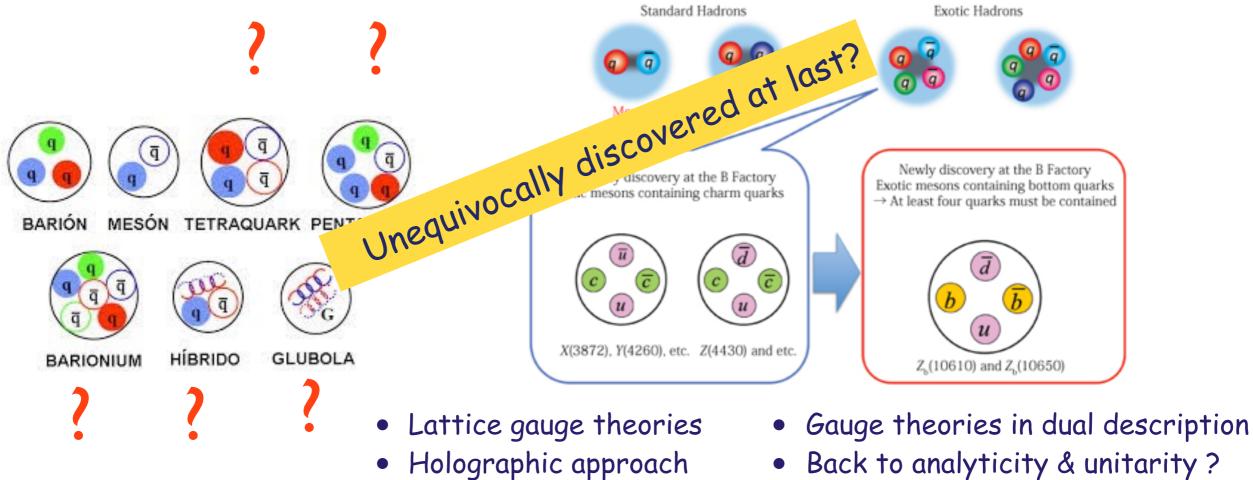
Do we understand confinement?

Challenging problem in particle physics well inside the SM

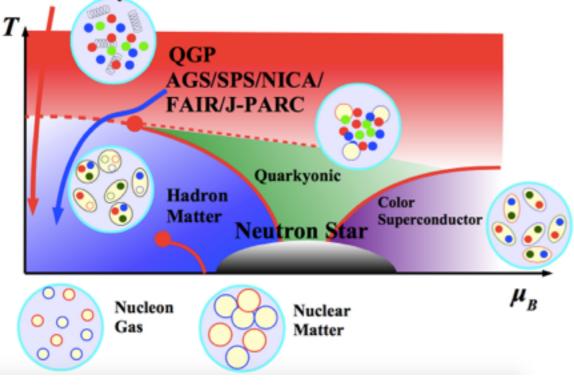
Time to come back?

- How confinement actually works?
- Why colourless states?
- Which bound states exist in Nature?

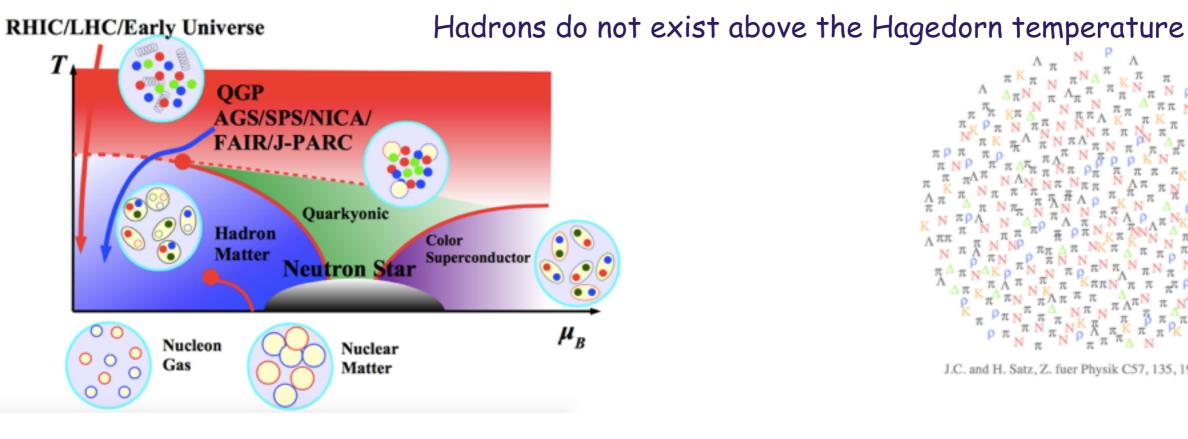


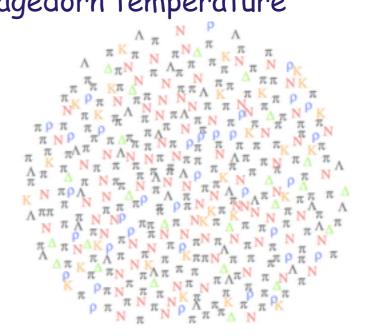


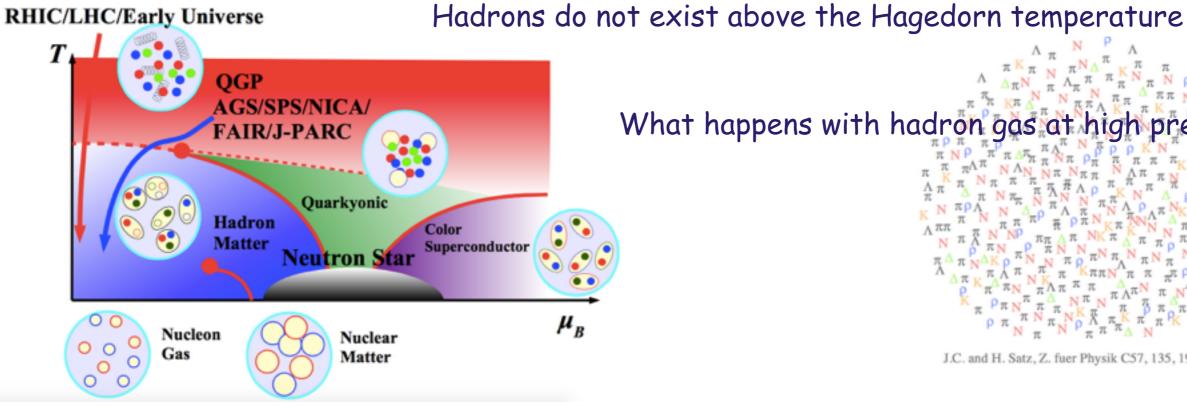
RHIC/LHC/Early Universe



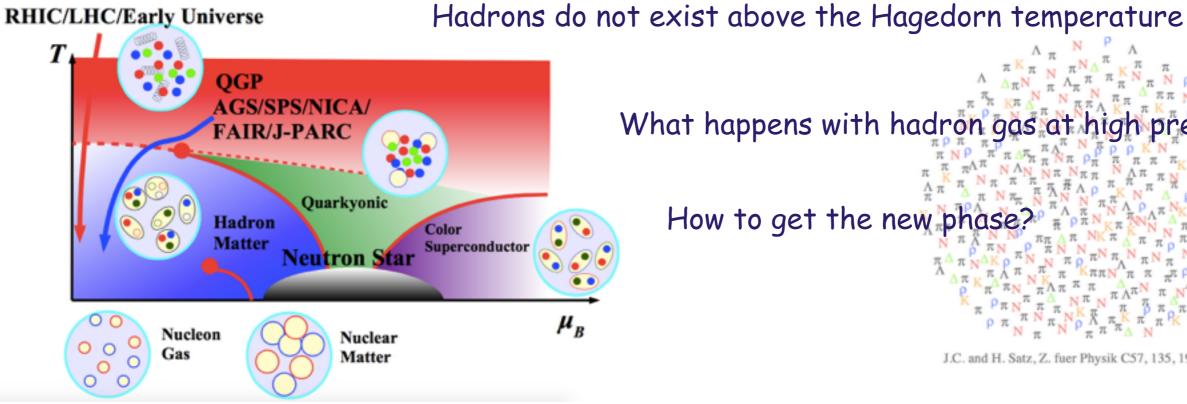




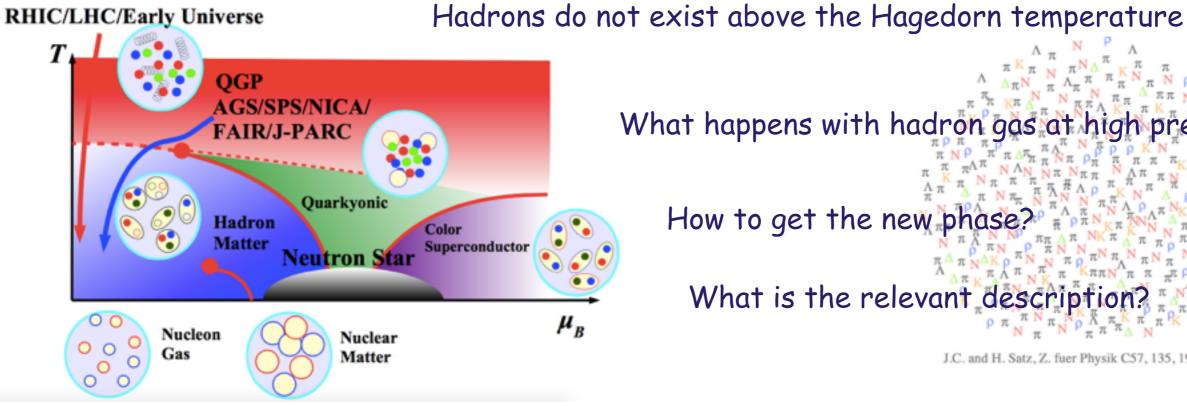




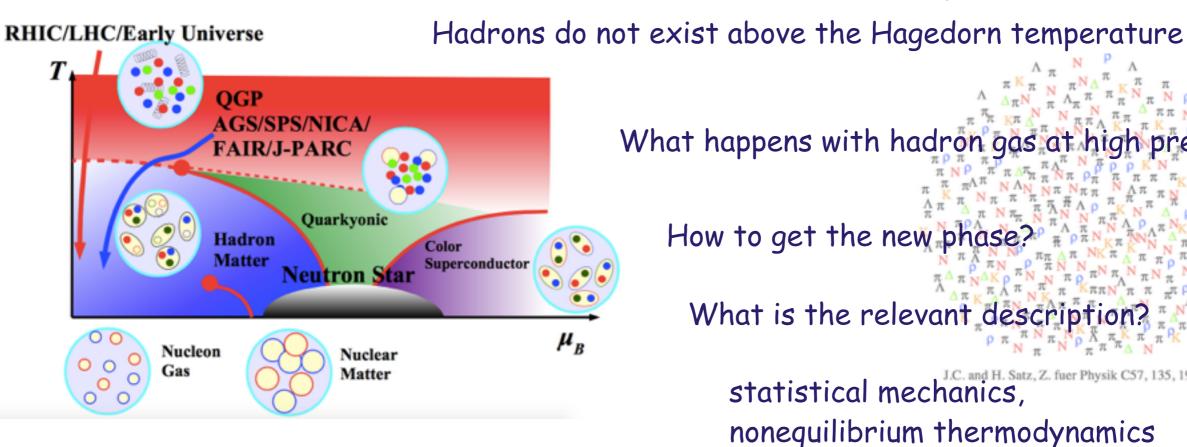
What happens with hadron gas at high pressure?



What happens with hadron gas at high pressure? How to get the new phase?

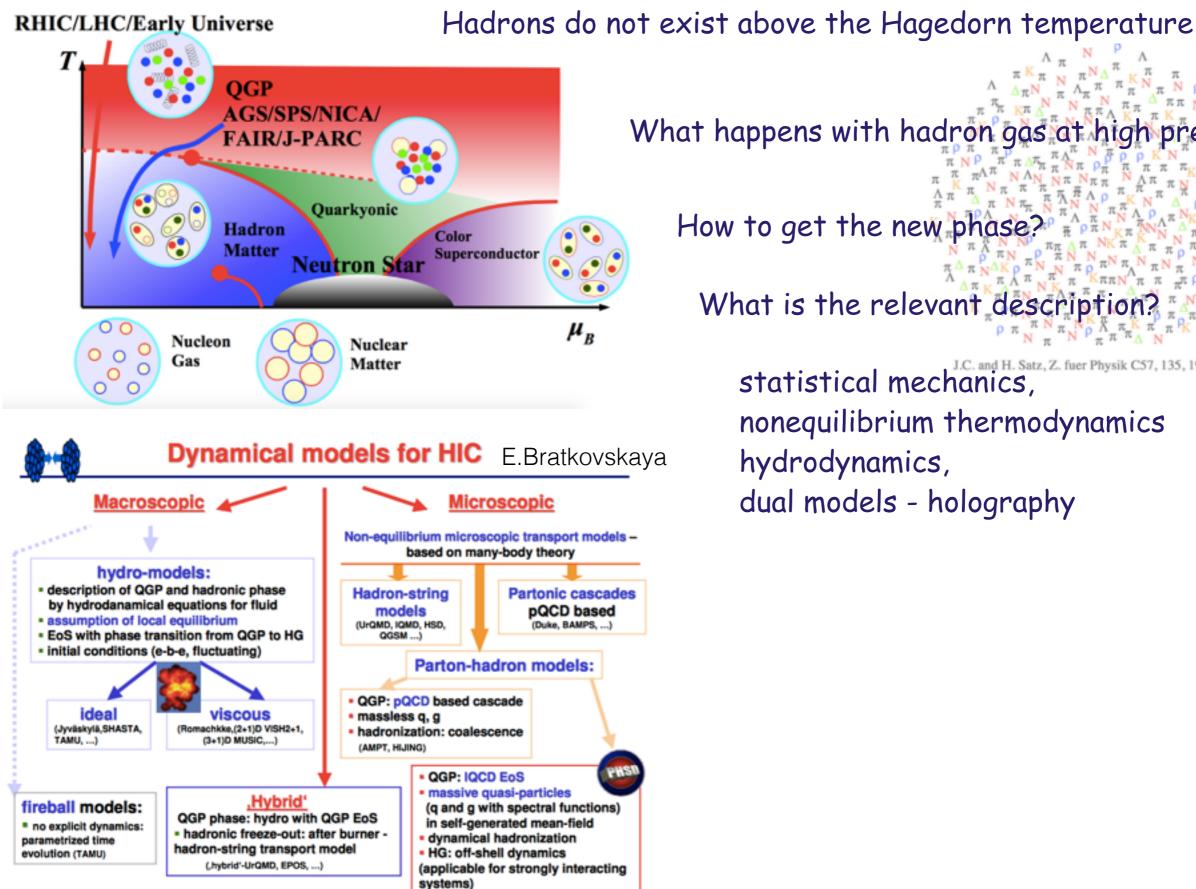


What happens with hadron gas at high pressure? How to get the new phase? What is the relevant description J.C. and H. Satz, Z. fuer Physik C57, 135, 1993.



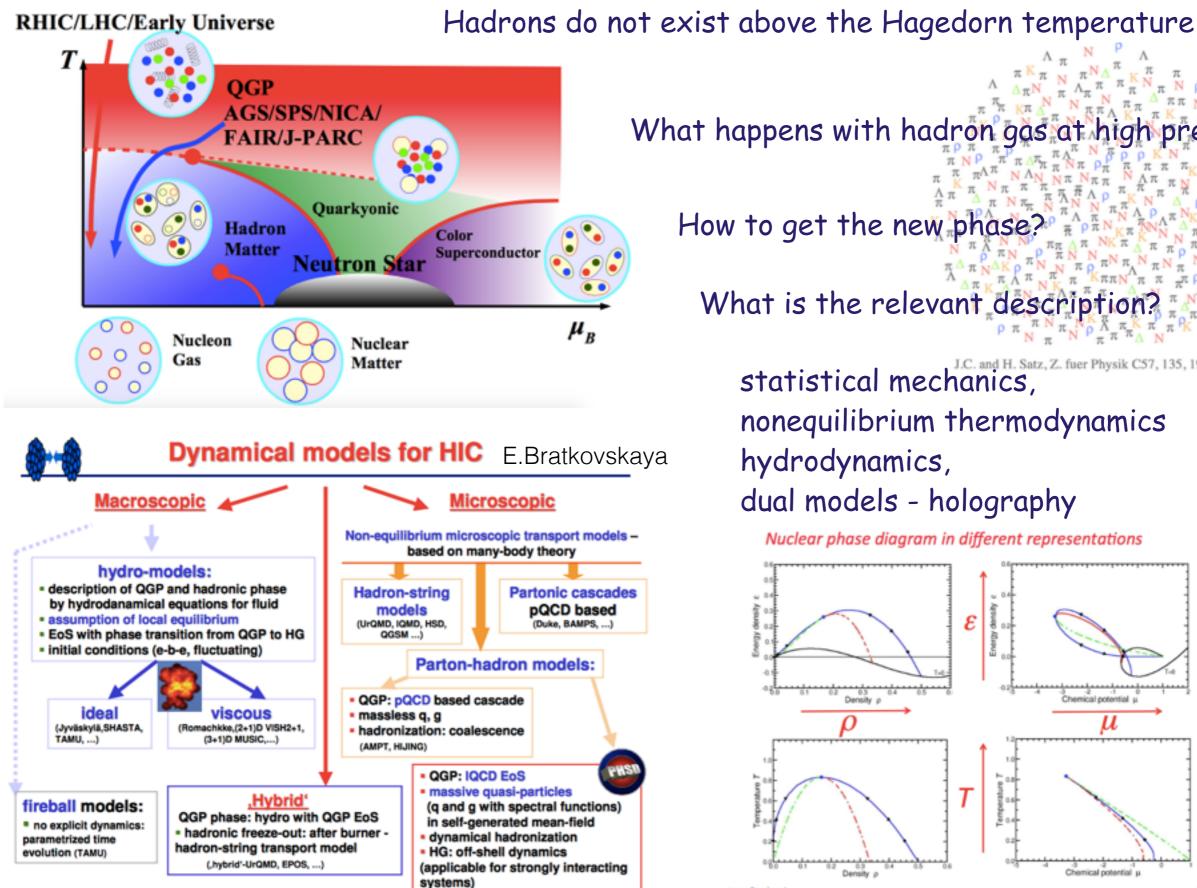
What happens with hadron gas at high pressure? How to get the new phase? What is the relevant description statistical mechanics,

nonequilibrium thermodynamics hydrodynamics, dual models - holography



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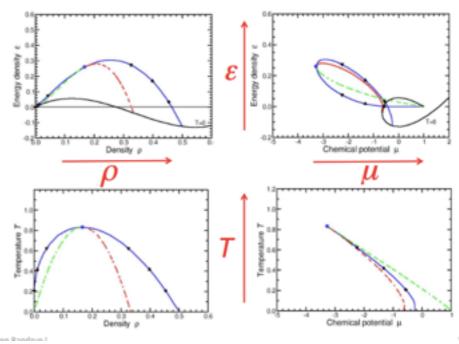
statistical mechanics, nonequilibrium thermodynamics hydrodynamics, dual models - holography

