

LHC - Why Terascale?

Stabilization of the Higgs mechanism

$$\rightarrow \Lambda \sim 1 \text{ TeV}$$

Unitarization of EW vector boson and heavy quark amplitudes

$$\rightarrow \Lambda \sim 1 \text{ TeV}$$

If $M_h \sim 1 \text{ TeV} \rightarrow$ SM Higgs width $\sim 0.5 \text{ TeV}$, strong coupling regime

Dark Matter density: in most popular scenarios masses of DM candidates are less than 1 TeV

$$\Omega_{\text{WIMP}} \sim 0.2 \left(\frac{m_\chi}{200 \text{ GeV}} \right)^2 \left(\frac{0.1}{g^2} \right)^2$$

LHC - Why Terascale?

After Higgs discovery

Stabilization of the Higgs mechanism

→ $\Lambda \sim 1 \text{ TeV}$

Unitarization of EW vector boson and heavy quark amplitudes

→ $\Lambda \sim 1 \text{ TeV}$

If $M_h \sim 1 \text{ TeV} \rightarrow$ ~~SM Higgs width $\sim 0.5 \text{ TeV}$, strong coupling regime~~

Dark Matter density: in most popular scenarios masses of DM candidates are less than 1 TeV

$$\Omega_{\text{WIMP}} \sim 0.2 \left(\frac{m_\chi}{200 \text{ GeV}} \right)^2 \left(\frac{0.1}{g^2} \right)^2$$

LHC - Why Terascale?

After Higgs discovery

Stabilization of the Higgs mechanism

→ $\Lambda \sim 1 \text{ TeV}$

?

Unitarization of EW vector boson and heavy quark amplitudes

→ $\Lambda \sim 1 \text{ TeV}$

If $M_h \sim 1 \text{ TeV}$ → ~~SM~~ Higgs width $\sim 0.5 \text{ TeV}$, strong coupling regime

Dark Matter density: in most popular scenarios masses of DM candidates are less than 1 TeV

?

$$\Omega_{\text{WIMP}} \sim 0.2 \left(\frac{m_\chi}{200 \text{ GeV}} \right)^2 \left(\frac{0.1}{g^2} \right)^2$$

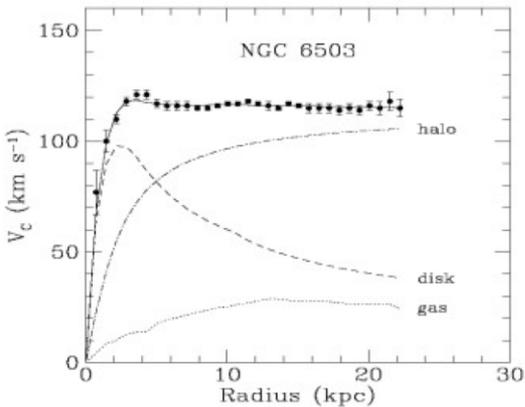
Facts which can not be explained in SM

- EW symmetry is broken - photon is massless, W and Z are massive particles
- Fermions have very much different masses
($M_{top} \approx 172 \text{ GeV}$, $m_e \approx 0.5 \text{ MeV}$, $\Delta m_\nu \approx 10^{-3} \text{ eV}$)

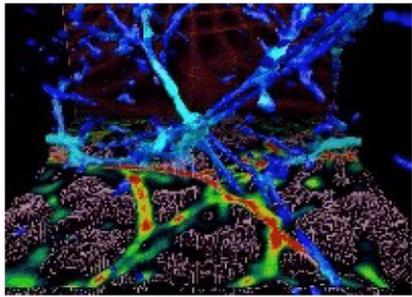
- Dark Matter exists in the Universe



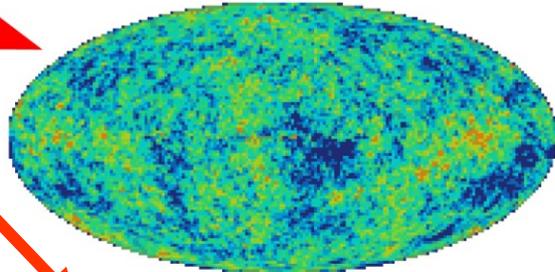
Rotation curves of galaxies



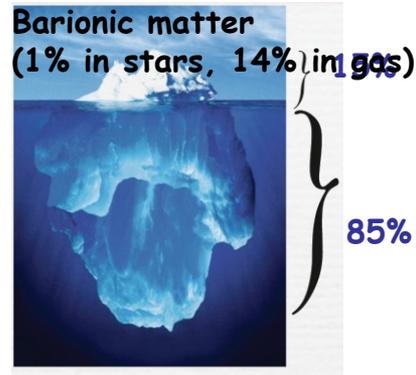
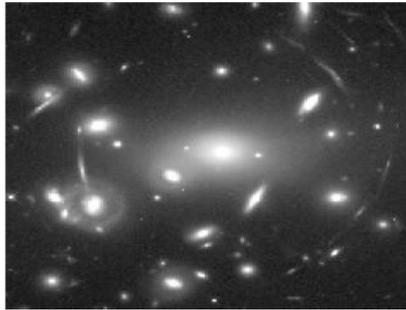
Large Scale Structure



CMB



Lensing



- $(g-2)_\mu$ (about 3.5σ)
- Neutrino masses, mixing, oscillations
- Particle - antiparticle asymmetry in the Universe,

CP violation

baryon asymmetry: $\frac{n_B - n_{\bar{B}}}{n_B + n_{\bar{B}}} \sim 10^{-10}$

CKM phase - too small effect

In addition to mentioned problems (naturalness/hierarchy, dark matter content, CP violation) SM does not give answers to many questions

What is a generation? Why there are only 3 generations?

How quarks and leptons related to each other, what is a nature of quark-lepton analogy?

What is responsible for gauge symmetries, why charges are quantize?
Are there additional gauge symmetries?

What is responsible for a formation of the Higgs potential?

To which accuracy the CPT symmetry is exact?

Why gravity is so weak comparing to other interactions?

.....

Main options beyond SM



1. Fundamental Higgs:

- Supersymmetric models
(MSSM, NMSSM...)

2. Composite Higgs:

- Models with new strong dynamics
(Chiral Lagrangians from holography, latest technicolor variants,
Little Higgs models, Twin Higgs models...)

3. Mixed cases:

- Models with extra space dimensions
- Partially composite models...

4. Many more (hidden valleys, landscape)

BSM searches

Collision energy $>$ particle production threshold

-Searches for new particles

strongly interacting new particles with large cross sections (squarks, gluinos...)

top partners motivated by naturalness (stop, sbottom, vector like quarks, t^* ...)

new resonances predicted by many BSM extensions (Z' , W' , π_T , ρ_T , KK states, ..)

extended Higgs sector (new neutral Higgses, charged Higgs)

Collision energy $<$ particle production threshold

-Anomalous/new interactions of SM particles (EFT)

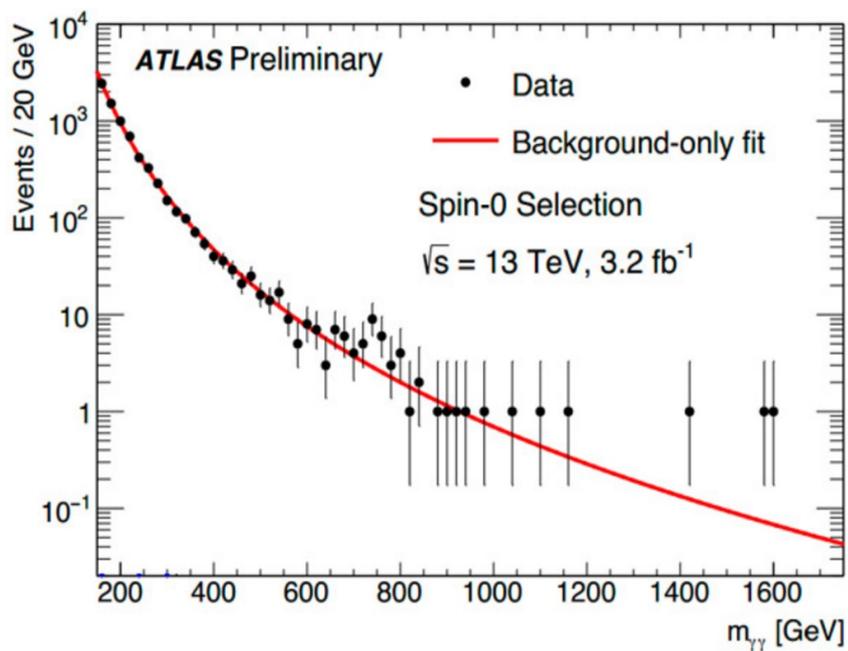
(anom. gauge boson couplings, anom. Wtb couplings, FCNC ...)

-New particle contributions via quantum loops

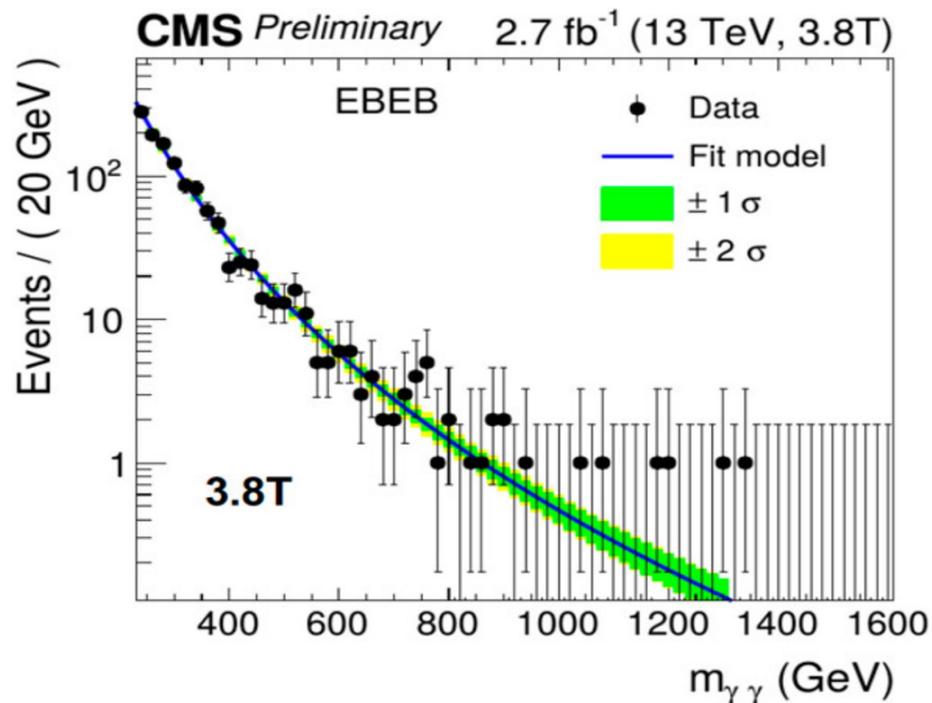


**More Higgses, more scalars
practically in all BSM models**

The 750 GeV diphoton excess presented by CMS and ATLAS experiments in Dec. 2015



Local significance = 3,9 σ
Global significance = 2 σ



Local significance = 3.4 σ
Global significance = 1.6 σ

- Lot of proposals but no generally accepted explanations of the 750 GeV diphoton excess (> 550 submissions to arXiv)

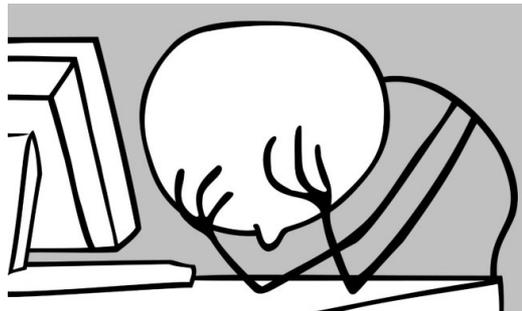
Composite Pseudo-Nambu-Goldstone boson, Dilaton, Radion, KK graviton, Quarkonium-like bound state, Sgoldstino, Heavy axion (axizilla), ...

- Important to rule out the scenarios, which cannot explain the excess

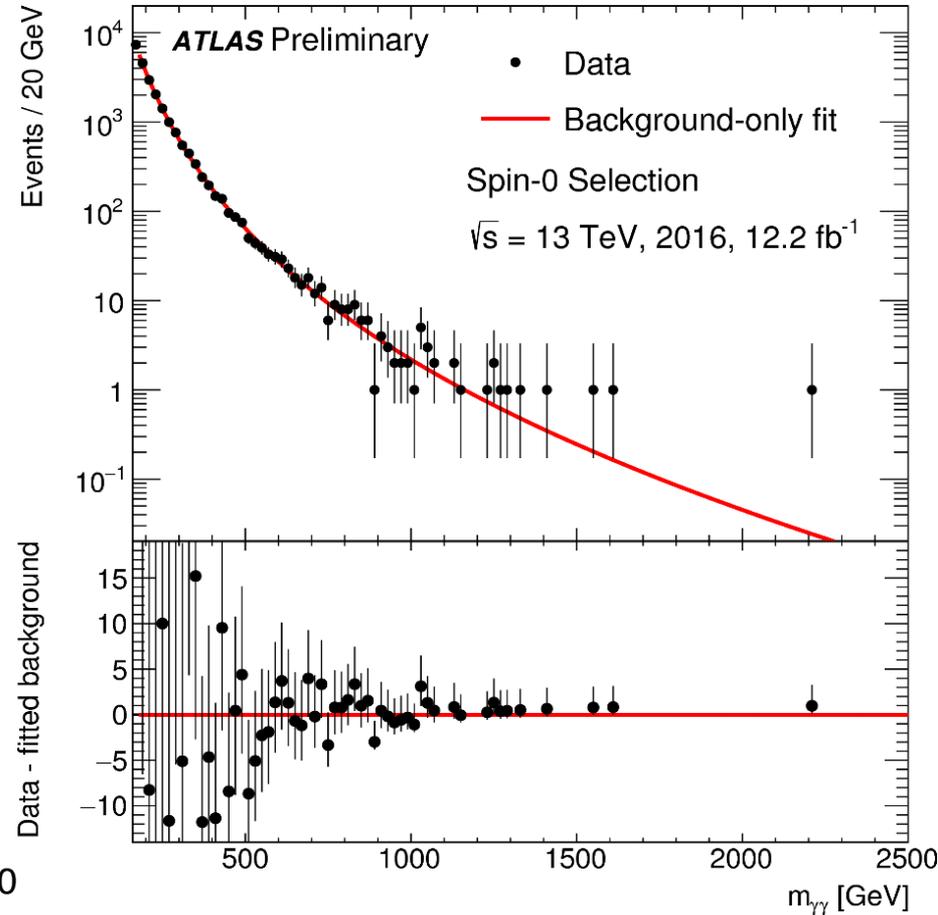
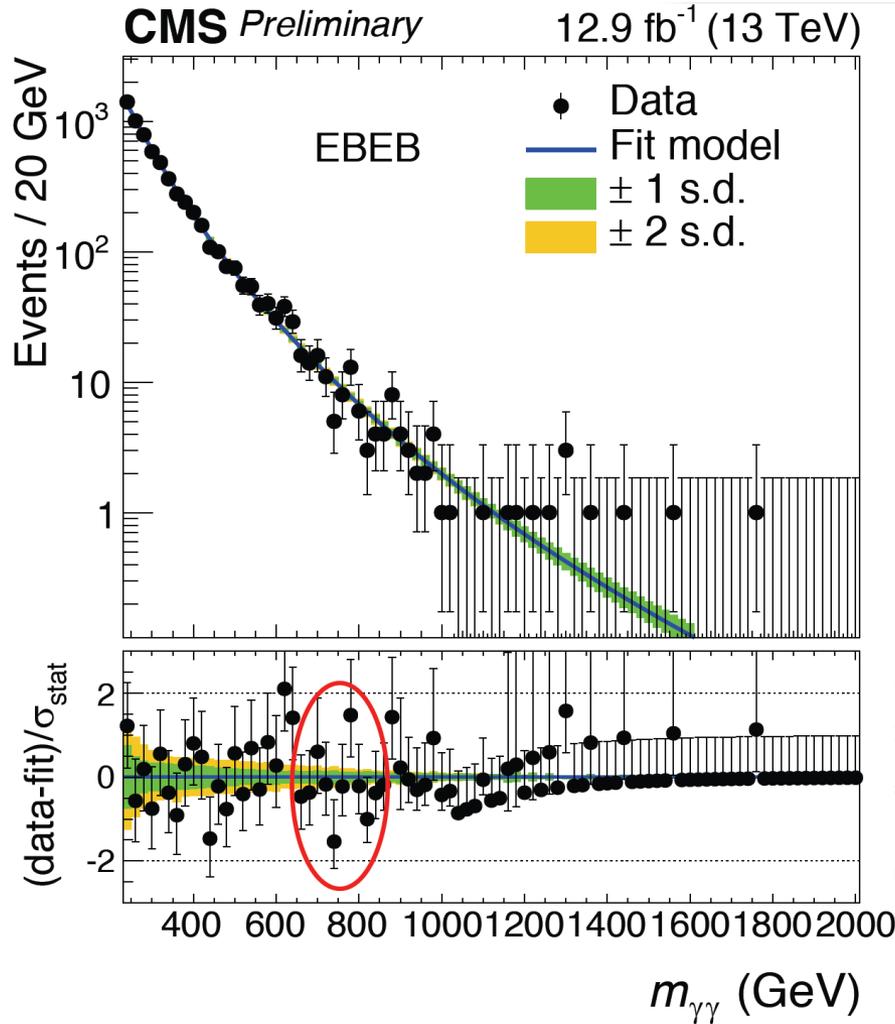
750 GeV excess



RS Radion



No signal with new much larger data set



Statistical fluctuation !

2HDM

Why the only one Higgs doublet?

- No fundamental reasons

Simple extension - two Higgs doublets (2HDM)

MSSM prototype, strong CP and axion, CP violation and baryogenesis

$$\langle \Phi_1 \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_1 \end{pmatrix}, \quad \langle \Phi_2 \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_2 \end{pmatrix} \quad v^2 \equiv v_1^2 + v_2^2 \quad \tan \beta \equiv \frac{v_2}{v_1}$$

2 complex scalar doublets => 8 degrees of freedom

$$\Phi_a = \begin{pmatrix} \phi_a^+ \\ (v_a + \rho_a + i\eta_a) / \sqrt{2} \end{pmatrix}, \quad a = 1, 2$$

As in the SM 3 Goldstone bosons are absorbed ("eaten") by W^\pm and Z

5 physics degrees of freedom

h, H - CP even scalars,
 A - CP odd scalar,
 H^\pm - charged scalars

Generic Higgs potential is not that simple

$$\begin{aligned}
 V = & m_{11}^2 \Phi_1^\dagger \Phi_1 + m_{22}^2 \Phi_2^\dagger \Phi_2 - m_{12}^2 (\Phi_1^\dagger \Phi_2 + \text{h.c.}) + \frac{1}{2} \lambda_1 (\Phi_1^\dagger \Phi_1)^2 + \frac{1}{2} \lambda_2 (\Phi_2^\dagger \Phi_2)^2 \\
 & + \lambda_3 (\Phi_1^\dagger \Phi_1) (\Phi_2^\dagger \Phi_2) + \lambda_4 (\Phi_1^\dagger \Phi_2) (\Phi_2^\dagger \Phi_1) \\
 & + \left\{ \frac{1}{2} \lambda_5 (\Phi_1^\dagger \Phi_2)^2 + [\lambda_6 (\Phi_1^\dagger \Phi_1) + \lambda_7 (\Phi_2^\dagger \Phi_2)] \Phi_1^\dagger \Phi_2 + \text{h.c.} \right\},
 \end{aligned}$$

Gunion, Haber,
Kane, Dawson '00...

Mostly studied cases with Z_2 symmetry $\Phi_1 \rightarrow +\Phi_1 \quad \Phi_2 \rightarrow -\Phi_2 \quad \rightarrow \lambda_6, \lambda_7 = 0$

Physics states - the states with definite masses

$$\begin{pmatrix} m_H^2 & 0 \\ 0 & m_h^2 \end{pmatrix} = \begin{pmatrix} c_\alpha & s_\alpha \\ -s_\alpha & c_\alpha \end{pmatrix} \begin{pmatrix} \mathcal{M}_{11}^2 & \mathcal{M}_{12}^2 \\ \mathcal{M}_{12}^2 & \mathcal{M}_{22}^2 \end{pmatrix} \begin{pmatrix} c_\alpha & -s_\alpha \\ s_\alpha & c_\alpha \end{pmatrix}$$

$$\begin{aligned}
 H &= (\sqrt{2} \text{Re } \Phi_1^0 - v_1) c_\alpha + (\sqrt{2} \text{Re } \Phi_2^0 - v_2) s_\alpha, \\
 h &= -(\sqrt{2} \text{Re } \Phi_1^0 - v_1) s_\alpha + (\sqrt{2} \text{Re } \Phi_2^0 - v_2) c_\alpha
 \end{aligned}$$

Notations: $\cos \alpha = c_\alpha$, $\sin \alpha = s_\alpha$, $\cos \beta = c_\beta$, $\sin \beta = s_\beta$,
 $\cos(\beta - \alpha) = c_{\beta - \alpha}$, $\sin(\beta - \alpha) = s_{\beta - \alpha}$

Several types of 2HDM depending on Yukawa arrangement

Glashow, Weinberg, Paschos condition '77

Avoid FCNC: if all fermions with the same quantum numbers couple to the same Higgs multiplet, then FCNC are absent

Branco, Ferreira, Lavoura, Rebelo, Sher, Silva '11,12

Model	u_R^i	d_R^i	e_R^i
Type I	Φ_2	Φ_2	Φ_2
Type II	Φ_2	Φ_1	Φ_1
Lepton-specific	Φ_2	Φ_2	Φ_1
Flipped	Φ_2	Φ_1	Φ_2

MSSM like
Type III
Type IV

Yukawa couplings to the Higgs bosons normalized to SM Higgs

	Type I	Type II	Lepton-specific	Flipped
ξ_h^u	$\cos \alpha / \sin \beta$	$\cos \alpha / \sin \beta$	$\cos \alpha / \sin \beta$	$\cos \alpha / \sin \beta$
ξ_h^d	$\cos \alpha / \sin \beta$	$-\sin \alpha / \cos \beta$	$\cos \alpha / \sin \beta$	$-\sin \alpha / \cos \beta$
ξ_h^ℓ	$\cos \alpha / \sin \beta$	$-\sin \alpha / \cos \beta$	$-\sin \alpha / \cos \beta$	$\cos \alpha / \sin \beta$
ξ_H^u	$\sin \alpha / \sin \beta$	$\sin \alpha / \sin \beta$	$\sin \alpha / \sin \beta$	$\sin \alpha / \sin \beta$
ξ_H^d	$\sin \alpha / \sin \beta$	$\cos \alpha / \cos \beta$	$\sin \alpha / \sin \beta$	$\cos \alpha / \cos \beta$
ξ_H^ℓ	$\sin \alpha / \sin \beta$	$\cos \alpha / \cos \beta$	$\cos \alpha / \cos \beta$	$\sin \alpha / \sin \beta$
ξ_A^u	$\cot \beta$	$\cot \beta$	$\cot \beta$	$\cot \beta$
ξ_A^d	$-\cot \beta$	$\tan \beta$	$-\cot \beta$	$\tan \beta$
ξ_A^ℓ	$-\cot \beta$	$\tan \beta$	$\tan \beta$	$-\cot \beta$

$s_{(\beta-\alpha)} + c_{(\beta-\alpha)}/t_\beta$

$s_{(\beta-\alpha)} - t_\beta^* c_{(\beta-\alpha)}$

MSSM

MSSM potential after supersymmetry breaking

$$V(H_1, H_2) = m_1^2 H_1^\dagger H_1 + m_2^2 H_2^\dagger H_2 + m_3^2 (H_1^T i\tau_2 H_2 + h.c.) + \frac{\lambda_1}{2} (H_1^\dagger H_1)^2 + \frac{\lambda_2}{2} (H_2^\dagger H_2)^2 + \lambda_3 (H_1^\dagger H_1) (H_2^\dagger H_2) + \lambda_4 |(H_1^T i\tau_2 H_2)|^2$$

2HDM type II with quartic couplings fixed due to the gauge nature

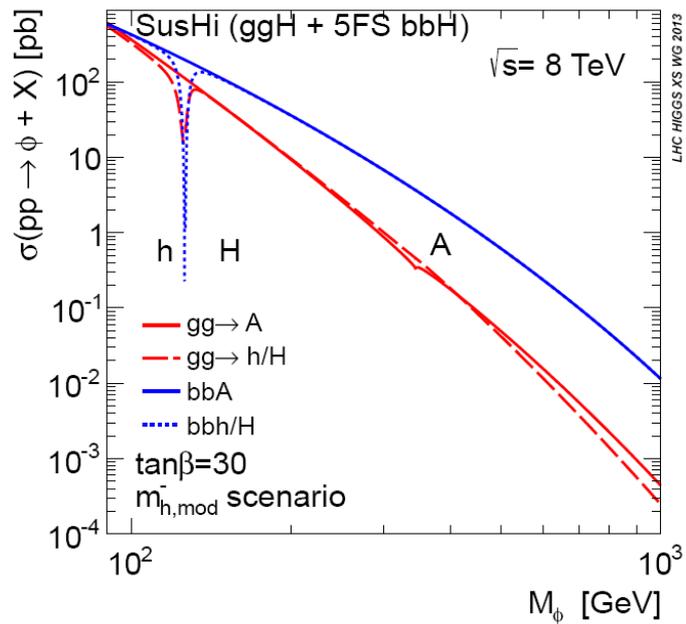
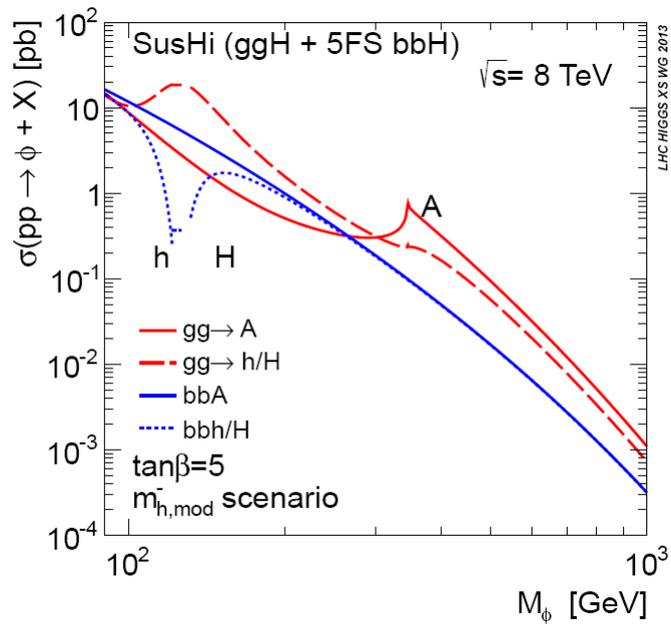
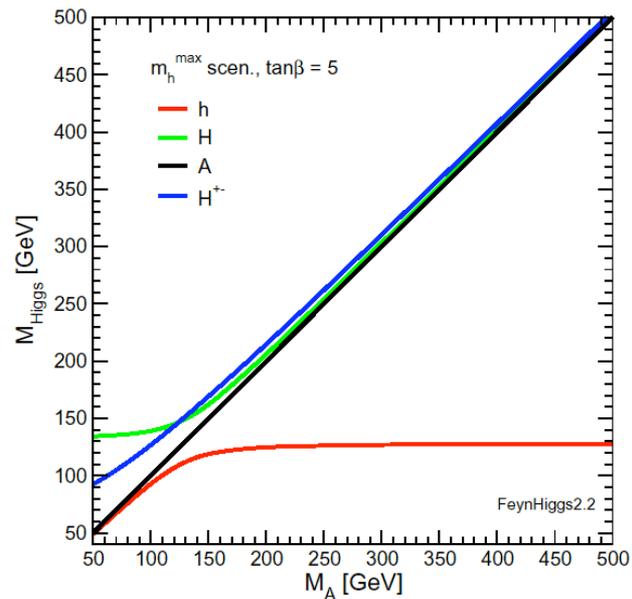
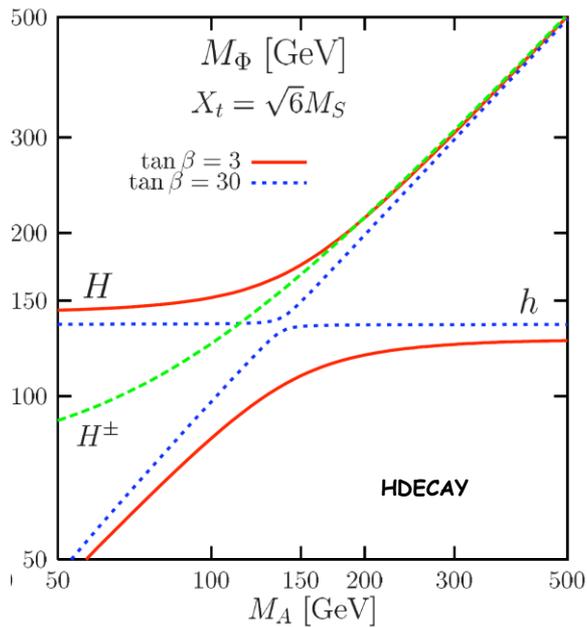
$$\lambda_1 = \lambda_2 = \frac{g_1^2 + g_2^2}{4}, \quad \lambda_3 = \frac{g_2^2 - g_1^2}{4}, \quad \lambda_4 = -\frac{g_2^2}{2}$$

8-3=5 physics states

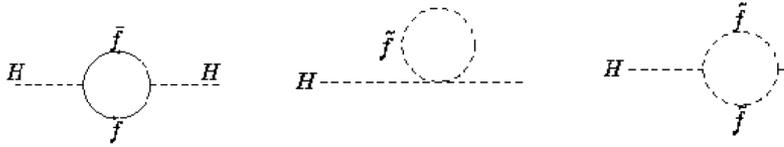
h, H - CP even scalars,
A - CP odd scalar,
H[±] - charged scalars

Φ	$g_{\Phi\bar{u}u}$	$g_{\Phi\bar{d}d}$	$g_{\Phi VV}$	$g_{\Phi AZ}/g_{\Phi H^+W^-}$
h	$\cos \alpha / \sin \beta$	$-\sin \alpha / \cos \beta$	$\sin(\beta - \alpha)$	$\propto \cos(\beta - \alpha)$
H	$\sin \alpha / \sin \beta$	$\cos \alpha / \cos \beta$	$\cos(\beta - \alpha)$	$\propto \sin(\beta - \alpha)$
A	$\cot \beta$	$\tan \beta$	0	$\propto 0/1$

Couplings are shared between the Higgses: $\sum_i g_{H_i VV}^2 = (g_{HVV}^{\text{SM}})^2$



M_H is protected due to cancellation of Λ^2 dependence!



$$\Delta m_h^2 = \frac{3m_t^4}{4\pi^2 v^2} \left[\ln \left(\frac{M_{\text{SUSY}}^2}{m_t^2} \right) + \frac{X_t^2}{M_{\text{SUSY}}^2} \left(1 - \frac{X_t^2}{12M_{\text{SUSY}}^2} \right) \right]$$

($X_t = \sqrt{6} M_{\text{SUSY}}$ **Maximal mixing scenario**)

$$M_{\text{SUSY}} \equiv \sqrt{m_{\tilde{t}_1} m_{\tilde{t}_2}}$$

$$X_t = A_t - \mu \cot \beta$$

Only two parameters at tree level

$$\tan \beta \equiv \frac{v_2}{v_1}, \quad M_A$$

But large loop correction

$$M_h^2 \leq M_Z^2 + \Delta m_h^2$$

\swarrow \downarrow \searrow
 125 GeV² 91 GeV² 86 GeV²

Available parameter range after all constrains ?

Carena, Heinemeyer, Stal, Wagner, Weiglein'13

Parameter	m_h^{\max}	$m_h^{\text{mod}+}$	$m_h^{\text{mod}-}$	<i>light stop</i>	<i>light stau</i>	τ - <i>phobic</i>	<i>low-M_H</i>
m_t	173.2	173.2	173.2	173.2	173.2	173.2	173.2
M_A	varied	varied	varied	varied	varied	varied	110
$\tan \beta$	varied	varied	varied	varied	varied	varied	varied
M_{SUSY}	1000	1000	1000	500	1000	1500	1500
$M_{\tilde{t}_3}$	1000	1000	1000	1000	245 (250)	500	1000
$X_t^{\text{OS}}/M_{\text{SUSY}}$	2.0	1.5	-1.9	2.0	1.6	2.45	2.45
$X_t^{\text{MS}}/M_{\text{SUSY}}$	$\sqrt{6}$	1.6	-2.2	2.2	1.7	2.9	2.9
A_t	Given by $A_t = X_t + \mu \cot \beta$						
A_b	$= A_t$	$= A_t$	$= A_t$	$= A_t$	$= A_t$	$= A_t$	$= A_t$
A_τ	$= A_t$	$= A_t$	$= A_t$	$= A_t$	0	0	$= A_t$
μ	200	200	200	350	500 (450)	2000	varied
M_1	Fixed by GUT relation to M_2						
M_2	200	200	200	350	200 (400)	200	200
$m_{\tilde{g}}$	1500	1500	1500	1500	1500	1500	1500
$M_{\tilde{q}_{1,2}}$	1500	1500	1500	1500	1500	1500	1500
$M_{\tilde{l}_{1,2}}$	500	500	500	500	500	500	500
$A_{f \neq t, b, \tau}$	0	0	0	0	0	0	0

Intensively used in experimental analyses

hMSSM

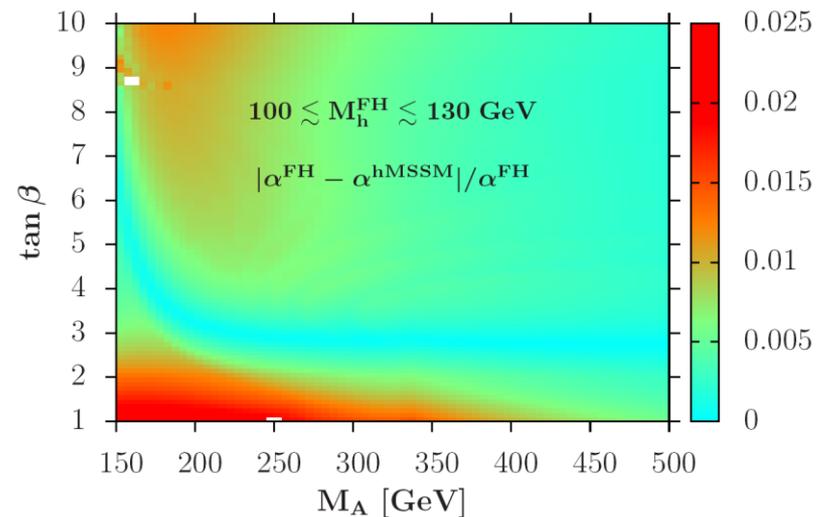
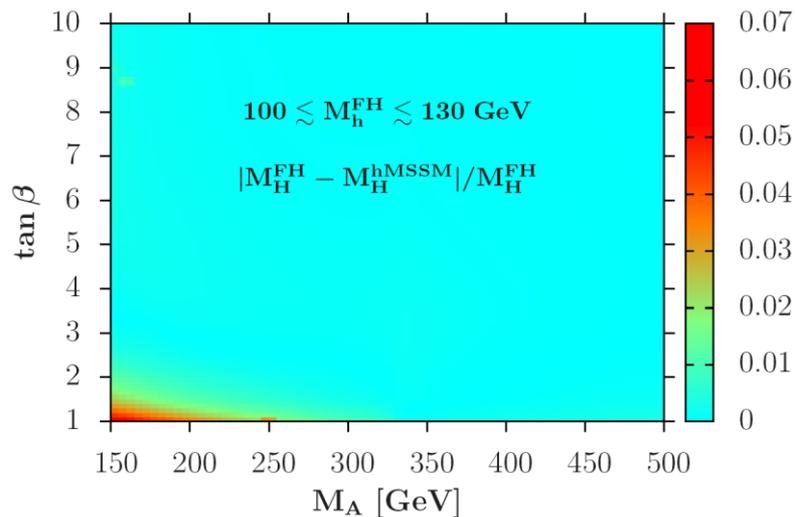
M_h is fixed to be 125 GeV

With few simplified assumptions one gets (including leading loops)

$$M_H^2 = \frac{(M_A^2 + M_Z^2 - M_h^2)(M_Z^2 \cos^2 \beta + M_A^2 \sin^2 \beta) - M_A^2 M_Z^2 \cos^2 2\beta}{M_Z^2 \cos^2 \beta + M_A^2 \sin^2 \beta - M_h^2}$$
$$\alpha = -\arctan \left(\frac{(M_Z^2 + M_A^2) \cos \beta \sin \beta}{M_Z^2 \cos^2 \beta + M_A^2 \sin^2 \beta - M_h^2} \right)$$

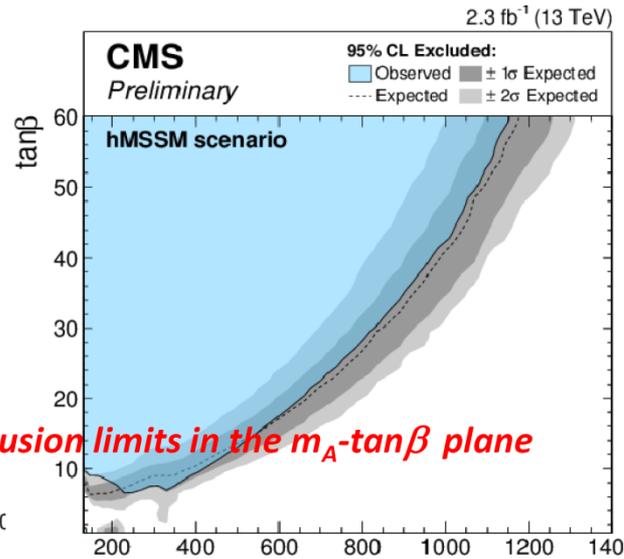
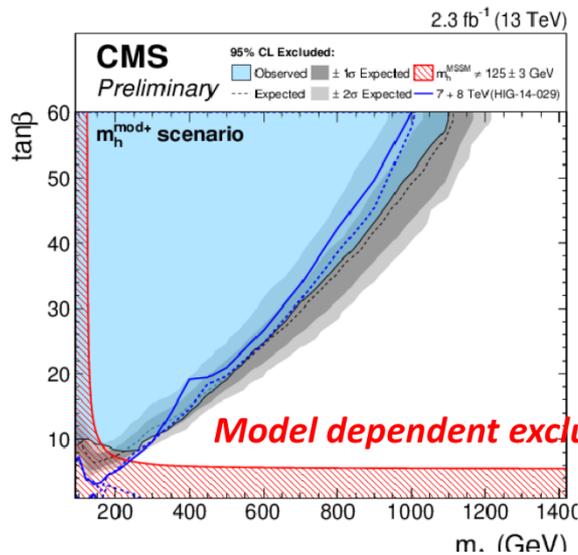
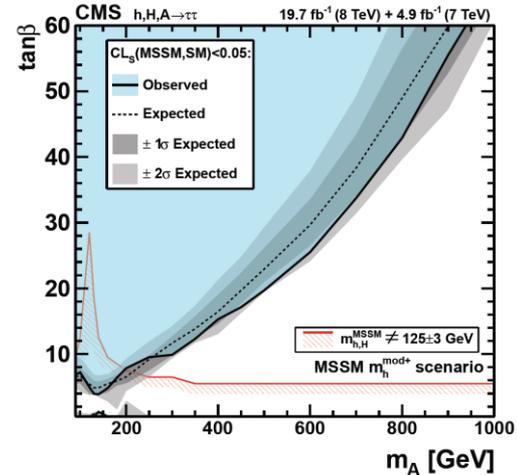
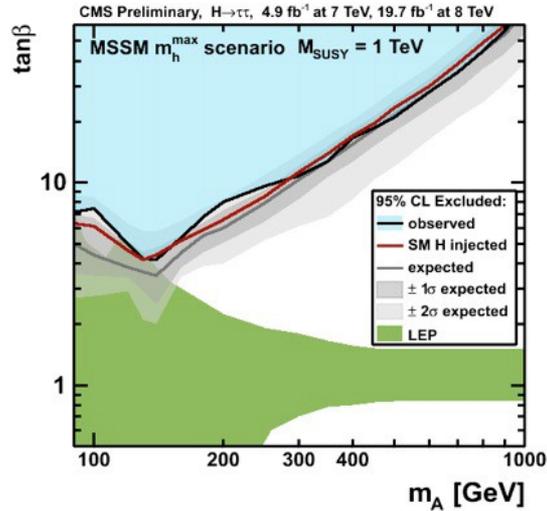
Djouadi, Maiani, Moreau,
Polosa, Quevillon, Riquer
(1502.05653)

Validation with FeynHiggs



CMS searches as an example

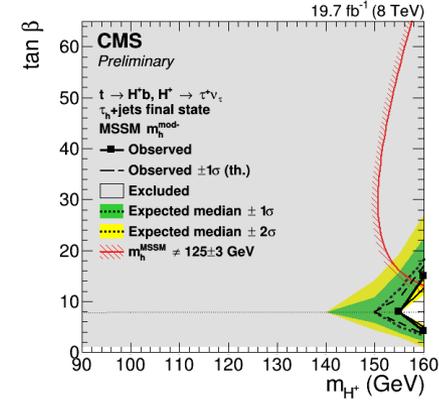
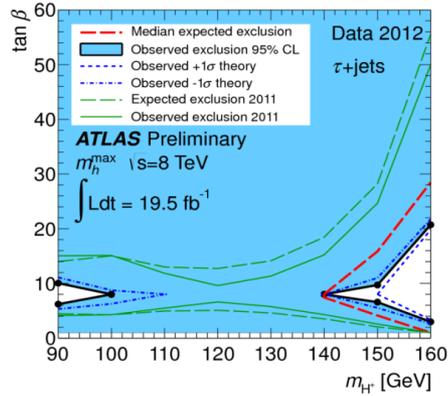
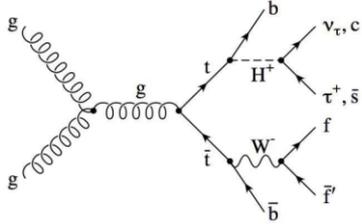
1408.3316



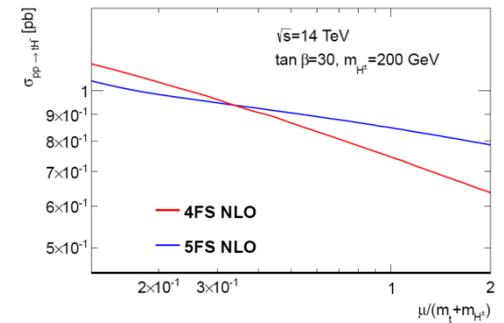
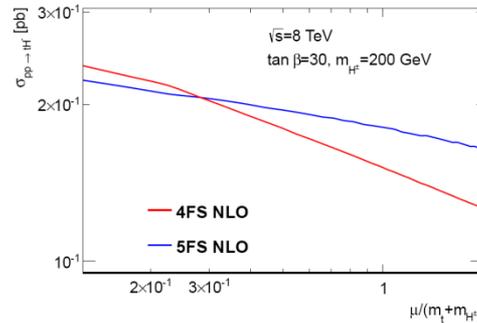
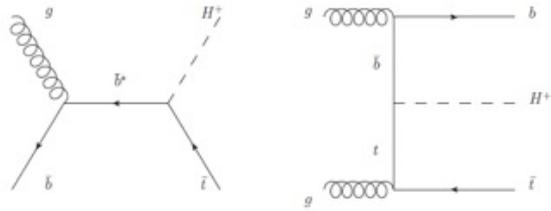
Model dependent exclusion limits in the m_A - $\tan\beta$ plane

Charged Higgses are predicted in many BSM (2HDM, MSSM, NMSSM...)