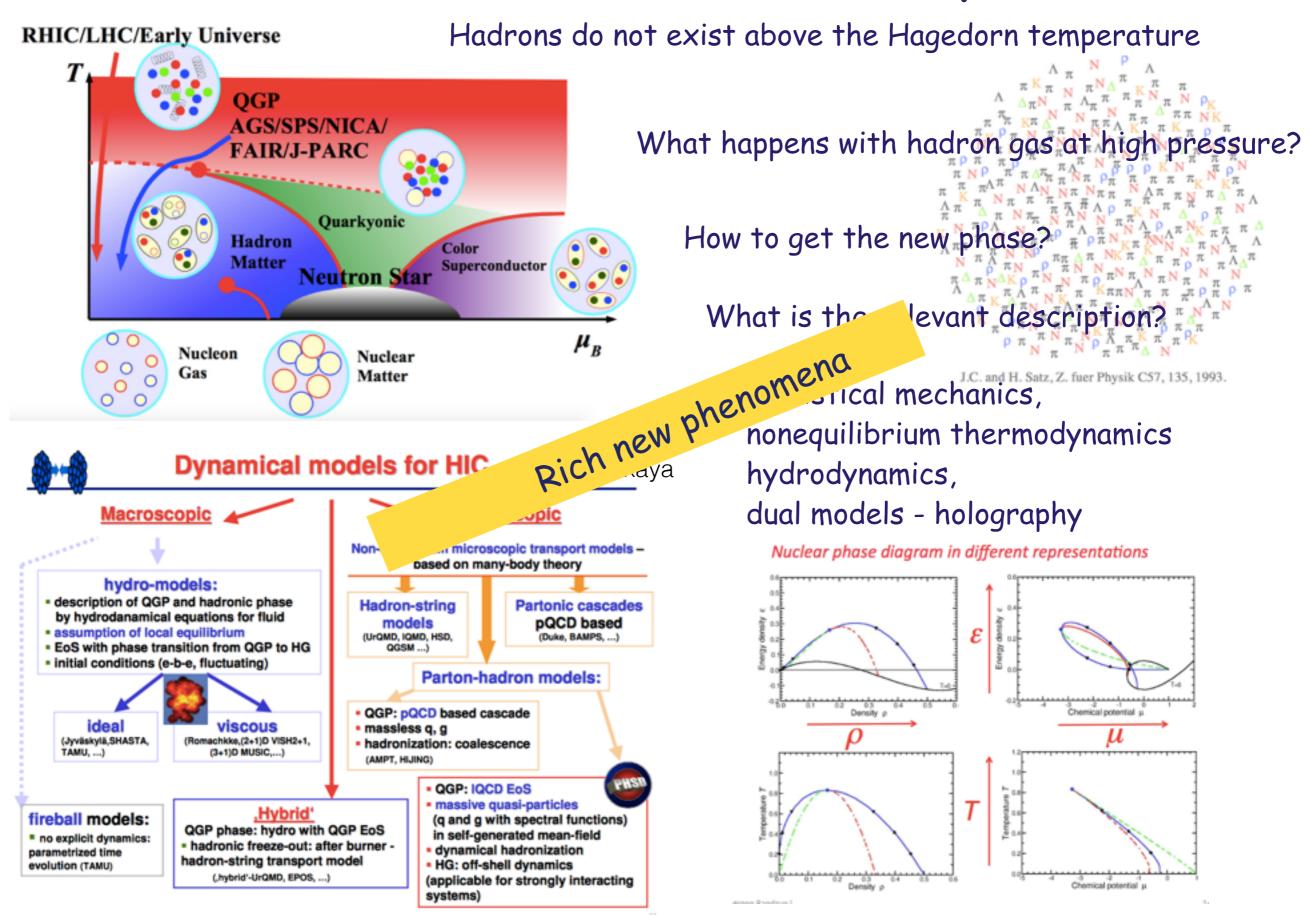


What happens with hadron gas at high pressure? How to get the new phase? What is the relevant description?

J.C. and H. Satz, Z. fuer Physik C57, 135, 1993. statistical mechanics, nonequilibrium thermodynamics hydrodynamics, dual models - holography

Nuclear phase diagram in different representations









The Higgs Boson - Target #1

Is it the SM Higgs boson or not?



The Higgs Boson - Target #1

Is it the SM Higgs boson or not?

What are the alternatives?



Is it the SM Higgs boson or not?

What are the alternatives?

- A. Singlet extension
- B. Higgs doublet extension
- C. Higgs triplet extension



- A. Singlet extension
- B. Higgs doublet extension
- C. Higgs triplet extension

Custodial symmetry as guiding principle for extensions

indicates that an approximate global symmetry exists,

$$\rho = \frac{M_W^2}{M_Z^2 \cos^2 \theta_W} = 1$$

broken by the vev to the diagonal 'custodial' symmetry

group $SU(2)_L \times SU(2)_R \rightarrow SU(2)_{L+R}$





Is it the SM Higgs boson or not? What are the alternatives?

- A. Singlet extension
- B. Higgs doublet extension
- C. Higgs triplet extension

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Thus the Higgs field transforms under $SU(2)_L \times SU(2)_R : \Phi \to L\Phi R^{\dagger}$

 $\rho = \frac{\sum_{i=1}^{n} [I_i(I_i+1) - \frac{1}{4}Y_i^2]v_i}{\sum_{i=1}^{n} \frac{1}{2}Y_i^2v_i} \sim 1$ For both SU(2)-singlet with Y=0
M.Spannowsky
and SU(2) doublet with Y=+-1



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 \boldsymbol{n}

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$$\text{For both SU(2)-singlet with Y=0}$$

$$M.Spannowsky$$

Any number of singlets and doublets respects custodial symmetry at tree level. Not so for arbitrary triplet models ...



The Higgs Boson - Target #1

Is it the SM Higgs boson or not?



The Higgs Boson - Target #1

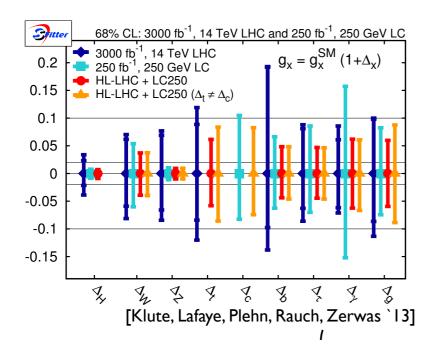
Is it the SM Higgs boson or not?

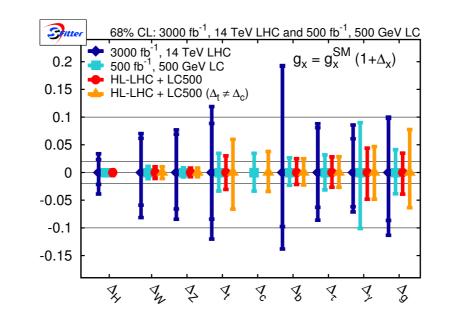
How to probe?



The Higgs Boson - Target #1

- Is it the SM Higgs boson or not?
 - How to probe?
 - Probe deviations from the SM Higgs couplings

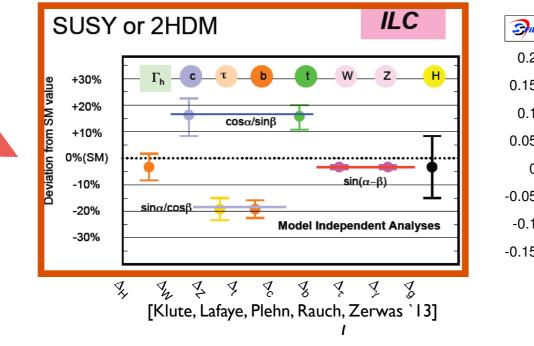


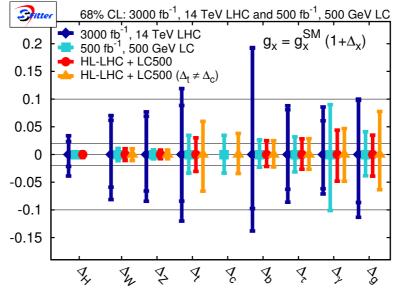




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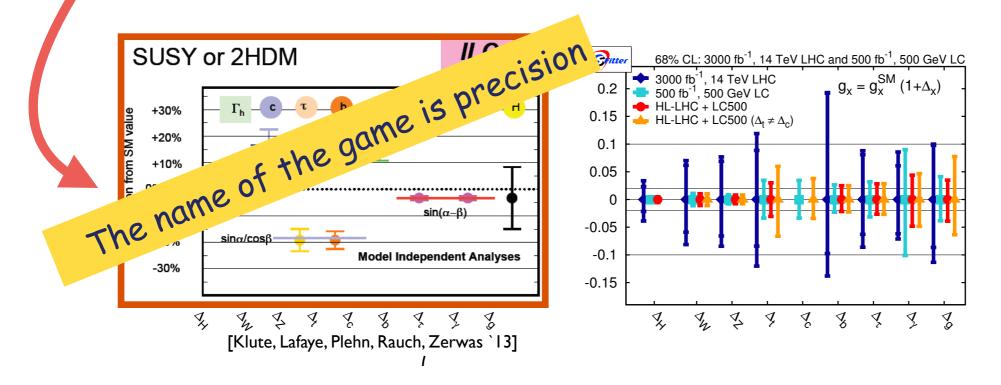


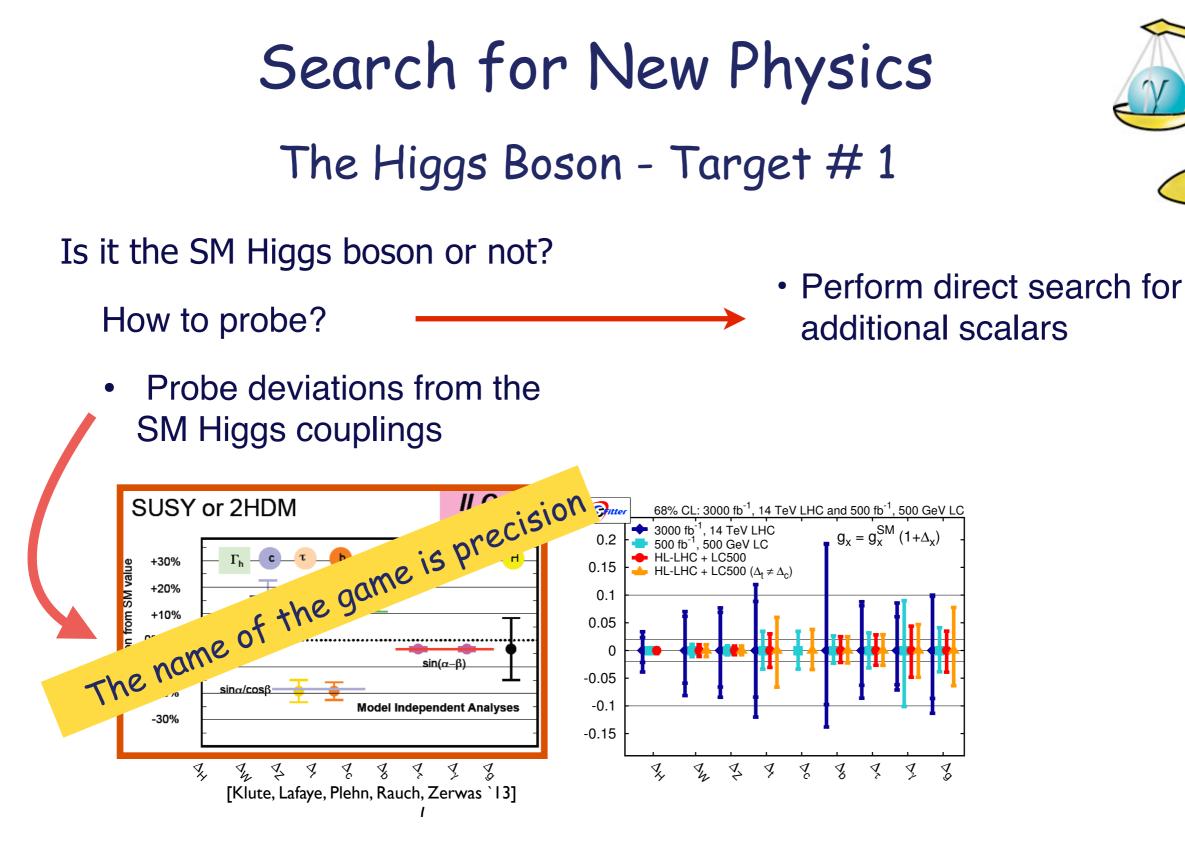




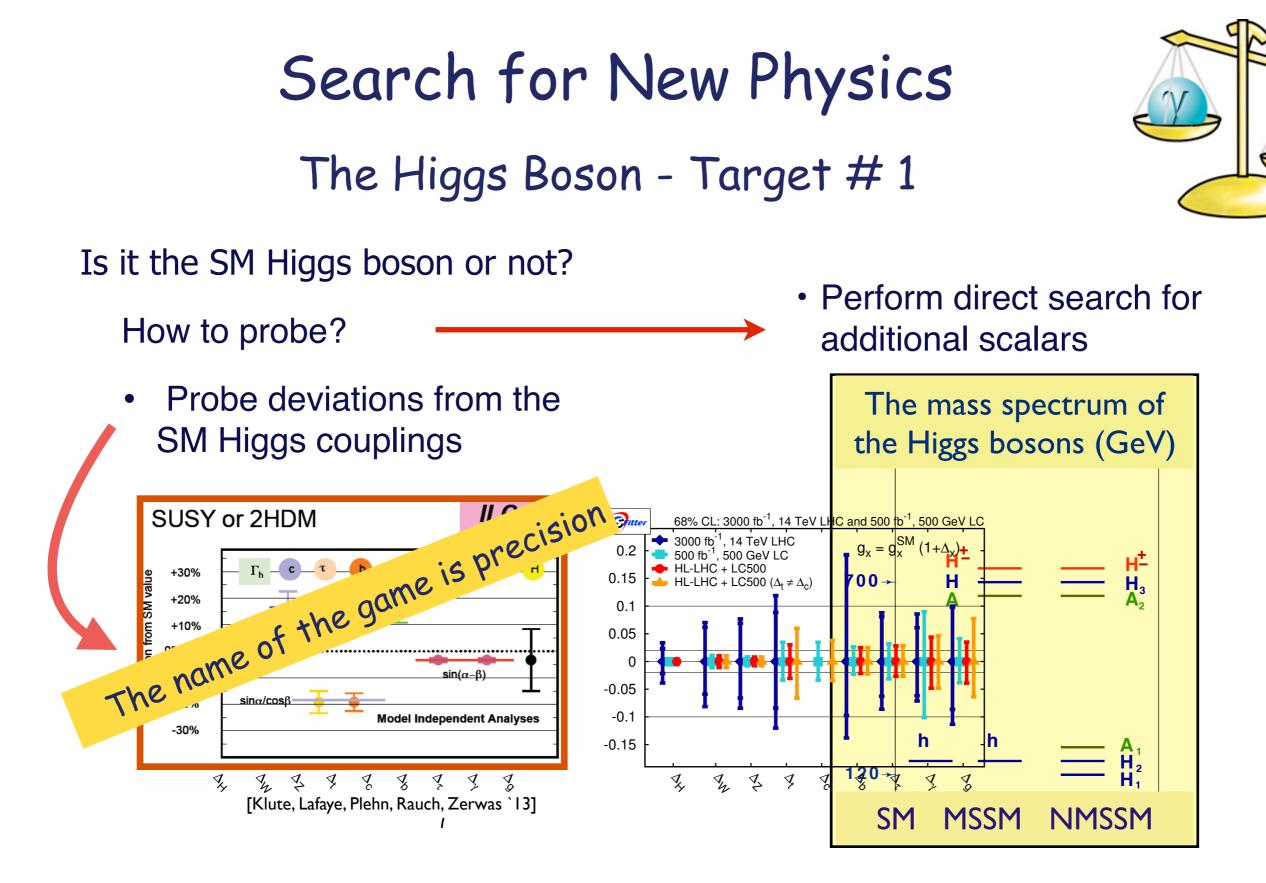
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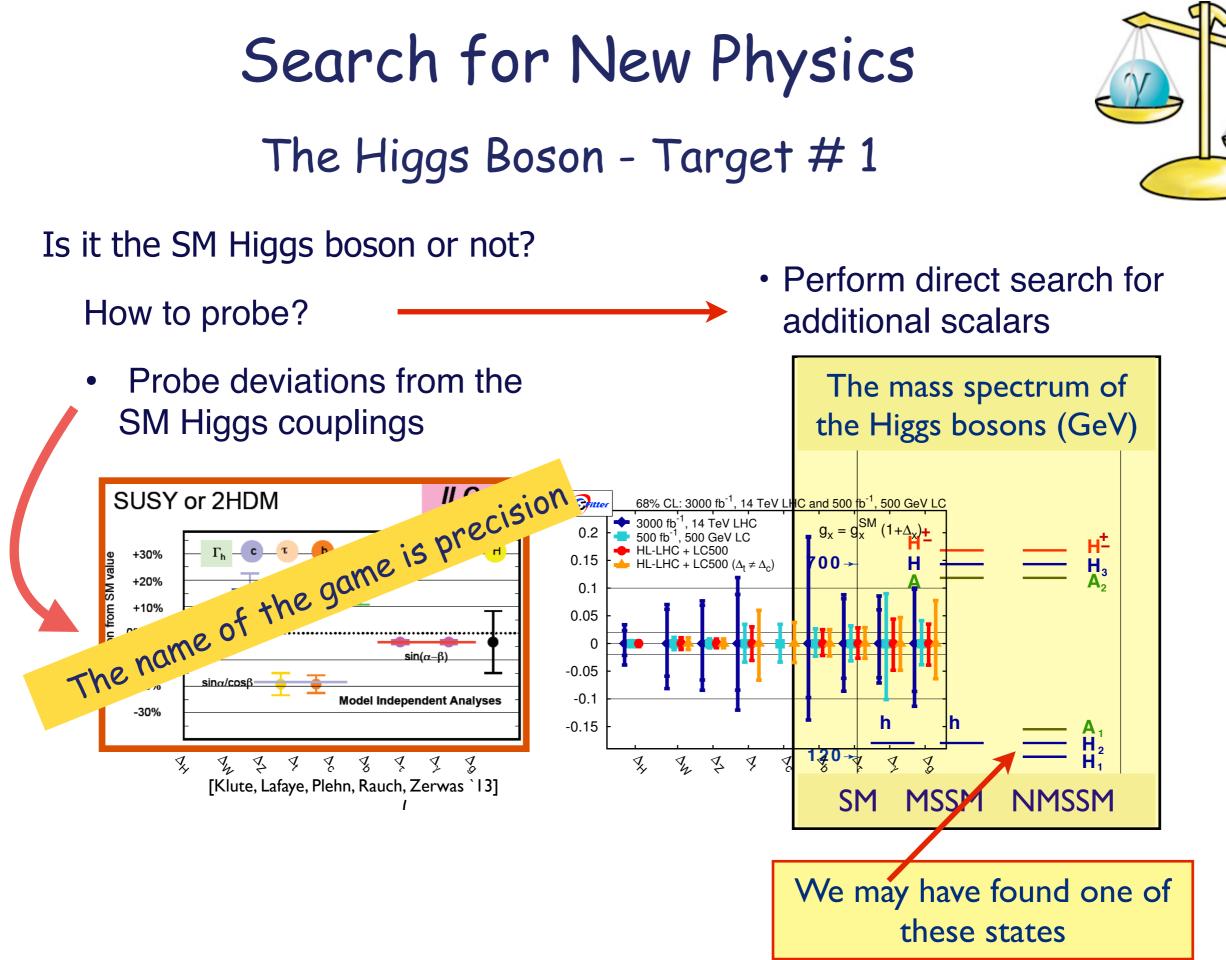
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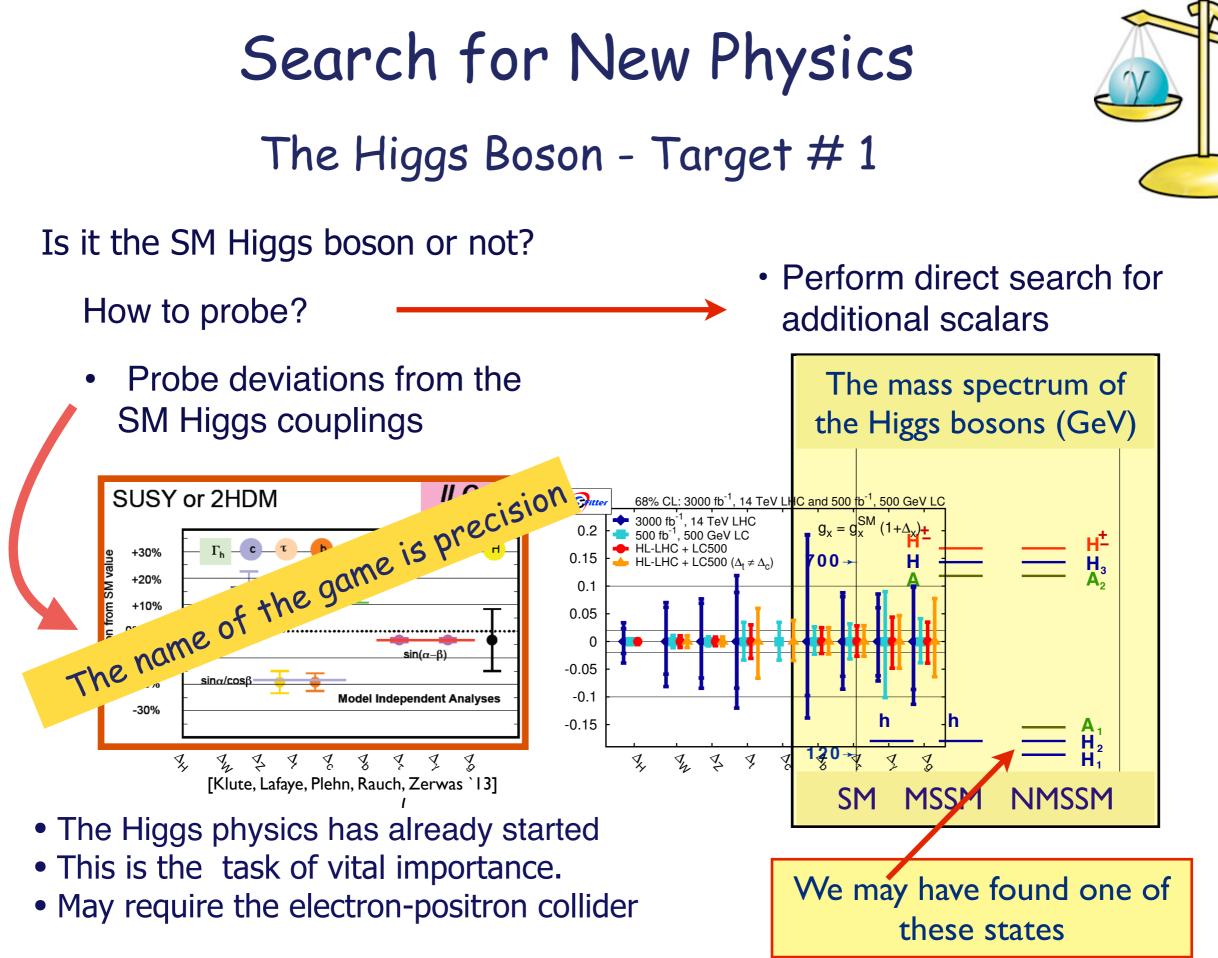




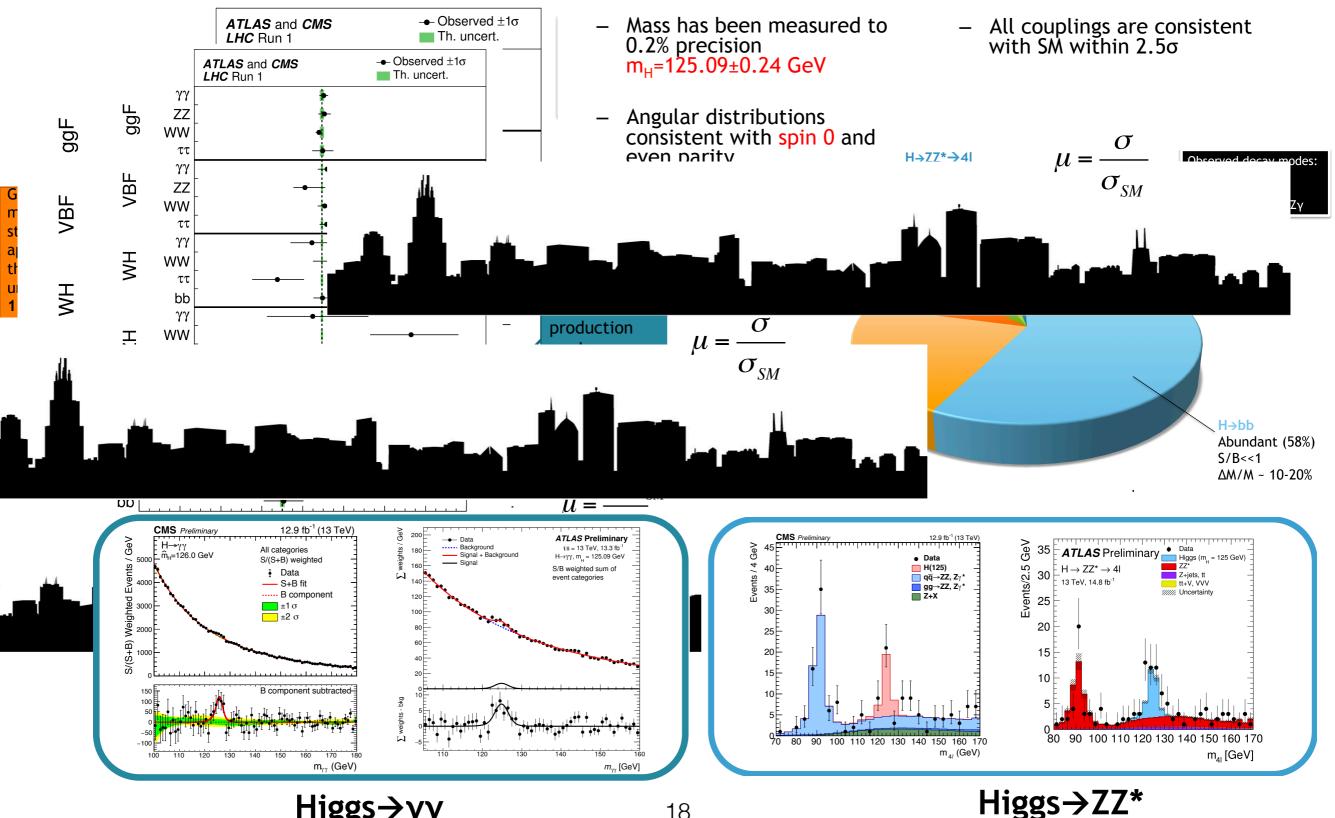








Higgs Boson (125)



Higgs→γγ

18

Extra Higgs Bosons

Heavy Higgs \rightarrow ZZ \rightarrow 4l Higgs→hh→bbττ CMS Preliminary 12.9 fb⁻¹ (13 TeV) Resonance search Non-resonance search GeV ATLAS Preliminary Data Data 20 Only SM $H \rightarrow ZZ^* \rightarrow 4I$, inclusive ZZ* New Physics 20 H(125) tī+V. VVV 13TeV, 14.8 fb⁻¹ □ qq→ZZ, Zγ* □ gg→ZZ, Zγ* □ Z+X $\overline{\mathbf{w}}$ 70F $\overline{}$ Z+Jets, tī ///// Uncertainty Interference 000 مع ATLAS Preliminary Data $\overline{\mathbf{w}}$ ZZ* $H \rightarrow ZZ^{\star} \rightarrow 4I,$ inclusive tī+V, VVV 13TeV, 14.8 fb⁻¹ Z+Jets, tť Uncertainty معف h 10 300 400 500 200 600 700 800 m₄₁ (GeV) Non-Resonant Charged Resonant 12.9 fb⁻¹ (13 TeV) 12.9 fb⁻¹ (13 TeV) bbचt) [fb] 10 10 BR(H→ hh→ bbπτ) [pb] CMS CMS bb $\mu\tau_{h}$ + bb $e\tau_{h}$ + bb $\tau_{h}\tau_{h}$ bb $\mu\tau_{h}$ + bb $e\tau_{h}$ + bb $\tau_{h}\tau_{h}$ ~60 m preliminary preliminary Combined channels Combined channels <u> ഫ</u>60 പ 200 300 400 500 600 700 800 m, [GeV] 10 षेत्र इ BR(hh→ Search for $H^{\pm} \rightarrow tb$ 50 300<m_H[±]<1000 Ge 10 45 š 40 10 u 10² дX 35 ATLAS Preliminary Ч ATLAS Preliminary Observed x SM Ы 200 Observed \s= 13 TeV, 14.7 fb⁻¹ Expected CLs 30 95% ---- Expected CLs ਹੋ¹⁰ \s= 13 TeV, 14.7 fb⁻¹ 10 $H^+ \rightarrow \tau v$: hMSSM scenario Expected $\pm 1\sigma$ 25 <u>~</u>60 $H^+ \rightarrow \tau v$; hMSSM scenario Expected $\pm 2\sigma$ Observed exclusion 95% 2015 result ···· Expected exclusion 20 Observed exclusion <u>क</u> 55 10 Observed ± 1σ 300 400 500 600 700 800 90(..... Expected exclusion 15 E ---- Expected <u>±</u> 2σ m_H [GeV] 50 10 200 250 300 350 400 450 500 550 600 m_μ· [GeV] ± 1σ <u>±</u> 2σ Heavy Higgs→ττ ···· 550 600 n...⊧ [GeV] ≯ττ)[pb] 35 ATLAS Preliminary Events / GeV ATLAS Preliminary - Data ATLAS Preliminary \s= 13 TeV, 14.7 fb⁻¹ Η/Α→ττ 30 10E √s = 13 TeV, 13.2 fb⁻ $m_A = 600 \text{ GeV}, \tan \beta = 20$ $H^+ \rightarrow \tau v$; hMSSM scenario MSSM m Observed exclusion 10 $H/A \to \tau_{had} \tau_{had}$ Multi-jet 25 5.2 fb⁻¹ (13 TeV) Observed exclusion --- Expected exclusion $Z \rightarrow \tau \tau$ BR(H/Ah-vetc H⁺→ tb W→tv + jets 2015 result Expected exclusion - Observed 20 ± 1σ tī, single top vs = 13 TeV, 13.2 fb --- Observed ± 1σ -- Expected 10 Others ± 20 15E **± 1**σ Uncertainty ---- Expected <u>±</u> 2σ Pre-fit background š <u>±</u>2σ 10 •تت)[pb] 10-ATLAS Preliminary m_{H⁺} [GeV] ATLAS Preliminary — Data Observed --- Expected $H/A \rightarrow \tau\tau$, 95 % CL limits $'10^{3}$ √s = 13 TeV, 13.2 fb 400 500 600 700 800 900 300 1000^{H±} m_{a} = 600 GeV, tan β = 20 ±10 √s = 13 TeV, ≤ 13.3 fb⁻¹ $H/A \rightarrow \tau_{had} \tau_{had}$ stu 10² 10^{-2} ± 20 m_{H*} [GeV] aluon-aluon fusion 2015, 3.2 fb⁻¹ (Obs BR(H/Ab-veto tī, single to х 600 Data/Pred 1.0 2.0 ь ATLAS Preliminary ğ 400 1 m^{mod} Observed exclusion 200 30 200 10-1 ······ Expected exclusion b $m_{\rm T}^{\rm tot}$ 200 600 400 800 1000 ± 1σ m_{µ⁼} [GeV] τ_{had}τ_{had} (Exp.) 3 TeV, 13.2 fb⁻¹

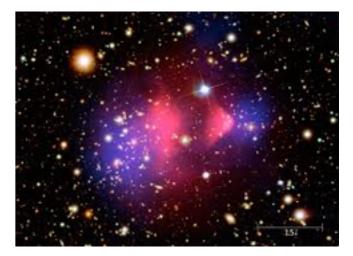
The Dark Matter - Target # 2

- The Dark Matter is made of:
- Macro objects Not seen

Not from

the SM

- New particles right heavy neutrino
 - axion (axino)
 - neutralino mSUGRA
 - sneutrino
 - gravitino
 - heavy photon
 - heavy pseudo-goldstonelight sterile higgs



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not favorable but possible



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WIMP

Not from

the SM

The Dark Matter - Target # 2

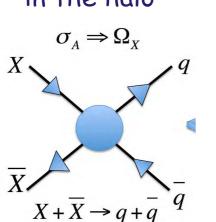
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 - gravitino
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Annihilation in the halo

WIMP

Not from

the SM



not favorable but possible might be invisible (?) detectable in 3 spheres less theory favorable

might be undetectable (?)

possible, but not related to the other models



The Dark Matter - Target # 2

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 - gravitino

Annihilation

- heavy photon
- heavy pseudo-goldstone
- light sterile higgs

not favorable but possible might be invisible (?) detectable in 3 spheres less theory favorable

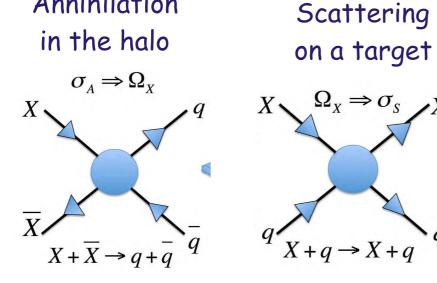
might be undetectable (?)

possible, but not related to the other models



Not from

the SM





The Dark Matter - Target # 2

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- Macro objects Not seen

Not from

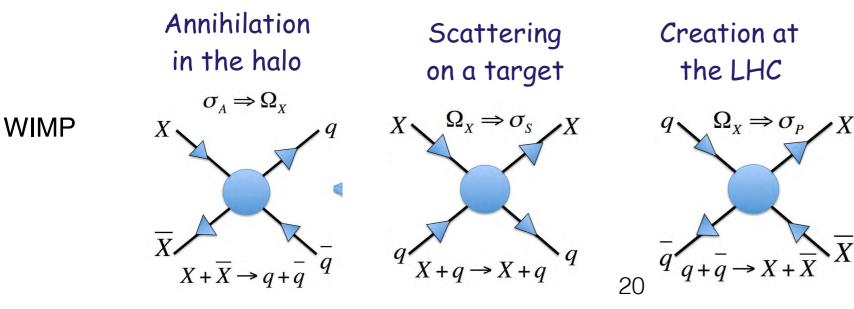
the SM

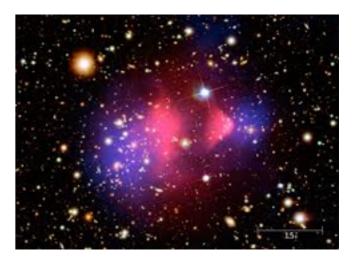
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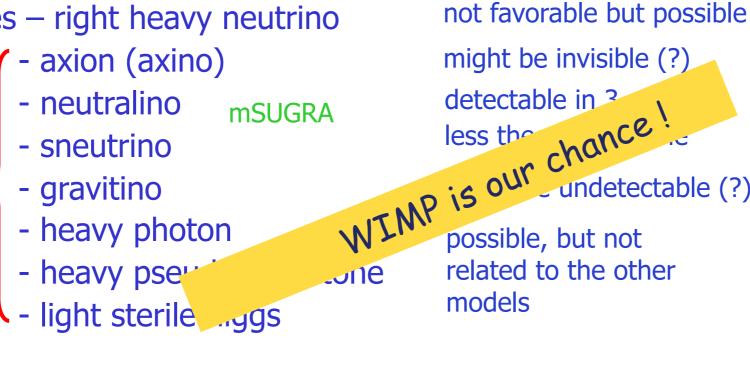
The Dark Matter - Target # 2

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X

New particles – right heavy neutrino

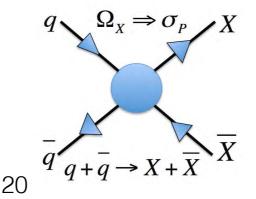




Annihilation Scattering in the halo on a target $\sigma_A \Rightarrow \Omega_X$ $\Omega_x \Rightarrow \sigma_s X$ $X + \overline{X} \rightarrow q + q$ $X + a \rightarrow X + a$

Creation at the LHC

undetectable (?)