Light H^\pm in top decays



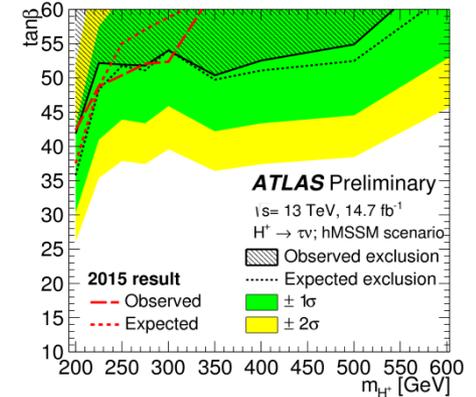
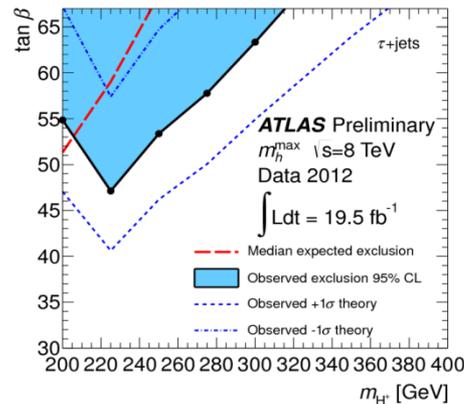
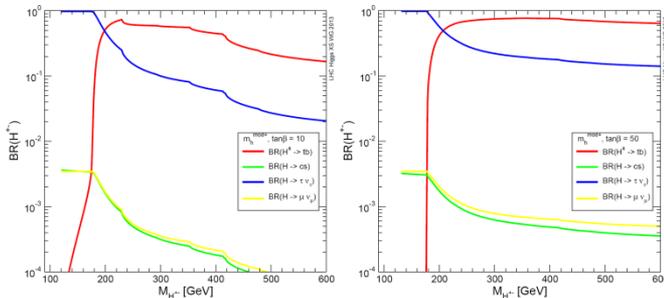
Heavy H^\pm in associated production



Flechl, Klees, Kraemer, Spira, Ubiali '14

Decay modes

Heinemeyer et al. '13



Supersymmetry is one of the most favorite BSM ideas, relating spin $\frac{1}{2}$ fermions with spin 0,1 bosons

$$Q|\text{Boson}\rangle = |\text{Fermion}\rangle \quad Q^\dagger|\text{Boson}\rangle = |\text{Fermion}\rangle$$

Fermion degrees of freedom \leftrightarrow boson degrees of freedom

Minimal particle content

□ Gauge / Gaugino Sector

Standard Bosons	Supersymmetric Partners
$W^\pm \quad H^\pm$	Charginos $\chi_1^\pm \chi_2^\pm$
$g \quad Z$ $h \quad H \quad A$	Neutralinos $\chi_1^0 \chi_2^0 \chi_3^0 \chi_4^0$
g_i	Gluinos \tilde{g}_i

[Two Higgs doublets]

[All fermions]

And also ...

Graviton G	Gravitino \tilde{G}
-------------------	---

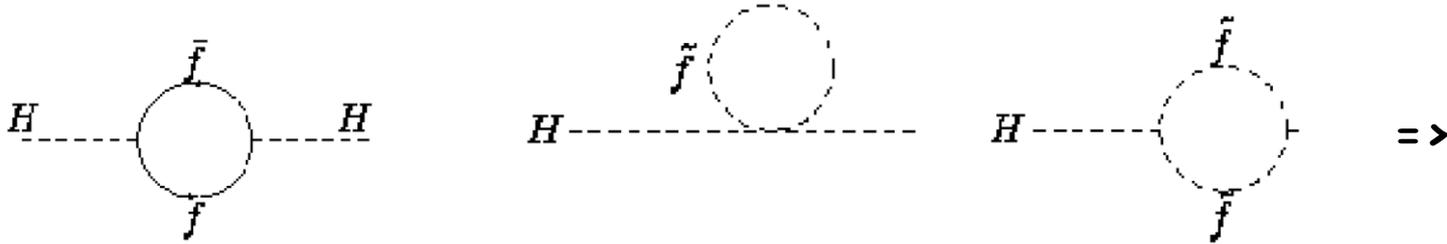
□ Particle / Sparticle Sector

Standard Particles	Supersymmetric Partners
Leptons l	Sleptons $\tilde{l}_{R,L}$
Neutrinos ν_l	Sneutrinos $\tilde{\nu}_l$
Quarks q	Squarks $\tilde{q}_{R,L}$

[All scalars]

SUSY

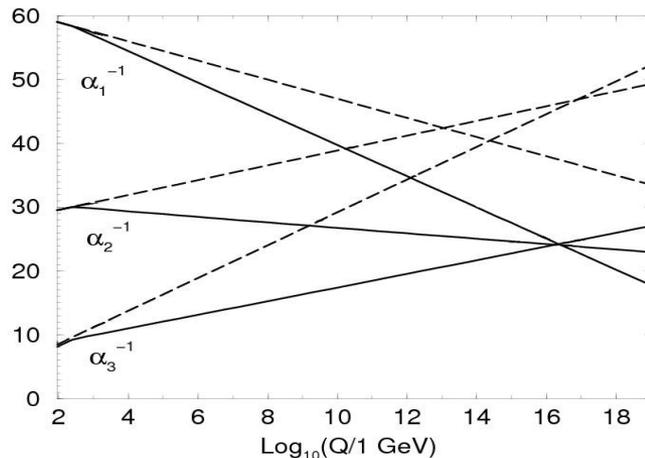
1. Cancellation of Λ^2 dependence



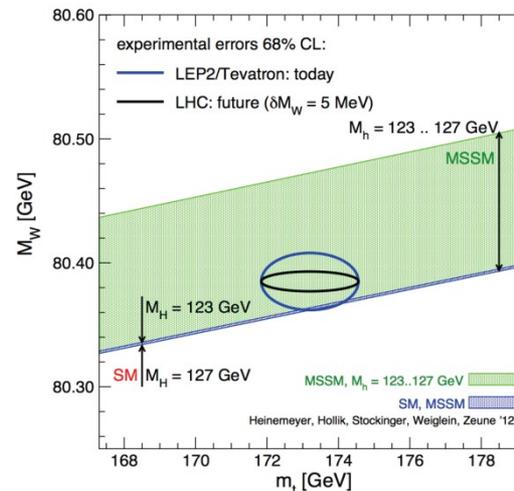
$$\Delta M_H^2|_{\text{tot}} = \frac{\lambda_f^2 N_f}{4\pi^2} \left[(m_f^2 - m_S^2) \log\left(\frac{\Lambda}{m_S}\right) + 3m_f^2 \log\left(\frac{m_S}{m_f}\right) \right] \quad M_H \text{ is protected!}$$

2. Lightest SUSY particle is stable (if R-parity) - Dark Matter candidate

3. Unification of couplings in contrast to SM

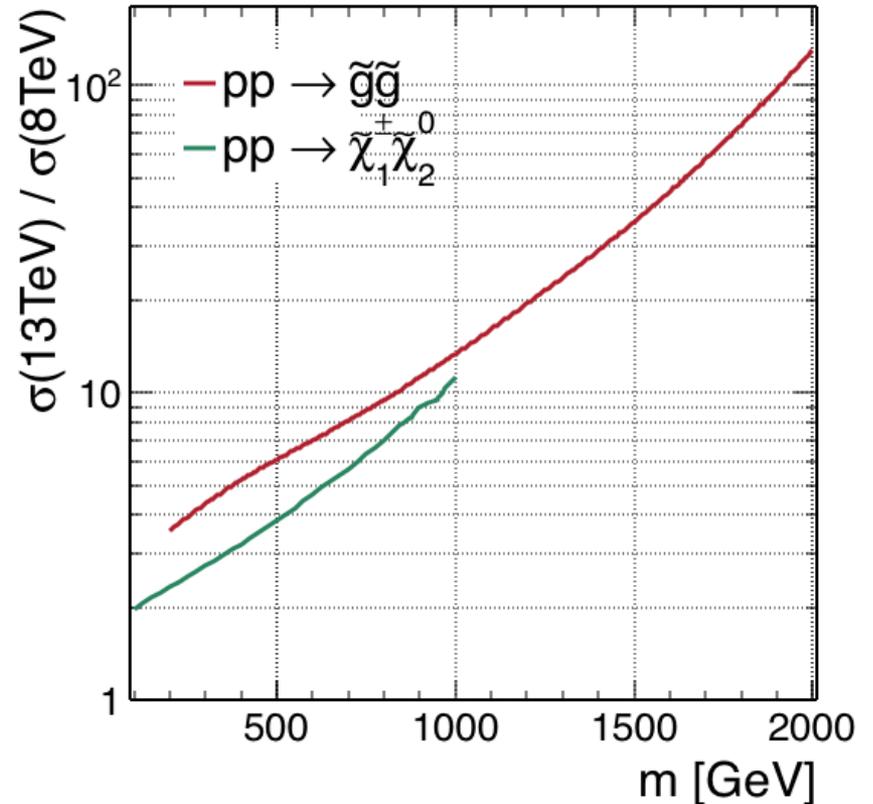
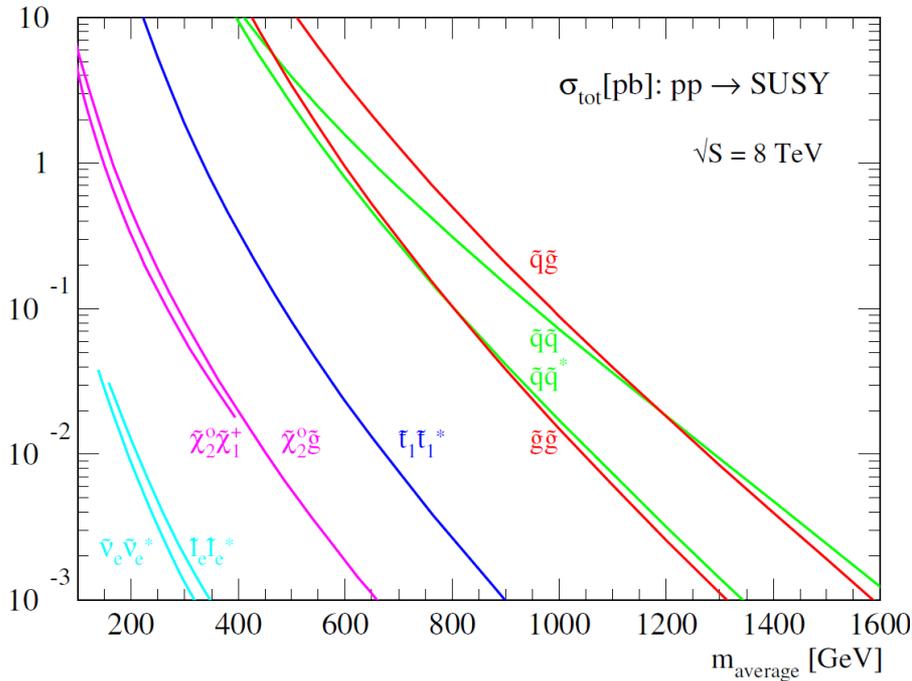
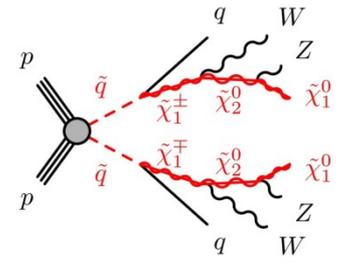
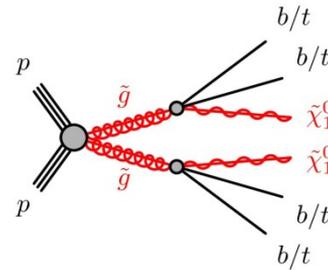
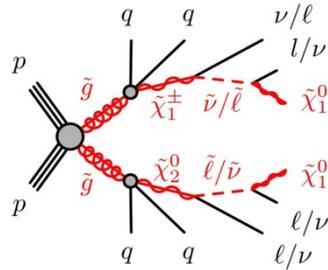


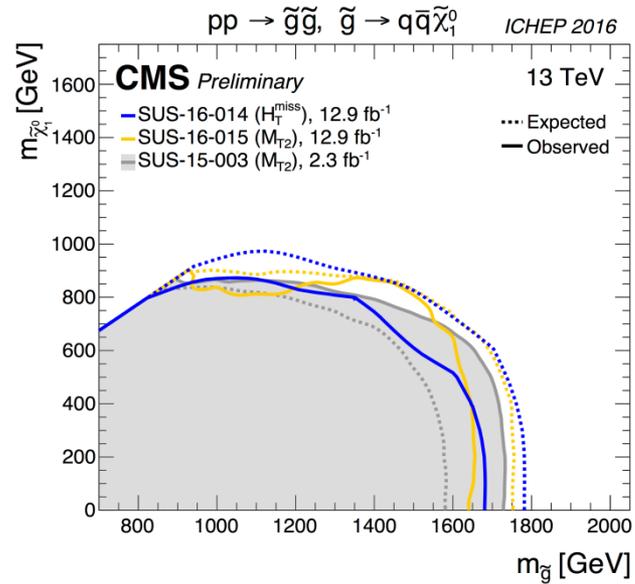
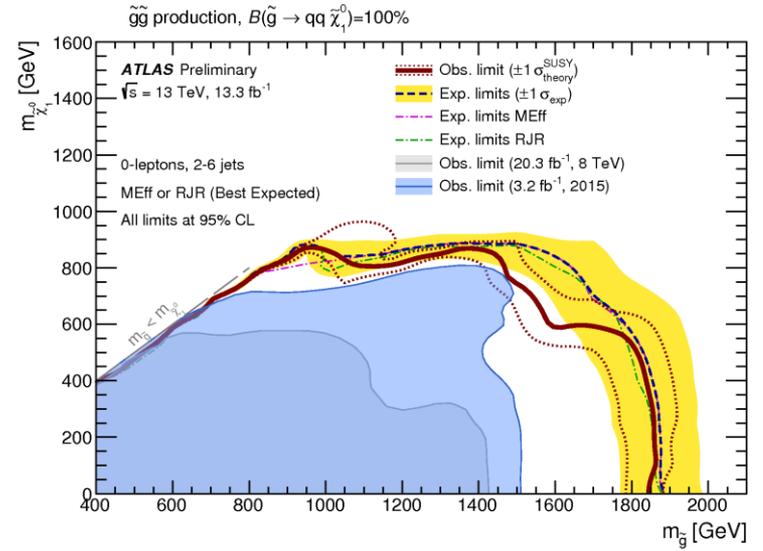
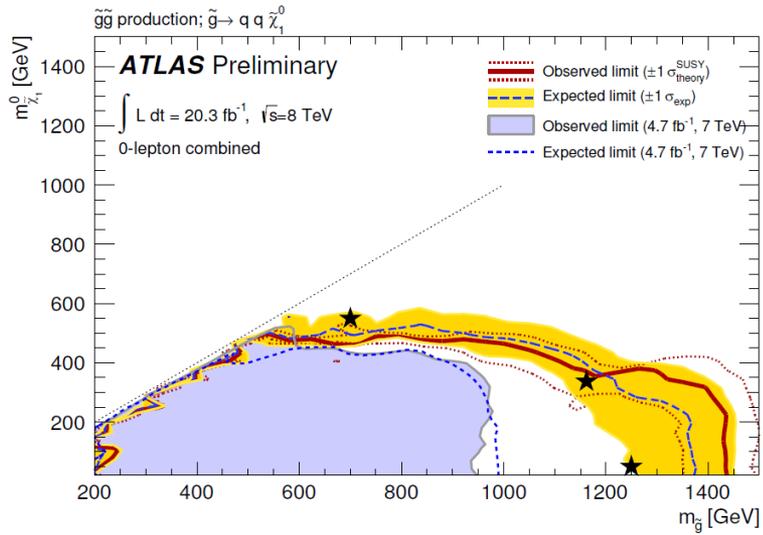
4. Fit of EW precision data



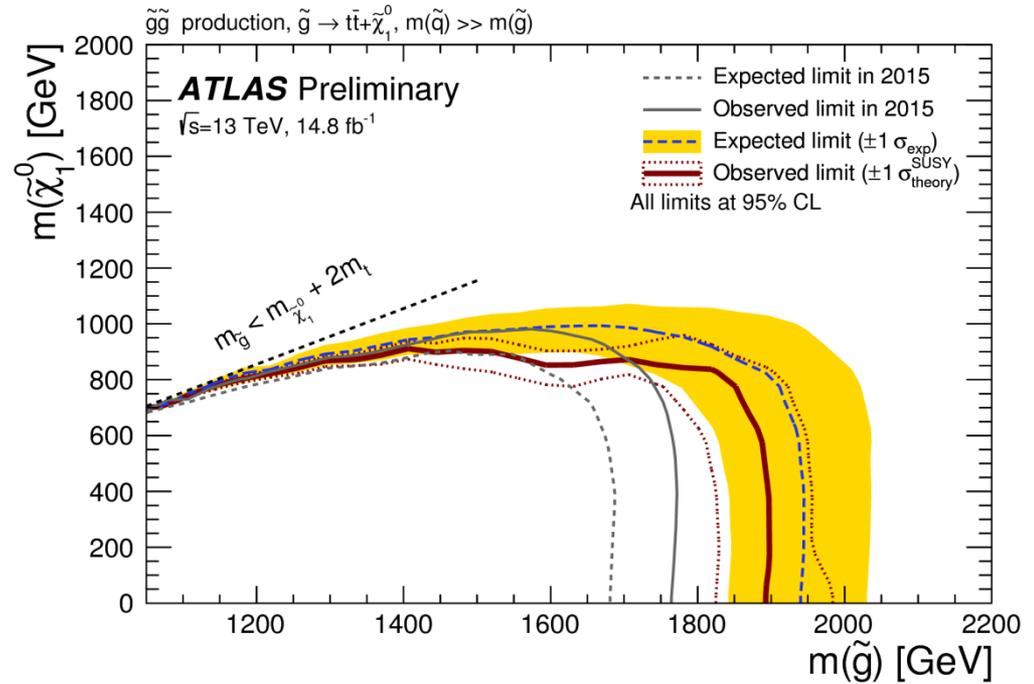
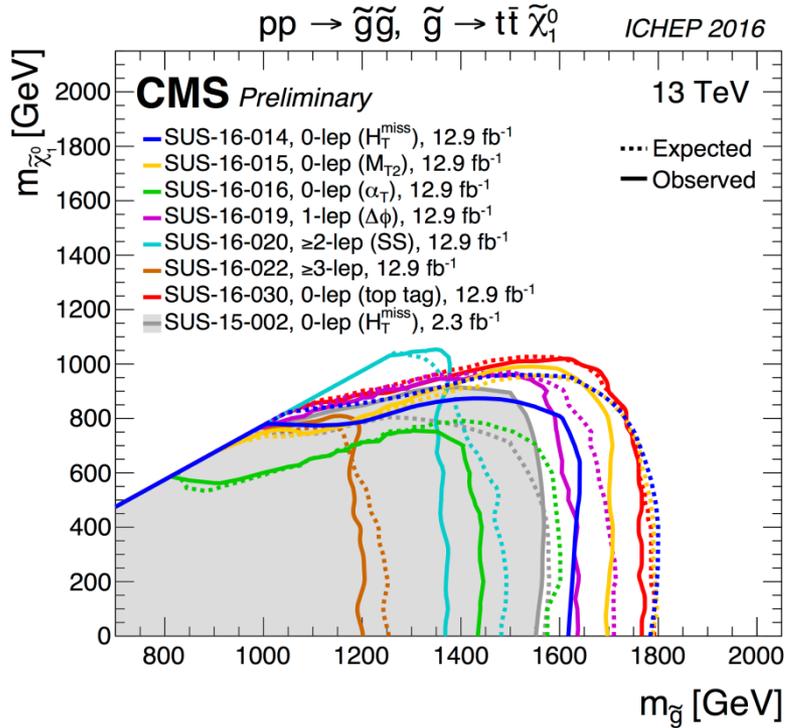
Searches for strongly interacting superpartners