WIMP

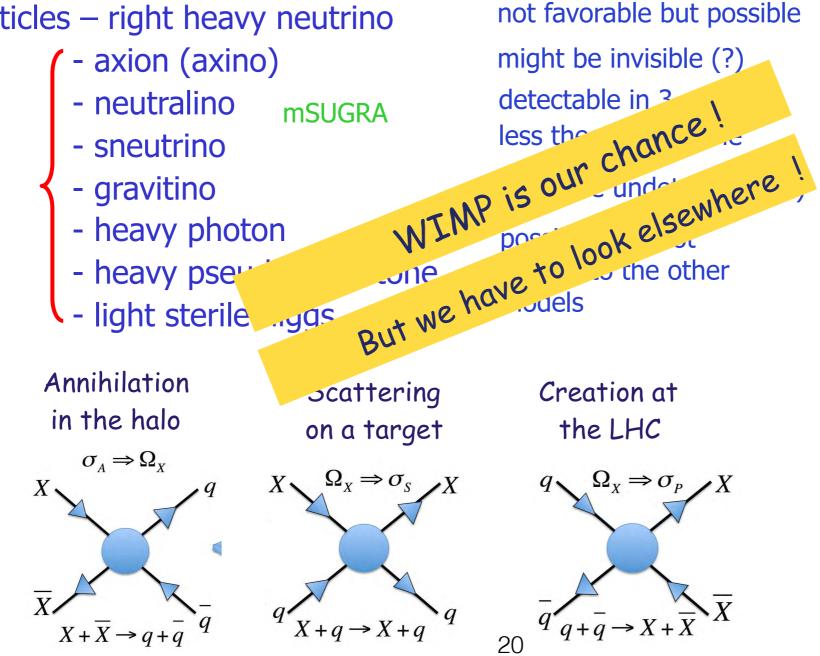
The Dark Matter - Target # 2



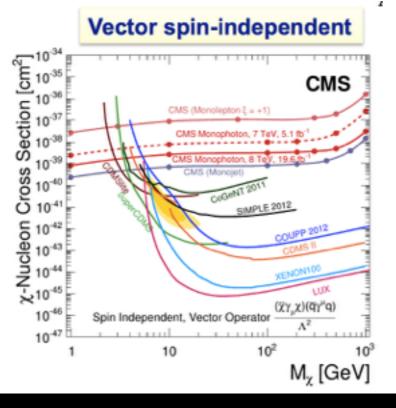
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WIMP

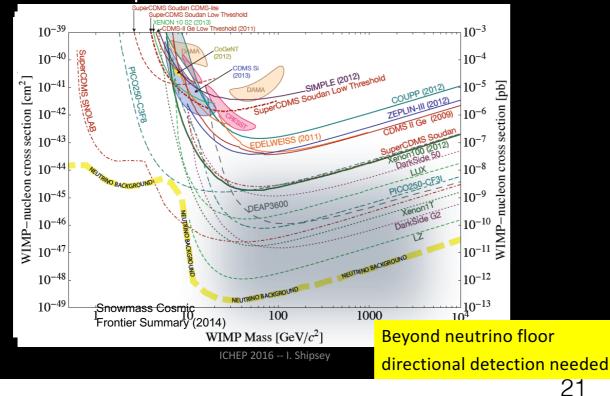




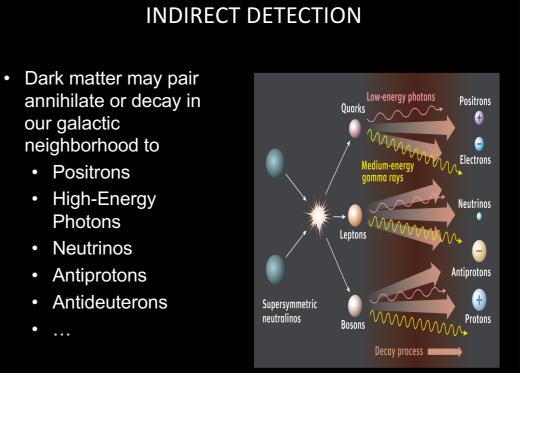


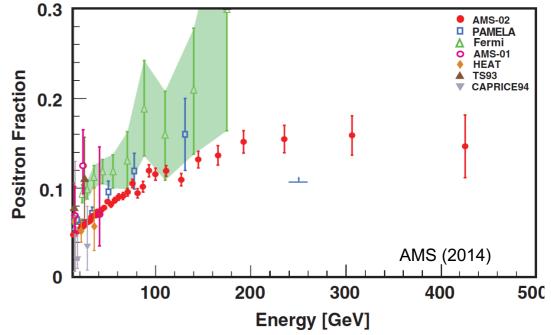
DIRECT DETECTION: STATUS AND PROSPECTS

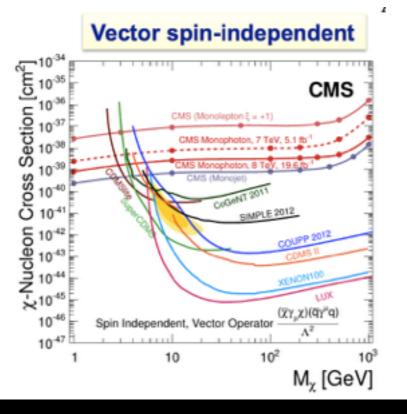
- Since 2010, sensitivity improved by ~100 (for m ~ 100 GeV)
- Further improvements by 2-3 orders of magnitude expected by a suite of experiments world-wide



DM Searches







DM Searches

our galactic

Positrons

Photons

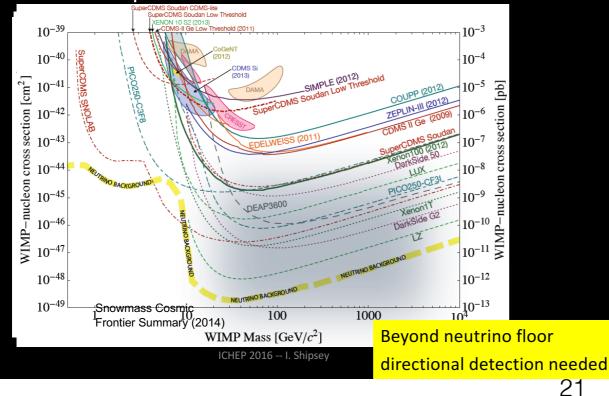
Neutrinos

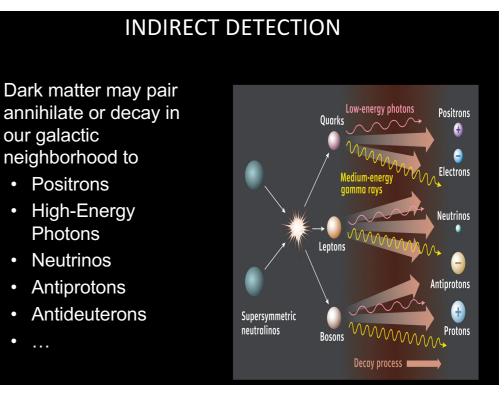
0

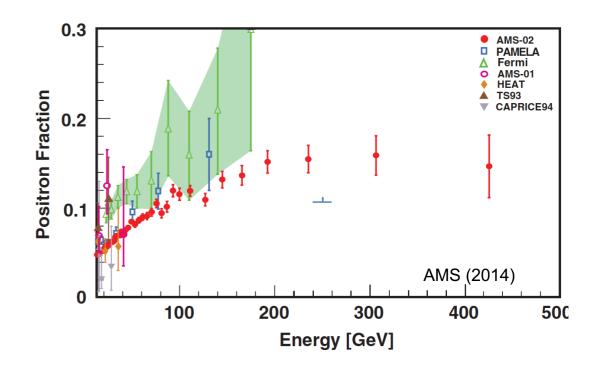
Already close to neutrino floor

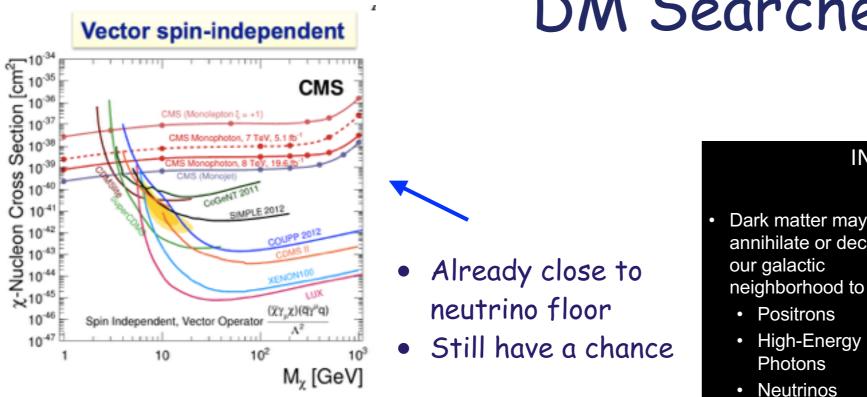
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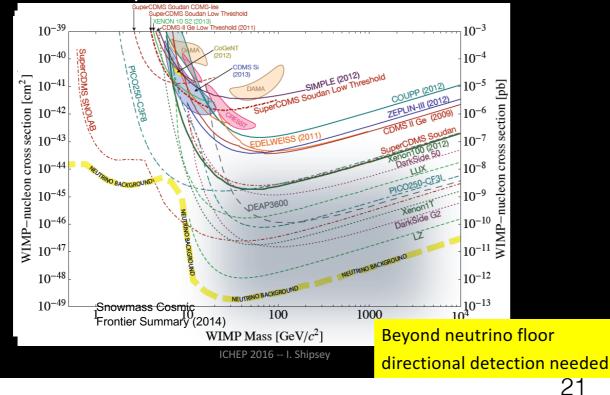






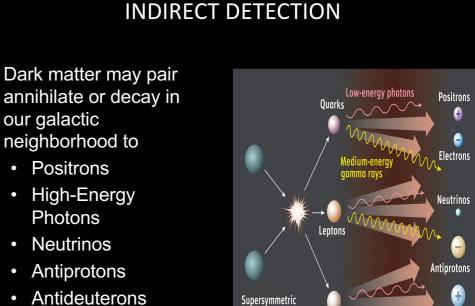
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DM Searches

•

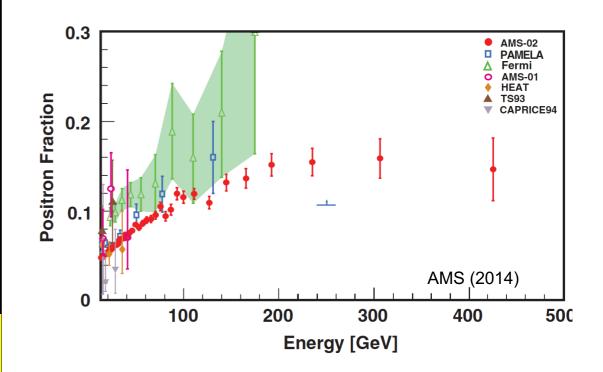


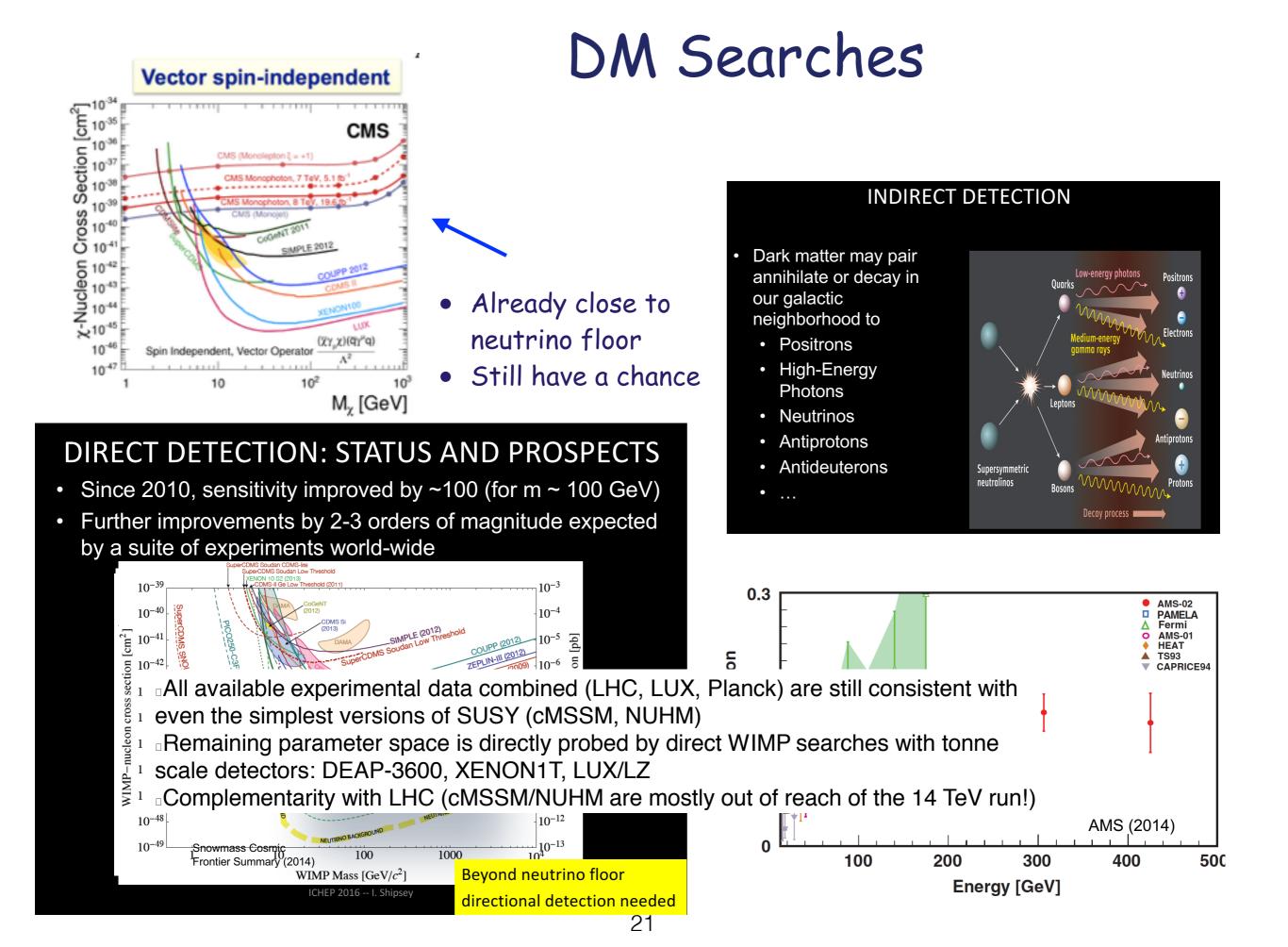
neutralinos

Rosons

Decay process 💼

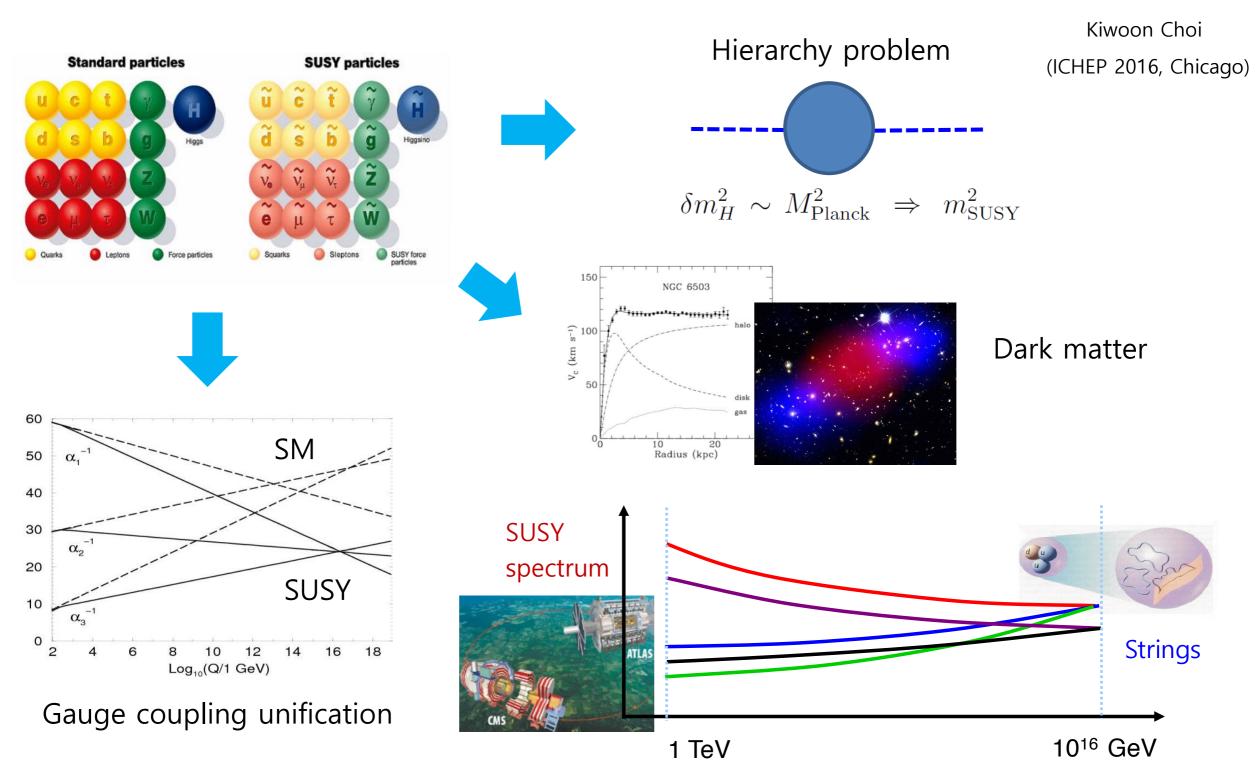
Protons





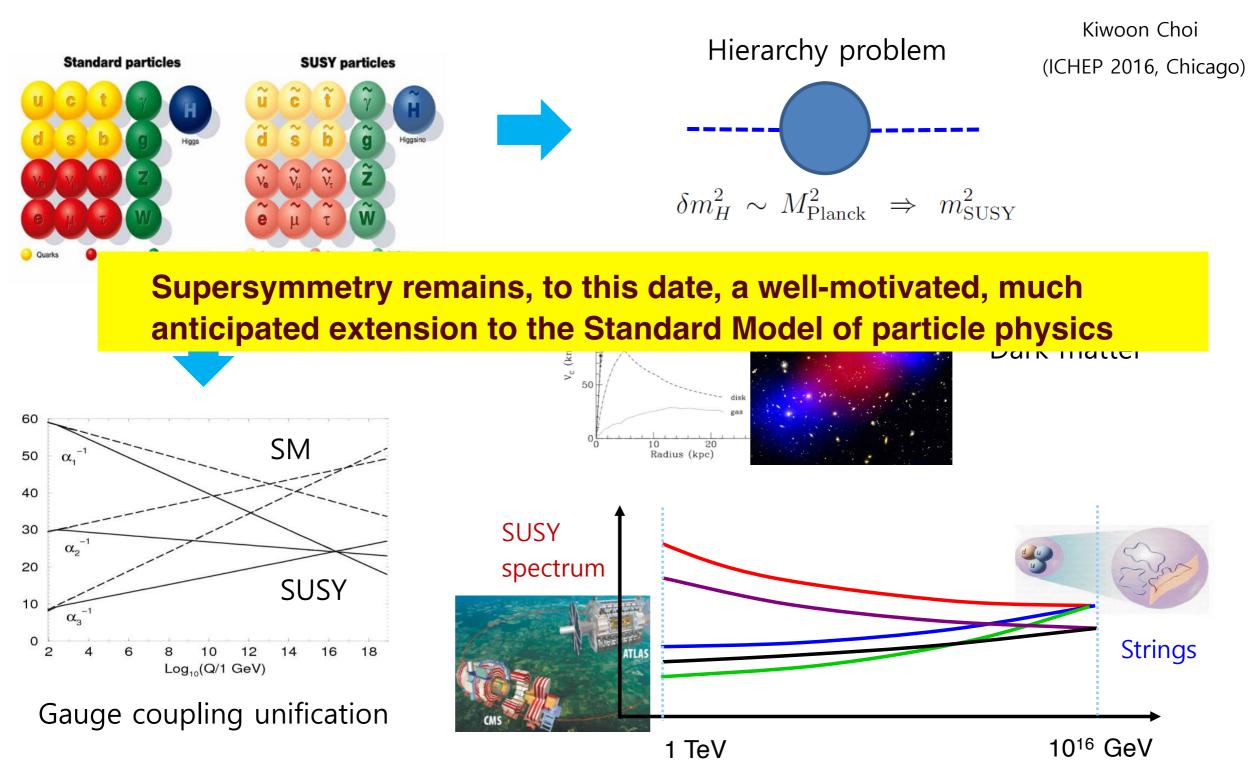
Supersymmetry - Target # 3

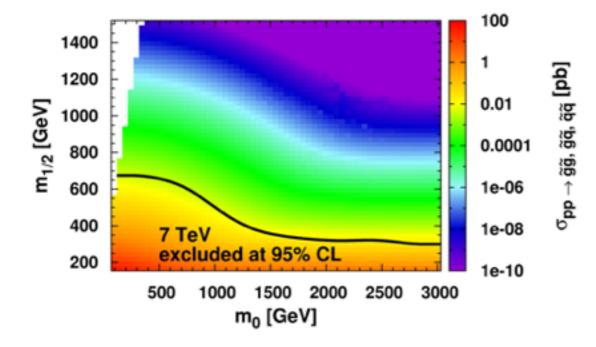
SUSY has been the prime candidate for BSM physics near the TeV scale.

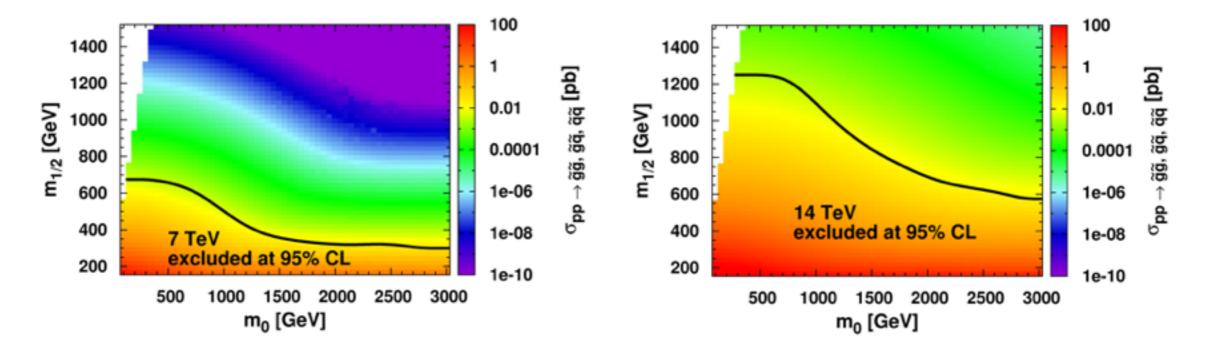


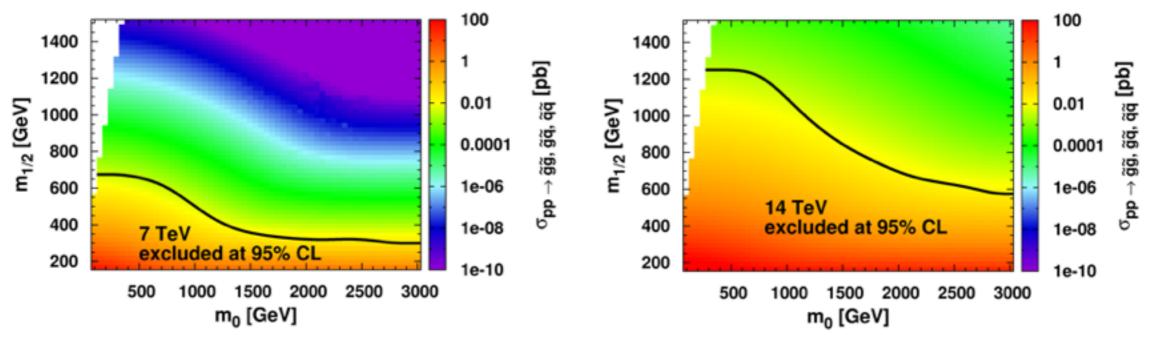
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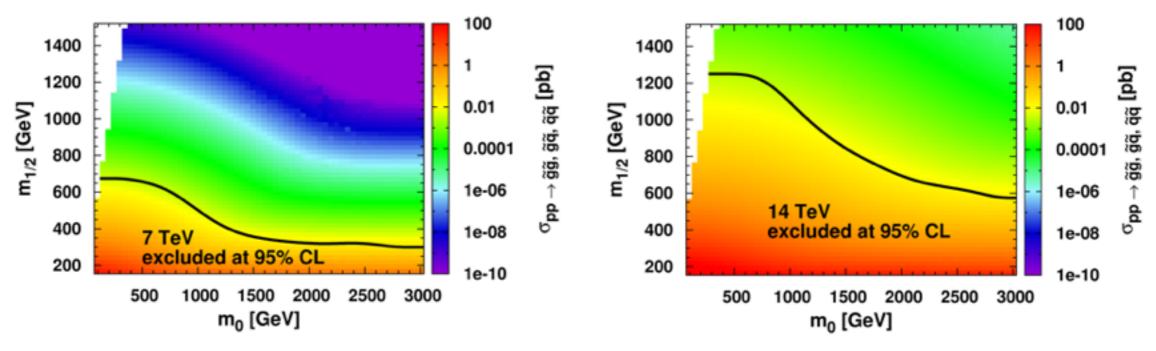




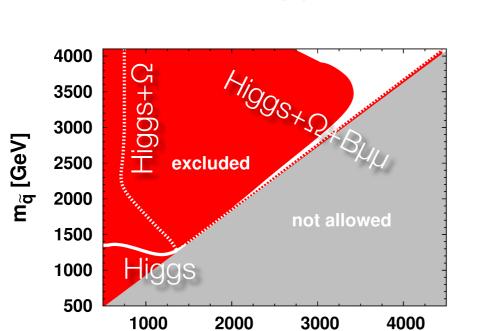


Masses of superpartners

Universal scenario



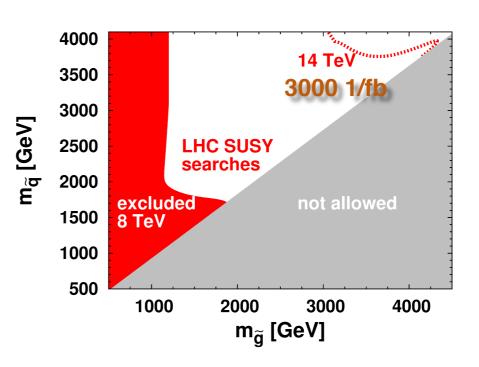
Masses of superpartners



 $m_{\tilde{a}}$ [GeV]