Glino and squark signatures:



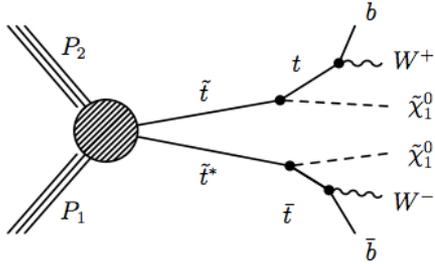


Glauino decays to $t\bar{t} + \text{LSP}$

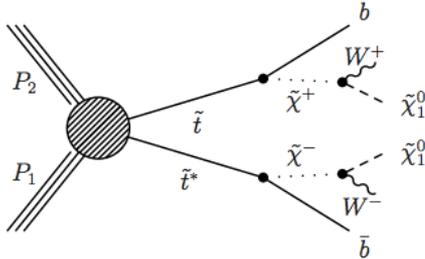


Searches for Stops

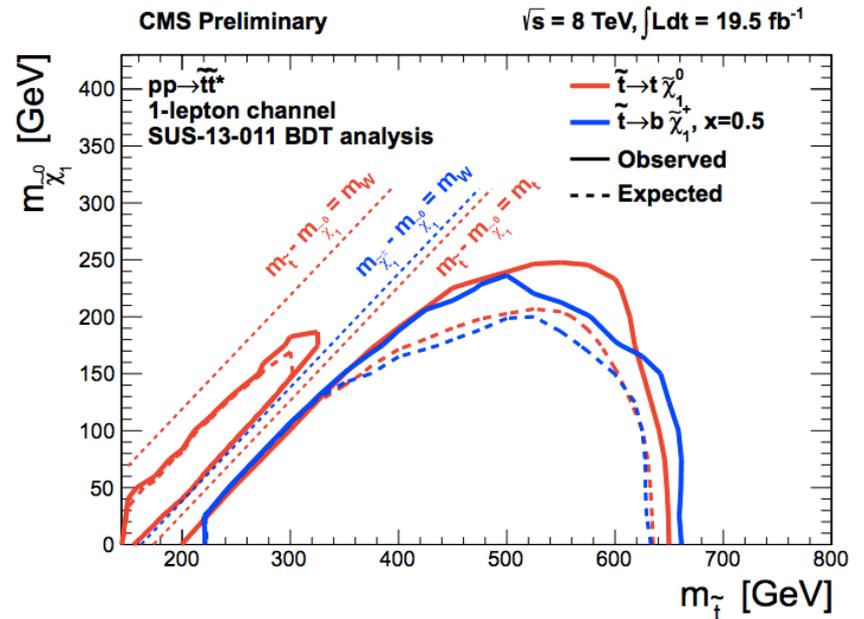
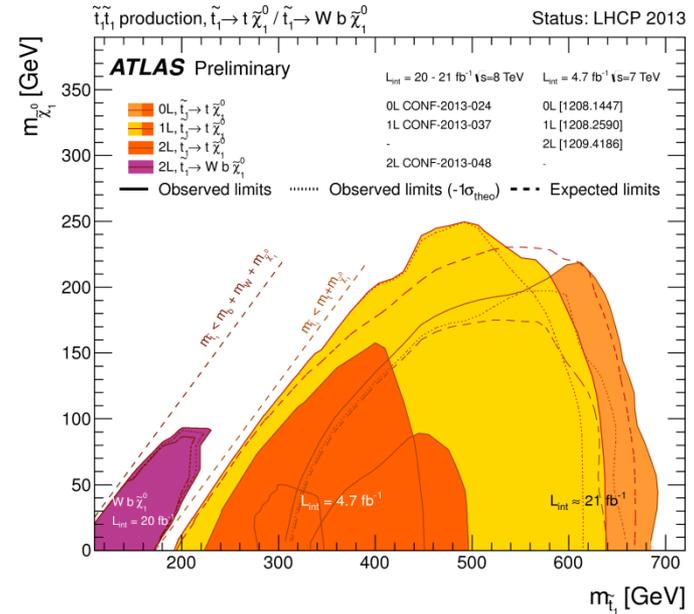
$$\tilde{t}\tilde{t} \rightarrow t\bar{t}\tilde{\chi}_1^0\tilde{\chi}_1^0$$

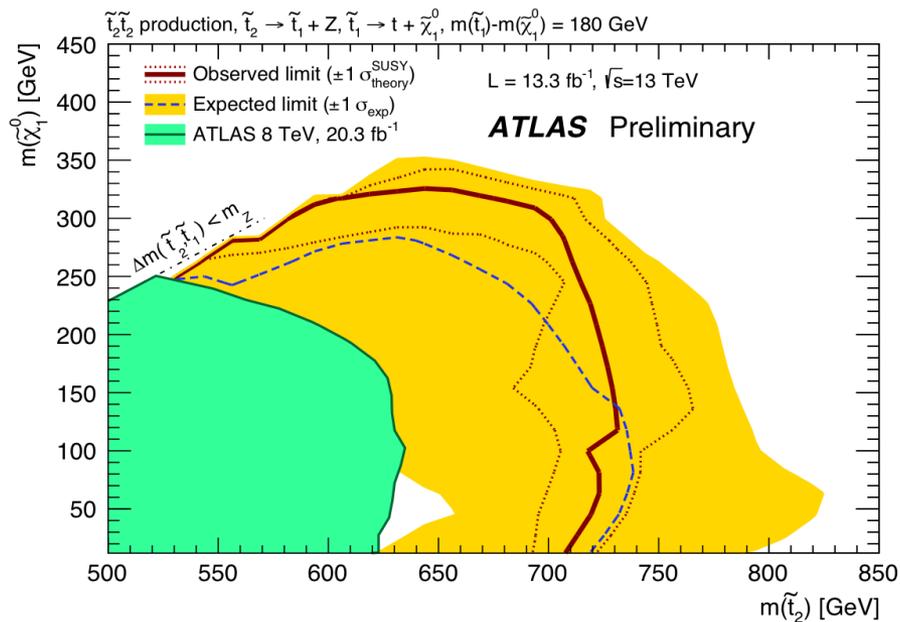
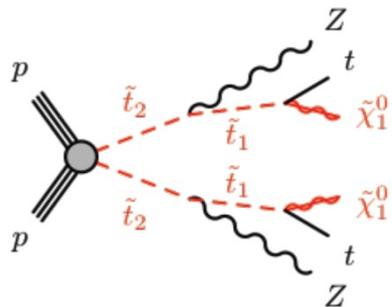
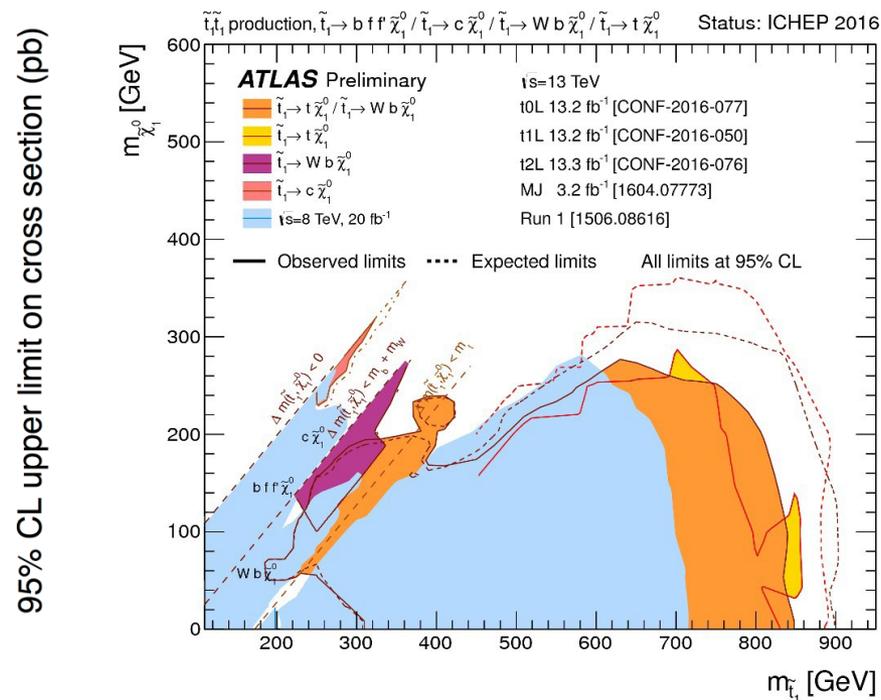
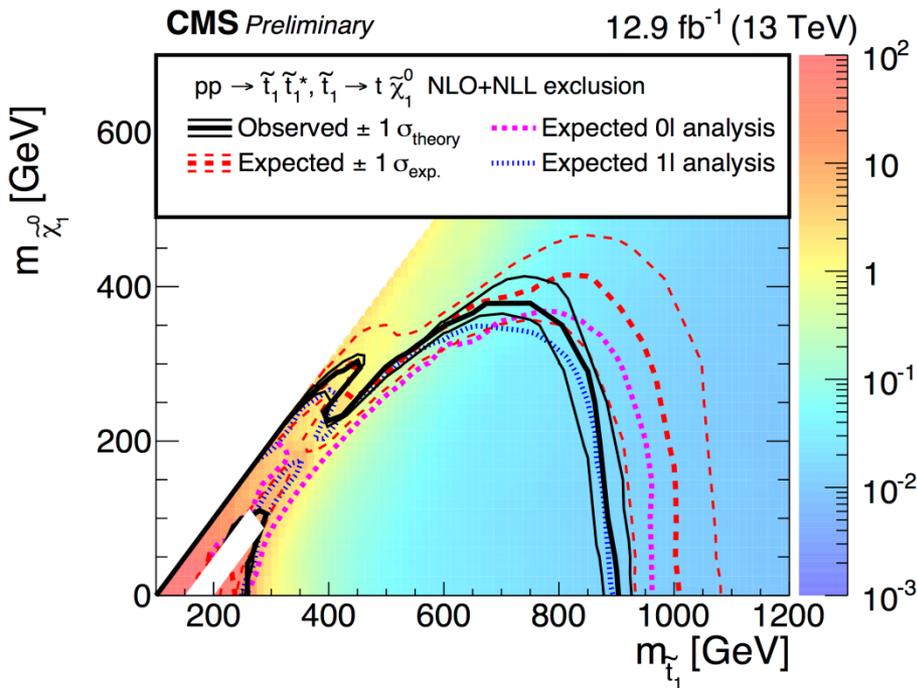


$$\tilde{t}\tilde{t} \rightarrow b\bar{b}\tilde{\chi}_1^+\tilde{\chi}_1^- \rightarrow b\bar{b}W^+W^-\tilde{\chi}_1^0\tilde{\chi}_1^0$$



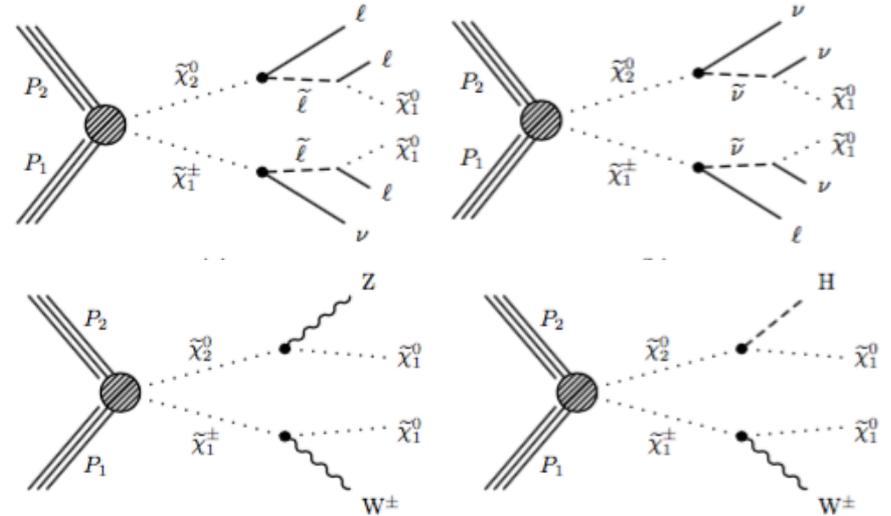
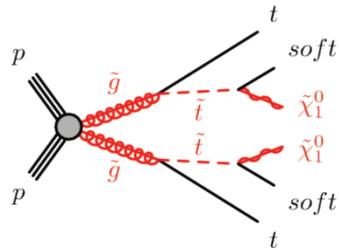
Mass exclusion limits:
 $M_{\text{stop}} \sim 660 \text{ GeV}$ and
 $M_{\text{bottom}} \sim 630 \text{ GeV}$



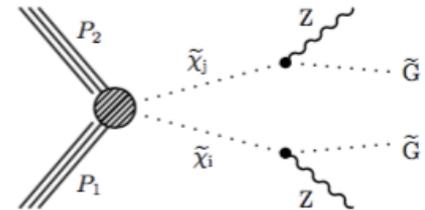
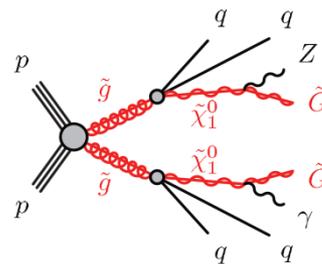


Many other searches for superpartners

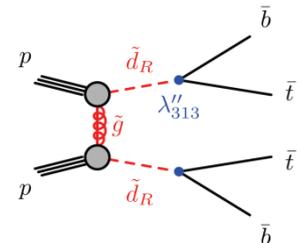
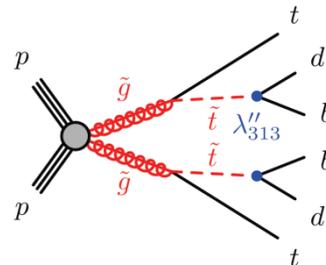
R-parity conserving scenarios



Gauge mediated scenarios



R-parity violating scenarios



ATLAS SUSY Searches* - 95% CL Lower Limits

Status: August 2016

ATLAS Preliminary

$\sqrt{s} = 7, 8, 13$ TeV

	Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt (fb^{-1})$	Mass limit	$\sqrt{s} = 7, 8$ TeV	$\sqrt{s} = 13$ TeV	Reference	
Inclusive Searches	MSUGRA/CMSSM	0-3 e, μ	1-2 τ	2-10 jets/3 b	Yes	20.3	\tilde{g}, \tilde{Z}	1.85 TeV	$m(\tilde{g})=m(\tilde{Z})$	1507.05525
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}$	0	2-6 jets	Yes	13.3		1.35 TeV	$m(\tilde{g}) < 200$ GeV, $m(1^{st} \text{ gen. } \tilde{q})=m(2^{nd} \text{ gen. } \tilde{q})$	ATLAS-CONF-2016-078	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}$ (compressed)	mono-jet	1-3 jets	Yes	3.2	608 GeV		$m(\tilde{g})-m(\tilde{Z}) < 5$ GeV	1604.07773	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{Z}$	0	2-6 jets	Yes	13.3			$m(\tilde{Z})=0$ GeV	ATLAS-CONF-2016-078	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{Z}$	0	2-6 jets	Yes	13.3			$m(\tilde{Z}) < 400$ GeV, $m(\tilde{Z})=0.5(m(\tilde{Z})+m(\tilde{Z}))$	ATLAS-CONF-2016-078	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}(\ell\ell/\nu\nu)\tilde{Z}$	3 e, μ	4 jets	-	13.2			$m(\tilde{Z}) < 400$ GeV	ATLAS-CONF-2016-037	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}WZ\tilde{Z}$	2 e, μ (SS)	0-3 jets	Yes	13.2			$m(\tilde{Z}) < 500$ GeV	ATLAS-CONF-2016-037	
	GMSB (\tilde{Z} NLSP)	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	3.2			2.0 TeV	1607.05979	
	GGM (bino NLSP)	2 γ	-	Yes	3.2			1.65 TeV	1606.09150	
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	20.3			1.37 TeV	1507.05493	
	GGM (higgsino-bino NLSP)	γ	2 jets	Yes	13.3			1.8 TeV	ATLAS-CONF-2016-066	
	GGM (higgsino NLSP)	2 e, μ (Z)	2 jets	Yes	20.3			900 GeV	1503.03290	
Gravitino LSP	0	mono-jet	Yes	20.3			865 GeV	1502.01518		
3 rd gen. \tilde{g} med.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\bar{b}\tilde{Z}$	0	3 b	Yes	14.8			1.89 TeV	$m(\tilde{Z})=0$ GeV	ATLAS-CONF-2016-052
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}\tilde{Z}$	0-1 e, μ	3 b	Yes	14.8			1.89 TeV	$m(\tilde{Z})=0$ GeV	ATLAS-CONF-2016-052
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\bar{b}\tilde{Z}$	0-1 e, μ	3 b	Yes	20.1			1.37 TeV	$m(\tilde{Z}) < 300$ GeV	1407.06900
3 rd gen. squarks direct production	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\bar{b}\tilde{Z}$	0	2 b	Yes	3.2			840 GeV	$m(\tilde{Z}) < 100$ GeV	1606.06772
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\bar{t}\tilde{Z}$	2 e, μ (SS)	1 b	Yes	13.2			325-685 GeV	$m(\tilde{Z}) < 150$ GeV, $m(\tilde{Z}) = m(\tilde{Z}) + 100$ GeV	ATLAS-CONF-2016-037
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\bar{b}\tilde{Z}$	0-2 e, μ	1-2 b	Yes	4.7/13.3			200-720 GeV	$m(\tilde{Z}) = 2m(\tilde{Z}), m(\tilde{Z}) = 55$ GeV	1209.2102, ATLAS-CONF-2016-077
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{Z}$ or $\tilde{Z}\tilde{Z}$	0-2 e, μ	0-2 jets/1-2 b	Yes	4.7/13.3			90-198 GeV	$m(\tilde{Z})=1$ GeV	1506.08616, ATLAS-CONF-2016-077
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow t\bar{t}\tilde{Z}$	0	mono-jet	Yes	3.2			90-323 GeV	$m(\tilde{Z})-m(\tilde{Z})=5$ GeV	1604.07773
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.3			150-600 GeV	$m(\tilde{Z}) > 150$ GeV	1403.5222
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow t\bar{t} + Z$	3 e, μ (Z)	1 b	Yes	13.3			290-700 GeV	$m(\tilde{Z}) < 300$ GeV	ATLAS-CONF-2016-038
$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow t\bar{t} + h$	1 e, μ	6 jets + 2 b	Yes	20.3			320-620 GeV	$m(\tilde{Z})=0$ GeV	1506.08616	
EW direct	$\tilde{\chi}_1^0\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \ell\bar{\ell}\tilde{Z}$	2 e, μ	0	Yes	20.3			90-335 GeV	$m(\tilde{Z})=0$ GeV	1403.5294
	$\tilde{\chi}_1^0\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \ell\nu(\ell\nu)$	2 e, μ	0	Yes	20.3			140-475 GeV	$m(\tilde{Z})=0$ GeV, $m(\tilde{Z}, \nu)=0.5(m(\tilde{Z})+m(\tilde{Z}))$	1403.5294
	$\tilde{\chi}_1^0\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\nu(\tau\nu)$	2 τ	-	Yes	20.3			355 GeV	$m(\tilde{Z})=0$ GeV, $m(\tilde{Z}, \nu)=0.5(m(\tilde{Z})+m(\tilde{Z}))$	1407.0350
	$\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow \tilde{\chi}_1^0\tilde{\chi}_1^0, \ell\nu\tilde{Z}, \ell(\nu\nu)$	3 e, μ	0	Yes	20.3			715 GeV	$m(\tilde{Z})=m(\tilde{Z}), m(\tilde{Z})=0, m(\tilde{Z}, \nu)=0.5(m(\tilde{Z})+m(\tilde{Z}))$	1402.7029
	$\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow W\tilde{Z}$	2-3 e, μ	0-2 jets	Yes	20.3			425 GeV	$m(\tilde{Z})=m(\tilde{Z}), m(\tilde{Z})=0, \tilde{Z}$ decoupled	1403.5294, 1402.7029
	$\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow W\tilde{Z}, h\tilde{Z}, h\tilde{Z} \rightarrow WW/\tau\tau/\gamma\gamma$	e, μ, γ	0-2 b	Yes	20.3			270 GeV	$m(\tilde{Z})=m(\tilde{Z}), m(\tilde{Z})=0, \tilde{Z}$ decoupled	1501.07110
	$\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow \tilde{Z}\tilde{Z}$	4 e, μ	0	Yes	20.3			635 GeV	$m(\tilde{Z})=m(\tilde{Z}), m(\tilde{Z})=0, \tilde{Z}$ decoupled	1405.5086
	GGM (wino NLSP) weak prod.	1 $e, \mu + \gamma$	-	Yes	20.3			115-370 GeV	$m(\tilde{Z})=0, m(\tilde{Z}, \nu)=0.5(m(\tilde{Z})+m(\tilde{Z}))$	1507.05493
	GGM (bino NLSP) weak prod.	2 γ	-	Yes	20.3			590 GeV	$c\tau < 1$ mm	1507.05493
	Long-lived particles	Direct $\tilde{\chi}_1^0\tilde{\chi}_1^0$ prod., long-lived $\tilde{\chi}_1^0$	Disapp. trk	1 jet	Yes	20.3			270 GeV	$m(\tilde{Z})-m(\tilde{Z})=160$ MeV, $\tau(\tilde{\chi}_1^0)=0.2$ ns
Direct $\tilde{\chi}_1^0\tilde{\chi}_1^0$ prod., long-lived $\tilde{\chi}_1^0$		dE/dx trk	-	Yes	18.4			495 GeV	$m(\tilde{Z})+m(\tilde{Z})=160$ MeV, $\tau(\tilde{\chi}_1^0) < 15$ ns	1506.05332
Stable, stopped \tilde{g} R-hadron		0	1-5 jets	Yes	27.9			850 GeV	$m(\tilde{Z})=100$ GeV, $10 \mu\text{s} < \tau(\tilde{g}) < 1000$ s	1310.8584
Stable \tilde{g} R-hadron		trk	-	-	3.2			1.58 TeV	$c\tau < 1$ mm	1606.05129
Metastable \tilde{g} R-hadron		dE/dx trk	-	-	3.2			1.57 TeV	$m(\tilde{Z})=100$ GeV, $\tau > 10$ ns	1604.04520
GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{Z}, \tilde{\mu})\tau(e, \mu)$		1-2 γ	-	-	19.1			537 GeV	$10 < \text{ctan}\beta < 50$	1411.6795
GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{Z}$, long-lived $\tilde{\chi}_1^0$		2 γ	-	Yes	20.3			440 GeV	$1 < \tau(\tilde{\chi}_1^0) < 3$ ns, SPS8 model	1409.5542
$\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow e\bar{e}\nu/\mu\bar{\mu}\nu$		displ. ee/ $\mu\bar{\mu}$	-	-	20.3			1.0 TeV	$7 < c\tau(\tilde{\chi}_1^0) < 740$ mm, $m(\tilde{g})=1.3$ TeV	1504.05162
GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow ZG$		displ. vtx + jets	-	-	20.3			1.0 TeV	$6 < c\tau(\tilde{\chi}_1^0) < 480$ mm, $m(\tilde{g})=1.1$ TeV	1504.05162
RPV		LFV $pp \rightarrow \tilde{\nu}_i + X, \tilde{\nu}_i \rightarrow e\mu/\tau\nu/\mu\tau$	$e\mu, \tau\nu, \mu\tau$	-	-	3.2			1.9 TeV	$\tilde{\chi}_{1,2}^0 = -0.11, \tilde{\chi}_{3,2}(1/2/3)=0.07$
	Billinear RPV CMSSM	2 e, μ (SS)	0-3 b	Yes	20.3			1.45 TeV	$m(\tilde{g})=m(\tilde{Z}), c\tau_{LSP} < 1$ mm	1404.2500
	$\tilde{\chi}_1^0\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow W\tilde{Z}, \tilde{\chi}_1^0 \rightarrow e\bar{e}\nu, \mu\bar{\mu}\nu$	4 e, μ	-	Yes	13.3			1.14 TeV	$m(\tilde{Z}) > 400$ GeV, $A_{1,2} \neq 0$ ($k=1, 2$)	ATLAS-CONF-2016-075
	$\tilde{\chi}_1^0\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow W\tilde{Z}, \tilde{\chi}_1^0 \rightarrow \tau\nu_e, e\nu_\tau$	3 $e, \mu + \tau$	-	Yes	20.3			450 GeV	$m(\tilde{Z}) > 0.2m(\tilde{Z}), A_{1,2} \neq 0$	1405.5086
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}g$	0	4-5 large- R jets	-	14.8			1.08 TeV	$BR(\tilde{g})-BR(\tilde{g})=BR(\tilde{Z})=0\%$	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{Z}, \tilde{Z} \rightarrow q\bar{q}g$	0	4-5 large- R jets	-	14.8			1.55 TeV	$m(\tilde{Z})=800$ GeV	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\bar{t}, \tilde{t}_1 \rightarrow b\bar{s}$	2 e, μ (SS)	0-3 b	Yes	13.2			1.3 TeV	$m(\tilde{Z}) < 750$ GeV	ATLAS-CONF-2016-037
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\bar{s}$	0	2 jets + 2 b	-	15.4			410 GeV	$m(\tilde{Z}) < 480$ GeV	ATLAS-CONF-2016-022, ATLAS-CONF-2016-084
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\bar{t}$	2 e, μ	2 b	-	20.3			0.4-1.0 TeV	$BR(\tilde{t}_1 \rightarrow b\bar{e}/\mu) > 20\%$	ATLAS-CONF-2015-015	
Other	Scalar charm, $\tilde{Z} \rightarrow c\bar{c}\tilde{Z}$	0	2 c	Yes	20.3			510 GeV	$m(\tilde{Z}) < 200$ GeV	1501.01325