CMSSM





NMSSM

LHC Run2

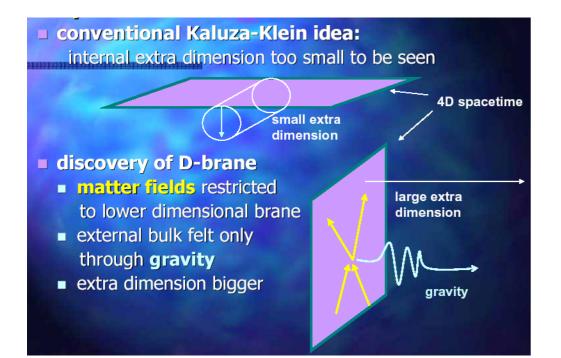
	TLAS SUSY Se atus: August 2016 Model		s* - 95 / _{Jets}	12			VI = 7, 8 TeV	ATLAS Preliminary $\sqrt{s} = 7, 8, 13 \text{ TeV}$ Reference
Inclusive Searches	$\begin{array}{l} MSUGRA/CMSSM \\ \bar{q}\bar{q}, \bar{q} \rightarrow q\bar{\chi}_{1}^{0} \\ \bar{q}\bar{q}, \bar{q} \rightarrow q\bar{\chi}_{1}^{0} (compressed) \\ \bar{g}\bar{g}, \bar{g} \rightarrow q\bar{q}\bar{\chi}_{1}^{0} (compressed) \\ \bar{g}\bar{g}, \bar{g} \rightarrow q\bar{q}\bar{\chi}_{1}^{0} \rightarrow q\bar{q}W^{\pm}\bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}, \bar{g} \rightarrow q\bar{q}\bar{\chi}_{1}^{\pm} \rightarrow q\bar{q}W^{\pm}\bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}, \bar{g} \rightarrow q\bar{q}WZ\bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}\bar{g}, \bar{g} \rightarrow q\bar{q}\bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}\bar{\chi}, \bar{g} \rightarrow q\bar{q}\bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}\bar{\chi}, \bar{g} \rightarrow q\bar{q}\bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}\bar{\chi}, \bar{g} \rightarrow q\bar{\chi}_{1}^{0} \\ \bar{g}\bar{\chi}, \bar{g} \rightarrow q\bar{\chi}, \bar{\chi}, \bar{\chi},$	$\begin{array}{c} 0\text{-}3 \ e, \mu/1\text{-}2 \ \tau \\ 0 \\ \text{mono-jet} \\ 0 \\ 0 \\ 3 \ e, \mu \\ 2 \ e, \mu \ (\text{SS}) \\ 1\text{-}2 \ \tau + 0\text{-}1 \\ 2 \ \gamma \\ \gamma \\ \gamma \\ 2 \ e, \mu \ (\text{Z}) \\ 0 \\ \end{array}$	2-10 jets/3 <i>b</i> 2-6 jets 1-3 jets 2-6 jets 2-6 jets 4 jets 0-3 jets	•	20.3 13.3 13.3 13.3 13.2 13.2 3.2 3.2 20.3 13.3 20.3 20.3	.2 1.3 608 GeV	$\begin{array}{c c} \textbf{1.85 TeV} & m(\hat{q}) = m(\hat{q}) \\ \textbf{1.85 TeV} & m(\hat{x}_1^0) < 200 \text{ GeV}, \ m(1^d \ \text{gen}, \hat{q}) = m(2^{ad} \ \text{gen}, \hat{q}) \\ & m(\hat{q}) - m(\hat{x}_1^0) < 5 \ \text{GeV} \\ \textbf{1.86 TeV} & m(\hat{x}_1^0) < 00 \ \text{GeV}, \ m(\hat{x}^{-1}) = 0.5(m(\hat{x}_1^0) + m(\hat{g})) \\ \textbf{1.7 TeV} & m(\hat{x}_1^0) < 400 \ \text{GeV} \\ \textbf{1.6 TeV} & m(\hat{x}_1^0) < 400 \ \text{GeV} \\ \textbf{1.6 TeV} & m(\hat{x}_1^0) < 500 \ \text{GeV} \\ \textbf{2.0 TeV} \\ \textbf{1.65 TeV} & cr(\text{NLSP}) < 0.1 \ \text{mm} \\ \textbf{37 TeV} & m(\hat{x}_1^0) < 500 \ \text{GeV}, \ cr(\text{NLSP}) < 0.1 \ \text{mm}, \ \mu < 0 \\ \textbf{1.8 TeV} & m(\hat{x}_1^0) > 680 \ \text{GeV}, \ cr(\text{NLSP}) < 0.1 \ \text{mm}, \ \mu > 0 \\ \textbf{m}(\hat{x}_1^0) > 680 \ \text{GeV} \\ \textbf{m}(\hat{x}_1^0) > 1.8 \times 10^{-4} \ \text{eV}, \ m(\hat{y}) = m(\hat{y}) = 1.5 \ \text{TeV} \\ \end{array}$	1507.05525 ATLAS-CONF-2016-078 1604.07773 ATLAS-CONF-2016-078 ATLAS-CONF-2016-078 ATLAS-CONF-2016-037 ATLAS-CONF-2016-037 1607.05979 1606.09150 1507.05493 ATLAS-CONF-2016-066 1503.03290 1502.01518
3 rd gen. § med.	$\begin{array}{c} \bar{g}\bar{g}, \bar{g} \rightarrow b \bar{b} \bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}, \bar{g} \rightarrow c \bar{c} \bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}, \bar{g} \rightarrow c \bar{c} \bar{\chi}_{1}^{0} \end{array}$	0 0-1 e,µ 0-1 e,µ	3 b 3 b 3 b	Yes Yes Yes	14.8 14.8 20.1	1.3	1.89 TeV m(k ^b ₁)=0 GeV 1.89 TeV m(k ^b ₁)=0 GeV 37 TeV m(k ^c ₁)=0 GeV	ATLAS-CONF-2016-052 ATLAS-CONF-2016-052 1407.0600
3 rd gen. squarks direct production	$ \begin{array}{l} \bar{b}_{1}\bar{b}_{1}, \ \bar{b}_{1} \rightarrow b\bar{\chi}_{1}^{0} \\ \bar{b}_{1}\bar{b}_{1}, \ \bar{b}_{1} \rightarrow b\bar{\chi}_{1}^{0} \\ \bar{t}_{1}\bar{t}_{1}, \ \bar{t}_{1} \rightarrow b\bar{\chi}_{1}^{0} \\ \bar{t}_{1}\bar{t}_{1}, \ \bar{t}_{1} \rightarrow b\bar{\chi}_{1}^{0} \\ \bar{t}_{1}\bar{t}_{1}, \ \bar{t}_{1} \rightarrow c\bar{\chi}_{1}^{0} \\ \bar{t}_{1}\bar{t}_{1}, \ \bar{t}_{1} \rightarrow c\bar{\chi}_{1}^{0} \\ \bar{t}_{1}\bar{t}_{1}, \ \bar{t}_{1} \rightarrow c\bar{\chi}_{1}^{0} \\ \bar{t}_{2}\bar{t}_{2}, \ \bar{t}_{2} \rightarrow \bar{t}_{1} + Z \\ \bar{t}_{2}\bar{t}_{2}, \ \bar{t}_{2} \rightarrow \bar{t}_{1} + h \end{array} $	$\begin{array}{c} 0 \\ 2 \ e, \mu (SS) \\ 0 - 2 \ e, \mu \\ 0 - 2 \ e, \mu \\ 0 \\ 2 \ e, \mu (Z) \\ 3 \ e, \mu (Z) \\ 1 \ e, \mu \end{array}$	2 b 1 b 1-2 b 0-2 jets/1-2 b mono-jet 1 b 1 b 6 jets + 2 b	Yes 4 Yes Yes Yes	3.2 13.2 .7/13.3 .7/13.3 3.2 20.3 13.3 20.3	840 GeV 325-685 GeV -170 GeV 90-198 GeV 90-323 GeV 150-600 GeV 290-700 GeV 320-620 GeV	$\begin{array}{l} m(\bar{k}_{1}^{0}) < 100 {\rm GeV} \\ m(\bar{k}_{1}^{0}) < 150 {\rm GeV}, m(\bar{k}_{1}^{0}) = m(\bar{k}_{1}^{0}) + 100 {\rm GeV} \\ m(\bar{k}_{1}^{1}) = 2m(\bar{k}_{1}^{0}), m(\bar{k}_{1}^{0}) - 55 {\rm GeV} \\ m(\bar{k}_{1}^{0}) = 1 {\rm GeV} \\ m(\bar{k}_{1}^{0}) = 1 {\rm GeV} \\ m(\bar{k}_{1}^{0}) > 150 {\rm GeV} \\ m(\bar{k}_{1}^{0}) < 150 {\rm GeV} \\ m(\bar{k}_{1}^{0}) < 300 {\rm GeV} \\ m(\bar{k}_{1}^{0}) = 0 {\rm GeV} \end{array}$	1606.08772 ATLAS-CONF-2016-037 1209.2102, ATLAS-CONF-2016-077 1506.08616, ATLAS-CONF-2016-077 1604.07773 1403.5222 ATLAS-CONF-2016-038 1506.08616
EW direct	$ \begin{array}{l} \tilde{\ell}_{LR} \tilde{\ell}_{LR}, \tilde{\ell} \rightarrow \tilde{\ell} \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{1}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\ell} \nu (\ell \bar{\nu}) \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\tau} \nu (r \bar{\nu}) \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{L} \nu \tilde{\ell}_{L} \ell (\bar{\nu} \nu), \ell \bar{\nu} \tilde{\ell}_{L} \ell (\bar{\nu} \nu) \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} L \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} L \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} L \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{2}^{+} \tilde{\chi}_{3}^{0}, \tilde{\chi}_{2,3}^{0} \rightarrow \tilde{\ell}_{R} \ell \\ GGM (wino NLSP) weak prod. \\ GGM (bino NLSP) weak prod. \end{array} $	2 e,μ 2 e,μ 2 τ 3 e,μ 2-3 e,μ τ/γγ e,μ,γ 4 e,μ 1 e,μ + γ 2 γ	0 0 0-2 jets 0-2 <i>b</i> 0	Yes Yes Yes Yes Yes Yes Yes Yes	20.3 20.3 20.3 20.3 20.3 20.3 20.3 20.3	90-335 GeV 1 140-475 GeV 355 GeV 1 140-475 GeV 1 140-475 GeV 715 GeV 1 140-475 GeV 715 GeV 715 GeV 715 GeV 1 15-370 GeV 590 GeV	$\begin{array}{c} m(\tilde{k}_1^0){=}0~{\rm GeV} \\ m(\tilde{k}_1^0){=}0~{\rm GeV}, m(\tilde{\ell},\tilde{\nu}){=}0.5(m(\tilde{k}_1^0){+}m(\tilde{k}_1^0)) \\ m(\tilde{k}_1^0){=}0~{\rm GeV}, m(\tilde{\ell},\tilde{\nu}){=}0.5(m(\tilde{k}_1^0){+}m(\tilde{k}_1^0)) \\ m(\tilde{k}_1^0){=}m(\tilde{k}_2^0), m(\tilde{\ell},0){=}0, m(\tilde{\ell},\tilde{\tau}){+}m(\tilde{k}_1^0)) \\ m(\tilde{k}_1^0){=}m(\tilde{k}_2^0), m(\tilde{k}_1^0){=}0, \tilde{\ell}~{\rm decoupled} \\ m(\tilde{k}_1^0){=}m(\tilde{k}_2^0), m(\tilde{k}_1^0){=}0, \tilde{\ell}~{\rm decoupled} \\ m(\tilde{k}_2^0){=}m(\tilde{k}_2^0), m(\tilde{k}_1^0){=}0, m(\tilde{k},\tilde{\nu}){=}0.5(m(\tilde{k}_2^0){+}m(\tilde{k}_2^0)) \\ cr{<}1~{\rm mm} \\ cr{<}1~{\rm mm} \end{array}$	1403.5294 1403.5294 1407.0350 1402.7029 1403.5294, 1402.7029 1501.07110 1405.5086 1507.05493 1507.05493
	Direct $\hat{x}_1^+ \hat{x}_1^-$ prod., long-lived \hat{x}_1^+ Direct $\hat{x}_1^+ \hat{x}_1^-$ prod., long-lived \hat{x}_1^+	Disapp. trk	t jet	Yes Yes	20.3 18.4	270 GeV 495 GeV	m(ξ [±] ₁)-m(ξ ⁰ ₁)−160 MeV, τ(ξ [±] ₁)=0.2 ns m(ξ [±] ₁)-m(ξ ⁰ ₁)−160 MeV, τ(ξ [±] ₁)<15 ns	1310.3675 1506.05332
Long-lived particles	Stable, stopped \tilde{g} R-hadron Stable \tilde{g} R-hadron Metastable \tilde{g} R-hadron GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{\epsilon}, \tilde{\mu}) + r(GMSB, \tilde{\chi}_1^0 \rightarrow \gamma G, \log-ived \tilde{\chi}_1^0)$ $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow eev(e\mu v/\mu\mu v)$ GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow ZG$	0 trk dE/dx trk		Yes - - Yes -	27.9 3.2 19.1 20.3 20.3 20.3	850 GeV 537 GeV 440 GeV 1.0 TeV 1.0 TeV	$\begin{array}{l} \mbox{m}(k_1^{a}) = 100 \ {\rm GeV}, \ 10 \ \mu s < r(\tilde{g}) < 1000 \ s \\ \mbox{m}(\tilde{k}_1^{b}) = 100 \ {\rm GeV}, \ 10 \ \mu s < r(\tilde{g}) < 1000 \ s \\ \mbox{m}(\tilde{k}_1^{b}) = 100 \ {\rm GeV}, \ r > 10 \ {\rm ns} \\ \ 10 < tang < 50 \\ \ 1 < r(\tilde{k}_1^{b}) < 3 \ {\rm ns}, \ {\rm SPS8 \ model} \\ \ 7 < cr(\tilde{k}_1^{b}) < 740 \ {\rm mm}, \ m(\tilde{g}) = 1.3 \ {\rm TeV} \\ \ 6 < cr(\tilde{k}_1^{b}) < 480 \ {\rm mm}, \ m(\tilde{g}) = 1.1 \ {\rm TeV} \end{array}$	1310.6584 1606.05129 1604.04520 1411.6795 1409.5542 1504.05162 1504.05162
RPV Long-lived	Stable \tilde{g} R-hadron Metastable \tilde{g} R-hadron GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{\epsilon}, \tilde{\mu}) \star \tau(GMSB, \tilde{\chi}_1^0 \rightarrow \gamma G, \log - lived \tilde{\chi}_1^0)$	0 trk dE/dx trk (e, μ) 1-2 μ 2γ displ. $ee/e\mu/\mu$ displ. vtx + je $e\mu, er, \mu\tau$ $2 e, \mu$ (SS) $\mu\mu\nu$ 4 e, μ $3 e, \mu + \tau$ 0 4	- 	Yes - Ves - Yes Yes S S - Yes S S - Yes	3.2 3.2 19.1 20.3 20.3	537 GeV 440 GeV 1.0 TeV 1.0 TeV 1.0 TeV 1.0 TeV 1.14 Te 450 GeV 1.08 TeV	$\begin{array}{c} m(\tilde{k}_{1}^{b}){=}100 \; \mathrm{GeV}, \; 10 \; \mu \mathrm{scr}(\tilde{g}){<}1000 \; \mathrm{s} \\ \texttt{1.58 TeV} \\ \texttt{1.57 TeV} \\ \texttt{m}(\tilde{k}_{1}^{b}){=}100 \; \mathrm{GeV}, \; \texttt{r}{>}10 \; \mathrm{ns} \\ 10{<}\mathrm{tan}\beta{<}50 \\ 1{<}\mathrm{r}(\tilde{k}_{1}^{b}){<}3 \; \mathrm{ns}, \; \mathrm{SPS8 \; model} \\ 7{<}\mathrm{cr}(\tilde{k}_{1}^{b}){<}740 \; \mathrm{mm}, \; \mathrm{m}(\tilde{g}){=}1.3 \; \mathrm{TeV} \\ 6{<}\mathrm{cr}(\tilde{k}_{1}^{b}){<}480 \; \mathrm{mm}, \; \mathrm{m}(\tilde{g}){=}1.1 \; \mathrm{TeV} \\ \texttt{1.9 TeV} \; \mathcal{L}_{311}{=}0.11, \; \mathcal{L}_{132/133/239}{=}0.07 \\ \texttt{M} \\ \texttt{M} \\ \texttt{M} \\ \texttt{M} \\ \texttt{M} \\ \texttt{N} \\ \texttt{N} \\ \texttt{N} \\ \texttt{M} \\ \texttt{N} \\ $	1310.6584 1606.05129 1604.04520 1411.6795 1409.5542 1504.05162

LHC Run2

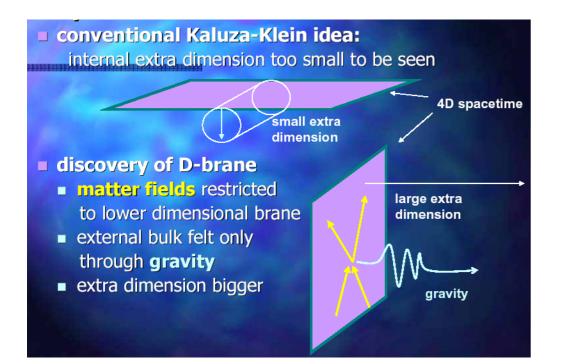
5395	atus: August 2016 Model	e, μ, τ, γ	Jets	$E_{\mathrm{T}}^{\mathrm{miss}}$	∫£ dt[fb	Mass limit $\sqrt{s} = 7$	8 TeV 🗸 = 13 TeV	$\sqrt{s} = 7, 8, 13 \text{ TeV}$ Reference
Inclusive Searches	$\begin{array}{l} MSUGRA/CMSSM\\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0}\\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0} \ (\text{compressed})\\ \tilde{g}\tilde{e}, \tilde{e} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{0} \ (\text{compressed})\\ \tilde{g}\tilde{e}, \tilde{e} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{0} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0}\\ \tilde{g}\tilde{e}, \tilde{e} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0}\\ \tilde{g}\tilde{e}, \tilde{e} \rightarrow qqWZ\tilde{\chi}_{1}^{0} \ (\tilde{g}\tilde{e}, \tilde{e} \rightarrow qqWZ\tilde{\chi}_{1}^{0})\\ \tilde{g}\tilde{e}, \tilde{e} \rightarrow qqWZ\tilde{\chi}_{1}^{0} \ (\tilde{g}\tilde{e}, \tilde{e} \rightarrow qqWZ\tilde{\chi}_{1}^{0})\\ \tilde{g}\tilde{e}, \tilde{e} \rightarrow qqWZ\tilde{\chi}_{1}^{0} \ (\tilde{g}\tilde{e}, \tilde{e} \rightarrow qqWZ\tilde{\chi}_{1}^{0})\\ \tilde{g}GMSB(\ell NLSP)\\ GGM(binoNLSP)\\ GGM(higgsino-binoNLSP)\\ GGM(higgsinobinoNLSP)\\ GGM(higgsinoNLSP)\\ GravitinoLSP \end{array}$	$\begin{array}{c} 0.3 \ e, \mu/1-2 \ \tau \\ 0 \\ mono-jet \\ 0 \\ 3 \ e, \mu \\ 2 \ e, \mu \ (SS) \\ 1-2 \ \tau + 0.1 \ \ell \\ 2 \ \gamma \\ \gamma \\ 2 \ e, \mu \ (Z) \\ 0 \end{array}$	2-6 jets 1-3 jets 2-6 jets 2-6 jets 4 jets 0-3 jets	Yes Yes Yes Yes	20.3 13.3 13.3 13.3 13.2 13.2 13.2 3.2 20.3 13.3 20.3 20.3 20.3	1.35 TeV 608 GeV 1.85 Te 1.83 Te 1.7 TeV 1.6 TeV	m(R ^b ₁) <500 GeV	1507.05525 ATLAS-CONF-2016-078 1604.07773 ATLAS-CONF-2016-078 ATLAS-CONF-2016-078 ATLAS-CONF-2016-037 ATLAS-CONF-2016-037 1607.05979 1806.09150 1507.05493 3-CONF-2016-066 1503.03290 1502.01518
§ med.	$\begin{array}{c} \bar{g}\bar{g}, \bar{g} \rightarrow b \bar{b} \bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}, \bar{g} \rightarrow b \bar{t} \bar{\chi}_{1}^{0} \\ \bar{g}\bar{g}, \bar{g} \rightarrow b \bar{t} \bar{\chi}_{1}^{+} \end{array}$	0 0-1 e, µ 0-1 e, µ	3 b 3 b 3 b	Yes Yes Yes	14.8 14.8 20.1	alists:	IL SEE IL	ATLAS-CONF-2016-052 ATLAS-CONF-2016-052 1407.0600
direct production	$ \begin{split} \vec{b}_{1}\vec{b}_{1}, \vec{b}_{1} \rightarrow b\vec{k}_{1}^{0} \\ \vec{b}_{1}\vec{b}_{1}, \vec{b}_{1} \rightarrow t\vec{k}_{1}^{-1} \\ \vec{t}_{1}\vec{t}_{1}, \vec{t}_{1} \rightarrow b\vec{k}_{1}^{0} \\ \vec{t}_{1}\vec{t}_{1}, \vec{t}_{1} \rightarrow b\vec{k}_{1}^{0} \\ \vec{t}_{1}\vec{t}_{1}, \vec{t}_{1} \rightarrow c\vec{k}_{1}^{0} \\ \vec{t}_{1}\vec{t}_{1}, \vec{t}_{1} \rightarrow c\vec{k}_{1}^{0} \\ \vec{t}_{2}\vec{t}_{2}, \vec{t}_{2} \rightarrow \vec{t}_{1} + Z \\ \vec{t}_{2}\vec{t}_{2}, \vec{t}_{2} \rightarrow \vec{t}_{1} + h \end{split} $	$\begin{array}{c} 0 \\ 2 e, \mu (\text{SS}) \\ 0 - 2 e, \mu \\ 0 - 2 e, \mu \\ 0 \\ 2 e, \mu (Z) \\ 3 e, \mu (Z) \\ 1 e, \mu \end{array}$	2 b 1 b 1-2 b 0-2 jets/1-2 mono-jet 1 b 1 b 6 jets + 2 b	Yes Yes Yes Yes Yes Yes Yes Yes	3.2 13.2 1.7/13.3 1.7/13.3 3.2 20.3 13.5	experimentane't y	$\begin{array}{c} & & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ \\$	1606.08772 ATLAS-CONF-2016-037 1209.2102, ATLAS-CONF-2016-077 1506.08616, ATLAS-CONF-2016-077 1604.07773 1403.5222 ATLAS-CONF-2016-038 1506.08616
direct	$ \begin{array}{l} \tilde{\ell}_{L,R} \tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}, \tilde{\chi}_{1}^{\dagger} \rightarrow \tilde{\ell} \nu (\ell \bar{\nu}) \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{\dagger} \rightarrow \tilde{\ell} \nu (\ell \bar{\nu}) \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{L} \nu \tilde{\ell}_{L} \ell (\bar{\nu} \nu), \ell \bar{\nu} \tilde{\ell}_{L} \ell (\bar{\nu} \nu) \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{L}^{0} \ell \chi_{1}^{0} \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{L}^{0} h \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{1}^{\dagger} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} h \tilde{\chi}_{1}^{0} \\ \tilde{\chi}_{2}^{\dagger} \tilde{\chi}_{3}^{0}, \tilde{\chi}_{2,3}^{0} \rightarrow \tilde{\ell}_{R} \ell \\ GGM (wino NLSP) weak prod. \\ GGM (bino NLSP) weak prod. \end{array} $	2 ε,μ 2 ε,μ 2 τ 3 2-3 τ/γγ ε,μ, 4 ε,μ 1 ε,μ + , 2 γ	jue SU	sti	0n (is 20.3	2.0 1.85 TeV 1.37 TeV ** scale 900 GeV 865 GeV ** scale 900 GeV 900 GeV 900 GeV 900 GeV 900 GeV 90 GeV 115-370 GeV 90 GeV	$\begin{array}{l} m(\tilde{k}_{1}^{0}) = 0 \ \text{GeV} \\ m(\tilde{k}_{1}^{0}) = 0 \ \text{GeV}, m(\tilde{\ell},\tilde{r}) = 0.5(m(\tilde{k}_{1}^{0}) + m(\tilde{k}_{1}^{0})) \\ m(\tilde{k}_{1}^{0}) = 0 \ \text{GeV}, m(\tilde{\ell},\tilde{r}) = 0.5(m(\tilde{k}_{1}^{0}) + m(\tilde{k}_{1}^{0})) \\ = m(\tilde{k}_{2}^{0}), m(\tilde{r}_{1}^{0}) = 0, m(\tilde{\ell},\tilde{r}) = 0.5(m(\tilde{k}_{1}^{0}) + m(\tilde{k}_{1}^{0})) \\ m(\tilde{k}_{1}^{0}) = m(\tilde{k}_{2}^{0}), m(\tilde{k}_{1}^{0}) = 0, \tilde{\ell} \ \text{decoupled} \\ m(\tilde{k}_{1}^{0}) = m(\tilde{k}_{2}^{0}), m(\tilde{k}_{1}^{0}) = 0, \tilde{\ell} \ \text{decoupled} \\ = m(\tilde{k}_{2}^{0}), m(\tilde{k}_{1}^{0}) = 0, m(\tilde{\ell},\tilde{r}) = 0.5(m(\tilde{k}_{2}^{0}) + m(\tilde{k}_{1}^{0})) \\ c\tau < 1 \ \text{mm} \\ c\tau < 1 \ \text{mm} \end{array}$	1403.5294 1403.5294 1407.0350 1402.7029 1403.5294, 1402.7029 1501.07110 1405.5096 1507.05493 1507.05493
particles	Direct $\hat{x}_{1}^{\dagger}\hat{x}_{1}^{\dagger}$ prod., long-lived \hat{x} Direct $\hat{x}_{1}^{\dagger}\hat{x}_{1}^{\dagger}$ prod., long-lived \hat{x} Stable, stopped \hat{g} R-hadron Stable \hat{g} R-hadron Metastable \hat{g} R-hadron GMSB, stable $\hat{\tau}, \hat{x}_{1}^{0} \rightarrow \hat{\tau}(\hat{v}, \hat{\mu}) + r$ GMSB, $\hat{x}_{1}^{0} \rightarrow \gamma \hat{G}$, long-lived \hat{x}_{1}^{0} $\hat{g}\hat{g}, \hat{x}_{1}^{0} \rightarrow eev/e\muv/\mu\muv$ GGM $\hat{g}\hat{g}, \hat{x}_{1}^{0} \rightarrow Z\hat{G}$	0 trk dE/dx trk	1-5 jøts - - - - -	Yes Yes Yes Yes	20.3 18.4 27.9 3.2 19.1 20.3 20.3 20.3	270 GeV 495 GeV 850 GeV 1.58 TeV 1.57 TeV 537 GeV 440 GeV 1.0 TeV 1.0 TeV	$\begin{split} & m(\tilde{k}_{1}^{4}) - m(\tilde{k}_{1}^{0}) - 160 \text{ MeV}, \ \tau(\tilde{k}_{1}^{*}) = 0.2 \text{ ns} \\ & m(\tilde{k}_{1}^{4}) - m(\tilde{k}_{1}^{0}) - 160 \text{ MeV}, \ \tau(\tilde{k}_{1}^{*}) < 15 \text{ ns} \\ & m(\tilde{k}_{1}^{0}) = 100 \text{ GeV}, \ 10 \ \mu \text{s} < \tau(\tilde{g}) < 1000 \text{ s} \\ & m(\tilde{k}_{1}^{0}) = 100 \text{ GeV}, \ \tau > 10 \text{ ns} \\ & 10 < \tan \theta < 50 \\ & 1 < \tau(\tilde{k}_{1}^{0}) < 3 \text{ ns}, \text{ SPS8 model} \\ & 7 < c\tau(\tilde{k}_{1}^{0}) < 740 \text{ mm}, \ m(\tilde{g}) = 1.3 \text{ TeV} \\ & 6 < c\tau(\tilde{k}_{1}^{0}) < 480 \text{ mm}, \ m(\tilde{g}) = 1.1 \text{ TeV} \end{split}$	1310.3675 1506.05332 1310.6584 1606.05129 1804.04520 1411.6795 1409.5542 1504.05162 1504.05162
RPV	LFV $pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e\mu/e\tau/\mu\tau$ Bilinear RPV CMSSM $\tilde{X}_{1}^{\dagger}\tilde{X}_{1}^{-}, \tilde{X}_{1}^{\dagger} \rightarrow W\tilde{X}_{1}^{0}, \tilde{X}_{1}^{0} \rightarrow eev, e\mu v, \mu$ $\tilde{x}_{1}^{\dagger}\tilde{X}_{1}^{-}, \tilde{X}_{1}^{\dagger} \rightarrow W\tilde{X}_{1}^{0}, \tilde{X}_{1}^{0} \rightarrow \tau\tau v_{e}, e\tau v$ $\bar{g}\bar{g}, \bar{g} \rightarrow qq\bar{g}$ $\tilde{g}\bar{g}, \bar{g} \rightarrow qq\bar{g}$ $\tilde{g}\bar{g}, \bar{g} \rightarrow \bar{q}q\bar{g}$ $\tilde{g}\bar{g}, \bar{g} \rightarrow \bar{t}_{1}t, \bar{t}_{1} \rightarrow bs$ $\bar{t}_{1}\bar{t}_{1}, \bar{t}_{1} \rightarrow bs$ $\bar{t}_{1}\bar{t}_{1}, \bar{t}_{1} \rightarrow b\ell$	$2 e, \mu$ (SS) $\mu\nu$ $4 e, \mu$ $3 e, \mu + \tau$ 0 4 0 4 $2 e, \mu$ (SS)	- 0-3 b -5 large-R je -5 large-R je 0-3 b 2 jets + 2 b 2 b	ets - Yes	3.2 20.3 13.3 20.3 14.8 14.8 13.2 15.4 20.3	1.9 T 1.45 TeV 1.45 TeV 1.14 TeV 450 GeV 1.08 TeV 1.55 TeV 1.3 TeV 410 GeV 0.4-1.0 TeV	eV $\lambda'_{311}=0.11, \lambda_{132(13)(23)}=0.07$ $m(\tilde{g})=m(\tilde{g}), c\tau_{LSF}<1 mm$ $m(\tilde{k}^0_1)>400GeV, \lambda_{132}\neq0 (k = 1, 2)$ $m(\tilde{k}^0_1)>0.2\times m(\tilde{k}^1_1), \lambda_{133}\neq0$ BR(r)=BR(b)=BR(c)=0% $m(\tilde{k}^0_1)=800 GeV$ $m(\tilde{k}^0_1)=600 GeV$ $BR(\tilde{r}_1 \rightarrow be/\mu)>20\%$	1607.06079 1404.2500 ATLAS-CONF-2016-075 1405.5086 ATLAS-CONF-2016-057 ATLAS-CONF-2016-057 ATLAS-CONF-2016-037 ATLAS-CONF-2016-022, ATLAS-CONF-2016-00 ATLAS-CONF-2015-015
ther	Scalar charm, $\tilde{c} \rightarrow c \tilde{\ell}_1^0$	0	2 c	Yes	20.3	510 GeV	m(ℓ ² 1)<200 GeV	1501.01325

Extra Dimensions/ Exotics

Extra Dimensions/ Exotics

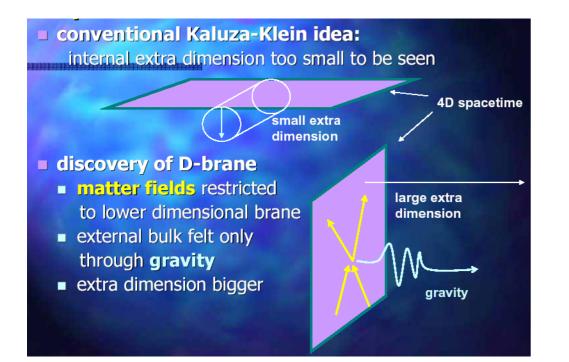


Extra Dimensions/ Exotics



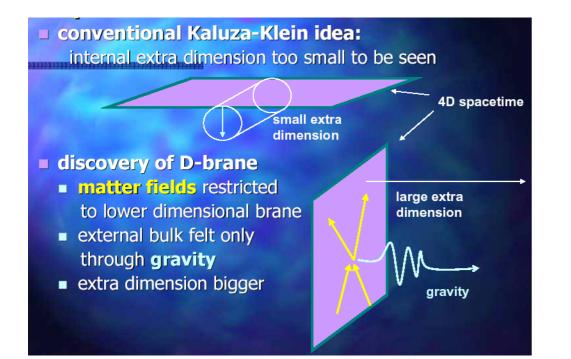
Q: Do we really live on a brane?

Extra Dimensions/ Exotics



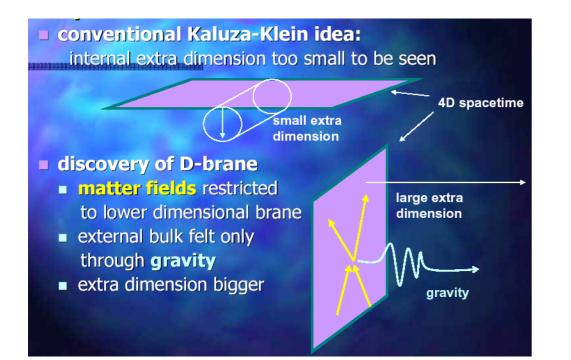
Q: Do we really live on a brane?A: We have to check it

Extra Dimensions/ Exotics



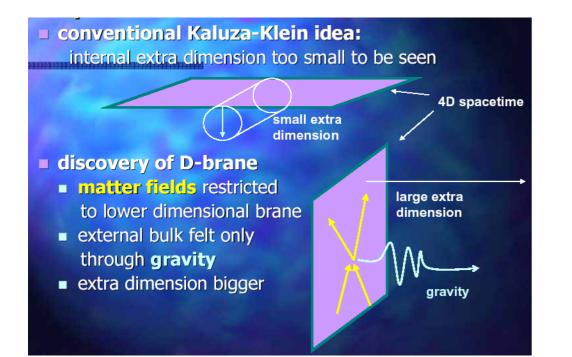
Q: Do we really live on a brane?A: We have to check itQ: Do we have good reasons to believe in it?

Extra Dimensions/ Exotics



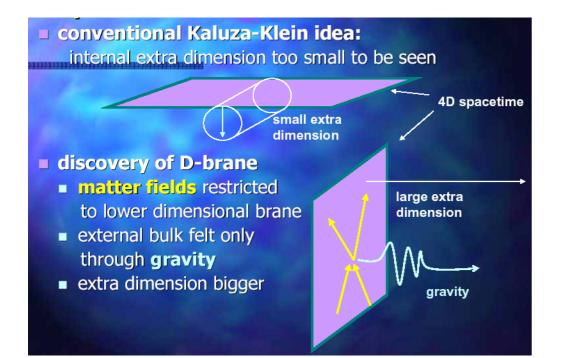
Q: Do we really live on a brane?A: We have to check itQ: Do we have good reasons to believe in it?A: No, but it is appealing

Extra Dimensions/ Exotics



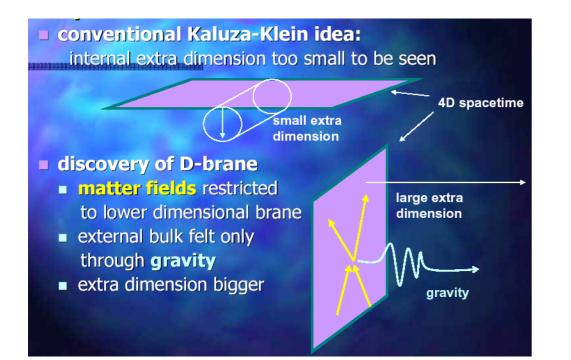
Q: Do we really live on a brane?A: We have to check itQ: Do we have good reasons to believe in it?A: No, but it is appealingQ: Why D>4?

Extra Dimensions/ Exotics

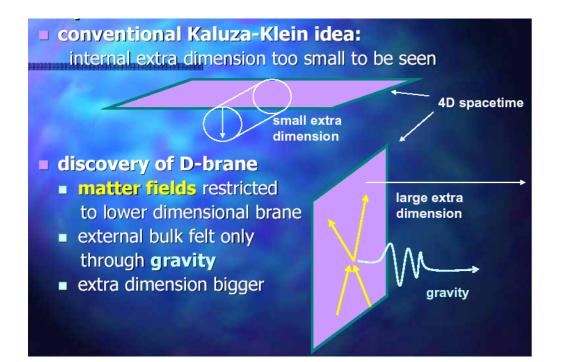


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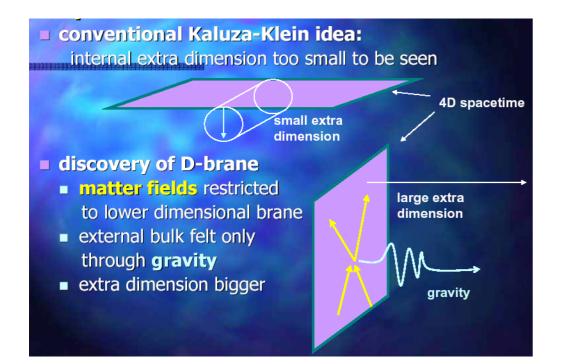
Extra Dimensions/ Exotics



Extra Dimensions/ Exotics

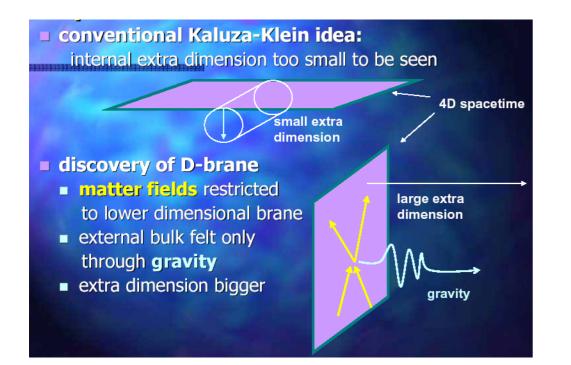


Extra Dimensions/ Exotics



Experiment

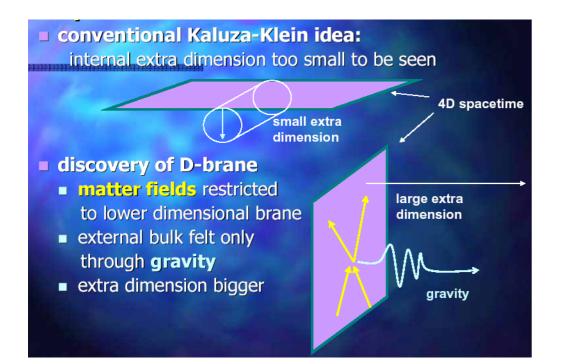
Extra Dimensions/ Exotics



Experiment

Search for Z' (Di-muon events)

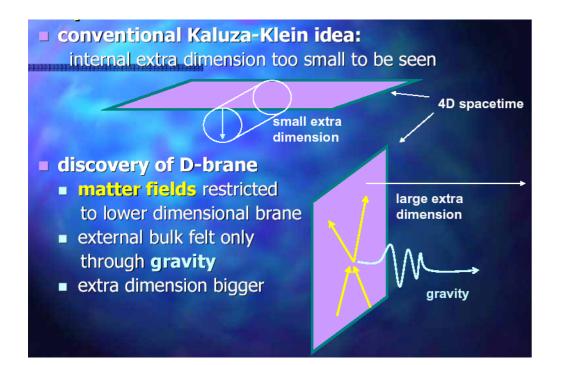
Extra Dimensions/ Exotics



Experiment

- Search for Z' (Di-muon events)
- Search for W' (single muon/ jets)

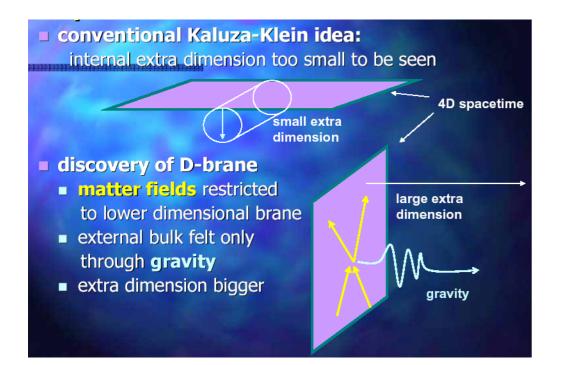
Extra Dimensions/ Exotics



Experiment

- Search for Z' (Di-muon events)
- Search for W' (single muon/ jets)
- Search for resonance decaying to t-tbar

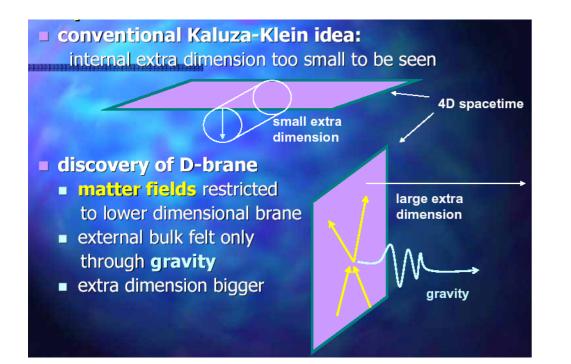
Extra Dimensions/ Exotics



Experiment

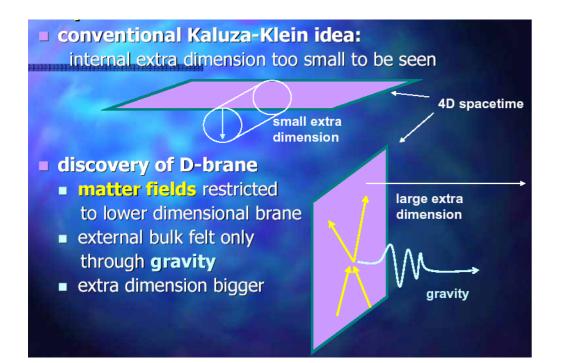
- Search for Z' (Di-muon events)
- Search for W' (single muon/ jets)
- Search for resonance decaying to t-tbar
- Search for diboson resonances