*Only a selection of the available mass limits on new states or phenomena is shown.

10⁻¹

1

Mass scale [TeV]

Searches for various resonances

Simplified models

Vector like top partners

Vector-like quarks - spin 1/2 particles with the same colour (triplet) and electroweak quantum numbers for left and right components

Masses not from the BEH mechanism

$\bar{Q}_L Q_R$ mass terms are allowed by EW gauge symmetry

VLQ appear in many BSM extensions

Matsedonskyia, Panicob, Wulzer

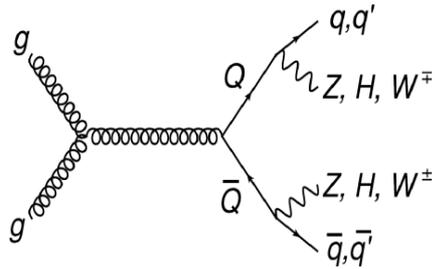
partner (MG name)	Q	W^\pm	Z	h	$W^\pm W^\pm$
$T_{2/3}$ (T23)	2/3	c_L^{TW}, c_R^{TW}	c_L^{TZ}, c_R^{TZ}	c_L^{Th}, c_R^{Th}	—
$B_{1/3}$ (B13)	-1/3	c_L^{BW}, c_R^{BW}	c_L^{BZ}, c_R^{BZ}	c_L^{Bh}, c_R^{Bh}	—
$X_{5/3}$ (X53)	5/3	c_L^{XW}, c_R^{XW}	—	—	—
$Y_{4/3}$ (Y43)	-4/3	c_L^{YW}, c_R^{YW}	—	—	—
$V_{8/3}$ (V83)	8/3	—	—	—	c_L^{VW}, c_R^{VW}

Example of the simplified model Lagrangian
(after mixing and mass matrix diagonalization)

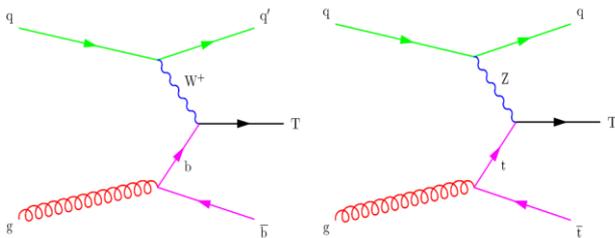
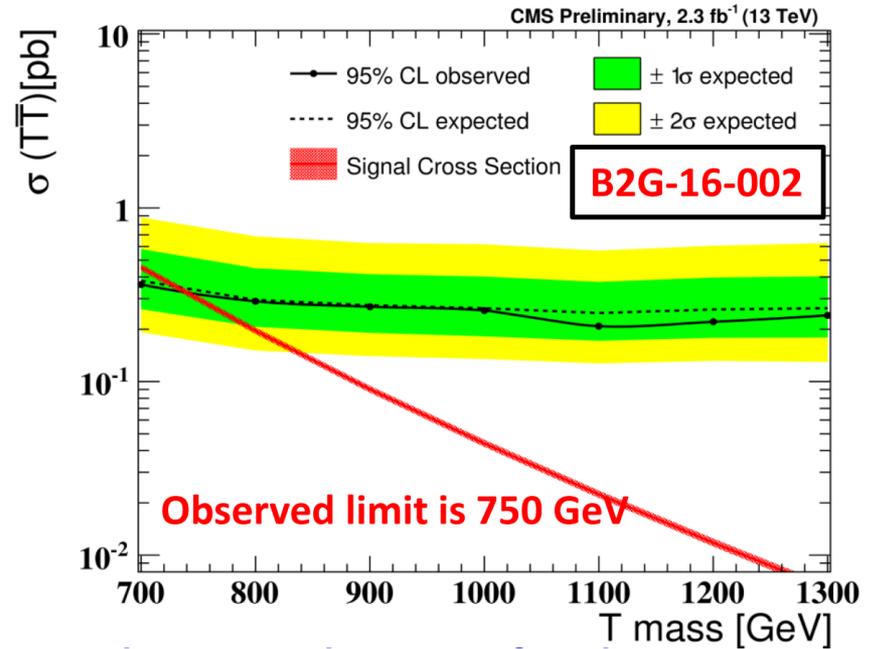
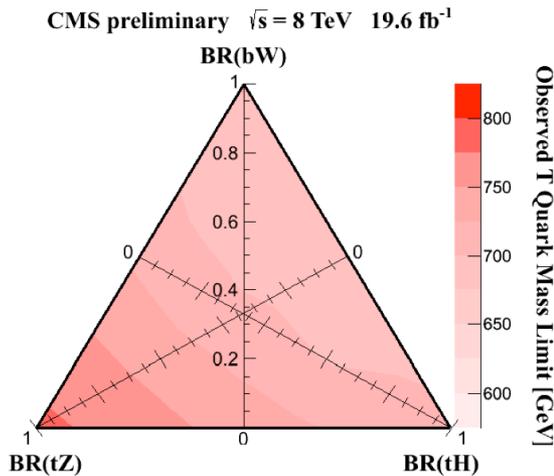
$$\frac{g_w}{2} c_L^{TW} \left[\bar{T}_L \gamma_\mu b_L W^\mu \right] + h.c.$$

$$c_L^{Th} \left[\bar{T}_R t_L h \right] + h.c.$$

Similar limit at 13 TeV with just 2.3 fb⁻¹ integrated luminosity



$T \rightarrow Zt, T \rightarrow Wb, T \rightarrow Ht$ for $Q_T=2/3$

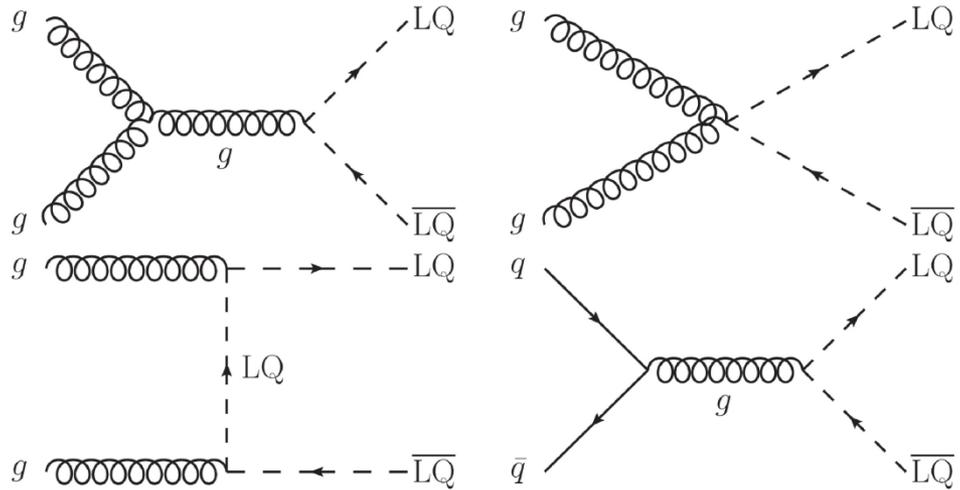


Also some limits are set on couplings and masses from single VLQ production

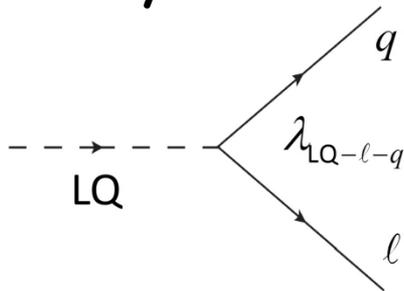
Leptoquark searches

LQs are predicted by composite models, GUT ...

Production channels



Decays



Final states for leptoquarks of three generations

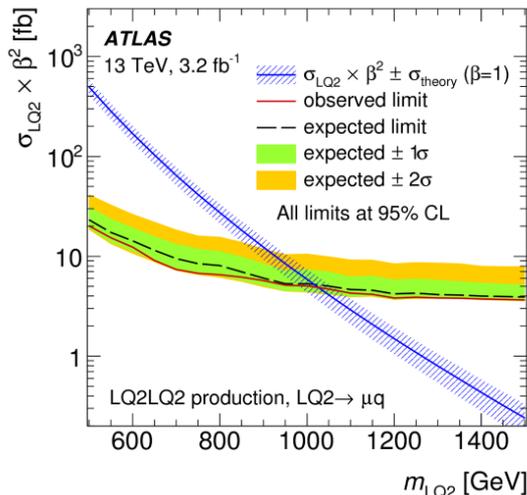
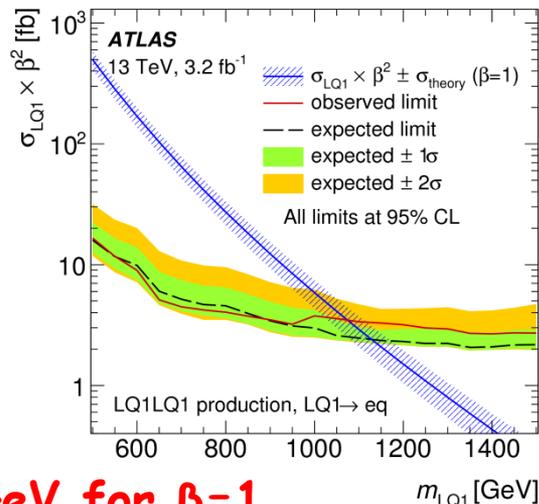
LQ1 \rightarrow eu, ed, ν_e u, ν_e d

LQ2 \rightarrow μ c, μ s, ν_μ c, ν_μ s

LQ3 \rightarrow τ t, τ b, ν_τ t, ν_τ b

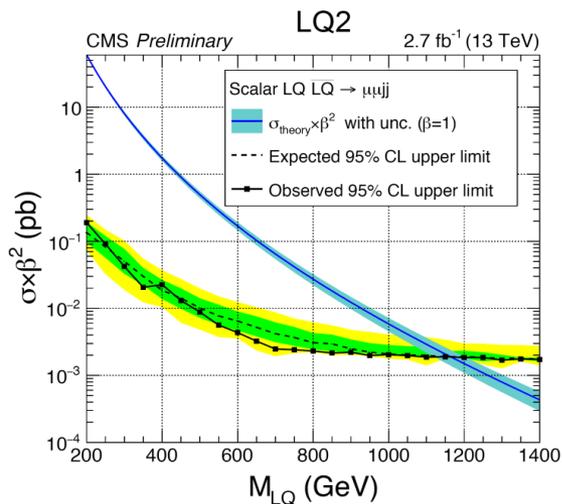
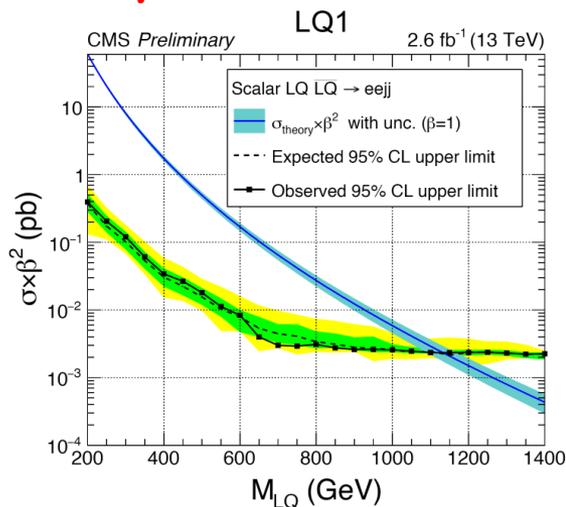
1st generation LQ mass limit for $\beta=1$
 2nd generation LQ mass limit for $\beta=1$

RUN1 • CMS: 1010 GeV; ATLAS 1060 GeV
 RUN1 • CMS: 1080 GeV; ATLAS 1050 GeV



1130 GeV for $\beta=1$

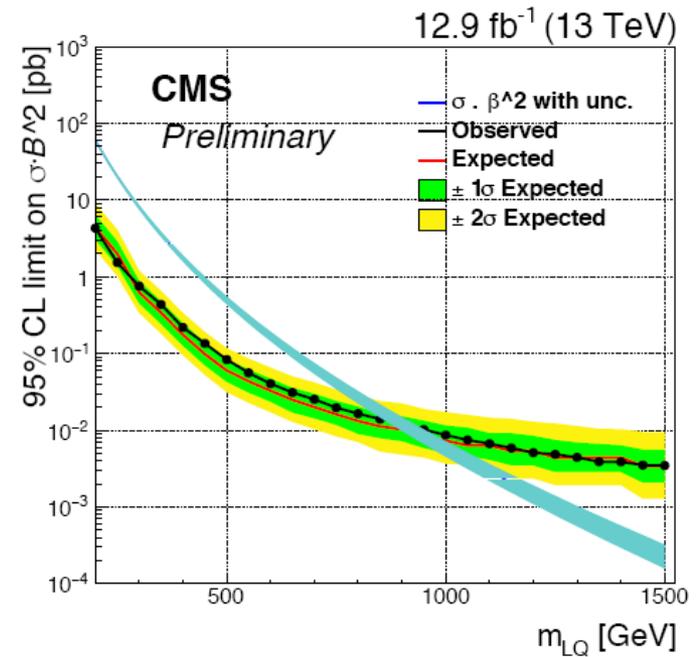
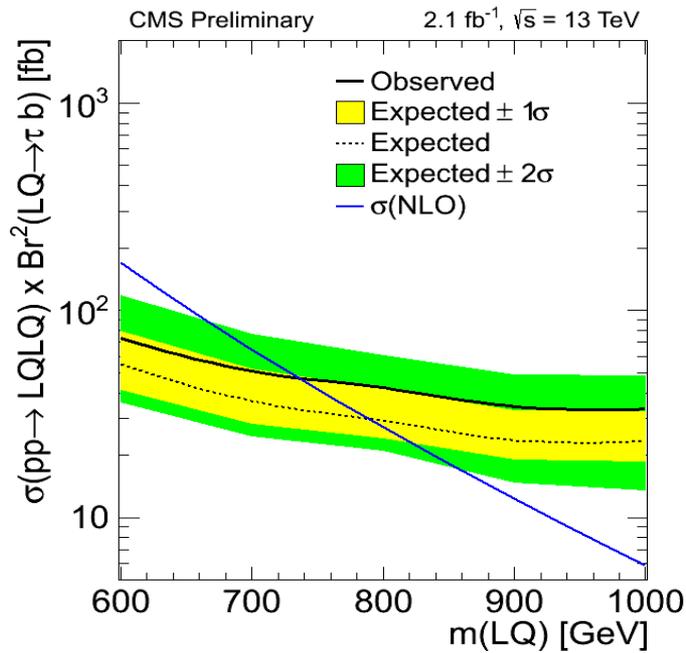
1165 GeV for $\beta=1$



Sensitivity is similar for $\sim 20 \text{ fb}^{-1}$ at $\sqrt{s} = 8 \text{ TeV}$ and for $\sim 3 \text{ fb}^{-1}$ at $\sqrt{s} = 13 \text{ TeV}$