Extra Dimensions/ Exotics



- Experiment
- Search for Z' (Di-muon events)
- Search for W' (single muon/ jets)
- Search for resonance decaying to t-tbar
- Search for diboson resonances
- Monojets + invisible

Extra Dimensions/ Exotics



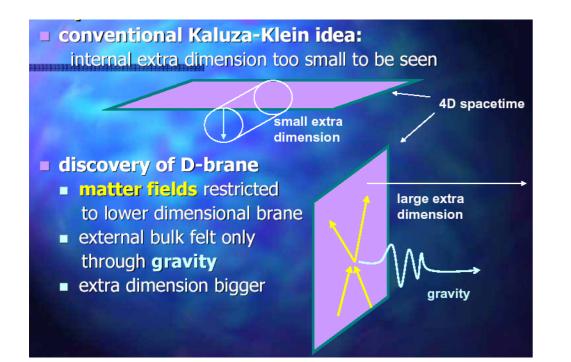
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Exotics

Extra Dimensions/ Exotics



Experiment

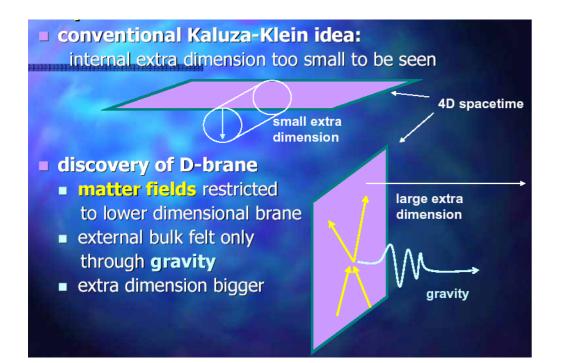
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Exotics

Leptoquarks

Extra Dimensions/ Exotics



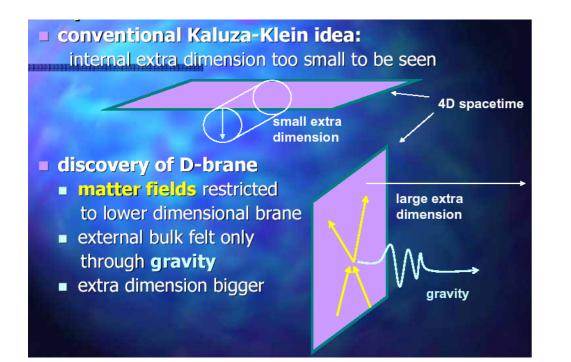
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- Leptoquarks
- Long-lived particles

Extra Dimensions/ Exotics



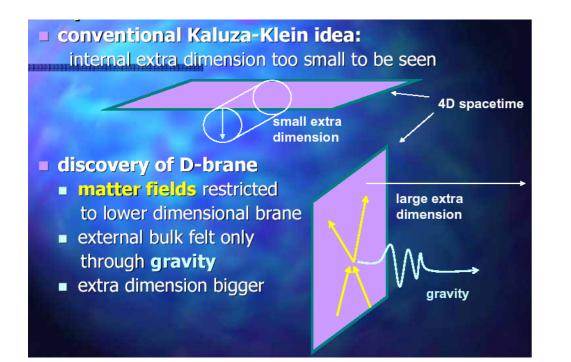
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- Leptoquarks
- Long-lived particles
- Off-pointing photons

Extra Dimensions/ Exotics



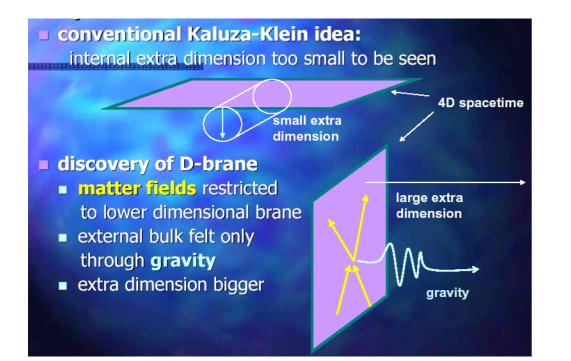
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Extra Dimensions/ Exotics



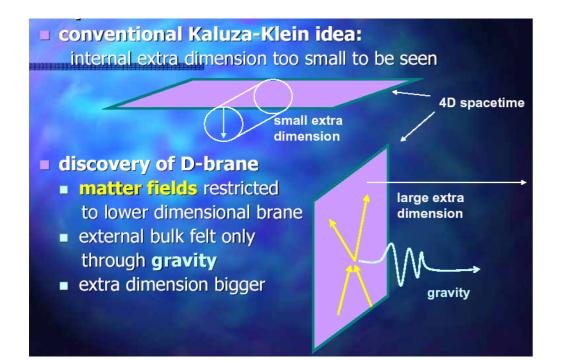
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- Off-pointing photons
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- Contact interactions

Extra Dimensions/ Exotics



Experiment

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Exotics

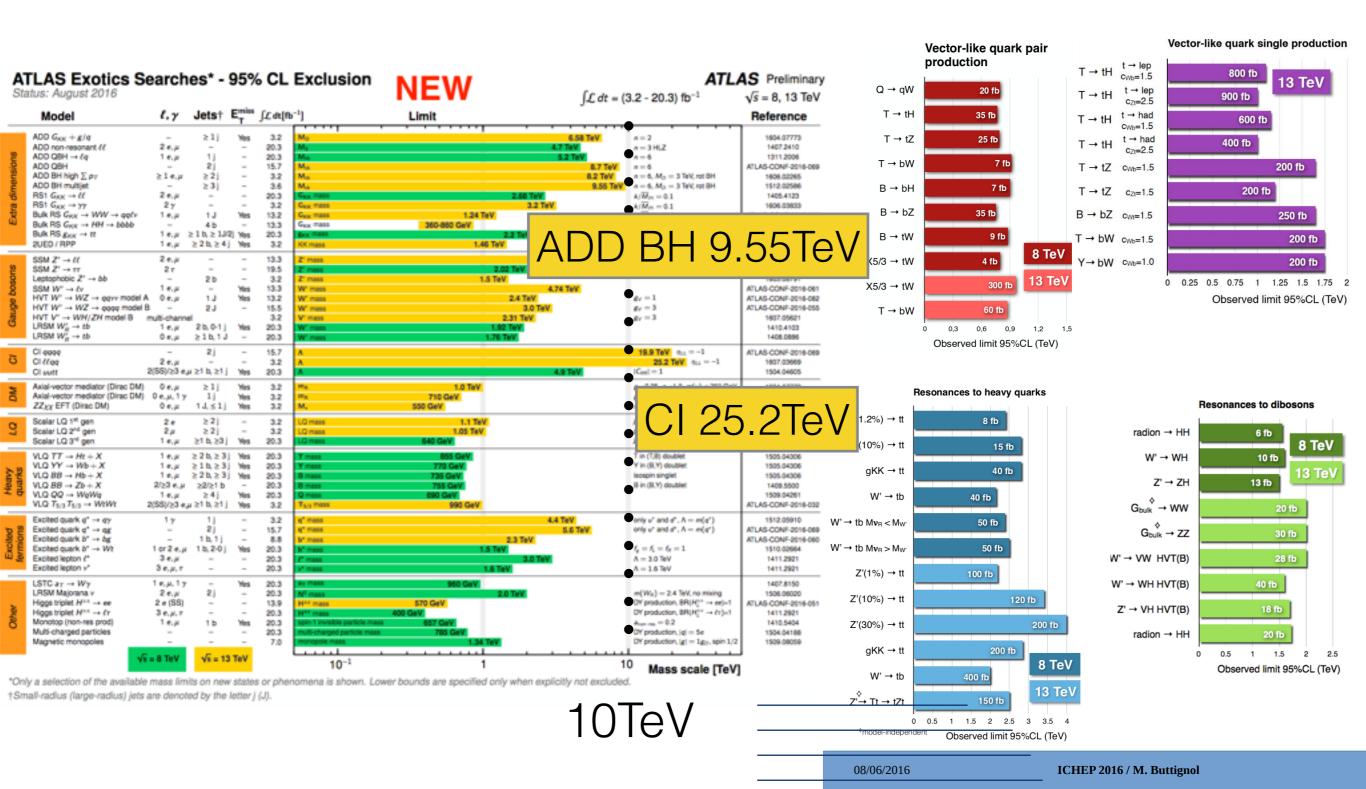
- Leptoquarks
- Long-lived particles
- Off-pointing photons
- Excited fermions
- Contact interactions

Drawback: No real motivation -> Unknown scale



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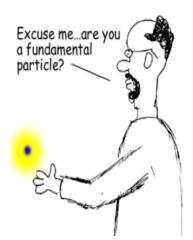
- Up to 25% mass limit increase by extending 2015 to 2016
- ~50% of the analyses updated to Run2

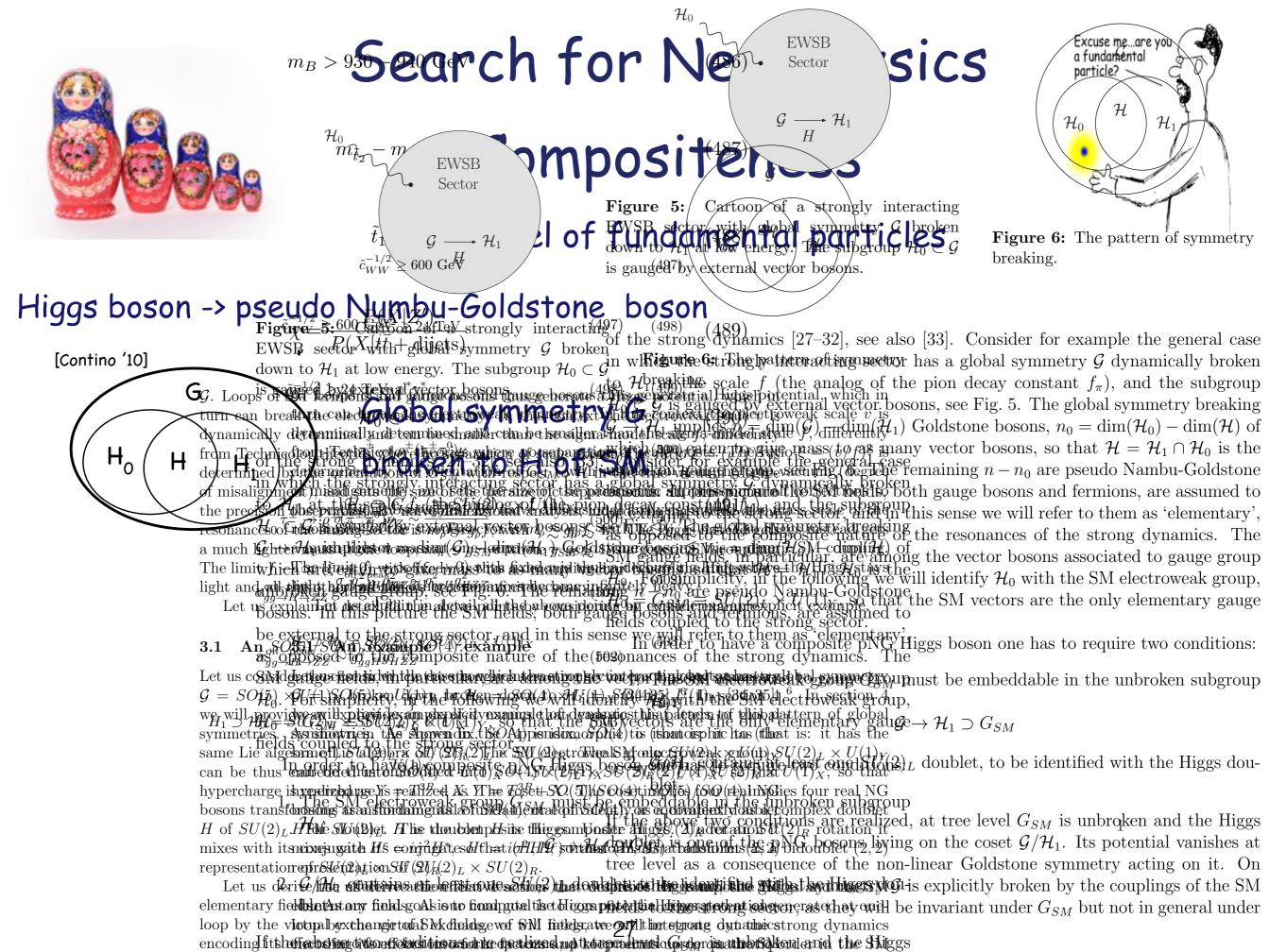




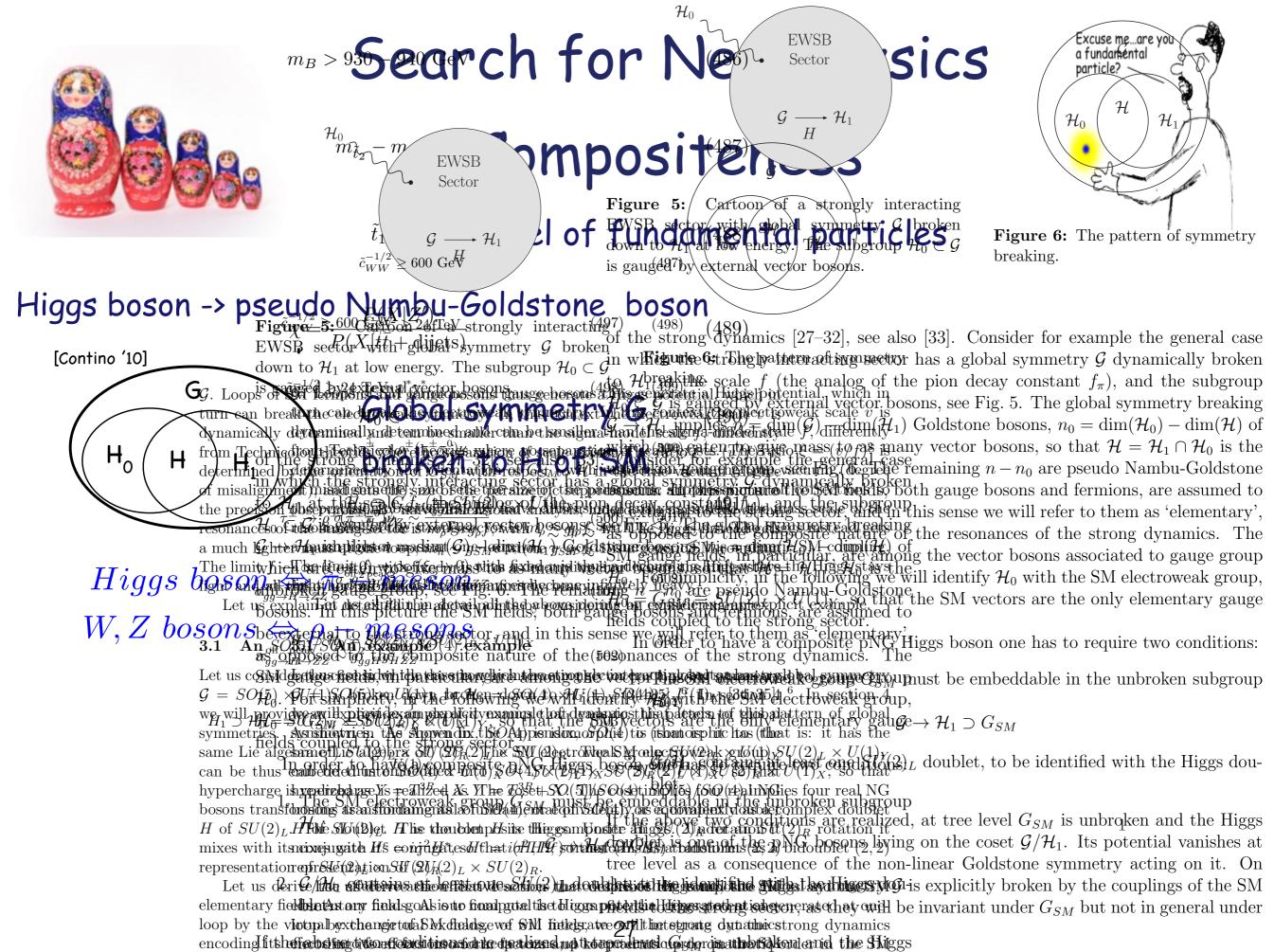
Compositeness

New level of fundamental particles





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Target #1: Higgs sector

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Matter Matter

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Mathebra Constant and Series and

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Target #3: New physics (supersymmetry)

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MLHC experiments are at the front line of mystery land: be patient

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Future development of HEP crucially depends on LHC outcome

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Future development of HEP crucially depends on LHC outcome

MLHC experiments are at the front line of mystery land: be patient

Target #1: Higgs sector

Target #2: Dark Matter

Target #3: New physics (supersymmetry)

Sector of the se

Complimentary searches for dark matter and insights in neutrino physics are of extreme importance

MLHC experiments are at the front line of mystery land: be patient

Target #1: Higgs sector

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Complimentary searches for dark matter and insights in neutrino physics are of extreme importance

The areas that were left behind come to the front: confinement, exotic hadrons, dense hadron matter

MLHC experiments are at the front line of mystery land: be patient

Target #1: Higgs sector

Arget #2: Dark Matter
 Target #3: New physics (superstries Will come!
 Future development discoveries Will come that discoveries on LHC outcome
 Complime bet that of ark matter and insights in neutrino physics are 1 extreme importance

The areas that were left behind come to the front: confinement, exotic hadrons, dense hadron matter