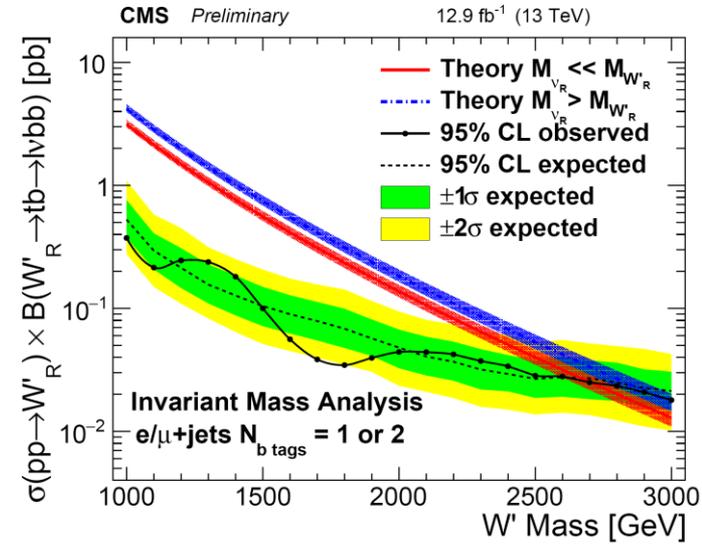
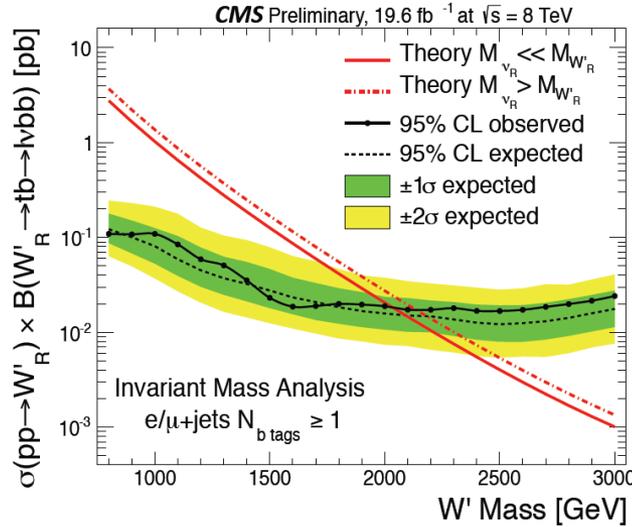
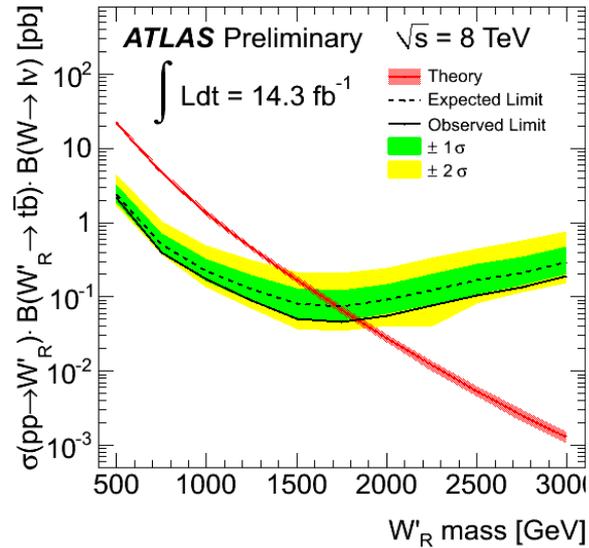
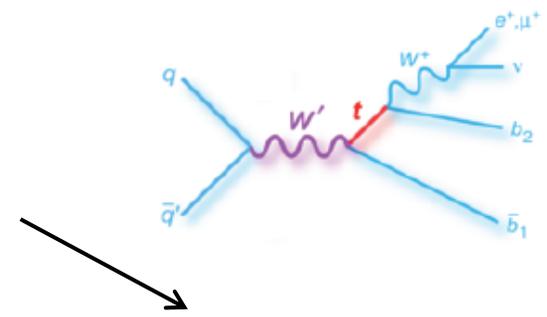
Limits on 3d generation leptoquarks

$$\beta = \text{BR}(\text{LQ} \rightarrow lq)$$

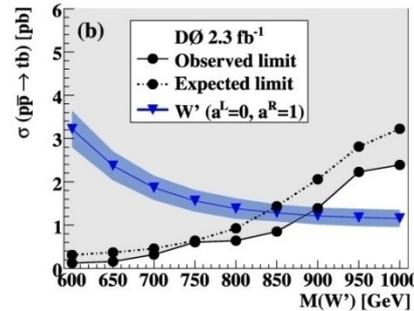
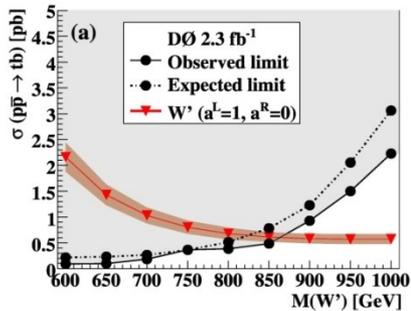


Limit: $M(\text{LQ3}) > \sim 900 \text{ GeV}$

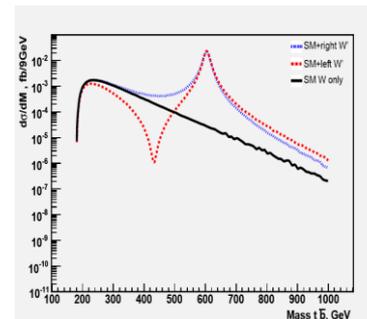
Searches for W' in top+b



D0 limits: $M_{W'} > 830$ (860) GeV L(R)

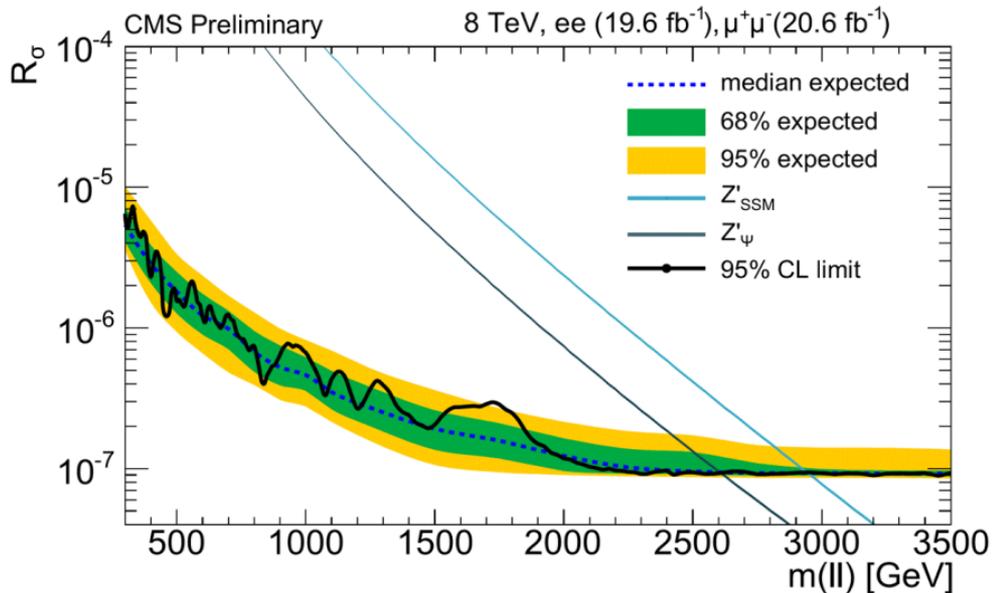


E.B., Bunichev, Dudko, Perfilov

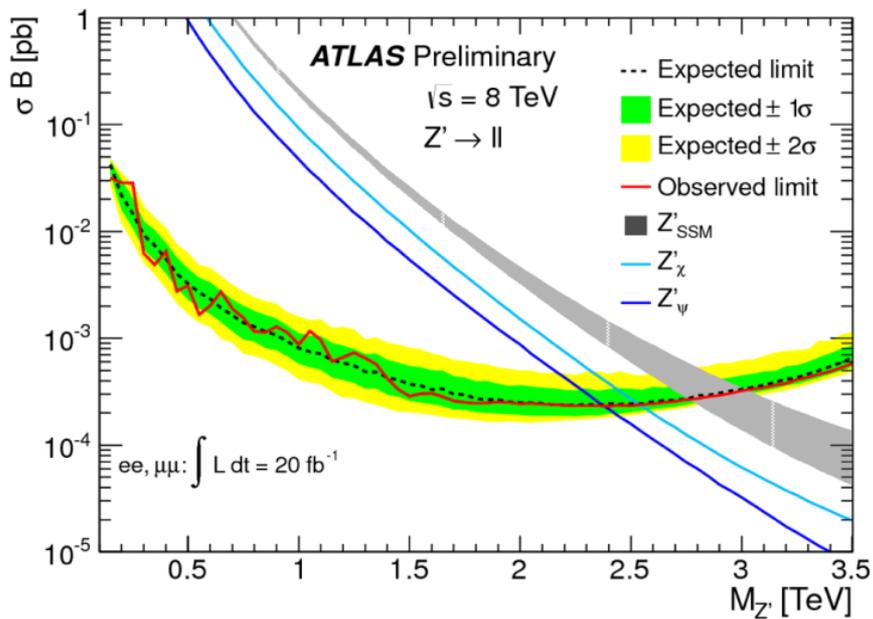


Negative interference

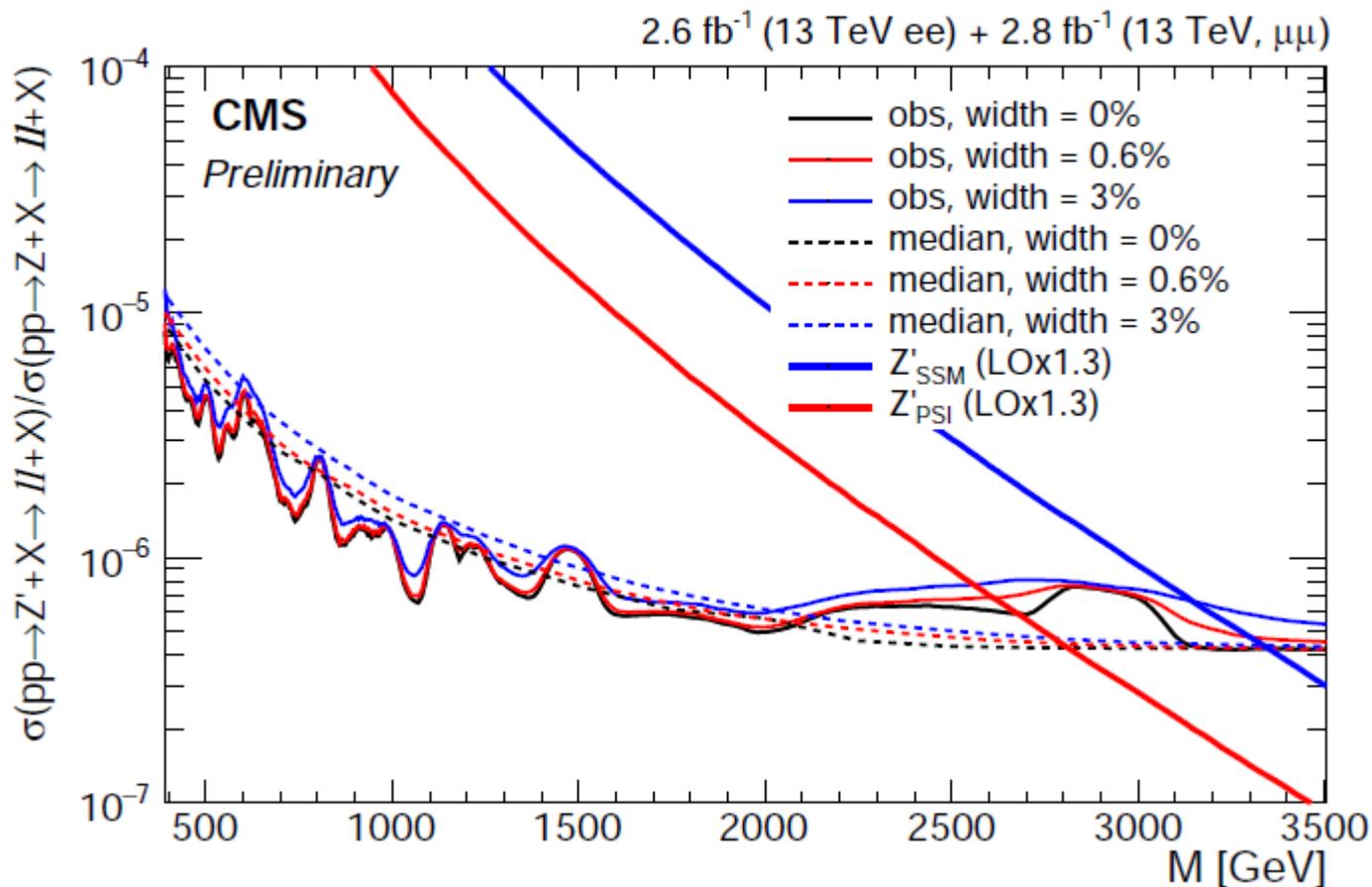
Searches for Z' in dileptons



$$R_\sigma = \frac{\sigma(\text{pp} \rightarrow Z' + X \rightarrow \ell\ell + X)}{\sigma(\text{pp} \rightarrow Z + X \rightarrow \ell\ell + X)}$$

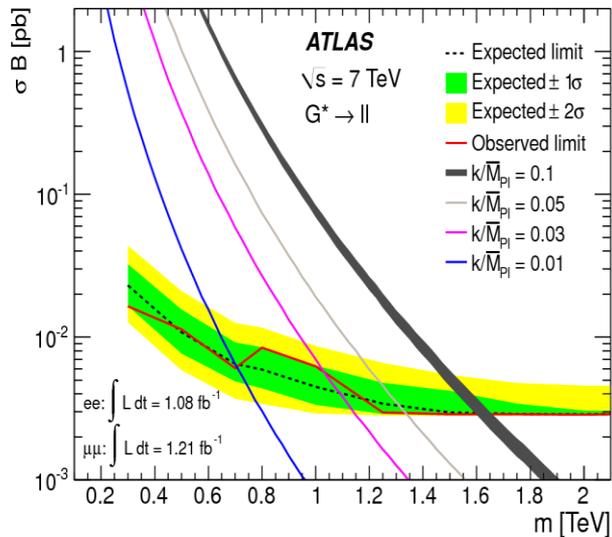
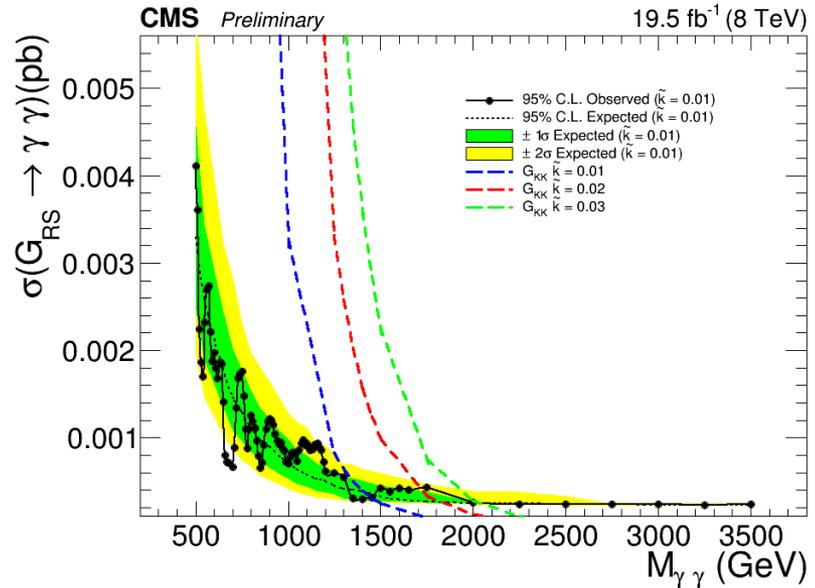
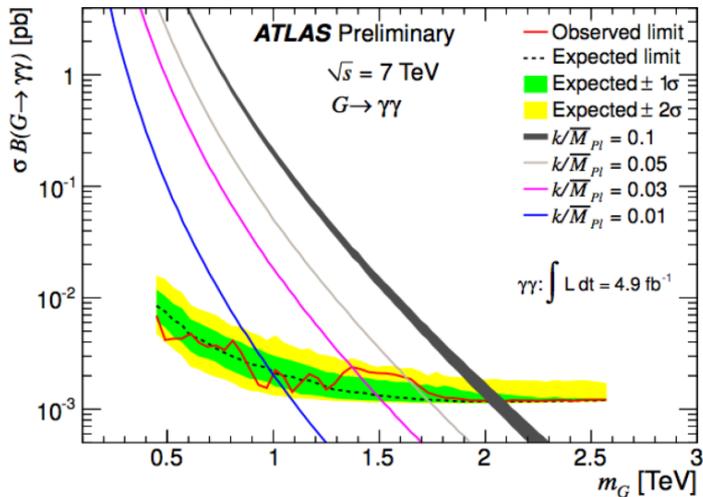


M(Z' _{SSM})	expected	observed
CMS	> 2.96 TeV	> 2.96 TeV
ATLAS	> 2.85 TeV	> 2.86 TeV



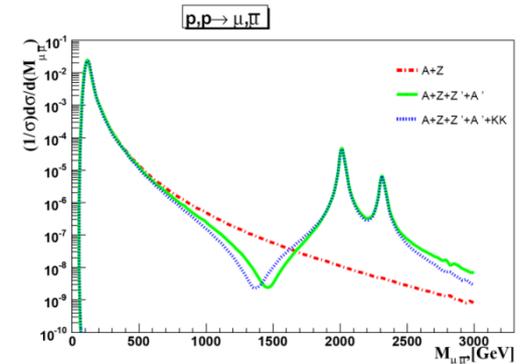
channel	Z'_ψ		Z'_{SSM}	
	obs (TeV)	expected (TeV)	obs (TeV)	expected (TeV)
ee	2.40	2.45	2.75	2.95
$\mu^+ \mu^-$	2.40	2.55	3.00	3.05
$ee + \mu^+ \mu^-$	2.60	2.80	3.15	3.35

Searches for RS gravitons



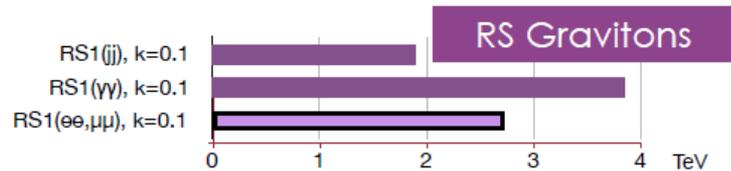
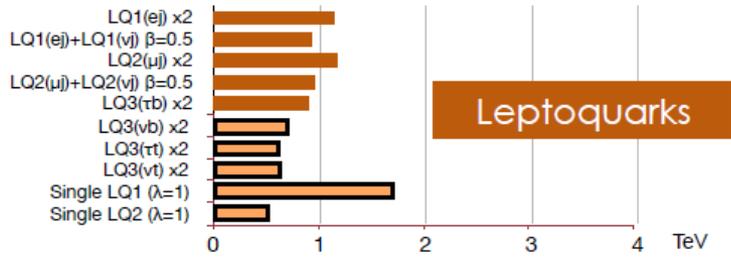
E.B., Bunichev, Smolyakov, Volobuev

Interferences

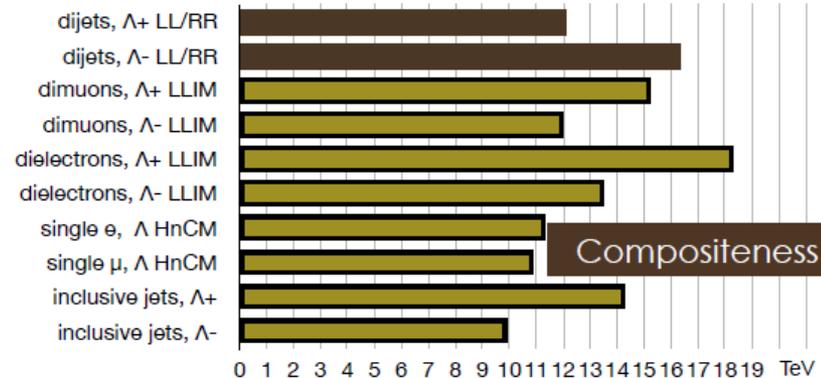
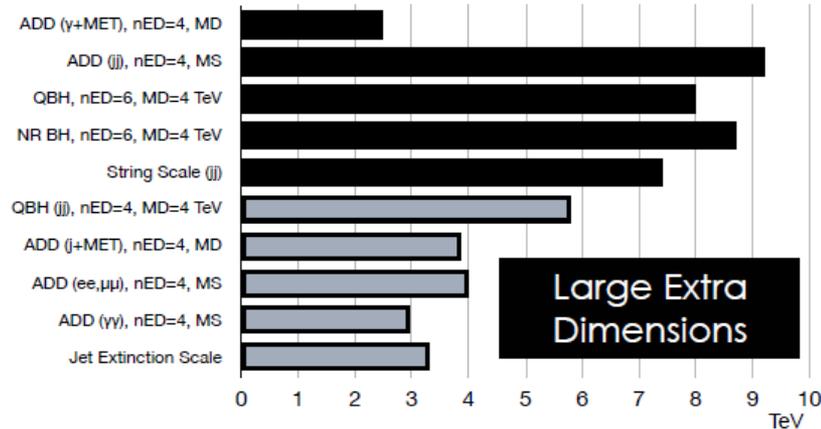
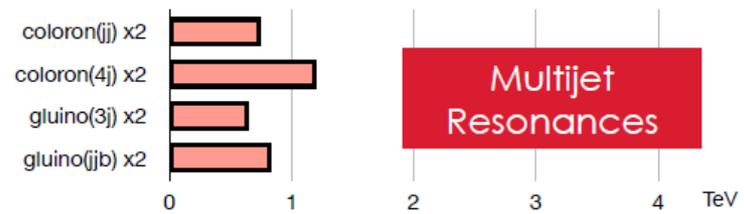
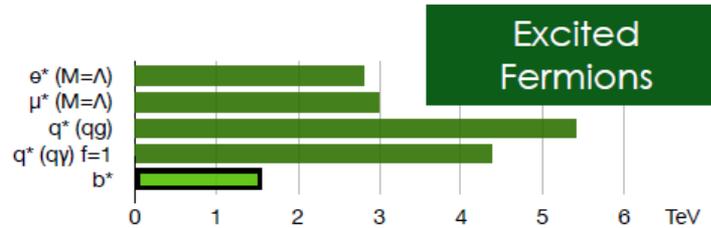
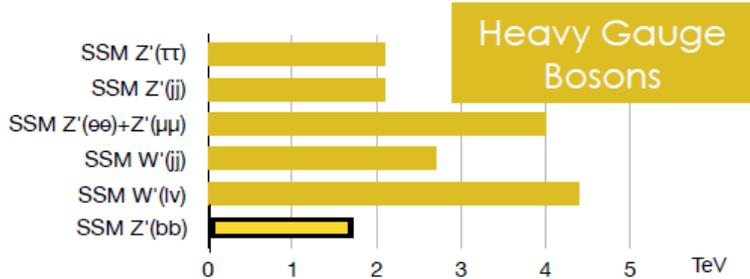


Excluding Dark Matter and Long Lived particles searches

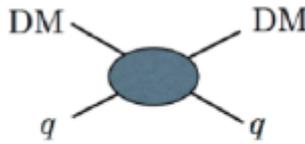
13 TeV 8 TeV (Similar limits from ATLAS)



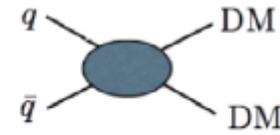
CMS Preliminary



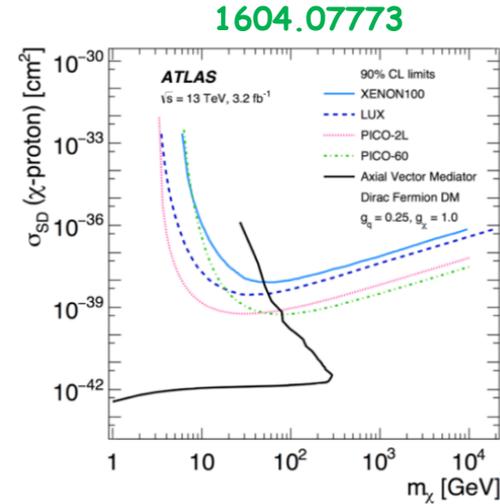
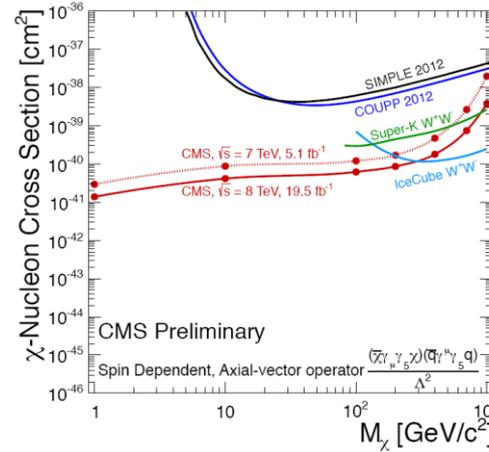
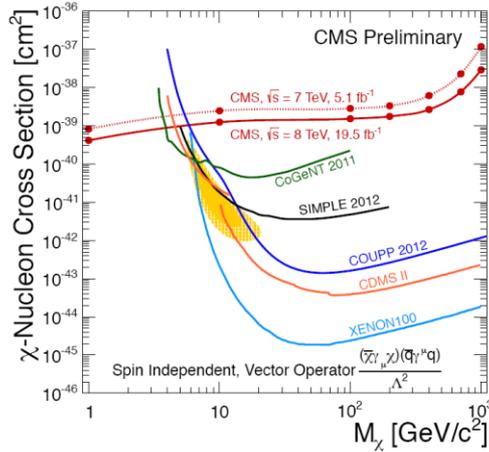
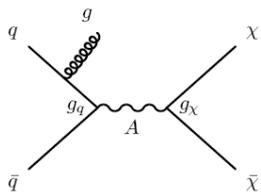
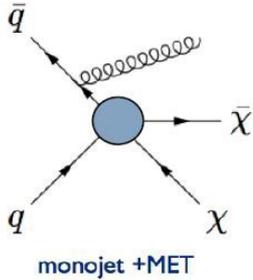
Dark Matter searches



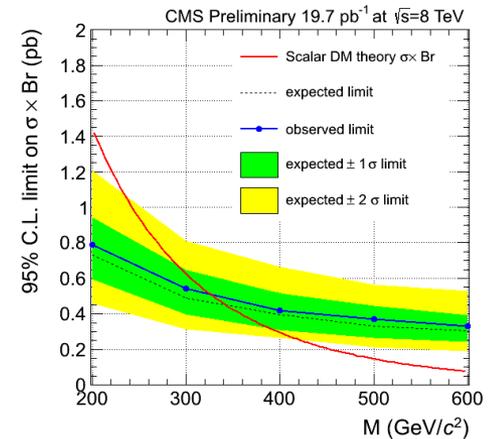
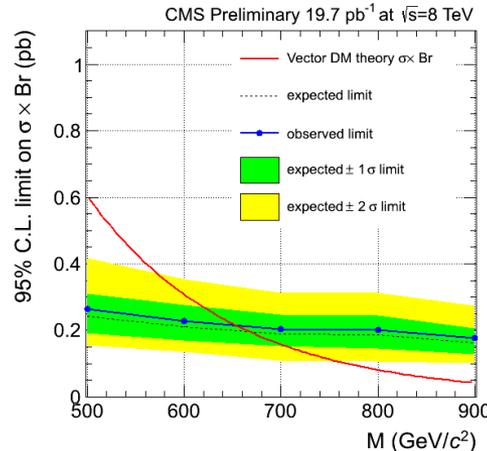
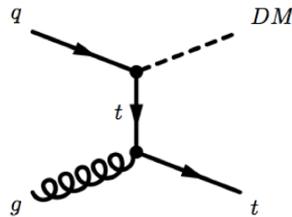
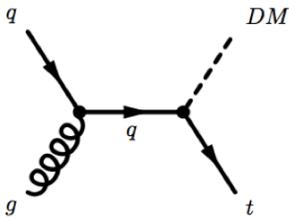
Direct Detection (t-channel)



Collider Searches (s-channel)



Dark Matter searches in association with single top



Mass of vector (scalar) DM candidate less than 655 (327) is excluded at 95% C.L.

CMS Preliminary

Dark Matter Summary - ICHEP 2016

DM + jets/ $V(q\bar{q})$
 $g_{DM}=1, g_q=0.25$

DM + γ
 $g_{DM}=1, g_q=0.25$

DM + $Z(\ell^+\ell^-)$
 $g_{DM}=1, g_q=0.25$

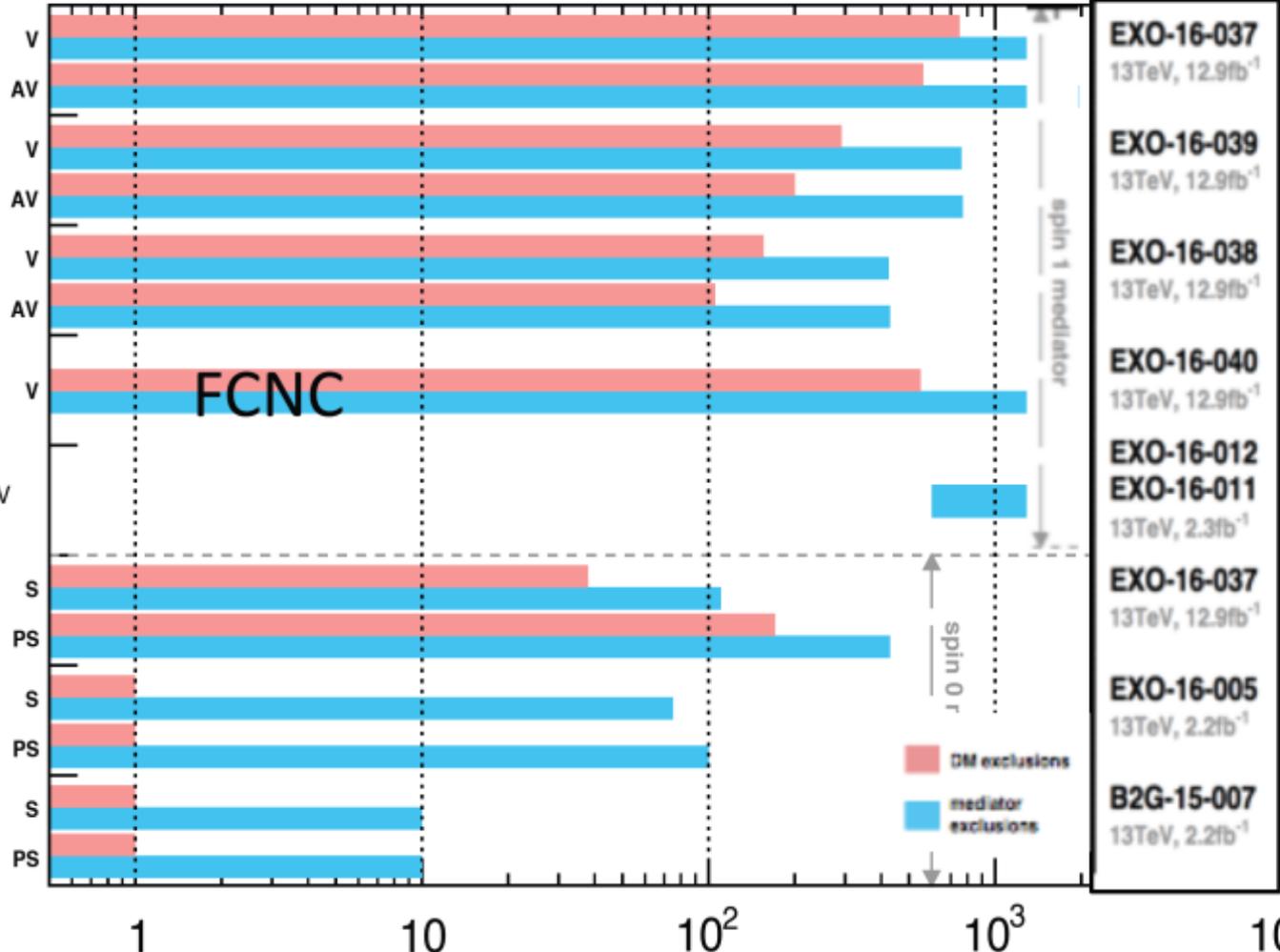
DM + t
 $g_{DM}=1, a_{FC}=b_{FC}=0.25$

DM + $H(bb/\gamma\gamma)$
 $m_{A_0}=300\text{GeV}; m_{DM}=100\text{GeV}$
 $g_Z=0.8$

DM + jets/ $V(q\bar{q})$
 $g_{DM}=g_q=1$

DM + $t\bar{t}$
 $g_{DM}=g_q=1$ $\alpha/\sigma_0 = 2$

DM + $b\bar{b}/t\bar{t}$
 $g_{DM}=g_q=1$ $\alpha/\sigma_0 = 5$
 $\alpha/\sigma_0 = 30$



m_{DM} [GeV]

Observed limits at 95%CL
 for considered simplified models
 Theory uncertainties not included

V = vector ; AV = axial-vector
 S = scalar ; PS = pseudoscalar

Maximal excluded mass [GeV]

EXO-16-037
 13TeV, 12.9fb⁻¹

EXO-16-039
 13TeV, 12.9fb⁻¹

EXO-16-038
 13TeV, 12.9fb⁻¹

EXO-16-040
 13TeV, 12.9fb⁻¹

EXO-16-012
 EXO-16-011
 13TeV, 2.3fb⁻¹

EXO-16-037
 13TeV, 12.9fb⁻¹

EXO-16-005
 13TeV, 2.2fb⁻¹

B2G-15-007
 13TeV, 2.2fb⁻¹

Searches below threshold

Effective field theory approach

$$\mathcal{L} = \mathcal{L}_{SM} + \sum \frac{c_i}{\Lambda^2} \mathcal{O}_i + \dots$$

c_i - dimensionless coefficients

\mathcal{O}_i - operators constructed from SM fields preserving SM gauge invariance

Anomalous Top Couplings

The top quark interactions of dimension 4:

$$\begin{aligned}\mathcal{L}_4 = & -g_s \bar{t} \gamma^\mu T^a t G_\mu^a - \frac{g}{\sqrt{2}} \sum_{q=d,s,b} \bar{t} \gamma^\mu (v_{tq}^W - a_{tq}^W \gamma_5) q W_\mu^+ \\ & - \frac{2}{3} e \bar{t} \gamma^\mu t A_\mu - \frac{g}{2 \cos \theta_W} \sum_{q=u,c,t} \bar{t} \gamma^\mu (v_{tq}^Z - a_{tq}^Z \gamma_5) q Z_\mu\end{aligned}$$

The dimension 5 couplings have the generic form:

$$\begin{aligned}\mathcal{L}_5 = & -g_s \sum_{q=u,c,t} \frac{\kappa_{tq}^g}{\Lambda} \bar{t} \sigma^{\mu\nu} T^a (f_{tq}^g + i h_{tq}^g \gamma_5) q G_{\mu\nu}^a - \frac{g}{\sqrt{2}} \sum_{q=d,s,b} \frac{\kappa_{tq}^W}{\Lambda} \bar{t} \sigma^{\mu\nu} (f_{tq}^W + i h_{tq}^W \gamma_5) q W_{\mu\nu}^+ \\ & - e \sum_{q=u,c,t} \frac{\kappa_{tq}^\gamma}{\Lambda} \bar{t} \sigma^{\mu\nu} (f_{tq}^\gamma + i h_{tq}^\gamma \gamma_5) q A_{\mu\nu} - \frac{g}{2 \cos \theta_W} \sum_{q=u,c,t} \frac{\kappa_{tq}^Z}{\Lambda} \bar{t} \sigma^{\mu\nu} (f_{tq}^Z + i h_{tq}^Z \gamma_5) q Z_{\mu\nu}\end{aligned}$$

Natural size $k \sim v/\Lambda$

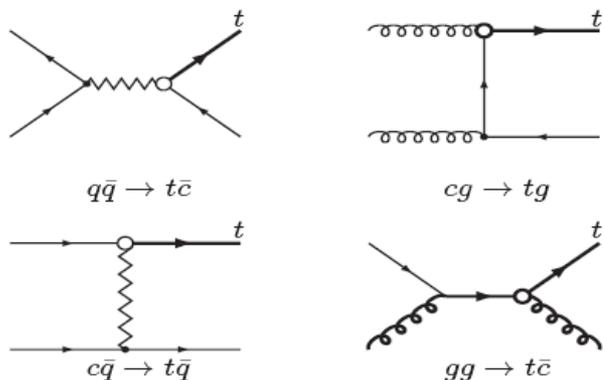
Present constrains come from

- Low energy data via loop contributions
 $K_L \rightarrow \mu^+ \mu^-$, $K_L - K_S$ mass difference, $b \rightarrow l^+ l^- X$, $b \rightarrow s \gamma$
- LEP2
- Tevatron Run1,2
- HERA
- Unitarity violation bounds

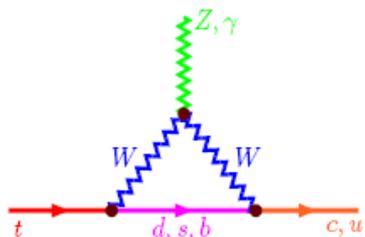
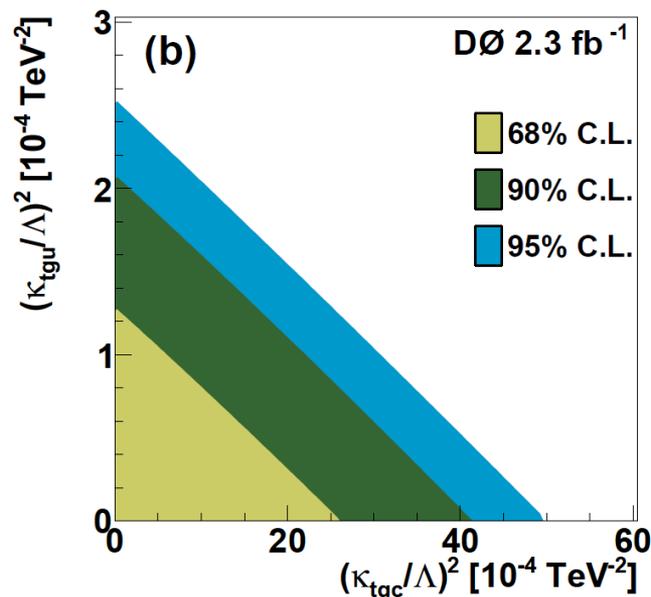
FCNC anomalous top couplings

- Couplings: tqg , $tq\gamma$, tqZ , where $q = u, c$

$$\Delta\mathcal{L}^{eff} = \frac{1}{\Lambda} [\kappa_{tq}^{\gamma,Z} e\bar{t}\sigma_{\mu\nu}qF_{\gamma,Z}^{\mu\nu} + \kappa_{tq}^g g_s\bar{t}\sigma_{\mu\nu}\frac{\lambda^i}{2}qG^{i\mu\nu}] + h.c.$$



W' boson and FCNC MC event samples from SingleTop (CompHEP) generator



	SM	two-Higgs	SUSY
BR($t \rightarrow cg$)	$5 \cdot 10^{-11}$	10^{-6}	10^{-3}
BR($t \rightarrow c\gamma$)	$5 \cdot 10^{-13}$	10^{-6}	10^{-5}
BR($t \rightarrow cZ$)	$\sim 10^{-13}$	10^{-9}	10^{-4}

FCNC decays are highly suppressed in SM $t \rightarrow qg$, $t \rightarrow q\gamma$, $t \rightarrow qZ$

To compare FCNC limits from top decays and top production one can express limits on FCNC couplings in term of Br fractions

$$\Gamma(t \rightarrow qg) = \left(\frac{\kappa_{tq}^g}{\Lambda}\right)^2 \frac{8}{3} \alpha_s m_t^3, \quad \Gamma(t \rightarrow q\gamma) = \left(\frac{\kappa_{tq}^\gamma}{\Lambda}\right)^2 2\alpha m_t^3,$$

$$\Gamma(t \rightarrow qZ)_\gamma = (|v_{tq}^Z|^2 + |a_{tq}^Z|^2) \alpha m_t^3 \frac{1}{4M_Z^2 \sin^2 2\theta_W} \left(1 - \frac{M_Z^2}{m_t^2}\right)^2 \left(1 + 2\frac{M_Z^2}{m_t^2}\right),$$

$$\Gamma(t \rightarrow qZ)_\sigma = \left(\frac{\kappa_{tq}^Z}{\Lambda}\right)^2 \alpha m_t^3 \frac{1}{\sin^2 2\theta_W} \left(1 - \frac{M_Z^2}{m_t^2}\right)^2 \left(2 + \frac{M_Z^2}{m_t^2}\right)$$

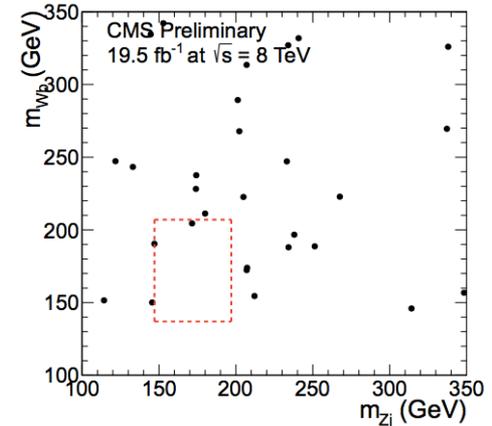
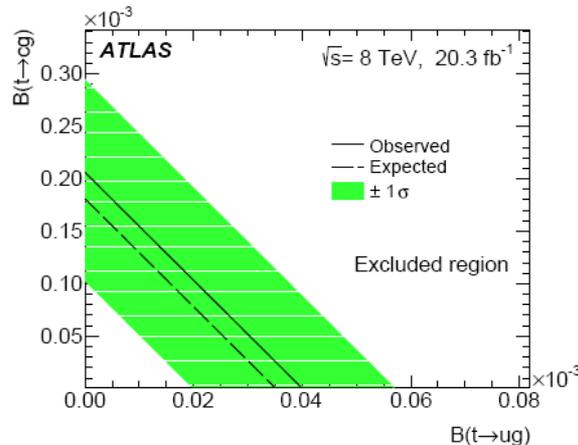
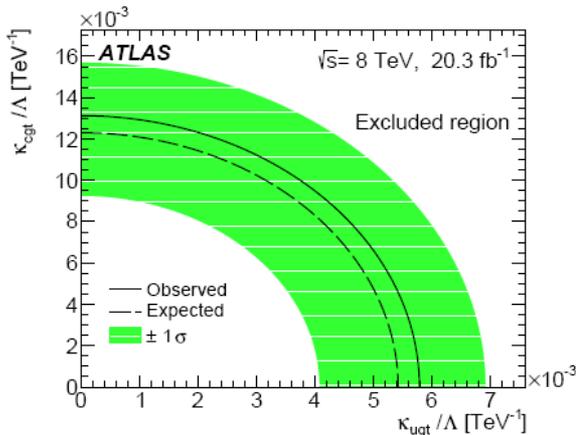
DO:

	<i>tgu</i>	<i>tgc</i>
Cross section	0.20 pb	0.27 pb
κ_{tgf}/Λ	0.013 TeV ⁻¹	0.057 TeV ⁻¹
$\mathcal{B}(t \rightarrow qg)$	2.0×10^{-4}	3.9×10^{-3}

CDF:

$$\mathcal{B}(t \rightarrow u + g) < 3.9 \cdot 10^{-4}$$

$$\mathcal{B}(t \rightarrow c + g) < 5.7 \cdot 10^{-3}$$



CMS limit:

$$\mathcal{B}(t \rightarrow Zq) < 0.07\% \text{ @ } 95\% \text{ C.L.}$$

Many interesting new results

Indirect search for BSM physics - the main goal of the LHCb experiment.

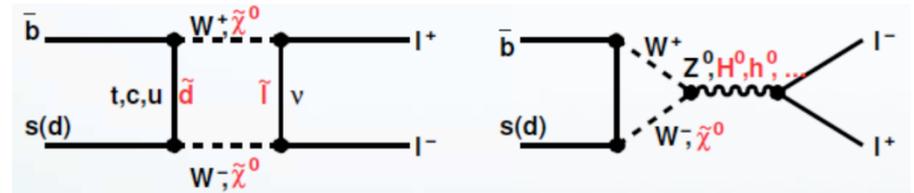
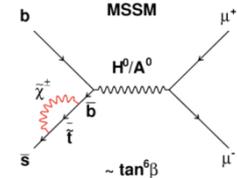
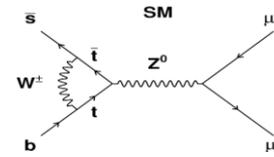
Rare $B_s^0 \rightarrow \mu^+ \mu^-$ decay

LHCb:

$$Br(B_s^0 \rightarrow \mu\mu) = (3.2_{-1.2}^{+1.4}(\text{stat})_{-0.3}^{+0.5}(\text{syst})) \times 10^{-9}$$

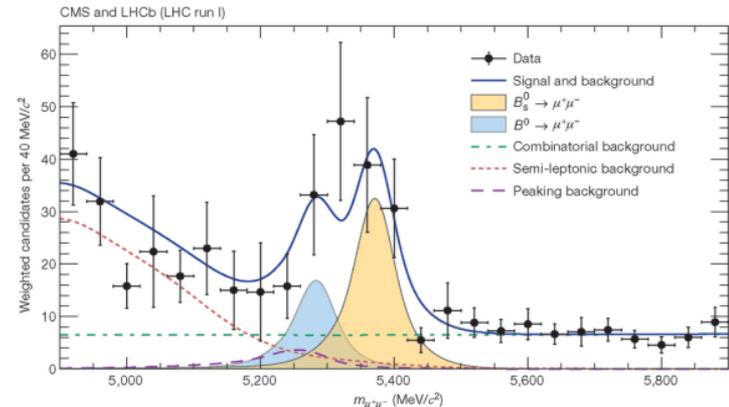
CMS:

$$\mathcal{B}(B_s^0 \rightarrow \mu^+ \mu^-) = (3.0_{-0.9}^{+1.0}) \times 10^{-9}$$



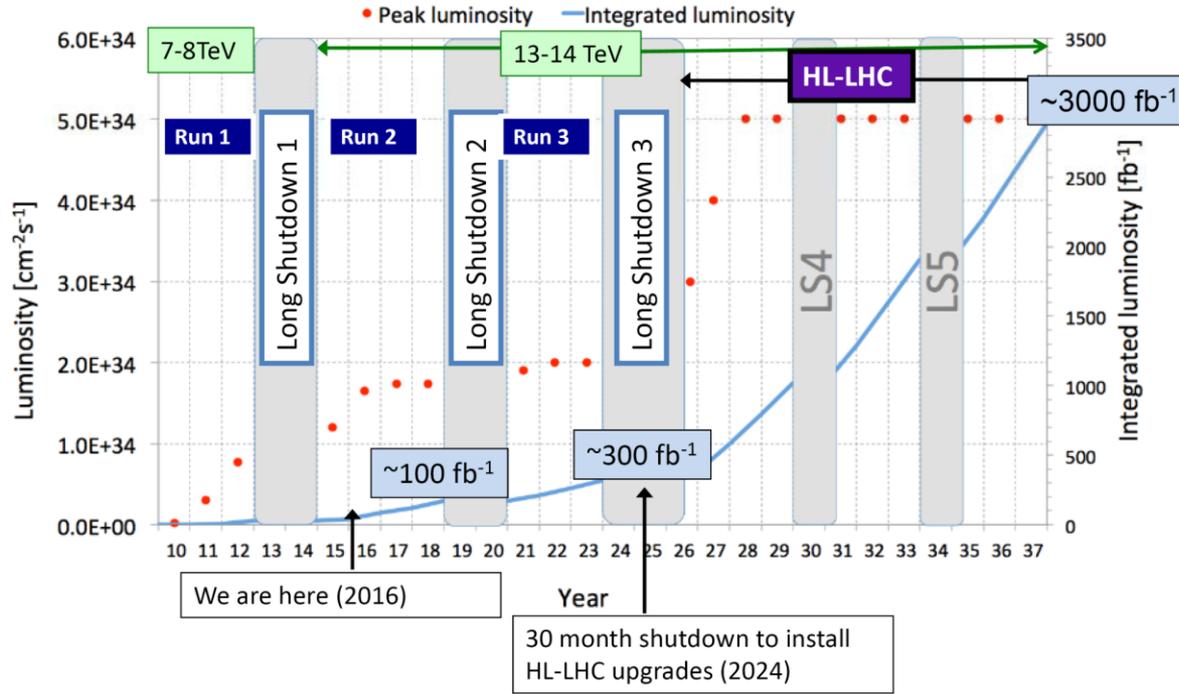
Combined result (**Nature 522 (2015) 68**)

$$Br(B_s^0 \rightarrow \mu^+ \mu^-)_{\text{exp}} = (2.8_{-0.8}^{+0.7}) \times 10^{-9}$$



Observed decay rate is compatible with the SM expectation:

$$Br(B_s^0 \rightarrow \mu^+ \mu^-)_{\text{theory}} = (3.66 \pm 0.23) \times 10^{-9}$$



International conceptual design study ~100 km ring:

- ❑ pp collider (FCC-hh)

$\sqrt{s} \sim 100 \text{ TeV}, L \sim 2 \times 10^{35}$

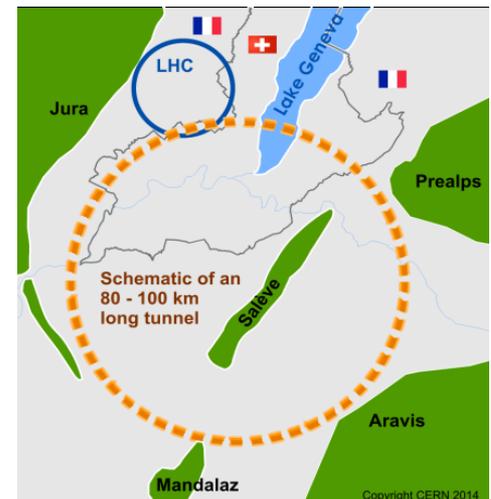
- ❑ e^+e^- collider (FCC-ee)

$\sqrt{s} = 90\text{-}350 \text{ GeV}, L \sim 200\text{-}2 \times 10^{34} \quad -1$

- ❑ pe collider (FCC-he): option

$\sqrt{s} \sim 3.5 \text{ TeV}, L \sim 10^{34}$

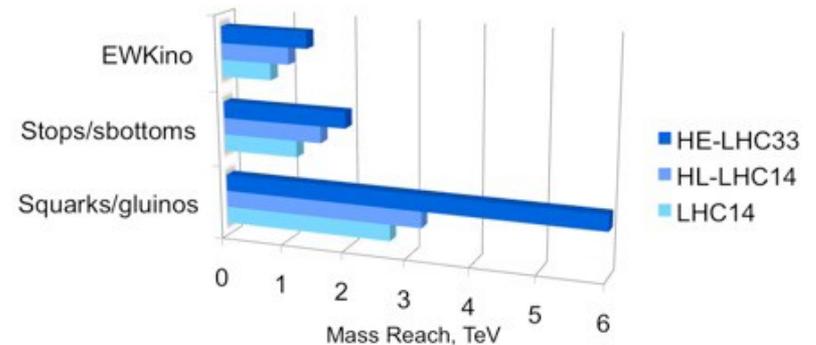
Future Circular Colliders (FCC)

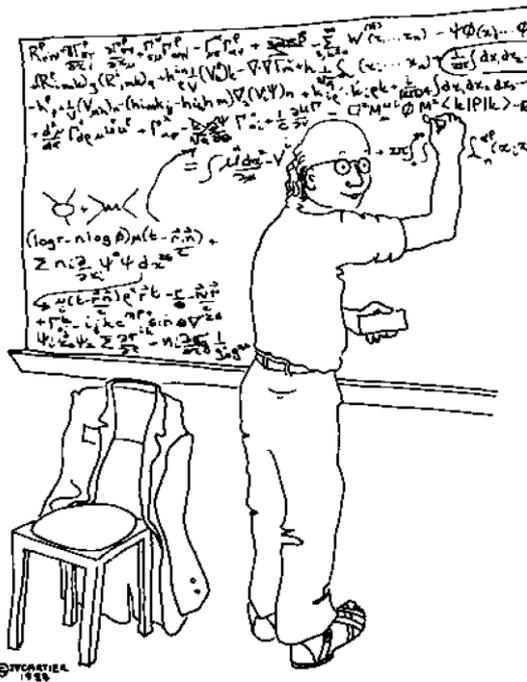


Expected precisions for Higgs couplings

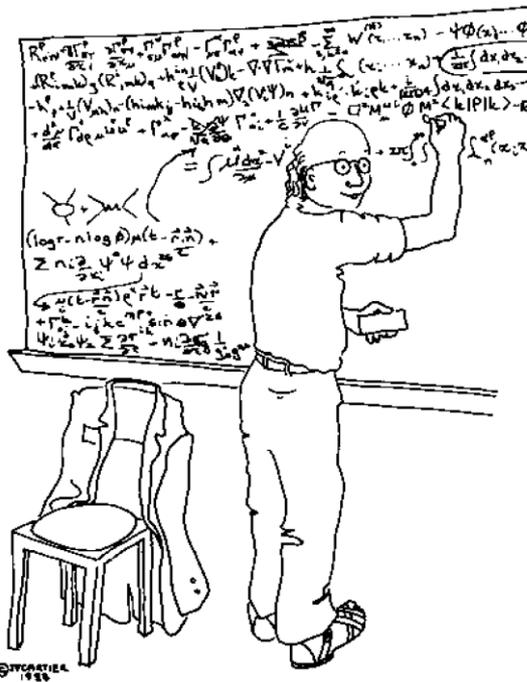
Facility	LHC	HL-LHC	ILC500	ILC500-up	ILC1000	ILC1000-up	CLIC	TLEP (4 IPs)
\sqrt{s} (GeV)	14,000	14,000	250/500	250/500	250/500/1000	250/500/1000	350/1400/3000	240/350
$\int \mathcal{L} dt$ (fb $^{-1}$)	300/expt	3000/expt	250+500	1150+1600	250+500+1000	1150+1600+2500	500+1500+2000	10,000+2600
κ_γ	5 – 7%	2 – 5%	8.3%	4.4%	3.8%	2.3%	–/5.5/<5.5%	1.45%
κ_g	6 – 8%	3 – 5%	2.0%	1.1%	1.1%	0.67%	3.6/0.79/0.56%	0.79%
κ_W	4 – 6%	2 – 5%	0.39%	0.21%	0.21%	0.13%	1.5/0.15/0.11%	0.10%
κ_Z	4 – 6%	2 – 4%	0.49%	0.24%	0.44%	0.22%	0.49/0.33/0.24%	0.05%
κ_ℓ	6 – 8%	2 – 5%	1.9%	0.98%	1.3%	0.72%	3.5/1.4/<1.3%	0.51%
κ_d	10 – 13%	4 – 7%	0.93%	0.51%	0.51%	0.31%	1.7/0.32/0.19%	0.39%
κ_u	14 – 15%	7 – 10%	2.5%	1.3%	1.3%	0.76%	3.1/1.0/0.7%	0.69%

$$\begin{aligned}
 \mathcal{L} = & \kappa_3 \frac{m_H^2}{2v} H^3 + \kappa_Z \frac{m_Z^2}{v} Z_\mu Z^\mu H + \kappa_W \frac{2m_W^2}{v} W_\mu^+ W^{-\mu} H \\
 & + \kappa_g \frac{\alpha_s}{12\pi v} G_{\mu\nu}^a G^{a\mu\nu} H + \kappa_\gamma \frac{\alpha}{2\pi v} A_{\mu\nu} A^{\mu\nu} H + \kappa_{Z\gamma} \frac{\alpha}{\pi v} A_{\mu\nu} Z^{\mu\nu} H \\
 & - \left(\kappa_t \sum_{f=u,c,t} \frac{m_f}{v} f\bar{f} + \kappa_b \sum_{f=d,s,b} \frac{m_f}{v} f\bar{f} + \kappa_\tau \sum_{f=e,\mu,\tau} \frac{m_f}{v} f\bar{f} \right) H
 \end{aligned}$$





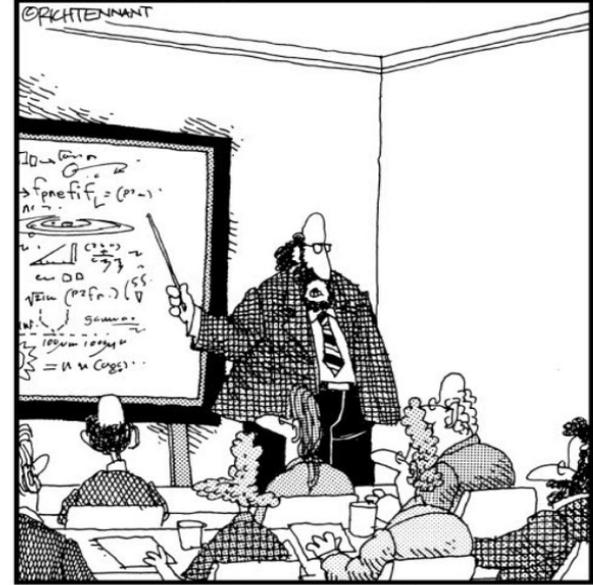
"At this point we notice that this equation is beautifully simplified if we assume that space-time has 92 dimensions."



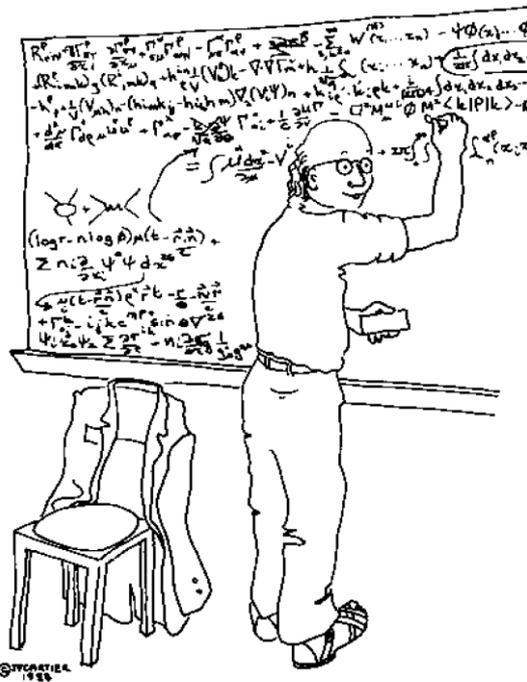
"At this point we notice that this equation is beautifully simplified if we assume that space-time has 92 dimensions."

The 5th Wave

By Rich Tennant



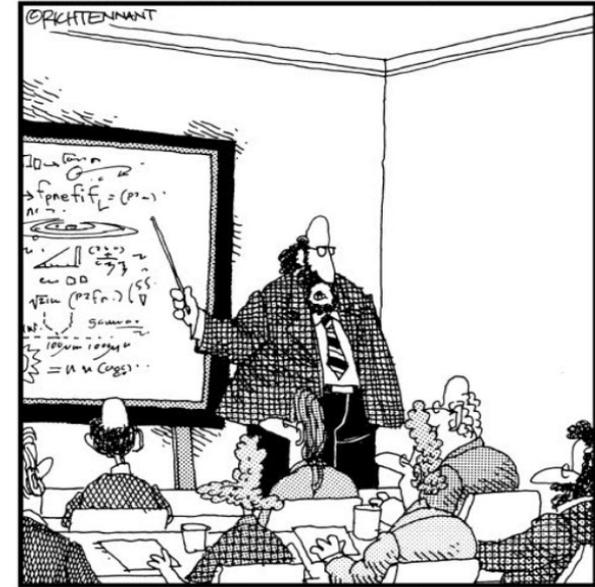
"Along with 'Antimatter,' and 'Dark Matter,' we've recently discovered the existence of 'Doesn't Matter,' which appears to have no effect on the universe whatsoever."



"At this point we notice that this equation is beautifully simplified if we assume that space-time has 92 dimensions."

The 5th Wave

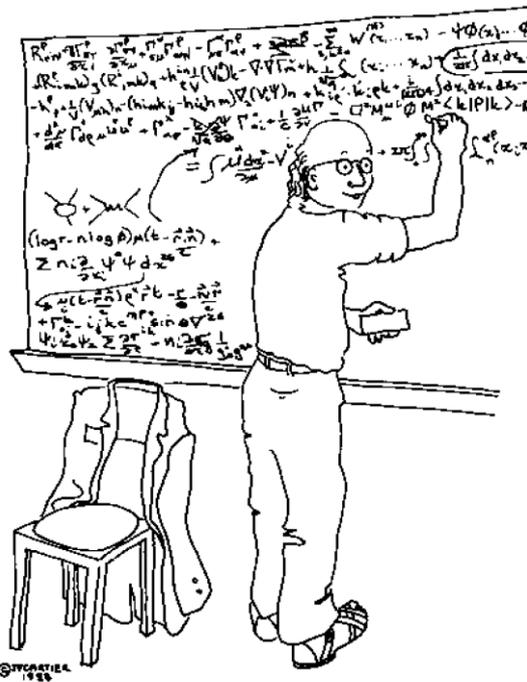
By Rich Tennant



"Along with 'Antimatter,' and 'Dark Matter,' we've recently discovered the existence of 'Doesn't Matter,' which appears to have no effect on the universe whatsoever."



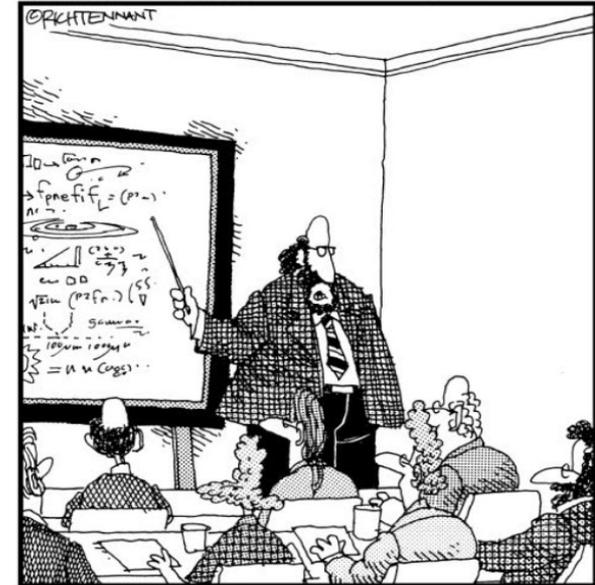
**"It doesn't matter how beautiful your theory is,
it doesn't matter how smart you are.
If it doesn't agree with experiment, it's wrong".
Richard P. Feynman**



"At this point we notice that this equation is beautifully simplified if we assume that space-time has 92 dimensions."

The 5th Wave

By Rich Tennant



"Along with 'Antimatter,' and 'Dark Matter,' we've recently discovered the existence of 'Doesn't Matter,' which appears to have no effect on the universe whatsoever."



**"It doesn't matter how beautiful your theory is,
it doesn't matter how smart you are.
If it doesn't agree with experiment, it's wrong".
Richard P. Feynman**

But with correct (confirmed) experiment...